Complete Temporal Characterization of Asymmetric Pulse Compression in a Laser Wakefield

J. Schreiber,^{1,*} C. Bellei,^{1,†} S. P. D. Mangles,¹ C. Kamperidis,^{1,‡} S. Kneip,¹ S. R. Nagel,¹ C. A. J. Palmer,¹

P. P. Rajeev,² M. J. V. Streeter,¹ and Z. Najmudin¹

¹Blackett Laboratory, Imperial College, London SW7 2AZ, United Kingdom

²Central Laser Facility, Rutherford-Appleton Laboratory, Chilton, Oxon, OX11 0QX, United Kingdom

(Received 3 August 2010; published 2 December 2010)

We present complete experimental characterization of the temporal shape of an intense ultrashort 200-TW laser pulse driving a laser wakefield. The phase of the pulse was uniquely measured by using (second-order) frequency-resolved optical gating. The pulses are asymmetrically compressed and exhibit a positive chirp consistent with the expected asymmetric self-phase-modulation due to photon acceleration or deceleration in a relativistic plasma wave. The measured pulse duration decreases linearly with increasing length and density of the plasma, in quantitative agreement with the intensity-dependent group velocity variation in the plasma wave.

DOI: 10.1103/PhysRevLett.105.235003

PACS numbers: 52.38.Hb, 42.65.Jx, 42.65.Re

Laser wakefield accelerators have now demonstrated the production of quasimonoenergetic electron bunches to GeV-scale energies over only \sim cm lengths [1,2]. Though remarkably short for an accelerator, these distances are still many times the Rayleigh length z_R of the driving laser pulse, so that the nonlinear interplay between laser and plasma becomes important. This leads to interesting phenomena such as self-focusing [3,4], self-phase-modulation [5,6], or photon acceleration [7] and pulse shortening [8,9]. These processes can result in a^2 (intensity) amplification of the laser pulse [2,8]. Here $a = eE/(m_e\omega_0 c)$ is the normalized vector potential for a laser with electric field E and frequency ω_0 . Laser wakefields are driven by the ponderomotive force of the laser $F_p = -\frac{1}{\gamma}m_ec^2\nabla a^2$, where $\gamma^2 =$ $1 + a^2/2$ is the relativistic factor due to the transverse quiver of electrons in the laser field. Hence, determining the evolution of a^2 is of vital importance in understanding wakefield accelerators. For example, a^2 amplification was essential in the first demonstrations of monoenergetic beam production [10].

For the interaction to extend to many Rayleigh lengths, the laser must be prevented from diverging. This can be through the action of an external guiding structure [1] or simply through ponderomotive and relativistic self-focusing [2]. For sufficiently intense ultrashort laser pulses (a > 1), the laser will quickly focus to a matched spot radius $w_m \approx \frac{1}{\pi} \sqrt{a_0} \lambda_p$, where $\lambda_p \equiv 2\pi c/\omega_p$, with a_0 the peak vacuum value of a [4,11]. Once matched, transverse variations in spot size play little role in increasing a^2 .

However, as the laser sits in the density variation created by the plasma wave, it experiences a varying refractive index $\eta = [1 - \omega_p^2/(\gamma \omega_0^2)]^{1/2}$. At the front of the plasma wave, the density increase as plasma electrons are pushed forward is mostly compensated by the intensity dependence of η (through the dependence on γ) [12]. However, beyond this, both density decrease and further increase of γ cause η to increase. This results in an increase in group velocity $v_g \approx \eta c$ and thus compression of the pulse increasing towards the back of the plasma wave.

Associated with pulse compression must be an increase in spectral bandwidth. At the front of the pulse, density and intensity variations cancel to limit the amount of frequency shifts. However, beyond the first maximum of the plasma wave, both decreasing density and increasing intensity cause an increasing η and thus a reciprocal decrease in phase velocity v_{φ} . Hence, the majority of the pulse can be redshifted. For longer pulses $(c\tau \approx \lambda_p)$, both increasing plasma density and relativistic self-phase-modulation (decreasing γ) act together to cause strong blueshifts at the back of the pulse.

At high intensity and including multidimensional effects, this picture is complicated further. The plasma wave becomes nonlinearly steepened, and the center of the plasma wave evolves into a cavitonlike structure with relatively small variations in η within it [13]. As a result, laser energy mostly moves forward towards the front of the caviton, where it is rapidly redshifted and thus slips quickly back within the wave frame. This can cause pulse steepening at the *front* of the pulse and reduced steepening at the rear.

