Elimination of numerical Cherenkov instability in flowing-plasma particle-in-cell simulations by using Galilean coordinates

Remi Lehe,^{1,*} Manuel Kirchen,² Brendan B. Godfrey,^{1,3} Andreas R. Maier,² and Jean-Luc Vay¹

¹Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

²Center for Free-Electron Laser Science & Department of Physics, University of Hamburg, 22761 Hamburg, Germany

³University of Maryland, College Park, Maryland 20742, USA

(Received 14 July 2016; published 14 November 2016)

Particle-in-cell (PIC) simulations of relativistic flowing plasmas are of key interest to several fields of physics (including, e.g., laser-wakefield acceleration, when viewed in a Lorentz-boosted frame) but remain sometimes infeasible due to the well-known numerical Cherenkov instability (NCI). In this article, we show that, for a plasma drifting at a uniform relativistic velocity, the NCI can be eliminated by simply integrating the PIC equations in *Galilean coordinates* that follow the plasma (also sometimes known as *comoving coordinates*) within a spectral analytical framework. The elimination of the NCI is verified empirically and confirmed by a theoretical analysis of the instability. Moreover, it is shown that this method is applicable both to Cartesian geometry and to cylindrical geometry with azimuthal Fourier decomposition.

DOI: 10.1103/PhysRevE.94.053305

I. INTRODUCTION

Simulating relativistic flowing plasmas is of importance in several fields of physics, including relativistic astrophysics (e.g., Refs. [1–5]) and laser-plasma acceleration [6]. More precisely, although in laser-plasma acceleration the plasma is typically at rest in the laboratory frame, it was shown [7] that simulating the interaction in a Lorentz-boosted frame—where the plasma is flowing with relativistic speed—reduces computational demands by orders of magnitude.

However, despite the interest surrounding simulations of relativistic flowing plasmas, performing these simulations with particle-in-cell (PIC) algorithms [8,9] remains a challenge. This is because a violent numerical instability, known as the numerical Cherenkov instability (NCI) [10–17], quickly develops for relativistic plasmas and disrupts the simulation.

Several solutions have been proposed to mitigate the NCI [18–25]. Although these solutions efficiently reduce the numerical instability, they typically introduce either strong smoothing of the currents and fields or arbitrary numerical corrections, which are tuned specifically against the NCI and go beyond the natural discretization of the underlying physical equation. Therefore, it is sometimes unclear to what extent these added corrections could impact the physics at stake.

For instance, NCI-specific corrections include periodically smoothing the electromagnetic field components [10], using a special time step [11,12] or applying a wideband smoothing of the current components [11–13]. Another set of mitigation methods involves scaling the deposited currents by a carefully designed wave-number-dependent factor [18,19] or slightly modifying the ratio of electric and magnetic fields (E/B) before gathering their value onto the macroparticles [20–22]. Yet another set of NCI-specific corrections [23–25] consists in combining a small time step Δt , a sharp low-pass spatial filter, and a spectral or high-order scheme that is tuned to create a small, artificial "bump" in the dispersion relation [23]. While most mitigation methods have only been applied to Cartesian

By contrast, in Ref. [28], we propose that the NCI can be eliminated—with no arbitrary correction—by simply integrating the PIC equations in Galilean coordinates (also known as comoving coordinates). More precisely, in our method, the Maxwell equations in Galilean coordinates are integrated analytically, using only natural hypotheses, within the pseudospectral analytical time domain (PSATD) framework [29,30]). In the present article, we present the mathematical derivation and implementation of this Galilean PSATD scheme. Moreover, we conduct a detailed empirical and theoretical stability analysis for a *uniform* flowing plasma. On the other hand, the practical application of this algorithm to realistic, nonuniform plasmas (such as, e.g., in laser-wakefield acceleration) is presented in Ref. [28]. Overall, our method intrinsically supresses the NCI, does not require a small time step, and applies to both Cartesian and quasicylindrical geometry.

The outline of the present article is the following. We give an intuitive explanation of the Galilean scheme in Sec. I and then detail its exact implementation for Cartesian geometry in Sec. II. We also show *empirically*, in Sec. II, that, for a plasma drifting at a uniform relativistic velocity, the Galilean scheme suppresses the NCI. This fact is then confirmed and explained by a theoretical stability analysis in Sec. III (again, for Cartesian geometry). Finally, Sec. IV shows that the Galilean scheme can also be applied to quasicylindrical geometry and that it is then equally effective at suppressing the NCI.

II. AN INTUITIVE EXPLANATION OF THE GALILEAN SCHEME

The idea of the proposed scheme is to perform a Galilean change of coordinates and to carry out the simulation in the

geometry, this last set of methods ([23–25]) has the remarkable property that it can be applied [24] to both Cartesian geometry and quasicylindrical geometry (i.e., cylindrical geometry with azimuthal Fourier decomposition [26,27]). However, the use of a small time step proportionally slows down the progress of the simulation, and the artificial "bump" is again an arbitrary correction that departs from the underlying physics.

^{*}rlehe@lbl.gov

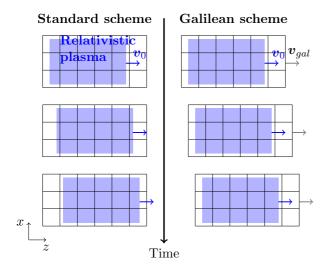


FIG. 1. Schematic representation of the Galilean scheme. As explained in the text, the Galilean scheme is *not* equivalent to a moving window.

new coordinates:

$$x' = x - v_{\text{gal}}t, \tag{1}$$

where $\mathbf{x} = x \mathbf{u}_x + y \mathbf{u}_y + z \mathbf{u}_z$ and $\mathbf{x}' = x' \mathbf{u}_x + y' \mathbf{u}_y + z' \mathbf{u}_z$ are the position vectors in the standard and Galilean coordinates, respectively.

We typically choose $\mathbf{v}_{gal} = \mathbf{v}_0$, where \mathbf{v}_0 is the speed of the bulk of the relativistic plasma. In this case, in the Galilean coordinates \mathbf{x}' , the plasma does not move with respect to the grid or, equivalently, in the standard coordinates \mathbf{x} , the grid moves along with the plasma (as represented in Fig. 1). The heuristic intuition behind this scheme is that these coordinates should prevent the discrepancy between the Lagrangian and Eulerian points of view, which gives rise to the NCI [15].

While in the standard coordinates x, the equations of particle motion and the Maxwell equations have the familiar form

$$\frac{d\mathbf{x}}{dt} = \frac{\mathbf{p}}{\gamma m},\tag{2a}$$

$$\frac{d\mathbf{p}}{dt} = q\left(\mathbf{E} + \frac{\mathbf{p}}{vm} \times \mathbf{B}\right),\tag{2b}$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E},\tag{2c}$$

$$\frac{1}{c^2} \frac{\partial \mathbf{E}}{\partial t} = \nabla \times \mathbf{B} - \mu_0 \mathbf{j},\tag{2d}$$

in the Galilean coordinates x', these equations become

$$\frac{d\mathbf{x}'}{dt} = \frac{\mathbf{p}}{vm} - \mathbf{v}_{\text{gal}},\tag{3a}$$

$$\frac{d\mathbf{p}}{dt} = q\left(\mathbf{E} + \frac{\mathbf{p}}{\gamma m} \times \mathbf{B}\right), \quad (3b)$$

$$\left(\frac{\partial}{\partial t} - v_{\text{gal}} \cdot \nabla'\right) \mathbf{B} = -\nabla' \times \mathbf{E},\tag{3c}$$

$$\frac{1}{c^2} \left(\frac{\partial}{\partial t} - \boldsymbol{v}_{\text{gal}} \cdot \boldsymbol{\nabla}' \right) \boldsymbol{E} = \boldsymbol{\nabla}' \times \boldsymbol{B} - \mu_0 \boldsymbol{j}, \tag{3d}$$

where ∇' denotes a spatial derivative with respect to the Galilean coordinates x'. The idea of the Galilean scheme is to design a PIC code which integrates the equations Eqs. (3a) to (3d) instead of Eqs. (2a) to (2d). Of course, physically, these two sets of equations are equivalent, as they are simply connected by a change of variables. However, we show in this paper that, numerically, these sets of equations have different stability properties when integrated with the PSATD scheme. Indeed, as shown in the next section, one of the unique feature of the PSATD scheme is that it takes into account the assumed time evolution of the current j within one time step. This allows us to push the idea of the coordinate change Eq. (1) further by embedding it into the assumed time evolution of j [see Eqs. (6) and (7) in the next section]. As we will show in Sec. III, this turns out to be key for the elimination of the NCI.

Before going further, let us remark that the Galilean change of coordinates Eq. (1) is a simple translation. Thus, when used in the context of Lorentz-boosted simulations [7], it does of course preserve the relativistic dilatation of space and time, which gives rise to the characteristic computational speedup of the boosted-frame technique.

Another important remark is that the Galilean scheme is *not* equivalent to a moving window (and, in fact, the Galilean scheme can be independently *combined* with a moving window). Whereas, in a moving window, grid points are added and removed to effectively translate the boundaries, in the Galilean scheme the grid points *themselves* are translated (and, again, in this case the physical equations are modified accordingly). In addition, the assumed time evolution of j within one time step [see Eqs. (6) and (7) in the next section] differs in a standard PSATD scheme with moving window and in a Galilean PSATD scheme.

III. THE GALILEAN PSATD SCHEME IN CARTESIAN GEOMETRY

While the previous section gave an intuitive description of the Galilean scheme, in the present section we introduce the exact numerical scheme that corresponds to this intuitive description—in the case of Cartesian geometry.

We start by deriving the update equations for the fields (Sec. IIA). The resulting PIC loop is then briefly described in Sec. IIB. Finally, in Sec. IIC, we show empirically that this PIC scheme has better stability than the standard PSATD in the case of a relativistic plasma.

A. Derivation of the discretized Maxwell equations in the Galilean PSATD scheme

In the PSATD scheme, the Maxwell equations are advanced by tranforming the fields \boldsymbol{E} and \boldsymbol{B} into Fourier space and then by integrating the Maxwell equations *analytically* over one time step.

In the case of the Galilean PSATD scheme, in order to analytically integrate the Maxwell equations in *Galilean coordinates* Eqs. (3c) and (3d), we first decouple the equations for *E* and *B* by combining Eqs. (3c) and (3d) into second-order

differential equations:

$$\left(\frac{\partial}{\partial t} - \boldsymbol{v}_{\text{gal}} \cdot \boldsymbol{\nabla}'\right)^{2} \boldsymbol{B} - c^{2} \boldsymbol{\nabla}'^{2} \boldsymbol{B} = \frac{1}{\epsilon_{0}} \boldsymbol{\nabla}' \times \boldsymbol{j}, \tag{4a}$$

$$\left(\frac{\partial}{\partial t} - \boldsymbol{v}_{\text{gal}} \cdot \boldsymbol{\nabla}'\right)^{2} \boldsymbol{E} - c^{2} \boldsymbol{\nabla}'^{2} \boldsymbol{E} = -\frac{c^{2}}{\epsilon_{0}} \boldsymbol{\nabla}' \rho$$

$$-\frac{1}{\epsilon_{0}} \left(\frac{\partial}{\partial t} - \boldsymbol{v}_{\text{gal}} \cdot \boldsymbol{\nabla}'\right) \boldsymbol{j}. \tag{4b}$$

Note that we used the equations $\nabla' \cdot \mathbf{E} = \rho/\epsilon_0$, $\nabla' \cdot \mathbf{B} = 0$, and $\epsilon_0 \mu_0 c^2 = 1$ in order to obtain the above equations.

In Fourier space, these equations become:

$$\left(\frac{\partial}{\partial t} - i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\right)^{2} \hat{\boldsymbol{\mathcal{B}}} + c^{2}\boldsymbol{k}^{2} \hat{\boldsymbol{\mathcal{B}}} = \frac{1}{\epsilon_{0}} i\boldsymbol{k} \times \hat{\boldsymbol{\mathcal{J}}}, \tag{5a}$$

$$\left(\frac{\partial}{\partial t} - i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\right)^{2} \hat{\boldsymbol{\mathcal{E}}} + c^{2}\boldsymbol{k}^{2} \hat{\boldsymbol{\mathcal{E}}} = -\frac{c^{2}}{\epsilon_{0}} \hat{\rho} i\boldsymbol{k}$$

$$-\frac{1}{\epsilon_{0}} \left(\frac{\partial}{\partial t} - i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\right) \hat{\boldsymbol{\mathcal{J}}}, \tag{5b}$$

where the Fourier components are defined by $\hat{\mathcal{F}}(\mathbf{k},t) = \int d^3x' F(x',t) e^{-i\mathbf{k}\cdot\mathbf{x}'}$, and F is \mathbf{E} , \mathbf{B} , \mathbf{j} , or ρ .

Equations (5a) and (5b) are linear ordinary differential equations in t, and they can be integrated analytically over one time step [i.e., from $t = n\Delta t$ to $t = (n+1)\Delta t$], provided that the time evolutions of the source terms $\hat{\mathcal{J}}(k,t)$ and $\hat{\rho}(k,t)$ are known over this time step—or, equivalently, provided that the time evolutions of the corresponding real-space function j(x',t) and $\rho(x',t)$ are known.

However, in a PIC code, j(x',t) and $\rho(x',t)$ are obtained from the deposition of the macroparticles' charge and current onto the grid. For this reason, j(x',t) and $\rho(x',t)$ are only known at a few discrete times between $n\Delta t$ and $(n+1)\Delta t$. For instance, in a typical PSATD PIC cycle, j is only computed at time $(n+1/2)\Delta t$ and ρ at times $n\Delta t$ and $(n+1)\Delta t$.

Thus, in order to analytically integrate equations Eqs. (5a) and (5b), one needs to make explicit assumptions on the time evolution of j and ρ between these known times. In the standard PSATD scheme (i.e., when the Maxwell equations are integrated in the standard coordinates x), one typically assumes the current j to be constant over one time step:

$$j(x,t) = j(x,(n+1/2)\Delta t) \quad \forall t \in [n\Delta t,(n+1)\Delta t].$$
(6)

However, when integrating the Maxwell equations in the Galilean variables x', it is more natural to assume

$$\mathbf{j}(\mathbf{x}',t) = \mathbf{j}(\mathbf{x}',(n+1/2)\Delta t) \quad \forall t \in [n\Delta t,(n+1)\Delta t],$$
(7)

i.e., that the current is constant over one time step *in the Galilean coordinates*. Because of the definition of x' [see Eq. (1)], the assumptions Eq. (6) and Eq. (7) are not equivalent. In fact, assuming Eq. (7) instead of Eq. (6) is one of the *key*

differences between the Galilean PSATD scheme described here and the standard PSATD scheme.

