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Free energy of singular sticky-sphere clusters

Yoav Kallus*

Santa Fe Institute, 1399 Hyde Park Road, Santa Fe, New Mexico 87501, USA

Miranda Holmes-Cerfon*

Courant Institute of Mathematical Sciences, New York University, New York, New York 10012, USA (Received 16 August 2016; published 22 February 2017)

Networks of particles connected by springs model many condensed-matter systems, from colloids interacting with a short-range potential and complex fluids near jamming, to self-assembled lattices and various metamaterials. Under small thermal fluctuations the vibrational entropy of a ground state is given by the harmonic approximation if it has no zero-frequency vibrational modes, yet such singular modes are at the epicenter of many interesting behaviors in the systems above. We consider a system of N spherical particles, and directly account for the singularities that arise in the sticky limit where the pairwise interaction is strong and short ranged. Although the contribution to the partition function from singular clusters diverges in the limit, its asymptotic value can be calculated and depends on only two parameters, characterizing the depth and range of the potential. The result holds for systems that are second-order rigid, a geometric characterization that describes all known ground-state (rigid) sticky clusters. To illustrate the applications of our theory we address the question of emergence: how does crystalline order arise in large systems when it is strongly disfavored in small ones? We calculate the partition functions of all known rigid clusters up to $N \leq 21$ and show the cluster landscape is dominated by hyperstatic clusters (those with more than 3N-6 contacts); singular and isostatic clusters are far less frequent, despite their extra vibrational and configurational entropies. Since the most hyperstatic clusters are close to fragments of a close-packed lattice, this underlies the emergence of order in sticky-sphere systems, even those as small as N=10.

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I. INTRODUCTION

Particles that live on the mesoscale commonly interact over ranges much shorter than their diameters. Naturally occurring systems include C_{60} molecules [1,2] and colloids interacting via depletion [3,4]. A widely studied synthetic system is colloidal particles coated with strands of DNA or other functional tethers, which create highly specific interactions with range roughly the length of a tether [5–9]. Rather than model the details of the interparticle interaction, which are often not important for macroscale observations, it is convenient to model the interaction in the *sticky limit*, which considers a central-force potential with a single well that is narrow and deep [10,11].

As the range of the interaction goes to zero, the space of energy-preserving motions available to a finite system of particles is identical to that of a bar framework, with nodes located at the particle centers and bars connecting pairs of particles that are touching, as in Fig. 1. The energy of the system is proportional to the number of contacts, or bars in the framework. When thermal fluctuations allow the structure to slightly deviate from this space, the resulting entropy can be derived from the harmonic response of the network, provided all vibrational modes have nonzero frequency. The result contains a dimensionless geometric factor and a single physical parameter, characterizing the temperature and stiffness of the interaction potential [11].

This limit is appealing because it clearly separates the effects of geometry of the particles from those of the interaction potential [12,13], allowing one to more easily vary the

latter to study self-assembly processes, for example. It was originally considered by Baxter to study phase transitions in fluids of particles with short-range interactions [10], though singularities related to the ones we will solve for have prevented it from being more widely used [14]. The sticky limit can also be used as a controlled approximation for the entropy of a system of hard particles without attraction near close packing [15].

A natural starting point for investigating the consequences of the sticky limit is to consider a collection of finitely many spheres. In the sticky limit the lowest-energy states have the maximum number of contacts, and they are typically mechanically rigid. A recent body of work has focused on computing the set of rigid clusters of N spheres and understanding their thermodynamic properties [16–20]. This program aims partly at identifying a set of geometrical possibilities for processes like self-assembly [21,22] and self-replication [23], and partly at understanding the question of emergence: how does a finite system of particles transition from a small size, where the preferred order might be incompatible with crystalline order, to a large one where it assembles into a highly structured close-packed lattice [24–29]?

A major roadblock for such a program is the treatment of *singular* clusters, those with vibrational modes with zero frequency that do not extend to finite motions (singular modes), so-called because they correspond to singularities of a system of algebraic equations. Current methods for treating the sticky limit do not work directly for these systems because the entropy associated with the singular modes diverges. The smallest singular cluster occurs at a mere N = 9 [17], preventing the sticky limit from being applied to systems that size or larger with current methods. Yet singular clusters are not a rarity that can be glibly ignored: they account for about 2.5% of the known rigid clusters of N spheres [20]. A related problem is

^{*}Corresponding authors: yoav@santafe.edu; holmes@cims.nyu.edu

hyperstatic rigid clusters, those with more than the 3N-6 contacts required for a cluster to be rigid generically. The free energy of hyperstatic clusters also diverges in the traditional sticky limit, a fact that has hindered the limit from being more widely applied to study bulk systems [14]. The smallest hyperstatic cluster has N=10 spheres [17]. Observations of colloids with a short-range depletion interaction found the N=9 singular cluster occurred with disproportionate frequency, and for $N \geqslant 10$ the majority of observed clusters were either singular or hyperstatic [16]. This suggests there could be a competition between the higher entropy of a singular mode and the lower energy of an additional contact. We wish to evaluate this competition and determine if singular clusters could be thermodynamically stable as the number of spheres increases into the bulk regime.

In this paper we extend the sticky limit to systems with singular modes and nonisostatic numbers of contacts. We show that, provided the system is *second-order rigid*, its partition function depends on only two parameters, characterizing the depth and range of the pair potential. This is only one more parameter than is required in the nonsingular case. The property of being second-order rigid is explained in the text and characterizes all known rigid clusters. Our approach is to calculate the partition function from an expansion of the potential energy function up to fourth-order in the particle displacements. The result diverges as the range of the potential approaches zero, but by using the leading asymptotic term, we obtain a finite result if the range is taken as a small, nonzero parameter.

This computation makes specific predictions for the probabilities of observing singular clusters in equilibrium, which we confirm by comparing to simulations of particles interacting with a short but finite-range potential. We then calculate the free energies of all known rigid clusters of spheres up to $N \leq 14$ and of hyperstatic clusters for $15 \leq N \leq 21$, and show that the added entropy of singular clusters never beats the lower energy of the maximally hyperstatic clusters at realistic values of stickiness, except when the singular clusters are maximally hyperstatic. Since the most hyperstatic clusters are fragments of a close-packed lattice (or close to such), this observation explicitly and quantitatively demonstrates the emergence of order even at small system size. Finally, our results suggest a universal scaling, not yet explained, of the configurational and vibrational entropies of sticky sphere clusters as a function of the number of contacts.

The problem of calculating the effects of singular modes extends far beyond the system specifically under study in this paper. Any system whose motions are limited by soft constraints, in the limit that the constraints become very stiff, could exhibit infinitesimally free motions that do not extend to finite degrees of freedom. In particular, singular modes exist and seem to play an important role in the mechanical and thermodynamic stability of many metamaterials, a class of systems of growing interest in material science, composed of mesoscopic or macroscopic building blocks and designed to have properties that are hard or impossible to engineer in traditional materials. For example, metamaterials based on ball-and-spring networks may be designed to have singular modes on purpose so they fail under stress at desired locations [30]. Origami mechanisms designed to have a single

extended degree of freedom may, as has been observed, end up with extensively many infinitesimal degrees of freedom, with important implications to the mechanics of actuating the mechanism [31]. As low-frequency modes seem to play an important role in the rheology of athermal fluids made of isotropic particles [32,33], zero-frequency modes, which occur in extensive number at the jamming point of anisotropic particles [34], could also have a significant effect.

While the mechanical effects above are athermal, zerotemperature effects, in this paper we focus on the thermal effects of singular modes. Such effects are already relevant for some metamaterials and will become relevant for others as the trend toward miniaturization is followed and such systems are fabricated at smaller and smaller scales. An abundance of low-frequency vibrational modes increases the entropy of a system and can often be a crucial factor stabilizing one structure over another when they are energetically equivalent: for example, such an effect favors a face-centered-cubic closepacking structure over a hexagonally close-packed structure in many systems [15], stabilizes marginally coordinated lattices in patchy colloids [35], and favors an ordered structure over a zigzag structure in a rhombus lattice [36]. If given two energetically equivalent structures, one has singular modes, this effect becomes even stronger: the entropic factor favoring the singular structure grows larger and larger as the constraints become stiffer, as we show in this paper. It is conceivable that in such cases the singular structure could be thermodynamically favored in an appropriate limit even when it is energetically disfavored. We derive the conditions for such a scenario to occur, but we find that they are not present in the case of a system of N sticky spheres for any $N \leq 21$.

II. THE STICKY LIMIT

A. Rigid clusters

We start by explaining how to calculate the free energy in the sticky limit for a cluster of N identical spheres with diameter d, forming a cluster $\mathbf{r} = \mathbf{r}_1 \oplus \mathbf{r}_2 \oplus \ldots \oplus \mathbf{r}_N \in \mathbb{R}^{3N}$, where $\mathbf{r}_i \in \mathbb{R}^3$ is a vector specifying the position of the center of the ith sphere, and \mathbf{r} is the corresponding point in a 3N-dimensional configuration space. We suppose the cluster has B bonds between pairs of interacting spheres $E = \{(i_1, j_1), \ldots, (i_B, j_B)\}$. When the range of interaction is infinitesimally short, each bonded pair will be exactly touching, so the cluster lies in the solution set to the system of equations,

$$|\mathbf{r}_i - \mathbf{r}_j|^2 = d^2, \quad (i, j) \in E. \tag{1}$$

We focus for illustration on clusters that are rigid, though the calculations extend naturally to floppy ones. A cluster is defined to be *rigid* if it lies on a connected component of the solution set to Eq. (1) that contains only rotations and translations [37,38]. Physically, being rigid means one cannot continuously deform the cluster internally by any finite amount while maintaining all contacts. This is a nonlinear notion that is *not* equivalent to counting the number of infinitesimal degrees of freedom, as in the Maxwell-Calladine theorem [39].

An example of a cluster that is rigid but has an infinitesimal degree of freedom is shown in Fig. 1. This is a cluster of N = 9 spheres, consisting of two bipyramids that share a sphere at





FIG. 1. Left: the smallest singular cluster of sticky spheres, consisting of nine spheres. Right: a framework representation of the cluster with the singular mode indicated by red arrows. The singular mode is a vibrational mode with zero frequency, which cannot be extended to any finite deformation without incurring an energy cost, or equivalently without breaking a contact.

a common vertex and are connected by three parallel bonds. When one bipyramid is rotated along the axis going through the shared vertex and parallel to the bonds, the three bond lengths do not change to linear order in the displacement, so this rotation is an infinitesimal degree of freedom. The bond lengths do change at quadratic order in the displacement. In fact, any finite motion that is not a rigid body motion changes at least one bond length, and the cluster is rigid.

We suppose the potential energy of the system is $U(\mathbf{r}) = \sum_{1 \leq i < j \leq N} V(|\mathbf{r}_i - \mathbf{r}_j|)$, a sum of pair potentials V(r) that depend on the distance r between each pair. The pair potential is assumed to have a minimum at d and approach zero rapidly as r-d exceeds a certain characteristic interaction length ϵ . For $r-d \ll -\epsilon$, we assume the potential is positive and much larger in magnitude than its minimum value. As a particular example, we consider the Morse potential $V(r) = \mathcal{E}e^{-\rho(r-1)}(e^{-\rho(r-1)}-2)$, where the position of the minimum of the pair potential determines the diameter, d=1, and the parameter ρ determines the range of the interaction, $\epsilon=1/\rho$.

