Reduced Hadronic Uncertainty in the Determination of V_{ud}

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We analyze the universal radiative correction Δ_R^V to neutron and superallowed nuclear β decay by expressing the hadronic γW -box contribution in terms of a dispersion relation, which we identify as an integral over the first Nachtmann moment of the γW interference structure function $F_3^{(0)}$. By connecting the needed input to existing data on neutrino and antineutrino scattering, we obtain an updated value of $\Delta_R^V = 0.02467(22)$, wherein the hadronic uncertainty is reduced. Assuming other standard model theoretical calculations and experimental measurements remain unchanged, we obtain an updated value of $|V_{ud}| = 0.97370(14)$, raising tension with the first row Cabibbo-Kobayashi-Maskawa unitarity constraint. We comment on ways current and future experiments can provide input to our dispersive analysis.

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The unitarity test of the Cabibbo-Kobayashi-Maskawa (CKM) matrix serves as one of the most important precision tests of the standard model. In particular, tests of first row CKM unitarity $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$ receive the most attention since these matrix elements are known with the highest precision, all with comparable uncertainties. The good agreement with unitarity [1] serves as a powerful tool to constrain new physics scenarios.

Currently, the most precise determination of $|V_{ud}|$ comes from measurements of half-lives of superallowed $0^+ \rightarrow 0^+$ nuclear β decays with a precision of 10^{-4} [2]. At tree level, these decays are mediated by the vector part of the weak charged current only, which is protected against renormalization by strong interactions due to a conserved vector current (CVC), making the extraction of $|V_{ud}|$ relatively clean. Beyond tree level, however, electroweak radiative corrections (EWRC) involving the axial current are not protected, and lead to a hadronic uncertainty that dominates the error in the determination of $|V_{ud}|$.

The master formula relating the CKM matrix element $|V_{ud}|$ to the superallowed nuclear β decay half-life is as follows [2]:

$$|V_{ud}|^2 = \frac{2984.432(3) \text{ s}}{\mathcal{F}t(1+\Delta_R^V)},\tag{1}$$

where the nucleus-independent $\mathcal{F}t$ value is obtained from the experimentally measured ft value by absorbing all nuclear-dependent corrections, and where Δ_R^V represents the nucleus-independent EWRC. Currently, an average of the 14 best measured half-lives yields an extraordinarily precise value of $\mathcal{F}t = 3072.07(63)$ s [3]. A similar master formula exists for free neutron β decay [4] depending additionally on the axial-to-vector nucleon coupling ratio $\lambda = g_A/g_V$, and is free of nuclear-structure uncertainties. But the much larger experimental errors in the measurement of its lifetime and the ratio λ [5] makes it less competitive in the extraction of $|V_{ud}|$. Regardless, if first row CKM unitarity is to be tested at a higher level of precision, improvement in the theoretical estimate of Δ_R^V by reducing hadronic uncertainties is essential.

The best determination of $\Delta_R^V = 0.02361(38)$ was obtained in 2006 by Marciano and Sirlin [6] (in the following, we refer to their work as [MS]). They were able to reduce the hadronic uncertainty by a factor of 2 over their earlier calculation [7] by using high order perturbative QCD corrections originally derived for the polarized Bjorken sum rule to precisely estimate the short distance contribution. At intermediate distances, an interpolating function motivated by vector meson dominance (VMD) was used to connect the long and short distances and was identified as the dominant source of theoretical uncertainty. This result leads to the current value of $|V_{ud}| = 0.97420(21)$ [1].

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In this Letter, we introduce a new approach for evaluating Δ_R^V based on dispersion relations that relate it to directly measurable inclusive lepton-hadron and neutrinohadron structure functions. Dispersion relations have proved crucial for evaluating the γZ -box correction to parity violating electron-hadron interaction in atoms and in scattering processes [8–20]. It led to a significant shift in the 1-loop SM prediction for the hadronic weak charges, and ensured a correct extraction of the weak mixing angle at low energy [21]. Using existing data on neutrino and antineutrino scattering, we obtain a more precise value of the nucleus-independent EWRC,

$$\Delta_R^V = 0.02467(22),\tag{2}$$

and therefore a new determination of $|V_{ud}|$,

$$|V_{ud}| = 0.97370(14). \tag{3}$$

We summarize in this Letter the essential features of our analysis that lead us to these values, and defer details to an upcoming longer paper [22].

