Search for Axionlike Particles Using a Variable-Baseline Photon-Regeneration Technique

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We report the first results of the GammeV experiment, a search for milli-eV mass particles with axionlike couplings to two photons. The search is performed using a "light shining through a wall" technique where incident photons oscillate into new weakly interacting particles that are able to pass through the wall and subsequently regenerate back into detectable photons. The oscillation baseline of the apparatus is variable, thus allowing probes of different values of particle mass. We find no excess of events above background and are able to constrain the two-photon couplings of possible new scalar (pseudo-scalar) particles to be less than $3.1 \times 10^{-7} \text{ GeV}^{-1}$ ($3.5 \times 10^{-7} \text{ GeV}^{-1}$) in the limit of massless particles.

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Recently, the PVLAS experiment reported a positive signal in a photon oscillation experiment [1]. In their "disappearance" experiment, a polarized 1064 nm laser beam was sent into an Fabry-Perot cavity which also contained a high transverse magnetic field. By measuring a net rotation of the laser polarization vector upon exiting the cavity, PVLAS found a nonzero relative attenuation of laser polarization components transverse to and parallel to the external magnetic field. This result was interpreted as evidence for the production of new spin-0 particles from one of the polarization components [2]. Newer PVLAS results reported at conferences [3] included the measurements of nonzero polarization rotation with 532 nm green light, and nonzero ellipticity at both wavelengths, where the relative phase delay is induced by loop effects.

The new particles may have scalar $g\phi(\vec{E} \cdot \vec{E} - \vec{B} \cdot \vec{B})$ or pseudoscalar $g\phi(\vec{E} \cdot \vec{B})$ interactions with photons. While PVLAS was not able to determine the parity of the new particle, their four measurements gave a consistent picture of a low scalar mass $m_{\phi} \sim 1.2$ meV and a rather large twophoton coupling $g \sim 2.5 \times 10^{-6}$ GeV⁻¹. The suggested PVLAS region of interest was not previously excluded by the pioneering BFRT laser oscillation experiment [4,5]. However, for reasons unknown, the PVLAS signal disappeared after the experimental apparatus was later rebuilt in an effort to improve the detection [6]. A number of experimental efforts worldwide have begun to test the hypothesis that the anomalous PVLAS observations are due to a new axionlike particle [7], and, in particular, the BMV experiment has recently excluded the pseudoscalar interpretation of the PVLAS signal [8,9].

New particles with masses smaller than the electron mass have yet to be directly detected, but models containing such light degrees of freedom have been motivated by the small neutrino mass differences [11,12], by the vacuum dark energy density of $(2 \text{ meV})^4$ [13], and by string theory

[14] and theories with large extra dimensions [15]. Couplings of light particles to photons are tightly constrained by star cooling considerations, including the experimental limits set by the CAST axion helioscope [16] whose recent limit $g < 8.8 \times 10^{-11}$ GeV⁻¹ strongly excludes PVLAS. Various methods have been postulated, however, to evade the star cooling limits [13,17,18].

In this Letter we report the first results from GammeV, a photon-regeneration experiment similar to that originally suggested in [19]. Our apparatus is specifically designed to quickly probe the region in parameter space suggested by the PVLAS results, at modest expense. In our "appearance" experiment, a positive signal would yield an unambiguous interpretation as oscillations of photons into new, weakly interacting particles.

Experimental design.—The key to this experiment is the short 5 ns, 160 mJ pulses of 532 nm light emitted with a repetition rate of 20 Hz by our light source, a frequency-doubled Continuum Surelite I-20 Nd:YAG laser. This small duty cycle enables a large reduction in the detector noise via coincidence counting. The laser light is vertically



FIG. 1. Diagram (not to scale) of the experimental apparatus. The initial vacuum chamber consists of a 10 m insulating warm bore which is offset by 1.6 m from the end of the 6 m magnetic field region, and is sealed to the sliding plunger via a double oring assembly. The plunger has a range of motion of 1.9 m, and contains an independent vacuum chamber. The vacuum window at the far end slides within a stationary dark box.

polarized and when needed, a half-wave plate is used to obtain horizontal polarization. The laser pulses are sent through a vacuum system (diagramed in Fig. 1) designed around an insulating warm bore inserted into a 6 m Tevatron superconducting dipole magnet. The magnet produces a 5 T vertical field uniform across the aperture of the 48 mm inner diameter warm bore. A "wall" consisting of a high-power laser mirror on the end of a long (7 m) hollow stainless steel "plunger" is inserted into the warm bore. The mirror may be placed at various positions within the magnet by sliding the plunger. The plunger mirror projects the reflected wave into a photon state and the transmitted wave into a (pseudo-)scalar state, provided that the scalar is sufficiently weakly interacting to pass through the material of the mirror. The mirror also functions to reflect the incident laser power out of the magnet to prevent heating of the magnet coils. The mirror is mounted on a welded stainless steel cap on the end of the plunger in order to prevent stray photons from passing through. Thus, the beam passing through the end of the plunger is a pure scalar beam. These scalars can then oscillate back into photons through the remaining magnetic field region inside the 35 mm inner diameter plunger. Upon exiting the magnetic field region, the interaction ceases and the photonscalar wave function is frozen. This combined wave function then propagates ~ 6 m into a dark box where a Hamamatsu H7422P-40 photomultiplier tube (PMT) module is used to detect single photons in coincidence with the laser pulses. As described below, a high signal-to-noise ratio is achieved due to the very short pulses emitted by the laser.

