



## Holographic Dual of an Einstein-Podolsky-Rosen Pair has a Wormhole

Kristan Jensen<sup>1</sup> and Andreas Karch<sup>2</sup>

<sup>1</sup>*Department of Physics and Astronomy, University of Victoria, Victoria, British Columbia V8W 3P6, Canada*

<sup>2</sup>*Department of Physics, University of Washington, Seattle, Washington 98195-1560, USA*

(Received 6 August 2013; published 20 November 2013)

We construct the holographic dual of two colored quasiparticles in maximally supersymmetric Yang-Mills theory entangled in a color singlet Einstein-Podolsky-Rosen (EPR) pair. In the holographic dual, the entanglement is encoded in a geometry of a nontraversable wormhole on the world sheet of the flux tube connecting the pair. This gives a simple example supporting the recent claim by Maldacena and Susskind that EPR pairs and nontraversable wormholes are equivalent descriptions of the same physics.

DOI: [10.1103/PhysRevLett.111.211602](https://doi.org/10.1103/PhysRevLett.111.211602)

PACS numbers: 11.25.Tq, 03.65.Ud

*Introduction.*—Quantum entanglement is one of the most perplexing consequences of quantum mechanics. Two particles, created in an entangled Einstein-Podolsky-Rosen (EPR) pair, are tied together by this famous “spooky action at a distance”. While the measurement of the properties of the pair as a whole comes with the usual quantum uncertainty, measurement of any one of the two forces the wave function of the system into a particular eigenstate of the measured observable completely determining the state of the EPR partner.

Recently, it has been argued by Maldacena and Susskind [1] (MS) that quantum entanglement is secretly tied to another, so far purely theoretical, construction: nontraversable wormholes, or Einstein-Rosen (ER) bridges. Their basic argument was built around the quantum mechanics of black holes. The standard Schwarzschild metric for a spherically symmetric black hole describes not just one, but two asymptotically flat black holes. The outside of the black holes are not in direct contact with each other: it is impossible to send a signal between two outside observers, Alice and Bob, through the black hole. The two outside observers can however influence what the other one observes when jumping into the black hole. Signals sent by Alice into her black hole can determine what Bob will experience after traversing the horizon and vice versa. These are the characteristic features of an ER bridge. MS argue that the geometry of the ER bridge is the geometric manifestation of the entanglement between the two black holes. Their argument may be made more precise in the context of eternal black holes in AdS (we reproduce the Penrose diagram for such a black hole in Fig. 1), where we may bring the power of anti-de Sitter-conformal field theory (AdS-CFT) [2] to bear.

MS go on to consider more complicated scenarios. One could have produced the entangled black hole pair by pair production, very much like a more conventional EPR pair. Alternatively, an outside observer of a one-sided black hole formed by collapse could have captured the Hawking radiation emitted by the black hole over a period of time and then collapsed the photons into their own black hole.

Clearly, the original Hawking radiation is entangled with the black hole that emitted it.

While all these examples involved entanglement between black holes as well as macroscopic wormholes, MS argued that entanglement in general should be associated with wormhole formation. Individual Hawking quanta are claimed to be connected to the black hole interior via Planck-scale wormholes encoding the entanglement. When collapsing the Hawking radiation into a second black hole, all these microwormholes combine into a single macroscopic ER bridge.

At first sight, such a claim sounds preposterous. Quantum entanglement is a property of any quantum mechanical system, even when gravity is absent. Why microscopic wormholes should play a role in nongravitational systems is far from obvious. What we will demonstrate in this Letter is that, in fact, a single EPR pair in maximally supersymmetric Yang-Mills theory (SYM) in the limit of a large number of colors and large 't Hooft coupling has an equivalent description in terms of an ER bridge. In this limit, SYM has a dual holographic description [2] in terms of string theory on asymptotically anti-de-Sitter (AdS) space. In this dual description, the EPR pair formed by a single (external) quark antiquark pair is described by a single string connecting the two quasiparticles [3]. We find that there is an ER bridge on the string world sheet. The two horizons shielding the wormhole from the two asymptotic regions carry an entropy, which can be identified with the entanglement entropy of the pair. EPR and ER, at least in this theory, can indeed be seen as two equivalent descriptions of one and the same physical reality.

