

Generalized Supersoft Supersymmetry Breaking and a Solution to the μ Problem

Ann E. Nelson¹ and Tuhin S. Roy^{2,3}

¹*Department of Physics, University of Washington, Seattle, Washington 98195, USA*

²*Department of Theoretical Physics, Tata Institute of Fundamental Research, Mumbai 400005, India*

³*Theory Division T-2, Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA*

(Received 19 February 2015; published 20 May 2015)

We propose the framework *generalized supersoft supersymmetry breaking*. “Supersoft” models, with D -type supersymmetry breaking and heavy Dirac gauginos, are considerably less constrained by the LHC searches than the well studied MSSM. These models also ameliorate the supersymmetric flavor and CP problems. However, previously considered mechanisms for obtaining a natural size Higgsino mass parameter (namely, μ) in supersoft models have been relatively complicated and contrived. Obtaining a 125 GeV for the mass of the lightest Higgs boson has also been difficult. Additional issues with the supersoft scenario arise from the fact that these models contain new scalars in the adjoint representation of the standard model, which may obtain negative squared-masses, breaking color and generating too large a T parameter. In this Letter, we introduce new operators into supersoft models which can potentially solve all these issues. A novel feature of this framework is that the new μ term can give unequal masses to the up and down type Higgs fields, and the Higgsinos can be much heavier than the Higgs boson without fine-tuning. However, unequal Higgs and Higgsino masses also remove some attractive features of supersoft supersymmetry.

DOI: 10.1103/PhysRevLett.114.201802

PACS numbers: 12.60.Jv

Supersymmetry (SUSY) at the electroweak scale offers potential solutions to the gauge hierarchy and dark matter problems, along with a route towards a grand unified theory (GUT)[1]. A crucial ingredient is the presence of the Higgsinos (the superpartners of the Higgs bosons) with masses at the electroweak scale. At first glance, this does not appear to be a critical issue, since a supersymmetric Higgs and Higgsino mass term, namely “ μ ”, is allowed. In fact, issues regarding Higgsino masses are often trivialized by evoking the argument that due to the nonrenormalization of the superpotential, any value of μ is technically natural. However, this response does not address the depth of the problem. The μ parameter needs to be of the order of the electroweak mass scale, which, in a supersymmetric theory, is not an input parameter in the ultraviolet (UV), but is rather generated in the infrared (IR), after the theory is renormalized down to the IR, and is naturally at the scale of the superpartner masses [3–7]. These masses, in turn, are functions of the two fundamental mass scales of the theory: (i) the scale of the SUSY breaking vacuum expectation value (VEV) in the hidden sector, and (ii) the mass scale associated with the messenger mechanism which connects the hidden sector and the visible sector fields. In models of dynamical supersymmetry breaking (DSB), the scale of SUSY breaking is generated via dimensional transmutation [8–11]. The messenger scale is often the Planck scale [7,12–18], or the GUT scale [3,15,18], or can be the scale of DSB [19]. Inclusion of a bare mass term, which is of the order of the electroweak scale by pure

coincidence makes the theory much less elegant and plausible.

The μ problem is often discussed in the context of the minimal supersymmetric standard model (MSSM), which is the most well-studied incarnation of weak scale SUSY. Note that the MSSM is the weak scale effective theory of an underlying supersymmetric theory, with SUSY being spontaneously broken by the nonzero VEV of the F component of a hidden sector chiral superfield. In this framework, a robust solution to the μ problem is provided by the Giudice-Masiero mechanism [20], whereby a manifestly supersymmetric higher dimensional operator involving the Higgs fields and the SUSY breaking hidden sector superfield becomes a μ term. This mechanism assures that the μ term is naturally of the order of the superpartner masses. Note that the SUSY breaking terms of the MSSM are known as “soft” [21–23], because the resulting theory has only logarithmic UV divergences. Such logarithmic divergences however mean that the soft terms are sensitive to short distance flavor and CP violating physics which could potentially lead to problematic flavor-changing neutral currents (FCNC) [23,24], and new phases that could make detectable and potentially excessive contributions to electric dipole moments [25–30]. More recently, the accumulated null observations have put severe constraints on the MSSM, the most serious of which arises from the lack of observation of excess events with jets + missing energy at the LHC. In weak scale SUSY, events with jets + missing energy are produced mostly due to the

