

Quantum Speed Limits across the Quantum-to-Classical Transition

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Quantum speed limits set an upper bound to the rate at which a quantum system can evolve. Adopting a phase-space approach, we explore quantum speed limits across the quantum-to-classical transition and identify equivalent bounds in the classical world. As a result, and contrary to common belief, we show that speed limits exist for both quantum and classical systems. As in the quantum domain, classical speed limits are set by a given norm of the generator of time evolution.

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The multifaceted nature of time makes its treatment challenging in the quantum world [1,2]. Nonetheless, the understanding of time-energy uncertainty relations is somewhat privileged [3,4]. To a great extent, this is due to their reformulation in terms of quantum speed limits (QSLs) concerning the ability to distinguish two quantum states connected via time evolution. While QSLs provide fundamental constraints to the pace at which quantum systems can change, a plethora of applications have been found that well extend beyond the realm of quantum dynamics. Indeed, QSLs provide limits to the computational capability of physical devices [5], the performance of quantum thermal machines in finite-time thermodynamics [6,7], parameter estimation in quantum metrology [8,9], quantum control [10–14], the decay of unstable quantum systems [15–18], and information scrambling [19], among other examples [3,4,20].

Specifically, QSLs are derived as upper bounds to the rate of change of the fidelity $F(\tau) = |\langle \psi_0 | \psi_\tau \rangle|^2 \in [0, 1]$ between an initial quantum state $|\psi_0\rangle$ and the corresponding time-evolving state $|\psi_\tau\rangle = \hat{U}(\tau, 0)|\psi_0\rangle$, where $\hat{U}(\tau, 0)$ is the time-evolution operator. More generally, quantum states need not be pure, and given two density matrices ρ_0 and $\rho_\tau = \hat{U}(\tau, 0)\rho_0\hat{U}(\tau, 0)^\dagger$ the fidelity reads

$$F(\tau) = \left[\text{Tr} \sqrt{\sqrt{\rho_0} \rho_\tau \sqrt{\rho_0}} \right]^2. \quad (1)$$

The fidelity is useful to define a metric between quantum states in Hilbert space, known as the Bures angle [21,22]:

$$\mathcal{L}(\rho_0, \rho_\tau) = \cos^{-1} \sqrt{F(\tau)}. \quad (2)$$

This gives a geometric interpretation of the speed limit as the minimum time required to sweep out the angle $\mathcal{L}(\rho_0, \rho_\tau)$ under a given dynamics [23].

For unitary processes, two seminal results are known. The Mandelstam-Tamm bound estimates the speed of evolution in terms of the energy dispersion of the initial state [15,16,22,24–27]. Its original derivation relies on the Heisenberg uncertainty relation. The second seminal result is named after Margolus and Levitin and provides an upper bound to the speed of evolution in term of the difference between the mean energy and the ground state energy [28,29]. Its original derivation relies on the study of the survival amplitude $\langle \psi_0 | \psi_\tau \rangle$. These bounds can be extended to driven and open quantum systems [30–35]. In addition, the two bounds can be unified [29] so that the time of evolution τ required to sweep an angle $\mathcal{L}(\rho_0, \rho_\tau)$ is lower bounded by

$$\tau \geq \tau_{\text{QSL}} = \hbar \mathcal{L}(\rho_0, \rho_\tau) \max \left\{ \frac{1}{E - E_0}, \frac{1}{\Delta E} \right\}, \quad (3)$$

where E_0 is the ground state of the system, E is its mean energy, and ΔE denotes the energy dispersion. Note, however, that there is an infinite family of bounds in terms of higher-order moments of the energy of the system [36].

