Enhanced Laboratory Sensitivity to Variation of the Fine-Structure Constant using Highly Charged Ions

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We study atomic systems that are in the frequency range of optical atomic clocks and have enhanced sensitivity to potential time variation of the fine-structure constant α . The high sensitivity is due to coherent contributions from three factors: high nuclear charge Z, high ionization degree, and significant differences in the configuration composition of the states involved. Configuration crossing keeps the frequencies in the optical range despite the large ionization energies. We discuss a few promising examples that have the largest α sensitivities seen in atomic systems.

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Introduction.—Theories that seek to unify gravity with the other fundamental interactions suggest that temporal variation of fundamental constants is a possibility, or even a necessity, in an expanding universe (see, e.g. [1]). Hints from quasar absorption spectra that the fine-structure constant, $\alpha = e^2/\hbar c$, may have been different in the distant past have been reported [2,3] but have not been confirmed by other groups working on a different telescope [4]. More recently, the methodology of [4] was questioned [5] and a reanalysis of the same data increased the reported error bars by a factor of 6 [6]. The observational status is therefore still unclear.

Atomic clocks provide a complementary method to search for temporal variation of fundamental constants in terrestrial laboratories [7]. Two different clocks are compared over the course of several years and any discrepancies are interpreted in terms of variation of α . For such a scheme it is critical that the two clocks have different sensitivities to any possible variation of α . The best current laboratory limit, $\dot{\alpha}/\alpha = (-1.6 \pm 2.3) \times 10^{-17} \text{ yr}^{-1}$, comes from comparison of Hg⁺ and Al⁺ optical clocks over the course of a year [8]. Here the Hg⁺ transition is strongly dependent on α , while the Al⁺ clock is practically insensitive [7,9].

The laboratory clock limits on $\dot{\alpha}$ may be improved by finding systems with enhanced sensitivity to α variation, usually denoted q. Candidates under consideration include the thorium "nuclear clock" [10], which would utilize the ~7.5 eV nuclear transition in ²²⁹Th [11]. Among other benefits, such a clock could have many orders-ofmagnitude larger q values than the optical Hg⁺ clock transition [12,13]. Optical atomic transitions with high sensitivity can be found in Yb⁺ [14] and Th³⁺ [15].

A different kind of sensitivity comes from transitions that have high relative sensitivity to α . For example, in the dysprosium atom there are two different transitions that are "accidentally" nearly degenerate and have q values with different signs [7,16,17]. Because the levels are so close together the relative sensitivity, defined by $K = 2\Delta q/\omega$ where ω is the frequency of transitions between the two levels, is extremely high (~ 10⁸). The first experiment to utilize this transition gave fairly tight limits on α variation [18], but the full enhancement was not realized because one of the levels is very broad.

In this Letter we show that both kinds of sensitivity (large q and large K) can be realized in highly ionized atomic systems. While atomic spectroscopy in electron beam ion traps is currently not competitive with optical frequency standards (see, e.g., [19,20] and review [21]) the technology continues to improve, and with the enhancements in sensitivity reported here, highly charged ions may prove to be a good system for detecting variation of α . We show, using the Ag isoelectronic sequence as an example, why high q values can occur in highly charged ions, and how the tendency of such systems towards large transition frequencies can be overcome. We then perform high-precision atomic calculations for some of the most promising ions in the sequence. In addition we identify a two-valenceelectron ion, Sm¹⁴⁺, which has optical transitions that are the most sensitive to potential variation of α ever found.

Theory.—Using a simple analytical estimate of the relativistic effects in transition frequencies, we can see that more highly charged ions have higher sensitivity to α variation. Note that a ratio of frequencies (the quantity that is actually measured) does not depend on the units one uses. In this Letter we use atomic units $e = m_e = \hbar = 1$ unless otherwise stated; in these units the atomic unit of energy is constant. Consider the relativistic corrections to the central-field Schrödinger equation for a valence electron, derived in the Pauli theory (see, e.g., [22])

$$\Delta = -\frac{\alpha^2}{2} \int R^2(r) (E_0 + V(r))^2 r^2 dr$$

+ $\frac{\alpha^2}{4} \int R(r) \left(\frac{dR}{dr} - X\frac{R}{r}\right) \frac{dV}{dr} r^2 dr,$ (1)