production of squarks and gluinos, which subsequently decay to jets and the lightest supersymmetric particles (LSPs). These cross sections are maximized for degenerate squarks and gluinos, which is a generic feature of the MSSM. Within its framework, squarks receive loop suppressed but log enhanced contribution from the gluino mass as the theory is renormalized down to the IR. Except in the case where the squarks start out to be hierarchically heavier than the gauginos at the UV (such as in split-SUSY [31–33]), the gluino mass is always comparable to the squark masses in the MSSM. Satisfying experimental constraints, therefore, requires the raising of the mass scale of all colored particles. Also note that, because of the restricted form of the Higgs potential in the MSSM, the top squarks are now required to be very heavy, with mass of order a TeV or more in order to obtain 125 GeV for the mass of the Higgs boson. Since renormalization of the soft Higgs mass-squared term is proportional to the top squark mass, a heavy top squark gives rise to a finely tuned cancellation in the Higgs mass squared parameter. Thus, in the MSSM, with SUSY breaking parameters run down from a high scale, SUSY's promise to explain the origin of the weak scale without fine-tuning, is fading in the light of the LHC Higgs boson discovery and in the absence of any SUSY discovery [34–37].

An alternative way to break supersymmetry is via a VEV for the D component of a hidden sector real superfield [26,40]. Such symmetry breaking may be mediated to the visible sector via a class of operators known as “supersoft”, as they do not induce even logarithmic ultraviolet divergences in squark and slepton masses [41]. The most important previously considered supersoft operators are those giving rise to Dirac gaugino masses [26,40,42,43]. In supersoft models, the radiatively generated squark and slepton masses are finite, flavor symmetric, positive, UV insensitive, and light compared to the gaugino masses [41]. Therefore, these models additionally avoid the flavor-changing neutral current, naturalness, and CP difficulties of the MSSM. A heavy gluino suppresses processes such as gluino pair production and squark-gluino production. Also, the pair production of squarks is reduced as the T -channel diagrams involving gluinos do not contribute. Therefore, Dirac masses allow for a reduction in the number of events with jets + missing energy for a given squark mass [34,44–51]. The μ problem is, however, severe in supersoft models. The Giudice-Masiero mechanism does not work, since SUSY breaking is not mediated by the F term of a chiral superfield, but by the D term of a real superfield instead. A solution was proposed in Ref. [41], where the conformal compensator generates masses for Higgsinos. To generate the right Higgsino masses, however, this approach requires a conspiracy among the SUSY breaking scale, the messenger scale, and the Planck scale. One could reintroduce the gauge singlet chiral superfield with an F term and use the Giudice-Masiero mechanism. However, such a

gauge singlet field may lead to power law UV sensitivity, and to additional flavor and CP violating SUSY breaking operators; thus spoiling the supersoft solution to the SUSY FCNC and CP problems [41,52,53]. It is also conceivable to generate a μ term via a supersymmetric VEV of a singlet superfield, again bringing in the possibility of new power law divergences in the singlet potential. If the singlet carries discrete symmetries, then there could be cosmological problems with the production of domain walls associated with breaking of the discrete symmetries. Another potential problem with supersoft models is that the D term contribution to the Higgs quartic coupling vanishes [41], and accommodating a 125 GeV Higgs boson becomes difficult.

In this Letter, we propose a complete and viable framework of weak scale SUSY, namely *generalized supersoft supersymmetry*, where all SUSY breaking effects are sourced by the D component of a real field or operator from the hidden sector. We include a new class of D term mediated soft (but not necessarily supersoft) operators that allow for a new solution to the μ problem, restore the Higgs quartic coupling, and provide considerable modification to supersoft phenomenology.

The visible sector of our supersoft model includes the superfields of the MSSM, as well as additional chiral superfields Σ_i in the adjoint representation of the SM gauge groups. The fermionic components of Σ_i , (namely, ψ_i), will obtain Dirac masses with the gauginos (λ_i). Supersymmetry is broken by a D term of a hidden sector real superfield V'

$$\mathcal{D} \equiv \frac{1}{8} \langle D^2 \bar{D}^2 V' \rangle > 0. \quad (1)$$

The messenger sector that connects the visible and hidden sector is assumed to be very heavy, and we may integrate it out at the messenger scale M_m , which, in turn, could be as high as the Planck scale. The operators generating the gaugino masses are [43]:

$$\int d^2\theta \frac{w_{1,i}}{4} \frac{\bar{D}^2 D^\alpha V'}{M_m} W_{i,\alpha} \Sigma_i \longrightarrow M_{D_i} \lambda_i \psi_i, \quad (2)$$

where $M_{D_i} = \frac{w_{1,i} g_i}{\sqrt{2}} \frac{\mathcal{D}}{M_m}$.

