လ္စာ Yoctosecond Photon Pulses from Quark-Gluon Plasmas

Andreas Ipp[,*](#page-3-0) Christoph H. Keitel, and Jörg Evers^{[†](#page-3-1)}

Max Planck Institute for Nuclear Physics, Saupfercheckweg 1, D-69117 Heidelberg, Germany

(Received 29 April 2009; published 5 October 2009)

Present ultrafast laser optics is at the frontier between atto- and zeptosecond photon pulses, giving rise to unprecedented applications. We show that high-energetic photon pulses down to the yoctosecond time scale can be produced in heavy-ion collisions. We focus on photons produced during the initial phase of the expanding quark-gluon plasma. We study how the time evolution and properties of the plasma may influence the duration and shape of the photon pulse. Prospects for achieving double-peak structures suitable for pump-probe experiments at the yoctosecond time scale are discussed.

DOI: [10.1103/PhysRevLett.103.152301](http://dx.doi.org/10.1103/PhysRevLett.103.152301) PACS numbers: 25.75.Cj, 12.38.Mh, 42.65.Re, 78.47.J

Recent advances in ultrafast laser optics aim at the creation of ultrashort pulses beyond the visible spectral range. Shorter pulses are desirable, since they provide better temporal resolution, while higher photon energy gives rise to better spatial resolution. High-order harmonics of femtosecond laser radiation have been shown to be sources of trains of attosecond extreme-ultraviolet pulses [\[1,](#page-3-2)[2](#page-3-3)] that can be used to produce single attosecond soft x-ray pulses [\[3](#page-3-4),[4](#page-3-5)]. For example, such single attosecond x-ray bursts [\[5](#page-3-6)] have applications in molecular imaging [\[6,](#page-3-7)[7](#page-3-8)], quantum control [\[8](#page-3-9)[,9](#page-3-10)], or Raman spectroscopy [[10\]](#page-3-11). By introducing a controlled delay between two such peaks, the dynamics of electron systems could be studied using pump-probe techniques [\[4](#page-3-5)]. Such techniques also allow for the direct time resolution of many-body dynamics, like the observation of the dressing process of charged particles [\[11\]](#page-3-12). In view of the obvious desire for even shorter pulses with higher photon energies, it is reasonable to search for alternative production methods. In this spirit, it has been suggested that zeptosecond pulses could be created by focusing intense laser pulses on subwavelength-size structures [[12](#page-3-13)], by the reflection of a relativistically intense femtosecond laser pulse from the oscillating boundary of an overdense plasma [[13\]](#page-3-14), or via nonlinear Thomson back-scattering [\[14\]](#page-3-15).

Among the shortest possible time scales that are available experimentally are those obtained through highenergy collisions. Particularly interesting in this context are heavy-ion collisions that can produce a quark-gluon plasma (QGP), because they expose a complex system with rich internal dynamics that lives on a very short time scale. Heavy-ion collisions at the Relativistic Heavy Ion Collider (RHIC) and soon at the CERN LHC produce this new state of matter up to the size of a nucleus $(\sim 15 \text{ fm})$ for a duration of a few tens of yoctoseconds (1 ys = 10^{-24} s \approx 0.3 fm/c). In such a collision, the plasma is produced initially in a very anisotropic state and reaches a hydrodynamic evolution through internal interactions only after some thermalization time τ_{therm} and isotropization time $\tau_{\rm iso}$. The observed particle spectra turned out to agree well with ideal hydrodynamical model predictions [\[15–](#page-3-16) [17](#page-3-17)], which led to the assumption that isotropization times are as low as $\tau_{\rm iso} \approx 1$ ys $(0.3 \text{ fm}/c)$. However, it has been pointed out recently that viscous hydrodynamic models are still consistent with RHIC data if isotropization times as large as $\tau_{\text{iso}} \approx 7$ ys(2 fm/c) are assumed [\[18](#page-3-18)], even if the expansion before isotropization is assumed to be collisionless (''free streaming''). Besides a plethora of particles that are created in such collisions, high-energetic photons are also produced and detected experimentally [\[19](#page-3-19)[–21\]](#page-3-20). So far, only time-integrated quantities have been considered, like the extraction of the spatiotemporal extent of the emitting source through photon pair correlations [\[22\]](#page-3-21) and the suggestion to measure the isotropization time τ_{iso} using the direct photon yield [\[23,](#page-3-22)[24\]](#page-3-23).