The sticky limit occurs when the well of the pair potential is both *narrow* and *deep*. This limit can be constructed from any interaction potential with reasonable decay, by simultaneously rescaling the lengthscale r-d and the depth of V(r) (see Fig. 2). The sticky limit has traditionally been considered under scalings such that the contribution of the well to the partition function for a single pair of spheres approaches a constant [10,14]. Using Laplace's method (see Appendix A 1) this constant is shown to be proportional to

$$\kappa = \lim_{\epsilon \to 0} \frac{\sqrt{2\pi}}{d} \frac{e^{\beta \mathcal{E}}}{(\beta a)^{1/2}},\tag{2}$$

where $\mathcal{E} = -V(d)$, $a = V''(d) \sim \mathcal{E}/\epsilon^2$, and $\beta = (k_B T)^{-1}$ is the inverse of temperature T times the Boltzmann constant. Since it represents the equilibrium tendency of spheres to stick to each other, it is called the *sticky parameter* [11].

The sticky parameter characterizes the thermodynamic tradeoff between the free energy of a particle in a bulk fluid of fixed packing fraction ϕ , $\exp(-\beta F_{\rm fluid}/N) \sim d^3/\phi$, and a particle in an isostatic network of bonded particles, $\exp(-\beta F_{\rm iso}/N) \sim (\beta a d^2)^{-3/2} e^{3\beta \mathcal{E}}$. Therefore, the limit $\epsilon \to 0$ as $0 < \kappa < \infty$ is relevant for studying the instability of a fluid toward forming an isostatic bonded network [40]. However, a close-packed network has six bonds per particle in the bulk

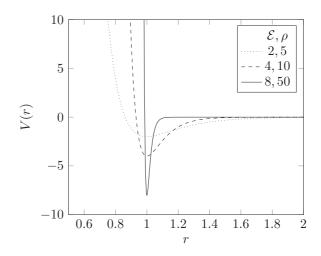


FIG. 2. The sticky sphere limit of a Morse pair potential, $V(r) = \mathcal{E}e^{-\rho(r-1)}[e^{-\rho(r-1)}-2]$, where the depth $\mathcal{E} \to \infty$ and the range $\epsilon = 1/\rho \to 0$.

limit and so the relevant limit for studying its formation is $0 < \lim_{\epsilon \to 0} (\beta a d^2)^{-1/2} e^{2\beta \mathcal{E}} < \infty$, which requires $\kappa \to 0$. As the number of spheres N in a finite system rises, the relevant clusters might vary from isostatic to close-packed, so one might wish to consider different limits. Our calculation allows the limits $\epsilon \to 0$ and $\mathcal{E} \to \infty$ to be taken in any way desired, as we show that it does not change the form of the leading term in the partition function.

To evaluate the free energy of the different rigid clusters that N spheres can form, we need to incorporate the contribution from vibrations about the minimum-energy configurations as well as from rigid body motions that do not change the energy. Such a calculation is familiar from the classical thermodynamic treatment of an ideal gas of polyatomic molecules [41]. For completeness, we go through the entire calculation in broad strokes in the main text and fully in Appendix A. However, our calculation has several differences from classical treatments: (i) we go beyond the harmonic approximation, which fails precisely for the clusters we are most interested in; (ii) we consider the partition function in the sticky limit of a fairly general pair potential, which leads to an expression depending on only two parameters; (iii) we show the connection between the vibrational modes of a cluster and its mechanical properties as a framework; and (iv) we take care to keep track of error terms and explicitly bound the scaling of subleading terms. Since we are interested in mesoscale particles, we do not take into consideration quantum effects. Moreover, our calculation is different from simply including all anharmonic corrections to the free energy, as has been done numerically in molecular dynamics studies of Lennard-Jones clusters [42]. Calculating these is usually a Sisyphean task, requiring heavy lifting and providing only minor corrections. However, when the vibrational spectrum contains zero modes, these corrections become crucial. Thankfully, as we show, the anharmonic corrections need only be calculated for the zero modes, though care must be taken to account for the coupling between the zero and nonzero modes, which has previously been neglected [16].

The contribution to the canonical partition function from each rigid cluster is given by

$$Z_{\mathbf{r}} = \int_{\mathcal{N}_{-}} \exp[-\beta U(\mathbf{r}')] d^{3N} \mathbf{r}', \tag{3}$$

where $\mathcal{N}_{\mathbf{r}}$ is an integration domain constructed by taking all configurations corresponding to the cluster of interest, including translations, rotations, permutations, and reflections, and fattening it by a length that is small compared to the particle diameter d and large compared to the interaction range ϵ (see Appendix A 3.) We take the total partition function to be the sum of the contributions from all the rigid clusters, ignoring floppy clusters in this paper. The free energy is $F_{\mathbf{r}} = -\beta^{-1} \log Z_{\mathbf{r}}$.

We may remove rigid-body degrees of freedom by a change of variables. If we let $W \subset \mathbb{R}^{3N}$ be the linear subspace of infinitesimal rigid-body motions, then we may restrict the integral to its orthogonal complement W^{\perp} , and include a factor I equal to the square root of the determinant of the moment of inertia tensor (Appendix A 3). Summing over permutations contributes a factor of $N!/\sigma$, where σ is the size of the Euclidean symmetry group of the cluster [43]. If the spheres lie in a container of volume Ω , then the free volume for each cluster is nearly Ω as long as $V \gg d^3$.

To evaluate the remaining parts of the integral in Eq. (3) we expand each term $V(\mathbf{r}_{ij} + \delta \mathbf{r}_{ij})$ in powers of $\delta \mathbf{r}_{ij}$, where we write $(\cdot)_{ij}$ for $(\cdot)_i - (\cdot)_j$. The expansion up to second order in the displacements yields (see Eq. (A4) for detailed expansions and error bounds)

$$U(\mathbf{r} + \delta \mathbf{r}) = -B\mathcal{E} + \frac{1}{2} \langle \delta \mathbf{r}, M \delta \mathbf{r} \rangle, \tag{4}$$

where M is a symmetric linear map, whose matrix representation is usually referred to as the *dynamical matrix*. When Eq. (4) is plugged into the integral in Eq. (3), the integral converges if and only if M, considered as a map $W^{\perp} \rightarrow W^{\perp}$, is positive definite, that is, all its eigenvalues are positive. The result gives the harmonic approximation for the vibrational partition function.

To see what the positive definiteness of the dynamical matrix implies about the rigidity of the cluster, let us consider how one might deform a cluster along a path $\mathbf{p}(t)$ with $\mathbf{p}(0) = \mathbf{r}$ without changing any of the bond lengths. Taking one derivative of Eq. (1) shows that

$$\mathbf{r}_{ij} \cdot \mathbf{p}'_{ij}(0) = 0 \quad \text{for all } (i,j) \in E.$$
 (5)

Defining the linear map $R: \mathbb{R}^{3N} \to \mathbb{R}^B$ that maps $\mathbf{p}'(0)$ to the B-dimensional vector with entries given by the left hand sides of Eq. (5), we see that $\mathbf{p}(t)$ maintains the constraints Eq. (1) only if $\mathbf{p}'(0)$ is in the null space of R. The matrix representation of R is known as the *rigidity matrix* in the mathematical study of framework rigidity. An element of the null space of R is called a *flex*, and it is *trivial* if it is a rigid-body motion (that is, in W). If the only flexes are trivial, then the cluster is said to be *infinitesimally rigid* or *first-order rigid*. This is sufficient for the cluster to be rigid in the nonlinear sense [38]. The dynamical map (matrix) and the rigidity map (matrix) are related as $R^*R = Md^2/a$, where R^* is the adjoint (transpose), so the harmonic approximation converges exactly when the cluster is first-order rigid.

The harmonic approximation fails when the null space of the dynamical matrix extends beyond W, or equivalently when nontrivial flexes of the rigidity matrix exist. These are infinitesimal degrees of freedom that may or may not be extendable to finite degrees of freedom. If they are not, we call them *singular* directions (or modes or flexes). The cluster in Fig. 1 has a singular direction corresponding the twisting motion described above.

To integrate along these singular directions we continue the expansion of the potential energy function. Let X be the space of singular direction: $X = \operatorname{ns}(M) \cap W^{\perp}$, where $\operatorname{ns}(\cdot)$ denotes the null space. Let $Y = [\operatorname{ns}(M)]^{\perp}$ be its orthogonal complement in W^{\perp} . For every $\delta \mathbf{r} \in W^{\perp}$ we may write $\delta \mathbf{r} = \mathbf{x} + \mathbf{y}$, where $\mathbf{x} \in X$, $\mathbf{y} \in Y$. We plug this decomposition into the potential energy function, use the fact that $\mathbf{x}_{ij} \cdot \mathbf{r}_{ij} = 0$ for all $(i, j) \in E$ by Eq. (5), and keep terms up to $O((|\mathbf{x}|^2 + |\mathbf{y}|)^2)$ (see Eq. (A5) for these calculations and error bounds). We obtain, neglecting the error terms,

$$U(\mathbf{x}, \mathbf{y}) = -B\mathcal{E} + ad^2 \tilde{U}_0(\mathbf{x}) + ad\langle \tilde{\mathbf{u}}_1(\mathbf{x}), \mathbf{y} \rangle + \frac{1}{2}a\langle \mathbf{y}, \tilde{M}\mathbf{y} \rangle.$$
(6)

Here $\tilde{M} = M/a$, $\tilde{U}_0(\mathbf{x})$ is a real scalar, and $\tilde{\mathbf{u}}_1(\mathbf{x}) \in Y$ is a 3N-dimensional vector, constructed as

$$\tilde{U}_0(\mathbf{x}) = \sum_{(i,j)\in E} \frac{1}{8d^4} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij})^2, \tag{7}$$

$$[\tilde{\mathbf{u}}_1(\mathbf{x})]_i = \sum_{j \text{ s.t. } (i,j) \in E} \frac{1}{2d^3} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) \mathbf{r}_{ij}. \tag{8}$$

