Semiclassical analysis of real atomic spectra beyond Gutzwiller's approximation

Benoît Grémaud*

Laboratoire Kastler Brossel, Université Pierre et Marie Curie, Case 74, 4, place Jussieu, 75252 Paris Cedex 05, France (Received 24 January 2005; published 13 October 2005)

Real atomic systems, like the hydrogen atom in a magnetic field or the helium atom, whose classical dynamics are chaotic, generally present both discrete and continuous symmetries. In this paper, we explain how these properties must be taken into account in order to obtain the proper (i.e., symmetry projected) \hbar expansion of semiclassical expressions like the Gutzwiller trace formula. In the case of the hydrogen atom in a magnetic field, we shed light on the excellent agreement between present theory and exact quantum results.

DOI: 10.1103/PhysRevE.72.046208

In the studies of the quantum properties of systems whose classical counterparts depict chaotic behavior, semiclassical formulas are essential links between the two worlds, emphasized by Gutzwiller's work [1]. More specifically, starting from Feynman's path formulation of quantum mechanics, he has been able to express the quantum density of states as a sum over all (isolated) periodic orbits of the classical dynamics. This formula, and extensions of it, have been widely used to understand and obtain properties of the energy levels of many classically chaotic systems, among which is the hydrogen atom in a magnetic field [2,3], the helium atom [4–6], or billiards [7–10].

At the same time, because the trace formula (and its variations) as derived by Gutzwiller only contained the leading term of the asymptotic expansion of the quantum level density, the systematic expansion of the semiclassical propagator in powers of \hbar has been the purpose of several studies [9–12], but which focused on billiards, for which both classical and quantum properties are easier to calculate.

In a recent paper [13], general equations for efficient computation of \hbar corrections in semiclassical formulas for a chaotic system with smooth dynamics were presented, together with explicit calculations for the hydrogen atom in a magnetic field. However, only the two-dimensional case was considered, because for the three-dimensional (3D) case, discrete symmetries and centrifugal terms had to be taken into account. Actually, this situation occurs in almost all real atomic systems depicting a chaotic behavior (molecules, two electron atoms...), for which experimental data involve levels having well defined parity, total angular momentum, and, if relevant, exchange between particles. In particular, semiclassical estimations of experimental signals like photoionization cross sections are calculated with closed orbits with vanishing total angular momentum, whereas they usually involve P(L=1) quantum states, whose positions in energy are shifted with respect to S(L=0) states. Furthermore, in recent years, the development of the harmonic inversion method makes it possible to extract the relevant quantities (position of peaks, complex amplitudes) from both theoretical and experimental data with a much higher accuracy than with the conventional Fourier transform [14]. In particular, it becomes possible to measure the deviation of the exact quantum results from the semiclassical leading order predictions. Thus a detailed semiclassical analysis of experimental results, beyond the leading order in \hbar , requires the understanding and the calculation of corrections due to both the discrete symmetries and centrifugal terms. In addition, we would like to stress that even if the present analysis is made with the density of states, it can also be made with the quantum Green function, which leads to expressions and numerical computations of the first order \hbar corrections for physical quantities like the photoionization cross section [15,16], which could either be compared to available experimental data [17,18], or become a starting point for refined experimental tests of the quantum-classical correspondence in the chaotic regime.

PACS number(s): 05.45.Mt, 03.65.Sq, 32.30.-r

 \hbar corrections and discrete symmetries have already been discussed, but only for billiards [9,10,12], whereas in the case of systems with smooth dynamics a detailed study is still lacking. Also, centrifugal terms and/or rotational symmetries have been considered by many authors, but either in the case of integrable systems [19,20], or for values of the angular momentum comparable to the action of classical orbits [1,21,22]. From this point of view, the present study, which focuses on fixed values of the quantum angular momentum and the effect of the centrifugal terms on \hbar corrections for systems with smooth chaotic dynamics, goes beyond the preceding considerations. More precisely, in this paper, we explain how to take into account both discrete symmetries and centrifugal terms in order to obtain a full semiclassical description of the first order \hbar corrections for the 3D hydrogen atom in a magnetic field.