In this Letter, we study the time-resolved production of high-energetic ultrashort photon pulses in the QGP. We demonstrate that the emission envelope depends strongly on the internal dynamics of the QGP. Under certain conditions, a double-peak structure in the emission envelope can be observed, which could be the first source for pumpprobe experiments at the yoctosecond time scale. We find that the delay between the peaks is directly related to the isotropization time, and the relative height between the peaks can be shaped by varying photon energy and emission angle. Such pulses could be utilized, for example, to resolve dynamics on the nuclear time scale such as that of baryon resonances [[25](#page-3-24)]. As an alternative interpretation of our results, a time-resolved study of the emitted photons could provide a window to the internal QGP dynamics throughout its expansion.

The QGP is formed in a collision of two heavy ions as illustrated in Figs. $1(a)-1(c)$. We study the emission of direct photons from the expanding QGP [[26](#page-3-25),[27](#page-3-26)]. The energy spectrum of such photons extends to the GeV range, and the upper limit for the temporal duration of the GeV photon pulse is given by the expansion dynamics of the QGP, which leads to yoctosecond pulses. Regarding the shape of the photon pulse, we find that the angle-resolved photon emission rate strongly depends on the internal state

FIG. 1 (color online). Early stages of a high-energy collision, involving preequilibrium (first two columns) and equilibrated QGP phases (last column). Parts (a)–(c) show three snapshots in time in position space. Shown are the two relativistically contracted colliding ions that create the quark-gluon plasma in the overlap region. Curly arrows denote photon emission and semicircles the detectors. Parts (d) – (f) are corresponding pictorial representations of the plasma in momentum space. In an intermediate stage of the preequilibrium phase, the momentum distribution is anisotropic, resulting in a change in the angular photon emission pattern that can give rise to double-peaked photon pulses.

of the plasma, characterized by the momentum distribution of the plasma constituents, see Figs. $1(d) - 1(f)$. For an intermediate time after the collision, a momentum anisotropy occurs due to the longitudinal expansion of the plasma, which leads to a preferential emission of photons perpendicular to the beam axis z . A photon detector placed towards the beam axis would therefore measure a timedependent photon flux. Based on a recent model [[28](#page-3-27),[29](#page-3-28)] for the internal plasma dynamics, we find that this mechanism can give rise to double-peaked pulse envelopes that could be utilized for pump-probe experiments.