In the above, $W_{i,\alpha}$ is the field-strength superfield of i th SM gauge group, with α being the spinor index. M_m is the messenger scale, w_1 are dimensionless coupling constants, and D and \bar{D} are superderivatives.

An additional class of supersoft terms gives mass to the scalar components of the Σ_i fields:

$$\int d^2\theta \frac{w_{3,i}}{4} \frac{(\frac{1}{4} \bar{D}^2 D V')^2}{M_m^2} \Sigma_i^2 \longrightarrow \left(\frac{w_{3,i}}{2} \frac{\mathcal{D}^2}{M_m^2} \right) \frac{\sigma_i^2}{2}. \quad (3)$$

In Eq. (3), σ_i denotes the scalar components of the Σ_i chiral superfields. Since these operators are generated at the

messenger scale, the scalar masses are of the order of the gaugino masses. Note that even though the gaugino mass operators in Eq. (2) give rise to masses for the real components of σ fields, Eq. (3) remains the only source of masses for the imaginary components at tree level. Also, given the fact that the squared masses generated in Eq. (3) are linear in the coupling constants $w_{3,i}$, these can be negative, giving rise to nonzero VEV for the color octet field, thus breaking color. The gaugino mediated squared masses for these fields are positive. However, as explained before, these masses are loop suppressed and *not* log enhanced and are, therefore, small with respect to the masses in Eq. (3). In gauge mediated supersoft models, some intricate model building is required to avoid negative masses squared for some of the adjoint scalars [41,54–56]. Both sets of terms are invariant under the hidden sector gauge symmetry $V' \rightarrow V' + \Lambda + \Lambda^\dagger$, where Λ is a chiral superfield. As discussed in Ref. [41], this hidden sector gauge invariance is key to the absence of UV-sensitive contributions to supersymmetry breaking scalar masses.

In this Letter, we propose a new class of operators which ameliorates all of the previously mentioned problems in this framework:

$$- \int d^2\theta \frac{1}{4} w_{2,\Phi_1\Phi_2} \frac{\bar{D}^2(D^\alpha V' D_\alpha \Phi_1)}{M_m} \Phi_2. \quad (4)$$

In Eq. (4), Φ_1 and Φ_2 are visible sector chiral superfields such that the bilinear $\Phi_1 \Phi_2$ is a gauge singlet. Examples of such bilinear gauge singlet in the weak scale supersymmetry are $H_u H_d$ and Σ_i^2 . Note that the operators as expressed in Eq. (4) are manifestly chiral (and part of the superpotential) because of the fact that $\bar{D}^3 = 0$. The terms in Eq. (4) can be given a gauge invariant form (but not supersymmetric), since if V' is set equal to its VEV, we find:

$$\bar{D}^2(D^\alpha V' D_\alpha \Phi_1) = (\bar{D}^2 D^\alpha V') D_\alpha \Phi_1 + \dots, \quad (5)$$

where \dots represent extra terms that do not contribute to the superpotential. When we treat our operators containing V' as a spurion, since it can come either from a supersymmetric or a gauge invariant operator, it will only generate gauge invariant corrections to SUSY breaking operators, and hence, cannot generate terms which require nongauge invariant counterterms. There are however other spurionic terms which share the feature of being either supersymmetric or gauge invariant, which can contribute to squark and slepton masses and nonsupersymmetric trilinears, so the new operators are not necessarily supersoft. One important aspect of this operator is that ordering of Φ_1 and Φ_2 in Eq. (4) matters in case these represent different fields. Expanding Eq. (4), we find masses for all the fermionic components of Φ_1 and Φ_2 , and for the scalar components of Φ_2 only. The scalar components of Φ_1 remain massless.

$$\frac{\mu_{\phi_2}}{2} (\tilde{\phi}_1 \tilde{\phi}_2 - 2F_{\phi_1} \phi_2) \rightarrow \frac{\mu_{\phi_2}}{2} \tilde{\phi}_1 \tilde{\phi}_2 + |\mu_{\phi_2}|^2 |\phi_2|^2, \\ \text{where } \mu_{\phi_2} = 2w_{2,\Phi_1\Phi_2} \frac{D}{M_m}, \quad (6)$$

where ϕ_i , $\tilde{\phi}_i$, and F_{ϕ_i} are the scalar, fermion, and auxiliary components of the chiral multiplet Φ_i respectively.