Once we adopt the assumption Eq. (7), our numerical scheme is fully determined. Equation (7) indeed results in:

$$\hat{\mathcal{J}}(\boldsymbol{k},t) = \hat{\mathcal{J}}(\boldsymbol{k},(n+1/2)\Delta t) \quad \forall t \in [n\Delta t, (n+1)\Delta t].$$
(8)

This equation in turn allows us to infer the time evolution of $\hat{\rho}(\boldsymbol{k},t)$ between $n\Delta t$ and $(n+1)\Delta t$. Indeed, in the Galilean coordinates, the equation of continuity reads $(\partial_t - \boldsymbol{v}_{\rm gal} \cdot \nabla')\rho + \nabla' \cdot \boldsymbol{j} = 0$, which becomes in Fourier space $(\partial_t - i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal})\hat{\rho} + i\boldsymbol{k} \cdot \hat{\mathcal{J}} = 0$. The solution $\hat{\rho}(\boldsymbol{k},t)$ of this equation for a constant $\hat{\mathcal{J}}$ is necessarily of the form:

$$\hat{\rho}(\mathbf{k},t) = \hat{\rho}(\mathbf{k},(n+1)\Delta t) \frac{1 - e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}(t-n\Delta t)}}{1 - e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}\Delta t}} - \hat{\rho}(\mathbf{k},n\Delta t) \frac{e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}\Delta t} - e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}(t-n\Delta t)}}{1 - e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}\Delta t}},$$
(9)

where we explicitly ensured that this solution satisfies the known initial and final conditions $\hat{\rho}(k, n\Delta t)$ and $\hat{\rho}(k, (n+1)\Delta t)$, which, again, are typically obtained from charge deposition during the PIC cycle. As a side note, notice that a necessary and sufficient condition for Eq. (9) to be a solution of the continuity equation with Eq. (8) is that the following relation be satisfied:

$$-i(\mathbf{k}\cdot\mathbf{v}_{\text{gal}})\frac{\hat{\rho}^{n+1}-\hat{\rho}^{n}e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}\Delta t}}{1-e^{i\mathbf{k}\cdot\mathbf{v}_{\text{gal}}\Delta t}}+i\mathbf{k}\cdot\hat{\mathbf{\mathcal{J}}}^{n+1/2}=0, \quad (10)$$

where we introduced the short-hand notations $\hat{\rho}^n \equiv \hat{\rho}(\boldsymbol{k}, n\Delta t)$, $\hat{\rho}^{n+1} \equiv \hat{\rho}(\boldsymbol{k}, (n+1)\Delta t)$, and $\hat{\mathcal{J}}^{n+1/2} \equiv \hat{\mathcal{J}}(\boldsymbol{k}, (n+1/2)\Delta t)$. Thus Eq. (10) is the discrete equation that $\hat{\rho}^{n+1}$, $\hat{\rho}^n$, $\hat{\mathcal{J}}^{n+1/2}$ should satisfy in order to satisfy the continuity equation—and therefore to ensure charge conservation—in the Galilean coordinates. [Notice that in the limit $\boldsymbol{v}_{\rm gal} = 0$ this equation reduces to $(\hat{\rho}^{n+1} - \hat{\rho}^n)/\Delta t + i\boldsymbol{k}\cdot\hat{\mathcal{J}}^{n+1/2} = 0$.] As such, it is also the equation that should be enforced during a PIC cycle, either through an Esirkepov-type deposition scheme or through a current correction scheme.

Finally, the time evolution of $\hat{\rho}$ and $\hat{\mathcal{J}}$ [Eqs. (8) and (9)] is inserted into the right-hand side of the Maxwell equations Eqs. (5a) and (5b). Again, these equations are linear ordinary differential equations, now with explicit expressions in their right-hand side, and they can be integrated analytically. Integrating these equations from $t = n\Delta t$ to $t = (n+1)\Delta t$ results in the following update equations (see Appendix A for the details of the derivation):

$$\hat{\mathbf{\mathcal{B}}}^{n+1} = \theta^2 C \hat{\mathbf{\mathcal{B}}}^n - \frac{\theta^2 S}{ck} i \mathbf{k} \times \hat{\mathbf{\mathcal{E}}}^n + \frac{\theta \chi_1}{\epsilon_0 c^2 k^2} i \mathbf{k} \times \hat{\mathbf{\mathcal{J}}}^{n+1/2},$$
(11a)
$$\hat{\mathbf{\mathcal{E}}}^{n+1} = \theta^2 C \hat{\mathbf{\mathcal{E}}}^n + \frac{\theta^2 S}{k} ci \mathbf{k} \times \hat{\mathbf{\mathcal{B}}}^n + \frac{i \nu \theta \chi_1 - \theta^2 S}{\epsilon_0 ck} \hat{\mathbf{\mathcal{J}}}^{n+1/2}$$

(11b)

 $-\frac{1}{\epsilon_0 k^2} (\chi_2 \ \hat{\rho}^{n+1} - \theta^2 \chi_3 \ \hat{\rho}^n) i \mathbf{k},$

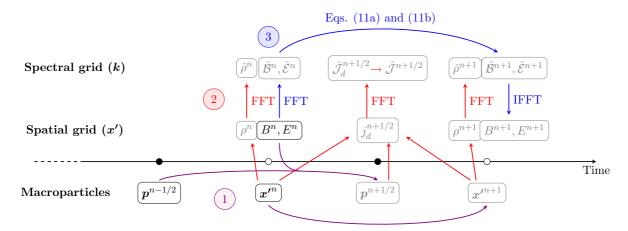


FIG. 2. Schematic representation of the PIC cycle. (FFT stands for fast Fourier transform, while IFFT stands for inverse fast Fourier transform.) The quantities that are known at the beginning of the PIC cycle are displayed in black, while the quantities that are computed during the PIC cycle are displayed in gray. The three successive steps of the PIC cycle—particle push (1), current deposition (2), and Maxwell solver (3)—are represented in purple, red, and blue, respectively.

where we used the short-hand notations $\hat{\mathcal{E}}^n \equiv \hat{\mathcal{E}}(k, n\Delta t), \hat{\mathcal{B}}^n \equiv \hat{\mathcal{B}}(k, n\Delta t)$, as well as

$$C = \cos(ck\Delta t)$$
 $S = \sin(ck\Delta t)$ $k = |\mathbf{k}|$, (12a)

$$v = \frac{\mathbf{k} \cdot \mathbf{v}_{\text{gal}}}{ck} \quad \theta = e^{i\mathbf{k} \cdot \mathbf{v}_{\text{gal}} \Delta t/2} \quad \theta^* = e^{-i\mathbf{k} \cdot \mathbf{v}_{\text{gal}} \Delta t/2}, \quad (12b)$$

$$\chi_1 = \frac{1}{1 - \nu^2} (\theta^* - C\theta + i\nu\theta S),\tag{12c}$$

$$\chi_2 = \frac{\chi_1 - \theta(1 - C)}{\theta^* - \theta} \quad \chi_3 = \frac{\chi_1 - \theta^*(1 - C)}{\theta^* - \theta}.$$
(12d)

Note that, in the limit $v_{gal} = 0$, Eqs. (11a) and (11b) reduce to the standard PSATD equations [29], as expected.

B. Overview of the PIC cycle for the Galilean PSATD scheme $\,$

Equations (10), (11a), and (11b) are the fundamental field equations of our PIC cycle. While Sec. II A emphasized the logical reasoning that leads to these equations, it did not give a precise description of their role within the PIC cycle. Therefore, the present section gives a concise overview of the different steps of the PIC cycle, for the Galilean PSATD scheme.

Apart from the fact that the simulation is performed in Galilean coordinates, our PIC cycle is very close to the standard PSATD scheme [29]. In particular, the fields E, B, and ρ and the macroparticles' positions x' are defined at integer times, whereas the field J and the macroparticles' momenta p are defined at half-integer times. All the fields are defined at the same points in space (i.e., the spatial grid is not staggered). The successive steps of the PIC cycle are represented in Fig. 2 and described below.

1. Particle push

The fields E and B are interpolated at time $t = n\Delta t$ from the spatial grid to the macroparticles' positions (naturally using the Galilean coordinates x' for the interpolation). The interpolated fields are then used to push the macroparticles' momenta from $t = (n - 1/2)\Delta t$ to $t = (n + 1/2)\Delta t$ using a

discretized version of the equation of motion Eq. (3b). Note that Eq. (3b) is familiar and can be discretized by using, e.g., the Boris pusher [31] or, as we chose in this paper, the Vay pusher [32]. Then the macroparticles' positions are pushed from $t = n\Delta t$ to $t = (n+1)\Delta t$ by using a trivial leap-frog discretization of Eq. (3a):

$$x'^{n+1} = x'^n + \Delta t \left(\frac{p^{n+1/2}}{\gamma^{n+1/2} m} - v_{\text{gal}} \right),$$
 (13)

where
$$\gamma^{n+1/2} = \sqrt{1 + (p^{n+1/2}/mc)^2}$$
.

2. Current and charge deposition

The charge density ρ is then computed on the spatial grid at $t = n\Delta t$ and $t = (n+1)\Delta t$ from the macroparticles' positions \mathbf{x}'^n and \mathbf{x}'^{n+1} , respectively. In addition, by using the intermediate positions $\mathbf{x}'^{n+1/2} \equiv (\mathbf{x}'^n + \mathbf{x}'^{n+1})/2$, the current \mathbf{j}_d is calculated on the spatial grid at $t = (n+1/2)\Delta t$. Here, the subscript d emphasizes the fact that we use a *direct* deposition scheme rather than a charge-conserving deposition scheme. As a consequence, the Fourier transform $\hat{\mathcal{J}}_d^{n+1/2}$ of the current $\mathbf{j}_d^{n+1/2}$ does not satisfy the discretized continuity equation Eq. (10) by default. For this reason, we use the Fourier transform of the charge density at $t = n\Delta t$ and $t = (n+1)\Delta t$ to compute a corrected current $\hat{\mathcal{J}}_{n+1/2}^{n+1/2}$ which does satisfy Eq. (10):

$$\hat{\mathcal{J}}^{n+1/2} = \hat{\mathcal{J}}_d^{n+1/2} + \frac{i\boldsymbol{k}}{k^2}\hat{\mathcal{G}}, \qquad (14a)$$

$$\hat{\mathcal{G}} = -i(\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}) \frac{\hat{\rho}^{n+1} - \hat{\rho}^n e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\Delta t}}{1 - e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\Delta t}} + i\boldsymbol{k} \cdot \hat{\mathcal{J}}_d^{n+1/2}.$$

(14b)

Finally, a light amount of spatial smoothing is applied to $\hat{\rho}^n$, $\hat{\rho}^{n+1}$, and $\hat{\mathcal{J}}^{n+1/2}$. More precisely, each of these fields is multiplied by a low-pass filter which is equivalent, in real-space, to a one-pass binomial smoother followed by a

TABLE I. Parameters of the test simulations. The simulations are scaled by $k_{p,r} \equiv k_p/\gamma_0^{1/2}$, with $k_p^2 = n_0 e^2/m_e \epsilon_0 c^2$.

Plasma density	n_0 (scales the simulation)
Lorentz factor	$\gamma_0 = 130$
Cell size along z	$\Delta z = 0.3868 \ k_{p,r}^{-1}$
Cell size along <i>x</i>	$\Delta x = 0.3868 \ k_{p,r}^{-1}$
Time step	$\Delta t = \Delta z/c$
Number of grid points	$N_x = N_z = 200$
Order of the shape factor	3 (both in x and z)

compensator [9]:

$$\hat{T}(\mathbf{k}) = [1 - \sin^2(k_x \Delta x/2)][1 + \sin^2(k_x \Delta x/2)]$$

$$\times [1 - \sin^2(k_y \Delta y/2)][1 + \sin^2(k_y \Delta y/2)]$$

$$\times [1 - \sin^2(k_z \Delta z/2)][1 + \sin^2(k_z \Delta z/2)], \quad (15)$$

where Δx , Δy , and Δz are the cell size of the spatial grid in each direction. [For two-dimensional (2D) simulations in the x-z plane, this expression is applied with $k_y = 0$.]

3. Maxwell solver

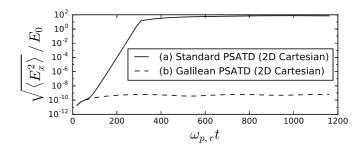
In order to update the values of the fields E and B from $t = n\Delta t$ to $t = (n+1)\Delta t$, we first transform them to Fourier space at $t = n\Delta t$. We then use the deposited fields $\hat{\rho}^n$, $\hat{\rho}^{n+1}$, and $J^{n+1/2}$ as well as Eqs. (11a) and (11b) to obtain the updated values $\hat{\mathcal{E}}^{n+1}$ and $\hat{\mathcal{B}}^{n+1}$ in spectral space. Finally, these fields are converted back to real space by using an inverse Fourier transform.

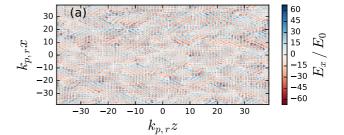
C. Stability of a uniform, relativistic plasma

We implemented the Galilean PSATD scheme described in Sec. IIB in the code WARP [33]. We then tested its stability for simulations of relativistic flowing plasmas in a 2D Cartesian geometry.

In the test simulations, a uniform plasma of density n_0 fills the entire periodic simulation box initially and flows towards the positive z with a relativistic speed v_0 . We ran this simulation both with the standard PSATD scheme and with the Galilean PSATD scheme, using $v_{\rm gal} = v_0$ in the latter case. (Thus in the standard PSATD simulation, the relativistic plasma cycles through the fixed periodic boundaries of the box, while in the Galilean PSATD simulation, the box moves along with the plasma.) The physical and numerical parameters of the simulation are summarized in Table I.

