Flux and storage of energy in nonequilibrium stationary states

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Systems kept out of equilibrium in stationary states by an external source of energy store an energy $\Delta U =$ $U - U_0$. U_0 is the internal energy at equilibrium state, obtained after the shutdown of energy input. We determine ΔU for two model systems: ideal gas and a Lennard-Jones fluid. ΔU depends not only on the total energy flux, J_U , but also on the mode of energy transfer into the system. We use three different modes of energy transfer where the energy flux per unit volume is (i) constant, (ii) proportional to the local temperature, and (iii) proportional to the local density. We show that $\Delta U/J_U = \tau$ is minimized in the stationary states formed in these systems, irrespective of the mode of energy transfer. τ is the characteristic timescale of energy outflow from the system immediately after the shutdown of energy flux. We prove that τ is minimized in stable states of the Rayleigh-Benard cell.

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I. INTRODUCTION

Systems out of equilibrium are notoriously difficult to describe in a single coherent methodology based on variational principles. Principles such as Prigogine minimum entropy production [\[1\]](#page-4-0), Attard second entropy variation [\[2\]](#page-4-0), or Ziegler maximum entropy production [\[3\]](#page-4-0) suggested over the last 100 years have not reached the same status as the maximum entropy principle known from equilibrium thermodynamics [\[4–6\]](#page-4-0). A new paradigm, such as the driven lattice gas system, is believed to become an "Ising model" for nonequilibrium statistical physics $[7-12]$. Steady state thermodynamics is yet another description framework for nonequilibrium stationary states, which is still being developed [\[13–16\]](#page-4-0). Here we present a different approach to stationary states, based on two quantities: the energy stored in nonequilibrium states, ΔU , and the total energy flux, J_U , in these states.

The second law of thermodynamics states that the entropy of a system has its maximum value at the equilibrium state. Entropy, *S*, is a function of state; thus for an isolated system of *N* molecules of total internal energy, *U*, enclosed in a volume, *V*, the entropy has a fixed value $S = S(U, V, N)$. Internal constraints make it possible to divide the system into *n* isolated subsystems of entropies $S_i = S(U_i, V_i, N_i)$ for $i = 1, ..., n$, where $\sum_{i}^{n} N_i = N$, $\sum_{i}^{n} V_i = V$, $\sum_{i}^{n} U_i = U$. The maximum entropy principle states [[17,18\]](#page-4-0) that $S(U, V, N) \geq$ $\sum_i^n S_i$. We develop a similar methodology for nonequilibrium stationary states, by introducing internal constraints in these states. Next we make a conjecture on variational principles for these states. We do not claim the generality of this conjecture but simply prove it for three studied cases. In order to illustrate

II. IDEAL GAS BETWEEN TWO PLANAR WALLS

An ideal gas is confined between two planar walls of surface area \mathcal{A} located at $z = \pm L$. The temperature of the walls is constant $T_{-L} = T_L = T_0$. Energy is supplied to the volume of the fluid (e.g., by microwaves). In stationary states the total flux of energy into the system, J_U , matches the total flux at the walls. In the hydrodynamic limit, the time evolution of the system is given by the conservation laws for mass, momentum, and energy supplemented by the relations between thermodynamic forces and fluxes and thermodynamic equations of state. For an ideal gas $U = (3/2)Nk_BT$, $p = \rho k_BT$, where *U* is the internal energy, *p* is the pressure, $\rho = N/V$ is the number density of the ideal gas, *V* is the volume, *N* is the number of gas particles, and k_B is the Boltzmann constant. We observed in our previous simulations and calculations

our methodology we study two systems: ideal gas and a Lennard-Jones (LJ) fluid. Next we test our methodology on two competing states in the Rayleigh-Benard (RB) cell [\[19\]](#page-4-0). Summarizing our results obtained for these systems, we find that ΔU is minimized in nonequilibrium stationary states for a fixed J_U . Minimization is with respect to all constrained states of the system, similarly as in equilibrium thermodynamics. Due to the lack of a general theoretical framework for describing the nonequilibrium systems under consideration, we are not able to formulate a general argument for the validity of the conjectured principle. Nevertheless, it is plausible that, under well-defined conditions, the energy storage obeys some sort of variational principle. In the search for this principle we employed dimensional analysis and considered the relevant physical quantities. This analysis suggests the ratio $\Delta U/J_U$, which has the dimension of time and a nice physical interpretation, as a candidate for the quantity to be extremized.