If $U(\mathbf{x}, \mathbf{y}) - U(0,0)$ is positive for all nonzero (\mathbf{x}, \mathbf{y}) we can calculate the leading-order term of the partition function (see Eq. (A18)):

$$Z_{\mathbf{r}} = \frac{I\Omega e^{\beta B \mathcal{E}}}{(\det \tilde{M}|_{Y})^{1/2} \sigma} \left(\frac{2\pi}{a\beta}\right)^{\frac{d_{Y}}{2}} \left(\frac{d^{2}}{a\beta}\right)^{\frac{d_{X}}{4}} \left(\int_{X} e^{-Q(\tilde{\mathbf{x}})} d\tilde{\mathbf{x}}\right), \quad (9)$$

where d_X, d_Y are the dimensions of X and Y, and

$$\begin{split} Q(\tilde{\mathbf{x}}) &= \frac{1}{ad^2} \min_{\mathbf{y} \in Y} U(d\tilde{\mathbf{x}}, \mathbf{y}) \\ &= \tilde{U}_0(d\tilde{\mathbf{x}}) - \frac{1}{2} \langle \tilde{\mathbf{u}}_1(d\tilde{\mathbf{x}}), \tilde{M}^{-1} \tilde{\mathbf{u}}_1(d\tilde{\mathbf{x}}) \rangle \end{split} \tag{10}$$

is a dimensionally reduced quartic form that gives the minimum of the potential at a fixed displacement along the singular directions. Letting $\gamma = e^{\beta \mathcal{E}}$, $\alpha = (a\beta d^2)^{1/4}$, and $B_{\rm ISO} = 3N - 6$, we can write Eq. (9) as

$$Z_{\mathbf{r}} = \Omega d^{3N-3} (2\pi)^{-\frac{B_{\rm ISO}}{2}} \gamma^B \alpha^{-2B_{\rm ISO} + d_X} z_{\mathbf{r}}, \tag{11}$$

where

$$z_{\mathbf{r}} = \frac{I}{\sigma d^3} (\det \tilde{M}|_{Y})^{-\frac{1}{2}} (2\pi)^{-\frac{d_X}{2}} \int_{X} e^{-Q(\tilde{\mathbf{x}})} d\tilde{\mathbf{x}}.$$
 (12)

Equations (11) and (12) are our main result. They express the partition function for a sticky rigid cluster as a product of a number of dimensional quantities that depend on the pair potential, temperature, and particle dimensions, times a dimensionless quantity $z_{\mathbf{r}}$. We call $z_{\mathbf{r}}$ the geometrical partition function because it depends only on the positions of the particles, which are given by solving the geometry problem defined by Eq. (1) and not on any externally controllable

parameters. The interaction potential and temperature enter *only* in combination through the two parameters γ and α , which measure, respectively, the contribution of each extra bond and each singular direction. Remarkably, only the second derivative of the interaction potential affects the free energy. All higher-order derivatives contribute to subleading order.

When does the remaining integral in Eq. (12) converge? We will show it converges for all clusters that are *second-order rigid*. To explain what this means, consider again the hypothetical deformation of the cluster $\mathbf{p}(t)$, and now take the second time derivative of Eq. (1):

$$\mathbf{r}_{ij} \cdot \mathbf{p}_{ij}''(0) + \|\mathbf{p}_{ij}'(0)\|^2 = 0, \tag{13}$$

for all $(i, j) \in E$. A given flex $\mathbf{p}'(0)$ can be extended to a second-order motion only if there is a solution $\mathbf{p}''(0)$ to the linear Eq. (13). If such a solution exists, the pair $(\mathbf{p}'(0), \mathbf{p}''(0))$ is called a second-order flex. If there is no nontrivial second-order flex, then the cluster is second-order rigid. Second-order rigidity is a sufficient condition for rigidity [38]. However, analogous higher-order versions do not necessarily imply rigidity [44].

Since $U(\mathbf{x}, \mathbf{y}) - U(0,0) = \sum_{(i,j) \in E} \frac{a}{2d^2} [(\mathbf{r}_{ij} \cdot \mathbf{y}_{ij}) + \frac{1}{2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij})]^2$, then $U(\mathbf{x}, \mathbf{y}) = U(0,0)$ for some $\mathbf{x} \in X$ and $\mathbf{y} \in Y$ if and only if $\mathbf{r}_{ij} \cdot \mathbf{y}_{ij} + \frac{1}{2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) = 0$ for all $(i,j) \in E$, namely, $(\mathbf{x}, 2\mathbf{y})$ is a second-order flex. If the cluster is second-order rigid then no such flex exists, so $U(\mathbf{x}, \mathbf{y}) > U(0,0)$ for all nonzero (\mathbf{x}, \mathbf{y}) , and Eq. (12) converges.

B. Square well potential

We have assumed a potential with nonzero second derivative at the minimum, but Baxter's original sticky-sphere limit considered a square well potential of depth $\mathcal E$ and width ϵ , where the derivatives $V^{(n)}(d)$ vanish for all n [10]. However, we can show (Appendix A 4) that while this difference changes the prefactors in the calculation, it does not change how the partition function scales with the parameters γ and α , which are defined for a square well potential to be $\gamma = e^{\beta \mathcal E}$, $\alpha = (d/\epsilon)^{1/2}$. Therefore, our calculations are a natural extension of Baxter's limit.

C. Floppy clusters

We have derived Eq. (11) for rigid clusters, but a similar calculation can be performed for floppy clusters, those with internal degrees of freedom along which the cluster can deform by some finite amount. In this case, some of the zero vibrational modes extend to finite degrees of freedom, and some do not so they are singular directions. We suppose the internal degrees of freedom form a manifold \mathcal{M} with dimension m, and that the number of singular modes d_X is constant over the manifold. Now X, Y are the linear subspaces along which the cluster is rigid, with X representing the singular directions and Y the nonzero vibrational modes. The zero vibrational modes are the union of X and the tangent space to \mathcal{M} . If \mathcal{M} has dimension m, then we must have $\dim(X) + \dim(Y) + m = 3n - 6$.

The additional calculations to deal with the floppy degrees of freedom closely resemble those performed in Ref. [11] and result in an integral over the internal degrees of freedom of a cluster. The resulting partition function still has the form of

Eq. (11), where the geometric partition function is now $z_{\mathbf{r}}^{\text{tot}}$ given by

$$z_{\mathbf{r}}^{\text{tot}} = \int_{\mathcal{M}} z_{\mathbf{r}}(\mathbf{x}) d\sigma_{\mathcal{M}}(\mathbf{x}), \tag{14}$$

with $z_{\mathbf{r}}(\mathbf{x})$ given by Eq. (12), and in both Eqs. (11) and (12) we must set $B_{\rm ISO} = 3N - 6 - m$. Here $d\sigma_{\mathcal{M}}$ is the natural surface measure on the manifold, described in detail in Ref. [11].

III. FREE ENERGY LANDSCAPE OF SMALL CLUSTERS

Equation (11) shows that extra bonds and singular directions contribute to the stability of a rigid cluster by factors γ and α , which both diverge in the sticky limit. This observation raises some natural questions: for finite values of these parameters, which kinds of cluster tend to dominate the partition function? How does this answer depend on N? And, with a sight toward emergence, how do close-packing fragments come to dominate the landscape as $N \to \infty$?

We address these questions by calculating the geometrical partition functions for all known rigid clusters (see Appendix B I for methods). We use the enumeration of rigid clusters produced by Holmes-Cerfon [20], which includes rigid clusters up to N=14, and rigid clusters containing a given number of contacts or more up to N=21 [45]. These are thought to be nearly complete lists of rigid clusters for each N and for each specified number of contacts. All clusters in the lists are second-order rigid to numerical tolerance, so Eq. (11) converges. We use the results to characterize the competition between singular and hyperstatic clusters at each value of N.

A.
$$N \leq 8$$

All clusters of these sizes are regular (isostatic and nonsingular). For $N \le 5$, there is only a single rigid cluster for each value of N: a single sphere, a dimer, a triangle, a tetrahedron, and a triangular bipyramid. For N = 6, 7, and 8, multiple rigid clusters exist, all nonsingular, and all with the same number of bonds. The most important factor in the partition function is the symmetry number, with low-symmetry clusters dominating high-symmetry clusters [16,46,47].

B.
$$N = 9$$

The smallest singular rigid cluster, illustrated in Fig. 1, occurs at N=9. Including this cluster, there are 52 rigid clusters, not counting enantiomers, all with the same number of bonds [17]. So, the partition function is given, up to constant factors, by $Z=\alpha z_1+\sum_{i=2}^{52}z_i$, where z_i is the geometric term in Eq. (11) for the *i*th cluster, the singular cluster being first. The equilibrium probability of the singular cluster is then

$$P_1 = \frac{\alpha}{K + \alpha},\tag{15}$$

with $K = \frac{1}{z_1} \sum_{i=2}^{52} z_i \approx 235$. In the sticky sphere limit, $\alpha \to \infty$, and this probability approaches one. However, we expect this estimate to hold even for finite α .

To test this prediction, we sampled from the canonical ensemble of nine spheres interacting via a Morse-harmonic potential with various ranges and depths using a Monte Carlo simulation (see Appendix B 2 for methods). The diameter is

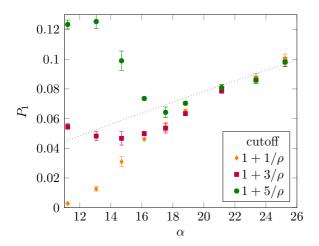


FIG. 3. Simulation results for N=9 sticky spheres with $\kappa\approx 220$ and $\rho=30,40,50,60,70,80,100,120$, and 140 interpreted using different bond cutoffs. The dashed line shows the theoretically predicted frequency of the singular cluster, Eq. (15). Error bars are standard deviation estimates.

taken to be 1, and the range of the interaction is characterized by a parameter ρ . For a given value of the range parameter ρ the pair potential is active over a distance of about $5/\rho$, after which it has decayed to about 1.3% of its minimum depth. For each sampled structure we construct an adjacency matrix by specifying a cutoff distance for bonds and identify the rigid cluster, if any, whose adjacency matrix is isomorphic to the constructed one. For small values of the range parameter ρ (corresponding to a large interaction range), this method was extremely sensitive to the cutoff being used, but for $\rho \gtrsim 70$ (corresponding to $\alpha \gtrsim 18$), the calculated cluster frequencies converged and agreed with the theoretical prediction, as illustrated in Fig. 3.

The discrepancy for small ρ likely arises because of interaction between nonnearest neighbors. The rigid cluster geometries are minima of $U(\mathbf{r})$ only in the limit $\rho^{-1} \to 0$ and are deformed for $\rho^{-1} > 0$ due to interactions between spheres that are not in contact in the ideal structure. The minimum gap in the rigid clusters of 9 spheres is $h_{\min} \approx 0.05d$, which occurs in seven of the 52 clusters, and 30 clusters have gaps of 0.09d or less. When $\rho^{-1}/h_{\min} \ll 1$, the deformations of the local minima are substantial. Small gaps also lead to problems in the identification of cluster geometries in our numerical sample, since they could be identified as bonds.

Even though Eq. (15) applied only in the limit that $\alpha \to \infty$, and so $P_1 \to 1$, we observe remarkable agreement even when P_1 is only around 8%.

Experiments that observed colloids interacting via a depletion interaction found the singular cluster occurred with frequency 11% (95% confidence interval 4%–27%) [16]. A calculation neglecting the coupling between the zero and nonzero modes predicted only a 3% probability for the singular cluster [16], and the excess stability was surmised to correspond to the free energy of about half of an extra bond. The range of attraction was estimated to be about 1.05 times the particle diameter, which corresponds to a range parameter of $\rho \approx 30$ [46,48]. The depth of interaction was estimated to be about $4k_BT$, but this gives a sticky parameter of $\kappa = 1.6$ for which clusters should melt; a depth of $\approx 8k_BT$ is probably

closer to the truth so that $\kappa \approx 60$, and is consistent with recent simulations [49]. These values give $\alpha \approx 11$ so the theoretically predicted probability is $P_1 \approx 0.045$, within the experimental confidence interval. These parameters are in the regime where non-nearest-neighbor interactions are relevant, so we cannot hope the calculation will agree exactly with the experimentally observed value, but it is still notable that the two are consistent.