Among the various contributions at $O(\alpha/\pi)$ to the neutron β decay amplitude, Sirlin established [23] that the only one sensitive to the hadronic scale is the part in the γW box amplitude (Fig. 1),

$$\mathcal{M}_{VA} = 2\sqrt{2}e^{2}G_{F}V_{ud}\int \frac{d^{4}q}{(2\pi)^{4}} \times \left(\frac{\bar{u}_{e}(\mathbf{k})\gamma_{\mu}(\not{k}-\not{q}+m_{e})\gamma_{\nu}P_{L}v_{\nu}(\mathbf{k})}{q^{2}[(k-q)^{2}-m_{e}^{2}]}\frac{M_{W}^{2}}{q^{2}-M_{W}^{2}}T_{VA}^{\mu\nu}\right),$$
(4)

involving the nucleon matrix element of the product of the electromagnetic (EM) and the axial part of the weak charged current

$$T_{VA}^{\mu\nu} = \frac{1}{2} \int d^4x e^{iqx} \langle p(\mathbf{p}) | T[J_{\rm em}^{\mu}(x) J_{W,A}^{\nu}(0)] | n(\mathbf{p}) \rangle.$$
(5)

After inserting the nucleon matrix element parametrized in terms of the *P*-odd invariant function $T_{VA}^{\mu\nu} = (i\epsilon^{\mu\nu\alpha\beta}p_{\alpha}q_{\beta}/2pq)T_3$ into the amplitude [Eq. (4)], the correction to the tree level amplitude is expressed as [23]

$$\Box_{\gamma W}^{VA} = \frac{\alpha}{8\pi} \int_0^\infty dQ^2 \frac{M_W^2}{M_W^2 + Q^2} \\ \times \int_{-i\sqrt{Q^2}}^{i\sqrt{Q^2}} \frac{d\nu}{\nu} \frac{4(Q^2 + \nu^2)^{3/2}}{\pi M Q^4} T_3(\nu, Q^2) \quad (6)$$

where, after Wick rotation, the azimuthal angles of the loop momentum have been integrated over and the remaining integrals have been expressed in terms of $Q^2 = -q^2$ and $\nu = (pq)/M$. With negligible error, we assume a common nucleon mass M in the isospin symmetric limit and we



FIG. 1. Feynman diagrams corresponding to the amplitude in Eq. (4), which contribute at order $\mathcal{O}(\alpha/\pi)$ to neutron β decay and are sensitive to the hadronic scale.

work in the recoil-free approximation. This contributes to the nucleus-independent EWRC as

$$\Delta_R^V = 2 \Box_{\gamma W}^{VA} +, \dots, \tag{7}$$

where the ellipsis denotes all other corrections insensitive to the hadronic scale.

Marciano and Sirlin estimate $\Box_{\gamma W}^{VA}$ by phenomenologically treating the ν integral $F_{M.S.}(Q^2) \equiv \int d\nu$, ..., in the second line of Eq. (6) as a function of Q^2 , and parametrizing it piecewise over three domains: in the short distance domain $Q^2 > (1.5 \text{ GeV})^2$, the leading term in the OPE corrected by high order perturbative QCD is used; in the long distance domain $Q^2 < (0.823 \text{ GeV})^2$, the elastic nucleon with dipole form factors is used with a 10% uncertainty; and at intermediate scales $(0.823 \text{ GeV})^2 < Q^2 < (1.5 \text{ GeV})^2$, an interpolating function inspired by VMD is used and is assigned a generous 100% uncertainty. Performing the integration over Q^2 in Eq. (6) yields their value of Δ_R^V quoted above.

Our evaluation of $\Box_{\gamma W}^{VA}$ begins by first separating the invariant amplitude T_3 with respect to isosinglet and isotriplet components of the EM current $T_3 = T_3^{(0)} + T_3^{(3)}$. Crossing symmetry implies $T_3^{(0)}$ is odd under $\nu \to -\nu$ while $T_3^{(3)}$ is even. Since the ν integration measure in Eq. (6) is odd, only $T_3^{(0)}$ contributes to $\Box_{\gamma W}^{VA}$. We then write a dispersion relation in ν for $T_3^{(0)}$, taking into account the physical sheet singularities. Poles at $\nu_B = \pm Q^2/(2M)$ correspond to the elastic single-nucleon intermediate state, and branch points at $\nu_{\pi} = \pm (m_{\pi}^2 + 2Mm_{\pi} + Q^2)/(2M)$ correspond to single pion production thresholds. We identify the discontinuity of $T_3^{(0)}$ across the cut as the γW -interference structure function, $4\pi F_3^{(0)}(\nu, Q^2) =$ $T_3^{(0)}(\nu + i\epsilon, Q^2) - T_3^{(0)}(\nu - i\epsilon, Q^2)$, so that the dispersion relation reads

