Volpe et al. Reply: In our Letter [\[1](#page-0-0)] we have studied a Brownian particle diffusing in front of a horizontal wall, whose movement  $z(t)$  is described by a stochastic differential equation (SDE) with multiplicative noise. This SDE can be solved according to various rules parametrized by  $\alpha$ , e.g.,  $\alpha = 0$  (Ito), 0.5 (Stratonovich), and 1, with different properties. This leads, in particular, to different predictions for the equilibrium distributions  $p(z)$  and drift fields  $d(z) = \langle \frac{\Delta z(z)}{\Delta t} \rangle$  depending on  $\alpha$  [\[2\]](#page-0-1). We can compare such predictions with our measurements in order to determine the correct value of  $\alpha$  (which turns out to be  $\alpha = 1$ ). We remark that this is only possible because we have *a priori* knowledge of the force  $F(z)$  acting on the particle, which, in particular, amounts to gravity (buoyancy) for  $z > 280$  nm [Fig. [1\(a\)](#page-0-2)], a fact overlooked by Mannella and McClintock in their Comment [\[3](#page-0-3)] because not clearly stated in our Letter [\[1\]](#page-0-0). The details follow.

 $F(z)$  has two components: gravity (buoyancy)  $-G_{\text{eff}} =$  $-5.9$  fN, which pushes the particle towards the wall and is constant for all z, and electrostatic interactions  $F_{el}$  =  $Be^{-\kappa z}$ , which prevent the particle from sticking to the wall and decays away from the wall with  $\kappa^{-1} = 18$  nm. For  $z > 280$  nm, in particular, there is only effective gravity, which is known without any fitting parameter [black line in Fig.  $1(a)$ ]. We remark that these forces act on the particle independently of the presence of Brownian noise; i.e., they would also be present if the Brownian noise were switched off, for example, by decreasing the temperature of the system towards  $T = 0$ . The corresponding SDE is

$$
dz = -\frac{F(z)}{\gamma(z)}dt + \sqrt{2D(z)}dW,\tag{1}
$$

<span id="page-0-4"></span>where  $D(z)$  is the position-dependent diffusion coefficient  $D(z)$ ,  $\gamma(z)$  the particle friction coefficient, and W a Wiener process. The corresponding  $\alpha$ -dependent  $p(z)$  and  $d(z)$  can be related to force measurements as explained in Refs. [\[1](#page-0-0)[,2](#page-0-1)]. In particular, the values of

$$
F_p(z) = \frac{k_B T}{p(z)} \frac{dp(z)}{dz}
$$
 (2)

(solid red line) and

$$
F_d(z) = \gamma(z)d(z) \tag{3}
$$

(dashed blue line), which have units of force, are shown for  $\alpha = 0$  in Figs. [1\(b\)](#page-0-2) and [1\(c\),](#page-0-2)  $\alpha = 0.5$  in Fig. [1\(d\)](#page-0-2) and [1\(e\)](#page-0-2), and for  $\alpha = 1$  in Fig. 1(f) and [1\(g\)](#page-0-2). We remark that, even though  $F_p(z)$  and  $F_d(z)$  are clearly different,  $F_d(z) - F_p(z)$ is independent from  $\alpha$ , as correctly pointed out by Mannella and McClintock [\[3\]](#page-0-3).

In Ref. [\[1](#page-0-0)], we experimentally measured  $F_p(z)$  (red squares) and  $F<sub>d</sub>(z)$  (blue circles) [Figs. [1\(b\)](#page-0-2)–[1\(e\)](#page-0-2)]; there is agreement with the solution of the SDE ([1](#page-0-4)) for  $\alpha = 1$ [Figs. 1(f) and [1\(g\)\]](#page-0-2), while the cases of  $\alpha = 0$  and 0.5



<span id="page-0-2"></span>FIG. 1 (color online). (a) Force acting on the Brownian particle studied in Ref. [[1](#page-0-0)]; the black line highlights the region where only gravity (buoyancy) acts. (b)–(g) Expected  $F_p(z)$  (solid red line) and  $F<sub>d</sub>(z)$  (dashed blue line) for  $\alpha = 0, 0.5,$  and 1; the symbols represent the experimental measurement of  $F_p(z)$  (red squares) and  $F_d(z)$  (blue circles) (from Fig. 2 in Ref. [[1](#page-0-0)]).

show clear deviations [Figs.  $1(b)-1(e)$ ]. This has the consequence that the force  $F(z) = F_p(z)$  and  $F(z) =$  $F_d(z) - \gamma(z) \frac{dD(z)}{dz}$ , which is the main result of Refs. [\[1](#page-0-0)[,2\]](#page-0-1). Finally, we remark that for other systems, which are not coupled to a heat bath [\[4](#page-0-5)], the relations between  $F(z)$ ,  $F_p(z)$ , and  $F_d(z)$  may be different.

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- <span id="page-0-3"></span>[3] R. Mannella and P.V.E. McClintock, preceding Comment, Phys. Rev. Lett. 107, 078901 (2011).
- <span id="page-0-5"></span>[4] Some examples of such systems are given, e.g., in R. Kupferman, G. A. Pavliotis, and A. M. Stuart, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevE.70.036120) E 70[, 036120 \(2004\)](http://dx.doi.org/10.1103/PhysRevE.70.036120); P. Ao et al., [Complexity](http://dx.doi.org/10.1002/cplx.20171) 12, 19 [\(2007\)](http://dx.doi.org/10.1002/cplx.20171), and references therein.