Steep Cliffs and Saturated Exponents in Three-Dimensional Scalar Turbulence

Kartik P. Iyer,^{1,*} Jörg Schumacher,^{1,2} Katepalli R. Sreenivasan,^{1,3} and P. K. Yeung⁴

¹Tandon School of Engineering, New York University, New York, New York 11201, USA

²Institut für Thermo- und Fluiddynamik, Technische Universität Ilmenau, Postfach 100565, D-98684 Ilmenau, Germany

³Department of Physics and the Courant Institute of Mathematical Sciences, New York, New York 10012, USA

⁴Schools of Aerospace Engineering and Mechanical Engineering, Georgia Institute of Technology, Atlanta, Georgia 30332, USA

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The intermittency of a passive scalar advected by three-dimensional Navier-Stokes turbulence at a Taylor-scale Reynolds number of 650 is studied using direct numerical simulations on a 4096³ grid; the Schmidt number is unity. By measuring scalar increment moments of high orders, while ensuring statistical convergence, we provide unambiguous evidence that the scaling exponents saturate to 1.2 for moment orders beyond about 12, indicating that scalar intermittency is dominated by the most singular shocklike cliffs in the scalar field. We show that the fractal dimension of the spatial support of steep cliffs is about 1.8, whose sum with the saturation exponent value of 1.2 adds up to the space dimension of 3, thus demonstrating a deep connection between the geometry and statistics in turbulent scalar mixing. The anomaly for the fourth and sixth order moments is comparable to that in the Kraichnan model for the roughness exponent of 4/3.

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The mixing of a substance in a complex turbulent flow is a generic and fundamental problem which serves as a paradigm for many processes in nature and technology [1–5]. The basic characteristic of such turbulent systems is intermittency, manifested as intense and sporadic fluctuations of the small scales, which are not captured by classical mean field theories [6,7]. Two important examples of intermittent systems are three-dimensional Navier-Stokes (NS) turbulence [8,9] and 3D scalar turbulence [10,11]—a short phrase for passive scalars mixed by NS turbulence.

Scalar turbulence provides a clear example of a generic feature of nonlinear multiscale phenomena, namely the connection between the multifractal scaling of statistical moments of the physical quantity or field under consideration to the geometric properties of the developing coherent structures. This intimate connection extends beyond fluid mechanics and can be found in various other fields of physics and beyond, such as in fracture mechanics of solids [12], nonlinear fiber optics [13], and bitcoin markets [14]. In all such multifractal statistical processes, quasidiscontinuous features characterized by steep *cliffs* or *fronts* abound, for instance, see Fig. 1. In the context of 3D turbulence, the influence of such cliffs on scalar intermittency has remained an open question.

In the related problems of Burgers turbulence [15–17] and the Kraichnan model [18] for a scalar advected by a synthetic velocity field with no temporal memory, much progress has been made on this particular subject [19–25]. However, the finite-time correlations of the advecting 3D NS turbulence have impeded theoretical progress, with the high spatial and temporal resolution requirements imposing



FIG. 1. Ramp-cliff structures in a scalar field, $\Theta \equiv \theta + Gx$, here θ is the scalar fluctuation and (G, 0, 0) is the mean gradient, at $R_{\lambda} = 650$ and the Schmidt number Sc $\equiv \nu/D = 1$, where ν is the kinematic viscosity of the fluid and *D* is the scalar diffusivity. *L* is the size of the computational cube in one direction. The main figure to the left plots four 1D profiles of Θ in the *x* direction, along which the mean gradient is imposed. Examples for ramps and cliffs are indicated by arrows as is the mean scalar concentration profile by dashed lines. Profiles are shifted in steps of 6 units with respect to each other for clarity. The vertical solid lines indicate the spatial positions for the magnifications of the scalar fluctuation profiles plotted to the right. Grid resolution and Kolmogorov length η_K are indicated.

considerable challenges on empirical work. A large number of experimental and numerical efforts continue to be made on understanding scalar intermittency [26–34], but the connection between the intermittent statistics and the spatial geometry in 3D scalar mixing has eluded a clear demonstration.