In any case, the outcome is generally compression of the laser pulse and possible increase in a^2 . Along with the exact pulse shape, this will determine F_p and thus wakefield growth. Compression of a 30 fs pulse to [9] ~10 fs was reported previously with a second-order autocorrelator. However, this method does not give phase information and thus does not allow a complete diagnosis of the pulse shape. In this Letter, we report on the complete temporal (amplitude and phase) characterization of short ($\tau \approx$ 45 fs), relativistically intense ($I \approx 2 \times 10^{19} \text{ W cm}^{-2}$) laser pulses after interaction with dilute plasmas such that the initial pulse was shorter than the excited plasma wave wavelength ($c\tau < \lambda_p$). Spectral broadening, photon acceleration, and asymmetric pulse shortening have been observed. The dependence of these nonlinear effects on plasma density and interaction length have been investigated, detailing the first parametric investigation of temporal pulse evolution in a laser wakefield.

The experiment was conducted with the Astra-Gemini laser at the Rutherford Appleton Laboratory, which has $\lambda_L = 800$ nm, energies up to 12 J, and duration of $\tau_0 =$ 45 fs. The 200 TW laser pulses were focused (f/20) to a spot size of $w_0 = (22.0 \pm 0.6) \ \mu \text{m}$ FWHM, with a corresponding confocal parameter of $z_R = 1$ mm. For these parameters, $a_0 \approx 3.0$ in vacuum. Three different gas jets of length l = 4, 6, and 8.5 mm were investigated for electron densities $n_e = 1.2$ –6.6 × 10¹⁸ cm⁻³. n_e and lwere determined by interferometry with a transverse optical probe [2]. Under these conditions the intensity modification was not sufficient to produce electron beams via self-injection.

The transmitted pulses were collimated by a large aperture (f/10) spherical mirror and imaged to confirm the quality of self-guiding. The full beam was also focused onto the slit of a spectrometer sensitive over 300-1000 nm that was absolutely calibrated by using a white light source. An absolute photodiode measured the transmitted laser energy. The temporal dependence of intensity and phase were measured by using second-order frequencyresolved optical gating (FROG) based on second harmonic generation (SHG). A small part of the transmitted laser (close to the center of the beam) was guided into a SHG FROG device (GRENOUILLE [14]) with 12 fs temporal and 4 nm spectral resolution. The amount of glass in the beam (vacuum windows, neutral density filters, and the recollimation lens) was kept to a minimum and did not exceed 4.5 mm in total.

In SHG FROG, a second-order autocorrelation is spectrally dispersed in the direction vertical to the time axis. The resulting 2D map—the FROG trace—thus contains both amplitude and phase information of the laser pulse. This allows the pulse shape to be uniquely determined, which is impossible from the autocorrelation alone. The pulse shape is established by using a multiparametric fitting algorithm, referred to as retrieval (FROG3 [15]). In the retrieval, the temporal axis is discretized, and, starting from an initial guess, the amplitude and phase of the laser at each point are varied until the measured FROG trace is best fitted by the synthetic one. The error is defined as the square root of the sum of the squared differences in each pixel of the measured and synthetic FROG.

Experimentally obtained SHG FROG traces can be asymmetric with respect to the time axis due to unwanted spatial chirp and pulse front tilt. Although we did not observe appreciable asymmetries, the FROG traces were symmetrized prior to the retrieval in order to improve the contrast and access a larger data set. The analysis was restricted to retrievals with an error smaller than 2% and for which the total spectrum measured in the spectrometer was in good agreement with the retrieved spectrum. This gives confidence that the FROG results are representative of the complete pulse and that no limitation due to spectral clipping occurred. Finally, fields were corrected for propagation through glass, vacuum window, and filters in the beam path. The pulse was attenuated by reflection off glass wedges before passing through any material so that only linear dispersion in the optics need be accounted for.

The complex laser fields are most usefully displayed in the form of a Wigner distribution $W(t, \omega)$ [16], the distribution of photons in (ω, t) phase space. The marginals of the Wigner distribution, $\int W(t, \omega)d\omega$ and $\int W(t, \omega)dt$, give the temporal intensity and spectrum of the pulse, and its first normalized moment gives the instantaneous frequency. Therefore, in a single 2D map it is possible to visualize the most relevant information about the pulse.

