

# Radiation torque on a birefringent sphere caused by an electromagnetic wave

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We present an exact *ab initio* calculation of the optical torque on a spherical uniaxially birefringent particle of arbitrary size illuminated by plane electromagnetic wave of arbitrary polarization mode and direction of propagation. The calculation is based on the extended Mie theory and the Maxwell stress tensor formalism. The expression for evaluating radiation torque is derived for arbitrary (absorbing and lossless) isotropic surrounding medium. The dependence of the optical torque on the incident angle, the polarization mode, the material birefringence, as well as the particle size, has been systematically investigated. For normal illumination, namely, with the incident wave vector  $\mathbf{k}_0$  perpendicular to the extraordinary axis (EA) of the particle, the optical torque  $\Gamma$  caused by a linearly polarized (LP) incident wave always shows the angle dependence  $\Gamma = \Gamma_0 \sin 2\varphi_e$ . Here,  $\varphi_e$  is the angle between the EA and the incident electric field, whereas  $\Gamma_0$  may take positive or negative values, dependent on  $n_e, n_o$ , and the particle size. In the small particle limit,  $\Gamma$  versus particle radius  $a$  displays different power law behaviors,  $\Gamma \sim a^3$  and  $\Gamma \sim a^6$ , for LP and circularly polarized (CP) incident waves, respectively, while for small material birefringence  $|\Delta n| = |n_e - n_o|$ , linear and square laws,  $\Gamma \sim |\Delta n|$  and  $\Gamma \sim |\Delta n|^2$ , are found for the LP and the CP incident modes, respectively.

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## I. INTRODUCTION

It is well known that electromagnetic wave transports angular momentum as well as linear momentum. When scattered, in addition to optical force caused by the transfer of linear momentum, the electromagnetic wave also exerts on the scatterer a torque due to the transfer of angular momentum. In a certain case, this torque may make the scatterer rotate [1], which, when combined with optical tweezers [2], introduces a unique and simple handle to control both the location and the orientation of a particular micro-sized particle. The advance has led to a wide variety of applications, including the possibility of rotating biological structure and developing optically driven and controlled micromachines. It has been shown [3] that the optical torque would vanish for an isotropic nonabsorbing spherical scatterer. So, early rotational micromanipulation was achieved by using absorbing particles [4,5]. To overcome the heating problem in the absorbing particle that limits the achievable rotation rate, optical systems have recently been designed with lossless geometrically anisotropic [6–11] and optically anisotropic (birefringent) particles [12–19]. In particular, optical apparatuses have been demonstrated [18,19] that can apply and accurately measure the torque exerted by laser beam on a birefringent particle.

Although intensively investigated experimentally, the optical torque on a birefringent particle was theoretically evaluated mostly in two limiting cases:  $\lambda \ll a$  [12,20] and  $\lambda \gg a$  [18]. Here,  $\lambda$  denotes the incident wavelength, and  $a$  the dimension of the scattering particle, e.g., the radius of spherical scatterer. Friese and co-workers [12] proposed that the optical torque per unit area is given by

$$\Gamma = -\frac{\epsilon c}{2\omega} E_0^2 \sin[k_0 d(n_o - n_e)] \cos 2\phi \sin 2\varphi_e + \frac{\epsilon c}{2\omega} E_0^2 \{1 - \cos[k_0 d(n_o - n_e)]\} \sin 2\phi, \quad (1)$$

where  $\epsilon$  is the electric permittivity,  $k_0$  the free-space wave number,  $\omega$  the angular frequency of incident wave,  $c$  the light velocity,  $d$  the thickness of the particle, and  $n_o$  and  $n_e$  are the refractive indices of the birefringent particle in, respectively, the ordinary and extraordinary directions, whereas  $\varphi_e$  is the angle between the extraordinary axis (EA) of the birefringent particle and the incident electric field, and finally,  $\phi$  describes the degree of ellipticity of the incident light [12], with  $\phi=0$  or  $\pi/2$  corresponding to linearly polarized (LP) and  $\phi=\pi/4$  circularly polarized (CP) light. The expression, ideally suited for a flat disk, is based on ray optics, which is valid in the limiting case  $\lambda \ll a$ . La Porta and Wang presented another expression [18], which was implicit in the earlier analysis of more elementary cases (see, e.g., Ref. [3])

$$\Gamma = \hat{q} \Gamma_0 \sin 2\varphi_e, \quad (2)$$

where  $\hat{q}$  is a unit vector normal to the electric field and the polarization induced on the scatterer, and  $\Gamma_0$  is the maximum magnitude of the optical torque. Equation (2) was derived in the limiting case  $\lambda \gg a$ . The two expressions meet difficulty for the intermediate case, where the size of the scattering particle  $a$  and the incident wavelength  $\lambda$  are of the same order. It is therefore desirable to have a complete solution to the optical torque problem on birefringent particle.

The purpose of this paper is to present an exact *ab initio* calculation of the radiation torque on a spherical uniaxially birefringent particle of arbitrary size, caused by an incident

plane electromagnetic wave with arbitrary polarization mode and direction of propagation. The dependence of the radiation torque on the incident angle, incident polarization mode, incident wavelength, as well as the material birefringence has been systematically investigated. The results verify the wave-plate behavior in the transfer of angular momentum from light to the scatterer. It is found that the angle dependence of Eq. (2), which is derived in the limiting case  $\lambda \gg a$ , remains valid for arbitrary particle size when the particle is subject to normal illumination, namely, with the incident wave vector  $\mathbf{k}_0$  normal to the EA, by the LP incident wave. However, the maximum magnitude  $\Gamma_0$  of the optical torque may change sign as particle size increases, in qualitative agreement with the first term of Eq. (1) derived based on ray optics. Due to the shape effect, at some particular ranges of particle size, the EA can be aligned by the LP incidence with the  $E$ -field direction regardless of  $n_e - n_o = \pm |\Delta n|$ , presenting a striking contrast to Eq. (1), since the latter implies that, for the LP incidence, particles with  $n_e - n_o = |\Delta n|$  and  $n_e - n_o = -|\Delta n|$  are always subject to torque in opposite directions. For different polarization modes of the incident plane wave, the radiation torque  $\Gamma$  versus the particle radius  $a$  may display different power-law behaviors in the small particle limit. In addition, the torque is found to exhibit linear and square law dependence on the material birefringence  $\Delta n = n_e - n_o$  for small  $|\Delta n|$ , regardless of the particle size.

The rest of the paper is organized as follows. In Sec. II, we describe a comprehensive formulation of exact *ab initio* calculation of radiation torque based on the Mie-type solution and the Maxwell stress tensor formalism. Systematic numerical results are presented in Sec. III. A summary is given in Sec. IV. Some technical results are relegated to the Appendix.

## II. GENERAL FORMULATION

The Mie-type solution to the scattering problem of plane electromagnetic waves by a sphere of arbitrary size and with gyromagnetic permeability tensor has recently been presented [21]. Here, we first outline the scheme of the solution to the scattering problem and then concentrate on the derivation of the radiation torque formulas based on the Maxwell stress tensor formalism.

### A. Mie-type solution of scattering by a birefringent sphere

The Maxwell equations for time-harmonic field inside the sourceless and homogeneous spherical particle read (assuming time dependence  $e^{-i\omega t}$ )

$$\nabla \times \mathbf{E}_I = i\omega \mathbf{B}_I, \quad (3a)$$

$$\nabla \times \mathbf{H}_I = -i\omega \mathbf{D}_I, \quad (3b)$$

$$\nabla \cdot \mathbf{D}_I = 0, \quad (3c)$$

$$\nabla \cdot \mathbf{B}_I = 0. \quad (3d)$$

The constitutive relations between the electric displacement vector  $\mathbf{D}_I$ , the magnetic induction  $\mathbf{B}_I$ , the electric field  $\mathbf{E}_I$ ,

and the magnetic field  $\mathbf{H}_I$  inside the particle are given, for a birefringent particle, by

$$\mathbf{D}_I = \epsilon_s \vec{\epsilon} \cdot \mathbf{E}_I, \quad \mathbf{B}_I = \mu_s \mathbf{H}_I, \quad (4)$$

where the permittivity tensor  $\epsilon_s \vec{\epsilon}$  is given by

$$\epsilon_s \vec{\epsilon} = \epsilon_s \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1+u \end{pmatrix}, \quad (5)$$

and  $\epsilon_s(\mu_s)$  is the scalar permittivity (permeability). For simplicity,  $\mu_s$  is set to be  $\mu_0$ , the permeability of the surrounding medium, while  $\epsilon_s = \epsilon_r \epsilon_0$ , with  $\epsilon_0$  denoting the permittivity outside particle. It follows from (3) and (4) that  $\mathbf{D}_I$  satisfies the wave equation

$$\nabla \times \nabla \times (\vec{\epsilon}^{-1} \cdot \mathbf{D}_I) - k_s^2 \mathbf{D}_I = \mathbf{0}, \quad (6)$$

with  $k_s^2 = \omega^2 \epsilon_s \mu_s$  and

$$\vec{\epsilon}^{-1} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & \frac{1}{1+u} \end{pmatrix}. \quad (7)$$

The divergenceless property (3c) suggests that  $\mathbf{D}_I$  be expanded in terms of the vector spherical wave functions (VSWFs)  $\mathbf{M}_{mn}^{(1)}(k, \mathbf{r})$  and  $\mathbf{N}_{mn}^{(1)}(k, \mathbf{r})$  [21–23]

$$\mathbf{D}_I = \sum_{n,m} E_{mn} [c_{mn} \mathbf{M}_{mn}^{(1)}(k, \mathbf{r}) + d_{mn} \mathbf{N}_{mn}^{(1)}(k, \mathbf{r})], \quad (8)$$

where  $c_{mn}$  and  $d_{mn}$  are the expansion coefficients, and  $k$  is as yet undetermined. In general, there are three kinds of VSWFs:  $\mathbf{M}_{mn}^{(J)}(k, \mathbf{r})$ ,  $\mathbf{N}_{mn}^{(J)}(k, \mathbf{r})$ , and  $\mathbf{L}_{mn}^{(J)}(k, \mathbf{r})$ . The divergenceless property of  $\mathbf{D}$  implies that it does not involve  $\mathbf{L}_{mn}$ , thereby simplifying the algebra involved. The three kinds of VSWFs are defined for  $J=1$  and 3 as in Refs. [21–23]. Except otherwise explicitly specified, hereinafter the summation  $\sum_{n,m}$  implies that  $n$  runs from 1 to  $+\infty$  and  $m$  from  $-n$  to  $+n$  for each  $n$ . The implication of  $\sum_{\nu,\mu}$  is similar. The prefactor  $E_{mn}$  is given by  $E_{mn} = i^n E_0 C_{mn}$ , with  $E_0$  characterizing the amplitude of the electric field of incident wave and [22]

$$C_{mn} = \left[ \frac{2n+1}{n(n+1)} \frac{(n-m)!}{(n+m)!} \right]^{1/2}. \quad (9)$$