At first, in the case of a chaotic system, whose Hamiltonian $H=\mathbf{p}^2/2+V(\mathbf{q})$ is invariant under a group \mathcal{S} of discrete transformations σ , the leading order of semiclassical approximation for the trace of the Green function G(E)=1/(E-H), restricted to the mth irreducible representation is given by [23]

$$g_m^{\text{sc}}(E) = \frac{d_m}{i\hbar} \sum_{l} \frac{T_l}{|K_l|} \sum_{n} \chi_m(\sigma_l^n) g_{(l,n)}^{(0)}(E)$$
 (1)

with

$$g_{(l,n)}^{(0)}(E) = \frac{1}{|\det(A_l^n - 1)|^{1/2}} \exp\left[\frac{i}{\hbar} nS_l - in\mu_l \frac{\pi}{2}\right], \quad (2)$$

where the l sum is taken over all primitive (isolated) orbits which become periodic through the symmetry operation σ_l

^{*}Electronic address: Benoit.Gremaud@spectro.jussieu.fr

[i.e., final position (respectively, velocity) is mapped back to initial position (respectively, velocity) by σ_l]. $\chi_m(\sigma_l^n)$ is the character of σ_l^n in the mth irreducible representation of dimension d_m . S_l is the action of the orbit l, μ_l is the Maslov index, T_l is the "period," A_l^n represents the Poincaré surface-of-section map linearized around the orbit, and K_l is the subgroup of \mathcal{S} leaving each point of the orbit l invariant. Adding first order \hbar corrections, the preceding equation (1) becomes

$$g_{m}^{\text{sc}}(E) = \frac{d_{m}}{i\hbar} \sum_{l} \frac{T_{l}}{|K_{l}|} \sum_{n} \chi_{m}(\sigma_{l}^{n}) g_{(l,n)}^{(0)}(E) \left(1 + i\hbar \mathcal{C}_{(l,n)}^{\text{tr}}\right). \tag{3}$$

 $C_{l,n}^{tr}$ can be derived by a detailed analysis of the stationary phase approximations starting from the Feynman path integral, following the same steps as in Refs. [10,13] and reads as follows:

$$C_{l,n}^{\text{tr}} = C_{l,n}^{T \to E} + \frac{1}{nT_l} \int_0^{nT_l} dt_0 C_{l,n}(t_0), \tag{4}$$

where $C_{l,n}^{T\to E}$ arises from the time to energy domain transformation. $C_{l,n}(t_0)$ (see Ref. [13] for the expressions) involves the classical Green functions $\mathcal{G}_{l,n}(t,t')$, i.e., the solutions of the equations controlling the linear stability around the classical trajectory $\mathbf{q}_{l,n}^{\rm cl}(t)$:

$$\left(-\frac{d^2}{dt^2}\mathbb{I} - \frac{\partial^2 V}{\partial \mathbf{q} \partial \mathbf{q}} \left[\mathbf{q}_{l,n}^{\text{cl}}(t)\right]\right) \mathcal{G}_{l,n}(t,t') = \mathbb{I} \delta(t-t'). \tag{5}$$

The fact that the orbits are periodic after the symmetry transformation σ_l^n determines the boundary conditions that the classical Green functions $\mathcal{G}_{l,n}(t,t')$ must fulfill, namely,

$$\begin{cases} \sigma_{l}^{-n}\mathcal{G}_{l,n}(nT_{l},t') = \mathcal{G}_{l,n}(0,t') \\ \mathcal{P}_{t_{0}}\mathcal{G}_{l,n}(0,t') = 0 & \forall t' \in [0,nT_{l}], \\ \mathcal{Q}_{t_{0}}\sigma_{l}^{-n}\dot{\mathcal{G}}_{l,n}(nT_{l},t') = \mathcal{Q}_{t_{0}}\dot{\mathcal{G}}_{l,n}(0,t') \end{cases}$$
(6)

where \mathcal{P}_{t_0} is the projector along the "periodic" orbit at the position depicted by time t_0 and $\mathcal{Q}_{t_0}=1-\mathcal{P}_{t_0}$. Of course, for $\sigma_l=1$, one recovers the boundary conditions given in Ref. [13]. Finally, all technical steps of Ref. [13] leading to efficient computation of $\mathcal{G}_{l,n}(t,t')$ and \hbar corrections, that is, solutions of sets of first order differential equations, can easily be adapted to take into account these modified boundary conditions.