which is accurate to order $(\nu/c)^2$. Here R(r) and E_0 are the nonrelativistic radial wave function and energy,

respectively, V(r) is the potential, and X = j(j + 1) - l(l + 1) - s(s + 1). Near the origin V = Z/r (it is unscreened by core electrons) and the integrals (1) converge as $1/r^3$. The radial *s*-wave wave function near the origin is given in the semiclassical approximation by (see, e.g., [22])

$$R_s^2(r) \approx \frac{4Z_a^2 Z}{\nu^3}, \qquad r \lesssim \frac{3}{2Z}, \tag{2}$$

where Z_a is the effective charge that an external electron "sees" and ν is the effective principal quantum number, defined by $E_0 = -\frac{Z_a^2}{2\nu^2}$. For a single-valence electron above closed shells, $Z_a = Z_i + 1$ where Z_i is the ion charge, while for hydrogenlike ions $Z_a = Z$ and $\nu = n$, the principal quantum number. For higher waves $R_l^2(r)$ is proportional to the same parameter $Z_a^2 Z/\nu^3$.

Noting that the integrals converge over a distance $r \leq 3/(2Z)$ and neglecting the small contribution of the E_0 term in (1) (for $\nu \gg Z_a^2/Z^2$) we obtain

$$\Delta_n = -\frac{Z_a^2}{2\nu^2} \frac{(Z\alpha)^2}{\nu} = -I_n \frac{(Z\alpha)^2}{\nu(j+1/2)},$$
 (3)

where I_n is the ionization energy of the orbital. Numerical verification of this formula is presented in Fig. 1. Actually, if we neglect the E_0 term, we can immediately see from (1) that the ratio of the relativistic corrections in our "distorted wave" case to the relativistic corrections in the Coulomb case is proportional to the ratio of the squared wave functions near the origin. This again leads to Eq. (3).

In principle we could restore the E_0 term in (1), however many-body corrections are much more important than small improvements to Eq. (3). As discussed in [9], for single-valence electrons the relativistic energy shift can be approximately described by the equation

$$\Delta_n = -I_n \frac{(Z\alpha)^2}{\nu} \left[\frac{1}{j+1/2} - C(Z, j, l) \right], \qquad (4)$$

where C(Z, j, l) depends on the atom and partial wave, but does not depend strongly on the principal quantum number.



FIG. 1. Ratio $-\Delta_n/(Z^2\alpha^2 I_n)$ for the 5s level along the Ag isoelectronic sequence, calculated using the Dirac-Fock theory. It is seen that as Z increases the ratio tends to a constant (which is close to 1/5) in accordance with (3).

In practice a more sophisticated numerical treatment is needed. We characterize the dependence of transition frequencies on small changes in α by the parameter q, defined by the formula

$$\omega = \omega_0 + qx, \tag{5}$$

where ω_0 is the energy at the present-day value of the fine-structure constant α_0 and $x = (\alpha/\alpha_0)^2 - 1 \approx 2(\alpha - \alpha_0)/\alpha_0$. Our definition of *q* reflects the fact that the relativistic shift scales as α^2 .

We have shown that sensitivity to α variation increases with ion charge as $Z_a^2 = (Z_i + 1)^2$. Unfortunately, the interval between different energy levels in an ion also increases as $\sim Z_a^2$, which can quickly take the transition frequency out of the range of lasers as Z_a increases. However, the phenomena of Coulomb degeneracy and configuration crossing can be used to combat this tendency. In a neutral atom, an electron orbital with a larger angular momentum is significantly higher than one with smaller angular momentum but with the same principal quantum number n. On the other hand, in the hydrogenlike limit orbitals with different angular momentum but the same principal quantum number are nearly degenerate. Therefore somewhere in between there can be a crossing point where two levels with different angular momentum and principal quantum number can come close together: in such cases the excitation energy may be within laser range.

Consider, for example, the Ag isoelectronic sequence. Neutral Ag (Z = 47) has a single-valence electron above closed shells. The ground state has the valence electron in the 5s orbital, while the 4f orbital forms an excited level. In Fig. 2 we present calculated Dirac-Fock ionization energies of the Ag isoelectronic sequence. One can see that as Z is increased, there is a crossing point where the 4f level becomes the ground state. At this point, around Z = 61, even though both levels have ionization energies of ~270 eV, the difference between them is very much



FIG. 2. Dirac-Fock ionization energies of 5s (solid) and $4f_{7/2}$ (dashed) levels for the Ag isoelectronic sequence.

smaller. In Table I we see that the excitation energy for ions near the crossing point is well within the range of optical lasers.