$$T_3^{(0)}(\nu, Q^2) = -4i\nu \int_0^\infty d\nu' \frac{F_3^{(0)}(\nu', Q^2)}{\nu'^2 - \nu^2}.$$
 (8)

where $F_3^{(0)}$ contains both the elastic and inelastic contributions. No subtraction constant appears since $T_3^{(0)}$ is an odd function of ν . Only I = 1/2 intermediate states contribute because the EM current is isoscalar. After

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inserting Eq. (8) into Eq. (6), performing the ν integration, and changing the integration variable $\nu' \rightarrow Q^2/(2Mx)$ we obtain

$$\Box_{\gamma W}^{VA} = \frac{3\alpha}{2\pi} \int_0^\infty \frac{dQ^2}{Q^2} \frac{M_W^2}{M_W^2 + Q^2} M_3^{(0)}(1, Q^2), \qquad (9)$$

where $M_3^{(0)}(1, Q^2)$ is the first Nachtmann moment of the structure function $F_3^{(0)}$ [24,25]

$$M_3^{(0)}(1,Q^2) = \frac{4}{3} \int_0^1 dx \frac{1+2r}{(1+r)^2} F_3^{(0)}(x,Q^2), \qquad (10)$$

and $r = \sqrt{1 + 4M^2 x^2/Q^2}$. To estimate $\Box_{\gamma W}^{VA}$, we require the functional form of $F_3^{(0)}$ depending on x and Q^2 , or equivalently, $W^2 = M^2 + (1-x)Q^2/x$ and Q^2 .

We draw attention to the fact that Eq. (9) relates [MS]'s phenomenological function to the first Nachtmann moment

$$F_{\rm MS}(Q^2) = \frac{12}{Q^2} M_3^{(0)}(1,Q^2), \qquad (11)$$

which will prove useful when comparing their results with ours. Furthermore, since $F_3^{(0)}$ depends directly on on-shell intermediate hadronic states, it provides a better handle on the physics that may enter at various scales. Figure 2 depicts the domain in the $W^2 - Q^2$ plane over which $F_3^{(0)}$ has support: the single-nucleon elastic pole is at $W^2 = M^2$, and the inelastic continuum covers the region above $W^2 > (M + m_{\pi})^2$. Our parametrization of $F_3^{(0)}$ is as follows:

$$F_{3}^{(0)} = F_{\text{Born}} + \begin{cases} F_{\text{pQCD}}, & Q^{2} \gtrsim 2 \text{ GeV}^{2} \\ F_{\pi N} + F_{\text{res}} + F_{\mathbb{R}}, & Q^{2} \lesssim 2 \text{ GeV}^{2}, \end{cases}$$
(12)

where each component is given by

$$F_{\rm Born} = -\frac{1}{4} (G_M^p + G_M^n) G_A \delta(1 - x)$$
(13)

$$\int_{0}^{1} dx F_{pQCD} = \frac{1}{12} [1 + pQCD]$$
(14)



FIG. 2. Phase space of the structure functions $F_3^{(0)}$ and $F_3^{\nu p + \bar{\nu} p}$ in the $W^2 - Q^2$ plane.

$$F_{\pi N} = F_{\chi \text{PT}} \times (F_1^p + F_1^n) \frac{|G_A|}{g_A}$$
(15)

$$F_{\rm res} = {\rm negligible}$$
 (16)

$$F_{\mathbb{R}} = C_{\gamma W} f_{\text{th}} \frac{m_{\omega}^2}{m_{\omega}^2 + Q^2} \frac{m_{a_1}^2}{m_{a_1}^2 + Q^2} \left(\frac{\nu}{\nu_0}\right)^{\alpha_0^{\rho}}, \qquad (17)$$

and supplies the dominant contribution to $F_3^{(0)}$ in various regions indicated in Fig. 2, which we describe next.