By using the properties of VSWFs, it can be worked out that [21]

$$\begin{aligned} \vec{\epsilon}^{-1} \cdot \mathbf{M}_{mn} &= \sum_{\nu=0}^{+\infty} \sum_{\mu=-\nu}^{+\nu} [\tilde{g}_{\mu\nu}^{mn} \mathbf{M}_{\mu\nu} + \tilde{e}_{\mu\nu}^{mn} \mathbf{N}_{\mu\nu} + \tilde{f}_{\mu\nu}^{mn} \mathbf{L}_{\mu\nu}], \\ \vec{\epsilon}^{-1} \cdot \mathbf{N}_{mn} &= \sum_{\nu=0}^{+\infty} \sum_{\mu=-\nu}^{+\nu} [\tilde{g}_{\mu\nu}^{mn} \mathbf{M}_{\mu\nu} + \tilde{e}_{\mu\nu}^{mn} \mathbf{N}_{\mu\nu} + \tilde{f}_{\mu\nu}^{mn} \mathbf{L}_{\mu\nu}]. \end{aligned} \quad (10)$$

Therefore, one has

$$\begin{aligned} \tilde{\epsilon}^{-1} \cdot \mathbf{D}_I = & \sum_{n,m} E_{mn} [\bar{c}_{mn} \mathbf{M}_{mn}^{(1)}(k, \mathbf{r}) + \bar{d}_{mn} \mathbf{N}_{mn}^{(1)}(k, \mathbf{r}) \\ & + w_{mn} \mathbf{L}_{mn}^{(1)}(k, \mathbf{r})] + w_{00} \mathbf{L}_{00}^{(1)}(k, \mathbf{r}), \end{aligned} \quad (11)$$

where the coefficients  $\tilde{g}_{\mu\nu}^{mn}, \tilde{e}_{\mu\nu}^{mn}, \tilde{f}_{\mu\nu}^{mn}$  etc., together with  $\bar{d}_{mn}, \bar{c}_{mn}, w_{mn}$ , and  $w_{00}$ , can be obtained in the same way as in Ref. [21].

Inserting (8) and (11) into the wave equation (6), after some algebra one gets

$$\sum_{n,m} E_{mn} [\bar{c}_{mn} \mathbf{M}_{mn}^{(1)}(k, \mathbf{r}) + \bar{d}_{mn} \mathbf{N}_{mn}^{(1)}(k, \mathbf{r})] = \mathbf{0}, \quad (12)$$

with

$$\begin{aligned} \bar{c}_{mn} = & k^2 \sum_{\nu,\mu} \frac{E_{\mu\nu}}{E_{mn}} [\tilde{g}_{mn}^{\mu\nu} c_{\mu\nu} + \tilde{g}_{mn}^{\nu\mu} d_{\mu\nu}] - k_s^2 c_{mn}, \\ \bar{d}_{mn} = & k^2 \sum_{\nu,\mu} \frac{E_{\mu\nu}}{E_{mn}} [\tilde{e}_{mn}^{\mu\nu} c_{\mu\nu} + \tilde{e}_{mn}^{\nu\mu} d_{\mu\nu}] - k_s^2 d_{mn}. \end{aligned} \quad (13)$$

Equations (12) and (13) imply an eigensystem governing the value of  $k$  for expansion (8)

$$\begin{pmatrix} \bar{\mathcal{E}} & \tilde{\mathcal{E}} \\ \bar{\mathcal{G}} & \tilde{\mathcal{G}} \end{pmatrix} \begin{pmatrix} d \\ c \end{pmatrix} = \lambda \begin{pmatrix} d \\ c \end{pmatrix}, \quad (14)$$

where  $\lambda = k_s^2 / k^2$ , and the matrices  $\tilde{\mathcal{G}}, \bar{\mathcal{G}}, \tilde{\mathcal{E}}$ , and  $\bar{\mathcal{E}}$  are given by

$$\begin{aligned} \bar{\mathcal{G}}_{mn,\mu\nu} = & \frac{E_{\mu\nu}}{E_{mn}} \tilde{g}_{mn}^{\mu\nu}, & \tilde{\mathcal{G}}_{mn,\mu\nu} = & \frac{E_{\mu\nu}}{E_{mn}} \bar{g}_{mn}^{\mu\nu}, \\ \tilde{\mathcal{E}}_{mn,\mu\nu} = & \frac{E_{\mu\nu}}{E_{mn}} \tilde{e}_{mn}^{\mu\nu}, & \bar{\mathcal{E}}_{mn,\mu\nu} = & \frac{E_{\mu\nu}}{E_{mn}} \bar{e}_{mn}^{\mu\nu}, \end{aligned} \quad (15)$$

with  $mn$  and  $\mu\nu$  denoting the row and column indices, respectively. Let  $\lambda_l$  and  $(d_{mn,l}, c_{mn,l})^T$  denote, respectively, the eigenvalues and the corresponding eigenvectors of eigensystem (14), with  $l$  representing the index of eigenvalues and corresponding eigenvectors. One can then construct a new set of vector functions  $\mathbf{V}_l$  based on the eigenvectors  $(d_{mn,l}, c_{mn,l})^T$

$$\mathbf{V}_l = -\frac{i\epsilon_s}{\lambda_l} \sum_{n,m} E_{mn} [c_{mn,l} \mathbf{M}_{mn}^{(1)}(k_l, \mathbf{r}) + d_{mn,l} \mathbf{N}_{mn}^{(1)}(k_l, \mathbf{r})], \quad (16)$$

with  $k_l = k_s / \sqrt{\lambda_l}$ . It follows directly from

$$\nabla \cdot \mathbf{M}_{mn} = \nabla \cdot \mathbf{N}_{mn} = 0 \quad (17)$$

that  $\mathbf{V}_l$  are divergenceless

$$\nabla \cdot \mathbf{V}_l = 0. \quad (18a)$$

In addition, they satisfy the wave equation for  $\mathbf{D}_I$  field (6)

$$\nabla \times \nabla \times (\tilde{\epsilon}^{-1} \cdot \mathbf{V}_l) - k_s^2 \mathbf{V}_l = \mathbf{0}. \quad (18b)$$

Thus,  $\mathbf{D}_I$  can be expanded in terms of  $\mathbf{V}_l$

$$\mathbf{D}_I = \sum_l \alpha_l \mathbf{V}_l, \quad (19)$$

where the expansion coefficients  $\alpha_l$  are to be determined by matching the boundary conditions at the surface of sphere. With  $\mathbf{D}_I$  given by (19), it follows from (3a) and (4) that  $\mathbf{E}_I$  and  $\mathbf{H}_I$  fields can be written as

$$\begin{aligned} \mathbf{E}_I = & \frac{1}{\epsilon_s} \tilde{\epsilon}^{-1} \cdot \mathbf{D}_I = -\sum_{n,m} iE_{mn} \sum_l \alpha_l \left[ c_{mn,l} \mathbf{M}_{mn}^{(1)}(k_l, \mathbf{r}) + d_{mn,l} \right. \\ & \left. \times \mathbf{N}_{mn}^{(1)}(k_l, \mathbf{r}) + \frac{w_{mn,l}}{\lambda_l} \mathbf{L}_{mn}^{(1)}(k_l, \mathbf{r}) \right] \\ & + \sum_l i\alpha_l \left[ \frac{w_{00,l}}{\lambda_l} \mathbf{L}_{00}^{(1)}(k_l, \mathbf{r}) \right], \\ \mathbf{H}_I = & \frac{-i}{\omega\mu_s} \nabla \times \mathbf{E}_I = -\frac{1}{\omega\mu_s} \sum_{n,m} E_{mn} \sum_l k_l \alpha_l [d_{mn,l} \mathbf{M}_{mn}^{(1)}(k_l, \mathbf{r}) \\ & + c_{mn,l} \mathbf{N}_{mn}^{(1)}(k_l, \mathbf{r})], \end{aligned} \quad (20)$$

with

$$\begin{aligned} w_{mn,l} = & \sum_{\nu,\mu} \frac{E_{\mu\nu}}{E_{mn}} [\tilde{f}_{mn}^{\mu\nu} d_{\mu\nu,l} + \tilde{f}_{mn}^{\nu\mu} c_{\mu\nu,l}], \\ w_{00,l} = & \sum_{\nu,\mu} E_{\mu\nu} [\tilde{f}_{00}^{\mu\nu} d_{\mu\nu,l} + \tilde{f}_{00}^{\nu\mu} c_{\mu\nu,l}] = -\sqrt{\frac{2}{15}} u d_{02,l}. \end{aligned}$$

Notice that, since  $\mathbf{E}_I$  no longer satisfies  $\nabla \cdot \mathbf{E}_I = 0$ , its expansion includes  $\mathbf{L}_{mn}$  terms that are absent in the isotropic case.

The scattered fields  $\mathbf{E}_s, \mathbf{H}_s$ , and incident fields  $\mathbf{E}_{\text{inc}}, \mathbf{H}_{\text{inc}}$  in the isotropic surrounding medium have the same form as in the Mie solution [23,24]. However, since the form of the permittivity tensor adopted [see (5)] implies that the EA of birefringent particle is along the  $z$  axis, for arbitrary incident direction of plane wave, the expansions of the field are not limited to  $m = \pm 1$  modes.

In terms of VSWFs, the scattered fields  $(\mathbf{E}_s, \mathbf{H}_s)$  are expanded as

$$\begin{aligned} \mathbf{E}_s = & \sum_{n,m} iE_{mn} [a_{mn} \mathbf{N}_{mn}^{(3)}(k_0, \mathbf{r}) + b_{mn} \mathbf{M}_{mn}^{(3)}(k_0, \mathbf{r})], \\ \mathbf{H}_s = & \frac{k_0}{\omega\mu_{0,n,m}} \sum_{n,m} E_{mn} [b_{mn} \mathbf{N}_{mn}^{(3)}(k_0, \mathbf{r}) + a_{mn} \mathbf{M}_{mn}^{(3)}(k_0, \mathbf{r})], \end{aligned} \quad (21)$$

where  $k_0^2 = \omega^2 \epsilon_0 \mu_0$  with  $\epsilon_0$  and  $\mu_0$  being, respectively, the scalar permittivity and permeability of the surrounding medium. The expansion coefficients  $a_{mn}$  and  $b_{mn}$  are to be determined by matching boundary conditions.