Higher-Z ions with many valence electrons can have more complex behavior as ion charge is increased. For example, in the neutral Th the 5s electron is a core electron while the 5f electron is a valence electron; therefore the 5f orbital is above the 5s orbital on the energy scale. Moreover, it even lies above the 7s orbital. As seen from the experimental spectrum of the energy levels of the neutral (four-valence) Th, the energy of the $6d^27s5f$ state is above the energy of the $6d^27s^2$ state. In the hydrogenlike Th the energy of the 5f orbital is approximately equal to the energy of the 5s orbital; i.e., it is significantly lower than the energy of the 7s state. Therefore, there should be an ion charge at which the 5f orbital "crosses" the 7s orbital, while at some higher charge it crosses the 6s orbital, etc.

Results and discussion.—To calculate energy levels we use a combination of the configuration interaction and many-body perturbation theory methods (CI + MBPT) [23] which has been extensively described in previous papers [24,25]. Once we have level energies we can obtain α sensitivity by repeating the calculation for different values of α and extracting q from the gradient of the transition energies (5).

In Table I we present the results of our method for ions of the Ag isoelectronic sequence near the 5s-4f orbital crossing point. It is seen that even though the ionization energies of these levels are very large ($\sim 270 \text{ eV}$), which is reflected in the very large q values, the transitions themselves can be within the optical regime. High sensitivity to variation of α can be achieved in clocks by comparing the very sensitive $4f \rightarrow 5s$ transition in Sm¹⁵⁺ to the 4f fine-structure transition, which in this case would be the anchor. Alternatively, if it is convenient to compare frequencies between ions with high accuracy, comparison of the $4f \rightarrow 5s$ negative-shifting transition in Sm¹⁵⁺ with one of the large positive shifting $5s \rightarrow 4f$ transitions in Pm¹⁴⁺ or Nd¹³⁺ would allow an even higher sensitivity to α variation.

TABLE I. Energies and sensitivity coefficients (q) for isoelectronic sequence Nd^{13+} , Pm^{14+} , and Sm^{15+} (cm⁻¹).

Ion	Ζ	Level	Energy	q
Nd ¹³⁺	60	$5s_{1/2}$	0	0
		$4f_{5/2}$	59279	106 000
		$4f_{7/2}$	64 654	111 000
Pm^{14+}	61	$5s_{1/2}$	0	0
		$4f_{5/2}$	3973	120 000
		$4f_{7/2}$	10351	126 000
${\rm Sm}^{15+}$	62	$4f_{5/2}$	0	0
		$4f_{7/2}$	7480	7000
		$5s_{1/2}$	56272	-136000

One of the most interesting cases is Sm^{14+} . It has two valence electrons above closed shells which are in the 5s and 4f states for the low-lying configurations. The states of all three configurations 4f5s, $4f^2$, and $5s^2$ are relatively close to each other on the energy scale and are probably readily accessible to modern lasers. The transitions between these states correspond to the $4f \rightarrow 5s$ or $5s \rightarrow 4f$ single-electron transitions, which ensures strong sensitivity to the variation of the fine-structure constant. The results of calculations for Sm^{14+} are presented in Table II. We are unaware of any experimental data for this ion; therefore, only theoretical values are presented. Lifetimes are calculated using electric-dipole transitions only (Table III).

Consider, for example, the levels marked as GS (ground state), A (495 cm⁻¹) and B (28248 cm⁻¹). Both of these states have long lifetimes: the decay of level B is dominated by the E2 transition to ${}^{3}F_{2}$, which we estimate has lifetime on the order of minutes, while the level A must decay by M2/E3 transition and has a lifetime of many years unless it is quenched by the hyperfine-allowed E1 transition (the odd isotopes 147 Sm and 149 Sm have nuclear spin I = 7/2, so there are suitable transitions). Therefore, the transitions are sufficiently narrow to ensure accuracy similar to that achieved in atomic clocks. If α varies in time, states A and B will move in opposite directions. Both of these states have some of the largest q values ever seen in an atomic system.