The results of this test are shown in Fig. 3. As shown in the top panel, in the case of the standard PSATD the fluctuations of the electric fields grow exponentially from an initial noise (which is generated by numerical round-off errors in the charge and current deposition of a neutral plasma) and saturate shortly after the beginning of the simulation. This a result of the well-known NCI. Conversely, with the Galilean PSATD the fluctuations of E_x remain at a low level and can be explained by a simple accumulation of numerical noise. This interpretation is confirmed by the maps of the electric field in the middle and bottom panels. While the standard PSATD simulation exhibits a high-wave-number pattern that is characteristic of the NCI,





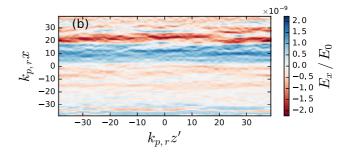


FIG. 3. Results of a relativistic flowing plasma simulation in WARP, with the standard PSATD scheme (a) and Galilean PSATD scheme (b). Top panel: RMS amplitude of the electric field in the simulation box versus time. (By definition, $\omega_{p,r} \equiv \omega_p/\gamma_0^{1/2}$ and $E_0 \equiv m_e c \omega_{p,r}/e$.) Middle and bottom panels: Maps of the electric field in the simulation box, at time $t = 329 \, \omega_{p,r}^{-1}$, for both the standard and Galilean PSATD schemes $(k_{p,r} \equiv k_p/\gamma_0^{1/2})$.

the Galilean PSATD simulation exhibits a random-looking pattern (with an amplitude that is lower by almost 10 orders of magnitude) that is consistent with numerical or thermal noise.

The Galilean PSATD scheme is thus empirically much more stable than the standard PSATD scheme. Again, a remarkable point is that we did not introduce any NCI-specific correction here. Instead, the Galilean scheme simply results from the natural analytical integration of the Maxwell equations in the Galilean coordinates, with no additional corrections.

IV. STABILITY ANALYSIS IN 2D CARTESIAN GEOMETRY

While the previous section showed *empirically* that the Galilean PSATD scheme is more stable, in the present section we confirm and explain these results by using the theoretical dispersion relation that corresponds to this numerical scheme.

A. Dispersion equation

More precisely, we start with a neutral, uniform plasma, flowing with a velocity $\mathbf{v}_0 = v_0 \mathbf{u}_z$ (and Lorentz factor γ_0) through a 2D periodic grid, and we consider the evolution of

a small perturbation to its fields of the form:

$$E, B \propto e^{ik \cdot x - i\omega t} = e^{ik \cdot x' - (\omega - k \cdot v_{\text{gal}})t}.$$
 (16)

Notice that, with the above definition, the physical interpretation of ω and k is the natural one, and in particular this interpretation does not depend on the choice of $v_{\rm gal}$.

By combining the perturbed Vlasov equation and the Maxwell equations, we obtain a dispersion equation that relates ω and k. Importantly, the analysis—and the resulting dispersion equation—incorporates all the numerical effects that are introduced by the PIC cycle from Sec. II B (including finite time step, finite spatial resolution, shape factors, current correction, etc.). Note, however, that the analysis has been restricted to the case where $v_{\rm gal}$ is along z (i.e., $v_{\rm gal} = v_{\rm gal} u_z$). The full derivation of the dispersion equation is given in Appendix B. Although this derivation builds on previous work [15–18], a number of important changes have been introduced in order to accommodate the specifics of the Galilean PSATD scheme.

The resulting dispersion relation is given in Eq. (19), along with the expression of the dimensionless coefficients

 ξ [Eqs. (20a) to (20c)], which represent the response of the plasma, and include the effects of spatial field smoothing $[\hat{T}(k)]$, spatial aliases $(K_m = m_x \frac{2\pi}{\Delta x} u_x + m_z \frac{2\pi}{\Delta z} u_z)$, and finite shape factor $[\hat{S}(k)]$. The factor $\hat{S}(k)$ indeed represents the Fourier transform of the macroparticle shape factor, so, e.g., for a shape factor of order ℓ_x and ℓ_z along x and z, respectively, one has:

$$\hat{S}(\mathbf{k}) = \operatorname{sinc}^{\ell_x + 1} \left(\frac{k_x \Delta x}{2} \right) \operatorname{sinc}^{\ell_z + 1} \left(\frac{k_z \Delta z}{2} \right) \tag{17}$$

with sinc(x) = sin(x)/x. In addition, in Eqs. (19) and (21) we also used the short-hand notations

$$s_x = \sin\left(\frac{x\Delta t}{2}\right)$$
 $c_x = \cos\left(\frac{x\Delta t}{2}\right)$ $t_x = \tan\left(\frac{x\Delta t}{2}\right)$ (18)

and we introduced χ_5 and χ_5' , which are coefficients that depend only on ω , k, and $\nu = \mathbf{k} \cdot \mathbf{v}_{\rm gal}/(ck)$, and whose mathematical expression results from the key hypothesis Eq. (7).

$$\left(s_{\omega}^{2} - t_{ck}^{2}c_{\omega}^{2}\right)\left[1 - \frac{1}{k}\left(k_{x}\xi_{3x} + \frac{k_{z}\xi_{3z}}{\gamma_{0}^{2}}\right)\right] - \xi_{1}\left[\frac{\chi_{5}}{k^{2}}\left(k_{z}^{2} + \frac{k_{x}^{2}}{\gamma_{0}^{2}}\right) + \chi_{5}'\frac{k_{z}v_{0}}{ck}\right] \\
+ \frac{k_{x}v_{0}}{ck}\left[\frac{\chi_{5}}{k}\left(k_{z}\xi_{2x} - \frac{k_{x}\xi_{2z}}{\gamma_{0}^{2}}\right) + \chi_{5}'\frac{\xi_{2x}v_{0}}{c}\right] + \frac{1}{\gamma_{0}^{2}}\left(\chi_{5} + \frac{k_{z}v_{0}}{ck}\chi_{5}'\right)\left[\xi_{1}\frac{k_{x}\xi_{3x} + k_{z}\xi_{3z}}{k} + \frac{k_{x}v_{0}}{ck}(\xi_{3x}\xi_{2z} - \xi_{3z}\xi_{2x})\right] = 0 \quad (19)$$

with

$$\xi_1 = \hat{T}(\mathbf{k}) \frac{\omega_p^2}{\gamma_0 c k} \sum_{m_x, m_z = -\infty}^{\infty} \frac{1}{\frac{2}{\Delta t} \sin\left[(\omega - k_z v_0) \frac{\Delta t}{2} + m_z \frac{\pi \Delta t}{\Delta z} (v_{\text{gal}} - v_0)\right]} \hat{S}^2(\mathbf{k} + \mathbf{K}_m), \tag{20a}$$

$$\boldsymbol{\xi}_{2} = \hat{T}(\boldsymbol{k}) \frac{\omega_{p}^{2}}{\gamma_{0}k} \sum_{m_{x},m_{z}=-\infty}^{\infty} \frac{\cos\left[(\omega - k_{z}v_{0})\frac{\Delta t}{2} + m_{z}\frac{\pi\Delta t}{\Delta z}(v_{\text{gal}} - v_{0})\right]}{\left(\frac{2}{\Delta t}\right)^{2} \sin^{2}\left[(\omega - k_{z}v_{0})\frac{\Delta t}{2} + m_{z}\frac{\pi\Delta t}{\Delta z}(v_{\text{gal}} - v_{0})\right]} \hat{S}^{2}(\boldsymbol{k} + \boldsymbol{K}_{m})(\boldsymbol{k} + \boldsymbol{K}_{m}),$$
(20b)

$$\boldsymbol{\xi}_{3} = \hat{T}(\boldsymbol{k}) \frac{\omega_{p}^{2}}{\gamma_{0} k} \sum_{m_{x}, m_{z} = -\infty}^{\infty} \frac{1}{\left(\frac{2}{\Delta t}\right)^{2} \sin^{2}\left[\left(\omega - k_{z} v_{0}\right) \frac{\Delta t}{2} + m_{z} \frac{\pi \Delta t}{\Delta z} \left(v_{\text{gal}} - v_{0}\right)\right]} \hat{S}^{2}(\boldsymbol{k} + \boldsymbol{K}_{m})(\boldsymbol{k} + \boldsymbol{K}_{m}), \tag{20c}$$

$$\chi_5 = \frac{c_{\omega}c_{vck}}{1 - v^2} [t_{\omega}(t_{ck} - vt_{vck}) - t_{ck}(t_{vck} - vt_{ck})] \quad \chi_5' = \frac{c_{\omega}c_{vck}}{1 - v^2} [t_{\omega}(t_{vck} - vt_{ck}) - t_{ck}(t_{ck} - vt_{vck})]. \tag{21}$$

Several remarks can be made on the dispersion equation Eq. (19). First, note that the set of equations (19), (20a)–(20c), and (21) is valid for any value of $v_{\rm gal}$, including $v_{\rm gal}=0$ (standard PSATD) and $v_{\rm gal}=v_0$ (optimal Galilean PSATD).

Another important point is that it can be verified (although only after some algebra) that Eq. (19) reduces, for any value of $v_{\rm gal}$, to

$$\frac{\Delta t^2}{4} \left(\omega^2 - c^2 k^2 - \frac{\omega_p^2}{\gamma_0} \right) \left(1 - \frac{\omega_p^2}{\gamma_0^3 (\omega - k_z v_0)^2} \right) = 0 \quad (22)$$

in the limit of infinitely small time step and cell size ($\omega \Delta t \ll 1$, $k\Delta x \ll 1$, $k\Delta z \ll 1$). Thus, as expected, in the limit of infinitely high resolution, the dispersion equation recovers the two independent physical modes of a relativistic plasma—

the relativistic plasma mode $\omega = k_z v_0 \pm \omega_p \gamma_0^{-3/2}$ and the relativistic electromagnetic mode $\omega^2 = c^2 k^2 + \omega_p^2/\gamma_0$.

Conversely, at finite resolution, Eq. (19) gives rise to distorted modes, which can potentially become unstable. This is particularly true near the numerical resonances of the plasma coefficients ξ_1, ξ_2, ξ_3 , i.e., whenever

$$\omega - k_z v_0 + m_z \frac{2\pi}{\Delta z} (v_{\text{gal}} - v_0) = 0 \quad \left(\text{modulo } \frac{2\pi}{\Delta t} \right), \quad (23)$$

so the sine term in the denominators of Eqs. (20a) to (20c) goes to 0. Since the resonance condition Eq. (23) depends on the alias number $m_z \in \mathbb{Z}$, this equation expresses the well-known fact that resonances occur at a set of different frequencies (aliased resonances) [16–18].

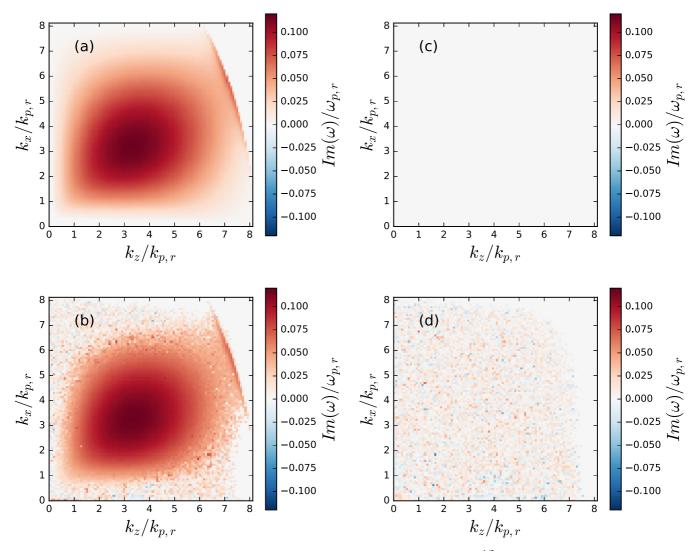


FIG. 4. Predicted and observed growth rates $\text{Im}(\omega)$ of the NCI, in k space (with $k_{p,r} = k_p/\gamma_0^{1/2}$), for the standard and Galilean PSATD schemes: (a) Standard PSATD, predicted; (b) Standard PSATD, observed in simulations; (c) Galilean PSATD, predicted; (d) Galilean PSATD, observed in simulations. The predicted growth rates (upper panels) are obtained from Eq. (19). All four panels use the parameters from Table I, and, in addition, the Galilean PSATD (right panels) uses $\boldsymbol{v}_{\text{gal}} = v_0 \boldsymbol{u}_z$.

In this regard, one consequence of the Galilean coordinates is clear: When choosing $v_{\rm gal}=v_0$, the term proportional to m_z in the resonance condition Eq. (23) vanishes, and thus all the aliased resonances are *relocated* to the same frequency: $\omega - k_z v_0 = 0$ (modulo $2\pi/\Delta t$). Interestingly, when tracking the corresponding terms throughout Appendix B, one realizes that this *relocation* of resonances is a direct consequence of the fact that the grid follows the plasma (as shown in Fig. 1) and that it does not depend on making the assumption Eq. (7) as opposed to Eq. (6).

Aside from this effect, the only other impact of $v_{\rm gal}$ on the dispersion equation Eq. (19) is in the expression of the coefficients χ_5 and χ_5' [through $v = \mathbf{k} \cdot \mathbf{v}_{\rm gal}/(ck)$], which, on the other hand, does result from the assumption of Eq. (7). This leads us to think that, in the case $v_{\rm gal} = v_0$, there are special relationships between χ_5 and χ_5' which effectively cancel the relocated resonance. For instance, it can be shown that, in the case $v_{\rm gal} = v_0$, the factor $[\chi_5 + (k_z v_0/ck)\chi_5']$ in the last term of Eq. (19) cancels at the resonance, whereas this is not true for $v_{\rm gal} = 0$.

Beyond these first remarks, it is difficult to analytically extract more insights from Eq. (19), and thus this equation needs to be solved *numerically* in order to actually predict the stability of a given situation.

B. Numerical solution and comparison with simulations

We solved the dispersion equation Eq. (19) numerically, with the physical and numerical parameters from Table I. In particular, when solving Eq. (19) for ω , we allowed, of course, for a nonzero imaginary part—since $\text{Im}(\omega)$ corresponds the growth rate of the instability.

These predicted growth rates were calculated for $v_{\rm gal}=0$ (standard PSATD) and $v_{\rm gal}=v_0$ (optimal Galilean PSATD), and they were compared with the corresponding WARP simulations from Sec. IIC. The results of these comparisons are shown in Fig. 4. Note that in the WARP simulations, the growth rate was estimated by taking the Fourier transform of the fields at $\omega_{p,r}t\simeq 120$ and $\omega_{p,r}t\simeq 170$ (i.e., within the linear growth

phase; see Fig. 3) and, for each Fourier mode, by calculating the difference in amplitude between those two times.