C.
$$10 \le N \le 21$$

To analyze the landscape for higher N, we partition the rigid clusters based on the number of extra bonds $\Delta B = B - (3N - 6)$ above or below the isostatic number, and the number of singular zero modes d_X . The canonical partition function is then

$$Z = \kappa^{3N-6} \sum_{\Delta B, d_X} \alpha^{d_X} \gamma^{\Delta B} z_{\Delta B, d_X}, \tag{16}$$

where $z_{\Delta B,d_X}$ is the sum of the geometric contributions of all the clusters in a given partition (see Appendix C for values), and we have used the fact that the sticky parameter Eq. (2) is

$$\kappa = \frac{\sqrt{2\pi}\gamma}{\alpha^2}.\tag{17}$$

We analyze the landscapes for $N \ge 10$ by determining which term in Eq. (16) is the largest for different values of α and γ .

The phase diagram for N = 10, shown in Fig. 4, indicates where on the parameter space each term dominates. There are four terms in the partition function, corresponding to clusters that are regular ($\Delta B = 0$, $d_X = 0$), singular isostatic ($\Delta B =$ 0, $d_X = 1$), hyperstatic ($\Delta B = 1$, $d_X = 0$), and hypostatic $(\Delta B = -1, d_X = 2)$. When γ is large enough and α is fixed or grows more slowly than γ , the hyperstatic cluster dominates. We also see that for fixed γ and large enough α , the hypostatic cluster appears to dominate. However, in that region of parameter space, the sticky parameter κ tends to zero and we do not expect the spheres to leave the fluid phase and stick to each other. Figure 4 shows lines of constant κ for $\kappa = 10$ and 1000; these are values between which we could reasonably expect to see clusters that equilibrate over experimental timescales, rather than a gas or a glass. Figure 4 also shows the relative frequencies as a function of α for $\kappa = 40$, a reasonable value to observe in experiments [12], where we see the transition from the region dominated by regular clusters, favored by the higher number of distinct clusters, to that dominated by the hyperstatic cluster, favored by lower energy. The singular clusters peak in frequency near this transition, but never rise above 3.6%. The hypostatic cluster is extremely rare and never reaches above 0.03% despite its high vibrational entropy.

For N=11 to 14, we use the enumeration of rigid clusters to construct similar phase diagrams, shown in Fig. 5. Again, in the region of κ where we can expect to observe the formation of equilibrated clusters, we primarily see a transition between regular clusters and the most hyperstatic ones. The exception is N=13, for which two of the eight most hyperstatic clusters are singular and dominate in the sticky limit.

For N = 15 to 21, the maximum value of ΔB is 5, 6, 7, 8, 9, 10, and 11, respectively, and only clusters with fewer than four missing bonds compared to the maximum were enumerated [20]. The phase diagrams computed in Fig. 6

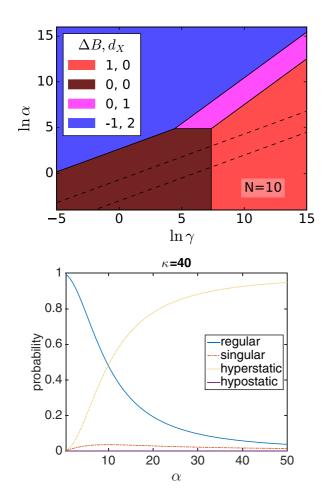


FIG. 4. Top: phase diagram for N=10 spheres. Colors (shading) indicate which kind of cluster has the largest partition function for each value of γ , α , where the clusters are grouped by the number of extra contacts ΔB and the number of singular modes d_X . Dashed lines are $\kappa=10$ (top), $\kappa=1000$ (bottom). Bottom: Relative frequency of the types of rigid clusters of 10 spheres as a function of α (roughly the inverse square root width of the potential), at $\kappa=40$.

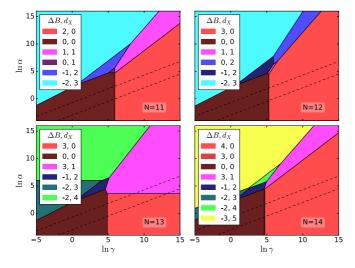


FIG. 5. Phase diagrams for rigid clusters of N = 11-14 spheres. Labeling is the same as described in the caption of Fig. 4.

are therefore not a complete description of rigid clusters, but their common appearance and patterns is a telling signal of emergence. In particular, as in most previous cases, at fixed $\kappa \gtrsim 1$ and increasing α , we observe a transition from dominance of the least hyperstatic nonsingular clusters among those enumerated to dominance of the most hyperstatic. For N=11 to 14, the former are regular clusters, but regular clusters are too numerous to have been enumerated for $N \geqslant 15$. This transition occurs in a narrow range of γ , with possibly a few intermediate regions.

IV. DISCUSSION

A. Toward larger N

Our observations about the free-energy landscape of clusters of up to 21 sticky spheres suggest two conjectures about the landscape as *N* increases:

- (1) For most values of N, singular clusters are rare at experimentally relevant values of κ .
- (2) The phase diagrams of rigid clusters have a universal shape, which is approached even at small N.

The first conjecture, if true, goes much of the way toward explaining the phenomenon of emergence. It implies that despite the apparent competition between the higher vibrational entropy of singular clusters and the lower energy of hyperstatic clusters, the latter always wins. Since the most hyperstatic clusters are close to fragments of a close-packed lattice with defects only on the surface, close-packing order should arise for sticky spheres even for small clusters.

Based on Eq. (11), we have already noted that each extra bond contributes to the stability of a rigid cluster by a factor γ , and each singular mode contributes by a factor α . This surmise neglects the geometric factor, which might be correlated with the number of bonds or singular modes, and we return to such correlations shortly. Since $\gamma \propto \kappa \alpha^2$, and assuming that $\kappa \gtrsim 1$, it follows that as $\alpha \to \infty$, each extra bond contributes at least twice as much as each singular mode. In this sense, the observation of Meng *et al.* that a zero mode contributes to stability about half as much as an extra bond turns out to be rather prescient, even as it was surmised from experiments far from the Baxter limit [16].

Therefore, the only way a rigid cluster with fewer than the maximum number of bonds can become dominant at large α is if it had two singular modes for each missing bond. This scenario does not seem to occur for any value of N. In fact, the only exceptions to the rarity of singular clusters are when singular clusters exist among the most hyperstatic ones, as they do for N = 9 and 13. We do not know if such clusters might exist for larger values of N. Such "magic numbers," associated with special clusters and deviating from the trend toward bulk behavior, are familiar in other systems, such as atomic clusters [50–53].

Motivated by the first conjecture, we limit our attention to nonsingular clusters and consider the second conjecture. The transition from a region dominated by regular clusters to one dominated by hyperstatic ones occurs not perfectly abruptly but nevertheless over a small range of γ . Such behavior implies that the total geometric partition function $z_{\Delta B}$ (we drop the second subscript for the remainder of the section, implicitly

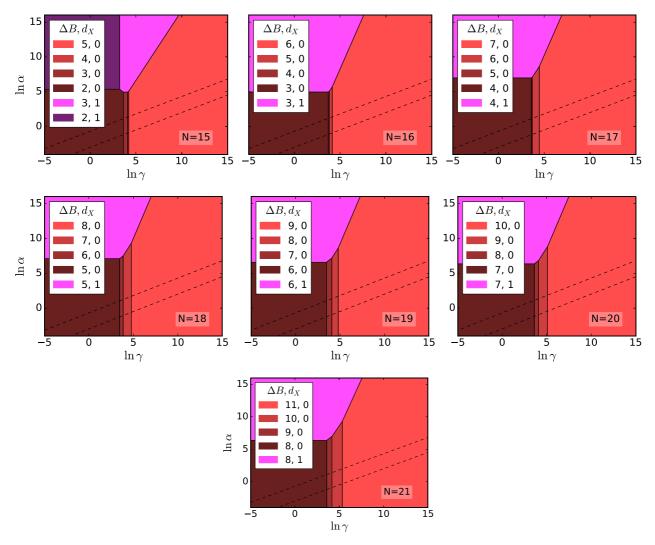


FIG. 6. Phase diagrams for rigid clusters of 15–21 spheres. Labeling is the same as described in the caption of Fig. 4.

taking $d_X = 0$) decreases exponentially as a function of ΔB with a roughly consistent rate $z_{\Delta B} \sim \exp(-\lambda \Delta B)$, as can in fact be seen in Fig. 7. The rate, and therefore also the center of the transition, at $\gamma = \exp(\lambda)$, seems to trend upwards as a function of N. We do not know if this rate approaches a constant as N increases, though we hypothesize it might, as such a limiting rate appears necessary for the total number of clusters to grow exponentially with N, an issue we will discuss in a future publication.

Which factor in the partition function contributes most strongly to this exponential scaling? Part of it is due to the number of distinct clusters with a given number of bonds, $n_{\Delta B}$, whose logarithm can be thought of as a configurational entropy. Figure 7 shows $n_{\Delta B}$ decreases exponentially with ΔB , and the gain in configuration entropy for each lost bond appears to grow with N. The average geometric partition function of a typical single cluster with a certain number of bonds also appears to decrease exponentially as a function of ΔB , with most of this decreases accounted for by the decrease in the average vibrational factor $(\det \tilde{M})^{-1/2}$ per cluster. The deviations from the exponential fits for these individual components are more pronounced than in the case

of the total geometric partition function, but the consistent exponential scaling is notable. The rotational contributions, from the moment of inertia and symmetry factors, do not depend strongly on the number of bonds (Appendix B 3.)

To further investigate how these observations bear on the question of emergence of long-range order, we consider a metric for the bond-orientational order. We use the order metric Q_6 , which is widely used in the study of the structure of liquids (see Appendix B 4 for a definition) [54]. Figure 8 shows that bond-orientational order is strongly correlated with the existence and number of extra bonds. Isostatic clusters have particularly low bond-orientation order. Deviation from isostaticity in either direction seems to require the introduction of partial order, since only when going away from the generic situation can the algebraic singularities associated with hypostaticity and hyperstaticity arise. The presence of even a single extra bond seems to require the structure to contain two octahedra sharing an edge, and as more bonds are added, this substructure can grow to incorporate larger and larger close-packing fragments. The most hyperstatic clusters observed for each N are fragments of close-packing lattices with defects only on the surface.

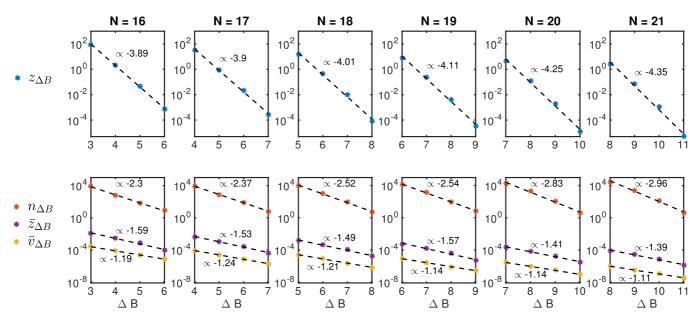


FIG. 7. Top row: the total geometric partition function of rigid clusters with ΔB bonds above the isostatic number of bonds as a function of ΔB . The dashed line is an exponential fit, $\log z_{\Delta B} = c_1 + c_2 \Delta B$, with the best fit value of c_2 indicated. The bottom row shows different contributions to the value of $z_{\Delta B}$: $n_{\Delta B}$ is the number of distinct rigid clusters, $\bar{z}_{\Delta B}$ is the mean geometric partition function of such a cluster, and $\bar{v}_{\Delta B}$ is the mean vibrational contribution to the latter.