The relative change of the ratio of the frequencies of two transitions can be written as

$$\frac{\Delta(\omega_1/\omega_2)}{(\omega_1/\omega_2)} = (K_1 - K_2)\frac{\Delta\alpha}{\alpha},\tag{6}$$

where $K = 2q/\omega$. Substituting numbers from Table II one obtains for levels A (495 cm⁻¹) and B (28 248 cm⁻¹):

$$K_A = 525, \qquad K_B = -8.8, \qquad \frac{\Delta(\omega_A/\omega_B)}{(\omega_A/\omega_B)} = 534 \frac{\Delta \alpha}{\alpha}.$$
(7)

TABLE II. Energy levels, sensitivity coefficients (q) (cm⁻¹), and lifetimes (τ) for lower states of Sm¹⁴⁺.

Configuration		J	Energy	q	au
GS	$5s4f^3F^0$	2	0	0	
	v	3	1814	987	920 s
		4	7282	6350	4.2 s
Α	$4f^{23}H$	4	495	129 898	
	-	5	6288	135 389	
		6	12067	140 213	
	$4f^{23}F$	2	9771	131 140	16 ms
		3	13746	135 303	12 ms
		4	14377	134910	35 ms
	$5s4f^{1}F^{0}$	3	13 047	6819	50 ms
	$4f^{21}G$	4	21 850	141 015	11 ms
В	$5s^{21}S$	0	28 2 48	-124689	

TABLE III. Reduced electric-dipole matrix elements for transitions between lower states of Sm^{14+} (a.u.).

Level 1		Level 2		Amplitude
Configuration	J	Configuration	J	$\langle 2 d 1\rangle$
$4f^{23}F$	2	$5s4f^{3}F^{0}$	2	0.011 800
		$5s4f^{3}F^{0}$	3	0.007480
$4f^{23}F$	3	$5s4f^3F^0$	2	0.002 030
		$5s4f^{3}F^{0}$	3	0.012 300
		$5s4f^{3}F^{0}$	4	0.006 650
		$5s4f^{1}F^{0}$	3	0.005 530
$4f^{23}F$	4	$5s4f^{3}F^{0}$	3	0.004 450
		$5s4f^3F^0$	4	0.015 500
		$5s4f^{1}F^{0}$	3	0.017 900
$4f^{21}G$	4	$5s4f^3F^0$	3	0.001 850
		$5s4f^{3}F^{0}$	4	0.006 820
		$5s4f^{1}F^{0}$	3	0.018 100
$5s4f^3F^0$	3	$4f^{23}H$	4	0.001 280
$5s4f^{1}F^{0}$	3	$4f^{23}H$	4	0.005 890
		$4f^{23}F$	2	0.001 970
$5s4f^3F^0$	4	$4f^{23}H$	4	0.001 840
-		$4f^{23}H$	5	0.000 078

This is more than 2 orders of magnitude higher relative sensitivity than that of the Hg⁺ frequency standard (K = -3.19), the system in which the current strongest constraint on the present-day time variation of α was obtained [8]. The very high sensitivity K_A comes from large q and small ω , and we should note that our theoretical transition frequencies result from a cancellation of energy levels at the 0.01%–0.1% level, and may only be accurate to perhaps ~2000 cm⁻¹, enough to reduce the relative sensitivity K_A considerably. On the other hand the q values in Table II are very stable in our calculation. Comparison of any of the lines in Table II with B will give values of $q_1 - q_2 \approx 260\,000$ cm⁻¹. This represents an absolute enhancement in α sensitivity 5 times that of the Hg⁺/Al⁺ clock comparison.

The lifetimes depend more strongly on the transition frequencies than on the dipole amplitudes. On the other hand, the accuracy of the transition frequencies may be low since they are obtained as the difference of the twoelectron removal energies of two levels. Therefore, along with the lifetimes of the levels, we also present the calculated values of the transition amplitudes (Table III). If frequencies are measured, then the lifetimes can be recalculated using these amplitudes. Note that the amplitudes are small. This is because they correspond to s - f single-electron transitions, which cannot be an electric-dipole transition. Therefore, the *E*1 transition amplitudes in the two-electron states are due to configuration mixing with suitable states, and this mixing is small.