In the case of the standard PSATD (left panels in Fig. 4), one can see that the dispersion equation Eq. (19) correctly predicts that the simulation is unstable [existence of positive, nonzero $\text{Im}(\omega)$]. Moreover, the predicted growth rates from Eq. (19) are in excellent quantitative agreement with the growth rates observed in the WARP simulation. Notice also that, in the left panels of Fig. 4, the unstable modes cluster in two areas of k space: on a fine line at high k, which corresponds to the resonance $m_z = 0$ from Eq. (23), and on a broader area at lower k. This second, broader area corresponds to a nonresonant instability, which has also been predicted and observed in previous work [18,19,22].

In the case of the Galilean PSATD ($v_{\rm gal} = v_0$; right panels in Fig. 4), the dispersion equation Eq. (19) predicts that all modes are stable [Im(ω) = 0 across all k space]. This is again consistent with the observations from the WARP simulation, since the lower right panel in Fig. 4 displays only noise, with both positive and negative values of Im(ω). Again, in our understanding from the dispersion relation Eq. (19), this elimination of both the resonant and nonresonant NCI is due to mathematical expression of χ_5 and χ'_5 , which result from the assumption on the time evolution of j [Eq. (7)].

On the whole, this section confirms that the Galilean PSATD scheme eliminates the NCI, since the absence of NCI was both predicted theoretically and observed in simulations. Remarkably, the Galilean scheme simultaneously suppresses both the high-*k* resonant instability and low-*k* nonresonant instability. This contrasts with some of the previous mitigation techniques, which typically introduced two separate numerical corrections in order to handle the resonant and nonresonant instabilities, respectively.

C. Influence of $v_{\rm gal}$ on the growth rate

An interesting question is whether $v_{\rm gal} = v_0$ is indeed the optimal value of the Galilean scheme. To answer this question, we solved the dispersion equation Eq. (19) for a range of value of $v_{\rm gal}$, spanning from -1.5 c to +1.5 c, still with the parameters from Table I. (Note that, because we are performing a Galilean transformation—not a Lorentz transformation— v_{gal} is not necessarily restricted to the range -c to +c.) In addition, in order to evaluate the robustness of our scheme with respect to the time step Δt , we repeated this procedure for different values of Δt . The corresponding growth rates are plotted in Fig. 5, as a function of $v_{\rm gal}$. As shown in this figure, the growth of rate of the instability only goes to 0 for $v_{\rm gal} \simeq c$, thereby confirming that the optimal Galilean scheme has $v_{
m gal} \simeq v_0$ (in the case of an ultrarelativistic plasma). Remarkably, this behavior is observed for all the values of Δt that were tested, thereby indicating that the Galilean scheme eliminates the NCI independently of the value of Δt .

Another important feature of Fig. 5 is that the growth rate does not go to 0 for $v_{\rm gal} = -c$. In other words, the NCI is not suppressed when the relativistic plasma and Galilean grid move with *opposite* velocities. While this fact is to be expected from the intuitive picture of Sec. I, it can potentially have important implications for practical simulations. For instance, in Lorentz-boosted simulations of laser-wakefield acceleration

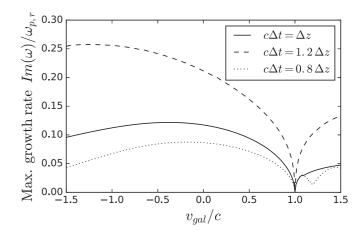


FIG. 5. Maximum growth rate of the NCI across all spectral modes (i.e., across all k space), as a function of the velocity $v_{\rm gal}$, and for different time steps Δt . The growth rates were calculated by solving Eq. (19) with the parameters from Table I.

(e.g., Refs. [34,35]), the optimal Galilean scheme would be that which follows the relativistically flowing background plasma (in typical conventions, this plasma flows to the left). However, in this case, the accelerated electron beam (which typically moves to the right) would counterpropagate with respect to the Galilean grid—thereby potentially triggering the NCI. Nevertheless, in this particular case, we see no evidence of the NCI in practical simulations (see Ref. [28]), including when analyzing the emittance of the accelerated beam. This absence of NCI is probably related to the lack of charge neutrality and limited spatial extent of the beam and will be investigated further in the future. In this regard, one important effect is the fact that the NCI modes often have a group velocity that is lower than c, and thus they rapidly slip behind the beam and stop growing.

V. THE GALILEAN PSATD SCHEME IN QUASI-CYLINDRICAL GEOMETRY

In Secs. II and III, we discussed the Galilean scheme in the context of a spectral *Cartesian* PIC code. Recently, two spectral *quasicylindrical* PIC codes were developed [36,37]. As shown in Ref. [26], simulations of physical systems with close-to-cylindrical symmetry can be made faster by orders of magnitude, when using a quasicylindrical grid instead of a 3D Cartesian grid. Therefore, in the present section, we extend the Galilean PSATD scheme to the spectral quasicylindrical framework of Ref. [36].

A. Numerical scheme in quasicylindrical geometry

It was shown in Ref. [36] that a PSATD algorithm could be derived in quasicylindrical geometry by expressing any scalar field S(x) as a sum of Fourier-Bessel modes:

$$S(\mathbf{x}) = \sum_{m=-\infty}^{\infty} \int_{-\infty}^{\infty} dk_z \int_{0}^{\infty} \frac{k_{\perp} dk_{\perp}}{(2\pi)^2} \times \hat{S}_m(k_z, k_{\perp}) J_m(k_{\perp} r) e^{ik_z z - im\theta}$$
(24)

TABLE II. Representation of common differential operators in a spectral Cartesian framework and spectral quasicylindrical framework [Eqs. (24) and (25a) to (25c)]. The expression of the spectral quasicylindrical representation can be derived by performing the same type of calculation as in Appendix B of Ref. [36].

Operator	Gradient: $F = \nabla S$	Curl: $F = \nabla \times V$	Divergence: $F = \nabla \cdot V$
Spectral Cartesian representation Spectral cylindrical representation	$\hat{\mathcal{F}} = i\mathbf{k}\hat{\mathcal{S}}$ i.e., $\begin{cases} \hat{\mathcal{F}}_x = ik_x\hat{\mathcal{S}} \\ \hat{\mathcal{F}}_y = ik_y\hat{\mathcal{S}} \\ \hat{\mathcal{F}}_z = ik_z\hat{\mathcal{S}} \end{cases}$ $\begin{cases} \hat{\mathcal{F}}_{+,m} = -k_\perp\hat{\mathcal{S}}_m/2 \\ \hat{\mathcal{F}}_{-,m} = k_\perp\hat{\mathcal{S}}_m/2 \\ \hat{\mathcal{F}}_{z,m} = ik_z\hat{\mathcal{S}}_m \end{cases}$	$\hat{\mathcal{F}} = i\mathbf{k} \times \hat{\mathbf{V}}$ i.e. $\begin{cases} \hat{\mathcal{F}}_x = ik_y \hat{\mathcal{V}}_z - ik_z \hat{\mathcal{V}}_y \\ \hat{\mathcal{F}}_y = ik_z \hat{\mathcal{V}}_x - ik_x \hat{\mathcal{V}}_z \\ \hat{\mathcal{F}}_z = ik_x \hat{\mathcal{V}}_y - ik_y \hat{\mathcal{V}}_x \end{cases}$ $\begin{cases} \hat{\mathcal{F}}_{+,m} = k_z \hat{\mathcal{V}}_{+,m} - ik_\perp \hat{\mathcal{V}}_{z,m}/2 \\ \hat{\mathcal{F}}_{-,m} = -k_z \hat{\mathcal{V}}_{-,m} - ik_\perp \hat{\mathcal{V}}_{z,m}/2 \\ \hat{\mathcal{F}}_{z,m} = ik_\perp \hat{\mathcal{V}}_{+,m} + ik_\perp \hat{\mathcal{V}}_{-,m} \end{cases}$	$\hat{\mathcal{F}} = i\mathbf{k} \cdot \hat{\mathbf{V}}$ i.e., $\hat{\mathcal{F}} = ik_x \hat{\mathcal{V}}_x + ik_y \hat{\mathcal{V}}_y + ik_z \hat{\mathcal{V}}_z$ $\hat{\mathcal{F}}_m = k_{\perp} (\hat{\mathcal{V}}_{+,m} - \hat{\mathcal{V}}_{-,m}) + ik_z \hat{\mathcal{V}}_{z,m}$

and, similarly, by expressing any vector field V(x) as

$$V_{r}(\mathbf{x}) = \sum_{m=-\infty}^{\infty} \int_{-\infty}^{\infty} dk_{z} \int_{0}^{\infty} \frac{k_{\perp} dk_{\perp}}{(2\pi)^{2}} [\hat{\mathcal{V}}_{+,m}(k_{z},k_{\perp}) J_{m+1}(k_{\perp}r) + \hat{\mathcal{V}}_{-,m}(k_{z},k_{\perp}) J_{m-1}(k_{\perp}r)] e^{ik_{z}z-im\theta}, \qquad (25a)$$

$$V_{\theta}(\mathbf{x}) = \sum_{m=-\infty}^{\infty} \int_{-\infty}^{\infty} dk_{z} \int_{0}^{\infty} \frac{k_{\perp} dk_{\perp}}{(2\pi)^{2}} i [\hat{\mathcal{V}}_{+,m}(k_{z},k_{\perp}) J_{m+1}(k_{\perp}r) - \hat{\mathcal{V}}_{-,m}(k_{z},k_{\perp}) J_{m-1}(k_{\perp}r)] e^{ik_{z}z-im\theta}, \qquad (25b)$$

$$V_{z}(\mathbf{x}) = \sum_{m=-\infty}^{\infty} \int_{-\infty}^{\infty} dk_{z} \int_{0}^{\infty} \frac{k_{\perp} dk_{\perp}}{(2\pi)^{2}} + \hat{\mathcal{V}}_{z,m}(k_{z},k_{\perp}) J_{m}(k_{\perp}r) e^{ik_{z}z-im\theta}, \qquad (25c)$$

where r is the radial coordinate, J_m is the Bessel function of order m, and the sum over m is a sum over azimuthal modes. (In a practical PIC simulation, this sum is truncated to a low number of modes, depending on the degree of cylindrical symmetry of the physical problem.) The terms $\hat{\mathcal{V}}_{+,m}$, $\hat{\mathcal{V}}_{-,m}$, $\hat{\mathcal{V}}_{z,m}$, and $\hat{\mathcal{S}}_m$ represent the spectral components of the fields V(x) and S(x).

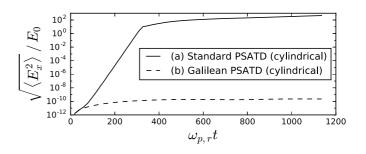
Within this formalism, the equations of the *standard quasicylindrical* PSATD are very similar to those of the *standard Cartesian* PSATD. In fact, although the *quasicylindrical* equations were derived from first principle in Ref. [36], they can alternatively be obtained by using a formal analogy (see Table II) between the representation of the differential operators in a spectral Cartesian and spectral quasicylindrical framework. More precisely, starting from the equations of the *standard Cartesian* PSATD, one can obtain the *standard quasicylindrical* PSATD scheme by simply replacing the expressions in the second line of Table II by those in the third line.

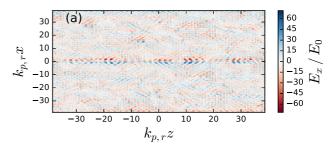
Therefore here, using the same heuristic procedure, we obtain the equations of the *Galilean quasicylindrical* PSATD (see Appendix C for their full expression) from the equations of the *Galilean Cartesian* PSATD [Eqs. (10), (11a), and (11b)], by simply replacing the expressions of the differential operators. Note that, in this context, both $v_{\rm gal}$ and the velocity of the relativistic plasma v_0 are necessarily along z.

Apart from these modified equations, the structure of our PIC cycle in quasicylindrical geometry is identical to that presented in Sec. II B for Cartesian geometry.

B. Stability of a uniform, relativistic plasma

The Galilean PSATD scheme described in the previous section was implemented in the spectral quasicylindrical code FBPIC [36]. We then performed test simulations featuring a uniform relativistic plasma. Apart from the shape factor (which was set to order 1), the numerical and physical parameters of the simulations are the same as in Table I (where Δx





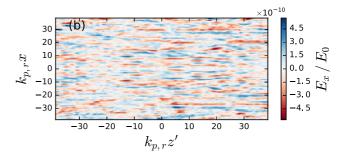


FIG. 6. Results of a relativistic flowing plasma simulation in the quasicylindrical PIC code FBPIC—with both the standard PSATD scheme (a) and Galilean PSATD scheme (b). Top panel: Evolution of the RMS amplitude of the electric field in the simulation box. Middle and bottom panels: Maps of the electric field in the simulation box, at time $t = 346 \, \omega_{p,r}^{-1}$.

and N_x are to be replaced with the corresponding radial parameters $\Delta r = 0.3868 \, k_{p,r}^{-1}$ and $N_r = 100$). In addition, the spatial smoothing function \hat{T} was set to $\hat{T}(k_z, k_\perp) = \cos^2(k_z \Delta z/2) \cos^2(k_\perp \Delta r/2)$ as in Ref. [36].

The results of these simulations are represented in Fig. 6, using a similar layout as for the corresponding Cartesian simulation (see Fig. 3). These quasicylindrical simulations support the same conclusions as their Cartesian counterpart: The standard PSATD scheme is unstable due to the NCI (as evidenced by the solid line in the upper panel of Fig. 6 and by the high-frequency pattern in the corresponding field map in the middle panel), while the Galilean PSATD scheme remains stable (see the dashed line in the upper panel and the corresponding field map on the bottom panel, which are consistent with numerical and thermal noise).

VI. CONCLUSION AND DISCUSSION

In this article, we showed that integrating the PIC equations in Galilean coordinates supresses the NCI, for a plasma drifting at a uniform relativistic velocity—in both Cartesian and quasicylindrical geometry. This new numerical scheme opens promising possibilities, especially for Lorentz-boosted simulations of laser-wakefield acceleration—as shown in Ref. [28].

Since the supression of the NCI is the aim of a number of previous schemes [18–25], it is worth discussing here the advantages and drawbacks of the Galilean PSATD scheme in relation to previous work, as well as areas of possible improvements.