The remaining uncertainty of the direction of time due to the symmetric nature of the SHG FROG was removed by physical considerations. By measuring the FROG trace of the pulse in vacuum with an *additional* 4 mm of glass in the beam path [Fig. 1(a)] so that we know the pulse was positively chirped, the ambiguity is removed. Figure 1(e) shows the Wigner distribution after correcting for the glass. We obtain an essentially transform-limited reference pulse with its temporal profile shown in black. The wings in the profile are a common artifact in high power ultrashort laser systems and have little influence on the physical processes to be described. By varying the parameters slowly, so that the transition from the previous conditions was smooth, the direction of time could then be determined for each retrieved pulse.

The pulse was measured after propagation through a plasma with relatively constant density $n_e \in 1.2-2.3 \times 10^{18} \text{ cm}^{-3}$, low enough that the laser $(c\tau \sim 13.5 \ \mu\text{m})$ did not extend beyond the first plasma wave period ($\lambda_p \in 30-22 \ \mu\text{m}$). For the shortest length l = 4 mm [Figs. 1(b) and 1(f)], we observe a positive chirp as expected. Noticeable, though, is the faster steepening at the back of the pulse (at positive times). The total pulse length is reduced. For a longer l = 6 mm [Figs. 1(c) and 1(g)],



FIG. 1 (color online). Length scan at low electron densities $n_e(1.2-2.3) \times 10^{18} \text{ cm}^{-3}$. FROG traces (first row) and Wigner distributions of retrieved and corrected fields (second row) for l = 0, 4, 6, and 8.5 mm from left to right. Overlaid black lines correspond to temporal intensity distribution. t < 0 represents the front of the pulse.

the front is also depleted, and for l = 8 mm [Figs. 1(d) and 1(h)], the pulse becomes near-transform limited with a reduced pulse duration ($\tau = 27 \pm 3$ fs).

Similar compression is observed for increasing n_{ρ} for a fixed *l*. Figure 2 shows the Wigner distributions and pulse profiles for the shortest interaction length (l = 4 mm) but this time for n_e increasing from 2.3 to 6.6×10^{18} cm⁻³ $(\lambda_p \in 22-13 \ \mu \text{m})$. The lowest density $n_e = 2.3 \times$ 10^{18} cm⁻³ [Fig. 2(a)] corresponds to the situation described in Figs. 1(b) and 1(d). With increasing n_e , the pulses show increased bandwidth and shorter pulse duration. In contrast with the length scan at low density, the redshift at the leading edge of the pulse is now more prominent. At $n_e = 6.6 \times 10^{18} \text{ cm}^{-3}$, the pulse is again almost transform limited with a broadened spectrum that supports the pulse duration $\tau = 18 \pm 3$ fs. The pulses also exhibit enhanced structure with increasing density, with additional pre- and postpulses observed. As a guide, the duration of one plasma period is indicated by the vertical lines at the top of each panel in Fig. 2. The appearance of multiple pulses and pulselets indicates the modulation of wings of the initial pulse by the plasma wave.

Figures 3(a) and 3(b) quantify the pulse shortening. For constant density $n_e \sim 1.8 \times 10^{18}$ cm⁻³, the FWHM duration reduces linearly from $\tau = 44$ to 27 fs when the plasma length is stepwise increased to l = 8.5 mm. The rate of compression is $\approx 2.0 \pm 0.7$ fs/mm. Similarly, pulses are shortened from 44 to 18 fs for a constant l = 4 mm and n_e increased from 2.3 to 6.6×10^{18} cm⁻³, giving a constant compression rate of $\approx 4.4 \pm \frac{1.8}{1.2}$ fs/(10¹⁸ cm⁻³).

Figures 3(c) and 3(d) show the variation in energy transmission E_L obtained from two complementary measurements: the integrated spectrum and the absolutely calibrated diode. E_L decreases, with increasing l and n_e . Denoting the normalized temporal pulse profile as $f(l, n_e, t)$ ($l = n_e = 0$ being the reference pulse), the transmitted energy $E_L(l, n_e) = \int P_p(l, n_e)f(l, n_e, t)dt$ can then be used to calculate the relative peak power



FIG. 2 (color online). Pulse profiles for $n_e = (a) 2.3$, (b) 3.8, (c) 4.7, and (d) $6.6 \times 10^{18} \text{ cm}^{-3}$ for l = 4 mm. The black vertical lines mark the plasma period $(2\pi/\omega_p)$, i.e., expected positions of electron density maxima for the respective densities.

$$\frac{P_p(l, n_e)}{P_p(0, 0)} = T(l, n_e) \frac{\int f(l, n_e, t) dt}{\int f(0, 0, t) dt},$$
(1)

where $T(l, n_e) = E_L(l, n_e)/E_L(0, 0)$ is the energy transmission. Figures 3(c) and 3(d) show a substantial initial drop in P_p to $\approx 30\%$ of its original value for l = 4 mm. This can be attributed to the initial self-focusing, which traps only a fraction of the original laser energy in the matched guided pulse, especially for nonideal beams [4]. However, for further propagation, the peak power changes only slowly over the range investigated. Under these conditions, pulse compression compensates for the energy loss of the pulse in continuously driving the wakefield.