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[\[20,21\]](#page-4-0) for gas-liquid evaporating systems that mechanical equilibrium is established very fast (in comparison to heat flow). Therefore we also expect a constant pressure across the system in this case. The stationary state satisfies $v = 0$ (gas velocity), $p = \text{const}$ [\[20,21\]](#page-4-0). These two equations come from the conservation of mass and momentum. (The same results, i.e., $\mathbf{v} = 0$ and $p = \text{const}$ in the stationary states are also obtained in MD simulations of the Lennard-Jones system described later.) The conservation of energy is given by $\nabla \cdot \mathbf{j}_{\varepsilon}(\mathbf{r}) = \sigma_{\varepsilon}$, where $\mathbf{j}_{\varepsilon}(\mathbf{r})$ is local heat flux and σ_{ε} is the external heat source. The integral of σ_{ε} over the volume is $\int_V \sigma_{\epsilon} = J_U$. The local flux of energy per unit area is proportional to the local temperature gradient $\mathbf{j}_{\varepsilon} = -\kappa \nabla T$, where κ is the thermal conductivity and thus

$$
-\kappa \nabla^2 T(\mathbf{r}) = \sigma_{\varepsilon}.
$$
 (1)

We solve Eq. (i) for three forms of the source term describing different manners of energy supply: (i) $\sigma_{\varepsilon} = \lambda_1$, (ii) $\sigma_{\varepsilon} =$ $\lambda_2 T(z)$, and (iii) $\sigma_{\varepsilon} = \lambda_3 \rho(z)$. The equations are transformed into dimensionless form by rescaling the variables $\tilde{z} = z/L$, $\tilde{T} = T/T_0$ and $\tilde{\rho} = \rho/\rho_0$ together with (i) $\tilde{\lambda}_1 = \lambda_1 L^2/(\kappa T_0)$, (ii) $\tilde{\lambda}_2 = \lambda_2 L^2/\kappa$, and (iii) $\tilde{\lambda}_3 = \lambda_3 L^2 \rho_0/(\kappa T_0)$, with $\lambda_i >$ $0, i = 1, 2, 3$. Here ρ_0 is the number density of the ideal gas at temperature T_0 and pressure p_0 . The temperature is a function of coordinate *z* only. In all these cases we introduce energy directly into the volume. This mode of energy transfer is well known by, e.g., electromagnetic waves (microwaves, visible light, etc.). If we couple a microwave device with a thermovision camera we can in principle add energy to hotter places (where density is low) or to less hot spots, where density is high. Thus the three modes of energy transfer are physically feasible.

For case (i) the equation for the stationary temperature profile is $d^2\tilde{T}(\tilde{z})/d\tilde{z}^2 = -\tilde{\lambda}_1$ with symmetric boundary conditions $\tilde{T}(-1) = \tilde{T}(1) = 1$. This equation has the solution $\tilde{T}(\tilde{z}) = -\frac{\tilde{\lambda}_1}{2}(\tilde{z}^2 - 1) + 1$. The energy of this stationary state is calculated using the assumption of local equilibrium. The energy density field $U/(V \rho_0 T_0) = \varepsilon [T(\tilde{z}), \tilde{\rho}(\tilde{z})] =$ $(3/2)k_B\tilde{\rho}(\tilde{z})\tilde{T}(\tilde{z})$ for an ideal gas. Upon rescaling by the equilibrium value $U_0/(V \rho_0 T_0) = \varepsilon_0 = (3/2)k_B$, the dimensionless local energy density $\tilde{\varepsilon} = \varepsilon/\varepsilon_0$ obeys $\tilde{\varepsilon}(\tilde{T}(\tilde{z}), \tilde{\rho}(\tilde{z})) / \tilde{T}(\tilde{z}) =$ $\tilde{\rho}(\tilde{z})$. Our system does not exchange molecules with the environment; therefore the number of particles is constant. This condition is given by the equation $\int_{-1}^{1} \tilde{\rho}(\tilde{z}) d\tilde{z} = 2$ and thus implies $\int_{-1}^{1} d\tilde{z} \tilde{\varepsilon}(\tilde{T}(\tilde{z}), \tilde{\rho}(\tilde{z})) / \tilde{T}(\tilde{z}) = 2$. The reduced pressure obeys $\tilde{p} = \tilde{T}(\tilde{z})\tilde{\rho}(\tilde{z}) = \tilde{\varepsilon}(\tilde{T}(\tilde{z}), \tilde{\rho}(\tilde{z}))$, thus since \tilde{p} is constant, so is $\tilde{\varepsilon}$. We obtain $\tilde{\varepsilon}$ as

$$
\tilde{\varepsilon} = \frac{2}{\int_{-1}^{1} d\tilde{z}/\tilde{T}(\tilde{z})}.
$$
 (2)

