

Observation of a Space-Time Crystal in a Superfluid Quantum GasJ. Smits,¹ L. Liao,² H. T. C. Stoof,² and P. van der Straten^{1,*}¹*Debye Institute for Nanomaterials Science and Center for Extreme Matter and Emergent Phenomena, Utrecht University, P.O. Box 80.000, 3508 TA Utrecht, Netherlands*²*Institute for Theoretical Physics and Center for Extreme Matter and Emergent Phenomena, Utrecht University, P.O. Box 80.000, 3508 TA Utrecht, Netherlands*

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Time crystals are a phase of matter, for which the discrete time symmetry of the driving Hamiltonian is spontaneously broken. The breaking of discrete time symmetry has been observed in several experiments in driven spin systems. Here, we show the observation of a space-time crystal using ultracold atoms, where the periodic structure in both space and time is directly visible in the experimental images. The underlying physics in our superfluid can be described *ab initio* and allows for a clear identification of the mechanism that causes the spontaneous symmetry breaking. Our results pave the way for the usage of space-time crystals for the discovery of novel nonequilibrium phases of matter.

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Wilczek proposed the idea of time crystals in 2012 [1], where in analogy to space crystals continuous time symmetry is broken spontaneously. Since that time there has been discussion on what should constitute a time crystal [2,3] and how to create them. Watanabe and Oshikawa [4] showed that, in principle, continuous time symmetry cannot be broken spontaneously into discrete symmetry in the ground state. However, there have been proposals to realize instead a discrete time crystal by breaking of discrete time translation symmetry [5–9]. Following a theoretical model by Yao *et al.* [10], several experiments [11–14] realized this particular symmetry breaking in driven spin systems. These experiments were limited to probing a very restricted number of particles [11] or an ensemble of particles without any spatial resolution [12–14], preventing the direct observation of spatial ordering.

In this Letter, we report the direct observation of a space-time crystal exhibiting not only periodic oscillations in time with double the period of the driving force, but also an oscillatory spatial structure; i.e., both discrete time translation symmetry as well as continuous spatial translation symmetry are broken. Because of the small dissipation in our superfluid gas, we can study the space-time crystal over an extensive period of time, showing the collapse and revival of the oscillating long-lived spatially ordered state. Superfluid quantum gases are the ideal system to study discrete time-crystals. Because of the low viscosity and heat conduction, excitations in the system can be induced without the associated heating of the system. Periodic driving of the excitations in the system can easily be arranged due to the harmonic confinement of the atoms in the trap. Crucial in the driven spin systems [11,12,15] has been the occurrence of strong disorder, where either

many-body localization or some other mechanism is the cause for the small dissipation in the experiments.

However, as shown by Else, Bauer, and Nayak [16], time crystals can also exist in the prethermal regime, if the drive frequency is sufficiently large compared to the excitation frequency. Following these experiments, there have been a large number of proposals [17–22] for the observation of time crystals using several different systems (see also the review in Ref. [23]). In superfluid quantum gases, disorder is absent. Since superfluid quantum gases can be imaged using phase-contrast techniques, which allows the accumulation of several tens of images of the same superfluid cloud, the dynamics of the system can be studied over many cycles. Moreover, as the conditions of the space-time crystal are not very sensitive to the initial drive of the excitations, the superfluid cloud can be studied over a prolonged period of time by combining multiple measurement series together, extending the observation period to several seconds. Finally, the dynamics of the superfluid quantum gas in a radial symmetric trap can be simulated using time-splitting spectral methods [24], which allows us to compare our experimental findings with simulations to elucidate the mechanisms behind the space-time crystal formation.