As a numerical example, we have considered the 2D hydrogen atom in a magnetic field, at scaled energy ϵ =-0.1 [2]. More precisely, we have computed the trace of the quantum Green function, using roughly 8000 states belonging to the *EEE* representation [24] of the group D_4 , corresponding to effective $1/\hbar$ values ranging from 0 to 124 (see Ref. [13] for further details). In that case, the periodic orbit 1234 [25,26] (see inset of the top of Fig. 1 for the trajectory in semiparabolic coordinates), being (globally) invariant under a rotation of angle $\pi/2$, gives rise to contributions in the semiclassical approximation of the trace at all multiples of $S_{\overline{1234}}/4$. In the same way, the periodic orbit 1243 (see middle inset of Fig. 1) being invariant under a rotation of angle π , contributions are

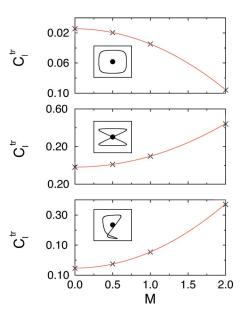


FIG. 1. (Color online) First order \hbar correction to the semiclassical approximation of the trace of the quantum Green function for the hydrogen atom in a magnetic field for different values of the magnetic number M, M=1/2 corresponding to the 2D case [13]. Crosses depict the values extracted from the exact quantum function using harmonic inversion, whereas the solid line corresponds to the classical results given by Eq. (10). For the three different periodic orbits, whose trajectories in the (u,v) plane are plotted (the nucleus being depicted by the black dot), the agreement is excellent, thus emphasizing the validity of Eqs. (9) and (10).

present at all multiples of $S_{\overline{1243}}/2$. For both these orbits, Table I displays the comparison of the present theoretical calculation and the numerical coefficient $C_{l,n}^{HI}$, extracted from the exact quantum Green function, using harmonic inversion [13,14]. As one can notice, the agreement is excellent for the amplitudes and rather good for the phases, which is the usual behavior of harmonic inversion. Furthermore, the same agreement has also been found for the other representations, thus emphasizing the present approach for the calculation of the first order \hbar corrections when taking into account discrete symmetries.

Contrary to the preceding, calculating first order \hbar corrections due to centrifugal terms is more complicated and is best explained in the case of the 3D hydrogen atom in a magnetic

TABLE I. Numerical comparison between the theoretical \hbar corrections $\mathcal{C}_l^{\mathrm{Ir}}$ for the trace of the quantum Green function, restricted to the *EEE* representation, of the 2D hydrogen atom in a magnetic field and the numerical coefficients \mathcal{C}_l^{HI} extracted from exact quantum function using harmonic inversion. The agreement is excellent for the amplitudes and rather good on the phases, thus emphasizing the validity of the present approach.

Code	$\mathcal{C}_l^{ ext{tr}}$	$ \mathcal{C}_l^{HI} $	Rel. error	arg \mathcal{C}_l^{HI}
$\frac{1}{4}$ 1234	-0.094 430	0.09445	$\approx 2 \times 10^{-4}$	$0.9996 \times \pi$
$\frac{1}{2}$ 1234	-0.361 689	0.3611	$\approx 2 \times 10^{-3}$	$0.996 \times \pi$
$\frac{3}{4}\overline{1234}$	-0.400 555	0.3992	$\approx 3 \times 10^{-3}$	$1.005 \times \pi$
$\frac{1}{2}\overline{1243}$	0.049 399	0.0493	$\approx 8 \times 10^{-4}$	$-0.075 \times \pi$

field. The regularized Hamiltonian in semiparabolic coordinates, for fixed value M of the projection of the angular momentum along the field axis, is given by [2]

$$H = -\frac{\hbar^2}{2} \left\{ \frac{\partial^2}{\partial u^2} + \frac{\partial^2}{\partial v^2} + \left(\frac{1}{4} - |M|^2 \right) \left[\frac{1}{u^2} + \frac{1}{v^2} \right] \right\}$$
$$-\epsilon (u^2 + v^2) + \frac{1}{8} u^2 v^2 (u^2 + v^2)$$
$$= H_0 + \frac{\hbar^2}{2} \left(|M|^2 - \frac{1}{4} \right) U(u, v). \tag{7}$$