B. Relation to bulk behavior

In addition to being physically important for their own sake, finite clusters also carry importance as a formal tool in low-density expansions of the fluid behavior of bulk systems. Baxter used the expansion up to the second term to predict a gas-liquid transition in a fluid of sticky spheres [10]. It has been previously observed that Baxter's calculation is theoretically problematic because later terms in the expansion diverge in Baxter's limit [14]. Specifically, Stell and Williams showed that for N = 12 spheres the partition function diverges, due to the existence of rigid clusters which remain rigid even when turning off the interaction between some bonded pair [14]. This divergence in fact occurs for any hyperstatic cluster, leading to divergence even for N = 10. Here, we have shown that singular clusters also lead to a divergence in Baxter's limit and so the theoretical problems for the expansion begin even earlier, at N = 9.

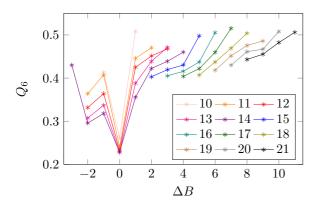


FIG. 8. Mean value of the bond-orientational order parameter Q_6 as a function of the number of extra bonds ΔB . The mean includes all clusters, regular and singular, weighted uniformly.

Our numerical results show a sharp transition to the prevalence of isostatic structures over highly hyperstatic, close-packed ones when the depth of the potential well is less than a critical value, and we expect it to extend into the bulk regime.

V. CONCLUSIONS AND FUTURE WORK

We have shown that the free energies of systems of spheres interacting with a very short-ranged pair potential can be described by only two parameters, characterizing the depth and width of the potential. This is true even when the system is in a singular energy minimum, where the harmonic approximation diverges, provided the associated framework is second-order rigid. We used our results to study the free-energy landscapes of clusters of a finite number $N \leq 21$ of spheres. Previous work has shown that for $N \leq 8$, the most likely clusters to be observed in equilibrium are the least symmetric ones. We showed that for larger N, the free-energy landscape is dominated by the most hyperstatic clusters. These clusters are usually close to fragments of a close-packed lattice with defects only on the surface, so we see the emergence of crystalline order even for N as small as 10. Only for the magic numbers, N = 9 and N = 13, is the maximum number of contacts achieved by both nonsingular and singular clusters, and the latter would form with near unit probability in the limit of zero range interactions. It would be interesting to know if there are other, larger values of N with singular clusters having the maximum number of contacts.

We found empirically that the partition function of rigid clusters of size N with B contacts scales exponentially with B. If this scaling continues, it leads to a sharp transition between a region of parameter space where clusters with the largest possible number of contacts dominate and a region where the expected number of contacts is extensively less than the

maximum. This sharp transition may be a finite-size incarnation of the thermodynamic observation that particles with a short-range interaction have little or no liquid phase, a fact that has made it hard to create liquids of biomolecules or other nanoscale objects [55]. We analyzed the factors contributing to this scaling, but do not have an explanation from first principles that would suggest the scaling continues for larger sizes, or that the rates approach a constant as *N* increases. We leave this as an open problem that could be addressed using ideas from geometry and statistical mechanics.

We have focused our calculations on rigid clusters, but one could use Eq. (14) to compute the partition functions of floppy clusters or other networks with floppy modes of deformation (see, e.g., Ref. [56]). This is more challenging because it involves computing integrals over manifolds, but is still tractable using specialized parametrizations as in Refs. [11,57] or thermodynamic integration techniques [58]. Moreover, current numerical and analytic methods for classifying the rigidity of bar frameworks focus on testing rigidity, not floppiness, and new methods are required for determining what subspace of the space of infinitesimal flexes (zero modes) extends to finite degrees of freedom in a floppy cluster.

One possible issue with using our result is the evaluation of the integral $\int_{\mathbb{R}^n} \exp[-Q(\mathbf{x})] d\mathbf{x}$, where Q is a positive quartic form. Unlike its Gaussian relative, this integral has no general analytic solution for even moderate values of N [59]. A future avenue of research is to simplify the integral under stronger rigidity assumptions, such as prestress stability.

We showed via simulations that the sticky limit describes real, finite-range potentials, provided the range of the potential is much shorter than minimum gap between noncontacting spheres. However, this minimum gap appears to become arbitrarily small as N increases [20]. Nevertheless, we expect our calculations to give qualitative insights into real systems and it would be interesting to test them with experiments on clusters larger than N=9. For example, it could still be the case that the type of clusters that dominate in equilibrium depends on only a small number of parameters as the interaction potential is varied, or that the ones that are close to singular become more probable as the range of the potential is decreased.

Although the gap size poses a problem for quantitative agreements in real systems, it may be possible to build on the present results by including perturbations that account for non-nearest-neighbor interactions. As non-nearest-neighbor interactions increase, the energy minima coming from distinct rigid clusters are expected to merge, leading to critical points that can be analyzed using catastrophe theory [60] or related ideas. Singular clusters themselves can be thought of as multiple, colocated minima that might separate as they are perturbed by non-nearest-neighbor interactions. By starting with the sticky limit and slowly increasing the range of the interaction potential, one might create a bifurcation diagram of energy landscapes, which shows a finite or low-dimensional number of possible landscapes; indeed, in related systems it has been shown that the space of energy-minimizing configurations is sometimes much lower-dimensional than the space of interaction potentials [61].

Many systems beyond clusters, particularly metamaterials and systems near jamming or close packing, are often modeled as subject to soft constraints in the limit that the stiffness

of the constraint is taken to infinity. As we have shown here, this limit is qualitatively different than modeling the constraints as hard constraints, particularly when the excitation spectrum includes zero modes. In this paper, we have focused on the thermodynamic effect of zero modes, calculating their contribution to the free energy. Such thermal effects are already observed in many metamaterials and self-assembly systems and are likely to become important for more of them as they are fabricated on smaller scales. The zero-temperature effects of zero modes are also potentially important in metamaterials and athermal fluids and more work will need to be done to determine their effect on mechanical stability and yielding behavior. We hope that the present work will become the first step in a broader program to provide analytic tools for going beyond the harmonic approximation and calculating the thermodynamic and mechanical effects of zero-frequency vibration modes.

ACKNOWLEDGMENTS

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APPENDIX A: DETAILS ON THE PARTITION FUNCTION CALCULATIONS

1. Sticky limit

Starting from an arbitrary potential $V_0(r)$, we can define a one-parameter family of rescaled potentials, parameterized by the variable ϵ , representing the interaction range. Let $V(r) = C(\epsilon)V_0[(r-d)d/\epsilon+d]$. We need to extend V_0 for negative arguments in some way such that $V_0(r) \to \infty$ as $r \to -\infty$. Assuming such scaling is used to approach the sticky limit, then all derivatives of V(r) at r=d satisfy the scaling $d^nV(r)/dr^n|_{r=d} = O(\mathcal{E}/\epsilon^n)$, where $\mathcal{E} = -V(d)$. We assume this scaling of the derivatives for deriving all the big-O error bounds in our calculation. We also assume that $V_0(r) \to 0$ as $r \to \infty$ in such a way that $V(r) \to 0$ as $\epsilon \to 0$ pointwise for all r > d.

The traditional sticky sphere limit of Baxter [10] is defined so that the difference between the partition function for two spheres interacting via V(r) and the partition function for two hard spheres approaches a constant. Namely,

$$0 < \lim_{\epsilon \to 0} \frac{1}{d^3} \int_{\mathbf{r} \in \mathbb{R}^3} (e^{-\beta V(|\mathbf{r}|)} - 1) d^3 \mathbf{r} + \frac{4\pi}{3} < \infty, \quad (A1)$$

where $\beta = (k_B T)^{-1}$ is the inverse of temperature T times the Boltzmann constant. The limit can be evaluated using Laplace's method as

$$2^{5/2} \pi^{3/2} \lim_{\epsilon \to 0} (a\beta d^2)^{-1/2} \exp(\beta \mathcal{E}),$$
 (A2)

where $\mathcal{E} = -V(d) = -C(\epsilon)V_0(d)$ and $a = V''(d) = C(\epsilon)V_0''(d)d^2/\epsilon^2$.

In our calculation we do not assume this limit, letting ϵ and C approach 0 and ∞ simultaneously in any

way desirable. We define the sticky parameter to be $\kappa = (2\pi)^{1/2} (a\beta d^2)^{-1/2} \exp(\beta \mathcal{E}) = (2\pi)^{1/2} \alpha^{-2} \gamma$ even when the limit is not taken.

2. Expansion of the pairwise potential

We consider the Taylor expansion of $V(|\mathbf{r}_{ij} + \delta \mathbf{r}_{ij}|)$ in powers of $\delta \mathbf{r}_{ij}$, where $|\mathbf{r}_{ij}| = d$. Applying the chain rule and using the Taylor expansion of the square root, we get

$$V(|\mathbf{r}_{ij} + \delta \mathbf{r}_{ij}|) = V(\sqrt{d^2 + |\delta \mathbf{r}_{ij}|^2 + 2\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij}})$$

$$= \sum_{l=0}^{\infty} \frac{1}{l!} V^{(l)}(d) \left[\sum_{k=0}^{\infty} {1/2 \choose k} d^{1-2k} (|\delta \mathbf{r}_{ij}|^2 + 2\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^k \right]^l, \tag{A3}$$

where $\binom{1/2}{k} = \frac{1}{k!} \prod_{n=0}^{k-1} (\frac{1}{2} - n)$ is the generalized binomial coefficient.

When we expand the sum and keep terms up to fourth degree in $\delta \mathbf{r}_{ij}$, recalling that V'(d) = 0, we get

$$V(\|\mathbf{r}_{ij} + \delta \mathbf{r}_{ij}\|)$$

$$= -\mathcal{E} + \frac{a}{2d^2} (\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^2 + \frac{b}{6d^3} (\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^3$$

$$+ \frac{a}{2d^2} (\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij}) (\delta \mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij}) + \frac{c}{24d^4} (\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^4$$

$$+ \frac{b}{4d^3} (\mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^2 (\delta \mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})$$

$$+ \frac{a}{8d^2} (\delta \mathbf{r}_{ij} \cdot \delta \mathbf{r}_{ij})^2 + O(\mathcal{E} \|\delta \mathbf{r}_{ij}\|^5 / \epsilon^5), \tag{A4}$$

where $\mathcal{E} = -V(d)$, a = V''(d), b = V'''(d) - 3V''(d)/d, and $c = V^{(4)}(d) - 6V'''(d)/d + 15V''(d)/d^2$.