We base our model on the one-dimensional expansion model by Bjorken [\[30\]](#page-3-29). This assumes boost-invariant evolution of the quark-gluon plasma in a central region of the collision. The detector is placed away from the beam axis by an angle θ within the reaction plane. Since we want to study hard photons produced in the GeV energy range, we restrict ourselves to the production of direct photons emitted from the preequilibrium and equilibrated phases of the QGP. We will not consider photons produced in later stages of the collision where, on average, lower energy photons are emitted. For the same reason, we will omit higher order soft scattering processes like bremsstrahlung or inelastic pair annihilation, whose contributions are again more important at lower energies [\[21](#page-3-20)[,23\]](#page-3-22). The leading contribution to the photon production rate R originates from quarkgluon Compton scattering and quark-antiquark annihilation [\[31\]](#page-3-30). For anisotropic momentum distributions, the photon production rate R has to be calculated numerically [\[21\]](#page-3-20). This rate, obtained in the form Ed^3R/d^3k , depends on the temperature T of the medium, the photon energy E and momentum k, the fine structure constant α , and the corresponding quantity for the strong force α_s (we use $\hbar = c$ = $k_B = 1$). The rate further depends on the anisotropy, which is described by a parameter $\xi = \langle p_T^2 \rangle / (2 \langle p_L^2 \rangle) - 1$ that relates the mean longitudinal and transverse momenta p_L and p_T [\[23](#page-3-22)[,28,](#page-3-27)[29](#page-3-28)]. To integrate this rate over time, we use a recent model to describe the time evolution of the preequilibrium and equilibrated QGP [[28](#page-3-27),[29](#page-3-28)]. This model specifies the time evolution for the energy density $\mathcal{E} = \mathcal{E}(\tau)$, for the hard scale $p_{\text{hard}} = p_{\text{hard}}(\tau)$ (which corresponds to T in the isotropic case), and for the anisotropy parameter ξ = $\xi(\tau)$ as a function of the proper time τ . These quantities scale with different powers depending on the QGP expansion dynamics. For example, the time evolution of ξ can be written as $\xi(\tau) = (\tau/\tau_0)^{\delta} - 1$. In the free-streaming phase, $\delta = 2$, while in the ideal hydrodynamic phase, $\delta =$ 0. The model in Ref. [[29](#page-3-28)] essentially introduces a smeared step function for the exponent $\delta = \delta(\tau)$ to interpolate between $\delta = 2$ at early times $\tau \ll \tau_{\text{iso}}$ and $\delta = 0$ at late times $\tau \gg \tau_{\text{iso}}$, where the duration of the transition is controlled by τ_{iso}/γ with dimensionless parameter γ . In this way, both the preequilibrium phase and the equilibrated QGP phase of the expansion can be described by a single model. In the model, thermalization and isotropization happen concurrently, $\tau_{\text{therm}} = \tau_{\text{iso}}$.

In the radial direction, we use a straightforward generalization of the temperature dependence $T(r, \tau_0) =$ $T_0[2(1 - r^2/R_T^2)]^{1/4}$ [\[33,33\]](#page-3-31) for noncentral collisions, which is valid in the overlap region of the colliding nuclei. As in Ref. [\[29\]](#page-3-28), we neglect the expansion of the QGP into transverse directions, since it is small in the initial stage throughout which the high-energetic photons are emitted. T_0 is the initial temperature, r is the distance from the center in radial direction, and R_T is the transverse radius of the nucleus. The number of photons of frequency ω_d that arrive at the detector at time t_d in the laboratory system can be obtained by integrating the photon rate along all possible light paths. The world line of the photons is parametrized as $\mathbf{z}(t; t_d, x_k, y_k) = x_k \hat{\mathbf{k}}_1^{\perp} + y_k \hat{\mathbf{k}}_2^{\perp} + \hat{\mathbf{k}}(t - t_d + d),$ where $\hat{\mathbf{k}}$ is the spatial direction of the light wave vector $k =$ (ω, \mathbf{k}) , $\hat{\mathbf{k}}_1^{\perp}$ and $\hat{\mathbf{k}}_2^{\perp}$ are two direction vectors orthogonal to k and to each other, and d is the distance to the detector. The signal f_d that arrives at the detector at time t_d is given by integrating the photon emission rate along the photon world line $f_d(k, t_d, x_k, y_k) = \int_0^{t_d} dt f(k, t, \mathbf{z}(t; t_d, x_k, y_k)),$ where $f(k, t, z)$ is the Lorentz transformed photon emission rate Ed^3R/d^3k inside the QGP and vanishing outside. One finally obtains the photon rate per unit time and frequency interval with A_d the detector area,

$$
\frac{d^2N(\omega_d, t_d)}{dt_d d\omega_d} = \frac{A_d}{8\pi d^2} \int dx_k dy_k \omega_d f_d(\omega_d, t_d, x_k, y_k). \tag{1}
$$