As mentioned in the Introduction, one advantage of the Galilean scheme is that it is built on the natural integration of the Maxwell equations and does not introduce strong smoothing or arbitrary or manually tuned numerical corrections. This contrasts, for instance, with Refs. [18–21] but also with the methods from Refs. [23-25] in which both the time step and "bump" in k space need to be tuned in relation with the plasma density [23] (making it potentially difficult to simulate plasmas with longitudinally or transversally varying density profiles). On the other hand, while the methods from Refs. [18–21,23–25] can in some cases simulate relativistically crossing plasmas (such as those occurring in astrophysical shocks), in the present formulation of the Galilean scheme this could trigger the NCI for one of the two counterstreaming populations of macroparticles (see Sec. III C). In future works, we will present an alternate formulation of the Galilean scheme, using multiple grids for the current J, which relaxes the restrictions on simulations of crossing plasmas.

Another important point is that the usual advantages of the standard PSATD scheme naturally carry over to the Galilean PSATD scheme, including dispersion-free wave propagation (in all directions) and suppression of staggered interpolation artifacts (see, e.g., Ref. [36]). This is not the case for methods based on finite-difference time domain (FDTD), pseudospectral time domain (PSTD) or hybrid schemes (PSTD longitudinally and FDTD transversally), as in Refs. [20–25]. On the other hand, the methods based on FDTD or hybrid schemes can be more easily scaled to multiple computing nodes (see especially Ref. [25]). In order to mitigate this limitation, the Galilean scheme is using domain decomposition, as proposed in Ref. [30] in both WARP and FBPIC. To remove the limitation

further, this was extended in WARP (and will be extended in FBPIC in the future) to incorporate a spectral representation of *finite-difference* high-order operators (as discussed, e.g., in Refs. [25,38–40]), which have better scalability than purely spectral operators.

Finally, let us discuss the potential computational overhead of the Galilean PSATD scheme, when compared with alternative methods based on PSTD or FDTD [20-25]. While no quantitative comparisons of performance exist, we can still heuristically estimate the overhead based on the algorithmic operations involved. First, let us notice that the field advance of the Galilean PSATD scheme should have a comparable computational cost as that of the PSTD-based methods. This is because both the PSATD and PSTD schemes involve a set of FFTs (which usually dominate the computational cost) and a set of element-wise additions and multiplications in spectral space [which, for the Galilean PSATD are given by Eqs. (11a) and (11b)].]. It is also important to note that, although the Galilean PSATD seems to require the calculation of costly coefficients [e.g., C, S, χ_1 , χ_2 , and χ_3 in Eqs. (12a), (12c), and (12d)], these coefficients are in fact computed at the beginning of the simulation and reused at each iteration.

Although the Galilean PSATD scheme should have a comparable computational cost as the PSTD-based methods, it may have a different cost than an FDTD-based method such as that described in Ref. [25]. In principle, because the method of Ref. [25] only uses FFTs for the current correction, a single PIC iteration should be significantly faster than in the Galilean PSATD scheme (assuming that the computational time is dominated by the field advance and not by the particle operations). However, it should also be noted that the method of Ref. [25] (and of Refs. [23,24]) requires a time step which is substantially smaller than $\Delta z/c$, in order to avoid the NCI. This is not the case for the Galilean PSATD scheme, and thus the PSATD scheme will need fewer time steps for a given simulation. Whether one method is faster than the other will then depend on the trade-off between the overhead of the FFT and the required time step of the FDTD-based method.

ACKNOWLEDGMENTS

The simulation results were stored and visualized using the new open-source format openPMD [41]. The authors thank the openPMD contributors and in particular its creator, Axel Huebl (HZDR, Germany). The authors also thank Patrick Lee (U. Paris-Sud, France) for interesting discussions and for performing additional tests of the Galilean PSATD scheme (not presented here). This work was partly supported by the Director, Office of Science, Office of High Energy Physics, U.S. Dept. of Energy under Contract No. DE-AC02-05CH11231, including from the Laboratory Directed Research and Development (LDRD) funding from Berkeley Lab. The research used resources of the National Energy Research Scientific Computing Center, a DOE Office of Science User Facility supported by the Office of Science of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.

This document was prepared as an account of work sponsored in part by the United States Government. While this document is believed to contain correct information, neither the United States Government nor any agency thereof, nor The Regents of the University of California, nor any of their employees, nor the authors makes any warranty, express or implied, or assumes any legal responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service

by its trade name, trademark, manufacturer, or otherwise, does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or any agency thereof, or The Regents of the University of California. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or any agency thereof or The Regents of the University of California.

APPENDIX A: ANALYTICAL INTEGRATION OF THE MAXWELL EQUATIONS FROM $t = n\Delta t$ TO $t = (n + 1)\Delta t$

After inserting Eqs. (8) and (9) into Eqs. (5a) and (5b), we obtain the following equations:

$$\left(\frac{\partial}{\partial t} - i\boldsymbol{k} \cdot \boldsymbol{v}_{\text{gal}}\right)^2 \hat{\boldsymbol{\beta}} + c^2 \boldsymbol{k}^2 \hat{\boldsymbol{\beta}} = \frac{1}{\epsilon_0} i\boldsymbol{k} \times \hat{\boldsymbol{\mathcal{J}}}^{n+1/2},\tag{A1a}$$

$$\left(\frac{\partial}{\partial t} - i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}\right)^2 \hat{\boldsymbol{\mathcal{E}}} + c^2 \boldsymbol{k}^2 \hat{\boldsymbol{\mathcal{E}}} = \frac{1}{\epsilon_0} i(\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}) \hat{\boldsymbol{\mathcal{J}}}^{n+1/2} - \frac{c^2}{\epsilon_0} \hat{\rho}^{n+1} \frac{1 - e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}(t - n\Delta t)}}{1 - e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}\Delta t}} i\boldsymbol{k} + \frac{c^2}{\epsilon_0} \hat{\rho}^n \frac{e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}\Delta t} - e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}(t - n\Delta t)}}{1 - e^{i\boldsymbol{k} \cdot \boldsymbol{v}_{\rm gal}\Delta t}} i\boldsymbol{k}, \text{ (A1b)}$$

where $\hat{\rho}^{n+1}$, $\hat{\rho}^n$, and $\hat{\mathcal{J}}^{n+1/2}$ are constant. Notice in particular that the time derivative of $\hat{\mathcal{J}}$ from Eq. (5b) vanishes in Eq. (A1b) due to Eq. (8).

For the purpose of time integration, both equations can be cast into the following general form:

$$\left(\frac{\partial}{\partial t} - i\nu ck\right)^2 f + c^2 k^2 f = \alpha + \beta e^{i\nu ck(t - n\Delta t)},\tag{A2}$$

where $v \equiv \mathbf{k} \cdot \mathbf{v}_{\rm gal}/ck$ as in Eq. (12b) and where α and β are constants. For instance, in the case of Eq. (A1a), one has:

$$\alpha = \frac{1}{\epsilon_0} i \mathbf{k} \times \hat{\mathbf{J}}^{n+1/2} \quad \beta = 0 \quad f = \hat{\mathbf{B}}, \tag{A3}$$

while in the case of Eq. (A1b):

$$\alpha = \frac{i v c k}{\epsilon_0} \hat{\mathcal{J}}^{n+1/2} - \frac{c^2}{\epsilon_0} \frac{\hat{\rho}^{n+1} - \hat{\rho}^n e^{i v c k \Delta t}}{1 - e^{i v c k \Delta t}} i \mathbf{k} \quad \beta = \frac{c^2}{\epsilon_0} \frac{\hat{\rho}^{n+1} - \hat{\rho}^n}{1 - e^{i v c k \Delta t}} i \mathbf{k} \quad f = \hat{\mathbf{\mathcal{E}}}.$$
(A4)

The general solution of Eq. (A2) is

$$f(t) = \kappa_1 \cos[ck(t - n\Delta t)]e^{i\nu ck(t - n\Delta t)} + \kappa_2 \sin[ck(t - n\Delta t)]e^{i\nu ck(t - n\Delta t)} + \frac{\alpha}{c^2k^2(1 - \nu^2)} + \frac{\beta}{c^2k^2}e^{i\nu ck(t - n\Delta t)}, \quad (A5)$$

where κ_1 and κ_2 are integration constants. These integration constants can be determined from the initial condition $f(n\Delta t)$ and $\partial_t f(n\Delta t)$, in which case Eq. (A5) becomes:

$$f(t) = \left[f(n\Delta t) - \frac{\alpha}{c^2 k^2 (1 - v^2)} - \frac{\beta}{c^2 k^2} \right] \cos[ck(t - n\Delta t)] e^{ivck(t - n\Delta t)} + \frac{\alpha}{c^2 k^2 (1 - v^2)} + \frac{\beta}{c^2 k^2} e^{ivck(t - n\Delta t)}$$

$$+ \frac{1}{ck} \left[\partial_t f(n\Delta t) - ivck f(n\Delta t) + ivck \frac{\alpha}{c^2 k^2 (1 - v^2)} \right] \sin[ck(t - n\Delta t)] e^{ivck(t - n\Delta t)}.$$
(A6)

Finally, since our purpose is to integrate (A1a) and (A1b) from $t = n\Delta t$ to $t = (n+1)\Delta t$, let us evaluate the above equation at $t = (n+1)\Delta t$:

$$f((n+1)\Delta t) = C\theta^2 f(n\Delta t) + \frac{\theta \chi_1}{c^2 k^2} \alpha + \frac{\theta^2 (1-C)}{c^2 k^2} \beta + \frac{S\theta^2}{ck} [\partial_t f(n\Delta t) - i\nu ck f(n\Delta t)], \tag{A7}$$

where C, S, θ , and χ_1 have the same definition as in Eqs. (12a) to (12c). The integrated Maxwell equations [Eqs. (11a) and (11b)] are then obtained by combining (A7) with (A3) and (A4), respectively. In particular, in order to evaluate the last term in Eq. (A7), we used the equations:

$$\left(\frac{\partial}{\partial t} - i\nu ck\right)\hat{\mathcal{B}} = -i\mathbf{k} \times \hat{\mathcal{E}},\tag{A8a}$$

$$\left(\frac{\partial}{\partial t} - i\nu ck\right)\hat{\mathcal{E}} = c^2 i \mathbf{k} \times \hat{\mathbf{B}} - \frac{1}{\epsilon_0}\hat{\mathbf{J}},\tag{A8b}$$

which are the spectral representations of the Maxwell equations [Eqs. (3c) and (3d)].

APPENDIX B: DERIVATION OF THE DISPERSION RELATION FOR THE GALILEAN PSATD SCHEME

The dispersion relation typically results from combining the Vlasov equation and Maxwell equations. Here, we use a discretized version of the Vlasov equation and Maxwell equation that take into account all the numerical effects described in Sec. II B (interpolation to grid; current correction, etc.).

We consider a periodic box, and a uniform plasma having a density n_0 and a relativistic factor γ_0 . We will treat perturbations δf to the distribution function, as well as the fields E and B, as small quantities.

1. Notations and definitions

Let us consider a 2D Cartesian grid with $N_x \times N_z$ grid points and periodic boundaries. We will denote the position of the grid points x'_i , i.e.,

$$\mathbf{x}'_{j} = j_{x} \Delta x \, \mathbf{u}_{x} + j_{z} \Delta z \, \mathbf{u}_{z} \quad \begin{aligned} j_{x} &\in [0, N_{x} - 1] \\ j_{z} &\in [0, N_{z} - 1] \end{aligned}$$
 (B1)

where j_x and j_z are integers. In addition, we will denote X_{ℓ} the vectors of periodicity of the grid, i.e.,

$$X_{\ell} = \ell_x N_x \Delta x \, \boldsymbol{u}_x + \ell_z N_z \Delta z \, \boldsymbol{u}_z \quad \ell_x \in \mathbb{Z}, \ell_z \in \mathbb{Z}. \quad (B2)$$

With these notations, any vector of the reciprocal lattice can be written as $k_m = k + K_m$, where k is a vector of the first Brillouin zone, and K_m is a vector of periodicity of the reciprocal lattice, i.e., k and K_m are of the form:

$$\mathbf{K}_{m} = m_{x} \frac{2\pi}{\Delta x} \mathbf{u}_{x} + m_{z} \frac{2\pi}{\Delta z} \mathbf{u}_{z} \quad m_{x} \in \mathbb{Z}, m_{z} \in \mathbb{Z},$$
 (B3)

$$\mathbf{k} = j_x \frac{2\pi}{N_x \Delta x} \mathbf{u}_x + j_z \frac{2\pi}{N_z \Delta z} \mathbf{u}_z \quad \begin{aligned} j_x &\in [-N_x/2, N_x/2 - 1] \\ j_z &\in [-N_z/2, N_z/2 - 1] \end{aligned},$$
(B4)

where j_x and j_z are integers.

With these definitions, the expressions of the discrete Fourier transforms of the grid fields E and B (which are defined *exclusively* on the grid points x'_i) are

$$\hat{\mathcal{E}}(\mathbf{k}) = \frac{\sum_{j} e^{-i\mathbf{k}\cdot\mathbf{x}_{j}'} \mathbf{E}(\mathbf{x}_{j}')}{N_{x}N_{z}} \quad \hat{\mathcal{B}}(\mathbf{k}) = \frac{\sum_{j} e^{-i\mathbf{k}\cdot\mathbf{x}_{j}'} \mathbf{B}(\mathbf{x}_{j}')}{N_{x}N_{z}}.$$
(B5)

By contrast, the expression of the Fourier transform for the distribution function f(x', p) (which is defined also *in between* grid points and is periodic) is

$$\hat{f}(\mathbf{k} + \mathbf{K}_{m}, \mathbf{p}) = \frac{1}{N_{x} N_{z} \Delta x \Delta z} \int_{\text{box}} d\mathbf{x}' e^{-i(\mathbf{k} + \mathbf{K}_{m}) \cdot \mathbf{x}'} f(\mathbf{x}', \mathbf{p}),$$
(B6)

where the integration is performed over the (finite) extent of the box.