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We obtain the elastic Born contribution at $W^2 = M^2$ in Eq. (13) by using the updated values of the magnetic Sachs form factor G_M and the axial form factor G_A for the nucleon [26,27]. Above the threshold, $W^2 \ge (M + m_\pi)^2$, we consider the dominant physics operating in various of domains in the $Q^2 - W^2$ plane separately. At large $Q^2 \gtrsim 2 \text{ GeV}^2$, the Nachtmann moment $M_3^{(0)}$ reduces to the Mellin moment and is fixed by the sum rule corrected by pQCD in Eq. (14) by an analogy with that of the polarized Bjorken sum rule [MS]. At small $Q^2 \lesssim 2 \text{ GeV}^2$, we estimate the contribution [Eq. (15)] near the inelastic threshold by computing the single pion production contribution $F_{\gamma PT}$ in chiral perturbation theory (χ PT) at a leading order. To improve the behavior of $F_{\chi PT}$ at a larger Q^2 , we replace the pointlike nucleon vertices with measured Dirac and axial nucleon form factors, F_1 and G_A . At higher W^2 , we investigated the impact of several low-lying I = 1/2 resonances based on a few models [28–30], and found their contributions to \Box_{vW}^{VA} to be negligible. Note that Δ resonances do not contribute since only isoscalar electromagnetic transitions enter $F_3^{(0)}$.

Finally, at a large W^2 , we use the form in Eq. (17) inspired by Regge phenomenology together with VMD [31] as illustrated in Fig. 3(a). In this picture, the Regge behavior $(\nu/
u_0)^{lpha_0^{
ho}}$ arises from the exchange of the hotrajectory with an intercept $\alpha_0^{\rho} = 0.477$ [32], and is coupled to the external currents via a_1 and ω mesons encoded by the VMD factors $m_V^2/(m_V^2+Q^2)$. We include a threshold function $f_{\text{th}} = \Theta(W^2 - W^2_{\text{th}})(1 - \exp[(W^2_{\text{th}} - W^2)/\Lambda^2_{\text{th}}])$, which smoothly vanishes at the two-pion threshold point $W_{\rm th}^2 = (M + 2m_\pi)^2$ to model the smooth background in the resonance region [11]. We choose equal values for the Regge and threshold scales of $\nu_0 = \Lambda_{th} = 1$ GeV to ensure that Regge behavior sets in around $W^2 \sim$ $(2.5 \text{ GeV})^2$. The function $C_{\chi W}(Q^2)$ accounts for residual



FIG. 3. Regge exchange model (a) for $F_3^{(0)}$ and (b) for $F_3^{\nu p + \bar{\nu} p}$ using vector meson dominance.

 Q^2 dependence beyond that of the VMD, which we infer from experimental data as explained below.

Since the isospin structure of $F_3^{(0)}$ is $(I = 0) \times (I = 1)$, it is not directly accessible experimentally. However, information about the *P*-odd structure function with a different isospin structure $(I = 1) \times (I = 1)$ is available from ν and $\bar{\nu}$ scattering. In particular, data exist on the first Nachtmann moment $M_3^{\nu p + \bar{\nu}p}$ for the combination $F_3^{\nu p + \bar{\nu}p} =$ $(F_3^{W^-} + F_3^{W^+})/2$ derived from the difference of νp and $\bar{\nu} p$ differential cross sections. The data by the CCFR [33,34], BEBC and Gargamelle [35], and WA25 [36] collaborations cover a wide region of Q^2 from 0.15 to 600 GeV² (see Fig. 4). Although the precision below $Q^2 \approx 1.4$ GeV² is less satisfactory, we are able to use it to collect information about the form of the analogous Regge coefficient function $C_{WW}(Q^2)$ for this structure function, and thereby infer the form of the required $C_{\gamma W}(Q^2)$ as follows.

We parametrize the structure function $F_3^{\nu p+\bar{\nu}p}$ in precisely the same way as in Eq. (12) for $F_3^{(0)}$, and establish $F_{\text{Born}}^{\nu p+\bar{\nu}p}$, $F_{p\text{QCD}}^{\nu p+\bar{\nu}p}$, $F_{\text{res}}^{\nu p+\bar{\nu}p}$, and $F_{\mathbb{R}}^{\nu p+\bar{\nu}p}$ along similar lines. In this case, $\int_0^1 dx F_{p\text{QCD}}^{\nu p+\bar{\nu}p}$ satisfies the Gross-Llewellyn-Smith sum rule [37] corrected by pQCD [38], while at low Q^2 , the Δ resonance and the Born contribution saturate the Nachtmann moment [35]. At a large W^2 , the ω trajectory controls the leading behavior, and couples to the external currents by the a_1 and ρ mesons [see Fig. 3(b)], leading to

$$F_{\mathbb{R}}^{\nu p + \bar{\nu} p} = C_{WW} f_{\text{th}} \frac{m_{\rho}^2}{m_{\rho}^2 + Q^2} \frac{m_{a_1}^2}{m_{a_1}^2 + Q^2} \left(\frac{\nu}{\nu_0}\right)^{\alpha_0^{\omega}}.$$
 (18)