Later on in the calculation, we decompose the displacement into orthogonal components $\delta \mathbf{r} = \mathbf{x} + \mathbf{y}$ with the property that $\mathbf{r}_{ij} \cdot \mathbf{x}_{ij} = 0$ for all $(i,j) \in E$. When we plug this decomposition into Eq. (A4) and keep terms below $O(\frac{\mathcal{E}}{\epsilon^3 d^3} (\|\mathbf{x}\|^2 + d\|\mathbf{y}\|)^3)$, we get

$$V(|\mathbf{r}_{ij} + \mathbf{x}_{ij} + \mathbf{y}_{ij}|)$$

$$= -\mathcal{E} + \frac{a}{8d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij})^2 + \frac{a}{2d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) (\mathbf{r}_{ij} \cdot \mathbf{y}_{ij})$$

$$+ \frac{a}{2d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) (\mathbf{x}_{ij} \cdot \mathbf{y}_{ij}) + \frac{a}{2d^2} (\mathbf{r}_{ij} \cdot \mathbf{y}_{ij})^2$$

$$+ \frac{a}{d^2} (\mathbf{x}_{ij} \cdot \mathbf{y}_{ij}) (\mathbf{r}_{ij} \cdot \mathbf{y}_{ij}) + \frac{a}{4d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) (\mathbf{y}_{ij} \cdot \mathbf{y}_{ij})$$

$$+ \frac{a}{2d^2} (\mathbf{x}_{ij} \cdot \mathbf{y}_{ij})^2 + O\left[\frac{\mathcal{E}}{\epsilon^3 d^3} (\|\mathbf{x}\|^2 + d\|\mathbf{y}\|)^3\right]. \quad (A5)$$

3. Free energy of a rigid cluster

Consider the coordinates of the centers of N spheres, $\mathbf{r}_i \in \mathbb{R}^3$, i = 1, ..., N, which we consider as a single point $\mathbf{r} = \mathbf{r}_1 \oplus ... \oplus \mathbf{r}_N \in \mathbb{R}^{3N}$ of the configuration space. Let $E = \{(i_1, j_1), ..., (i_B, j_B)\}$ be the complete list of pairs such that

$$|\mathbf{r}_i - \mathbf{r}_j|^2 = d^2, \quad (i, j) \in E.$$
 (A6)

We say that the point \mathbf{r} describes a rigid cluster if any other point in the connected component of the solution space to Eq. (A6) that contains \mathbf{r} is related to \mathbf{r} by translation and rotation.

a. Partition function. We ignore momenta in our phase space because they contribute a constant term to the partition function. Therefore, the partition function is defined as

$$Z = \frac{1}{N!} \int_{\mathbf{r}' \in \Omega^n} \exp[-\beta U(\mathbf{r}')] d\mathbf{r}', \tag{A7}$$

where Ω is the region of space to which the particles are confined. We use Ω to denote also the volume of this region.

b. Domain of integration. The free energy of a particular rigid cluster \mathbf{r} is defined to be $-\beta \log Z_{\mathbf{r}}$, where $Z_{\mathbf{r}}$ is the contribution to the total partition function due to the configurations that are associated with the rigid cluster \mathbf{r} . The space of such configurations, denoted $\mathcal{N}_{\mathbf{r}}$, is defined to be the union over all permutations $\rho \in S_N$ of domains of the form

$$\{\mathbf{r}' \in \Omega^N : \mathbf{r}'_i = U(\mathbf{r}_{\rho(i)}) + \mathbf{t} + \delta \mathbf{r}_i$$
 for some $U \in O(3), \mathbf{t} \in \mathbb{R}^3, \delta \mathbf{r} \in \mathbb{R}^{3N}$ such that $|\delta \mathbf{r}_i| \leq l\}$, (A8)

where l is chosen such that $\epsilon \ll l \ll d$.

c. Change of variables. We can remove the rigid-body degrees of freedom and the permutations from the calculation by performing a change of variables. Fix \mathbf{r} to be some configuration corresponding to the cluster of interest. Let $W = \{\delta \mathbf{r} \in \mathbb{R}^{3N} : \delta \mathbf{r}_i = \mathbf{t} + \mathbf{s} \times \mathbf{r}_i \text{ for some } \mathbf{t}, \mathbf{s}\} \subseteq \mathbb{R}^{3N}, \text{ and let } W^{\perp} \text{ be its orthogonal complement as a linear subspace of } \mathbb{R}^{3N}$. The transformation $\phi : W^{\perp} \times \mathbf{R}^3 \times \mathrm{SO}(3) \to \mathbf{R}^{3N}$ that maps $\phi : (\delta \mathbf{r}_i, \mathbf{t}, U) \mapsto U(\mathbf{r}_i + \delta \mathbf{r}_i) + \mathbf{t} - \mathbf{r}_i$ is nonsingular and its Jacobian determinant is

$$J_{\phi}(\delta \mathbf{r}) = I + O(d^2 \|\delta \mathbf{r}\|), \tag{A9}$$

where I is the square root of the determinant of the moment of inertia tensor about the center of mass. Summing over permutations contributes a factor of $N!/\sigma$, where σ is the order of the symmetry group of the cluster, that is, the number of Euclidean motions (combinations of reflections, rotations, and translations) that map \mathbf{r} to a permutation of itself. In the new variables, the integrand is independent of U and \mathbf{t} , and we can integrate those variables. The only complication is that the range over which \mathbf{t} varies, the free volume available to the cluster, can be slightly different for different clusters due to boundary effects. However, these differences are of order $\Omega^{2/3}d$, compared to the total volume Ω , so they are negligible if $\Omega \gg d^3$.

Following the change of variables and the integration over the rigid body degrees of freedom, $Z_{\mathbf{r}}$ reduces to

$$Z_{\mathbf{r}} = \frac{\Omega}{\sigma} \int_{\delta \mathbf{r} \in W^{\perp}, |\delta \mathbf{r}_{i}| \leq l} \exp[-\beta U(\mathbf{r} + \delta \mathbf{r})] J_{\phi}(\delta \mathbf{r}) d(\delta \mathbf{r}), \text{ (A10)}$$

where we have removed a factor of N! and the dimensionless factor obtained from integrating over SO(3). The inclusion of N! depends on whether we wish to treat the spheres as distinguishable or not, but in either case it will not change the results as it is common to all clusters. The free energy is given by $F_{\bf r} = -\beta^{-1} \log Z_{\bf r}$. Finally, we can replace $U({\bf r} + \delta {\bf r})$ by

 $U_E(\mathbf{r} + \delta \mathbf{r}) = \sum_{(i,j) \in E} V(\mathbf{r}_{ij} + \delta \mathbf{r}_{ij})$ and extend the domain of integration to the entire linear space W^{\perp} . This contributes an error of order $U_{\rm nb} = \sum_{(i,j) \notin E} V(\mathbf{r}_{ij})$ in the energy, whose exact order compared to other error terms depends on how the pair potential $V_0(r)$ used to construct the sticky limit approaches zero as $r \to \infty$:

$$Z_{\mathbf{r}} = \frac{\Omega}{\sigma} \int_{\delta \mathbf{r} \in W^{\perp}} \exp[-\beta U_{E}(\mathbf{r} + \delta \mathbf{r}) + O(\beta U_{\text{nb}})] J_{\phi}(\delta \mathbf{r}) d(\delta \mathbf{r}).$$
(A11)

d. Harmonic approximation. To evaluate the integral in Eq. (A11), we use the expansion Eq. (A4). Keeping terms up to second order in the displacements yields

$$U_E(\mathbf{r} + \delta \mathbf{r}) = -B\mathcal{E} + \frac{1}{2} \langle \delta \mathbf{r}, M \delta \mathbf{r} \rangle + O(\mathcal{E} \|\delta \mathbf{r}\|^3 / \epsilon^3),$$
 (A12)

where M is a symmetric linear map, whose matrix representation is usually referred to as the *dynamical matrix*. When Eq. (A12) is plugged into the integral in Eq. (A11) without the error term, the integral converges if and only if M, considered as a map $W^{\perp} \rightarrow W^{\perp}$ is positive definite, that is, all its eigenvalues are positive. The result gives the harmonic approximation for the vibrational partition function and the error term contributes a multiplicative correction of the form $1 + O(\beta^{-1/2}\mathcal{E}^{-1/2} + \beta U_{\rm nb})$.

e. Fourth-order approximation. The harmonic approximation fails when the null space of the dynamical matrix extends beyond W, or equivalently when nontrivial flexes of the rigidity matrix exist. These are infinitesimal degrees of freedom that, if the cluster is rigid, are not extendable to finite degrees of freedom. We call them singular directions. Let $X = ns(M) \cap W^{\perp}$ and $Y = [ns(M)]^{\perp}$, so that $W^{\perp} = X \oplus Y$ is an orthogonal decomposition of W^{\perp} and X is the null space of M restricted to W^{\perp} . For every $\delta \mathbf{r} \in W^{\perp}$ we may write $\delta \mathbf{r} = \mathbf{x} + \mathbf{y}$ where $\mathbf{x} \in X$, $\mathbf{y} \in Y$. Using the expansion Eq. (A5), we get

$$U_E(\mathbf{r} + \mathbf{x} + \mathbf{y}) = -B\mathcal{E} + U_0(\mathbf{x}) + \langle \mathbf{u}_1(\mathbf{x}), \mathbf{y} \rangle$$

+ $\frac{1}{2} \langle \mathbf{y}, U_2(\mathbf{x})\mathbf{y} \rangle + O(\mathcal{E} \|\mathbf{y}\|^3 / \epsilon^3).$ (A13)

Here $U_0(\mathbf{x})$ is a real scalar, $\mathbf{u}_1(\mathbf{x}) \in Y$ is a 3N-dimensional vector, and $U_2(\mathbf{x}): Y \to Y$ is a real symmetric map. These have the explicit forms

$$U_0(\mathbf{x}) = \sum_{(ij)\in E} \frac{a}{8d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij})^2 + O(\mathcal{E}\|x\|^6/d^3\epsilon^3), \quad (A14)$$
$$[\mathbf{u}_1(\mathbf{x})]_i = \sum_{j \text{ s.t. } (i,j)\in E} \frac{a}{2d^2} (\mathbf{x}_{ij} \cdot \mathbf{x}_{ij}) \mathbf{r}_{ij}$$
$$+ O(\mathcal{E}\|x\|^3/d^2\epsilon^2 + \mathcal{E}\|x\|^4/d^2\epsilon^3), \quad (A15)$$

$$U_2(\mathbf{x}) = M + O(\mathcal{E} \|\mathbf{x}\|^2 / d\epsilon^3). \tag{A16}$$

We write $U_0(\mathbf{x}) = ad^2\tilde{U}_0(\mathbf{x}) + O(\mathcal{E}\|\mathbf{x}\|^6/d^3\epsilon^3)$, $\mathbf{u}_1(\mathbf{x}) = ad\tilde{\mathbf{u}}_1(\mathbf{x}) + O(\mathcal{E}\|\mathbf{x}\|^3/d^2\epsilon^2 + \mathcal{E}\|\mathbf{x}\|^4/d^2\epsilon^3)$, and $U_2(\mathbf{x}) = a\tilde{M} + O(\mathcal{E}\|\mathbf{x}\|^2/d\epsilon^3)$, where $\tilde{U}_0(\mathbf{x})$ is a homogeneous quartic function of \mathbf{x} , $\tilde{\mathbf{u}}_1(\mathbf{x})$ is a homogeneous quadratic function of \mathbf{x} , and $\tilde{M} = M/a$ is the geometric part of the dynamical matrix. Let

$$\tilde{U}(\mathbf{x}, \mathbf{y}) = \tilde{U}_0(\mathbf{x}) + \frac{1}{d} \langle \tilde{\mathbf{u}}_1(\mathbf{x}), \mathbf{y} \rangle + \frac{1}{2d^2} \langle \mathbf{y}, \tilde{M}\mathbf{y} \rangle.$$
 (A17)

As discussed in the text, $\tilde{U}(\mathbf{x}, \mathbf{y})$ is positive for all nonzero (\mathbf{x}, \mathbf{y}) if and only if the cluster is *second-order rigid*.