Finally, the Fourier transform of the particle shape factor S [which is defined over \mathbb{R}^2 and is not periodic—since, for instance, $S(\mathbf{x}) = (1 - |\mathbf{x}|/\Delta x)\Theta(\Delta x - |\mathbf{x}|) \times (1 - |\mathbf{z}|/\Delta z)\Theta(\Delta z - |\mathbf{z}|)$ for order 1 shape factor] is:

$$\hat{S}(\mathbf{k} + \mathbf{K}_{m}) = \frac{1}{\Delta x \Delta z} \int_{\mathbb{R}^{2}} d\mathbf{x}' \, e^{-i\mathbf{k} + \mathbf{K}_{m} \cdot \mathbf{x}'} S(\mathbf{x}'). \tag{B7}$$

These distinctions regarding the Fourier transform are important in order to correctly derive the space aliases.

2. Discretized Vlasov equation

Let us define $f^{n+1/2}(\mathbf{x}', \mathbf{p})$ as the distribution function of positions and momenta at half-integer step $(n + 1/2)\Delta t$, and let us derive the evolution of the f from one half-integer step to the next.

From the equations of motion of the particles Eqs. (3b) and (13), the evolution of position and momenta of one given particle from one half-integer time step to the next is

$$\frac{\boldsymbol{x}^{\prime n+1/2} - \boldsymbol{x}^{\prime n-1/2}}{\Delta t} = \frac{\boldsymbol{p}^{n+1/2}}{2m\gamma^{n+1/2}} + \frac{\boldsymbol{p}^{n-1/2}}{2m\gamma^{n-1/2}} - \boldsymbol{v}_{\text{gal}}, \quad \text{(B8a)}$$

$$\frac{\boldsymbol{p}^{n+1/2} - \boldsymbol{p}^{n-1/2}}{\Delta t} = q \boldsymbol{E}_{\text{interp}}^{n}$$

$$+ q \left(\frac{\boldsymbol{p}^{n+1/2}}{2m\gamma^{n+1/2}} + \frac{\boldsymbol{p}^{n-1/2}}{2m\gamma^{n-1/2}} \right) \cdot \boldsymbol{B}_{\text{interp}}^{n},$$
(B8b)

where by definition $x'^{n+1/2} \equiv (x'^{n+1} + x'^n)/2$ and where E_{interp}^n and B_{interp}^n are the interpolated fields at the particle's position, at time $n\Delta t$.

Since the fields E and B are treated as small quantities, the modifications of p over one time step is small, and thus these equations can be approximated to:

$$\frac{x^{m+1/2} - x^{m-1/2}}{\Delta t} = \frac{p^{n-1/2}}{m\gamma^{n-1/2}} - v_{\text{gal}},$$
 (B9a)

$$\frac{\boldsymbol{p}^{n+1/2} - \boldsymbol{p}^{n-1/2}}{\Delta t} = q \boldsymbol{E}_{\text{interp}}^n + q \frac{\boldsymbol{p}^{n-1/2}}{m \gamma^{n-1/2}} \times \boldsymbol{B}_{\text{interp}}^n. \text{ (B9b)}$$

Because the volume in phase space is conserved during this evolution, the corresponding evolution of the distribution function f is

$$f^{n+1/2} \left[\mathbf{x}' + \left(\frac{\mathbf{p}}{m\gamma} - \mathbf{v}_{\text{gal}} \right) \Delta t,$$

$$\mathbf{p} + q \Delta t \left(\mathbf{E}_{\text{interp}}^n + \frac{\mathbf{p}}{m\gamma} \times \mathbf{B}_{\text{interp}}^n \right) \right] = f^{n-1/2}(\mathbf{x}', \mathbf{p}).$$
(B10)

Now $f^{n+1/2}(\mathbf{x}', \mathbf{p})$ is of the form $f_0(\mathbf{p}) + \delta f^{n+1/2}(\mathbf{x}', \mathbf{p})$, where f_0 is the distribution function of a uniform, stationary plasma and δf is a perturbation. Since δf , \mathbf{E} , and \mathbf{B} are treated as small perturbations, the above equation can be Taylor expanded to first order:

$$\delta f^{n+1/2}(\mathbf{x}' + \mathbf{v}\Delta t - \mathbf{v}_{\text{gal}}\Delta t, \mathbf{p}) - \delta f^{n-1/2}(\mathbf{x}', \mathbf{p})$$

$$+ q\Delta t \left(\mathbf{E}_{\text{interp}}^{n} + \mathbf{v} \times \mathbf{B}_{\text{interp}}^{n}\right) \cdot \frac{\partial f_{0}}{\partial \mathbf{p}} = 0,$$
(B11)

where we used the short-hand notation $v \equiv p/(\gamma m)$.

Now since E_{interp}^n is interpolated to the macroparticles at time $n\Delta t$, its expression for a given macroparticle at position

 $x^{\prime n}$ is

$$E_{\text{interp}}^{n} = \sum_{j,\ell} S(x'^{n} - x'_{j} - X_{\ell}) E^{n}(x'_{j} + X_{\ell})$$

$$= \sum_{j,\ell} S(x'^{n} - x'_{j} - X_{\ell}) E^{n}(x'_{j}), \tag{B12}$$

where X_{ℓ} are vectors of periodicity of the grid and x'_{j} denote grid points [see (B1) and (B2)] and where the sum over j corresponds to a sum over the whole (finite) grid. S is the shape factor of the macroparticle. Finally $E^{n}(x'_{j})$ is the expression of E on the grid. Since B is defined on the same grid and at the same time as E (i.e., the grid is not staggered here), the equation for B^{n}_{interp} is similar to Eq. (B12). By combining Eq. (B12) with Eq. (B11) and the equation $x'^{n} = x'^{n-1/2} + [p^{n-1/2}/(\gamma^{n-1/2}m) - v_{\text{gal}}]\Delta t/2$, we obtain:

$$\delta f^{n+1/2}(\mathbf{x}' + \mathbf{v}\Delta t - \mathbf{v}_{gal}\Delta t, \mathbf{p}) - \delta f^{n-1/2}(\mathbf{x}', \mathbf{p})$$

$$+ q\Delta t \sum_{j,\ell} S \left[\mathbf{x}' + \frac{(\mathbf{v} - \mathbf{v}_{gal})\Delta t}{2} - \mathbf{x}'_j - \mathbf{X}_{\ell} \right]$$

$$\times \left[\mathbf{E}^n(\mathbf{x}'_j) + \mathbf{v} \times \mathbf{B}^n(\mathbf{x}'_j) \right] \cdot \frac{\partial f_0}{\partial \mathbf{p}} = 0.$$
(B13)

Let us evaluate the Fourier transform of the above equation, at a vector of the reciprocal lattice $k_m = k + K_m$ [see (B3) and (B4)]. By using the definition Eq. (B6), we have:

$$\delta \hat{f}^{n+1/2}(\boldsymbol{k_m}, \boldsymbol{p}) e^{i\boldsymbol{k_m} \cdot (\boldsymbol{v} - \boldsymbol{v_{\text{gal}}})\Delta t} - \delta \hat{f}^{n-1/2}(\boldsymbol{k_m}, \boldsymbol{p})$$

$$+ \frac{q \Delta t}{\Delta x \Delta z} \sum_{j} \left\{ \int_{\mathbb{R}^2} d\boldsymbol{x}' e^{-i\boldsymbol{k_m} \cdot \boldsymbol{x}'} S \left[\boldsymbol{x}' + \frac{(\boldsymbol{v} - \boldsymbol{v_{\text{gal}}})\Delta t}{2} - \boldsymbol{x}'_{j} \right] \right\}$$

$$\times \frac{1}{N_x N_z} [\boldsymbol{E}^{n}(\boldsymbol{x}'_{j}) + \boldsymbol{v} \times \boldsymbol{B}^{n}(\boldsymbol{x}'_{j})] \cdot \frac{\partial f_0}{\partial \boldsymbol{p}} = 0.$$
(B14)

And, finally, by using (B5) and (B7):

$$\delta \hat{f}^{n+1/2}(\boldsymbol{k}_{m},\boldsymbol{p}) e^{i\boldsymbol{k}_{m}\cdot(\boldsymbol{v}-\boldsymbol{v}_{\mathrm{gal}})\Delta t/2}$$

$$-\delta \hat{f}^{n-1/2}(\boldsymbol{k}_{m},\boldsymbol{p}) e^{-i\boldsymbol{k}_{m}\cdot(\boldsymbol{v}-\boldsymbol{v}_{\mathrm{gal}})\Delta t/2}$$

$$+q\Delta t \,\hat{\mathcal{S}}(\boldsymbol{k}_{m})[\hat{\boldsymbol{\mathcal{E}}}^{n}(\boldsymbol{k})+\boldsymbol{v}\times\hat{\boldsymbol{\mathcal{B}}}^{n}(\boldsymbol{k})] \cdot \frac{\partial f_{0}}{\partial \boldsymbol{p}} = 0. \quad (B15)$$

3. Discretized Maxwell equation

Let us now derive an expression of the discretized Maxwell equations, where the source terms are expressed as a function of δf . Let us first remark that, in Eqs. (11a) and (11b), the terms $\hat{\rho}^n$, $\hat{\rho}^{n+1}$, and $\hat{\mathcal{J}}^{n+1/2}$ are the charge and current obtained *after* current correction and smoothing, as described in Sec. IIB. After inserting the explicit expression for current correction and smoothing Eqs. (14a) and (15), the discretized Maxwell equations become

$$\hat{\mathcal{B}}^{n+1} = \theta^2 C \hat{\mathcal{B}}^n - \frac{\theta^2 S}{ck} i \mathbf{k} \times \hat{\mathcal{E}}^n + \frac{\theta \chi_1 \hat{\mathcal{T}}}{\epsilon_0 c^2 k^2} i \mathbf{k} \times \hat{\mathcal{J}}_d^{n+1/2},$$
(B16a)

$$\hat{\mathcal{E}}^{n+1} = \theta^2 C \hat{\mathcal{E}}^n + \frac{\theta^2 S}{k} cik \times \hat{\mathcal{B}}^n - \frac{\hat{T}ik}{\epsilon_0 k^2} (\hat{\rho}^{n+1} - \theta^2 C \hat{\rho}^n)$$

$$+\frac{i\nu\theta\chi_{1}-\theta^{2}S}{\epsilon_{0}ck}\hat{\mathcal{T}}\left[\hat{\mathcal{J}}_{d}^{n+1/2}-\frac{\left(\mathbf{k}\cdot\hat{\mathcal{J}}_{d}^{n+1/2}\right)\mathbf{k}}{k^{2}}\right],\tag{B16b}$$

where $\hat{\rho}^n$, $\hat{\rho}^{n+1}$, and $\hat{\mathcal{J}}_d^{n+1/2}$ are the charge and current obtained just after deposition and *before* current correction and smoothing, and where $\hat{\mathcal{T}}$ is the smoothing factor defined in Eq. (15) (with $k_v = 0$ in the 2D Cartesian case).

In addition, the above equations can be rewritten in a time-symmetrical form, which is more convenient for the analysis in the rest of this Appendix. It can indeed be verified that, if $\hat{\mathcal{E}}$ and $\hat{\mathcal{B}}$ satisfy (B16a) and (B16b), as well as the associated conservation equations $i\mathbf{k} \cdot \hat{\mathcal{B}}^n = 0$ and $i\mathbf{k} \cdot \hat{\mathcal{E}}^n = \hat{T} \hat{\rho}^n/\epsilon_0$, then they also satisfy

$$\theta^* c \hat{\mathbf{\mathcal{B}}}^{n+1} - \theta c \hat{\mathbf{\mathcal{B}}}^n = -t_{ck} \frac{i \mathbf{k} \times (\theta^* \hat{\mathbf{\mathcal{E}}}^{n+1} + \theta \hat{\mathbf{\mathcal{E}}}^n)}{k} + 2\chi_4' \frac{\hat{\mathcal{T}}}{\epsilon_0 c k} \frac{\mathbf{k} \times \hat{\mathcal{J}}_d^{n+1/2}}{k}, \quad (B17)$$

$$\theta^* \hat{\mathbf{\mathcal{E}}}^{n+1} - \theta \hat{\mathbf{\mathcal{E}}}^n = t_{ck} \frac{i \mathbf{k} \times (\theta^* c \hat{\mathbf{\mathcal{B}}}^{n+1} + \theta c \hat{\mathbf{\mathcal{B}}}^n)}{k} - \frac{\hat{\mathcal{T}} i \mathbf{k}}{\epsilon_0 k^2} (\theta^* \hat{\rho}^{n+1} - \theta \hat{\rho}^n)$$

$$-2\chi_4 \frac{\hat{\mathcal{T}}}{\epsilon_0 c k} \left[\hat{\mathcal{J}}_d^{n+1/2} - \frac{(\mathbf{k} \cdot \hat{\mathcal{J}}_d^{n+1/2}) \mathbf{k}}{k^2} \right], \quad (B18)$$

where * denotes the complex conjugate and where

$$\chi_4 = \frac{c_{vck}(t_{ck} - vt_{vck})}{(1 - v^2)} \quad \chi_4' = \frac{c_{vck}(t_{vck} - vt_{ck})}{(1 - v^2)}, \quad (B19)$$

$$t_{ck} \equiv \tan\left(\frac{ck\Delta t}{2}\right) \quad c_{ck} \equiv \cos\left(\frac{ck\Delta t}{2}\right),$$
 (B20)

$$t_{vck} \equiv \tan\left(\frac{vck\Delta t}{2}\right) \quad c_{vck} \equiv \cos\left(\frac{vck\Delta t}{2}\right).$$
 (B21)

Let us now express the deposited charge and currents as a function of δf . The current density, which is deposited at half-integer time, is given within one given cell by

$$J_d^{n+\frac{1}{2}}(\mathbf{x}_j') = \frac{1}{\Delta x \Delta z} \int d\mathbf{p} \int_{\mathbb{R}^2} d\mathbf{x}' \, q\mathbf{v} \, S(\mathbf{x}_j' - \mathbf{x}') \delta f^{n+1/2}(\mathbf{x}', \mathbf{p}).$$
(B22)