*The holographic EPR pair.*—A single heavy test quark in SYM is holographically dual to a fundamental string stretching from the Poincaré patch horizon to the boundary of AdS. The string endpoint on the boundary represents the quark. The action of the fundamental string is proportional to  $(\alpha')^{-1} \sim \sqrt{\lambda}$ , where  $\lambda$  is the 't Hooft coupling. Consequently, its free energy, energy, and entropy are all proportional to  $\sqrt{\lambda}$ . Clearly, these are not the properties of

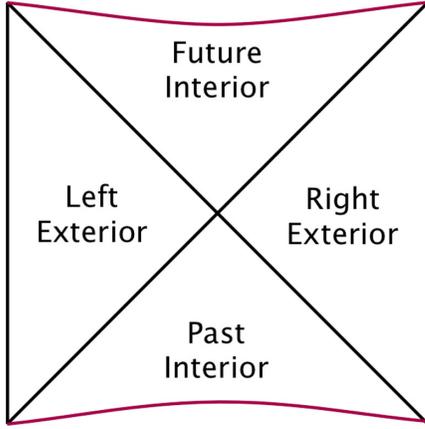


FIG. 1 (color online). The Penrose diagram of an eternal AdS black hole.

a single quark in free SYM theory. In the strongly coupled SYM, the quark is really a colored quasiparticle formed by the heavy test quark surrounded by a cloud of order  $\sqrt{\lambda}$  gluons.

A color neutral state can be formed by making a quark antiquark ( $q\bar{q}$ ) pair. By forcing the  $q\bar{q}$  state to be a color singlet, we automatically demand that the two quasiparticles are entangled. In terms of the dual string, such color singlet “meson” like states are described by a single open string with both endpoints at the boundary. String configurations dual to a separating  $q\bar{q}$  had first been numerically constructed in [5,6]. Similar numerical solutions for light quarks, this time dual to falling strings, have been obtained in [7].

The string connecting the quasiparticles is the holographic dual of the color fluxtube between the two. Unlike in a confining theory, in the scale invariant SYM, the  $1/r$  falloff of the  $q\bar{q}$  potential allows the pair to separate arbitrarily far despite the flux between them.

An analytic solution for an accelerating  $q\bar{q}$  pair was found in [8]. The geometry of this solution is depicted schematically in Fig. 2. In this solution, the quark and antiquark are accelerated so that their velocity asymptotically approaches the speed of light. This is a crucial feature, insofar as the two entangled particles are out of causal contact. This is a property this system shares with the original EPR pair. No signal emitted from particle

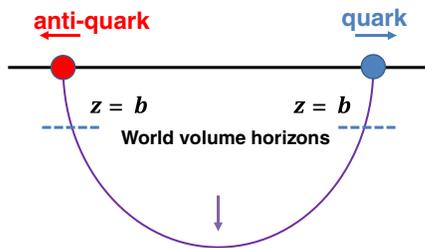


FIG. 2 (color online). The holographic  $q\bar{q}$  system entangled into a color-neutral EPR pair.

Alice can reach particle Bob in a finite amount of time. It is exactly in this situation that MS claimed that entanglement is encoded in the geometry of an ER bridge.

Using Poincaré patch coordinates in which the background AdS<sub>5</sub> metric is

$$ds^2 = \frac{R^2}{z^2}(-dt^2 + d\vec{x}^2 + dz^2), \quad (1)$$

the embedding of the string is given by the expanding semicircle

$$x^2 = t^2 + b^2 - z^2. \quad (2)$$

The quark and antiquark are located at  $x = \pm\sqrt{t^2 + b^2}$ , accelerating away from each other for all time. They initially travel toward each other, until they turn around at  $t = 0$  at  $x = \pm b$ , then fly away from each other. At late times, they go to  $x = \pm\infty$  near the speed of light. For infinitely heavy test quarks, one can simply prescribe their trajectory as an external boundary condition. If the quarks are very heavy dynamical objects, the string needs to end on a flavor probe brane [9] at a small but finite  $z_m < b$ , which is related to the quark mass by  $m = \sqrt{\lambda}/(2\pi z_m)$ . In this case, the boundary conditions on the string require a constant electric field  $E = m/b$  on the flavor brane. This electric field is responsible for accelerating the quasiparticles.

The most important aspect of the world sheet metric for us is that it has two horizons located at  $z = b$ , as indicated in Fig. 2. To understand the causal structure on the string world sheet, we have mapped out lightlike geodesics in the two-dimensional universe living on the string world sheet in Fig. 3. One look at the picture shows that this causal structure is identical to that of the eternal AdS black hole pictured in Fig. 1. The holographic dual of the EPR pair then has two horizons and an Einstein-Rosen bridge connecting them. By an ER bridge here we simply refer to a geometry with spacelike paths connecting causally disconnected regions on the world sheet.