In this case, we can calculate the leading-order term of the partition function. We first integrate over *Y* by completing the square and computing the simple Gaussian integral. The remaining integral over *X* is nontrivial [59], and we leave it unevaluated, but dimensionally reduce it:

$$Z = \frac{I\Omega e^{\beta B\mathcal{E}}}{\sigma} \int_{X} d\mathbf{x} \int_{Y} d\mathbf{y} \exp[-a\beta d^{2}\tilde{U}(\mathbf{x}, \mathbf{y}) + O(\epsilon_{1})]$$

$$= \frac{I\Omega e^{\beta B\mathcal{E}}}{(\det \tilde{M}|_{Y})^{1/2}\sigma} \left(\frac{2\pi}{a\beta}\right)^{\frac{d_{Y}}{2}}$$

$$\times \int_{X} d\mathbf{x} \exp[-a\beta d^{2}\tilde{U}_{\min}(\mathbf{x}) + O(\epsilon_{2})]$$

$$= \frac{I\Omega e^{\beta B\mathcal{E}}}{(\det \tilde{M}|_{Y})^{1/2}\sigma} \left(\frac{2\pi}{a\beta}\right)^{\frac{d_{Y}}{2}} \left(\frac{d^{2}}{a\beta}\right)^{\frac{d_{X}}{4}} \left(\int_{X} e^{-Q(\tilde{\mathbf{x}})} d\tilde{\mathbf{x}}\right)$$

$$\times [1 + O(\epsilon_{3})], \tag{A18}$$

where $\tilde{U}_{\min}(\mathbf{x}) = \tilde{U}_0(\mathbf{x}) - \frac{1}{2} \langle \tilde{\mathbf{u}}_1(\mathbf{x}), \tilde{M}^{-1} \tilde{\mathbf{u}}_1(\mathbf{x}) \rangle$ is the minimum of $\tilde{U}(\mathbf{x}, \mathbf{y})$ at fixed \mathbf{x} , $Q(\tilde{\mathbf{x}}) = \tilde{U}_{\min}(d\tilde{\mathbf{x}})$ is its geometric part, d_X and d_Y are the dimensions of X and Y, and the error terms are $\epsilon_1 = \frac{\beta \mathcal{E}}{\epsilon^3 d^3} (\|\mathbf{x}\|^2 + d\|\mathbf{y}\|)^3 + \frac{\beta \mathcal{E}}{d^2 \epsilon^2} \|x\|^3 \|y\| + \frac{1}{d} (\|\mathbf{x}\| + \|\mathbf{y}\|) + \beta U_{\text{nb}}$, and ϵ_2 is given by ϵ_1 by substituting $(a\beta)^{-1/2}$ for $\|\mathbf{y}\|$, and ϵ_3 by substituting $(a\beta/d^2)^{-1/4}$ for $\|\mathbf{x}\|$. The final error term is $\epsilon_3 = \beta^{-1/4} \mathcal{E}^{-1/4} d^{-1/2} \epsilon^{1/2} + \beta^{-1/2} \mathcal{E}^{-1/2} + \beta U_{\text{nb}}$.

Neglecting the error, this can be written as

$$Z = \Omega d^{3N-3} \gamma^B \alpha^{-2(3N-6)+d_x} z,$$

$$z = (I/d^3) \sigma^{-1} (\det \tilde{M}|_Y)^{-1/2} (2\pi)^{(3N-6-d_X)/2} \int_X e^{-Q(\tilde{\mathbf{x}})} d\tilde{\mathbf{x}},$$
where $\alpha = (a\beta d^2)^{1/4}$ and $\gamma = e^{\beta \mathcal{E}}$. (A19)

4. Relation to a square-well potential

The calculation presented in the text assumes that the pair potential V(r) has a nonzero second derivative at the minimum. However, many models of sticky spheres, including Baxter's original calculation, use a square-well pair potential given by

$$V_{\text{square}}(r) = \begin{cases} +\infty & r < d - \epsilon \\ -\mathcal{E} & d - \epsilon \leqslant r < d + \epsilon \\ 0 & r \geqslant d + \epsilon \end{cases}$$
 (A20)

Here we show that if the partition function of a rigid cluster for an analytic potential is given by $Z_{\rm analytic} \sim \exp(-\beta U_{\rm min})\alpha^{-M}$, where $\alpha = (a\beta d^2)^{1/4}$, a = V''(r), and M is a real exponent, then the partition function for the same cluster under a square-well pair potential is given by $Z_{\rm square} \sim \exp(-\beta U_{\rm min})\alpha^{-M}$, where $\alpha = (d/\epsilon)^{1/2}$, and M is the same exponent as for $Z_{\rm analytic}$. Explicitly, the notation $f \sim g$ above is used to mean ag < f < bg asymptotically for some constants a and b.

Let us show a more general result first. Let f_1, \ldots, f_B : $\mathbb{R}^N \to \mathbb{R}$ be real continuous functions such that $f_i(0) = 0$.

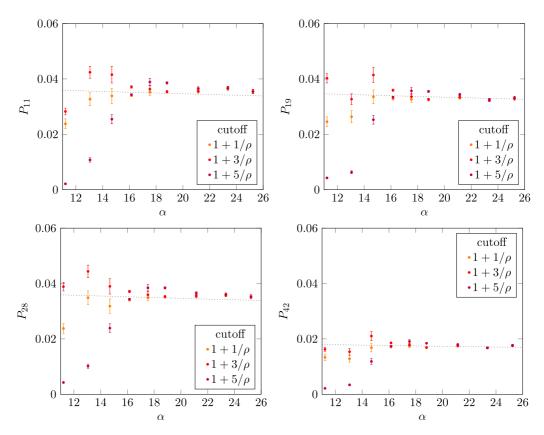


FIG. 9. Simulation results for N=9 sticky spheres with $\kappa\approx 220$ and $\rho=30,40,50,60,70,80,100,120$, and 140 interpreted using different bond cutoffs. We show the observed frequencies of four of the 51 nonsingular clusters. The dashed line shows the theoretically predicted frequency. Compare to Fig. 3, which plots the same data for the singular cluster.

Define the following integrals:

$$Z_{1} = \int_{\mathbb{R}^{N}} \prod_{i=1}^{B} \exp\left(-\frac{1}{2}\beta a f_{i}(\mathbf{r})^{2}\right) d\mathbf{r},$$

$$Z_{2} = \int_{\mathbb{R}^{N}} \prod_{i=1}^{B} \mathbf{1}(-\epsilon \leqslant f_{i}(\mathbf{r}) \leqslant \epsilon) d\mathbf{r},$$

$$D_{1}(t) = \int_{\mathbb{R}^{N}} \delta\left(t - \sum_{i=1}^{B} f_{i}(\mathbf{r})^{2}\right) d\mathbf{r},$$

$$D_{2}(t) = \int_{\mathbb{R}^{N}} \delta\left(t - \max_{i=1,\dots,B} f_{i}(\mathbf{r})^{2}\right) d\mathbf{r}.$$
(A21)

We have

$$Z_{1} = \int_{t=0}^{\infty} D_{1}(t) \exp\left(-\frac{1}{2}\beta at\right) dt$$

$$Z_{2} = \int_{t=0}^{\epsilon^{2}} D_{2}(t) dt.$$
(A22)

We also have that

$$\int_{t'=0}^{t/B} D_2(t')dt' \leqslant \int_{t'=0}^{t} D_1(t')dt' \leqslant \int_{t'=0}^{t} D_2(t')dt', \quad (A23)$$

since $\max f_i^2 \leq \sum f_i^2 \leq B \max f_i^2$. Therefore, we have the following sequence of implications:

$$Z_1 \sim (\beta a)^{-M} \Leftrightarrow D_1(t) \sim t^{M-1} \Leftrightarrow D_2(t) \sim t^{M-1} \Leftrightarrow Z_2 \sim \epsilon^{2M}.$$
(A24)

As a special case, we have that if

$$Z_{\rm analytic} = \Omega d^{3N-3} (2\pi)^{-B_{\rm ISO}/2} \exp(B\beta \mathcal{E}) (a\beta d^2)^{-M} z,$$

then

$$Z_{\text{square}} = \Omega d^{3N-3} (2\pi)^{-B_{\text{ISO}}/2} \exp(B\beta \mathcal{E}) (\epsilon/d)^{2M} z'.$$

However, the geometric parts z and z' can be different from each other for the two potentials.

APPENDIX B: CALCULATING THE PARTITION FUNCTIONS

1. Numerical method to calculate the geometric partition functions

Evaluating most of the quantities in Eq. (12) is straightforward. The two quantities that are worth discussing are the integral over the singular subspace, and the symmetry number.

a. Integrating the exponential of a quartic

Here is how we numerically calculate

$$\int_{\mathbf{v}} e^{-Q(\tilde{\mathbf{x}})} d\tilde{\mathbf{x}}.$$
 (B1)

If $d_X = 1$, then we use the fact that $\int_{-\infty}^{\infty} e^{-x^4} dx = 2\Gamma(5/4)$ to write Eq. (B1) as $2\Gamma(5/4)(Q(\mathbf{v}))^{-1/4}$, where $\mathbf{v} \in X$ is a unit vector.

If $d_X \ge 2$ we integrate Eq. (B1) numerically. While one can derive analytic expressions for the integral, even for a

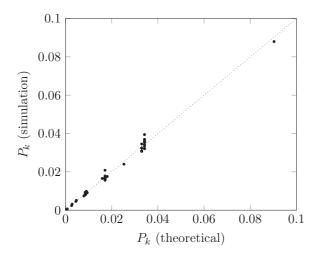


FIG. 10. Predicted vs. observed frequencies for the 52 different rigid cluster geometries for N=9 sticky spheres with $\kappa \approx 220$ and $\rho=120$, using a cutoff of $1+1/\rho$.

two-dimensional integral these are unwieldy [59, see, e.g., Table 2, formula for n=2,r=3]. Our numerical method is as follows: form an orthogonal basis $\{\mathbf{v}_i\}_{i=1}^{d_X}$ of X. For each direction \mathbf{v}_i , determine a value s_i such that $Q(s_i\mathbf{v}_i) \in [10^{-19}, 10^{-16}]$. Then for $2 \le d_X \le 4$ we integrate over a box $[-s_1, s_1] \times \cdots \times [-s_{d_X}, s_{d_X}]$ using the trapezoidal rule with equally spaced points. For $d_X = 5$ we use a Monte Carlo method where we choose points either uniformly, or do importance sampling from a Gaussian with standard deviation $s_i/3$ in each direction.