A nonzero value of either or both of $w_{2,H_u H_d}$, or $w_{2,H_d H_u}$ generates masses for the Higgsinos. A nice feature of these Higgsino masses is that the masses are naturally of the order of the gaugino masses and are sourced by a single mass scale (i.e., VEV of the D component of the hidden sector field). These new operators are also phenomenologically important. Equation (6) implies that unlike the conventional μ term, $w_{2,H_u H_d}$ only gives rise to down-type Higgs soft masses. The general contributions to the Higgs sector from these unconventional operators (with both $w_{3,H_u H_d}$ and $w_{3,H_d H_u}$) are then characterized by not one μ parameter, but rather by two separate mass parameters (namely, μ_u and μ_d):

$$\frac{1}{2} (\mu_u + \mu_d) \tilde{H}_u \tilde{H}_d + |\mu_u|^2 |h_u|^2 + |\mu_d|^2 |h_d|^2. \quad (7)$$

Only in the limit $\mu_u = \mu_d = \mu$, the mass terms become identical to that of the conventional μ term. A large mass term for H_d , will result in large $\tan\beta$ but a potentially natural spectrum. It is, therefore, possible to consider a model in which the Higgsinos and additional scalar bosons are substantially heavier than the Higgs bosons without fine-tuning. This setup also challenges the conventional wisdom regarding fine-tuning in models of weak scale SUSY. Since there is no observable that directly gives a measure of the messenger scale of the theory (and the size of the large logarithmic contribution to the Higgs mass), measuring masses of the Higgsinos seems to be the best way of estimating the size of cancellation needed in order to produce the electroweak scale. Even though exceptions were constructed, where the cancellation is the result of dynamics [57–59], not fine-tuning, (therefore, the naive interpretation of Higgsino masses being the measure of fine-tuning is incorrect) the belief remains widespread. Equation (7) provides an explicit example, where the Higgsino mass can be made large (because of large μ_d), without contributing to soft mass of the up-type Higgs boson. However, too large a $(|\mu_d|^2 - |\mu_u|^2)$, generates a log divergent, though loop suppressed Hypercharge D term, which, if too large, can give some scalars tachyonic masses [60]. Also, $\mu_u \neq \mu_d$, can give rise to additional log divergent contributions to scalar soft masses². For consistency, we assume that all terms which are needed for renormalization are present, and so, in the case $\mu_u \neq \mu_d$, squark and slepton masses squared must also receive non supersoft contributions, however such terms can naturally be smaller than the supersoft contributions.

The operator in Eq. (4), with Φ replaced by the Σ_i fields, can also provide potential solutions associated with the scalar adjoints. Operators with w_{3,Σ_i^2} generate positive definite squared masses for the scalar components and Majorana masses for the fermionic components of the Σ_i fields.

$$\frac{1}{2}M_{N_i}\psi_i^2 + \frac{1}{2}|M_{N_i}|^2|\sigma_i|^2, \quad M_{N_i} = 2w_{2,\Sigma_i^2} \frac{D}{M_m}. \quad (8)$$

Color breaking can be easily avoided (at tree level) for large enough $w_{3,3}$. As mentioned earlier, the gaugino mediated contributions to scalar soft masses at one loop are already positive definite.