It is widely believed that these bounds are quantum in nature and that, as a result, exist only in the quantum world [29]. Indeed, in the limit of vanishing \hbar , the right-hand side of (3) equals zero, and one is led to conclude that no “classical” speed limit exists as the inequality becomes trivial:

$$\tau \geq \lim_{\hbar \rightarrow 0} \tau_{\text{QSL}} = 0. \quad (4)$$

This conclusion is further supported by the aforementioned derivations of QSLs, which strongly rely on the framework of the quantum theory. In particular, the Mandelstam-Tamm bound follows from the Heisenberg uncertainty relation [3,15], and the Margolus-Levitin inequality exploits the notion of the transition probability amplitude between two quantum states in Hilbert space [28,29]. We note, however, that recent developments on the generalization of QSLs to

open quantum systems and arbitrary quantum channels have provided new derivations and an alternative understanding of QSLs [30–35]. As a result of these works, given an equation of motion for the state of the system, QSLs are derived in terms of a given norm of the generator of evolution acting on the initial state of the system ρ_0 or the time-dependent state ρ_t (with $0 \leq t \leq \tau$). Such a formulation appears not to be restricted to quantum mechanical systems, as we show here.

In this Letter, we focus on the existence and characterization of QSLs across the quantum-to-classical transition. We show that the conclusion on the quantum nature of QSLs is unjustified. We demonstrate that, contrary to common belief, similar speed limits hold in the classical world. To this end, we adopt a phase-space formulation of quantum mechanics and derive quantum speed limits for quasiprobability distributions: the Wigner function. We find that the speed of evolution is determined by a certain norm of the Moyal product of the Hamiltonian and the Wigner function. Using a semiclassical expansion, we then identify a classical speed limit and show that the resulting bound does indeed govern the evolution of the classical phase-space probability distribution. As a result, we establish the universal existence of fundamental limits to the pace of the evolution of a physical system, independently of its classical or quantum nature.

Quantum speed limits in phase space.—For simplicity and without loss of generality, we consider a one-dimensional system for which the phase-space representation is given by the Wigner function defined as [37,38]

$$W_t(q, p) = \frac{1}{\pi\hbar} \int \langle q - y | \hat{\rho}_t | q + y \rangle e^{2ipy/\hbar} dy, \quad (5)$$

where $\langle q | \hat{\rho}_t | q' \rangle$ denotes a density matrix in the coordinate representation. It is well known that W_t is a quasiprobability distribution that takes real but possibly negative values. We consider the Wigner function of the initial state W_0 and of the time-dependent state W_t generated via unitary dynamics with a time-independent Hamiltonian. The fidelity between any two pure states with respective density matrices $\hat{\rho}_0$ and $\hat{\rho}_t$ can be obtained as the trace in phase space of the corresponding Wigner functions:

$$F(t) = \text{Tr}(\hat{\rho}_0 \hat{\rho}_t) = \int d^2\Gamma W_0 W_t, \quad (6)$$

where $d^2\Gamma = 2\pi\hbar dq dp$, for short.

To derive a QSL, we compute the instantaneous rate of change of the fidelity as a function of time. This can be done using the equation of motion of the Wigner function:

$$\frac{\partial W_t}{\partial t} = \{\{H, W_t\}\} = \frac{1}{i\hbar} (H_{qp} \star W_t - W_t \star H_{qp}), \quad (7)$$

where the Moyal bracket $\{\{A, B\}\}$ can be explicitly written in terms of the Moyal product

$$H_{qp} \star W_t \equiv H_{qp} \exp\left(\frac{i\hbar}{2} \overleftarrow{\partial}_q \overrightarrow{\partial}_p - \frac{i\hbar}{2} \overleftarrow{\partial}_p \overrightarrow{\partial}_q\right) W_t(q, p) \quad (8)$$

and where $H_{qp} = \int dx \langle q - x/2 | \hat{H} | q + x/2 \rangle \exp(ipx/\hbar)$ denotes the Weyl ordered Hamiltonian operator in phase space. From Eqs. (6) and (7), it follows that the rate of change of the fidelity is set by

$$\begin{aligned} \dot{F}(t) &= \int d^2\Gamma W_0 \{\{H, W_t\}\} \\ &= \int d^2\Gamma \{\{H, W_0\}\} W_t, \end{aligned} \quad (9)$$

where we have used integration by parts to derive the second line. Using the Cauchy-Schwarz inequality, one finds

$$|\dot{F}(t)| \leq \left(\int d^2\Gamma W_t^2 \int d^2\Gamma \{\{H, W_0\}\}^2 \right)^{1/2}. \quad (10)$$