Integrating the temperature profile we obtain

$$
\tilde{\varepsilon}(\tilde{\lambda}_1) = \frac{\sqrt{\tilde{\lambda}_1(\tilde{\lambda}_1 + 2)}}{2 \tanh^{-1} \left(\sqrt{\frac{\tilde{\lambda}_1}{\tilde{\lambda}_1 + 2}} \right)}.
$$
\n(3)

For cases (ii) and (iii) of the external flux see the Supplemental Material [\[22\]](#page-4-0).

We introduce a rigid, impenetrable, adiabatic wall in the system at $z = z_1$. The wall divides the system into two subsystems (1) and (2). The internal energy at equilibrium is $U_0 = 2L\mathcal{A}\varepsilon_0$, the same as in the unconstrained system. We calculate the storage of energy over its equilibrium value, $\Delta U_1 + \Delta U_2 = U_1 + U_2 - U_0$ in the same way as presented above. The division into two subsystems satisfies the following conditions in our methodology: (a) The subsystems are in mutual equilibrium after the shutdown of energy input, so that no additional fluxes appear after removal of the constraints at equilibrium. (b) The subsystems reach a stationary state characterized by U_1 , U_2 and fluxes J_{U_1} , J_{U_2} . (c) The mode of energy transfer to each subsystem is the same. Since $J_U =$ $\mathcal{A} \int_{-L}^{L} \sigma_{\varepsilon} dz$, we obtain for case (1)

 ΔU

and

$$
\frac{\Delta U_1 + \Delta U_2 - U_0}{J_{U_1} + J_{U_2}} = \frac{\varepsilon_0}{\lambda_1} [f_1(\tilde{z}_1) + f_2(\tilde{z}_1) - 1] \tag{5}
$$

 $\frac{\Delta U}{J_U} = \frac{\varepsilon_0}{\lambda_1} (\tilde{\varepsilon} - 1)$ (4)

with

$$
f_1(\tilde{z}_1) = \frac{(1 - \tilde{z}_1)\sqrt{\Lambda_{-}^{(1)}(\Lambda_{-}^{(1)} + 2)}}{4\tanh^{-1}(\sqrt{\frac{\Lambda_{-}^{(1)}}{\Lambda_{-}^{(1)} + 2}})} \text{ and}
$$

$$
f_2(\tilde{z}_1) = \frac{(1 + \tilde{z}_1)\sqrt{\Lambda_{+}^{(1)}(\Lambda_{+}^{(1)} + 2)}}{4\tanh^{-1}(\sqrt{\frac{\Lambda_{+}^{(1)}}{\Lambda_{+}^{(1)} + 2}})}, \tag{6}
$$

where $\tilde{z}_1 = L_1/L$, $\Lambda_{-1}^{(1)} = \tilde{\lambda}_1 (1 - \tilde{z}_1)^2$, and $\Lambda_{+}^{(1)} = \tilde{\lambda}_1 (1 + \tilde{z}_1)^2$ $(z_1)^2$. By inspection, in the range $0 \le z_1 < 1$ the functions $f_i(\tilde{z}_1)$, $i = 1, 2$ are positive and lie above their tangent lines at $\tilde{z}_1 = 0$, i.e., $f_i(\tilde{z}_1) > f'_i(\tilde{z}_1)|_{\tilde{z}_1 = 0} \tilde{z}_1 + \frac{\varepsilon}{2}$, $i = 1, 2$ for fixed $\lambda_1 > 0$. Since $f_1(\tilde{z}_1) = f_2(-\tilde{z}_1)$ one has $f'_1(\tilde{z}_1)|_{\tilde{z}_1=0} =$ $-f'_{2}(\tilde{z}_1)|_{\tilde{z}_1=0}$ and hence $f_1(\tilde{z}_1) + f_2(\tilde{z}_1) > 2\frac{\varepsilon}{2}$, which proves that

$$
\frac{\Delta U_1 + \Delta U_2}{J_{U_1} + J_{U_2}} \geqslant \frac{\Delta U}{J_U}.\tag{7}
$$