In the above expression, the integration is carried out over all space (\mathbb{R}^2) because the shape factor S may extend beyond the finite grid. Let us now expand the periodic function δf in Fourier series:

$$\mathbf{J}_{d}^{n+\frac{1}{2}}(\mathbf{x}_{j}') = \frac{1}{\Delta x \Delta z} \int d\mathbf{p} \int_{\mathbb{R}^{2}} d\mathbf{x}' \, q\mathbf{v} \, S(\mathbf{x}_{j}' - \mathbf{x}') \\
\times \sum_{k} \delta \hat{f}^{n+1/2}(\mathbf{k}_{m}, \mathbf{p}) e^{i\mathbf{k}_{m} \cdot \mathbf{x}'}, \tag{B23}$$

where the sum is over all vectors $k_m = k + K_m$ of the reciprocal lattice [see (B3) and (B4)]. With some algebra, this

can be rewritten as:

$$J_d^{n+\frac{1}{2}}(\mathbf{x}_j') = \sum_{\mathbf{k}} e^{i\mathbf{k}\cdot\mathbf{x}_j'} \sum_{\mathbf{m}} \hat{\mathcal{S}}(\mathbf{k}_m) \int d\mathbf{p} \, q\mathbf{v} \, \delta \, \hat{f}^{n+1/2}(\mathbf{k}_m, \mathbf{p}),$$
(B2)

where we used the relation $e^{iK_m \cdot x'_j} = 1$, which comes from (B1) and (B4). By identification, we have

$$\hat{\mathcal{J}}_d^{n+\frac{1}{2}}(k) = \sum_{m} \hat{\mathcal{S}}(k_m) \int d\, p \, q\, v \, \delta \hat{f}^{n+1/2}(k_m, p). \quad (B25)$$

Similarly, since the charge density ρ^n is deposited from the particle position $\mathbf{x'}^n = \mathbf{x'}^{n+1/2} - (\mathbf{v}^{n+1/2} - \mathbf{v}_{\rm gal})\Delta t/2$, its expression is

$$\rho^{n}(\mathbf{x}'_{j}) = \frac{1}{\Delta x \Delta z} \int d\mathbf{p} \int_{\mathbb{R}^{2}} d\mathbf{x}' \ q$$

$$\times S \left[\mathbf{x}'_{j} - \mathbf{x}' + \frac{(\mathbf{v} - \mathbf{v}_{\text{gal}}) \Delta t}{2} \right] \delta f^{n+1/2}(\mathbf{x}', \mathbf{p}),$$
(B26)

and thus the expression of $\hat{\rho}^n$ is

$$\hat{\rho}^{n}(\mathbf{k}) = \sum_{\mathbf{m}} \hat{\mathcal{S}}(\mathbf{k}_{\mathbf{m}}) \int d\mathbf{p} \ q \ \delta \hat{f}^{n+1/2}(\mathbf{k}_{\mathbf{m}}, \mathbf{p}) e^{\frac{i\mathbf{k}_{\mathbf{m}} \cdot (\mathbf{v} - \mathbf{v}_{\text{gal}})\Delta t}{2}}.$$
(B27)

Similarly, the expression of $\hat{\rho}^{n+1}$ is

$$\hat{\rho}^{n+1}(\mathbf{k}) = \sum_{\mathbf{m}} \hat{\mathcal{S}}(\mathbf{k}_{\mathbf{m}}) \int d\mathbf{p} \ q \ \delta \hat{f}^{n+1/2}(\mathbf{k}_{\mathbf{m}}, \mathbf{p}) e^{-\frac{i\mathbf{k}_{\mathbf{m}} \cdot (\mathbf{v} - \mathbf{v}_{\text{gal}})\Delta t}{2}}.$$
(B28)

4. Eigenmodes and eigensystem

The discretized Vlasov equation [Eq. (B15)], the discretized Maxwell equations (B17) and (B18), and the expression of the source terms (B25), (B27), and (B28) form a set of coupled equations of evolution. Let us look for eigenmodes of this set of equations, where we assume all perturbations to be of the form $e^{ik \cdot x - i\omega t} = e^{ik \cdot x' - i(\omega - k \cdot v_{\rm gal})t}$, so the definition of ω is independent of $v_{\rm gal}$ and corresponds to physical intuition. This results in the following expressions:

$$\hat{\mathcal{E}}^{n}(\mathbf{k}) = \hat{\mathcal{E}}(\mathbf{k})e^{-i(\omega - \mathbf{k} \cdot \mathbf{v}_{\text{gal}})n\Delta t},$$
 (B29a)

$$\hat{\mathcal{B}}^{n}(k) = \hat{\mathcal{B}}(k)e^{-i(\omega - k \cdot v_{\text{gal}})n\Delta t},$$
 (B29b)

$$\delta \hat{f}^{n+\frac{1}{2}}(\boldsymbol{k}_{m},\boldsymbol{p}) = \delta \hat{f}(\boldsymbol{k}_{m},\boldsymbol{p})e^{-i(\omega-\boldsymbol{k}\cdot\boldsymbol{v}_{\mathrm{gal}})(n+1/2)\Delta t}.$$
 (B29c)

Notice that we used k instead of $k_m \equiv k + K_m$ in the expression of the time evolution of $\delta \hat{f}^{n+\frac{1}{2}}(k_m, p)$. This is because, by definition of an eigenmode, all quantities (in this case $\hat{\mathcal{E}}$, $\hat{\mathcal{B}}$, and $\delta \hat{f}$) should have the same time evolution.

With these expressions, the discretized Vlasov equation Eq. (B15) yields:

$$\delta \hat{f}(\mathbf{k}_{m}, \mathbf{p}) = -i \frac{q \Delta t}{2} \hat{\mathcal{S}}(\mathbf{k}_{m}) \frac{[\hat{\mathcal{E}}(\mathbf{k}) + \mathbf{v} \times \hat{\mathcal{B}}(\mathbf{k})] \cdot \frac{\partial f_{0}}{\partial \mathbf{p}}}{\sin \left\{ \frac{[\omega - \mathbf{k} \cdot \mathbf{v} - \mathbf{K}_{m} \cdot (\mathbf{v} - \mathbf{v}_{\text{gal}})] \Delta t}{2} \right\}}.$$
(B30)

And, after some algebra, inserting the above expression into

(B25), (B27), and (B28) results in:

$$\hat{\mathcal{J}}_{d}^{n+\frac{1}{2}} = i \frac{\epsilon_{0} \omega_{p}^{2}}{\gamma_{0}} e^{-i(\omega - \mathbf{k} \cdot \mathbf{v}_{\text{gal}})(n+1/2)\Delta t}
\times \sum_{m} \hat{\mathcal{S}}^{2}(\mathbf{k}_{m}) \left(\frac{\hat{\mathcal{F}}}{\frac{2}{\Delta t} s_{\omega'}} + \frac{c_{\omega'}(\mathbf{k}_{m} \cdot \hat{\mathcal{F}}) \mathbf{v}_{0}}{\left[\frac{2}{\Delta t} s_{\omega'}\right]^{2}} \right)$$

$$\theta^{*} \hat{\rho}^{n+1} - \theta \hat{\rho}^{n} = \frac{2\epsilon_{0} \omega_{p}^{2}}{\gamma_{0}} s_{\omega} e^{-i(\omega - \mathbf{k} \cdot \mathbf{v}_{\text{gal}})(n+1/2)\Delta t}
\times \sum_{m} \hat{\mathcal{S}}^{2}(\mathbf{k}_{m}) \frac{(\hat{\mathcal{F}} \cdot \mathbf{k}_{m})}{\left[\frac{2}{\Delta t} s_{\omega'}\right]^{2}},$$
(B32)

where

$$\hat{\mathcal{F}} \equiv \hat{\mathcal{E}}(k) + v_0 \times \hat{\mathcal{B}}(k) - \frac{(v_0 \cdot \hat{\mathcal{E}}(k))v_0}{c^2},$$
 (B33)

$$\omega_p^2 = \frac{n_0 q^2}{m\epsilon_0}$$
 $s_\omega = \sin\left(\frac{\omega \Delta t}{2}\right)$, (B34)

$$s_{\omega'} = \sin \left\{ \frac{[\omega - \mathbf{k} \cdot \mathbf{v}_0 - \mathbf{K}_m \cdot (\mathbf{v}_0 - \mathbf{v}_{\text{gal}})] \Delta t}{2} \right\}, \quad (B35)$$

$$c_{\omega'} = \cos \left\{ \frac{[\omega - \mathbf{k} \cdot \mathbf{v}_0 - \mathbf{K}_m \cdot (\mathbf{v}_0 - \mathbf{v}_{\text{gal}})] \Delta t}{2} \right\}.$$
 (B36)

In the derivation of (B31) and (B32), we used integration by parts and the fact that the distribution function of the unperturbed background plasma is $f_0(\mathbf{p}) = n_0 \delta(\mathbf{p} - \gamma_0 m \mathbf{v}_0)$, and we also made use of the relation $\partial_{\mathbf{p}} \cdot (\mathbf{v} \times \mathbf{B}) = 0$ when $\mathbf{v} = \mathbf{p} (1 + (\mathbf{p}/mc)^2)^{-1/2}/m$.

Finally, inserting (B31), (B32), (B29a), and (B29b) into the time-symmetrical discrete Maxwell equations (B17) and (B18) results in the eigensystem:

$$s_{\omega}c\hat{\mathcal{B}} - t_{ck}c_{\omega}\frac{\mathbf{k}\times\hat{\mathcal{E}}}{k} = -\chi_{4}'\xi_{1}\frac{\mathbf{k}\times\hat{\mathcal{F}}}{k} - \chi_{4}'(\xi_{2}\cdot\hat{\mathcal{F}})\frac{\mathbf{k}\times\mathbf{v}_{0}}{ck},$$
(B37a)

$$s_{\omega}\hat{\mathcal{E}} + t_{ck}c_{\omega}\frac{\mathbf{k} \times c\hat{\mathcal{B}}}{k} = s_{\omega}(\mathbf{\xi}_{3} \cdot \hat{\mathcal{F}})\frac{\mathbf{k}}{k} + \chi_{4}\xi_{1}\frac{\mathbf{k} \times \hat{\mathcal{F}}}{k} \times \frac{\mathbf{k}}{k}$$
$$+\chi_{4}(\mathbf{\xi}_{2} \cdot \hat{\mathcal{F}})\frac{\mathbf{k} \times \mathbf{v}_{0}}{ck} \times \frac{\mathbf{k}}{k}, \quad (B37b)$$

where the ξ coefficients represent the response of the plasma

$$\xi_1 = \frac{\hat{T}\omega_p^2}{\gamma_0 ck} \left[\sum_{m} \frac{\hat{S}^2(k_m)}{\frac{2}{\Delta t} s_{\omega'}} \right],$$
 (B38a)

$$\boldsymbol{\xi}_{2} = \frac{\hat{T}\omega_{p}^{2}}{\gamma_{0}k} \left[\sum_{m} \frac{c_{\omega'}\hat{S}^{2}(\boldsymbol{k}_{m})}{\left[\frac{2}{\Delta T}s_{\omega'}\right]^{2}} \boldsymbol{k}_{m} \right], \quad (B38b)$$

$$\xi_3 = \frac{\hat{T}\omega_p^2}{\gamma_0 k} \left[\sum_{m} \frac{\hat{S}^2(k_m)}{\left[\frac{2}{\Delta_0} s_{o'}\right]^2} k_m \right].$$
 (B38c)

5. Dispersion relation

Finally, let us simplify the above eigensystem in the case where both \mathbf{v}_{gal} and the velocity of the unperturbed plasma \mathbf{v}_0 are along \mathbf{u}_z (i.e., $\mathbf{v}_{\text{gal}} = v_{\text{gal}}\mathbf{u}_z$ and $\mathbf{v}_0 = v_0\mathbf{u}_z$). In this case, projecting Eq. (B37a) along y and Eq. (B37b) along x and z

[as well as using the expression of $\hat{\mathcal{F}}$ from Eq. (B33)] results in the eigensystem:

$$M\begin{pmatrix} c\hat{\mathcal{B}}_{y} \\ \hat{\mathcal{E}}_{z} \\ \hat{\mathcal{E}}_{x} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix}, \tag{B39}$$

where the matrix M can be expressed as

$$M = M_{\text{vacuum}} + U_1(V_1^T + V_2^T) + s_{\omega} U_3 V_3^T,$$
 (B40)

where the superscript T denotes the transpose operation and where

$$M_{\text{vacuum}} = \begin{pmatrix} s_{\omega} & t_{ck}c_{\omega}k_{x}/k & -t_{ck}c_{\omega}k_{z}/k \\ t_{ck}c_{\omega}k_{x}/k & s_{\omega} & 0 \\ -t_{ck}c_{\omega}k_{z}/k & 0 & s_{\omega} \end{pmatrix},$$
(B41)

$$U_{1} = \begin{pmatrix} \chi_{4}' \\ \chi_{4}k_{x}/k \\ -\chi_{4}k_{z}/k \end{pmatrix} \quad V_{1} = \xi_{1} \begin{pmatrix} -k_{z}v_{0}/(ck) \\ -k_{x}/(k\gamma_{0}^{2}) \\ k_{z}/k \end{pmatrix}, \quad (B42)$$

$$U_{2} = \begin{pmatrix} \chi_{4} \\ \chi_{4}' k_{x} / k \\ -\chi_{4}' k_{z} / k \end{pmatrix} \quad V_{2} = \frac{k_{x} v_{0}}{ck} \begin{pmatrix} \xi_{2x} v_{0} / c \\ -\xi_{2z} / \gamma_{0}^{2} \\ -\xi_{2x} \end{pmatrix}, \quad (B43)$$

$$U_{3} = \begin{pmatrix} 0 \\ k_{z}/k \\ k_{x}/k \end{pmatrix} \quad V_{3} = \begin{pmatrix} \xi_{3x} v_{0}/c \\ -\xi_{3z}/\gamma_{0}^{2} \\ -\xi_{3x} \end{pmatrix}, \tag{B44}$$

where we also introduced an additional vector U_2 which is not used in Eq. (B40) but will be useful below.