An additional effect of the large masses for the σ fields is the (partial) recovery of the Higgs quartic coupling. Take, for example, the on shell Lagrangian in the presence of the σ_2 fields, and the effective Lagrangian after the real components of σ_2 are integrated out:

$$\mathcal{L}_{\text{on shell}} \supset \frac{1}{2} \left(2M_{D_2}\sigma_{2_R} + \frac{g_2^2}{2} \sum_k q_k^* t_a q_k \right)^2 + \frac{1}{2}M_{N_2}(\sigma_{2_R}^2 + \sigma_{2_I}^2), \quad (9)$$

$$\mathcal{L}_{\text{eff}} \supset \frac{M_{N_2}^2}{M_{N_2}^2 + 4M_{D_2}^2} \frac{g_2^2}{8} \sum_k (q_k^* t_a q_k)^2. \quad (10)$$

We use the notation σ_{2_R} and σ_{2_I} to designate the real and the imaginary parts of σ_2 . Equations (9) and (10) are also useful for demonstrating the fact that unlike in the MSSM, D terms of the gauge fields do not contribute to the Higgs quartic in supersoft SUSY. Since the mass term M_{N_2} gets generated only by the operator in Eq. (8), the supersoft limit can be achieved by taking $M_{N_2} \rightarrow 0$, when the D term containing the Higgs quartic vanishes. In the opposite limit, namely $M_{N_2} \gg M_{D_2}$, one recovers the full MSSM strength quartic at the tree level.

The gauginos are no longer Dirac particles once the operators of Eq. (8) are included. For instance, the gluinos \tilde{g} and their Dirac partners ψ_3 obtain masses from two independent sources:

$$\mathcal{L}_{\text{gluinos}} \supset \frac{1}{2} \begin{pmatrix} \tilde{g} & \psi_3 \end{pmatrix} \begin{pmatrix} 0 & M_{D_3} \\ M_{D_3} & M_{N_3} \end{pmatrix} \begin{pmatrix} \tilde{g} \\ \psi_3 \end{pmatrix}. \quad (11)$$

Based on the relative strength of the Dirac mass of gluino and the Majorana mass of ψ_3 , three qualitatively distinct IR spectra emerge: (i) $M_{N_3} \gg M_{D_3}$: the gluino mass matrix has the “seesaw” texture. The ψ_3 field (in fact, the entire Σ_3 superfield) is integrated out at the scale M_{N_3} . The resultant light gluino (light with respect to M_{N_3}) is a Majorana fermion with a mass inversely proportional to M_{N_3} . The IR effective theory below M_{N_3} is the MSSM, with an added

feature of all scalar masses being still supersoft—in the sense that these masses do not get big log contribution from UV scales (although they are sensitive to $\log M_{N_3}$). (ii) $M_{N_3} \ll M_{D_3}$: gluinos are “pseudo-Dirac”, with two nearly degenerate Majorana color octet fermions and a small mass splitting. (iii) $M_{N_3} \sim M_{D_3}$: gluinos are mixed Majorana-Dirac [61], with two Majorana color octet fermions and a mass splitting of order their mass. The squark–quark–(lighter) gluino coupling deviates from the usual strong coupling constant ($\alpha_s \rightarrow \alpha_s \cos^2 \theta_g$, where θ_g is the mixing angle in the gluino mass matrix). The associated squark-gluino production cross section, for example, thus contains an additional factor of $\cos^2 \theta_g$ which deviates from one at the leading order.

The neutralino and chargino mass matrices are more complicated, and we leave a complete description for future work [62]. Here, we make a few remarks. In supersoft SUSY, the gauginos, Higgsinos, and additional Higgs bosons can naturally be substantially heavier than the squarks and sleptons without fine-tuning. In fact, a charged right-handed slepton is often predicted to be the lightest supersymmetric particle (LSP) in supersoft models. This, however, is problematic since a stable slepton is not cosmologically viable. In models with a low messenger scale, the gravitino becomes the LSP, thereby resolving this issue by allowing the slepton to decay into a lepton and gravitino. Depending on the gravitino mass and the reheating scale after inflation, the gravitino may provide a cold or warm dark matter candidate.

In the scenario we provide, a mostly binolike Majorana fermion could be the LSP. If its mass is close to the mass of the right-handed charged sleptons, then it can become a thermal relic with the right density due to coannihilation [63]. Consider the case, where $M_{D_1} \ll M_{N_1}, M_{D_2}, M_{N_2}, \mu_u, \mu_d$. Since $M_{D_1} \ll M_{N_1}$, there is a potentially light mass eigenstate which is mostly a binolike Majorana fermion, which can be chosen to yield the right thermal relic abundance. The right-handed charged slepton receives loop suppressed and finite mass which, at one loop, is of the order of $(g_1/2\pi)M_{N_1}^2 \log(M_{D_1}/M_{N_1})/M_{D_1}$. We may, without affecting naturalness, add flavor universal soft slepton mass squared terms which are large enough that the right-handed slepton mass is similar in size to the bino mass.