The photon-scalar transition probability [20] may be written in convenient units as

$$P_{\gamma \to \phi} = \frac{4B^2 \omega^2}{M^2 (\Delta m^2)^2} \sin^2 \left(\frac{\Delta m^2 L}{4\omega}\right) \tag{1}$$

$$= 1.5 \times 10^{-11} \frac{(B/\text{Tesla})^2 (\omega/\text{eV})^2}{(M/10^5 \text{ GeV})^2 (m_{\phi}/10^{-3} \text{ eV})^4} \\ \times \sin^2 \left(1.267 \frac{(m_{\phi}/10^{-3} \text{ eV})^2 (L/m)}{(\omega/\text{eV})} \right),$$
(2)

where M = 1/g, *B* is the strength of the external magnetic field, ω is the initial photon energy, and *L* is the magnetic oscillation baseline. The mass-squared difference between the scalar mass and the effective photon mass, $\Delta m^2 =$ $m_{\phi}^2 - m_{\gamma}^2 \approx m_{\phi}^2$, characterizes the mismatch of the phase velocities of the photon wave and the massive scalar wave and determines the characteristic oscillation length. While the photon does not really gain a mass in a normal dielectric medium, the phase advance may be modeled with an effective imaginary mass $m_{\gamma}^2 = -2\omega^2(n-1)$, where *n* is the index of refraction [21]. Both the warm bore and the interior of the plunger are pumped to moderate vacuum pressures of less than 10^{-4} Torr, and a conservative estimate gives $\sqrt{-m_{\gamma}^2} < 10^{-4}$ eV.

As seen in Eq. (2), the meter scale baseline provided by typical accelerator magnets is well suited for probing the milli-eV range of particle masses. This fact can be a curse as well as a boon because for a monochromatic laser beam, a fixed magnet length may accidentally coincide with a minimum in the oscillation rather than a maximum. Indeed, this may be why the BFRT experiment [5] did not see the PVLAS signal even though they had similar sensitivity. GammeV's plunger design allows us to change the oscillation baseline and thus scan through the milli-eV range of scalar mass without any regions of diminished sensitivity. The total conversion and regeneration probability contains two factors of Eq. (1), corresponding to the premirror and postmirror magnetic field regions of lengths L_1 and L_2 . The total probability varies as $\sin^2(\frac{m_{\phi}^2 L_1}{4\omega})\sin^2(\frac{m_{\phi}^2 L_2}{4\omega})$ where $L_1 + L_2 = 6$ m.

To detect regenerated photons we use a 51 mm diameter lens to focus the beam onto the 5 mm diameter GaAsP photocathode of the PMT. The alignment is performed using a low power green helium-neon alignment laser and a mock target. The alignment is verified both before and after each data-taking period by replacing the sealed plunger with an open-ended plunger, reestablishing the vacuum, and firing the Nd:YAG laser onto a flash paper target. An optical transport efficiency of 92% is measured using the ratio of laser power transmitted through the openended plunger and through the various optics and vacuum windows, to the initial laser power, using the same power meter in both cases to remove systematic effects. The quantum efficiency of the photocathode is factory measured to be 38.7% while the collection efficiency of the metal package PMT is estimated to be 70%. The PMT pulses are amplified by 46 dB and then sent into a NIM discriminator. Using a highly attenuated LED flasher as a single photon source, the discriminator threshold is optimized to give 99.4% efficiency for triggering on single photoelectron pulses while also efficiently rejecting the lower amplitude noise. By studying the trigger time distribution, the deadtime fraction due to possible multiple rapid PMT pulses is found to be negligible (0.001%). Thus, we estimate the total photon transport and counting efficiency to be $(25 \pm 3)\%$. Using this threshold, and the built-in cooler to cool the photocathode to 0 °C, we measure a typical dark count rate of 130 Hz.

To perform the coincidence counting we use two Quarknet boards [22,23] with 1.25 ns timing precision, referenced to a global positioning system (GPS) clock. The Quarknet boards determine the absolute time of the leading edge of time-over-threshold triggers from the PMT and from a monitoring photodiode that is located inside the laser box. The clocks on the laser board and on the PMT board are synchronized using an external trigger from a signal generator.

The absolute timing between the laser pulses and the PMT traces is established by removing the plunger with the mirror, and allowing the laser to shine on the PMT through



FIG. 2. PMT trigger times for the four run configurations, shown relative to the expected time distribution of photons as calibrated from the leaky-mirror data.

several attenuation stages consisting of two partially reflective ("leaky") mirrors, a pinhole, and multiple absorptive filters mounted directly on the aperture of the PMT module. The laser power is thus attenuated to a corresponding PMT trigger rate of less than 0.1 Hz for this timing calibration and to provide an *in situ* test of the data acquisition system.