We then fit the unknown function $C_{WW}(Q^2)$ to the data for $M_3^{\nu p + \bar{\nu}p}(1, Q^2)$ in the range $Q^2 \leq 2$ GeV². Due to the quality of the data, we choose the simple linear form



FIG. 4. World data of the first Nachtmann moment $M_3^{\nu p + \bar{\nu} p}(1, Q^2)$. The red curve is the pQCD-corrected GLS sum rule above $Q^2 \approx 2 \text{ GeV}^2$, and the blue curve is the result of the fit for A_{WW} and B_{WW} in Eq. (19).

$$C_{WW}(Q^2) = A_{WW}(1 + B_{WW}Q^2)$$
(19)

and obtain $A_{WW} = 5.2 \pm 1.5$ and $B_{WW} = 1.08^{+0.48}_{-0.28}$ GeV⁻². The result of the fit is shown by the blue curve in Fig. 4. The solid curve corresponds to the central value of the fit, and the dotted curve indicates the maximum variation in $M_3^{\nu p + \bar{\nu} p}$ allowed by the errors in the fit. We do not fit the three data points below $Q^2 = 0.1$ GeV² where Born and resonance contributions dominate the GLS sum rule: rather, we use the resonance parameters obtained in Ref. [28] from a fit to modern neutrino data.

Finally, to obtain $C_{\gamma W}(Q^2)$, we require the ratio of Nachtmann moments $M_{3,\mathbb{R}}^{(0)}(1,Q^2)/M_{3,\mathbb{R}}^{\nu p+\bar{\nu}p}(1,Q^2)$ to agree with the value predicted by VMD at $Q^2 = 0$, and the QCDcorrected parton model at $Q^2 = 2$ GeV². Since the ρ and ω Regge trajectories are nearly degenerate [32], the two conditions predict the same ratio [22]

$$\frac{M_{3,\mathbb{R}}^{(0)}(1,0)}{M_{3,\mathbb{R}}^{\nu p+\bar{\nu}p}(1,0)} \approx \frac{M_{3,\mathbb{R}}^{(0)}(1,2 \text{ GeV}^2)}{M_{3,\mathbb{R}}^{\nu p+\bar{\nu}p}(1,2 \text{ GeV}^2)} \approx \frac{1}{36}.$$
 (20)

For the linear parametrization in Eq. (19), this implies

$$C_{\gamma W}(Q^2) = \frac{1}{36} C_{WW}(Q^2),$$
 (21)

providing us with the final piece of $F_{\mathbb{R}}$ in Eq. (17).

Upon inserting our parameterization [Eq. (12)] for the structure function $F_3^{(0)}$ into Eqs. (9) and (10) and performing the integrations, we obtain the following contributions to $\Box_{\gamma W}^{VA}$ in units of 10^{-3} : 2.17(0) from parton + pQCD, 1.06(6) from Born, and 0.56(8) from Regge + resonance + πN , the digit in parentheses indicating the uncertainty. Combining them with the remaining known contributions [MS] gives our new values, $\Delta_R^V = 0.02467(22)$ and $|V_{ud}| = 0.97370(14)$. Our reevaluation of Δ_R^V represents a reduction in theoretical uncertainty over the previous [MS] result by nearly a factor of 2. However, it also leads to a substantial upward shift in the central value of Δ_R^V and a corresponding downward shift of $|V_{ud}|$ by nearly three times their quoted error, now raising tension with the first-row CKM unitarity constraint: $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9984(4)$. We pause to comment on the origin of the large shift in