For completeness, we elaborate briefly on the causal structure. All lightlike geodesics on the world sheet hit either the left antiquark or right quark exactly once at a time we denote as  $t_0$ . In terms of  $t_0$ , the world sheet light rays are

$$x = \pm \frac{t_0 t + b^2}{\sqrt{t_0^2 + b^2}}, \quad (3)$$

where we take the plus (minus) sign for a light ray which hits the right (left) quark. One might worry that the causal structure one obtains from the world sheet is not quite the causal structure one obtains from the bulk, pulled back to the world sheet. That is, perhaps we could send information to a point on the world sheet via bulk fields faster than we could by sending the signal along the string. However, a quick calculation shows that these world sheet geodesics

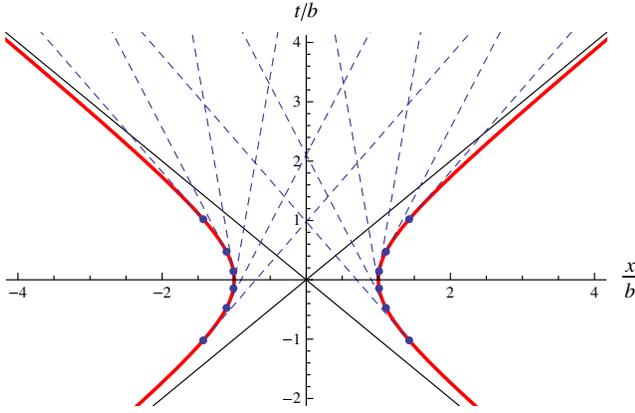


FIG. 3 (color online). The casual structure on the string world sheet. The thick red lines indicate the worldlines of the quark and antiquark, and the string world sheet fills the universe in between. The solid lines indicate the world sheet horizons, which happens to be the location of the Rindler horizons for each of the quarks. The solid dots denote events where light rays are emitted from the quark and antiquark into the dual world sheet, and the dotted lines indicate the resulting lightlike trajectories. The string world sheet clearly has the same causal structure of an eternal AdS black hole as in Fig. 1.

are precisely the bulk lightlike geodesics that propagate in the  $(t, x, z)$  directions.

At least in this particular example, we have demonstrated that the physics of entanglement in a single EPR pair is equivalently captured by the geometry of an ER bridge. The fact that the world sheet causal structure is inherited from the bulk causal structure makes it easy to see what is happening in a more general scenario. As long as the  $q\text{-}\bar{q}$  pair separates fast enough to ensure that the pair loses causal contact, we will inherit two world sheet horizons connected by an ER bridge from the bulk Rindler horizons of the accelerated charges. For other solutions, e.g., where the pair oscillates around the center of mass, the causal structure on the world sheet is trivial [10].

*Global perspective.*—There is a thermal entropy associated with the world sheet horizons. We would like to understand the relation between this thermal entropy and the entanglement of the  $q\text{-}\bar{q}$  EPR pair. To do so, we make an intermediate step and show that the accelerating string of [8] can be written as a static string stretching straight across global AdS. The horizon entropy and entanglement may be easily understood in this global picture. Recall that global AdS is the geodesic completion of the Poincaré patch. Its boundary is time cross a spatial sphere. To go from global AdS to the Poincaré patch, we represent  $\text{AdS}_{d+1}$  as a hypersurface in one higher dimension satisfying

$$-(X^0)^2 - (X^{d+1})^2 + \sum_{i=1}^d X^i X^i = -1. \quad (4)$$

We can get global AdS by defining

$$X^0 = \cosh\rho \sin\tau, \quad X^{d+2} = \cosh\rho \cos\tau, \\ X^i = \sinh\rho e^i,$$

where  $e^i$  are vectors on the  $\mathbb{S}^{d-1}$  unit sphere. These coordinates cover the entire hypersurface. Now consider the static string stretched in the  $\rho$  direction (and evolving in time  $\tau$ ), from one pole of the boundary sphere to the other:

$$X^1 = \pm \sinh\rho \cos\theta_0, \quad X^2 = \pm \sinh\rho \sin\theta_0, \\ X^{3,\dots,d} = 0. \quad (5)$$

Next, we restrict to the Poincaré patch, defining

$$X^{1,d+1} = \frac{\pm 1 - z^2 + t^2 - \vec{y}^2}{2z}, \quad X^0 = \frac{t}{z}, \\ X^2 = \frac{x}{z}, \quad X^{3,\dots,d} = \frac{\vec{y}}{z}.$$

The string sits at  $\vec{y} = \vec{0}$ , so that points on the string satisfy

$$\frac{X^1}{X^2} = \cot(\theta_0) = \frac{1 - z^2 + t^2 - x^2}{2x}.$$

Solving for  $x$  gives (up to a constant shift in  $x$ ) the expanding string of [8].