The superfluid is produced in the trap in a cigar-shaped form, where the ratio between the trap frequencies causes the axial size to be about 40 times larger than the radial size. After sample preparation [25], the radial trap frequency is suddenly perturbed, and this induces a radial breathing mode of the cloud with a frequency of $f_D = 104.691(16)$ Hz, which is only weakly damped and has a decay time of several seconds. This radial breathing mode with a period $T_D = 1/f_D$ acts as the drive for the excitation of the cloud in the axial direction. After many radial oscillations, a high-order excitation emerges in the axial

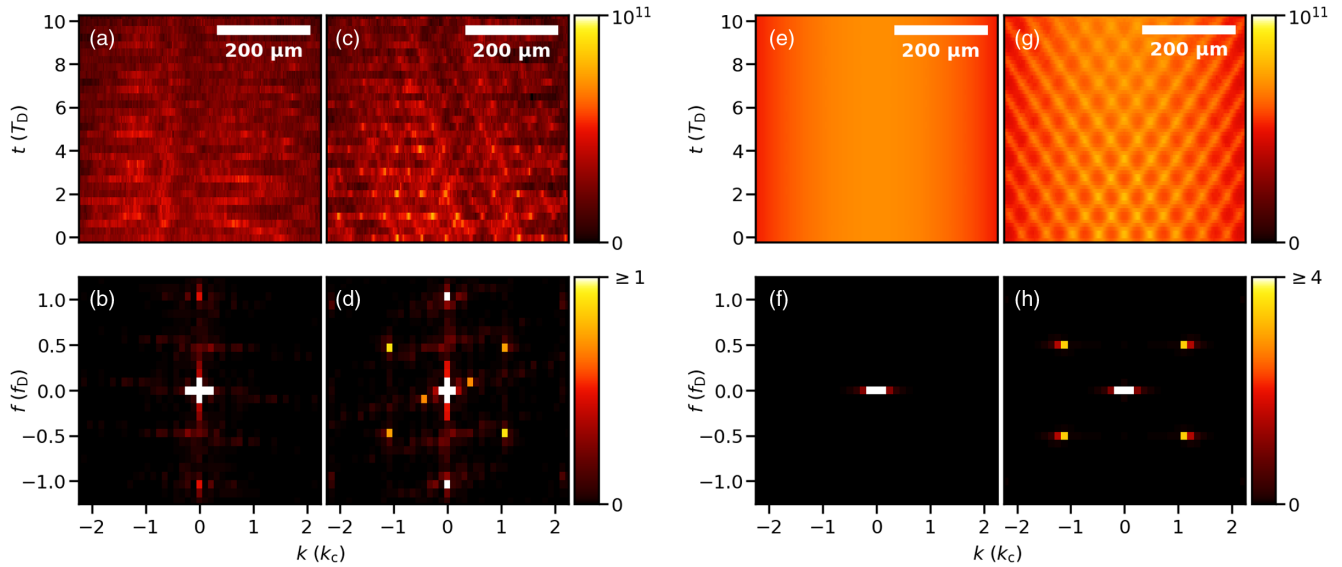


FIG. 3. Fourier analysis and comparison of the experiment with the simulation. (a) Line density at the center of the cloud before the onset of the space-time crystalline phase, directly after the excitation. (b) Fourier transform of the data in (a). Peaks at $f/f_D = \pm 1$ are associated with a weakly excited scissor mode. The signal around the origin is associated with the equilibrium profile of the condensate. (c) Take-out of Fig. 2. Line density at the center of the cloud after the transition to the space-time crystalline phase, after a driving time of 500 ms. A lattice has formed. (d) Fourier transform of the data in (c), with the appearance of four additional peaks due to the space-time crystal at $(k/k_c, f/f_D) = (\pm 1, \pm 0.5)$, where k_c is the center wavelength [25]. (e) Simulated line density for a modulation depth of 0.02, after a wait time of $25T_D$. (f) Fourier transform of the data in (e). Notice that in the simulation only the equilibrium profile is visible. (g) Simulated line density for a modulation depth of 0.2 after a wait time of $25T_D$. A pattern similar to the experimental data of (c) is observed. (h) Fourier transform of (g). Note the appearance of the four additional peaks at $(k/k_c, f/f_D) = (\pm 1, \pm 0.5)$ attributed to the space-time crystal. The line density in (a), (c), (e), and (g) is in units of 10^{11} atoms/m. Fourier images in (b), (d), (f), and (h) are truncated and normalized to 1 for the experimental data.