 H_0 is then the Hamiltonian of the 2D hydrogen atom in a magnetic field. If U(u,v) was regular, then the additional first order \hbar correction for the orbit l would simply be

$$-\frac{1}{2}\left(|M|^2 - \frac{1}{4}\right) \int_0^{T_l} dt \ U(u_l(t), v_l(t)). \tag{8}$$

One must mention that in this case, the Langer transformation [27] of the coordinates $(u,v) \rightarrow (\exp(-x), \exp(-y))$ gives rise to a Hamiltonian which does not separate into kinetic and potential energies and for which no expressions for \hbar corrections are available.

On the other hand, the fact that U(u,v) is singular imposes boundary conditions on both classical and quantum dynamics. The classical trajectories have to make (smooth) bounces near u=0 and v=0 and for vanishing values of \hbar , we expect the trajectories of H to be those of H_0 , but mapped onto the reduced phase space (u>0,v>0), i.e., making hard bounces on the (u,v) axis. From the quantum point of view, depending on the parity of M, only wave functions belonging to given representations of D_4 are allowed. Thus first order \hbar corrections due to the singular part of the potential U, are given by the preceding considerations on the symmetries, whereas remaining corrections are given by Eq. (8), where U has to be replaced by a smooth counterpart, namely,

$$\widetilde{U} = \lim_{\epsilon \to 0^+} \frac{1}{2} \left(\frac{1}{(u + i\epsilon)^2} + \frac{1}{(u - i\epsilon)^2} + \frac{1}{(v + i\epsilon)^2} + \frac{1}{(v - i\epsilon)^2} \right). \tag{9}$$

Actually, one can show that the preceding equation gives the right answers for \hbar expansion of the propagator of the free particle (up to \hbar^3) and the harmonic oscillator (up to \hbar^2), for which analytical expressions for classical trajectories, classical Green functions, and quantum propagators exist (higher orders have not been checked yet). However, even if a detailed analysis of the derivation of the trace formula in presence of centrifugal terms seems to show that the preceding

approach works in general cases, rigorous proof of Eq. (9) is lacking.

Nevertheless, in the case of the 3D hydrogen atom in a magnetic field, we have compared the first order \hbar corrections, for different periodic orbits and for different values of the magnetic number M, with the present prediction, namely,

$$C_l^{\text{tr}}(M) = C_l^{\text{tr}}(2D) - \frac{1}{8}(4|M|^2 - 1) \int_0^{T_l} dt \tilde{U}(u_l(t), v_l(t)).$$
(10)

The results are displayed in Fig. 1 for M = 0,1,2 and for three different orbits, namely 1234, 1243, and 12343, whose trajectories in the (u,v) plane are plotted. The solid line is the theoretical result given by Eq. (10), whereas the crosses are the values extracted from the trace of the exact quantum Green function, using harmonic inversion (for scaled energy ϵ =-0.1, roughly 8000 effective $1/\hbar$ values ranging from 0 to 124). As one can notice the agreement is excellent, thus giving strong support for the validity of Eqs. (9) and (10). Furthermore, the simplicity of the replacement \tilde{U} may serve as a guideline for a rigorous treatment of the \hbar corrections arising from the centrifugal terms. In particular, the calculation of higher orders involves products of the derivatives of these centrifugal terms and those of the potential V_0 , giving rise to nontrivial mixing between centrifugal and standard \hbar corrections

In conclusion, we have presented a semiclassical analysis, beyond the usual Gutzwiller approximation, including first order \hbar corrections, of the quantum properties of real chaotic systems. More specifically, we have explained the additional corrections arising when taking into account both discrete symmetries and centrifugal terms. In the case of the (3D) hydrogen in a magnetic field, the agreement between the theory and the numerical data extracted from exact quantum results is excellent, emphasizing the validity of the analysis, especially of Eqs. (9) and (10).