Suppose that the particle is illuminated by a plane wave characterized by the incident wave vector  $\mathbf{k}_0$ , with

$$\mathbf{k}_0 = k_0 (\sin \theta_k \cos \phi_k \mathbf{e}_x + \sin \theta_k \sin \phi_k \mathbf{e}_y + \cos \theta_k \mathbf{e}_z), \quad (22)$$

where  $\mathbf{e}_x, \mathbf{e}_y$ , and  $\mathbf{e}_z$  are three unit base vectors of the Cartesian coordinate system and  $\theta_k(\phi_k)$  is the polar (azimuthal) angle of  $\mathbf{k}_0$ , as schematically shown in Fig. 1. The electric and magnetic fields of the incident plane wave are then

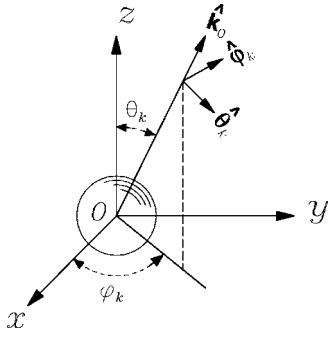


FIG. 1. Geometry of the scattering problem.

$$\mathbf{E}_{\text{inc}} = E_0(p_\theta \hat{\boldsymbol{\theta}}_k + p_\phi \hat{\boldsymbol{\phi}}_k) e^{i\mathbf{k}_0 \cdot \mathbf{r}},$$

$$\mathbf{H}_{\text{inc}} = \frac{k_0}{\omega \mu_0} E_0(p_\theta \hat{\boldsymbol{\phi}}_k - p_\phi \hat{\boldsymbol{\theta}}_k) e^{i\mathbf{k}_0 \cdot \mathbf{r}}, \quad (23)$$

where  $\hat{\mathbf{p}} = (p_\theta \hat{\boldsymbol{\theta}}_k + p_\phi \hat{\boldsymbol{\phi}}_k)$  is the normalized complex polarization vector, with  $|\hat{\mathbf{p}}| = 1$ , and the unit vectors  $\hat{\boldsymbol{\theta}}_k$  and  $\hat{\boldsymbol{\phi}}_k$  are defined in the direction of increasing  $\theta_k$  and  $\phi_k$  such as to constitute a right-hand base system together with  $\hat{\mathbf{k}}_0 = \mathbf{k}_0/k_0$ , as shown in Fig. 1, namely

$$\hat{\mathbf{k}}_0 \times \hat{\boldsymbol{\theta}}_k = \hat{\boldsymbol{\phi}}_k, \quad \hat{\boldsymbol{\theta}}_k \times \hat{\boldsymbol{\phi}}_k = \hat{\mathbf{k}}_0, \quad \hat{\boldsymbol{\phi}}_k \times \hat{\mathbf{k}}_0 = \hat{\boldsymbol{\theta}}_k. \quad (24)$$

In terms of VSWFs, the incident fields ( $\mathbf{E}_{\text{inc}}, \mathbf{H}_{\text{inc}}$ ) read

$$\begin{aligned} \mathbf{E}_{\text{inc}} &= - \sum_{n,m} i E_{mn} [p_{mn} \mathbf{N}_{mn}^{(1)}(k_0, \mathbf{r}) + q_{mn} \mathbf{M}_{mn}^{(1)}(k_0, \mathbf{r})], \\ \mathbf{H}_{\text{inc}} &= - \frac{k_0}{\omega \mu_0} \sum_{n,m} E_{mn} [q_{mn} \mathbf{N}_{mn}^{(1)}(k_0, \mathbf{r}) + p_{mn} \mathbf{M}_{mn}^{(1)}(k_0, \mathbf{r})]. \end{aligned} \quad (25)$$

The expansion coefficients  $p_{mn}$  and  $q_{mn}$  are

$$\begin{aligned} p_{mn} &= [p_\theta \tilde{\pi}_{mn}(\cos \theta_k) - i p_\phi \tilde{\tau}_{mn}(\cos \theta_k)] e^{-im\phi_k}, \\ q_{mn} &= [p_\theta \tilde{\pi}_{mn}(\cos \theta_k) - i p_\phi \tilde{\tau}_{mn}(\cos \theta_k)] e^{-im\phi_k}, \end{aligned} \quad (26)$$

where the regular angular functions  $\tilde{\pi}_{mn}(\cos \theta)$  and  $\tilde{\tau}_{mn}(\cos \theta)$  are defined based on the first kind of associated Legendre functions  $P_n^m(\cos \theta)$  [22–24]

$$\begin{aligned} \tilde{\pi}_{mn}(\cos \theta) &= C_{mn} \pi_{mn}(\cos \theta) = C_{mn} \frac{m}{\sin \theta} P_n^m(\cos \theta), \\ \tilde{\tau}_{mn}(\cos \theta) &= C_{mn} \tau_{mn}(\cos \theta) = C_{mn} \frac{d}{d\theta} P_n^m(\cos \theta), \end{aligned} \quad (27)$$

with  $C_{mn}$  defined in (9).

Matching boundary conditions at the surface of sphere, and after some algebra, one arrives at the equations that serve to determine the expansion coefficients  $\alpha_l$ ,  $a_{mn}$ , and  $b_{mn}$  based on  $p_{mn}$  and  $q_{mn}$

$$\begin{aligned} \frac{1}{m_s} \sum_l \frac{1}{\bar{k}_l \lambda_l} j_n(\bar{k}_l m_s \eta) w_{mn,l} \alpha_l + \xi_n'(\eta) a_{mn} \\ + \frac{1}{m_s} \sum_l \frac{1}{\bar{k}_l} \psi_n'(\bar{k}_l m_s \eta) d_{mn,l} \alpha_l = \psi_n'(\eta) p_{mn}, \end{aligned}$$

$$\xi_n(\eta) b_{mn} + \frac{1}{m_s} \sum_l \frac{1}{\bar{k}_l} \psi_n(\bar{k}_l m_s \eta) c_{mn,l} \alpha_l = \psi_n(\eta) q_{mn},$$

$$\xi_n(\eta) a_{mn} + \frac{\mu_0}{\mu_s} \sum_l \psi_n(\bar{k}_l m_s \eta) d_{mn,l} \alpha_l = \psi_n(\eta) p_{mn},$$

$$\xi_n'(\eta) b_{mn} + \frac{\mu_0}{\mu_s} \sum_l \psi_n'(\bar{k}_l m_s \eta) c_{mn,l} \alpha_l = \psi_n'(\eta) q_{mn}, \quad (28)$$

where

$$\eta = k_0 a, \quad m_s = \sqrt{1 + u} \frac{k_s}{k_0}, \quad \bar{k}_l = \frac{k_l}{k_s}, \quad \lambda_l = \frac{k_s^2}{k_l^2} = \frac{1}{\bar{k}_l^2}, \quad (29)$$

with  $a$  the radius of spherical particle. The Riccati-Bessel functions  $\psi_n(z)$ ,  $\xi_n(z)$ ,  $\chi_n(z)$  are given by [24]

$$\psi_n(z) = z j_n(z), \quad \xi_n(z) = z h_n^{(1)}(z), \quad \chi_n(z) = -z y_n(z). \quad (30)$$

Some details about numerical solution of (28) can be found in Ref. [21].

## B. Radiation torque

With the incident fields and the scattered field given by (25) and (21), respectively, the total external field outside the particle reads

$$\begin{aligned} \mathbf{E}_e &= \sum_{n,m} i E_{mn} [a_{mn} \mathbf{N}_{mn}^{(3)}(k_0, \mathbf{r}) + b_{mn} \mathbf{M}_{mn}^{(3)}(k_0, \mathbf{r}) - p_{mn} \\ &\quad \times \mathbf{N}_{mn}^{(1)}(k_0, \mathbf{r}) - q_{mn} \mathbf{M}_{mn}^{(1)}(k_0, \mathbf{r})], \\ \mathbf{H}_e &= \frac{k_0}{\omega \mu_0} \sum_{n,m} E_{mn} [b_{mn} \mathbf{N}_{mn}^{(3)}(k_0, \mathbf{r}) + a_{mn} \mathbf{M}_{mn}^{(3)}(k_0, \mathbf{r}) - q_{mn} \\ &\quad \times \mathbf{N}_{mn}^{(1)}(k_0, \mathbf{r}) - p_{mn} \mathbf{M}_{mn}^{(1)}(k_0, \mathbf{r})], \end{aligned} \quad (31)$$

which is explicitly reduced to

$$\begin{aligned} \mathbf{E}_e &= \frac{i}{k_0 r} \sum_{n,m} E_{mn} e^{im\phi} \left[ \frac{n(n+1)}{k_0 r} \mathcal{U}_{mn} P_n^m \mathbf{e}_r + (\mathcal{U}'_{mn} \tau_{mn} \right. \\ &\quad \left. + \mathcal{V}_{mn} i \pi_{mn}) \mathbf{e}_\theta + (\mathcal{U}'_{mn} i \pi_{mn} - \mathcal{V}_{mn} \tau_{mn}) \mathbf{e}_\phi \right], \end{aligned}$$

$$\mathbf{H}_e = \frac{k_0}{\omega\mu_0} \frac{1}{k_0 r} \sum_{n,m} E_{mn} e^{im\phi} \left[ \frac{n(n+1)}{k_0 r} \mathcal{V}_{mn} P_n^m \mathbf{e}_r + (\mathcal{V}'_{mn} \tau_{mn} + \mathcal{U}_{mn} i \pi_{mn}) \mathbf{e}_\theta + (\mathcal{V}'_{mn} i \pi_{mn} - \mathcal{U}_{mn} \tau_{mn}) \mathbf{e}_\phi \right], \quad (22)$$

where

$$\begin{aligned} \mathcal{U}_{mn} &= a_{mn} \xi_n(k_0 r) - p_{mn} \psi_n(k_0 r), \\ \mathcal{U}'_{mn} &= a_{mn} \xi'_n(k_0 r) - p_{mn} \psi'_n(k_0 r), \\ \mathcal{V}_{mn} &= b_{mn} \xi_n(k_0 r) - q_{mn} \psi_n(k_0 r), \\ \mathcal{V}'_{mn} &= b_{mn} \xi'_n(k_0 r) - q_{mn} \psi'_n(k_0 r). \end{aligned} \quad (33)$$

The time-averaged Maxwell stress tensor is then, for isotropic medium

$$\hat{\mathbf{T}} = \frac{1}{2} \text{Re} \left[ \mathbf{E}_e \mathbf{D}_e^* + \mathbf{H}_e \mathbf{B}_e^* - \frac{1}{2} (\mathbf{E}_e \cdot \mathbf{D}_e^* + \mathbf{H}_e \cdot \mathbf{B}_e^*) \hat{\mathbf{I}} \right], \quad (34)$$

where the superscript \* denotes the complex conjugate, and  $\hat{\mathbf{I}}$  is the unit dyadic. The time-averaged torque  $\mathbf{\Gamma}$  on the spherical scatterer can be evaluated by [25,26]

$$\mathbf{\Gamma} = - \int \int d\mathbf{S} \cdot \hat{\mathbf{K}} = - \int \int [\mathbf{e}_r \cdot \hat{\mathbf{K}}] dS, \quad (35)$$

where  $\mathbf{e}_r = \mathbf{r}/r$  is the unit vector in radial direction with  $r = |\mathbf{r}|$ , and the time-averaged angular momentum flux tensor  $\hat{\mathbf{K}}$  reads [26]

$$\hat{\mathbf{K}} = \hat{\mathbf{T}} \cdot [\mathbf{r} \times \hat{\mathbf{I}}] = \hat{\mathbf{T}} \times \mathbf{r}. \quad (36)$$

As a result, the time-averaged torque becomes

$$\begin{aligned} \mathbf{\Gamma} &= - \int \int [\mathbf{e}_r \cdot (\hat{\mathbf{T}} \times \mathbf{r})] dS = \int \int [\mathbf{r} \times (\hat{\mathbf{T}} \cdot \mathbf{e}_r)] dS \\ &= r^3 \int \int \mathbf{e}_r \times [\hat{\mathbf{T}} \cdot \mathbf{e}_r] d\Omega. \end{aligned} \quad (37)$$