In summary, we have shown that adding a new class of operators to models with supersoft supersymmetry breaking can offer a solution to the μ problem and have very attractive consequences. Gluinos in these models can be naturally heavy, several times the mass of the squarks, while the remainder of the sub TeV superpartner spectrum can be MSSM like, including the possibility of weakly interacting massive particles dark matter. With a heavy gluino, this scenario is less constrained by LHC searches and low energy observables than the MSSM, while still allowing a path towards unification and a dynamical solution to the hierarchy problem.

This work was supported in part by the US Department of Energy under Grant No. DE-SC0011637.

-
- [1] For a comprehensive review, see [2].
 - [2] S. P. Martin, *Adv. Ser. Dir. High Energy Phys.* **21**, 1 (2010).
 - [3] M. Dine and W. Fischler, *Phys. Lett.* **110B**, 227 (1982).
 - [4] L. Alvarez-Gaume, M. Claudson, and M. B. Wise, *Nucl. Phys.* **B207**, 96 (1982).
 - [5] L. E. Ibanez and G. G. Ross, *Phys. Lett.* **110B**, 215 (1982).
 - [6] M. Dine and W. Fischler, *Nucl. Phys.* **B204**, 346 (1982).
 - [7] L. Alvarez-Gaume, J. Polchinski, and M. B. Wise, *Nucl. Phys.* **B221**, 495 (1983).
 - [8] E. Witten, *Nucl. Phys.* **B188**, 513 (1981).
 - [9] S. Dimopoulos and S. Raby, *Nucl. Phys.* **B192**, 353 (1981).
 - [10] M. Dine, W. Fischler, and M. Srednicki, *Nucl. Phys.* **B189**, 575 (1981).
 - [11] I. Affleck, M. Dine, and N. Seiberg, *Phys. Rev. Lett.* **51**, 1026 (1983).
 - [12] E. Cremmer, S. Ferrara, L. Girardello, and A. Van Proeyen, *Nucl. Phys.* **B212**, 413 (1983).
 - [13] A. H. Chamseddine, R. L. Arnowitt, and P. Nath, *Phys. Rev. Lett.* **49**, 970 (1982).
 - [14] R. Barbieri, S. Ferrara, and C. A. Savoy, *Phys. Lett.* **119B**, 343 (1982).
 - [15] J. Polchinski and L. Susskind, *Phys. Rev. D* **26**, 3661 (1982).
 - [16] N. Ohta, *Prog. Theor. Phys.* **70**, 542 (1983).
 - [17] L. J. Hall, J. D. Lykken, and S. Weinberg, *Phys. Rev. D* **27**, 2359 (1983).
 - [18] L. E. Ibanez, *Phys. Lett.* **118B**, 73 (1982).
 - [19] A. E. Nelson and M. J. Strassler, *Phys. Rev. D* **60**, 015004 (1999).
 - [20] G. Giudice and A. Masiero, *Phys. Lett. B* **206**, 480 (1988).
 - [21] S. Dimopoulos and H. Georgi, *Nucl. Phys.* **B193**, 150 (1981).
 - [22] N. Sakai, *Z. Phys. C* **11**, 153 (1981).
 - [23] L. J. Hall, V. A. Kostelecky, and S. Raby, *Nucl. Phys.* **B267**, 415 (1986).
 - [24] H. Georgi, *Phys. Lett.* **169B**, 231 (1986).
 - [25] J. R. Ellis, S. Ferrara, and D. V. Nanopoulos, *Phys. Lett.* **114B**, 231 (1982).
 - [26] W. Buchmuller and D. Wyler, *Phys. Lett.* **121B**, 321 (1983).
 - [27] J. Polchinski and M. B. Wise, *Phys. Lett.* **125B**, 393 (1983).
 - [28] F. del Aguila, M. Gavela, J. Grifols, and A. Mendez, *Phys. Lett.* **126B**, 71 (1983).
 - [29] E. Franco and M. L. Mangano, *Phys. Lett.* **135B**, 445 (1984).