We would now like to quantify the entanglement of our  $q\text{-}\bar{q}$  pair. This is a somewhat perilous task. In a conventional EPR pair of two entangled spins, one may trace over only the spin degrees of freedom for one particle and thereby construct a reduced density matrix encoding the spin entanglement. Our setup is rather different. The entanglement is dominated by the large  $N$  color correlations between our quasiparticles so that we have a color singlet. We cannot trace over just the color states of a colored quasiparticle (which includes a cloud of  $\mathcal{N} = 4$  fields surrounding the test particle), and moreover, there is no gauge-invariant notion of a reduced density matrix encoding color entanglement. Thus, we must employ other observables to describe the entanglement. In what follows, we will content ourselves with position-space entanglement entropy (EE) for regions surrounding the quasiparticles. Our results suggest that this position-space EE is indeed describing the entanglement of the  $q\text{-}\bar{q}$  pair.

For the straight string in global AdS, it is straightforward to calculate the EE for certain regions around a single quasiparticle using a generalization [12] of the trick developed in [13]. Writing global AdS in yet another coordinate system, hyperbolic slicing, the Ryu-Takayanagi entangling surface [14] maps to a genuine horizon with a thermal entropy. The method of [12,13] shows that this entropy is equal to the EE one gets from tracing out everything below a line of constant latitude  $\theta$  on the boundary sphere (the  $q\text{-}\bar{q}$  reside at the north and south poles). By computing the string's contribution to the thermal entropy, we obtain the EE after tracing out the antiquark and everything else below  $\theta$  to be

$$S_{EE} = \frac{\sqrt{\lambda}}{3}, \quad (6)$$

notably independent of the latitude. For details of this calculation, see the Supplemental Material [15]. In fact, using the same hyperbolic slicing of the Poincaré patch, the author of [8] showed that the horizon has a temperature  $T = 1/(2\pi b)$ . From here, it is easy to compute that the thermal entropy of the world sheet horizon is also  $\sqrt{\lambda}/3$ . Furthermore, the trick of [13] shows that not only are the thermal and entanglement entropies equal, but the entire thermal and reduced density matrices are too. In the Supplemental Material [15], we consider a number of other configurations for the  $q-\bar{q}$  pair, finding that Eq. (6) is robust. We thereby argue that Eq. (6) describes the entanglement of the EPR pair.

Both the expanding string and the static string in global AdS lead to the same EE, but the expanding string has horizons and an ER bridge while the static string has neither. The primary difference between these solutions is that the expanding  $q-\bar{q}$  are out of causal contact, while the static  $q-\bar{q}$  pair are in causal contact. This is completely consistent with the claim of MS, who do not say that entanglement in general is geometrized into a wormhole, but only entanglement between causally disconnected degrees of freedom.

To summarize, the world sheet horizons of the expanding string carry a thermal entropy of  $\sqrt{\lambda}/3$ . We have argued that this is also the EE one would get from tracing over the degrees of freedom of a single quark placed on the south pole of a sphere (or [12,13] equivalently, the EE of a spherical region around a single quark). In this sense, our  $q-\bar{q}$  pair is maximally entangled like an ordinary EPR pair.

*Correlations.*—What are the physical consequences of the ER bridge? To answer this question, it is helpful to return to the EPR paradox and entanglement in field theory. In a conventional EPR pair, quantum entanglement is manifest through the violation of Bell’s inequalities. The latter may be understood as inequalities obeyed by the equal-time spin-spin correlator upon demanding local realism. However, a much simpler diagnostic of entanglement is a nonvanishing *connected* two-point function at space-like separation. Local realism would demand that the connected piece of, e.g., the equal-time spin-spin correlator vanishes. In the conventional EPR pair, the connected spin correlations may be used to quantify the spin entanglement. However, in a general context these connected correlations only signal the *existence* of entanglement.