After some algebra, the Cartesian components of the torque can be worked out as

$$\mathbf{\Gamma} = \text{Re} \{ [\Gamma_x^{(1)} + \Gamma_x^{(2)}] \mathbf{e}_x + [\Gamma_y^{(1)} + \Gamma_y^{(2)}] \mathbf{e}_y + [\Gamma_z^{(1)} + \Gamma_z^{(2)}] \mathbf{e}_z \}, \quad (38)$$

where

$$\begin{aligned} \Gamma_x^{(1)} &= \frac{1}{2} r^3 \epsilon_0^* \sum_{n,m} \sum_{\nu,\mu} [\mathcal{V}_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(x)} - \mathcal{U}'_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(x)}], \\ \Gamma_y^{(1)} &= \frac{1}{2} r^3 \epsilon_0^* \sum_{n,m} \sum_{\nu,\mu} [\mathcal{V}_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(y)} - \mathcal{U}'_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(y)}], \\ \Gamma_z^{(1)} &= \frac{1}{2} r^3 \epsilon_0^* \sum_{n,m} \sum_{\nu,\mu} [\mathcal{V}_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(z)} - \mathcal{U}'_{mn} \mathcal{U}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(z)}], \\ \Gamma_x^{(2)} &= \frac{1}{2} r^3 \frac{|\epsilon_0 \mu_0|}{\mu_0} \sum_{n,m} \sum_{\nu,\mu} [\mathcal{U}_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(x)} - \mathcal{V}'_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(x)}], \end{aligned} \quad (39)$$

$$\begin{aligned} \Gamma_y^{(2)} &= \frac{1}{2} r^3 \frac{|\epsilon_0 \mu_0|}{\mu_0} \sum_{n,m} \sum_{\nu,\mu} [\mathcal{U}_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(y)} - \mathcal{V}'_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(y)}], \\ \Gamma_z^{(2)} &= \frac{1}{2} r^3 \frac{|\epsilon_0 \mu_0|}{\mu_0} \sum_{n,m} \sum_{\nu,\mu} [\mathcal{U}_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\tau p}^{(z)} - \mathcal{V}'_{mn} \mathcal{V}_{\mu\nu}^* \mathcal{W}_{\pi p}^{(z)}], \end{aligned} \quad (40)$$

where  $\mathcal{W}_{\tau p}^{(x,y,z)}$  and  $\mathcal{W}_{\pi p}^{(x,y,z)}$  are integrals involving two regular auxiliary angular functions  $\pi_{mn}(\cos \theta)$  and  $\tau_{mn}(\cos \theta)$  given in the Appendix.

It should be noted that the integrations in Eqs. (35) and (37) are over the outer surface of the spherical particle, so  $\mathcal{U}_{mn}$ ,  $\mathcal{U}'_{mn}$ ,  $\mathcal{V}_{mn}$ , and  $\mathcal{V}'_{mn}$  given by Eq. (33), as well as  $\Gamma_{x,y,z}^{(1,2)}$  given in (39) and (40), should be calculated at the sphere surface  $r=a$ . If so evaluated, the formulas for radiation torque Eqs. (38)–(40) hold for both absorbing and lossless surrounding medium. They are presented here for arbitrary isotropic surrounding medium. If the background medium is lossless, with both permittivity and permeability being real numbers, then the integration (35) can be performed at spherical surface with arbitrary radius  $r > a$ , due to conservation of momentum and angular momentum. As a result, the integration is usually evaluated in the limit  $r \rightarrow \infty$  for lossless surrounding medium, where the field expressions become much simpler by using the asymptotical formulas for Riccati-Bessel functions

$$\begin{aligned} \xi_n(\rho) &\sim (-i)^{n+1} \exp(i\rho), \quad \zeta_n(\rho) \sim i^{n+1} \exp(-i\rho), \\ \psi_n(\rho) &\sim [\xi_n(\rho) + \zeta_n(\rho)]/2. \end{aligned} \quad (41)$$

Substituting Eq. (41) into Eqs. (39) and (40), after some algebra, the expressions for torque can be considerably simplified for lossless medium. They are

$$\Gamma_x = \text{Re}[\mathcal{N}_1], \quad \Gamma_y = \text{Im}[\mathcal{N}_1], \quad \Gamma_z = \text{Re}[\mathcal{N}_2], \quad (42)$$

where

$$\begin{aligned} \mathcal{N}_1 &= \frac{2\pi\epsilon_0 |E_0|^2}{k_0^3} \sum_{n,m} \rho_{mn} (\tilde{a}_{m,n} \tilde{a}_{m+1,n}^* + \tilde{b}_{m,n} \tilde{b}_{m+1,n}^* - \tilde{p}_{m,n} \tilde{p}_{m+1,n}^* \\ &\quad - \tilde{q}_{m,n} \tilde{q}_{m+1,n}^*), \\ \mathcal{N}_2 &= - \frac{2\pi\epsilon_0 |E_0|^2}{k_0^3} \sum_{n,m} m (|\tilde{a}_{m,n}|^2 + |\tilde{b}_{m,n}|^2 - |\tilde{p}_{m,n}|^2 - |\tilde{q}_{m,n}|^2), \end{aligned} \quad (43)$$

with  $\rho_{mn} = [(n-m)(n+m+1)]^{1/2}$ , whereas  $\tilde{a}_{m,n}$ ,  $\tilde{b}_{m,n}$ ,  $\tilde{p}_{m,n}$ , and  $\tilde{q}_{m,n}$  are given by

$$\begin{aligned} \tilde{a}_{m,n} &= a_{mn} - \frac{1}{2} p_{mn}, \quad \tilde{p}_{mn} = \frac{1}{2} p_{mn}, \\ \tilde{b}_{m,n} &= b_{mn} - \frac{1}{2} q_{mn}, \quad \tilde{q}_{mn} = \frac{1}{2} q_{mn}, \end{aligned} \quad (44)$$

with  $a_{mn}$ ,  $b_{mn}$ ,  $p_{mn}$ , and  $q_{mn}$  being coefficients in Eq. (31). We note that similar expressions for lossless background medium were presented by several authors [27–30].

### III. NUMERICAL RESULTS

We are now ready to present systematic numerical results. Except otherwise explicitly specified, the results shown are mostly for the case with  $\epsilon_r = \epsilon_y / \epsilon_0 = 2.56$ , roughly a typical value in most experiments. The material birefringence is thus due to the anisotropic parameter  $u$ . In our numerical calculation, the EA of the birefringent particle is set to be the  $z$  axis, without loss of generality. The axial symmetry of the permittivity (5) implies that the scattering is independent of the azimuthal angle  $\phi_k$  of the incident wave vector  $\mathbf{k}_0$ . As a result, for simplicity, we set  $\phi_k = 0$ , which means  $\hat{\phi}_k = \mathbf{e}_y$  and  $\mathbf{k}_0$  lies in the  $x$ - $y$  plane. The radiation torque is a function of incident angle  $\theta_k$ . In addition, as most experiments applied a laser beam with fixed incident wavelength while the size of the particle is changed (see, e.g., Ref. [16]), in our calculation the optical torque is presented in units of  $2I_0\lambda^3/c$ , with  $c$  the light velocity,  $\lambda$  the incident wavelength, and  $I_0 = \frac{1}{2}c\epsilon_0|E_0|^2$  the incident irradiance.

#### A. Oblique illumination

We first study the case of oblique illumination, namely, the case with  $\mathbf{k}_0$  not normal to the EA of the birefringent particle. Three different polarization modes of the incident plane wave are considered: (i) the TM mode, in which  $\hat{\mathbf{p}} = \hat{\theta}_k$  and the electric vector vibrates in the incident plane; (ii) the transverse electric (TE) mode, in which  $\hat{\mathbf{p}} = \hat{\phi}_k$  and thus the magnetic vector vibrates in the incident plane; and (iii) the left circularly polarized (LCP) mode with  $\hat{\mathbf{p}} = (1/\sqrt{2})(\hat{\theta}_k + i\hat{\phi}_k)$ . The incident plane is defined by the EA and  $\mathbf{k}_0$ . We do not study the right circularly polarized incident mode, since its results can be easily inferred from those for the LCP mode by symmetry.

The typical dependence of the optical torque on the incident angle  $\theta_k$  is shown, for  $0 \leq \theta_k \leq \pi/2$ , in Fig. 2 at size parameter  $\eta = k_0 a = 5$  with different values of birefringent parameter  $u$ . For  $\pi/2 \leq \theta_k \leq \pi$ , one has  $\Gamma_y(\theta_k) = -\Gamma_y(\pi - \theta_k)$  and  $\Gamma_x(\theta_k) = \Gamma_x(\pi - \theta_k)$  by symmetry. For TM and LCP incident modes, the  $y$  component of the optical torque  $\Gamma_y$  is found to display similar  $\theta_k$  dependence. When  $u > 0$  ( $u < 0$ ), corresponding to  $n_e > n_o$  ( $n_e < n_o$ ),  $\Gamma_y$  is negative (positive), indicating that the optical torque due to the TM incidence will align the EA with the direction parallel (normal) to the  $E$  field, and the CP incidence will make the EA parallel (normal) to the plane of the incident  $E$  field. Stable equilibrium is reached at  $\theta_k = \pi/2$  ( $\theta_k = 0$ ) for the  $u > 0$  ( $u < 0$ ) case. For the TM incident wave, the angle between the EA ( $z$  axis) and the incident electric field is  $\varphi_e = (\pi/2) - \theta_k$ , as shown in Fig. 2. However, the maximum value of  $\Gamma_y$  does not occur at  $\varphi_e = \theta_k = \pi/4$ , but instead depends on the material birefringence (the anisotropy parameter  $u$ ). This is different from the normal illumination case with  $\mathbf{k}_0$  perpendicular to the EA, where the torque always reaches its maximum at  $\varphi_e = \pi/4$  (see Fig. 5 below) independent of the value of  $u$ . For the TE case,  $\Gamma_y$  is much smaller than in the TM case, and the  $\theta_k$  dependence becomes somewhat complicated, especially for large  $u$ , as can be seen in Fig. 2(b). For the case with  $u = 0.4$ , e.g.,  $\Gamma_y$