The pulse compression can be understood by considering two points in the pulse separated by $c\tau$ having different group velocities v_{g1} and v_{g2} . Their separation changes after propagating in the z direction over a distance Δl according to $c\Delta \tau / \Delta l = (v_{g2} - v_{g1}) / c = (\eta_2 - \eta_1) \approx c \tau_0 \frac{\partial \eta}{\partial z}$, where τ_0 is the initial pulse duration. As a first approximation, $\frac{\partial \eta}{\partial z}$ can be considered to be constant over the part of the plasma period in which the majority of the laser energy sits for a pulse of duration $c\tau \approx \lambda_p$. Hence $\frac{\partial \eta}{\partial z} \approx (\eta_{\min} - \eta_{\max})/c\tau$. For $n_0 \ll n_c$, $\eta \approx 1 - \frac{1}{2}(n_e/\gamma n_c)$, where $n_c = m\varepsilon_0 \omega^2/e^2$ is the critical density. Hence for intense pulses a > 1, $\eta_{\text{max}} \approx$ 1, and $\eta_{\min} \approx 1 - (n_{e0}/2n_c)$. This gives a variation of pulse duration on initial density n_{e0} and l, $\tau = \tau_0 - \frac{n_{e0}l}{2cn_c}$. For the two scans performed, we find compression rates $\Delta \tau / \Delta l =$ $-n_{e0}/2cn_c = -1.7$ fs/mm for $n_{e0} = 1.8 \times 10^{18}$ cm⁻³ and $\Delta \tau / \Delta n_0 = -l/2cn_c = -3.8$ fs/10¹⁸ cm⁻³ with l =4 mm, in good agreement with the measured rates shown in Fig. 3.

The compression of a laser pulse in a laser wakefield has been studied through simulation in 1D and 2D for similar parameters [8]. These simulations display many of the



FIG. 3 (color online). Dependence of pulse duration τ on (a) propagation length l for $n_e = (1.2-2.3) \times 10^{18} \text{ cm}^{-3}$ and (b) electron density n_e for l = 4 mm. Best fit lines give compression rates of 2.0 fs/mm and 4.4 fs/(10¹⁸ cm⁻³), respectively. (c),(d) Transmission in terms of laser energy (full circles) and peak power (open circles).

same traits as our observations, in particular, a positive chirp for short propagation distance at low density. However, to gain better insight into the pulse evolution, we consider the one-dimensional quasistatic wave equation for a laser wakefield [12]. By writing the normalized vector potential a_L in terms of a slowly varying envelope and a carrier frequency term $k_0 = \omega_0/c$, $a_L(\xi, \tau) = \frac{1}{2}a(\xi, \tau) \times \exp(ik_0\xi) + \text{ c.c.}$, the wave equation for the complex laser envelope reads

$$2\frac{\partial^2 a}{\partial\xi\partial\tau} + 2ik_0\frac{\partial a}{\partial\tau} - \frac{\partial^2 a}{\partial\tau^2} = \frac{a}{1+\Phi},$$
 (2)

where $\xi = z - ct$ and $\tau = t$ are the spatial and temporal coordinates, respectively, in the comoving frame. Φ is the electrostatic potential in the plasma, which is self-consistently obtained by solving Poisson's equation [12]

$$\frac{\partial^2 \Phi}{\partial \xi^2} = \frac{1}{2} \left[\frac{\gamma^2}{(1+\Phi)^2} - 1 \right],\tag{3}$$

where $\gamma^2 = 1 + |a|^2/2$ as before. All quantities in Eqs. (2) and (3) are normalized to the characteristic plasma parameters, i.e., space to c/ω_p and time to $1/\omega_p$. The intensity distribution a^2 and the refractive index η obtained from the numerical integration of Eqs. (2) and (3) is shown in Fig. 4.