Returning to our holographic system, we note that any connected string world sheet, with or without world sheet horizons, will give to rise nonvanishing  $O(\sqrt{\lambda})$  connected correlations signaling entanglement. What is special about the world sheet with the wormhole geometry is that, due to the horizons, all causal Green’s functions between the quark and antiquark vanish. This is expected in the field theory, as the quark and antiquark are always spacelike

separated. The ER bridge on the world sheet however still allows connected correlations between the  $q-\bar{q}$  pair, thereby signaling entanglement. Conversely, the entanglement of the causally disconnected EPR pair should result in connected  $O(\sqrt{\lambda})$  correlations between the quark and antiquark. These spacelike correlations are only possible when the dual string has a wormhole connecting the causally disconnected regions near the string endpoints.

Note that for a disconnected configuration of two strings, each of which has one endpoint at the boundary and the other end reaching behind the horizon, the connected two-point correlation function of any operator dual to a string fluctuation vanishes at order  $\sqrt{\lambda}$ . Such correlators only receive contributions from bulk graviton exchange, which is suppressed by a factor of Newton’s constant, which is small [it is parametrically  $O(1/N^2)$  in the present instance] for holography to be applicable.

*Discussion.*—We have been able to demonstrate that the holographic dual of a single EPR pair in strongly coupled SYM has the geometry of an ER bridge. This takes a step towards establishing the conjecture of MS that even for single pairs entanglement can be equivalently viewed as formation of a microscopic wormhole geometry. In our construction, this wormhole lives on the string connecting the quark and antiquark in the dual AdS spacetime. While our construction does not rely on a macroscopic wormhole geometry, our analysis requires a classical world sheet description. This limit is reliable at large  $\lambda$ , when our quark quasiparticle has a large number  $O(\sqrt{\lambda})$  of internal degrees of freedom due to the gluon cloud surrounding the quark. Our holographic EPR pair should give an ideal laboratory for further theoretical studies of the connection between entanglement, ER bridges. For instance, imagine starting with a  $q-\bar{q}$  pair at rest, accelerating one quark away from the other, then bringing it to rest a long time later. The holographically dual world sheet description will involve two-sided horizon formation, followed by evaporation.

We would like to thank S. Minwalla, A. Nelson, and A. Ritz, and especially A. Yarom for useful discussions. We also thank J. Maldacena, T. Takayanagi, and M. van Raamsdonk for comments on an earlier version of this manuscript. A. K. would also like to thank the K. M. I. at Nagoya University for hospitality during the final stages of this work. The work of K. J. was supported by NSERC, Canada, while the work A. K. has been supported in part by the U.S. Department of Energy under Grant No. DE-FG02-96ER40956.

*Noted added in proof.*—see also a related Letter by Sonner [16].

- 
- [1] J. Maldacena and L. Susskind, [arXiv:1306.0533v2](https://arxiv.org/abs/1306.0533v2).
  - [2] J. M. Maldacena, *Adv. Theor. Math. Phys.* **2**, 231 (1998).
  - [3] A somewhat similar description of an EPR pair in terms of an AdS-BCFT construction has recently been given in [4].

- [4] M. Fujita, T. Takayanagi, and E. Tonni, *J. High Energy Phys.* **11** (2011) 043.
- [5] C. Herzog, A. Karch, P. Kovtun, C. Kozcaz, and L. Yaffe, *J. High Energy Phys.* **07** (2006) 013.
- [6] M. Chernicoff and A. Guijosa, *J. High Energy Phys.* **06** (2008) 005.
- [7] P. M. Chesler, K. Jensen, and A. Karch, *Phys. Rev. D* **79**, 025021 (2009).
- [8] B.-W. Xiao, *Phys. Lett. B* **665**, 173 (2008).
- [9] A. Karch and E. Katz, *J. High Energy Phys.* **06** (2002) 043.
- [10] For the closely related case of a holographic  $q$ - $\bar{q}$  pair moving in global AdS, the world sheet casual structure of the general solution has been worked out in [11].
- [11] M. Chernicoff and A. Paredes, *J. High Energy Phys.* **03** (2011) 063.
- [12] K. Jensen and A. O'Bannon, [arXiv:1309.4523](https://arxiv.org/abs/1309.4523).
- [13] H. Casini, M. Huerta, and R. C. Myers, *J. High Energy Phys.* **05** (2011) 036.
- [14] S. Ryu and T. Takayanagi, *Phys. Rev. Lett.* **96**, 181602 (2006).
- [15] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevLett.111.211602> for a derivation of the entanglement entropy in Eq. (6), as well as a discussion of other more complicated scenarios involving our quark-antiquark pair.
- [16] J. Sonner, following Letter, *Phys. Rev. Lett.* **111**, 211603 (2013).