 - [30] M. Dugan, B. Grinstein, and L. J. Hall, *Nucl. Phys.* **B255**, 413 (1985).
 - [31] N. Arkani-Hamed and S. Dimopoulos, *J. High Energy Phys.* **06** (2005) 073.
 - [32] G. Giudice and A. Romanino, *Nucl. Phys.* **B699**, 65 (2004).
 - [33] N. Arkani-Hamed, S. Dimopoulos, G. Giudice, and A. Romanino, *Nucl. Phys.* **B709**, 3 (2005).
 - [34] A. Arvanitaki, M. Baryakhtar, X. Huang, K. van Tilburg, and G. Villadoro, *J. High Energy Phys.* **03** (2014) 022.
 - [35] J. A. Evans, Y. Kats, D. Shih, and M. J. Strassler, *J. High Energy Phys.* **07** (2014) 101.
 - [36] H. Baer, V. Barger, D. Mickelson, and M. Padeffke-Kirkland, *Phys. Rev. D* **89**, 115019 (2014).
 - [37] Some viable parameter choices may still be considered natural [36,38,39], either because of cancellations in the renormalization group running, or because running from high scales is not considered.
 - [38] C. F. Berger, J. S. Gainer, J. L. Hewett, and T. G. Rizzo, *J. High Energy Phys.* **02** (2009) 023.
 - [39] J. L. Feng, *Annu. Rev. Nucl. Part. Sci.* **63**, 351 (2013).
 - [40] P. Fayet, *Phys. Lett.* **78B**, 417 (1978).
 - [41] P. J. Fox, A. E. Nelson, and N. Weiner, *J. High Energy Phys.* **08** (2002) 035.
 - [42] L. Hall and L. Randall, *Nucl. Phys.* **B352**, 289 (1991).
 - [43] M. Dine and D. A. MacIntire, *Phys. Rev. D* **46**, 2594 (1992).
 - [44] W. Beenakker, R. Hopker, M. Spira, and P. Zerwas, *Nucl. Phys.* **B492**, 51 (1997).
 - [45] A. Kulesza and L. Motyka, *Phys. Rev. Lett.* **102**, 111802 (2009).
 - [46] W. Beenakker, S. Brensing, M. Kramer, A. Kulesza, E. Laenen, and I. Niessen, *J. High Energy Phys.* **12** (2009) 041.
 - [47] A. Kulesza and L. Motyka, *Phys. Rev. D* **80**, 095004 (2009).
 - [48] W. Beenakker, S. Brensing, M. Kramer, A. Kulesza, E. Laenen, L. Motyka, and I. Niessen, *Int. J. Mod. Phys. A* **26**, 2637 (2011).
 - [49] M. Heikinheimo, M. Kellerstein, and V. Sanz, *J. High Energy Phys.* **04** (2012) 043.
 - [50] G. D. Kribs and A. Martin, *Phys. Rev. D* **85**, 115014 (2012).
 - [51] G. D. Kribs and A. Martin, *arXiv:1308.3468*.
 - [52] G. D. Kribs, E. Poppitz, and N. Weiner, *Phys. Rev. D* **78**, 055010 (2008).
 - [53] G. D. Kribs, T. Okui, and T. S. Roy, *Phys. Rev. D* **82**, 115010 (2010).
 - [54] K. Benakli and M. Goodsell, *Nucl. Phys.* **B830**, 315 (2010).
 - [55] K. Benakli and M. Goodsell, *Nucl. Phys.* **B840**, 1 (2010).
 - [56] C. Csaki, J. Goodman, R. Pavesi, and Y. Shirman, *Phys. Rev. D* **89**, 055005 (2014).
 - [57] T. S. Roy and M. Schmaltz, *Phys. Rev. D* **77**, 095008 (2008).
 - [58] H. Murayama, Y. Nomura, and D. Poland, *Phys. Rev. D* **77**, 015005 (2008).
 - [59] G. Perez, T. S. Roy, and M. Schmaltz, *Phys. Rev. D* **79**, 095016 (2009).
 - [60] We thank Andrew G. Cohen, and Martin Schmaltz for pointing this out to us.
 - [61] G. D. Kribs and N. Raj, *Phys. Rev. D* **89**, 055011 (2014).
 - [62] A. E. Nelson and T. S. Roy (to be published).
 - [63] K. Griest and D. Seckel, *Phys. Rev. D* **43**, 3191 (1991).