The model shows that, at low densities, pulse compression starts from the back of the pulse, as observed experimentally. At l = 8 mm, the pulse has compressed to $\tau \approx 21$ fs, in good agreement with our measured compression. Laser energy which sits in following buckets is trapped to form trailing pulses. In contrast, the front of the pulse evolves only slowly. Since for relativistic pulses it is the front edge of the pulse which determines the wakefield amplitude, this also evolves slowly. Hence in this regime (low n_e , high a), pulse compression can extend the growth of the plasma wave to distances well beyond a naive depletion length obtained by considering only energy transfer to the wake. Furthermore, here, depletion certainly cannot be modeled by pulse front erosion [11,17].

At these densities, there is little evidence for explosive increase in a^2 predicted for propagation longer than the nonlinear modification time [8], which can be stated as $l > c\tau_{nl} \approx (c\tau_0/|a|^2)(n_c/n_e)$ [17]. For $n_c/n_e \sim$ 1000, and our initial pulse length, $c\tau_{nl} > 10$ mm, and for l = 4 mm, this time is not reached for $n_e < 6 \times$ 10^{18} cm⁻³, provided in both cases that $|a|^2$ is not too much greater than 1. This supports the measurement, shown in Fig. 3, that for our conditions only a fraction of the laser energy is captured in the wakefield driving filament beyond z_R . This also explains the good agreement with the 1D calculations and also why there is not sufficient plasma wave growth to observe self-injected electron beams below this density. Pulse evolution leading to a^2 amplification and the resultant wave breaking of the plasma wave has been inferred from the properties of generated electron beam at the high end of this density range but over longer interaction lengths [2]. Though



FIG. 4 (color online). Numerical solution of 1D wave equation in quasistatic approximation. a^2 (blue filled area), n_e (black solid line), and η (red line) resulting from propagation of a pulse with $a_0 \approx 1.5$ at $n_{e0} = 2.5 \times 10^{18}$ cm⁻³ (a) initially and (b) at l = 8 mm. The reduced a_0 was chosen to account for the initial energy loss in the guided mode as in Fig. 3.

challenging, future studies into this nonlinear regime of pulse compression, which may produce extremely short laser pulses of higher power, may prove to be particularly rewarding.

The authors thank the staff of the Central Laser Facility for their assistance during the experiments and A.E. Dangor for useful discussions. J. S. acknowledges financial support from DAAD.

*j.schreiber@imperial.ac.uk

- [†]Present address: University of California, San Diego, CA, USA.
- [‡]Present address: Technical Institute of Crete, Greece.
- [1] W. P. Leemans *et al.*, Nature Phys. **2**, 696 (2006).
- [2] S. Kneip et al., Phys. Rev. Lett. 103, 035002 (2009).
- [3] G.Z. Sun *et al.*, Phys. Fluids **30**, 526 (1987); A. B. Borisov *et al.*, Phys. Rev. Lett. **68**, 2309 (1992).
- [4] A.G.R. Thomas *et al.*, Phys. Rev. Lett. **98**, 095004 (2007).
- [5] W.B. Mori, IEEE J. Quantum Electron. **33**, 1942 (1997).
- [6] I.F. Watts et al., Phys. Rev. E 66, 036409 (2002).
- [7] S. C. Wilks *et al.*, Phys. Rev. Lett. **62**, 2600 (1989); C. D.
 Murphy *et al.*, Phys. Plasmas **13**, 033108 (2006).
- [8] D.F. Gordon et al., Phys. Rev. Lett. 90, 215001 (2003).
- [9] J. Faure et al., Phys. Rev. Lett. 95, 205003 (2005).
- [10] S. P. D. Mangles *et al.*, Nature (London) **431**, 535 (2004);
 C. G. R. Geddes *et al.*, *ibid.* **431**, 538 (2004);
 J. Faure *et al.*, *ibid.* **431**, 541 (2004).
- [11] W. Lu *et al.*, Phys. Rev. ST Accel. Beams **10**, 061301 (2007).
- [12] P. Sprangle, E. Esarey, and A. Ting, Phys. Rev. Lett. 64, 2011 (1990); Phys. Rev. A 41, 4463 (1990).
- [13] A. Pukhov and J. Meyer-Ter-Vehn, Appl. Phys. B 74, 355 (2002).
- [14] P. O'Shea et al., Opt. Lett. 26, 932 (2001).
- [15] http://www.femtosoft.biz/; K. W. DeLong and R. Trebino,
 J. Opt. Soc. Am. A 11, 2429 (1994).
- [16] E. Wigner, Phys. Rev. 40, 749 (1932).
- [17] S. V. Bulanov *et al.*, Phys. Fluids B **4**, 1935 (1992).