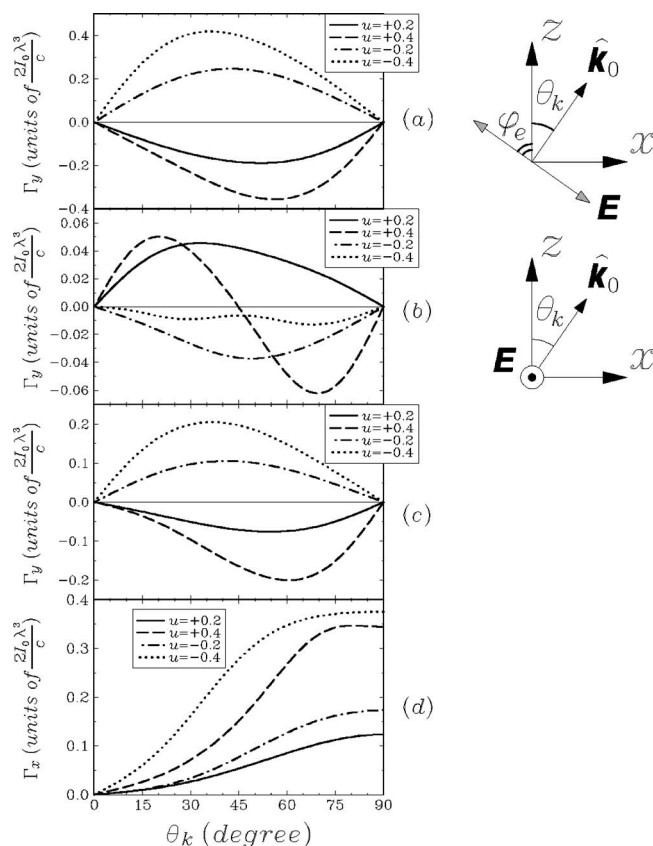


FIG. 2. The optical torque as a function of incident angle  $\theta_k$  at  $\eta = 5$  and different values of birefringent anisotropy parameter  $u$ . (a)  $\Gamma_y$  versus  $\theta_k$  for the TM incidence; (b)  $\Gamma_y$  versus  $\theta_k$  for the TE incidence; (c)  $\Gamma_y$  versus  $\theta_k$  for the LCP incidence; (d)  $\Gamma_x$  versus  $\theta_k$  for the LCP incidence. Also shown are the schematic plots of the directions of incident wave vectors and  $E$  fields for the TM and the TE incidences.

changes sign as  $\theta_k$  goes from 0 to  $\pi/2$ , resulting in two stable equilibrium states at both  $\theta_k = 0$  and  $\theta_k = \pi/2$ , quite different from the case of  $u = 0.2$ , where only  $\theta_k = 0$  is rotationally stable. While for the TE and TM cases,  $\Gamma_y$  is the only nonvanishing component of the radiation torque, for the LCP incidence, besides  $\Gamma_y$ , the incident wave also exerts a positive  $\Gamma_x$  on the particle, trying to make the EA rotate together with the  $E$  field. When  $\eta$  and  $u$  are not very large, the magnitude of  $\Gamma_x$  monotonically increases with the incident angle  $\theta_k$  in the region from 0 to  $\pi/2$ , and reaches a maximum at  $\theta_k = \pi/2$ ; then it decreases with  $\theta_k$  in the region from  $\pi/2$  to  $\pi$  due to symmetry  $\Gamma_x(\theta_k) = \Gamma_x(\pi - \theta_k)$ .

Figure 3 shows  $\Gamma_y / \eta^3$  and  $\Gamma_x / \eta^2$  versus  $\theta_k$  at  $u = 0.2$  for different values of size parameter  $\eta$ . Here,  $\Gamma_y / \eta^3$  and  $\Gamma_x / \eta^2$ , instead of  $\Gamma_y$  and  $\Gamma_x$ , versus  $\theta_k$  are exhibited for the convenience of displaying curves for different  $\eta$  in the same figure. For the TM and LCP incident waves, the  $\theta_k$  dependence of  $\Gamma_y$  is similar to the case of small  $\eta$ , implying that, irrespective of the value of  $\eta$ , for the TM case, the optical torque always tends to make the EA parallel (normal) to the incident  $E$  field for particle with  $n_e > n_o$  ( $n_e < n_o$ ), and the LCP incidence will align the EA with the directions parallel (normal) to the  $E$ -field plane for particle with  $n_e > n_o$  ( $n_e < n_o$ ). For the

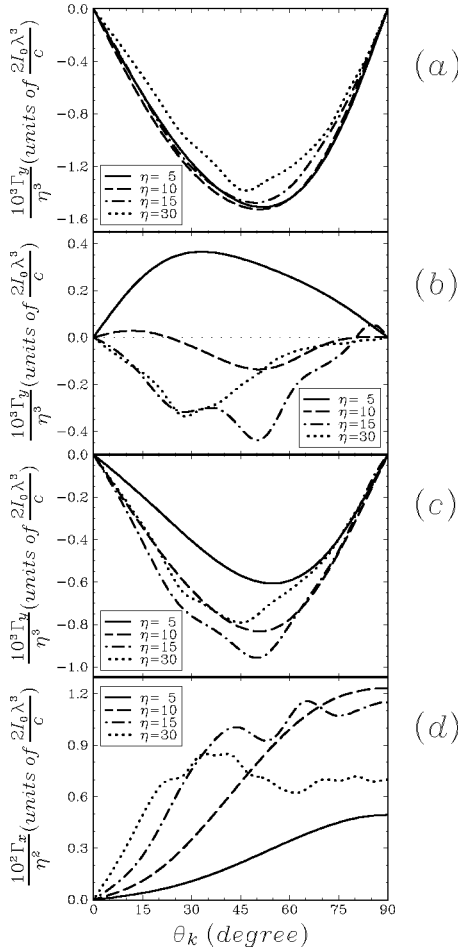


FIG. 3. Dependence of  $\Gamma_y/\eta^3$  and  $\Gamma_x/\eta^2$  on  $\theta_k$  at  $u=0.2$  and different values of  $\eta$ . (a)  $10^3\Gamma_y/\eta^3$  versus  $\theta_k$  for the TM incidence; (b)  $10^3\Gamma_y/\eta^3$  versus  $\theta_k$  for the TE incidence; (c)  $10^3\Gamma_y/\eta^3$  versus  $\theta_k$  for the LCP incidence; (d)  $10^2\Gamma_x/\eta^2$  versus  $\theta_k$  for the LCP incidence.

TE case, the  $\theta_k$  dependence becomes complicated as  $\eta$  increases. For example, when  $\eta=5$ ,  $\theta_k=0$  is the only rotationally stable equilibrium. For  $\eta=30$ , the only stable equilibrium is reached at  $\theta_k=\pi/2$ , in contrast to the case with  $\eta=5$ . If  $\eta=15$ , a stable equilibrium is achieved at  $\theta_k\approx 80^\circ$ . Neither  $\theta_k=0$  nor  $\theta_k=\pi/2$  is stable, although  $\Gamma_y$  vanishes at such values of  $\theta_k$ . While  $\eta=10$ , both  $\theta_k=0$  and  $\theta_k\approx 80^\circ$  are stable. Finally, for the LCP case, the  $\theta_k$  dependence of  $\Gamma_x$  is no longer monotonic in the region  $0\leq\theta_k\leq\pi/2$  when  $\eta$  is large, as exhibited in Fig. 3(d) for  $\eta=15$  and  $30$ .

To study the particle size dependence of the optical torque in the small particle limit, we consider also the general LP incident mode, in addition to the TM, TE, and LCP modes. Here, the general LP mode is characterized by the polarization vector  $\hat{\mathbf{p}}=p_\theta\hat{\boldsymbol{\theta}}_k+p_\phi\hat{\boldsymbol{\phi}}_k$ , with both  $p_\theta$  and  $p_\phi$  real and neither  $p_\theta$  nor  $p_\phi$  vanishing. It is taken into account in order to address the particle size dependence of  $\Gamma_x$ , which is vanishing for both the TM ( $p_\phi=0$ ) and the TE ( $p_\theta=0$ ) cases. Both  $\Gamma_x$  and  $\Gamma_y$  versus particle size are found to exhibit different power laws for different incident polarization modes, as exemplified in Fig. 4 for  $\theta_k=55^\circ$  and  $u=0.2$ . Figure 4(a) and Fig. 4(b) display  $\Gamma_x/\eta^6$  versus  $\log_{10}\eta$  for the LCP inci-

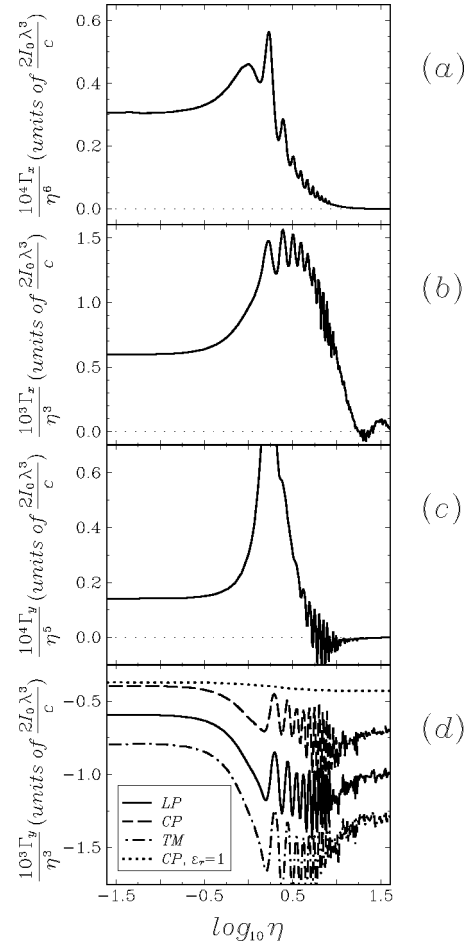


FIG. 4. (a)  $10^4\Gamma_x/\eta^6$  versus  $\log_{10}\eta$  at  $u=0.2$  and incident angle  $\theta_k=55^\circ$  for the LCP incidence with  $\hat{\mathbf{p}}=(1/\sqrt{2})(\hat{\boldsymbol{\theta}}_k+i\hat{\boldsymbol{\phi}}_k)$ ; (b)  $10^3\Gamma_x/\eta^3$  versus  $\log_{10}\eta$  at  $u=0.2$  and  $\theta_k=55^\circ$  for the LP incidence with  $\hat{\mathbf{p}}=\frac{1}{2}(\sqrt{3}\hat{\boldsymbol{\theta}}_k+\hat{\boldsymbol{\phi}}_k)$ ; (c)  $10^4\Gamma_y/\eta^5$  versus  $\log_{10}\eta$  at  $u=0.2$  and  $\theta_k=55^\circ$  for the TE incidence with  $\hat{\mathbf{p}}=\hat{\boldsymbol{\phi}}_k$ ; (d)  $10^3\Gamma_y/\eta^3$  versus  $\log_{10}\eta$  at  $u=0.2$  and  $\theta_k=55^\circ$  for the LP, LCP, and TM incidences. Also shown as dotted line is  $10^3\Gamma_y/\eta^3$  versus  $\log_{10}\eta$  for the LCP case at matching case  $\epsilon_r=1$ .

dent wave and  $\Gamma_x/\eta^3$  versus  $\log_{10}\eta$  for the LP incident wave with  $\hat{\mathbf{p}}=\frac{1}{2}(\sqrt{3}\hat{\boldsymbol{\theta}}_k+\hat{\boldsymbol{\phi}}_k)$ , respectively. The flat curves at small  $\eta$  implies that  $\Gamma_x\sim a^3$  for LP and  $\Gamma_x\sim a^6$  for LCP incidences. Figure 4(c) and Fig. 4(d) show, respectively, the plots of  $\Gamma_y/\eta^5$  and  $\Gamma_y/\eta^3$  versus  $\log_{10}\eta$  for the TE and other polarization modes studied. From Fig. 4(c),  $\Gamma_y\sim a^5$  law is observed for the TE case, while Fig. 4(d) indicates that  $\Gamma_y$  obeys  $\Gamma_y\sim a^3$  for all other incident polarization modes. In addition, although the  $\Gamma_y\sim a^3$  law is violated for the non-matching case at large  $\eta$ , our calculation shows that it still serves as a good approximation for the matching case  $\epsilon_r=1$  even for relatively large  $\eta$ , as can be seen from the dotted line in Fig. 4(d).