The purity of a density matrix is always lower than or equal to unity, so $\int d^2\Gamma W_t^2 \leq 1$, where the equality is reached for pure states or unitarity dynamics, as considered here. As a result,

$$|\dot{F}(t)| \leq v_\Gamma := \left(\int d^2\Gamma \{\{H, W_0\}\}^2 \right)^{1/2}, \quad (11)$$

and we find an upper bound v_Γ to the speed of evolution in phase space, with the dimension of frequency. This bound is, in fact, dictated by the energy variance of the initial state, and for pure states $v_\Gamma = \sqrt{2}\Delta E/\hbar$, with $\Delta E = \sqrt{\langle H^2 \rangle - \langle H \rangle^2}$, as we show in Ref. [39]. A time integration between $t = 0$ to $t = \tau$ readily gives

$$\frac{1 - F(\tau)}{v_\Gamma} = \tau_{\text{QSL}} \leq \tau, \quad (12)$$

which is already a QSL in phase space. Making use of the fact that $0 \leq F(t) \leq 1$ to parametrize the fidelity in terms of the Bures angle

$$\mathcal{L}(\rho_0, \rho_t) = \cos^{-1} \left(\sqrt{\int d^2\Gamma W_0 W_t} \right), \quad (13)$$

that satisfies $F(t) = 1 - \sin^2 \mathcal{L}_t$, we can rewrite the phase-space QSL as

$$\tau_{\text{QSL}} = \frac{\sin^2 \mathcal{L}(\rho_0, \rho_\tau)}{v_\Gamma} = \frac{1 - F(\tau)}{\sqrt{2}} \frac{\hbar}{\Delta E}. \quad (14)$$

Equation (14) constitutes a QSL of the Mandelstam-Tamm type for the Wigner function in phase-space quantum

mechanics. The upper bound to the speed of evolution in phase space v_Γ has units of frequency and is set by the action of the Moyal bracket on the initial Wigner function that is related to the energy variance of the initial state. The distance between states is defined by the Bures angle $\mathcal{L}(\rho_0, \rho_t)$ as a natural statistical distance [21] that is dimensionless and independent of \hbar . Note, however, that it is possible to derive alternative QSLs by considering other distances either in the space of density operators [35] or in phase space [40]. In what follows, we first use a semiclassical expansion to identify a semiclassical speed limit and then combine the results with an operational treatment of quantum dynamics to identify a classical speed limit.

Speed limits across the quantum-to-classical transition.—We recall that the Moyal bracket (7), in a \hbar expansion, reduces to the Poisson bracket so that

$$\{\{W_t, H\}\} = \{W_t, H\} + \mathcal{O}(\hbar^2), \quad (15)$$

where the action of the Poisson bracket on a function f is given by

$$\{f, H\} = \frac{\partial H}{\partial p} \frac{\partial f}{\partial q} - \frac{\partial H}{\partial q} \frac{\partial f}{\partial p}, \quad (16)$$

and rules the dynamics in classical statistical mechanics according to the (classical) Liouville equation. As a result, to leading order in the semiclassical \hbar expansion of the equation of motion for the Wigner function Eq. (7), the speed limit in phase space does not vanish. In particular, the semiclassical speed limit (SSL) reads

$$\begin{aligned} \tau \geq \tau_{\text{SSL}} &= \frac{\sin^2 \mathcal{L}(\rho_0, \rho_\tau)}{(\int d^2\Gamma \{H, W_0\}^2)^{1/2}} \\ &= \frac{\sin^2 \mathcal{L}(\rho_0, \rho_\tau)}{\|\{H, W_0\}\|_2}, \end{aligned} \quad (17)$$

where $\|f\|_2 = (\int |f|^2 d^2\Gamma)^{1/2}$ is the L^2 norm of f and we emphasize that $\|\{H, W_0\}\|_2$ has frequency units.