For the concrete calculation, we followed the parameters as given in [\[29\]](#page-3-28). We show results for LHC energies, where the initial conditions are given by a formation time τ_0 = 0.3 ys(= 0.088 fm/c), the initial temperature T_0 = 845 MeV, and transverse radius of the lead ion $R_T =$ 7:1 fm. The critical temperature, where the QGP ceases to exist, is taken as $T_c = 160$ MeV. We study the freestreaming model ($\delta = 2$) with parameter $\gamma = 2$. As in Ref. [[29](#page-3-28)], the isotropization time is varied in the range $\tau_{\rm iso} = \tau_0$ to $\tau_{\rm iso} = 6.7$ ys(= 2 fm/c). Both possibilities are not yet ruled out by RHIC data. We use the model with fixed final multiplicity, where the initial conditions are adjusted as a function of τ_{iso} in order to result in the same final entropy as for $\tau_{\text{iso}} = \tau_0$.

Figures [2\(a\)](#page-2-0) and [2\(b\)](#page-2-0) show the typical time evolution of the photon emission rate for a central collision with emission angle orthogonal to the beam axis ($\theta = \pi/2$). We present results at 2 GeV (3 GeV) energy, where the photon production from the QGP at midrapidity is 3 to 4 times as large as the production from the initial collisions and roughly 6 times (50 times) as large as the production from the hadron gas [\[27\]](#page-3-26). The origin of the abscissa is the time when a photon emitted from the center of the collision arrives at the detector. Photons arriving earlier originate from a part of the QGP that is closer to the detector. The pulse shape is mainly determined by the geometry of the nucleus with radius 7.1 fm.

FIG. 2 (color online). Photon emission rate as a function of detector time τ . Solid (blue) lines show a large isotropization time $\tau_{\text{iso}} = 6.7 \text{ ys} = 2 \text{ fm}/c \text{ with } \delta = 2 \text{ (free-streaming model)}$ while dashed (red) lines correspond to a short isotropization time $\tau_{\rm iso} = \tau_0 = 0.3$ ys = 0.088 fm/c. Parts (a) and (b) display emission at midrapidity ($\theta = \pi/2$) for a central collision with impact parameter $b = 0$. Parts (c)–(f) show double-peaked photon pulses obtained for $b = 9.2$ or 12.2 fm, and the vertical dotted line indicates the position of the larger $\tau_{\text{iso}} = 6.7$ ys. In parts (c) and (d), the detector direction is $\theta = \pi/4$; in (e) and (f), it is $\theta = \pi/8$.

Internal QGP dynamics occurs on a time scale of a few fm/c . Any structure of this order is washed out simply by the time for light to cross the size of the QGP. We apply two strategies to overcome this limit. First, we reduce the physical extent of the QGP by considering noncentral collisions with impact parameter b. Second, an optimization of the detection angle minimizes the traveling time through the plasma. In forward direction, the initial shape of the QGP is Lorentz contracted; hence, light travels through the initial shape quickly. This is partially spoiled due to the QGP expansion in the same direction. Thus intermediate emission angles are most promising for which the QGP appears partly Lorentz contracted but does not expand towards the detector.

Figures $2(c) - 2(f)$ show the photon emission in a direction away from midrapidity $\theta \neq \pi/2$ for a noncentral collision. The impact parameter $b = 9.2$ fm (b = 12:2 fm) corresponds to an overlap region of the size 5 fm \times 10.8 fm (2 fm \times 7.3 fm). Here, a striking new double-peak structure appears. Roughly speaking, the minimum between the peaks corresponds to the maximum of the anisotropy parameter $\xi(\tau)$, which grows with exponent $\delta = 2$ for $\tau < \tau_{\text{iso}}$ and vanishes for $\tau > \tau_{\text{iso}}$. This is because the photon emission rate is suppressed for larger values of ξ and smaller values of θ [\[23\]](#page-3-22). The distance between the two peaks is therefore approximately governed by τ_{iso} , indicated by the dotted line in Figs. [2\(c\)](#page-2-0)–2(f).