## B. Normal illumination

Next, we turn to the normal illumination with incident  $\mathbf{k}_0$  perpendicular to the EA, which has been intensively studied

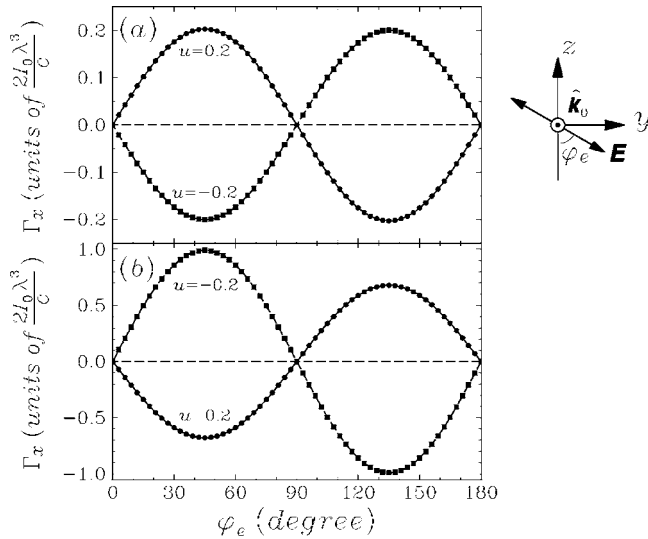


FIG. 5. Torque as a function of angle  $\varphi_e$  between the extraordinary axis and incident  $E$  field for normal illumination and  $u = \pm 0.2$  at  $\eta=5$  (a) and  $\eta=15$  (b). The circles and squares are calculated results, while the lines represent  $\Gamma_x = \Gamma_0 \sin 2\varphi_e$ . Also shown is the schematic plot for the directions of incident wave vector and  $E$  field.

experimentally [12–19]. To be specific, the incident wave is characterized by (22) with  $\theta_k = \pi/2$  and  $\phi_k = 0$ , namely,  $\mathbf{k}_0$  is in the  $x$  direction while keeping the EA in the  $z$  direction. For normal illumination, both the TE and the TM incident modes yield vanishing torque. We therefore consider two cases. The first case is the LP mode with the polarization vector given by  $\hat{\mathbf{p}} = \cos \varphi_e \hat{\boldsymbol{\theta}}_k + \sin \varphi_e \hat{\boldsymbol{\phi}}_k = -\cos \varphi_e \mathbf{e}_z + \sin \varphi_e \mathbf{e}_y$ , so that the angle between the EA and the incident  $E$  field is  $\varphi_e$ , as shown in Fig. 5. The TM and TE modes correspond to  $\varphi_e = 0$  and  $\varphi_e = \pi/2$ , respectively. The second case is the LCP incident wave with the polarization vector  $\hat{\mathbf{p}} = (1/\sqrt{2})(\hat{\boldsymbol{\theta}}_k + i\hat{\boldsymbol{\phi}}_k)$ .

Figure 5 shows the torque as a function of  $\varphi_e$  at  $\eta=5$  and 15 for the LP incident wave. The circles and squares denote the calculated optical torque, while the lines represent  $\Gamma_0 \sin 2\varphi_e$  curves, with  $\Gamma_0$  the maximum magnitude of the torque. The sine dependence of torque on  $2\varphi_e$  is found to hold up to numerical accuracy, with maximum magnitude of torque occurring at  $\varphi_e = \pi/4$ , as Eq. (2) proposed in Ref. [18] [see also the first term in (1)]. Our numerical results suggests that the  $\Gamma_0 \sin 2\varphi_e$  law is valid for arbitrary size parameter  $\eta$  and material birefringence, provided that  $\mathbf{k}_0$  is perpendicular to the EA. In addition, it is noted that when  $\eta$  is big enough, the EA of particle with  $n_e > n_o$  ( $u > 0$ ) is aligned with the direction normal to the incident  $E$  field, in contrast with the case of small  $\eta$  experimentally studied [18]. This is clearly seen from Fig. 5. When  $u=0.2$ , the maximum magnitude of the torque  $\Gamma_0 > 0$  for  $\eta=5$ , implying that the particle will rotate in the positive- $x$  direction. The stable equilibrium is achieved at  $\varphi_e = 0$  or  $\pi$ , resulting in the alignment of the EA with the  $E$  field for  $n_e > n_o$ . For  $u=-0.2$ ,  $\Gamma_0 < 0$ , a similar analysis leads to the conclusion that the EA is reoriented toward the direction normal to the  $E$  field for  $n_e < n_o$ . The situation becomes contrary for the case with  $\eta=15$ , as shown

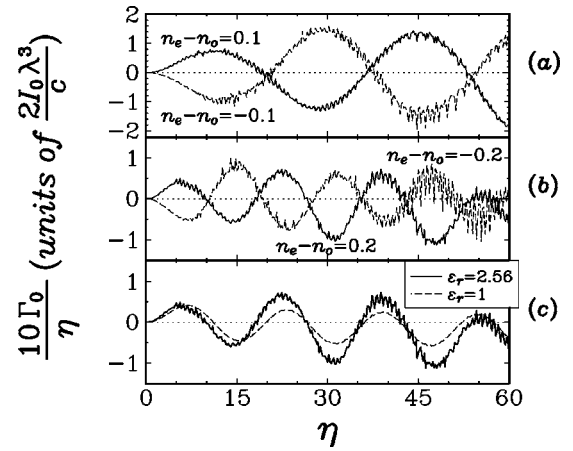


FIG. 6.  $10\Gamma_0/\eta$  versus  $\eta$  for the LP normal illumination at (a)  $n_e - n_o = \pm 0.1$ ; (b)  $n_e - n_o = \pm 0.2$ ; and (c)  $n_e - n_o = 0.2$  with  $\epsilon_r = 1$  and 2.56.

in Fig. 5(b), where  $\varphi_e = 0$  and  $\pi$  correspond to the stable equilibrium for  $n_e < n_o$ . For  $n_e > n_o$ ,  $\varphi_e = \pi/2$  is the stable equilibrium, implying that the optical torque tends to make the EA normal to the incident  $E$  field.

Figure 6 shows  $\Gamma_0$  for the LP incidence as a function of the size parameter  $\eta$ . It is found that, as the particle size increases, the torque will change its sign, implying that the particle with  $n_e > n_o$  is not always aligned with the  $E$  field. In addition,  $\Gamma_0/\eta$  shows an overall wavelike behavior between positive and negative values, somewhat resembling the first term in (1) based on ray optics for a simple, ideally flat disk. In most cases,  $\Gamma_0$  has different signs for  $n_e - n_o = \pm |\Delta n|$ , indicating that particles with  $n_e - n_o = +|\Delta n|$  and  $n_e - n_o = -|\Delta n|$  are subject to torque in opposite directions, which is in agreement with the first term in (1). However, it happens at some peculiar ranges of  $\eta$  that the torque may have the same sign for both  $n_e - n_o = \pm |\Delta n|$ , as shown, e.g., in Fig. 6(a) for  $|\Delta n| = 0.1$  at  $\eta \sim 20$  and  $\eta \sim 37$ . This suggests that the EA may be aligned with the directions parallel or normal to the  $E$  field for both  $n_e - n_o = \pm |\Delta n|$ . The situation presents a striking contrast to the first term in (1), since the latter implies that the particles with  $n_e - n_o = \pm |\Delta n|$  are always subject to torque in opposite directions. The difference comes from the effect of particle shape. Finally, it is also noted that the fluctuating behavior in  $\Gamma_x$  versus  $\eta$  curve is due to nonmatching  $\epsilon_r \neq 1$ . For the matching case  $\epsilon_r = 1$ , the wavelike behavior becomes quite smooth, provided that  $|\Delta n|$  is not very large, as can be seen in Fig. 6(c).

For the LCP incident wave, Fig. 7 shows plots of  $\Gamma_x/\eta^2$  versus  $\eta$  for different material birefringences. Qualitative agreement is observed with the second term in (1). The difference lies in that the torque will never vanish except in the limit  $\eta \rightarrow 0$ , and the amplitude of the wavelike behavior decays with  $\eta$  at relative small  $\eta$ . The overall  $\eta$  dependence is similar to the LP case, except that  $\Gamma_x$  is always positive instead of fluctuating between positive and negative values in the LP case. This is because the LCP wave always transports positive angular momentum to the scatterer, while the LP wave can transfer both positive and negative angular momentum to the scatterer. In contrast to the LP case shown in



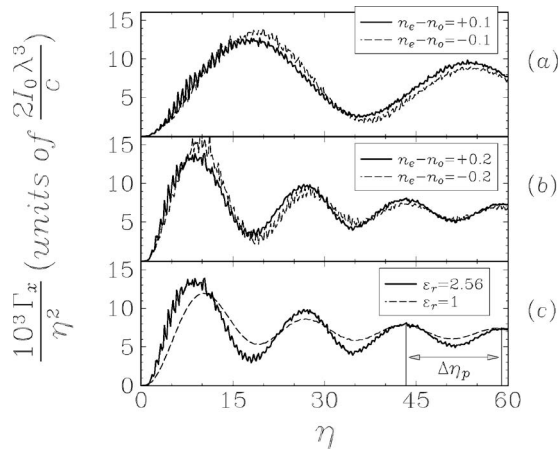


FIG. 7.  $10^3 \Gamma_x / \eta^2$  versus  $\eta$  for the LCP normal illumination at (a)  $n_e - n_o = \pm 0.1$ ; (b)  $n_e - n_o = \pm 0.2$ ; and (c)  $n_e - n_o = 0.2$  with  $\epsilon_r = 1$  and 2.56.

Fig. 6(a) and 6(b), for the LCP case,  $n_e - n_o = \pm |\Delta n|$  yield quite similar results. And also, as in the LP case, the zigzag behavior is due to  $\epsilon_r \neq 1$ . When  $\epsilon_r = 1$ , smooth wavelike dependence is observed, as shown in Fig. 7(c).

The wavelike behavior of  $\Gamma_x / \eta^2$  versus  $\eta$  curve for the LCP incidence leads to the appearance of peaks as  $\eta$  increases, as shown in Fig. 7 and Fig. 8(a). The peak separation  $\Delta \eta_p$  is dependent on  $|n_e - n_o|$ . In particular, it is found to fluctuate within 15% as  $\eta$  increases for all cases studied, as exemplified in Fig. 8(a). No decaying in amplitude of the wavelike behavior is found here for sufficiently large  $\eta$ . Figure 8(b) is a plot of peak separation  $\Delta \eta_p$  as a function of  $1/|n_e - n_o|$  for both  $n_e > n_o$  and  $n_o > n_e$ . The error bars denote the upper and lower bounds of  $\Delta \eta_p$  as  $\eta$  goes from 40 to 140. It is found that linear relation still provides a pretty fair approximation for  $\Delta \eta_p$  versus  $1/|n_e - n_o|$  irrespective of whether  $n_e > n_o$  or  $n_e < n_o$ , in agreement with the second term in (1) as well as the experimental result [16]. The wavelike behavior is thus verified in the transfer of the angular momentum from the incident light to the scatterer.