Let us discuss this expression in detail. The Moyal product provides a one-parameter deformation of the noncommutative algebra in quantum mechanics and of the commutative algebra in classical phase space according to Eq. (15). By reformulating QSLs in terms of Wigner functions, this correspondence leads to the identification of a SSL in phase space. The distance $\mathcal{L}(\rho_0, \rho_\tau)$ between states ρ_0 and ρ_τ is well defined whether these states are valid classical states (i.e., with a positive Wigner function) or not. As a result, Eq. (17) constitutes the semiclassical limit of the Mandelstam-Tamm time-energy uncertainty relation. Using Hamilton's equation of motion

$$\frac{\partial W_t}{\partial t} = \{H, W_t\}, \quad (18)$$

we interpret the upper bound to the speed of evolution as the root mean square of the initial rate of change of the Wigner function at $t = 0$ averaged over phase space, i.e.,

$$v_\Gamma^{\text{SSL}} = \|\{H, W_0\}\|_2 = \sqrt{\int d^2\Gamma (\partial_t W_t|_{t=0})^2}. \quad (19)$$

Alternatively, introducing the Liouvillian $i\hat{L}W_t = -\{H, W_t\}$, we can restate the SSL as

$$\tau_{\text{SSL}} = \frac{\sin^2 \mathcal{L}(W_0, W_\tau)}{\|\hat{L}W_0\|_2}. \quad (20)$$

As in the quantum case (14), the SSL is set by a given norm of the generator of evolution \hat{L} averaged over the initial state W_0 . We note that this expression still contains an explicit \hbar both in the integration measure and in the definition of the Wigner function.

Classical speed limit.—To identify a classical speed limit (CSL) from the semiclassical expression (20), we resort to the operational dynamic modeling developed by Bondar *et al.* [41,42]. The equivalence of the evolution of dynamical average values in the quantum and classical domain via Ehrenfest theorems yields a relation between the classical phase-space probability density $q_t(q, p)$ and the Wigner function $W_t(q, p)$:

$$q_t(q, p) = 2\pi\hbar W_t(q, p)^2. \quad (21)$$

Note that the factor $2\pi\hbar$, so far accounted for in $d^2\Gamma$, can be interpreted as dividing the phase-phase into cells of area $2\pi\hbar$ [43], which corresponds to the Böhr-Sommerfeld quantization rule in the “old” quantum theory. The normalization of a pure quantum state $|\psi_t\rangle$ carries over the classical distribution $\int 2\pi\hbar dq dp W_t(q, p)^2 = \int dq dp q_t(q, p) = 1$.

Accordingly, the fidelity (6) reduces to the Bhattacharyya coefficient [44]

$$B(t) = B(q_0, q_t) = \int dq dp \sqrt{q_0(q, p)q_t(q, p)} \quad (22)$$

that is related to the Hellinger distance $H(q_0, q_t)$ via the identity $B(t) = 1 - H(q_0, q_t)^2$. Note that $B(0) = 1$ due to the normalization condition. The Bures angle becomes

$$\mathcal{L}_B = \cos^{-1} \sqrt{B(t)}, \quad (23)$$

and the CSL thus reads $\tau \geq \tau_{\text{CSL}}$ with

$$\begin{aligned}\tau_{\text{CSL}} &= \frac{\sin^2 \mathcal{L}_B(q_0, q_\tau)}{\sqrt{\int dq dp (\partial_t \sqrt{q_t}|_{t=0})^2}} = \frac{\sin^2 \mathcal{L}_B(q_0, q_\tau)}{\sqrt{\int dq dp \{H, \sqrt{q_0}\}^2}} \\ &= \frac{1 - B(\tau)}{\|\hat{L} \sqrt{q_0}\|_2},\end{aligned}\quad (24)$$

where we identify the denominator with the upper bound to the classical speed of evolution $v_{\text{F}}^{\text{CSL}} = \|\hat{L} \sqrt{q_0}\|_2$ and \hat{L} is the classical Liouville operator satisfying $\partial q_t + i\hat{L}q_t = 0$. This is our main result and constitutes a classical version of the Mandelstam-Tamm bound.