The dynamics shown in Fig. [2](#page-2-1) requires a more detailed interpretation. In Figs. $2(c) - 2(f)$, the first peak corresponds to photons emitted from the blueshifted approaching part of the QGP, while the second peak corresponds to a slightly redshifted and time-dilated receding tail of the plasma. The relative size of the first and second peak is controlled both by the emission angle and the photon energy. At increasing emission angles, the second peak is suppressed, because the phase space for photons emitted from the receding part of the plasma is constricted. For smaller impact parameters or larger QGP sizes, the second peak overlaps with the first, thereby washing out this structure. Also for a short isotropization time $\tau_{\rm iso} = \tau_0$ (dashed lines), the separation into two peaks does not occur. Therefore, this effect is very sensitive to the internal dynamics of the QGP.

Our results are based on a number of model assumptions. For example, we omitted the transverse expansion, which would become important at later times. It could either enhance the double-peak structure through a further separation of red- and blueshifted parts of the QGP or reduce the effect through a prolonged photon passage through the QGP. These modifications may cause qualitative changes to the photon emission envelope but not to the yoctosecond time scale. In an actual experiment, there is a background of photons from different sources [\[32\]](#page-3-32). These include photons produced by a jet passing through the QGP [\[33\]](#page-3-31) and could dominate the effect that is expected from the QGP alone. Since these photons are produced on a similar yoctosecond time scale, they can be expected to have the positive effect of enhancing the photon rate of the pulse. But at the same time, they would constitute a background for the pulse shape determination. Other background photons are produced at different time scales, e.g., due to decay of pions produced in the hadronization of the QGP; therefore, they would not modify a time structure on the yoctosecond time scale. Photons produced from the initial collisions (pQCD photons) can be of comparable size, but they would only enhance the first peak of the double peaks depicted in Figs. $2(c)-2(f)$ $2(c)-2(f)$. A better theoretical understanding of the QGP dynamics would be necessary for a more quantitative prediction of the expected pulse structure.

We now turn to an estimate of the photon production rate. In the GeV energy range, these are of the order of a few photons per collision [[32](#page-3-32)]. Note that a single GeV photon pulse of 10 ys duration corresponds to a pulse energy of only about 100 pJ but to a power of 10 TW. A single photon per pulse is in principle sufficient to reconstruct a nontrivial pulse shape [\[34\]](#page-3-33). Decreasing the photon energy would enhance the photon yield but would also increase the number of unwanted background photons. Increasing the collision energy would further enhance the number of photons produced and could also increase the relative importance of the contribution of thermal photons compared to other kinds of photons [\[32\]](#page-3-32).

Concluding, we have studied the time evolution of the photon emission from the QGP. Since the QGP only exists on the yoctosecond time scale, it naturally emits photons at this time scale. We have shown that the emission envelope can be influenced by the geometry, emission angle, and internal dynamics like the isotropization time of the expanding QGP. For a particular parameter range, that is, noncentral collisions, large isotropization time, and an emission angle close to forward direction, a double-peak structure has been found within our model. Thus, pumpprobe experiments at the GeV energy scale could be envisioned. New detection schemes would be required, and it should be explored whether tools and ideas from attosecond metrology [\[3](#page-3-4)[,35\]](#page-3-34) can be scaled to zepto- or yoctosecond duration, utilizing physical processes in the GeV energy range, like electron-positron pair creation. Alternatively, determining the photon emission shape experimentally would give direct access to dynamic properties of the QGP, like its isotropization time.

We thank K. Z. Hatsagortsyan and M. Strickland for helpful discussions.