The ideas presented in this Letter may be extended to other isoelectronic sequences, Au, Hg, etc., which would benefit from a larger Z^2 enhancement, along with the $(Z_i + 1)^2$ enhancement we have discussed. We thank J. R. Crespo López-Urrutia for useful discussions. This work was supported in part by the Australian Research Council. We thank the NCI National Facility for valuable computer time.

- [1] J.-P. Uzan, Rev. Mod. Phys. 75, 403 (2003).
- [2] J. K. Webb, V. V. Flambaum, C. W. Churchill, M. J. Drinkwater, and J. D. Barrow, Phys. Rev. Lett. 82, 884 (1999).
- [3] M. T. Murphy, J. K. Webb, and V. V. Flambaum, Mon. Not. R. Astron. Soc. 345, 609 (2003).
- [4] R. Srianand, H. Chand, P. Petitjean, and B. Aracil, Phys. Rev. Lett. 92, 121302 (2004).
- [5] M. T. Murphy, J. K. Webb, and V. V. Flambaum, Phys. Rev. Lett. 99, 239001 (2007).
- [6] M. T. Murphy, J. K. Webb, and V. V. Flambaum, Mon. Not. R. Astron. Soc. 384, 1053 (2008).
- [7] V. A. Dzuba, V. V. Flambaum, and J. K. Webb, Phys. Rev. Lett. 82, 888 (1999).
- [8] T. Rosenband, D. B. Hume, P. O. Schmidt, C. W. Chou, A. Brusch, L. Lorini, W. H. Oskay, R. E. Drullinger, T. M. Fortier, J. E. Stalnaker, S. A. Diddams, W. C. Swann, N. R. Newbury, W. M. Itano, D. J. Wineland, and J. C. Bergquist, Science **319**, 1808 (2008).
- [9] V. A. Dzuba, V. V. Flambaum, and J. K. Webb, Phys. Rev. A 59, 230 (1999).
- [10] E. Peik and Chr. Tamm, Europhys. Lett. 61, 181 (2003).
- [11] B.R. Beck, J.A. Becker, P. Beiersdorfer, G.V. Brown, K.J. Moody, J.B. Wilhelmy, F.S. Porter, C.A. Kilbourne, and R.L. Kelley, Phys. Rev. Lett. 98, 142501 (2007).
- [12] V. V. Flambaum, Phys. Rev. Lett. 97, 092502 (2006).
- [13] J. C. Berengut, V. A. Dzuba, V. V. Flambaum, and S. G. Porsev, Phys. Rev. Lett. **102**, 210801 (2009).
- [14] S.G. Porsev, V. V. Flambaum, and J. R. Torgerson, Phys. Rev. A 80, 042503 (2009).
- [15] V. V. Flambaum and S. G. Porsev, Phys. Rev. A 80, 064502 (2009).
- [16] V. A. Dzuba, V. V. Flambaum, and M. V. Marchenko, Phys. Rev. A 68, 022506 (2003).
- [17] V. A. Dzuba and V. V. Flambaum, Phys. Rev. A 77, 012515 (2008).
- [18] A. Cingöz, A. Lapierre, A.-T. Nguyen, N. Leefer, D. Budker, S. K. Lamoreaux, and J. R. Torgerson, Phys. Rev. Lett. 98, 040801 (2007).
- [19] I. Draganić, J. R. Crespo López-Urrutia, R. DuBois, S. Fritzsche, V. M. Shabaev, R. S. Orts, I. I. Tupitsyn, Y. Zou, and J. Ullrich, Phys. Rev. Lett. **91**, 183001 (2003).
- [20] J. R. Crespo López-Urrutia, Can. J. Phys. 86, 111 (2008).
- [21] P. Beiersdorfer, Phys. Scr. T134, 014010 (2009).
- [22] I.I. Sobel'man, *Introduction to the Theory of Atomic Spectra* (Pergamon, New York, 1972).
- [23] V. A. Dzuba, V. V. Flambaum, and M. G. Kozlov, Phys. Rev. A 54, 3948 (1996).
- [24] J. C. Berengut, V. V. Flambaum, and M. G. Kozlov, Phys. Rev. A 72, 044501 (2005).
- [25] J. C. Berengut, V. V. Flambaum, and M. G. Kozlov, Phys. Rev. A 73, 012504 (2006).