The regenerated photons should arrive at the same time as the straight-through photons since milli-eV particles are also highly relativistic. For coincidence counting between the laser pulses and the PMT, a 10 ns wide window is chosen and includes 99% of the measured photon time distribution shown in Fig. 2. The coincident dark count rate can be estimated to be $R_{\text{noise}} = 20 \text{ Hz} \times 130 \text{ Hz} \times$ $10 \text{ ns} = 2.6 \times 10^{-5} \text{ Hz}$. This noise rate is negligible to the expected signal rate of $\sim 2 \times 10^{-3} \text{ Hz}$ estimated from the central values of the PVLAS parameters.

Data collection. —To cover the entire PVLAS range of m_{ϕ} , 20 h of data are collected in each of four configurations of plunger position and polarization. The plunger is placed either in the center of the magnet with $L_1 = 3.1$ m and $L_2 = 2.9$ m or near the far edge of the magnet with $L_1 = 5.0$ m and $L_2 = 1.0$ m. With each plunger position, data are taken separately with both vertically and horizontally polarized laser light in order to test pseudoscalar and scalar couplings, respectively. The central plunger position covers most of the PVLAS signal region but the corresponding oscillation probability goes to zero in Eq. (2) at $m_{\phi} \approx 1.4$ meV. Moving the plunger to the edge position restores sensitivity in this region.

The operating conditions are continuously monitored during each run. The reflected beam from the plunger mirror is slightly offset from the incident beam, and is directed into a calorimetric power meter by a pick-off mirror. The number of incident photons is determined with 3% accuracy from these measurements. The alignment of the laser is monitored using a fast solid state camera which takes images of the reflected laser spot on the pick-off mirror. Small ~mm scale transverse deviations are seen during the course of a typical 5 h run, due to small changes in the orientation of the plunger mirror as the plunger slowly cools through heat leaks in the warm bore insulation. The deviations are small enough that, were they due to actual changes in the laser alignment, the beam would still clear the aperture of the plunger. In addition, the focusing lens at the PMT makes the final light collection system insensitive to these potential sub-mrad angular deviations. Nevertheless, the alignment is double checked using the open-ended plunger after collecting data in each configuration, and no misalignment has ever exceeded our tolerances. The operation of the PMT is monitored using its dark rate. In addition, an LED flasher fires every 5 min to verify the integrity of the light collection system.

Results.—We count the PMT triggers within the 10 ns coincidence time window, defined *a priori* by the leakymirror data. The expected background is measured using the dark counts in time bins outside of the coincidence window. No excess counts above background are observed in any of the four configurations. The data are summarized in Table I. The PMT timing data, along with the leaky-mirror calibration data, are shown in an expanded time scale in Fig. 2. No excess counts above background are observed in any time bins near the laser pulse. For the central values of the PVLAS parameters one would expect ~ 150 excess counts.

We use the Rolke-Lopez method [24] to obtain limits on the regeneration probabilities, and use Eq. (1) to obtain the corresponding 3σ upper bounds on the coupling g as a function of m_{ϕ} . The relative systematic uncertainties of 12% on the photon transport and detection, and 3% on the laser power measurement are incorporated in the limits. The GammeV limits are shown in Figs. 3 and 4 along with the PVLAS 3σ signal region, the BFRT 3σ regeneration limits, and the 99.9% limit on pseudoscalar couplings from

TABLE I. Summary of data in each of the four configurations.

Configuration	Number of photons	Estimated background	Candidates	$g[\text{GeV}^{-1}]$
Horizontal, center	$6.3 imes 10^{23}$	1.6	1	3.4×10^{-7}
Horizontal, edge	6.4×10^{23}	1.7	0	$4.0 imes 10^{-7}$
Vertical, center	6.6×10^{23}	1.6	1	3.3×10^{-7}
Vertical, edge	7.1×10^{23}	1.5	2	$4.8 imes 10^{-7}$



FIG. 3 (color online). 3σ limit contours for scalar particles. The solid black line is the combined limit using data taken at both the central (red dot-dashed) and the edge (blue dashed) plunger positions. The PVLAS rotation signal (pink or dark gray) and the BFRT regeneration limit (tan or light gray) are also shown.

BMV. As expected, the regions of insensitivity for one plunger position are well covered by using the other plunger position. Data from both plunger positions are combined and analyzed jointly to produce the combined limit curve. The weakly interacting axionlike particle interpretation of the PVLAS data is excluded at more than 5σ by GammeV data for both scalar and pseudoscalar particles. The asymptotic 3σ upper bounds on *g* for small m_{ϕ} for each configuration are listed in Table I, and the combined analysis gives 3.1×10^{-7} GeV⁻¹ (3.5×10^{-7} GeV⁻¹) for the scalar (pseudoscalar) couplings. The GammeV exclusion region extends beyond the previous best limits and sets limits in regions where BFRT had reduced sensitivity.



FIG. 4 (color online). 3σ limit contours for pseudoscalar particles.

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