It is worth emphasizing that this bound can be derived independently of the semiclassical approach by making exclusive reference to the classical Hamiltonian formalism. Indeed, the rate of change of the Bhattacharyya coefficient is given by

$$\dot{B}(q_0, q_t) = \int dq dp \sqrt{\rho_0} \frac{\dot{q}_t}{2\sqrt{q_t}}. \quad (25)$$

Using Liouville's equation, we can rewrite the rate of change of the classical probability distribution to find

$$\frac{\dot{q}_t}{2\sqrt{q_t}} = \frac{\{H, q_t\}}{2\sqrt{q_t}} = \{H, \sqrt{q_t}\}. \quad (26)$$

To obtain a classical speed limit that depends only on the initial state, as opposed to its time evolution, it is convenient to shift the action of the Poisson bracket to the initial state q_0 . This is readily accomplished by an integration by parts, assuming q_t vanishes at the end points of the integration, that yields

$$\dot{B}(q_0, q_t) = - \int dq dp \{H, \sqrt{\rho_0}\} \sqrt{q_t}. \quad (27)$$

Use of the Cauchy-Schwarz inequality and the normalization condition $\int dq dp q_t = 1$ lead to

$$|\dot{B}(q_0, q_t)| \leq \left(\int dq dp \{H, \sqrt{\rho_0}\}^2 \right)^{1/2}, \quad (28)$$

which upon integration over the time variable from $t = 0$ to $t = \tau$ yields Eq. (24), given that $1 - B(\tau) = \sin^2 \mathcal{L}_B$.

Note that we consider only smooth classical phase-space distributions, for which $v_{\text{F}}^{\text{CSL}} = \|\partial_t \sqrt{q_t}|_{t=0}\|_2$ is well defined. For a singular distribution of the form $q_t(q, p) = \delta[q - q_{\text{cl}}(t)]\delta[p - p_{\text{cl}}(t)]$, characterizing a certain trajectory of a classical particle, the upper bound to the phase-space velocity $\|\hat{L} \sqrt{q_0}\|_2$ is singular and needs to be regularized. In this limit, the CSL is expected to vanish as the trajectories $q_t(q, p)$ and $q_t(q, p)' = q_t(q + \epsilon_q, p + \epsilon_p)$ are distinguishable for any ϵ_q, ϵ_p with $|\epsilon_q| > 0$ and $|\epsilon_p| > 0$ in the sense that $B(q_0, q_t') = 0$ and $\mathcal{L}_B = \pi/2$.

Quadratic Hamiltonians.—The existence of classical speed limits and their correspondence with their quantum

counterpart become self-evident whenever the Hamiltonian driving the evolution is quadratic in the position and momentum operators. The equation of motion of the Wigner function (7) simplifies, and the phase-space generators of evolution in classical and quantum dynamics are then equivalent. In the classical case, for a time-independent Hamiltonian, the corresponding canonical transformations

$$\begin{pmatrix} q \\ p \end{pmatrix} = \begin{pmatrix} \alpha & \beta \\ \gamma & \delta \end{pmatrix} \begin{pmatrix} q' \\ p' \end{pmatrix} \quad (29)$$

are elements of the two-dimensional real symplectic group $Sp(2, \mathbb{R})$. In the quantum case, the phase-space propagator that determines the evolution of the Wigner function via the identity

$$W_n(q, p; t) = \iint dq' dp' K(q, p|q', p') W_n(q', p'; 0) \quad (30)$$

becomes

$$K(q, p|q', p') = \delta[q' - (\alpha q + \beta p)]\delta[p' - (\gamma q + \delta p)], \quad (31)$$

and it is therefore identical to the classical one [45]. The quantum and semiclassical phase-space limits, Eqs. (14) and (17), are identical in this case. When the generator of evolution is explicitly time dependent, a representation of the corresponding canonical transformations is still possible.