The equality holds only for equal partition of the system into two subsystems $(z_1 = 0)$. This equation states that the energy stored in two subsystems is larger than in the unconstrained system [\[23\]](#page-4-0) for a fixed flux $J_U = J_{U_1} + J_{U_2}$. Cases (ii) and (iii) also satisfy Eq. (7) (discussed in Ref. [\[22\]](#page-4-0)). Additionally in the Supplemental Material we present calculations of case (i) with $\kappa = \text{const}\sqrt{T}$ (expected for the dilute gas), which further confirm Eq. (7) for this system. It is well understood that a true ideal gas would have noninteracting particles with a zero collision cross section. We use the equation of state of the ideal gas as an approximation for the interacting gas characterized by finite, temperature-dependent, κ .

III. LENNARD-JONES LIQUID IN A RECTANGULAR BOX

In order to perform analytical calculations for the ideal gas model we had to assume local equilibrium. This assumption is inherent in irreversible thermodynamics equations. In order to test our analytical results for the ideal gas model we performed molecular dynamics (MD) simulations of a

FIG. 1. LJ simulation box with the shaded region in the middle dividing the system into (1) and (2) subsystems of size L_1 and *L*2, respectively. The finite size of this wall precludes interactions between the molecules in different subsystems. Heating is indicated by the red arrows. Cooling is performed at the boundaries at a distance b_0 from two walls and shown schematically by the green arrows. The fluxes J_{U_1} and J_{U_2} in and out are equal in the stationary state. The total size of the system *L* is fixed.

LJ system. MD simulations provided qualitatively the same results as analytical calculations presented in the previous section. In the MD simulations Newton equations of motion are solved; thus no assumptions concerning local equilibrium or constancy of heat conductivity are made. Nevertheless, during MD simulations our system stays quite close to local equilibrium. In general we can expect a violation of the local equilibrium assumption only when the flux of the energy flowing across the system is faster than the process of local distribution of energy between all degrees of freedom. Such conditions are expected in, e.g., shock waves.

In all the simulations we set the flux constant for the system and subsystems: $J_U = J_{U_1} + J_{U_2}$. We also set the same *JU* for all the modes of energy transfer into the system. Thus we make a comparison between different cases using only one parameter, the energy ΔU stored in a system. MD simulations $[24,25]$ (Fig. 1) were performed in the system of fixed number $N = 266240$ of LJ atoms. The simulation box was divided into two subsystems (1) and (2) of sizes L_1 and L_2 , respectively. The total size $L_1 + L_2$ of the box was constant. Energy was added only to the regions (1) and (2) (Fig. 1) once per 10 time steps in three manners (i), (ii), and (iii) as described for the ideal gas model. When the flux of energy was proportional to the density each LJ atom received the same amount of kinetic energy. For flux proportional to temperature the portion of energy added to each atom was proportional to its kinetic energy. For energy flux proportional to density the same amount of energy was added to the same volume, i.e., all the atoms in a given volume received a given amount of energy equally shared between them. All temperature and density profiles in the stationary states are shown in Ref. [\[22\]](#page-4-0) for the system with and without internal walls. In Fig. $2(a)$ we show the dependence of the energy ΔU stored in the system, as a function of the size L_1 of one subsystem for fixed fluxes $J_U = J_{U_1} + J_{U_2}$. This figure contains results for all three scenarios of the energy transfer. All physical quantities are given in LJ units. When $L_1 = L_2$, ΔU reaches a minimal value

FIG. 2. (a) Energy stored $\Delta u_1 + \Delta u_2$ per LJ atom in subsystems (1) and (2) as a function of the size of one subsystem L_1 for a fixed flux $J_U = J_{U_1} + J_{U_2}$. The flux of energy per unit volume is constant (red triangles), proportional to temperature (black circles), or proportional to density(green squares). The minimal value of $\Delta U_1 + \Delta U_2 = \Delta U$ is equal to the value for the unconstrained system for $L_1 = L_2$ [see Eq. [\(7\)](#page-1-0)]. (b) The energy (per particle) in subsystem (1) as a function of *L*₁. This energy scales as $U_1/V \sim L_1^2$. It is not an extensive quantity, since the energy at equilibrium $U_0/V \sim$ const. (c) The characteristic time τ of energy out-flow from the subsystem (1) after the shutdown of energy flux into the subsystem. Energy decreases as $\exp(-t/\tau)$ (see Ref. [\[22\]](#page-4-0)) with an initial decay time $\tau = 2\Delta U_1 / J_{U_1}$. Factor 2 appears because the outflow of energy is only by one wall.