We pause to comment on the origin of the large shift in the central value for Δ_{X}^{P} with respect to [MS]. In Fig. 5 we plot the integrand $M_{3}^{(0)}(1, Q^{2})[1 + Q^{2}/M_{W}^{2}]^{-1}$ of Eq. (9) as a function of Q^{2} . In solid blue, we show the result of our parametrization [Eq. (12)] after integrating over x. In dashed red, we show the piecewise parametrization by [MS] obtained with the help of Eq. (11). The discontinuity in their parametrization at $Q^{2} = (1.5 \text{ GeV})^{2}$ arises from their choice of matching the Q^{2} integrals of pQCD and the interpolating function over the short distance domain, rather than matching the functions themselves. The log-linear scale conveniently accounts for the



FIG. 5. Log-linear plot of $M_3^{(0)}(1, Q^2)[1 + Q^2/M_W^2]^{-1}$ as a function of Q^2 . The blue curve is the result of our parametrization in Eq. (12), and the red curve is the piecewise parametrization used by [MS]. For a given parametrization, the contribution to \Box_{zW}^{VA} is proportional to the area under the curve, see Eq. (9).

integration measure dQ^2/Q^2 in Eq. (9) so that the correction $\Box_{\gamma W}^{VA}$ is directly proportional to the area under the curve. Although the line shapes are in agreement above $Q^2 \gtrsim 2 \text{ GeV}^2$ and below $Q^2 \lesssim 0.001 \text{ GeV}^2$, ours lies significantly above that of [MS] for an intermediate Q^2 . This difference is the origin of the discrepancy between our central values for Δ_R^V . By working with the two-variable structure function $F_3^{(0)}(x, Q^2)$, we were able to capture a broad variety of physics (Born, $N\pi$, Regge) operating at intermediate Q^2 in contrast with the one-variable analysis of $F_{\rm MS}(Q^2)$ by [MS]. We therefore believe our updated result provides a more realistic assessment of Δ_R^V , even though the difference with them is larger than their quoted theoretical uncertainty.

We conclude by discussing how new measurements could provide tests of our parametrization of $F_3^{(0)}$ and further reduce the uncertainty in Δ_R^V . In view of the upcoming high-intensity neutrino beam program at Fermilab, we wish to point out the potential impact which new, more precise measurements of $M_3^{\nu p+\bar{\nu}p}(1, Q^2)$ at low Q^2 can have on our fit, as evidenced by Fig. 4. That said, we have related $F_3^{(0)}$ and $F_3^{\nu p+\bar{\nu}p}$ within a model. However, by making use of isospin symmetry, we can establish a more robust relationship between $F_3^{(0)}$ and the *P*-odd structure function $F_{3,\gamma Z}^N$. The latter is accessible with parity-violating deep inelastic (inclusive) electron scattering. Since the axial component of the weak neutral current is predominantly isovector, we obtain

$$4F_{3}^{(0)} \approx F_{3,\gamma Z}^{p} - F_{3,\gamma Z}^{n} \approx 2F_{3,\gamma Z}^{p} - F_{3,\gamma Z}^{d}.$$
(22)

Thus, fixed target measurements using hydrogen and deuterium can in principle provide a more direct way to determine $\Box_{\gamma W}^{VA}$ from data. High quality data in the range 0.1 GeV² $\lesssim Q^2 \lesssim 1$ GeV² and $W^2 \gtrsim 5$ GeV² would be particularly advantageous, as our parametrization of $F_3^{(0)}$ admits the greatest model dependence and exhibits the largest difference from that of [MS] in this domain. Such an experimental program will however require a dedicated feasibility study, as the contribution of $F_{3,\gamma Z}$ to the parityviolating asymmetry with a polarized electron beam is suppressed by the small weak charge of the electron. Finally, with the reduction in the uncertainty of $|V_{ud}|$, the error in the first row CKM unitarity constraint is dominated by the uncertainty in $|V_{us}| = 0.2243(5)$. Combined with our results presented here, a commensurate reduction in the latter uncertainty would enhance the impact of first row CKM unitarity tests.

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- M. Tanabashi *et al.* (Particle Data Group), Phys. Rev. D 98, 030001 (2018).
- [2] J. C. Hardy and I. S. Towner, Phys. Rev. C 91, 025501 (2015).
- [3] This value is slightly lower than the one quoted in Ref. [2] and is a result of later updates by the same authors. It can be deduced from Eq. (1) and the values of Δ_R^V and V_{ud} quoted in the text.
- [4] A. Czarnecki, W. J. Marciano, and A. Sirlin, Phys. Rev. D 70, 093006 (2004).
- [5] A. Czarnecki, W. J. Marciano, and A. Sirlin, Phys. Rev. Lett. 120, 202002 (2018).