To study dependence of the radiation torque on particle size in the small particle limit, Fig. 9 shows the plots of  $\Gamma_0 / \eta^3$  and  $\Gamma_x / \eta^6$  versus  $\log_{10} \eta$  for, respectively, the LP and LCP incident waves. An incident polarization-dependent scaling law is clearly exhibited: the torque due to the LP incidence obeys the scaling law  $\Gamma_0 \sim a^3$ , as exhibited in Fig. 9(a) for some typical cases with  $u = \pm 0.2$  and  $\pm 0.4$ , while for the LCP incidence, the torque satisfies  $\Gamma_x \sim a^6$ , as shown typically in Fig. 9(b) for  $u = \pm 0.2$  and  $\pm 0.4$ .

Figure 10 displays the material birefringence dependence of the optical torque in the case of normal illumination for both the LP and the LCP incident modes. The change of the material birefringence is due solely to  $u$  while  $\epsilon_r$  is kept at 2.56. For the LP incidence, we present results for  $\Gamma_x$  at  $\varphi_e = \pi/4$ , namely,  $\Gamma_0$ . The  $\Gamma / \eta^2$  versus  $u$ , instead of  $\Gamma$  versus  $u$ , is displayed for the convenience of plotting curves for different values of  $\eta$  in the same figure. It is seen that there exists a monotonic increase region at small (absolute) value of  $u$ , in which the torque increases with the material birefringence, consistent with the experiment [16]. As  $|u|$  increases, how-

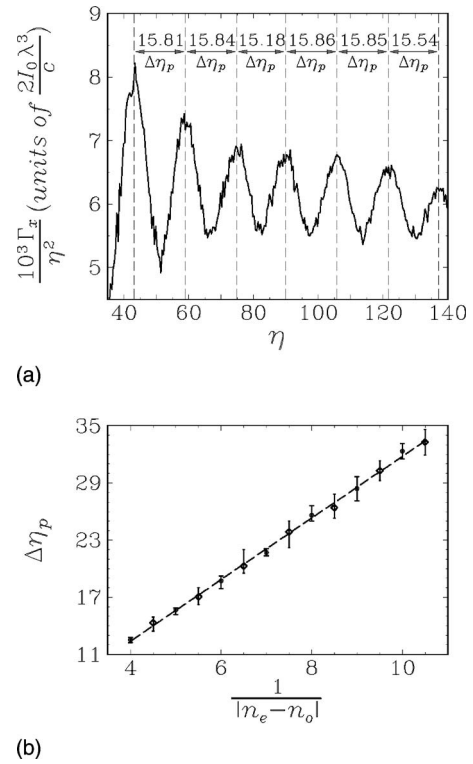


FIG. 8. (a)  $10^3 \Gamma_x / \eta^2$  versus  $\eta$  for the LCP normal illumination, showing that the peak separation  $\Delta \eta_p$  fluctuates within only a few percent as  $\eta$  increases. (b) Peak separation  $\Delta \eta_p$  as a function of  $1/|n_e - n_o|$  at  $\epsilon_r = 2.56$ . The error bars show the upper and lower bounds of  $\Delta \eta_p$  as  $\eta$  goes from 40 to 140. The change of the material birefringence is due solely to the change in value of  $u$ . Filled circles are obtained for the case with  $n_e > n_o$  and open diamonds for  $n_e < n_o$ . The dashed line is a guide to the eye showing the linear relation of  $\Delta \eta_p$  versus  $1/|n_e - n_o|$ .

ever, different oscillatory behaviors show up for different values of  $\eta$ , leading to appearance of many peaks in the range of  $u$  studied. As  $\eta$  increases, the monotonic increase region near  $u=0$  shrinks, in qualitative agreement with (1). In addition, it is seen that  $\Gamma_x$  is always positive for the LCP incident wave. For the LP case, the torque may change its sign as the material birefringence increases, indicating that the EA of a birefringent particle with  $n_e < n_o$  can also be oriented toward the  $E$ -field direction, even at relative small  $\eta$ , provided that the material birefringence is strong enough.

Figure 11 shows the typical torque versus the material birefringence behavior for small  $|\Delta n|$ . The torque due to the LP incident wave is found to display a linear dependence on  $|\Delta n|$ , while the torque by the CP incidence exhibits square law behavior  $\Gamma \sim |\Delta n|^2$ , both independent of the particle size. This dependence of  $\Gamma$  on  $\Delta n$  is in unexpected agreement with Eq. (1), although the latter is derived based on the ray optics for a flat disk.

Finally, in Fig. 12 the complex oscillatory behavior of  $\Gamma_0$  versus  $\epsilon_r = \epsilon_s / \epsilon_0$  is shown for the LP incidence at  $u = \pm 0.2$  and different values of  $\eta$ . The appearance of sharp peaks recalls the Mie-type resonances. For small  $\eta$ , in the range of  $\epsilon_s$  shown, it is seen that the torque does not change its sign. For large  $\eta$ , the torque may change its sign as  $\epsilon_r$  increases.

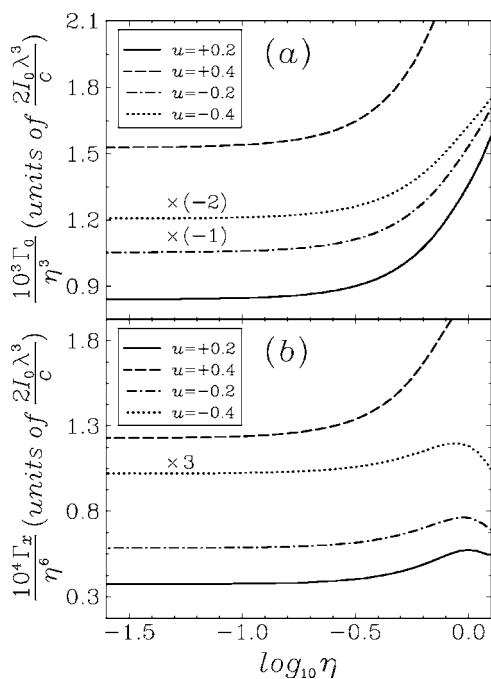


FIG. 9. (a)  $10^3 \Gamma_0 / \eta^3$  versus  $\log_{10} \eta$  at different values of  $u$  for the LP incidence; (b)  $10^4 \Gamma_x / \eta^6$  versus  $\log_{10} \eta$  at different values of  $u$  for the LCP incidence. To fit all curves for different  $u$  in the same figure, results for the LP cases with  $u = -0.2$  and  $u = -0.4$  have been multiplied by  $-1$  and  $-\frac{1}{2}$ , respectively, while the result for the LCP case with  $u = -0.4$  has been multiplied by  $\frac{1}{3}$ .

#### IV. SUMMARY

In summary, we have performed an exact *ab initio* calculation of the optical torque on a spherical birefringent particle of arbitrary size illuminated by plane electromagnetic wave with arbitrary polarization mode and direction of propagation, based on the extended Mie theory and the Maxwell stress tensor formalism. The expression for evaluating the radiation torque was derived for arbitrary (absorbing and lossless) isotropic surrounding medium. The radiation torque is found to exhibit miscellaneous dependences on the incident angle, the incident polarization, the material birefringence, as well as the particle size. The numerical results confirm the wave-plate mechanism in the transfer of angular momentum from the incident light to the scatterer. When a particle with  $n_e > n_o$  ( $n_e < n_o$ ) is subject to oblique illumination by a CP incident wave, its EA will always be aligned with the direction parallel (normal) to the  $E$  plane. An obliquely incident TM wave will reorient the EA of the scattering particle with  $n_e > n_o$  ( $n_e < n_o$ ) toward the direction parallel (normal) the incident  $E$  field irrespective of the particle size. When a particle is subject to normal illumination by a LP incident wave, the angle dependence (2), which is derived in the small particle limit, is found to hold for arbitrary particle size. The increase of the particle size affects only the value of  $\Gamma_0$ , making it vary between positive and negative numbers, indicating that even for a particle with  $n_e > n_o$  ( $n_e < n_o$ ), its EA is not always aligned with the direction parallel (normal) to the  $E$  field by the optical torque. Owing to the particle shape effect, in some peculiar ranges of  $\eta$ , the EA

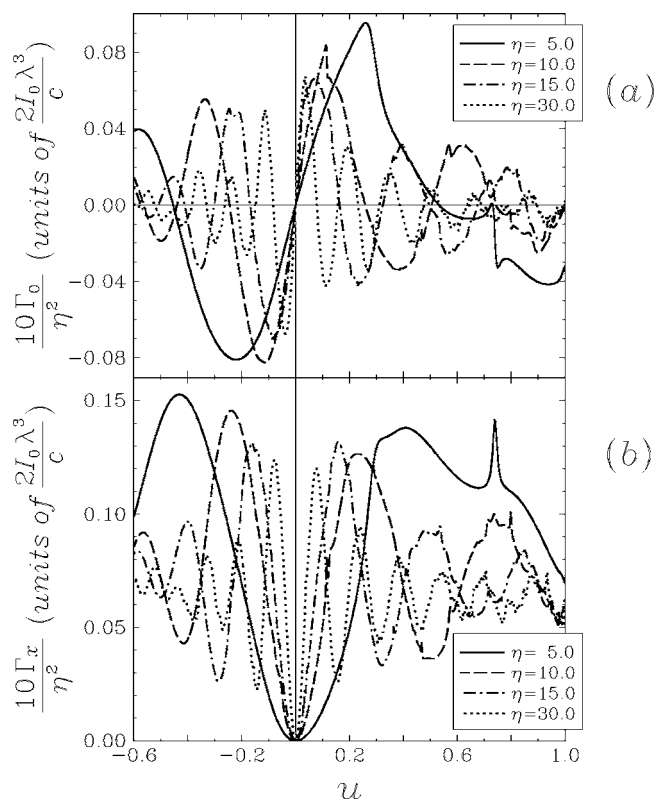


FIG. 10. The optical torque versus  $u$  at different values of  $\eta$  for the LP (a) and the LCP (b) normal illumination.

may be aligned with the  $E$ -field direction for both  $n_e - n_o = |\Delta n|$  and  $n_e - n_o = -|\Delta n|$ , and vice versa, at some other ranges of  $\eta$ , the torque may cause a reorientation of the EA toward the direction normal to the  $E$  field, irrespective of  $n_e - n_o = |\Delta n|$  or  $n_e - n_o = -|\Delta n|$ . In addition, dependent on the incident polarization mode as well as the incident direction, the radiation torque  $\Gamma$  exhibits different power-law dependences on the particle size in the small particle limit. To be specific, set the EA of the scattering particle to be the  $z$  axis and let the incident  $\mathbf{k}_0$  be in  $x$ - $z$  plane. For oblique illumination,  $\Gamma_x$  is found to satisfy  $\Gamma_x \sim a^6$  and  $\Gamma_x \sim a^3$  for the CP and the LP incidences, respectively, while  $\Gamma_y$  obeys  $\Gamma_y \sim a^3$  for all incident polarization modes except the TE case, where  $\Gamma_y$  displays  $\Gamma_y \sim a^5$  law, besides being much smaller than in other cases. For normal illumination, while  $\Gamma_y$  vanishes for any polarization mode,  $\Gamma_x$  behaves as in the case of oblique illumination, namely,  $\Gamma \sim a^\alpha$ , with the exponent  $\alpha = 3$  for the LP incidence and  $\alpha = 6$  for the CP incidence. Finally, when a particle with small material birefringence  $\Delta n$  is subject to normal LP and CP illuminations, the optical torque versus  $\Delta n$  displays linear and square law behaviors, respectively, regardless of the particle size.