(equal to the value obtained for the unconstrained system) as expected from Eq. (7) . In Fig. $2(b)$ we show the scaling of the energy per particle stored in the subsystem (1) as a function of the size of the subsystem L_1 . Finally we observe that after the shutdown of energy flux into the system, the energy decreases as $\exp(-t/\tau)$ (see Ref. [\[22\]](#page-4-0)) at short times, with a decay time $\tau = \Delta U / J_U$ [Fig. 2(c)]. In summary, the MD simulations of LJ fluid confirm Eq. [\(7\)](#page-1-0). In most of our simulations the system was in a single-phase state, but we also confirmed Eq. [\(7\)](#page-1-0) for the two-phase nonequilibrium system. We observed a spontaneous phase separation and the two phases present in the stationary state with a liquid at the cold boundary and a heated gas inside the simulation box. In these simulations we could not obtain stationary states with convection as in the RB cell irrespective of the value of the energy flux input into the system.

IV. TWO-DIMENSIONAL RAYLEIGH-BENARD SYSTEM OF HARD DISKS (HDS)

We further tested Eq. (7) in the RB cell (Fig. [3\)](#page-3-0) for two competing stationary states. One is called the conductive state Fig. $3(b)$ and second one the convective state as shown in Fig. $3(c)$. For small temperature gradients the conductive

FIG. 3. (a) $\Delta U / J_U$ as a function of the lower plate dimensionless temperature, T^* in the RB two-dimensional system. The vertical solid line marks RB instability at about $T^* = T^*_{RB} = 15.5$. The open squares are the MD data for the constrained system in the conductive state (for all T^*) with wall in the middle (see panel b). The open circles are the MD data for the unconstrained system in the conductive state (b) for $T < T_{RB}$ and convective state (c) for $T > T_{RB}$. (b) Snapshots of the stationary velocity field (at $T^* = 17$) in the conductive state (constrained state). (c) Snapshots of the stationary velocity field in the convective state, stable for $T^* > T^*_{RB}$.

state is stable, while for large temperature gradients it is the convective state that is stable. We use the internal walls to stabilize the conductive state above the transition to the convective state. This methodology allows for comparison of $\Delta U / J_U$ for both states at the same gravitational field and temperatures at the walls.

Disks of diameter σ and mass *m*, were placed in a rectangular simulation box of size $L_x d \times L_y d$ (where $d = \sigma / \sqrt{0.4}$). Two sizes were studied: $50d \times 100d$ (the small system) and $100d \times 100d$ (the large system). The dimensionless density, $\rho^* = N\sigma^2/(L_xL_yd^2) = 0.4$. The upper plate had a constant temperature T_0 , and the lower plate was heated to temperatures $T = T^* \times T_0$, with $T^* \ge 1$. All disks were subjected to a gravitational force $F = m(0, -g)$. The HD fluid was simulated with event-driven dynamics, where disks followed parabolic trajectories between collisions. The disks collided in an elastic manner with the side walls and between themselves. The collisions with the upper and lower plates allowed for a transfer the thermal energy to the system. Upon collision with the upper or lower plate the disc velocity was drawn from the Maxwell distribution corresponding to the temperature of the plate, and the direction was chosen randomly from 0 to 180° angles with respect to the plate [\[26\]](#page-4-0). The total energy of the system is the sum of the kinetic and potential energy of the disks (details are given in Ref. [\[22\]](#page-4-0)).