For the sake of illustration, we focus on the time-dependent harmonic oscillator:

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{1}{2} m \omega(t)^2 \hat{q}^2, \quad (32)$$

for which quantum speed limits have been reported with multiple applications including the characterization of control protocols [13,14,46–48] and the performance of quantum thermal machines [6]. As shown in Ref. [39], in the quantum case, the Wigner function of an eigenstate at $t = 0$ evolves under a modulation of the trapping frequency $\omega(t)$ according to

$$\begin{aligned}W_n(q, p; t) &= W_n\left(\frac{q}{b}, bp - mqb; 0\right) \\ &= \frac{(-1)^n}{\pi \hbar} e^{-(2/\hbar\omega_0)[(P^2/2m) + (1/2)m\omega_0^2 Q^2]} \\ &\quad \times L_n\left[\frac{4}{\hbar\omega_0} \left(\frac{P^2}{2m} + \frac{1}{2} m\omega_0^2 Q^2\right)\right],\end{aligned}\quad (33)$$

that we explicitly find in terms of the Laguerre polynomials $L_n(x)$ and the canonically conjugated pair of variables

$$Q := \frac{q}{b}, \quad P = bp - mqb, \quad (34)$$

associated with the matrix

$$\begin{pmatrix} \alpha & \beta \\ \gamma & \delta \end{pmatrix} = \begin{pmatrix} 1/b & 0 \\ -m\dot{b} & b \end{pmatrix}.$$

The time-dependent scaling factor $b(t) > 0$ is the solution of the Ermakov equation, $\ddot{b} + \omega(t)^2 b = \omega_0^2/b^3$, with the boundary conditions $b(0) = 1$ and $\dot{b}(0) = 0$; see, e.g., [49]. As a result, the dynamics arbitrarily far from equilibrium does not alter the form of the Wigner function and can be simply accounted for by the definition of the conjugated pair (34).

For the ground state of the harmonic oscillator with $n = 0$, $W_0(q, p, t) \geq 0$ is a smooth Gaussian distribution for all $0 \leq t \leq \tau$. When the classical distribution is chosen to be also of Gaussian form $\rho_0(q, p) = \exp(-q^2/\sigma_q^2 - p^2/\sigma_p^2)/(\pi\sigma_q\sigma_p)$, the CSL in Eq. (24) equals the quantum and semiclassical phase-space limits, Eqs. (14) and (17), provided that $\sigma_q = x_0/\sqrt{2}$ and $\sigma_p = \hbar/(x_0\sqrt{2})$ as dictated by the correspondence (21); see [39]. From the exact dynamics $\rho_t(q, p) = \rho_0(Q, P)$, we find the Bhattacharyya coefficient

$$B(q_0, q_t) = 2 \left[\frac{(1 + b^2)^2}{b^2} + \left(\frac{m\sigma_x \dot{b}}{\sigma_p} \right)^2 \right]^{-1/2}, \quad (35)$$

while the upper bound for the phase-space speed of evolution is set by

$$v_{\Gamma}^{\text{CSL}} = \|\{H, \sqrt{\rho_0}\}\|_2 = \frac{m\sigma_x |\dot{b}(0)|}{2\sigma_p}. \quad (36)$$

While the generalization of the CSL (24) to time-dependent generators is straightforward [39], we focus on the case when the driven Hamiltonian is constant for $t > 0$ and let the frequency of the trap be suddenly turned off at $t = 0$. It then follows that $b(t) = \sqrt{1 + \omega_0^2 t^2}$ and $\dot{b}(0) = \omega_0$. To illustrate these results, we show in Fig. 1 how the characteristic velocity in the phase space of v_{Γ}^{CSL} in (36) remains an upper bound to the instantaneous phase-space velocity set by the absolute value of the Bhattacharyya coefficient derivative during the course of the evolution.

In conclusion, we have shown that there exist fundamental speed limits to the pace of evolution of an arbitrary physical system, in both the classical and quantum worlds. To this end, we have introduced quantum speed limits in phase space and derived their semiclassical limit. Their comparison should be useful to identify scenarios in which the quantum dynamics provides a speedup over the classical evolution. From the semiclassical limit, we have further identified a family of classical speed limits that governs the classical Hamiltonian dynamics in phase space. In the quantum, semiclassical, and classical settings, speed limits are universally set by a given norm of the generator of