Our direct classical calculation is merited due to subtleties of both classical and quantum theories of electromagnetic angular momentum. Furthermore, it is believed to be relevant to many applications where birefringent spherical particles are implemented, typically in the case of making measurements of viscosity on a microscopic scale [14]. We note that, in reality, the experiments on optical torque are performed using a laser beam. In most cases, however, the re-

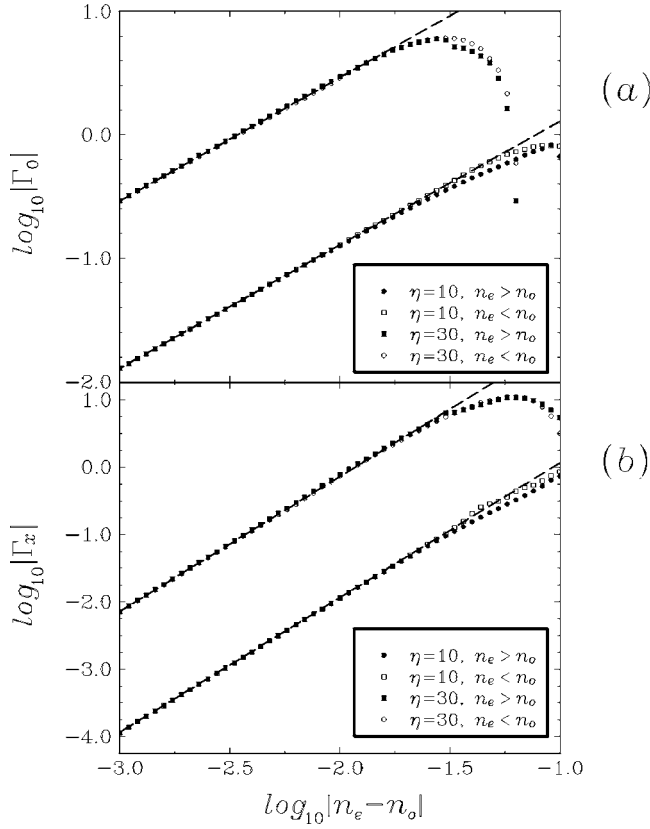


FIG. 11. The log-log plot of the optical torque versus the material birefringence  $|\Delta n|$  at  $\eta=10$  and  $30$  for the LP (a) and the LCP (b) normal illumination. The dashed lines are guides to the eye, showing  $|\Gamma| \sim |\Delta n|$  and  $|\Gamma| \sim |\Delta n|^2$  for the LP and the CP incident waves, respectively.

sults can be understood using the plane-wave picture [12], especially when the scattering particles are small compared with the beamwidth and located near the beam axis. In addition, our formulation provides a limiting case against which the solution of more complicated problems (e.g., torque on a particle illuminated by strongly focused laser beam) can be checked. Work along this line is in progress.

#### ACKNOWLEDGMENTS

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#### APPENDIX: SOME INTEGRALS INVOLVING $\pi_{mn}$ AND $\tau_{mn}$

Two auxiliary functions,  $\pi_{mn}(\cos \theta)$  and  $\tau_{mn}(\cos \theta)$ , are defined by

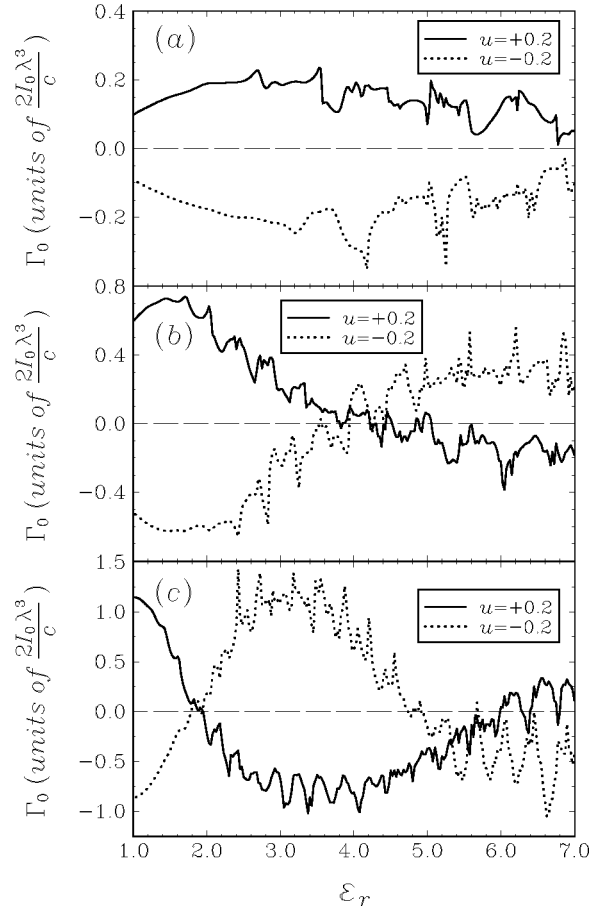


FIG. 12.  $\Gamma_0$  versus  $\epsilon_r = \epsilon_s / \epsilon_0$  for LP normal illumination at  $u = \pm 0.2$  and  $\eta=5$  (a),  $\eta=10$  (b), and  $\eta=15$  (c).

$$\pi_{mn}(\cos \theta) = \frac{m}{\sin \theta} P_n^m(\cos \theta),$$

$$\tau_{mn}(\cos \theta) = \frac{d}{d\theta} P_n^m(\cos \theta), \quad (\text{A1})$$

where the first kind of associated Legendre function  $P_n^m(x)$  is given by [22–24]

$$P_n^m(x) = \frac{1}{2^n n!} (1-x^2)^{m/2} \frac{d^{n+m}}{dx^{n+m}} [(x^2-1)^n]. \quad (\text{A2})$$

The integrals  $\mathcal{W}_{\pi p}^{(x,y,z)}$  and  $\mathcal{W}_{\tau p}^{(x,y,z)}$  are

$$\begin{aligned} \mathcal{W}_{\pi p}^{(x)} &= \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int [i\pi_{mn}(\cos \theta) P_v^\mu(\cos \theta) \cos \theta \cos \phi + \tau_{mn}(\cos \theta) P_v^\mu(\cos \theta) \sin \phi] e^{i(m-u)\phi} d\Omega \\ &= i \frac{2\pi}{|\eta|^2 \eta^*} |E_0|^2 \{ [(n-m)(v+u)]^{1/2} \delta_{v,n} \delta_{u,m+1} + [(n+m)(v-u)]^{1/2} \delta_{v,n} \delta_{u,m-1} \}, \end{aligned}$$

$$\begin{aligned}\mathcal{W}_{\pi p}^{(y)} &= \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int [i\pi_{mn}(\cos\theta)P_v^\mu(\cos\theta)\cos\theta\sin\phi - \tau_{mn}(\cos\theta)P_v^\mu(\cos\theta)\cos\phi] e^{i(m-u)\phi} d\Omega \\ &= + \frac{2\pi}{|\eta|^2 \eta^*} |E_0|^2 \{ [(n-m)(v+u)]^{1/2} \delta_{v,n} \delta_{u,m+1} - [(n+m)(v-u)]^{1/2} \delta_{v,n} \delta_{u,m-1} \},\end{aligned}$$

$$\mathcal{W}_{\pi p}^{(z)} = - \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int i\pi_{mn}(\cos\theta)P_v^\mu(\cos\theta) e^{i(m-u)\phi} \sin\theta d\Omega = -i \frac{4m\pi}{|\eta|^2 \eta^*} |E_0|^2 \delta_{v,n} \delta_{u,m},$$

$$\begin{aligned}\mathcal{W}_{\pi p}^{(x)} &= \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int [\tau_{mn}(\cos\theta)P_v^\mu(\cos\theta)\cos\theta\cos\phi - i\pi_{mn}(\cos\theta)P_v^\mu(\cos\theta)\sin\phi] e^{i(m-u)\phi} d\Omega \\ &= i \frac{2\pi}{|\eta|^2 \eta^*} |E_0|^2 \left\{ \left[ \frac{(n^2-1)(n+m)(v+m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n-1} \delta_{u,m-1} - \left[ \frac{(v^2-1)(n-u)(v-u)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n+1} \delta_{u,m-1} \right. \\ &\quad \left. - \left[ \frac{(n^2-1)(n-m)(v-m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n-1} \delta_{u,m+1} + \left[ \frac{(v^2-1)(n+u)(v+u)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n+1} \delta_{u,m+1} \right\},\end{aligned}$$

$$\begin{aligned}\mathcal{W}_{\pi p}^{(y)} &= \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int [\tau_{mn}(\cos\theta)P_v^\mu(\cos\theta)\cos\theta\sin\phi + i\pi_{mn}(\cos\theta)P_v^\mu(\cos\theta)\cos\phi] e^{i(m-u)\phi} d\Omega \\ &= - \frac{2\pi}{|\eta|^2 \eta^*} |E_0|^2 \left\{ \left[ \frac{(n^2-1)(n+m)(v+m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n-1} \delta_{u,m-1} - \left[ \frac{(v^2-1)(n-u)(v-u)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n+1} \delta_{u,m-1} \right. \\ &\quad \left. + \left[ \frac{(n^2-1)(n-m)(v-m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n-1} \delta_{u,m+1} - \left[ \frac{(v^2-1)(n+u)(v+u)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n+1} \delta_{u,m+1} \right\},\end{aligned}$$

$$\begin{aligned}\mathcal{W}_{\pi p}^{(z)} &= - \frac{v(v+1)}{|\eta|^2 \eta^*} E_{mn} E_{\mu\nu}^* \int \int \tau_{mn}(\cos\theta)P_v^\mu(\cos\theta) e^{i(m-u)\phi} \sin\theta d\Omega \\ &= i \frac{4\pi}{|\eta|^2 \eta^*} |E_0|^2 \left\{ \left[ \frac{(n^2-1)(n-m)(n+m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n-1} \delta_{u,m} + \left[ \frac{(v^2-1)(v-m)(v+m)}{(2n+1)(2v+1)} \right]^{1/2} \delta_{v,n+1} \delta_{u,m} \right\}.\end{aligned}$$

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