For $T^* = 1$ the system reaches the thermal equilibrium state with an internal energy U_0 . For $T^* > 1$ energy is being pumped into the system at the bottom plate until the system reaches a stationary state, characterized by the energy stored in the system above its equilibrium value, $\Delta U = U - U_0$. In the stationary state, this system is expected to conduct heat in a conductive manner up to a temperature which corresponds to the RB instability [\[19\]](#page-4-0), T_{RB}^* . This temperature (T_{RB}^*) marks the onset of the convective mode of heat transfer, characterized by the formation of convective rolls. In our system measuring $100d \times 100d$, the transition temperature T_{RB}^{*} is about 15.5. For $T^* > 15.5$ we observe a convective state with a single roll filling almost the whole area of the simulation box. However, for the system measuring $50d \times 100d$, the system is too small to develop rolls at all studied temperatures (up to $T^* = 25$). This observation makes it possible to constrain the system by the internal wall and stabilize the conductive state against convective instability. We introduced a constraint into the system by inserting a vertical adiabatic wall in the middle of the $100d \times 100d$ simulation box. The wall divides the system into two 50 $d \times 100d$ independent subsystems (1) and (2). The $50d \times 100d$ subsystem is too small to develop even a single convective roll for $T^* > 15.5$. Thus a conductive stationary state is observed for each T^* that is considered. Next, by removing the adiabatic wall for $T^* > 15.5$ we merge two subsystems in the conductive state into a single $100d \times 100d$ system in the convective state. For $T < T_{RB}^*$, the merging subsystems (1) and (2) do not change the conductive stationary state and in particular $\Delta U_1 + \Delta U_2 = \Delta U$ for $J_{U_1} + J_{U_2} = J_U$ at all temperatures $T^* < T^*_{RB}$.

Figure 3 shows the results of the simulations. The conductive state [shown in Fig. $3(b)$] is stable for $T^* < T^*_{RB}$. In this state the whole system and the subsystems satisfy the equations $\Delta U = \Delta U_1 + \Delta U_2$ and $J_U = J_{U_1} + J_{U_2}$. However, for T^* > T_{RB}^* the subsystems (1) and (2) are still in a conductive state, while the whole system without the constraint reaches a convective state [Fig. 3(c)]. We find $\Delta U/J_U < (\Delta U_1 +$ ΔU_2)/(*J*_{U₁} + *J*_{U₂}) for $T^* > T_{RB}^*$ as predicted by Eq. [\(7\)](#page-1-0). Concluding, in the RB cell $\Delta U / J_U$ is minimized in stationary states.

V. CONCLUSIONS AND FURTHER DISCUSSION

We have presented a methodology for the analysis of nonequilibrium states, based on internal constraints known from equilibrium thermodynamics. We have pointed out the importance of the mode of energy transfer into the system and introduced two quantities characterizing the nonequilibrium stationary state: ΔU the excess energy stored in the nonequilibrium state over the equilibrium value and $\tau = \Delta U/J_U$, the characteristic time of energy outflow from the system after shutdown of energy flux into the system. $\Delta U > 0$ in all examples discussed in this paper, because we used a reference equilibrium state of lower energy, rather than the energy of the stationary state.

We observed that in all cases studied the quantity $\Delta U/J_U$ is minimized in stationary states. However, the following counterexample suggests that this may not be a general principle [27]. Consider a huge box with adiabatic walls attached to our system by a thin heat-conducting wire ending at a point inside our system. The point is chosen in such a way that the temperature at this point is lower in the constrained system than in the unconstrained one. Such a point always exists in the systems studied. This box stores an extra energy during the energy flow from our system to the box. The flow stops when the temperature of the box reaches the stationary state temperature in the point of choice in our system. This sort of box does not influence the stationary state of the system but, nonetheless, changes the total amount of energy stored in the total system. Since in this way one an arbitrary large amount of energy can be stored, we cannot claim that for a fixed flux the energy stored in the system is minimized. We can make many different variants of this counterexample, in particular, allowing a small flux through this box and in this way affecting the final nonequilibrium state of the system. Since such situations are not eliminated by the current formulation of prerequisites, their further analysis is needed. Those prerequisites which are related to nonequilibrium states require special attention. Indeed, the division of a system into

subsystems, so obvious for the equilibrium state, is not at all obvious for nonequilibrium states. This is due to the crucial role of surfaces bounding the system establishing the final nonequilibrium, stationary state.

Our methodology and observations require further tests. Such tests can be performed for chemical systems with many competing stationary states or in hydrodynamic systems. One example, which we are going to study, is the reaction between nitrogen dioxide and nitrogen tetroxide [28]. A system consisting of these two chemical compounds is illuminated by light. Light is absorbed by nitrogen dioxide but not by nitrogen tetroxide. The absorption of light results in heating of the sample and in an increase in the backward reaction from tetroxide to dioxide. In this system many stationary states appear. We hope that such a test will further support our current observations.

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