The periodicity in the occurrence of the space-time crystal coincides approximately with the period that we extract from our simplified model describing the coupling between scissor and breathing mode [25]. The scissor mode has a period of about $T_D/2$ and is not linearly coupled to the axial excitation pattern due to parity conservation.

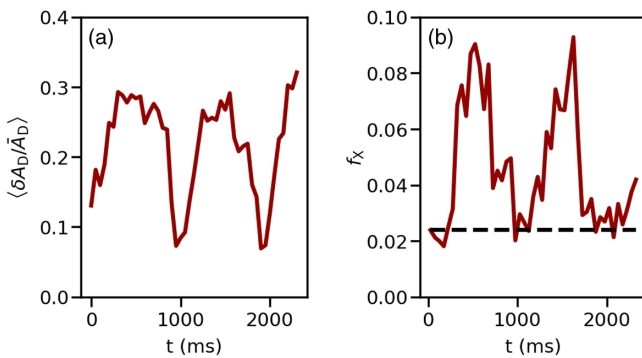


FIG. 4. Long-term behavior of the amplitudes of the drive and crystal fraction. (a) Relative amplitude of the radial breathing mode derived by fitting a two-dimensional profile to the data. (b) Crystal fraction determined from each measurement run. The dashed line indicates the background signal from shot-to-shot noise. Notice that the crystalline phase appears a certain time after the driving mode revives [25].

Theoretically, we treat the space-time crystal variationally as a multimode system with the mode functions $P_{4j+2}(\tilde{z}) - P_{4j}(\tilde{z})$, with $\tilde{z} = z/R_z(t)$ in terms of Legendre polynomials, and frequencies f_j excited by the drive due to the time dependence of the Thomas-Fermi radii $R_x(t)$, $R_y(t)$, and $R_z(t)$, for which $R_i(t) = R_i(t + T_D)$ and $i = x, y$, and z . After substituting this ansatz in the action for the Gross-Pitaevskii equation and neglecting nonlinear mode coupling, we ultimately obtain the Hamiltonian

$$\hat{H} = \sum_j [2\pi\hbar f_j a_j^\dagger a_j + g_j(t)(a_j^\dagger a_j^\dagger + a_j a_j + 2a_j^\dagger a_j)], \quad (1)$$

where $a_j^{(\dagger)}$ are the annihilation (creation) operators for quanta in the mode j and $g(t)$ is the coupling with the periodicity of the drive. By moving to the rotating frame and applying the rotating-wave approximation to eliminate the time dependence of the drive $g(t)$, we find the effective Hamiltonian

$$\hat{H}_{\text{eff}} = \sum_j [2\pi\hbar(f_j - f_D/2)a_j^\dagger a_j + g_{j,0}(a_j^\dagger a_j^\dagger + a_j a_j)], \quad (2)$$

where $g_{j,0}$ is proportional to the amplitude of the drive. Note that this yields a Hamiltonian which is time independent in