- [2] P. Paul et al., Science 292, 1689 (2001).
- [3] M. Hentschel et al., Nature (London) 414, 509 (2001).
- [4] Y. Silberberg, Nature (London) 414, 494 (2001).
- [5] T. Baeva et al., Laser Part. Beams 25, 339 (2007).
- [6] H. Niikura et al., Nature (London) 417, 917 (2002); J. Itatani et al., Nature (London) 432, 867 (2004); T. Ergler et al., Phys. Rev. Lett. 97, 193001 (2006).
- [7] M. Lein, Phys. Rev. Lett. 94, 053004 (2005); S. Baker et al., Science 312, 424 (2006); M. Lein et al., Phys. Rev. A 66, 023805 (2002).
- [8] H. Rabitz et al., Science 288, 824 (2000).
- [9] Y. Silberberg, Nature (London) 430, 624 (2004).
- [10] N. Dudovich, D. Oron, and Y. Silberberg, Nature (London) 418, 512 (2002).
- [11] R. Huber et al., Nature (London) **414**, 286 (2001); H. Haug, ibid. 414, 261 (2001).
- [12] A. E. Kaplan and P. L. Shkolnikov, Phys. Rev. Lett. 88, 074801 (2002); W.R. Garrett, ibid. 89, 279501 (2002); A.E. Kaplan and P.L. Shkolnikov, ibid. 89, 279502 (2002); G. Stupakov and M. Zolotorev, ibid. 89, 199501 (2002); A. E. Kaplan and P. L. Shkolnikov, ibid. 89, 199502 (2002).
- [13] S. Gordienko et al., Phys. Rev. Lett. 93, 115002 (2004).
- [14] P. Lan et al., Phys. Rev. E 72, 066501 (2005).
- [15] P. Huovinen et al., Phys. Lett. B 503, 58 (2001).
- [16] T. Hirano and K. Tsuda, Phys. Rev. C 66, 054905 (2002).
- [17] M.J. Tannenbaum, Rep. Prog. Phys. 69, 2005 (2006).
- [18] M. Luzum and P. Romatschke, Phys. Rev. C **78**, 034915 (2008).
- [19] J. Adams et al. (STAR Collaboration), Phys. Rev. C 70, 044902 (2004).
- [20] S. S. Adler et al. (PHENIX Collaboration), Phys. Rev. Lett. 94, 232301 (2005).
- [21] A. Ipp et al., Phys. Lett. B 666, 315 (2008).
- [22] M. Marqués et al., Phys. Rev. Lett. 73, 34 (1994); A. Aggarwal et al., Phys. Rev. Lett. 93, 022301 (2004).
- [23] B. Schenke and M. Strickland, Phys. Rev. D 76, 025023 (2007).
- [24] L. Bhattacharya and P. Roy, Phys. Rev. C **78**, 064904 (2008).
- [25] M. Dugger et al., Phys. Rev. C 76, 025211 (2007).
- [26] K. Reygers, AIP Conf. Proc. **892**, 413 (2007).
- [27] S. Turbide, R. Rapp, and C. Gale, Phys. Rev. C 69, 014903 (2004).
- [28] M. Martinez and M. Strickland, Phys. Rev. Lett. 100, 102301 (2008).
- [29] M. Martinez and M. Strickland, Phys. Rev. C 78, 034917 (2008).
- [30] J.D. Bjorken, Phys. Rev. D 27, 140 (1983).
- [31] J. I. Kapusta, P. Lichard, and D. Seibert, Phys. Rev. D 44, 2774 (1991).
- [32] S. Turbide et al., Phys. Rev. C 72, 014906 (2005).
- [33] R.J. Fries, B. Müller, and D.K. Srivastava, Phys. Rev. Lett. 90, 132301 (2003).
- [34] M. Keller et al., Nature (London) 431, 1075 (2004); P. Kolchin et al., Phys. Rev. Lett. 101, 103601 (2008).
- [35] M. Drescher et al., Science 291, 1923 (2001).

[^{*}a](#page-0-0)ndreas.ipp@mpi-hd.mpg.de

[[†]](#page-0-0) joerg.evers@mpi-hd.mpg.de

^[1] N. Papadogiannis et al., Phys. Rev. Lett. **83**, 4289 (1999).