We use 25 points per dimension for the deterministic integrals, which gives us an error between 10^{-6} and 10^{-7} . We test this using Matlab's built-in function integral, which has a default accuracy of about that amount. For the single integral with $d_X = 5$ we use 10^6 points, so the error is expected to be about 10^{-3} . Using the trapezoidal rule we obtain excellent

accuracy with very few points. For smooth periodic functions the trapezoidal rule achieves exponential convergence once enough discretization nodes are used so as to sample at the Nyquist rate. For nonperiodic functions, the error is dominated by the derivatives of the integrand at its endpoints, and achieves at least second-order convergence [62].

b. Symmetry numbers

The symmetry number accounts for the total number of distinct copies of each cluster obtained by permuting the labeling of the spheres. In our case we count two clusters as "distinct" if there is no rotation that maps the labeled spheres of one cluster to the spheres with the same labels in the other cluster.

The total number of copies of a rigid cluster of N spheres (including enantiomers) is

$$\frac{2N!}{a_k}$$
, (B2)

where a_k is the total number of permutations that map an adjacency matrix to itself and also preserve all the pairwise distances. The symmetry number for this method of counting is $\sigma = a_k$. We calculate a_k by first computing the automorphism group of the adjacency matrix using the function allgroup3 in the program nauty [63]. For each element in the automorphism group, we apply the corresponding permutation to the particles and check if the pairwise distance matrix is preserved. If so, we increase a_k by 1. We have to check the pairwise distances, because a permutation that preserves the adjacency matrix can yield a cluster that is distinct from the original. For example, consider a large octahedral shell, made of triangles glued together. Pick one triangle, and attach two spheres to it, one above and one below the triangle, to form a bipyramid. There is an automorphism of the adjacency matrix, namely switching the two spheres, that does not correspond to a rotation or a reflection.

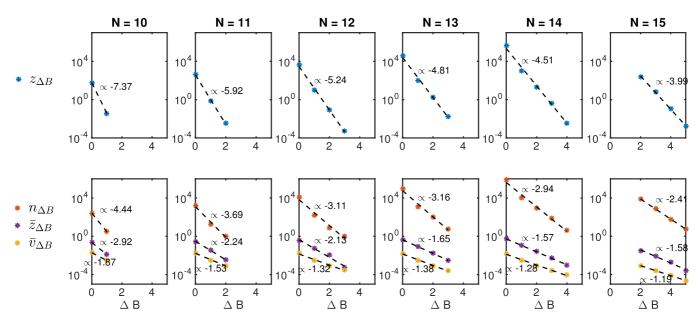


FIG. 11. Scaling laws for N = 10–14. Same labeling as described in the caption of Fig. 7.

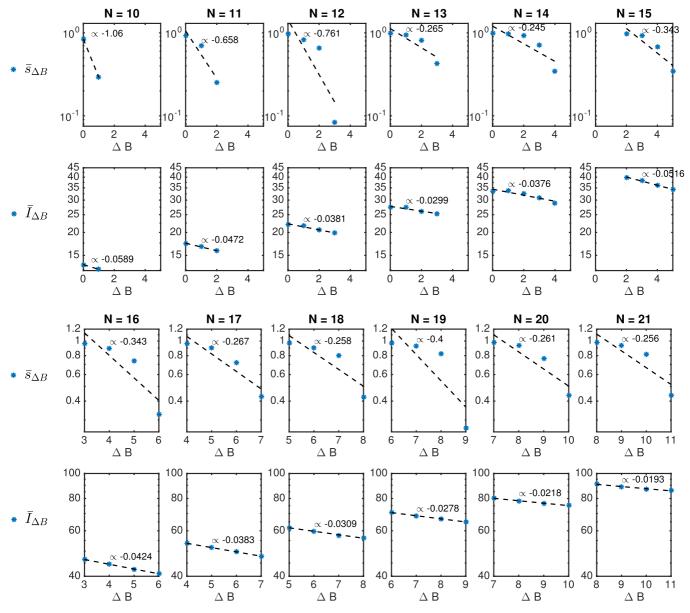


FIG. 12. Mean symmetry number $1/\sigma$ and mean moment of inertia factor $|I|^{1/2}$ for regular clusters as a function of ΔB , for each N = 10–21. The dashed line is the best-fit line with best-fit slope shown beside it. Vertical axis is a logarithmic scale. This shows that neither factor contributes strongly to the exponential scaling of $z_{\Delta B}$.

2. Simulation method

We performed metropolis Monte Carlo simulations for N=9 spheres interacting via a Morse-Harmonic potential with range parameter ρ of the form

$$V(r) = \begin{cases} \rho^2(r-1)^2 - 1 & r \leq 1\\ \exp[-\rho(r-1)][\exp(-\rho(r-1)) - 2] & r > 1 \end{cases}$$
(B3)

We replaced the Morse potential with a spring potential for r < 1 for numerical stability reasons. The entire matched potential is continuous with continuous first and second derivatives. We use a periodic simulation cell of size $6 \times 6 \times 6$.

Let $\mathbf{r}(t)$ be the coordinates sampled from the Monte Carlo trajectory at time t. We construct an adjacency matrix such that $a_{ij}(t) = 1$ if $r_{ij} < r_{\text{cutoff}}$ and $a_{ij} = 0$ otherwise. We set $b_k(t) = 1$ for the cluster $1 \le k \le 52$ (if any) whose adjacency matrix

 A_k is isomorphic to A(t) and $b_k(t)=0$ for all others. The frequency of cluster k is $P_k=\langle b_k\rangle/\sum_{k'=1}^{52}\langle b_{k'}\rangle$. We also measure the correlation function $C_k(\tau)=\langle b_k(t)b_k(t+\tau)\rangle/\langle b_k\rangle^2$. The estimated error for P_k is $\Delta P_k/P_k=(\langle b_k\rangle T/\tau_k)^{-1/2}$, where T is the length of the simulation and τ_k is the correlation time, which we estimate as $\tau_k=\sum_{\tau=0}^{\infty}C_k(\tau)$.

In the main text, we focus on the observed and predicted frequencies of the singular cluster. For completeness, we show in Figs. 9 and 10 some of the results for the nonsingular clusters. Figure 9 shows that the difficulty of identifying a cluster for larger values of the range parameter occurs just as much for regular clusters as for singular ones.

3. Scaling Laws

Figure 11 shows the same scaling laws as Fig. 7, but for smaller N.

Figure 12 shows the mean inverse symmetry number $\bar{s}_{\Delta B} = \frac{1}{n_{\Delta B}} \sum \sigma^{-1}$ for regular clusters with a given ΔB , and the mean moment of inertia factor $\bar{I}_{\Delta B} = \frac{1}{n_{\Delta B}} \sum I$. The mean inverse symmetry number does not follow an exponential scaling law. The mean moment of inertia does, but the slope is tiny. This demonstrates that neither the symmetry number, nor the moment of inertia, are important factors in determining why $\bar{z}_{\Delta B}$ follows an exponential scaling.

4. Bond-orientational order parameter

Given a cluster **r**, with bonds $E = \{(i_1, j_1), \dots, (i_B, j_B)\}$, we define the bond-orientational order parameter Q_l as follows:

$$q_{l,m} = \frac{1}{B} \sum_{(i,j)\in E} Y_{l,m}(\mathbf{r}_{ij}/d)$$

$$Q_l = \left(\frac{1}{2l+1} \sum_{m=-l}^{l} q_{l,m}^2\right)^{1/2},$$
(B4)

where $Y_{l,m}(\mathbf{u})$, $m=-l,\ldots,l$, is a basis for the spherical harmonics of degree l normalized such that $\int_{S^2} Y_{l,m}(\mathbf{u})^2 d\mathbf{u} = 4\pi$. The value of the bond-orientational order parameter Q_6 for a large fragment of the face-centered cubic lattice approaches the value 0.575 (the limit is 0.485 for the hexagonal close-packed lattice). On the other hand, if the bond directions are drawn randomly from a uniform measure on the sphere, the order parameter will tend to zero.

APPENDIX C: VALUES OF THE GEOMETRIC PARTITION FUNCTIONS

The following tables give the values of $z_{\Delta B, d_X}$ for each value of N. Each value is shown to three significant digits (if fewer are shown, it is because the trailing significant digits are zero).

(a) $N = 6$	$\frac{\overline{d_X \setminus \Delta B}}{0}$	0 0.0501					
(b) $N = 7$	$\frac{\overline{d_X \setminus \Delta B}}{0}$	0.222					
(c) $N = 8$	$\frac{\overline{d_X \setminus \Delta B}}{0}$	0 1.22					
(d) $N = 9$	$d_X \setminus \Delta B$ 0 1	7.88 0.0335					
(e) $N = 10$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ 0 \\ 1 \\ 2 \end{array} $	-1 - - 0.27	0 57.1 0.427 –	0.0362			
(f) $N = 11$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ 0 \\ 1 \\ 2 \\ 3 \end{array} $	-2 - - - 0.771	-1 - - 5.37 -	0 456 4.05 - -	1 0.701 0.0111 - -	2 0.00328 - - -	
(g) $N = 12$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \\ 2 \\ 3 \end{array} $	-2 - - - 9.5	-1 - - 59.5 -	0 4.06e + 0.35.8 0.158	1 3 9.24 0.225	2 0.0898 - - -	3 0.000501 - - -

(h) $N = 13$	$ \frac{\overline{d_X \setminus \Delta B}}{0} $ 1 2 3 4	-2 - - - 90.9 0.221	-1 - - 604 0.5 -	0 4.06e + 04 345 2.07	1 4 101 2.7 - - -	2 1.65 0.00517 - - -	3 0.0176 0.000482 - - -		
(i) $N = 14$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \\ 2 \\ 3 \\ 4 \\ 5 \end{array} $	- - -	- - - 6.41 018 1		0 4.56e + 05 3.71e + 03 21.7 - - -	1 1.04e + 03 29.8 - - - -	2 22 0.0791 - - - -	3 0.417 0.00488 - - - -	4 0.0037 - - - - - -
(j) $N = 15$	$ \frac{\overline{d_X \setminus \Delta B}}{0} $ 1	2 250 1.17	3 6.64 0.0471	4 0.112 –	5 0.00163 -				
(k) $N = 16$	$ \frac{\overline{d_X \setminus \Delta B}}{0} $ 1	3 87.8 0.611	4 2.14 –	5 0.0458 -	6 0.000727 –				
(1) $N = 17$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \end{array} $	33 0.0304	5 0.898 -	6 0.0223 -	7 0.00025	= 7 =			
(m) $N = 18$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \end{array} $	5 14.7 0.012	6 0.462 –	7 0.00989 –	8 8.38e-0 -	= 5 =			
(n) $N = 19$	$d_X \setminus \Delta B$ 0 1	6 7.8 0.0107	7 0.24 –	8 0.00408 -	9 3.39e-0	= 5 =			
(o) $N = 20$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \end{array} $	7 4.44 0.0077	8 0.121 –	9 0.00203 -	10 3 1.23e-				
(p) $N = 21$	$ \begin{array}{c c} \hline d_X \setminus \Delta B \\ \hline 0 \\ 1 \end{array} $	8 2.5 0.00421	9 0.696 –	10 0.0010 -	11 06 5.08e -				

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