In this Letter, we report the precise quantification of small-scale intermittency of a statistically stationary scalar field, advected by 3D isotropic NS turbulence using direct numerical simulations (DNS). We connect the statistical footprints of the well-mixed scalar regions known as ramps [35–41] to the steep cliff regions. The scaling exponents ζ_p^{θ} of the scalar correlations, which will be defined further below, saturate for moment orders above about 12, to a constant ζ_{∞}^{θ} , confirming that the almost-shocklike steep scalar fronts characterize scalar intermittency in 3D NS turbulence. We will also show that the spatial support of the cliffs with a fractal dimension of $D_F = 1.8$ combines with the saturated scaling exponent ζ_{∞}^{θ} to the space dimension, yielding the result $\zeta_{\infty}^{\theta} + D_F = 3$, where 3 is the space dimension, thus demonstrating the intimate link between the geometry and statistics in turbulent passive scalar mixing.

We use data from pseudospectral DNS of isotropic turbulence, computed using 4096³ mesh points in a periodic box of size L. A statistically steady state was obtained by forcing the low Fourier modes of the velocity field [42]. The Taylor-scale Reynolds number $R_{\lambda} = 650$, the Schmidt number Sc = 1, and the Taylor-scale Péclet number $Pe_{\lambda} \equiv R_{\lambda}Sc = 650$. The passive scalar (Θ) is evolved using the diffusion-advection equation in the presence of a uniform mean gradient $\mathbf{G} \equiv (G, 0, 0)$ along the x direction, where $G \neq 0$ is a constant, such that $\Theta = \theta + Gx$, θ here is the scalar fluctuation. The grid resolution $\Delta/\eta_K = 1.1$, Δ being the grid spacing and η_K the Kolmogorov length scale. The ratio of the magnitude of the largest gradient computed in the DNS to the largest gradient possible in the flow $(\theta_{\rm rms}/\eta_K \sim GL/\eta_K \sim R_{\lambda}^{3/2})$ is ~O(1), where $\theta_{\rm rms} \equiv \sqrt{\langle \theta^2 \rangle}$ and $\langle \cdot \rangle$ denotes combined space-time averages; hence the numerical resolution is adequate to resolve the largest scalar gradients. We have used 21 temporal snapshots over 10 eddy turnover times, with each snapshot rotated over 146 angular directions to extract the isotropic statistics of the anisotropic scalar field. In total, we have used 2×10^{12} data samples to obtain the results. For details on the exact laws of the velocity and mixed velocityscalar statistics and statistical convergence tests on the data, see the Supplemental Material [43] and Refs. [44–46].

The scalar signal organizes itself into conspicuous patterns, as shown in Fig. 1, consisting of two distinctive features: (i) ramp regions where the total scalar gradient $\nabla \theta + \mathbf{G}$ is quite small and (ii) high gradient cliffs which are interspersed between ramps. The small figures on the right demonstrate clearly that the scalar increment, $\delta_r \theta \equiv \theta(\mathbf{x} + \mathbf{r}) - \theta(\mathbf{x})$, can jump by the order *GL* over

 $r \equiv |\mathbf{r}|$ that is just a few multiples of the Kolmogorov scale η_K (which is also the smallest dynamically significant scale in the scalar field). The ramp-cliff structures are connected to the mean scalar gradient in the present DNS, and are known to cause the breakdown of local isotropy in the scalar field [10]. The cliffs are caused by the action of large scales in the scalar field, even in the absence of a mean gradient [10,47]. The generic existence of scalar cliffs in turbulence suggests that these local spatial barriers to scalar mixing have a significant impact on scalar intermittency.

In order to assess scalar intermittency, we define the *pth* order scalar structure function, $S^p_{\theta}(\mathbf{r}) \equiv \langle (\delta_r \theta)^p \rangle$. Because of the anisotropic mean scalar gradient, $S^p_{\theta}(\mathbf{r})$ depends on the separation vector **r**; however, the isotropic sector $\langle (\delta_r \theta)^p \rangle_0$ extracted from the SO(3) decomposition [48,49] of $S^p_{\theta}(\mathbf{r})$ depends solely on the separation distance r[50]. In the inertial range, $\eta_K \ll r \ll \ell$, where ℓ is the macroscale, $30 \le r/\eta_K \le 300$, $\langle (\delta_r \theta)^p \rangle_0$ follow power laws, $\langle (\delta_r \theta)^p \rangle_0 \sim r^{\zeta_p^{\theta}}$, where ζ_p^{θ} denote the *p*th order exponents (see Supplemental Material for details [43]). The higherorder exponents are determined using extended selfsimilarity [51], by plotting $\langle (\delta_r \theta)^p \rangle_0$ against $\langle (\delta_r \theta)^2 \rangle_0$ for p > 2. We have verified that estimating ζ_p^{θ} using local slopes, e.g., Ref. [34], or compensated structure functions, e.g., Ref. [29], yields results consistent with the extended self-similarity results.