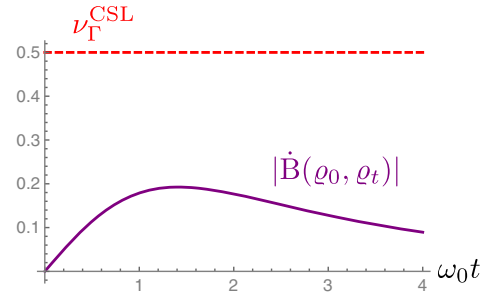


FIG. 1. Classical speed limit to the pace of evolution. Comparison of the upper bound to the phase-space speed of evolution v_{Γ}^{CSL} with the absolute value of the instantaneous rate of change of the Bhattacharyya coefficient $|\dot{B}(q_0, q_t)|$ as a function of time. The dynamics corresponds to a free expansion of a classical probability distribution of Gaussian form that is initially confined in a harmonic potential of frequency ω_0 , which is switched off for $t > 0$. The unit of time is set by ω_0^{-1} .

the dynamics and the state of the system under consideration. Our results provide further insight on the nature of time-energy uncertainty relations, speed limits in arbitrary physical process, and the limits of computation.

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Note added.—Recently, we learned about Ref. [50] which finds similar results.

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- [1] *Time in Quantum Mechanics—Vol. 1*, Lecture Notes in Physics Vol. 734, edited by J. G. Muga, R. Mayato, and I. L. Egusquiza (Springer, Heidelberg, 2002).
 - [2] *Time in Quantum Mechanics—Vol. 2*, Lecture Notes in Physics Vol. 789, edited by J. G. Muga, A. Ruschaupt, and A. del Campo (Springer, Heidelberg, 2009).
 - [3] P. Busch, *Lect. Notes Phys.* **734**, 73 (2008), Chap. 3 in Ref. [1].
 - [4] L. S. Schulman, *Lect. Notes Phys.* **734**, 107 (2008), Chap. 4 in Ref. [1].
 - [5] S. Lloyd, *Nature (London)* **406**, 1047 (2000); *Phys. Rev. Lett.* **88**, 237901 (2002); V. Giovannetti, S. Lloyd, and L. Maccone, *Phys. Rev. A* **67**, 052109 (2003).
 - [6] A. del Campo, J. Goold, and M. Paternostro, *Sci. Rep.* **4**, 6208 (2014).
 - [7] F. Campaioli, F. A. Pollock, F. C. Binder, L. C. Céleri, J. Goold, S. Vinjanampathy, and K. Modi, *Phys. Rev. Lett.* **118**, 150601 (2017).