Finally, since we know how to compute the \hbar corrections, it would be very interesting to work the other way around, that is, to perform the semiclassical quantization, thus getting \hbar corrections in the semiclassical estimations of the quantum quantities, like the eigenenergies. Of course, this represents a more considerable amount of work, since the $C_{l,n}^{tr}$ coefficients must be computed for all relevant orbits and then included in standard semiclassical quantization schemes, like the cycle expansion [5,11,28].

The author thanks D. Delande for his kind support during this work. Laboratoire Kastler Brossel is Laboratoire de l'Université Pierre et Marie Curie et de l'Ecole Normale Supérieure, unité mixte de recherche 8552 du CNRS.

^[1] M. C. Gutzwiller, *Chaos in Classical and Quantum Mechanics* (Springer, New York, 1990).

^[2] H. Friedrich and D. Wintgen, Phys. Rep. 183, 37 (1989).

^[3] D. Delande, in Chaos and Quantum Physics, edited by M.-J.

Giannoni, A. Voros, and J. Zinn-Justin, Les Houches Summer School, Session LII (North-Holland, Amsterdam, 1991).

^[4] G. S. Ezra, K. Richter, G. Tanner, and D. Wintgen, J. Phys. B 24, L413 (1991).

- [5] P. Gaspard and S. A. Rice, Phys. Rev. A 48, 54 (1993).
- [6] B. Grémaud and P. Gaspard, J. Phys. B 31, 1671 (1998).
- [7] L. A. Bunimovitch, Commun. Math. Phys. **65**, 295 (1979).
- [8] P. Cvitanović and B. Eckhardt, Phys. Rev. Lett. 63, 823 (1989).
- [9] P. Gaspard and D. Alonso, Phys. Rev. A 47, R3468 (1993).
- [10] P. Gaspard, D. Alonso, and I. Burghardt, Adv. Chem. Phys. 90, 105 (1995).
- [11] G. Vattay and P. E. Rosenqvist, Phys. Rev. Lett. 76, 335 (1996).
- [12] K. Weibert, J. Main, and G. Wunner, Eur. Phys. J. D 19, 379 (2002).
- [13] B. Grémaud, Phys. Rev. E 65, 056207 (2002).
- [14] J. Main, Phys. Rep. 316, 233 (1999).
- [15] E. P. Bogomolny, Zh. Eksp. Teor. Fiz. 96, 487 (1989) [Sov. Phys. JETP 69, 275 (1989)].
- [16] J. Gao and J. B. Delos, Phys. Rev. A 46, 1455 (1992).
- [17] A. Holle, J. Main, G. Wiebusch, H. Rottke, and K. H. Welge, Phys. Rev. Lett. **61**, 161 (1988).

- [18] J. Main, G. Wiebusch, K. Welge, J. Shaw, and J. B. Delos, Phys. Rev. A 49, 847 (1994).
- [19] L. S. Schulmann, *Techniques and Applications of Path Integration* (Wiley, New York, 1981).
- [20] K. C. Khandekar and S. V. Lawande, J. Math. Phys. 16, 384 (1975).
- [21] S. C. Creagh and R. G. Littlejohn, Phys. Rev. A 44, 836 (1991).
- [22] S. C. Creagh and R. G. Littlejohn, J. Phys. A 25, 1643 (1992).
- [23] J. M. Robbins, Phys. Rev. A 40, 2128 (1989).
- [24] C. C. Martens, R. L. Waterland, and W. P. Reinhardt, J. Chem. Phys. 90, 2328 (1989).
- [25] B. Eckhardt and D. Wintgen, J. Phys. B 23, 355 (1990).
- [26] K. T. Hansen, Phys. Rev. E 51, 1838 (1995).
- [27] R. E. Langer, Phys. Rev. 51, 669 (1937).
- [28] P. Cvitanović and B. Eckhardt, Phys. Rev. Lett. **63**, 823 (1989).