The scaling exponents ζ_p^{θ} are plotted against moment order p at $\text{Pe}_{\lambda} = 650$ in Fig. 2(a). The exponents saturate to $\zeta_{\infty}^{\theta} = 1.2$, indicated by the horizontal line, for $p \ge 12$. This is the clearest indication that the scalar fluctuations are limited in magnitude only by the largest allowable gradients in the field (largest temperature difference divided by the smallest length scale). The $\zeta_{\infty}^{\theta} = 1.2$ curve intersects the normal scaling curve at p = 3.6. In some sense, it is possible that this represents the situation for infinitely large Pe_{λ} . Figure 2(b) compares the present exponents with previous, lower Pe_{λ} , results. While our data robustly confirm that the exponents saturate, it is hard to reach a similar unambiguous conclusion from the previous results in the literature [29–32,34,52,53].

The statistical convergence of the moments of order p up to 20 was confirmed by the rapid decay of the moment integrands, $(\delta_r \theta)^p \mathcal{P}(\delta_r \theta)$, where \mathcal{P} denotes the probability density function (PDF). The integrands of moment orders 6 and 16 are shown in Fig. 3, each for r in the low end of the inertial range. The integrands peak before the tail contributions decay, ensuring statistical convergence of the moments. Saturation of exponents at higher orders implies that, for scalar jumps $|\delta_r \theta| \gtrsim \theta_{\rm rms}$, $\mathcal{P}(\delta_r \theta) \propto r^{\zeta_{\infty}^{\theta}}$ [23,54]. Figure 4 verifies that this is indeed the case, with $\mathcal{P}(\delta_r \theta) r^{-\zeta_{\infty}^{\theta}}$ collapsing for $|\delta_r \theta| \ge 3\theta_{\rm rms}$, for all inertial separations. The inference is that the saturation of exponents arises because of the dominance of the high-order moments by features that do not change with scale, suggesting that the gradients are of the order $\theta_{\rm rms}/\eta_K$.



FIG. 2. Scalar increment exponent ζ_p^{θ} versus moment order *p*. (a) Present DNS: $Pe_{\lambda} = 650$, dashed line at saturation exponent, $\zeta_{\infty}^{\theta} = 1.2$. Error bars indicate 95% confidence interval. (b) Comparison of present DNS (shaded region) with previous results: $(\nabla) Pe_{\lambda} = 220$ [52], (\bigcirc) $Pe_{\lambda} = 280$ [29], (\triangle) $Pe_{\lambda} = 396$ [32], $(\diamondsuit) Pe_{\lambda} = 580$ [34]. Dash-dotted line shows normal scaling $\zeta_p^{\theta} = p/3$ and dashed line is the model of Ref. [28]. A more extensive data comparison can be found in [43].

We now turn to quantifying the dimension of the spatial support of the cliffs where strong scalar gradients tend to be concentrated in sharp fronts (Fig. 1). The dimension of such fronts is estimated by the spatial support of regions of the strongest gradients of $\mathcal{O}(\theta_{\rm rms}/\eta_K)$ with cubes of edge size r, and counting their respective number N(r) for different r. As shown in the inset of Fig. 5, gradients greater than 20% of $\theta_{\rm rms}/\eta_K$ (marked by dotted lines) corresponding to $5\sqrt{\langle (\partial \theta / \partial x)^2 \rangle}$ are used to determine N(r). We chose the threshold of 20% as a good representative of gradients of the order $\theta_{\rm rms}/\eta_K$ occurring with low probability (see inset of Fig. 5). (The use of a somewhat different threshold alters the scaling range in Fig. 5 but does not alter the dimension itself.) The plot of N(r) versus r for such fronts shown in the main body of Fig. 5 is compensated by $r^{1.8}$ (see below for the rationale), and has three scaling regimes: (i) at the smallest scales, a slope of -2, which corresponds to flat fronts; (ii) at $r/\eta_K \in [4, 30]$, for which the slope from the least-squares fit is $D_F = 1.79 \pm 0.01$, corresponding to the spatial subset that supports the steep fronts in the scalar field; (iii) at the largest scales, the slope is -3, which corresponds to the Euclidean dimension of the flow. We confirm, for the first time in 3D NS flows, that the



FIG. 3. Integrands of scalar increment moments $[\mathcal{P}(\cdot)]$ denotes PDF of (\cdot)] as functions of scalar increments, for orders 6 and 16 at $r/\eta_K = 55$ (lower end of the inertial range), on lin-log scales; moments of orders up to 20 converge as well and confirm the saturation of exponents but are not shown here. The integrands are normalized by respective moments such that the area under each curve is unity.