- [8] R. Demkowicz-Dobrzanski, J. Kolodynski, and M. Guta, *Nat. Commun.* **3**, 1063 (2012).
- [9] M. Beau and A. del Campo, *Phys. Rev. Lett.* **119**, 010403 (2017).
- [10] M. Demirplak and S. A. Rice, *J. Chem. Phys.* **129**, 154111 (2008).
- [11] T. Caneva, M. Murphy, T. Calarco, R. Fazio, S. Montangero, V. Giovannetti, and G. E. Santoro, *Phys. Rev. Lett.* **103**, 240501 (2009).
- [12] A. del Campo, M. M. Rams, and W. H. Zurek, *Phys. Rev. Lett.* **109**, 115703 (2012).
- [13] S. Campbell and S. Deffner, *Phys. Rev. Lett.* **118**, 100601 (2017).
- [14] K. Funo, J.-N. Zhang, C. Chatou, K. Kim, M. Ueda, and A. del Campo, *Phys. Rev. Lett.* **118**, 100602 (2017).
- [15] L. Mandelstam and I. Tamm, *J. Phys. (Moscow)* **9**, 249 (1945).
- [16] K. Bhattacharyya, *J. Phys. A* **16**, 2993 (1983).
- [17] A. Chenu, M. Beau, J. Cao, and A. del Campo, *Phys. Rev. Lett.* **118**, 140403 (2017).
- [18] M. Beau, J. Kiukas, I. L. Egusquiza, and A. del Campo, *Phys. Rev. Lett.* **119**, 130401 (2017).
- [19] A. del Campo, J. Molina-Vilaplana, and J. Sonner, *Phys. Rev. D* **95**, 126008 (2017).
- [20] S. Deffner and S. Campbell, *J. Phys. A* **50**, 453001 (2017).
- [21] W. K. Wootters, *Phys. Rev. D* **23**, 357 (1981).
- [22] A. Uhlmann, *Phys. Lett. A* **161**, 329 (1992).
- [23] B. Russell and S. Stepney, *International Journal of Foundations of Computer Science* **28**, 321 (2017).
- [24] G. N. Fleming, *Nuovo Cimento A* **16**, 232 (1973).
- [25] J. Anandan and Y. Aharonov, *Phys. Rev. Lett.* **65**, 1697 (1990).
- [26] L. Vaidman, *Am. J. Phys.* **60**, 182 (1992).
- [27] P. Pfeifer, *Phys. Rev. Lett.* **70**, 3365 (1993).
- [28] N. Margolus and L. B. Levitin, *Physica (Amsterdam)* **120D**, 188 (1998).
- [29] L. B. Levitin and T. Toffoli, *Phys. Rev. Lett.* **103**, 160502 (2009).
- [30] M. M. Taddei, B. M. Escher, L. Davidovich, and R. L. de Matos Filho, *Phys. Rev. Lett.* **110**, 050402 (2013).
- [31] A. del Campo, I. L. Egusquiza, M. B. Plenio, and S. F. Huelga, *Phys. Rev. Lett.* **110**, 050403 (2013).
- [32] S. Deffner and E. Lutz, *Phys. Rev. Lett.* **111**, 010402 (2013).
- [33] Y.-J. Zhang, W. Han, Y.-J. Xia, J.-P. Cao, and H. Fan, *Sci. Rep.* **4**, 4890 (2014).
- [34] I. Marvian and D. A. Lidar, *Phys. Rev. Lett.* **115**, 210402 (2015).
- [35] D. P. Soares-Pinto, M. Cianciaruso, L. C. Céleri, G. Adesso, and D. O. Soares-Pinto, *Phys. Rev. X* **6**, 021031 (2016).
- [36] B. Zielinski and M. Zych, *Phys. Rev. A* **74**, 034301 (2006).
- [37] E. P. Wigner, *Phys. Rev.* **40**, 749 (1932).
- [38] M. Hillery, R. F. O'Connell, M. O. Scully, and E. P. Wigner, *Phys. Rep.* **106**, 121 (1984).
- [39] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevLett.120.070401> for additional details on the phase-space speed of evolution, the derivation of the fidelity in the time-dependent harmonic oscillator, and the generalization of classical speed limits to time-dependent generators.
- [40] S. Deffner, *New J. Phys.* **19**, 103018 (2017).
- [41] D. I. Bondar, R. Cabrera, R. R. Lompay, M. Y. Ivanov, and H. A. Rabitz, *Phys. Rev. Lett.* **109**, 190403 (2012).
- [42] D. I. Bondar, R. Cabrera, D. V. Zhdanov, and H. A. Rabitz, *Phys. Rev. A* **88**, 052108 (2013).
- [43] L. D. Landau and E. M. Lifshitz, *Quantum Mechanics*, 2nd ed. (Pergamon, New York, 1965), Vol. 3, Chap. VII.
- [44] A. Bhattacharyya, *Sankhyā: The Indian Journal of Statistics* **7**, 401 (1946).
- [45] G. García-Calderón and M. Moshinsky, *J. Phys. A* **13**, L185 (1980).
- [46] Xi Chen and J. G. Muga, *Phys. Rev. A* **82**, 053403 (2010).
- [47] Y.-Y. Cui, X. Chen, and J. G. Muga, *J. Phys. Chem. A* **120**, 2962 (2016).
- [48] Y. Zheng, S. Campbell, G. De Chiara, and D. Poletti, *Phys. Rev. A* **94**, 042132 (2016).
- [49] X. Chen, A. Ruschhaupt, S. Schmidt, A. del Campo, D. Guéry-Odelin, and J. G. Muga, *Phys. Rev. Lett.* **104**, 063002 (2010).
- [50] M. Okuyama and M. Ohzeki, following Letter, *Phys. Rev. Lett.* **120**, 070402 (2018).