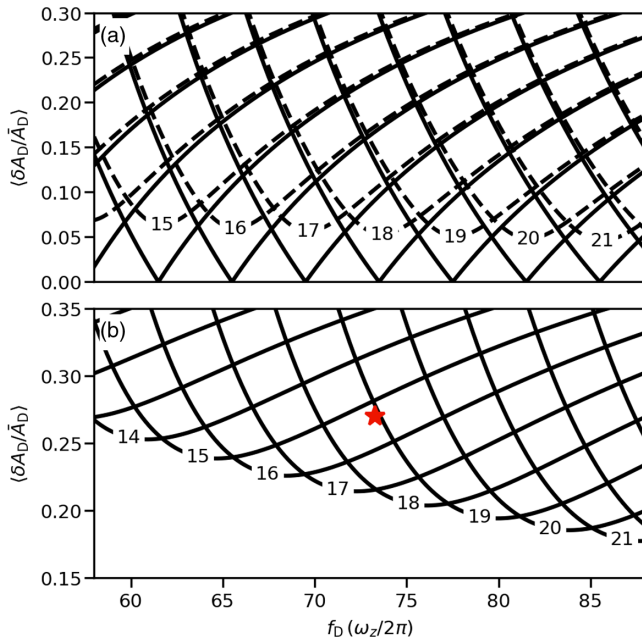


FIG. 5. Minimum required driving amplitude $\langle \delta A_D / \bar{A}_D \rangle$ as a function of the driving frequency f_D in a linear response analysis [36]. (a) Without damping (solid line) a mode j can be driven with an arbitrary small amplitude, if the driving frequency coincides with twice the mode frequency f_j , whereas for small damping (dashed line) there is for any drive frequency a threshold, below which the mode is not excited. (b) Under our experimental conditions (indicated by the red star), the damping is larger and the threshold for exciting modes becomes larger. In the case of linear response, the modes $j = 16, 17$, and 18 can be excited, and, depending on the competition between these modes, one of them dominates the pattern [36].

the rotating frame and that represents the appropriate Hamiltonian for prethermalization of the system.

The mode that is observed depends on the driving frequency f_D and the driving amplitude $\langle \delta A_D / \bar{A}_D \rangle$ [36]. In Fig. 5, the minimum required amplitude is shown as a function of the driving frequency. In the absence of damping, as shown in Fig. 5(a), a mode j can be driven with an arbitrary small amplitude, if the resonance condition $2f_j = f_D$ is fulfilled. In the case of damping, the threshold for exciting the pattern becomes finite. Applying the analysis of Ref. [36] to our experimental conditions [see Fig. 5(b)] shows that the driving amplitude used in our experiment is sufficient to excite several modes j and the competition between these modes causes one of the modes to grow exponentially and thus dominate the observed pattern.

The Hamiltonian of Eq. (2) explicitly breaks the $U(1)$ symmetry $a \rightarrow ae^{i\theta}$. This implies that in the laboratory frame $\langle a_j a_j \rangle \propto e^{-2\pi i f_D t}$ is always nonzero and oscillates with the period of the drive. However, there is an additional \mathbb{Z}_2 symmetry $a_j \rightarrow -a_j$, which is spontaneously broken when $\langle a_j \rangle \neq 0$, which occurs when the mode is Bose

condensed. This leads to the appearance of the time dependence $\langle a_j \rangle \propto e^{-\pi i f_D t}$ in the laboratory frame. The breaking of this \mathbb{Z}_2 symmetry thus leads to an oscillation with period $2T_D$. We propose that for low occupation ($\langle a_j \rangle \simeq 0$) the system is in a state dominated by a description based on the evolution of the pair correlation $\langle a_j a_j \rangle$. As occupation in the mode grows, i.e., the occupation number of the mode $\langle a_j \rangle$ goes up, there is a phase transition from the paired state to a state dominated by dynamics in $\langle a_j \rangle$, breaking the \mathbb{Z}_2 symmetry. We identify this transition as the phase transition to the time crystal.

In summary, we have shown the existence of a space-time crystal which is robust against fluctuations in experimental parameters and long-lived. Future experiments are aimed at studying elementary excitations such as solitons and sound in the presence of a space-time crystal, as our system is an excellent testing ground for these excitations. Moreover, it can be explored whether this spatially ordered state has supersolid properties, as this would allow the study of out-of-equilibrium supersolids [37,38], combining the fields of time crystals and supersolids and exploring a currently unknown corner of physics.

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