saturation exponent ζ_{∞}^{θ} and the box counting dimension of the steep fronts D_F are related to the space dimension, d = 3, as

$$\zeta^{\theta}_{\infty} + D_F = d. \tag{1}$$

The confirmation of this relation in Navier-Stokes turbulence is remarkable since it directly connects a property of the highly intermittent statistics of the scalar to the spatial geometry of mixing barriers in the flow. The geometrical features of the scalar cliffs shown here have interesting parallels beyond fluid turbulence; e.g., in fracture processes in solids, where a network of steep cliffs with a fractal dimension $D_F \approx 1.7$ has been detected on a fracture surface and related to the multifractal spectrum of height fluctuations [12].



FIG. 4. PDF of scalar increments across the inertial range, multiplied by $r^{-\zeta_{\infty}^{\theta}}$, where ζ_{∞}^{θ} is the saturation exponent (Fig. 2). The PDF tails collapse, confirming saturation of exponents.



FIG. 5. Log-log plot of the number N(r) of cubes of side r containing the steepest fronts versus size r. The ordinate is compensated by r^{D_F} , where $D_F = 1.8$ is the fractal codimension of the fronts. The plateau region (filled circle) corresponds to the scaling r^{-D_F} , indicated by the horizontal dashed line. At the smallest and largest r, the dimension of the fronts is 2 (corresponding to flat fronts) and 3 (which is the Euclidean dimension of the flow), respectively. Inset shows the PDF of the normalized gradient z, |z| > 0.2 (dotted lines) is used to calculate N(r).

The anomaly in the passive scalar field advected by a 3D NS flow at high R_{λ} is comparable for orders 4 and 6 to that advected by the δ -correlated 3D Kraichnan model [55,56], as seen in Fig. 6 where we quantify the degree of anomaly of exponents against the roughness parameter of the flow (which varies between 0 and 2 for the Kraichnan model and is 4/3 for NS turbulence). This observed agreement is plausible because, in a high- R_{λ} NS flow, the small scales evolve with temporal rapidity, and the normalized exponents, $\psi_p^{\theta} \equiv (p/2)\zeta_2^{\theta} - \zeta_p^{\theta}$, approach the Kraichnan limit of



FIG. 6. Flatness anomaly $(\psi_4^{\theta} = 2\zeta_2^{\theta} - \zeta_4^{\theta})$ (open symbols) and hyperflatness anomaly $(\psi_6^{\theta} = 3\zeta_2^{\theta} - \zeta_6^{\theta})$ (closed symbols) for the scalar versus flow roughness ξ . Circles correspond to 3D Kraichnan model, (()) [55] and (filled circle) [56], while triangles are for 3D NS flow at Pe_{λ} = 650, with the roughness parameter $\xi = 4/3$.

a flow without memory. Scalar exponents for the Kraichnan model saturate at different values for different roughness parameters [23,54], and the observed correspondence with the 3D NS results may not hold for high-order moments.

Our conclusive result here is that in a scalar field advected by 3D NS turbulence, the exponents ζ_p^{θ} saturate to ζ^{θ}_{∞} at large orders and that the saturation exponent is connected to the fractal dimension of the steep fronts. We do not expect ζ_{∞}^{θ} to be universal, since ζ_{p}^{θ} itself is nonuniversal [34,57–59], but the fact that scalar exponents saturate in 3D NS flows can have important consequences. For instance, the minimum Hölder exponent of θ , $h_{\min}^{\theta} := \lim_{p \to \infty} \zeta_p^{\theta} / p = \lim_{p \to \infty} \zeta_{\infty}^{\theta} / p = 0$, implies that shocklike quasidiscontinuities, or steep fronts, characterize the large gradients of the scalar field, reminiscent of 1D Burgers flow. However, while the Burgers flow displays a biscaling behavior, the lower-order scalar exponents appear to have a quadratic dependence on the order, similar to that derived for scalar advection in the high-dimensional Kraichnan model [22]. This work sets the stage for similar investigations in the low (Sc \ll 1) and high (Sc \gg 1) Schmidt number regimes, which have important physical applications [60-63]. We conjecture that the strong diffusion in the $Sc \ll 1$ limit may prevent such a saturation, whereas in the $Sc \gg 1$ case, the weak diffusion may enhance a saturation to the 1D Burgers limit. A careful analysis of the link between geometry and statistics in these two regimes is ongoing and will be reported as future work.

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^{*}kartik.iyer@nyu.edu

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