The final dispersion relation is obtained by solving the equation det(M) = 0. However, calculating the analytical expression of the determinant det(M) using, e.g., Sarrus's rule can be a daunting task. Instead, the calculation of det(M) can be faciltated by expressing the matrix M in the basis (U_1, U_2, U_3) (an operation which does not change its determinant). In other words, one has det(M) = det(M'), where M' is the expression of the matrix M in the basis (U_1, U_2, U_3) :

$$M' = \begin{bmatrix} s_{\omega} + (V_1^T + V_2^T)U_1 & t_{ck}c_{\omega} & s_{\omega}V_3^T U_1 \\ t_{ck}c_{\omega} + (V_1^T + V_2^T)U_2 & s_{\omega} & s_{\omega}V_3^T U_2 \\ (V_1^T + V_2^T)U_3 & 0 & s_{\omega}(1 + V_3^T U_3) \end{bmatrix}.$$
(B45)

Using this property, the equation det(M) = 0 becomes

$$(s_{\omega}^{2} - t_{ck}^{2} c_{\omega}^{2}) (1 + V_{3}^{T} U_{3}) + (V_{1}^{T} + V_{2}^{T}) U$$

$$+ (V_{3}^{T} U_{3}) (V_{1}^{T} U) - (V_{1}^{T} U_{3}) (V_{3}^{T} U)$$

$$+ (V_{3}^{T} U_{3}) (V_{2}^{T} U) - (V_{2}^{T} U_{3}) (V_{3}^{T} U) = 0,$$
 (B46)

where the trivial solution $s_{\omega} = 0$ has be discarded and where $U = s_{\omega}U_1 - t_{ck}c_{\omega}U_2$. After some algebra, this equation reduces to Eq. (19).

APPENDIX C: EXPRESSION OF THE PSATD EQUATIONS IN QUASICYLINDRICAL GEOMETRY

As mentioned in the text, the PSATD equations in quasicylindrical geometry are obtained from the equations in Cartesian geometry [Eqs. (10), (11a), and (11b)], using the correspondence table (Table II). Thus, with this method, the discretized continuity equation Eq. (10) becomes

$$-ik_z v_{\text{gal}} \frac{\hat{\rho}_m^{n+1} - \hat{\rho}_m^n e^{ik_z v_{\text{gal}} \Delta t}}{1 - e^{ik_z v_{\text{gal}} \Delta t}} + k_\perp (\hat{\mathcal{J}}_{+,m}^{n+1/2} - \hat{\mathcal{J}}_{-,m}^{n+1/2}) + ik_z \hat{\mathcal{J}}_{z,m}^{n+1/2} = 0, \tag{C1}$$

and the corresponding current correction [Eq. (14a) in Cartesian geometry] becomes

$$\hat{\mathcal{J}}_{+,m}^{n+1/2} = \hat{\mathcal{J}}_{d+,m}^{n+1/2} - k_{\perp} \hat{\mathcal{G}}_{m}/(2k^{2}), \tag{C2a}$$

$$\hat{\mathcal{J}}_{-,m}^{n+1/2} = \hat{\mathcal{J}}_{d-,m}^{n+1/2} + k_{\perp} \hat{\mathcal{G}}_{m}/(2k^{2}), \tag{C2b}$$

$$\hat{\mathcal{J}}_{z,m}^{n+1/2} = \hat{\mathcal{J}}_{dz,m}^{n+1/2} + ik_z \hat{\mathcal{G}}_m / k^2,$$
 (C2c)

where $k^2 = k_\perp^2 + k_z^2$ by definition. In the above equations, $\hat{\mathcal{J}}_d$ and $\hat{\mathcal{J}}$ are the deposited current and corrected current, respectively, and the expression of $\hat{\mathcal{G}}_m$ is obtained by replacing $\hat{\mathcal{J}}$ by $\hat{\mathcal{J}}_d$ in the left-hand side of Eq. (C1).

Similarly, the update equations for the $\hat{\mathcal{B}}$ and $\hat{\mathcal{E}}$ field [Eqs. (11a) and (11b)] become

$$\hat{\mathcal{B}}_{+,m}^{n+1} = \theta^2 C \hat{\mathcal{B}}_{+,m}^n - \frac{\theta^2 S}{ck} \left(k_z \hat{\mathcal{E}}_{+,m}^n - \frac{i k_\perp}{2} \hat{\mathcal{E}}_{z,m}^n \right) + \frac{\theta \chi_1}{\epsilon_0 c^2 k^2} \left(k_z \hat{\mathcal{J}}_{+,m}^{n+1/2} - \frac{i k_\perp}{2} \hat{\mathcal{J}}_{z,m}^{n+1/2} \right), \tag{C3a}$$

$$\hat{\mathcal{B}}_{-,m}^{n+1} = \theta^2 C \hat{\mathcal{B}}_{-,m}^n - \frac{\theta^2 S}{ck} \left(-k_z \hat{\mathcal{E}}_{-,m}^n - \frac{i k_\perp}{2} \hat{\mathcal{E}}_{z,m}^n \right) + \frac{\theta \chi_1}{\epsilon_0 c^2 k^2} \left(-k_z \hat{\mathcal{J}}_{-,m}^{n+1/2} - \frac{i k_\perp}{2} \hat{\mathcal{J}}_{z,m}^{n+1/2} \right), \tag{C3b}$$

$$\hat{\mathcal{B}}_{z,m}^{n+1} = \theta^2 C \hat{\mathcal{B}}_{z,m}^n - \frac{\theta^2 S}{ck} \left(i k_\perp \hat{\mathcal{E}}_{+,m}^n + i k_\perp \hat{\mathcal{E}}_{-,m}^n \right) + \frac{\theta \chi_1}{\epsilon_0 c^2 k^2} \left(i k_\perp \hat{\mathcal{J}}_{+,m}^{n+1/2} + i k_\perp \hat{\mathcal{J}}_{-,m}^{n+1/2} \right), \tag{C3c}$$

$$\hat{\mathcal{E}}_{+,m}^{n+1} = \theta^2 C \hat{\mathcal{E}}_{+,m}^n + \frac{\theta^2 S c}{k} \left(k_z \hat{\mathcal{B}}_{+,m}^n - \frac{i k_\perp}{2} \hat{\mathcal{B}}_{z,m}^n \right) + \frac{i \nu \theta \chi_1 - \theta^2 S}{\epsilon_0 c k} \hat{\mathcal{J}}_{+,m}^{n+1/2} + \frac{1}{\epsilon_0 k^2} \left(\chi_2 \; \hat{\rho}_m^{n+1} - \theta^2 \chi_3 \; \hat{\rho}_m^n \; \right) \frac{k_\perp}{2}, \tag{C4a}$$

$$\hat{\mathcal{E}}_{-,m}^{n+1} = \theta^2 C \hat{\mathcal{E}}_{-,m}^n + \frac{\theta^2 S c}{k} \left(-k_z \hat{\mathcal{B}}_{-,m}^n - \frac{i k_\perp}{2} \hat{\mathcal{B}}_{z,m}^n \right) + \frac{i \nu \theta \chi_1 - \theta^2 S}{\epsilon_0 c k} \hat{\mathcal{J}}_{-,m}^{n+1/2} - \frac{1}{\epsilon_0 k^2} \left(\chi_2 \hat{\rho}_m^{n+1} - \theta^2 \chi_3 \hat{\rho}_m^n \right) \frac{k_\perp}{2}, \quad (C4b)$$

$$\hat{\mathcal{E}}_{z,m}^{n+1} = \theta^2 C \hat{\mathcal{E}}_{z,m}^n + \frac{\theta^2 Sc}{k} \left(i k_\perp \hat{\mathcal{B}}_{+,m}^n + i k_\perp \hat{\mathcal{B}}_{-,m}^n \right) + \frac{i v \theta \chi_1 - \theta^2 S}{\epsilon_0 c k} \, \hat{\mathcal{J}}_{z,m}^{n+1/2} - \frac{1}{\epsilon_0 k^2} \left(\, \chi_2 \, \, \hat{\rho}_m^{n+1} - \theta^2 \chi_3 \, \, \hat{\rho}_m^n \, \right) i k_z, \tag{C4c}$$

In the above equations, the coefficients θ , C, S, ν , χ_1 , χ_2 , and χ_3 have the same expression as in Eqs. (12a) to (12d) (bearing in my mind, in quasicylindrical geometry, that $\boldsymbol{v}_{\rm gal}$ is necessarily along \boldsymbol{u}_z and that the expression of k is $\sqrt{k_\perp^2 + k_z^2}$).

- [1] Y. Kazimura, J. I. Sakai, T. Neubert, and S. V. Bulanov, Astrophys. J. Lett. 498, L183 (1998).
- [2] L. O. Silva, R. A. Fonseca, J. W. Tonge, J. M. Dawson, W. B. Mori, and M. V. Medvedev, Astrophys. J. Lett. 596, L121 (2003).
- [3] A. Spitkovsky, Astrophys. J. Lett. **682**, L5 (2008).
- [4] U. Keshet, B. Katz, A. Spitkovsky, and E. Waxman, Astrophys. J. Lett. 693, L127 (2009).
- [5] S. F. Martins, R. A. Fonseca, L. O. Silva, and W. B. Mori, Astrophys. J. Lett. 695, L189 (2009).
- [6] T. Tajima and J. M. Dawson, Phys. Rev. Lett. 43, 267 (1979).
- [7] J.-L. Vay, Phys. Rev. Lett. 98, 130405 (2007).
- [8] R. Hockney and J. Eastwood, Computer Simulation Using Particles (Taylor & Francis, London, 1988).
- [9] C. Birdsall and A. Langdon, *Plasma Physics via Computer Simulation*, *Appendix E*, Series in Plasma Physics (Taylor & Francis, London, 2004).
- [10] S. F. Martins, R. A. Fonseca, L. O. Silva, W. Lu, and W. B. Mori, Comput. Phys. Commun. 181, 869 (2010).
- [11] J. Vay, C. G. R. Geddes, C. Benedetti, D. L. Bruhwiler, E. Cormier-Michel, B. M. Cowan, J. R. Cary, and D. P. Grote, AIP Conf. Proc. 1299, 244 (2010).
- [12] J.-L. Vay, C. Geddes, E. Cormier-Michel, and D. Grote, J. Comput. Phys. 230, 5908 (2011).
- [13] J.-L. Vay, C. G. R. Geddes, E. Cormier-Michel, and D. P. Grote, Phys. Plasmas 18, 030701 (2011).
- [14] B. B. Godfrey, J. Comput. Phys. 15, 504 (1974).
- [15] B. B. Godfrey, J. Comput. Phys. 19, 58 (1975).
- [16] B. B. Godfrey and J.-L. Vay, J. Comput. Phys. 248, 33 (2013).
- [17] X. Xu, P. Yu, S. F. Martins, F. S. Tsung, V. K. Decyk, J. Vieira, R. A. Fonseca, W. Lu, L. O. Silva, and W. B. Mori, Comput. Phys. Commun. 184, 2503 (2013).
- [18] B. B. Godfrey, J.-L. Vay, and I. Haber, J. Comput. Phys. 258, 689 (2014).
- [19] B. Godfrey, J.-L. Vay, and I. Haber, IEEE Trans. Plasma Sci. 42, 1339 (2014).
- [20] B. B. Godfrey and J.-L. Vay, J. Comput. Phys. 267, 1 (2014).
- [21] B. B. Godfrey, AIP Conf. Proc. 1777, 020004 (2016).
- [22] B. B. Godfrey and J.-L. Vay, Comput. Phys. Commun. 196, 221 (2015).
- [23] P. Yu, X. Xu, V. K. Decyk, F. Fiuza, J. Vieira, F. S. Tsung, R. A. Fonseca, W. Lu, L. O. Silva, and W. B. Mori, Comput. Phys. Commun. 192, 32 (2015).

- [24] P. Yu, X. Xu, A. Tableman, V. K. Decyk, F. S. Tsung, F. Fiuza, A. Davidson, J. Vieira, R. A. Fonseca, W. Lu, L. O. Silva, and W. B. Mori, Comput. Phys. Commun. 197, 144 (2015).
- [25] F. Li, P. Yu, X. Xu, F. Fiuza, V. K. Decyk, T. Dalichaouch, A. Davidson, A. Tableman, W. An, F. S. Tsung, R. A. Fonseca, W. Lu, and W. B. Mori, arXiv:1605.01496 [physics.comp-ph].
- [26] A. F. Lifschitz, X. Davoine, E. Lefebvre, J. Faure, C. Rechatin, and V. Malka, J. Comput. Phys. 228, 1803 (2009).
- [27] A. Davidson, A. Tableman, W. An, F. S. Tsung, W. Lu, J. Vieira, R. A. Fonseca, L. O. Silva, and W. B. Mori, J. Comput. Phys. 281, 1063 (2015).
- [28] M. Kirchen, R. Lehe, B. B. Godfrey, I. Dornmair, S. Jalas, K. Peters, J.-L. Vay, and A. R. Maier, Phys. Plasmas 23, 100704 (2016).
- [29] I. Haber, R. Lee, H. Klein, and J. Boris, in *Proceedings of the Sixth Conference on Numerical Simulation of Plasmas* (Lawrence Livermore Laboratory, Livermore, CA, 1973), pp. 46–48.
- [30] J.-L. Vay, I. Haber, and B. B. Godfrey, J. Comput. Phys. **243**, 260 (2013).
- [31] J. Boris, in *Proceeding of the Fourth Conference on Numerical Simulations of Plasmas* (Naval Research Laboratory, Washington, DC, 1970).
- [32] J.-L. Vay, Phys. Plasmas (1994-present) 15, 056701 (2008).
- [33] J.-L. Vay, D. P. Grote, R. H. Cohen, and A. Friedman, Comput. Sci. Discov. 5, 014019 (2012).
- [34] D. Bruhwiler, J. Cary, B. Cowan, K. Paul, C. Geddes, P. Mullowney, P. Messmer, E. Esarey, E. Cormier-Michel, W. Leemans, and J.-L. Vay, AIP Conf. Proc. 1086, 29 (2009).
- [35] S. F. Martins, R. A. Fonseca, W. Lu, W. B. Mori, and L. O. Silva, Nat. Phys. 6, 311 (2010).
- [36] R. Lehe, M. Kirchen, I. A. Andriyash, B. B. Godfrey, and J.-L. Vay, Comput. Phys. Commun. **203**, 66 (2016).
- [37] I. A. Andriyash, R. Lehe, and A. Lifschitz, Phys. Plasmas 23, 033110 (2016).
- [38] C. Birdsall and A. Langdon, *Plasma Physics via Computer Simulation*, *Appendix E*, Series in Plasma Physics (Taylor & Francis, London, 2004).
- [39] J.-L. Vay and A. Arefiev, AIP Conf. Proc. 1777, 030002 (2016).
- [40] H. Vincenti and J.-L. Vay, Comput. Phys. Commun. 200, 147 (2016)
- [41] A. Huebl, R. Lehe, J.-L. Vay, D. P. Grote, I. Sbalzarini, S. Kuschel, M. Bussmann, and A. Huebl, "openpmd 1.0.0: Initial release" (2015).