- [6] W. J. Marciano and A. Sirlin, Phys. Rev. Lett. 96, 032002 (2006).
- [7] W.J. Marciano and A. Sirlin, Phys. Rev. Lett. 56, 22 (1986).
- [8] M. Gorchtein and C. J. Horowitz, Phys. Rev. Lett. 102, 091806 (2009).
- [9] A. Sibirtsev, P.G. Blunden, W. Melnitchouk, and A. W. Thomas, Phys. Rev. D 82, 013011 (2010).
- [10] B. C. Rislow and C. E. Carlson, Phys. Rev. D 83, 113007 (2011).
- [11] M. Gorchtein, C. J. Horowitz, and M. J. Ramsey-Musolf, Phys. Rev. C 84, 015502 (2011).
- [12] P.G. Blunden, W. Melnitchouk, and A. W. Thomas, Phys. Rev. Lett. **107**, 081801 (2011).
- [13] C. E. Carlson and B. C. Rislow, Phys. Rev. D 85, 073002 (2012).
- [14] P.G. Blunden, W. Melnitchouk, and A. W. Thomas, Phys. Rev. Lett. **109**, 262301 (2012).
- [15] N. L. Hall, P. G. Blunden, W. Melnitchouk, A. W. Thomas, and R. D. Young, Phys. Rev. D 88, 013011 (2013).
- [16] B. C. Rislow and C. E. Carlson, Phys. Rev. D 88, 013018 (2013).
- [17] N. L. Hall, P. G. Blunden, W. Melnitchouk, A. W. Thomas, and R. D. Young, Phys. Lett. B 731, 287 (2014); 733, 380 (E) (2014).
- [18] M. Gorchtein and X. Zhang, Phys. Lett. B 747, 305 (2015).
- [19] N. L. Hall, P. G. Blunden, W. Melnitchouk, A. W. Thomas, and R. D. Young, Phys. Lett. B **753**, 221 (2016).
- [20] M. Gorchtein, H. Spiesberger, and X. Zhang, Phys. Lett. B 752, 135 (2016).
- [21] D. Androić *et al.* (Qweak Collaboration), Nature (London) 557, 207 (2018).

- [22] C.-Y. Seng, M. Gorchtein, H. Patel, and M.J. Ramsey-Musolf (to be published).
- [23] A. Sirlin, Rev. Mod. Phys. 50, 573 (1978); 50, 905(E) (1978).
- [24] O. Nachtmann, Nucl. Phys. B63, 237 (1973).
- [25] O. Nachtmann, Nucl. Phys. B78, 455 (1974).
- [26] Z. Ye, J. Arrington, R. J. Hill, and G. Lee, Phys. Lett. B 777, 8 (2018).
- [27] B. Bhattacharya, R. J. Hill, and G. Paz, Phys. Rev. D 84, 073006 (2011).
- [28] O. Lalakulich, E. A. Paschos, and G. Piranishvili, Phys. Rev. D 74, 014009 (2006).
- [29] D. Drechsel, S. S. Kamalov, and L. Tiator, Eur. Phys. J. A 34, 69 (2007).
- [30] L. Tiator and M. Vanderhaeghen, Phys. Lett. B 672, 344 (2009).
- [31] C.A. Piketty and L. Stodolsky, Nucl. Phys. **B15**, 571 (1970).
- [32] V. L. Kashevarov, M. Ostrick, and L. Tiator, Phys. Rev. C 96, 035207 (2017).
- [33] A. L. Kataev and A. V. Sidorov, Phys. Lett. B **331**, 179 (1994).
- [34] J. H. Kim et al., Phys. Rev. Lett. 81, 3595 (1998).
- [35] T. Bolognese, P. Fritze, J. Morfin, D. H. Perkins, K. Powell, and W. G. Scott (Aachen-Bonn-CERN-Democritos-London-Oxford-Saclay Collaboration), Phys. Rev. Lett. 50, 224 (1983).
- [36] D. Allasia et al., Z. Phys. C 28, 321 (1985).
- [37] D. J. Gross and C. H. Llewellyn Smith, Nucl. Phys. B14, 337 (1969).
- [38] S. A. Larin and J. A. M. Vermaseren, Phys. Lett. B 259, 345 (1991).