# Reciprocity relations between ordinary temperature and the Frieden-Soffer Fisher temperature

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Frieden and Soffer conjectured some years ago [Phys. Rev. E **52**, 2274 (1995)] the existence of a "Fisher temperature"  $T_F$  that would play, with regards to Fisher's information measure *I*, the same role that the ordinary temperature *T* plays in relation to Shannon's logarithmic measure. Here we exhibit the existence of reciprocity relations between  $T_F$  and *T* and provide an interpretation with reference to the meaning of  $T_F$  for the canonical ensemble.

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## I. WHY A FISHER TEMPERATURE?

Frieden and Soffer conjectured some years ago [1,2] the existence of a "Fisher temperature"  $T_F$  that would play, with regards to Fisher's information measure *I*, the same role that the ordinary temperature *T* plays in relation to Shannon's logarithmic measure *S* [3,4]. In a series of more recent publications, this conjecture was amply validated by showing that the Legendre transform structure of thermodynamics can be replicated without changes if ones substitutes *I* for the Shannon entropy *S* [5–8], which yields then a "Fisher thermodynamics."

This Fisher thermodynamics is exactly equivalent to the conventional one, except that instead of the Shannon-Boltzmann-Gibbs (SBG) entropy *S* one uses Fisher's I [5–8]. A question still lingers, though: we have a SBG pair (*S*,*T*) and a Fisher pair ( $I, T_F$ ). What is the relation between *T* and  $T_F$ ?

In other words, in parallel to (1/T)=dS/dU (*U* is the mean energy) [9], we have  $(1/T_F)=dI/dU$  [5–8]. We need a thermometer to measure  $T_F$ , and this is best achieved by finding a relationship between the two temperatures. In this Brief Report we purport to provide a first answer with respects to the relation between *T* and  $T_F$ .

#### **II. BRIEF FISHER CONSIDERATIONS**

Estimation theory [2] provides one with a powerful result with reference to a system that is specified by a physical parameter  $\theta$ . Let **x** be a stochastic variable and  $p_{\theta}(\mathbf{x})$  the probability density for this variable, which depends on the parameter  $\theta$ . If an observer (i) makes a measurement of **x** and wishes to best infer  $\theta$  from this measurement, calling the resulting estimate  $\tilde{\theta} = \tilde{\theta}(\mathbf{x})$  and (ii) wonders how well  $\theta$  can be determined, then estimation theory asserts [2] that the best possible estimator  $\tilde{\theta}(\mathbf{x})$ , after a very large number of **x** samples is examined, suffers a mean-square error  $e^2$  from  $\theta$ , which obeys a relationship involving Fisher's *I*—namely,  $Ie^2=1$ —where the Fisher information measure (FIM) *I* is of the form

$$I = \left\langle \left(\frac{\partial \ln p_{\theta}}{\partial \theta}\right)^2 \right\rangle. \tag{1}$$

The FIM is additive [2]. If we have *n* independent parameters  $\theta_i$ , Eq. (1) becomes a sum of *n* terms of the form given

above [2]. The "best" estimator is called the *efficient* estimator. Any other estimator must have a larger mean-square error. The only proviso to the above result is that all estimators be unbiased—i.e., satisfy  $\langle \tilde{\theta}(\mathbf{x}) \rangle = \theta$ . Thus, Fisher's information measure has a lower bound, in the sense that, no matter what parameter of the system we choose to measure, *I* has to be larger or equal than the inverse of the mean-square error associated with the concomitant experiment. This result, i.e.,

$$Ie^2 \ge 1$$
, (2)

is referred to as the Cramer-Rao bound and constitutes a very powerful statistical result [2].

### **III. FORMALISM**

We start by defining the well-known density operator that describes a system at equilibrium [3,4]:

$$\hat{\rho} = \frac{1}{Z} \exp\left(-\sum_{i=1}^{M} \chi_i \hat{A}_i\right). \tag{3}$$

The  $\chi_i$  are Lagrangian multipliers associated to the *M* observables  $\hat{A}_i$ , whose expectation values are given by

$$\langle \hat{A}_i \rangle = \operatorname{Tr} \hat{\rho} \hat{A}_i \quad (i = 1, \dots, M),$$
 (4)

where the partition function Z has the form  $Z(\chi_i) = \text{Tr}[\exp(-\sum_{i=1}^{M} \chi_i \hat{A}_i)]$  [10]. In our present considerations we assume that these multipliers *have already been determined*.

Following Mandelbrot [11–13] we (i) *associate* the above Lagrange multipliers to parameters to be estimated via *Fisher considerations* involving a FIM that depends upon  $\hat{\rho}$  and (ii) write down this FIM as a sum of *M* terms, each one associated to the estimation of the parameter  $\chi_i$ , i.e.,

$$I = \sum_{i=1}^{M} \Gamma_i \left\langle \left( \frac{\partial \ln \hat{\rho}}{\partial \chi_i} \right)^2 \right\rangle, \tag{5}$$

where the  $\Gamma_i$  are suitable constants related to the (conventional) wish of having a dimensionless *I*, as discussed in [14,15]. After replacing Eq. (3) into Eq. (5) we then find that *I* is intimately connected to our observables' fluctuations, as pointed out long ago by Mandelbrot [11,16]:

$$I = \sum_{i=1}^{M} \Gamma_i \langle (\hat{A}_i - \langle \hat{A}_i \rangle)^2 \rangle.$$
 (6)

If we wish to have a dimensionless I,  $\Gamma_i$  has the dimension of  $[1/\dim(\hat{A}_i)^2]$ . Now, it is well known (and straightforwardly verified) that the statistical fluctuations of an observable obey the relation [11,16]

$$\langle (\hat{A}_i - \langle \hat{A}_i \rangle)^2 \rangle = -\frac{\partial \langle A_i \rangle}{\partial \chi_i} \tag{7}$$

(the  $\chi_i$  have dimension of  $[1/\dim(\hat{A}_i)]$ , which allows us to recast the Fisher measure in the fashion

$$I = -\sum_{i=1}^{M} \Gamma_i \frac{\partial \langle \hat{A}_i \rangle}{\partial \chi_i}.$$
 (8)

### IV. EXTREMIZATION OF *I* SUBJECT TO CONSTRAINTS

As stated above, the thermodynamics Legendre structure can be neatly re-obtained if one extremizes FIM subject to constraints instead of doing this using the Boltzmann entropy [5–8]. We deal then with the same mean values  $\langle \hat{A}_i \rangle$  used above, but, of course, different Lagrange multipliers will ensue. Let us call these new Fisher multipliers  $\gamma_i$  and borrow from the well-known thermodynamic relation that links information measure, Lagrange multipliers (here the Fisher ones), and expectation values [3,5]:

$$\gamma_i = \frac{\partial I}{\partial \langle \hat{A}_i \rangle}.$$
 (9)

It is now clear that, introducing the above result into Eq. (8), we get an expression for the Fisher multipliers  $\gamma_i$  in terms of the Shannon ones  $(\chi_i)$ :

$$\gamma_i = -\sum_{j=1}^{M} \Gamma_j \frac{\partial}{\partial \langle \hat{A}_i \rangle} \frac{\partial \langle \hat{A}_j \rangle}{\partial \chi_j}; \quad \dim(\gamma_i) = \dim(1/\hat{A}_i), \quad (10)$$

a relation which could be used to determine them. It might seem at this point natural to ask what happens if we consider a canonical distribution in which the Lagrange multipliers are the  $\gamma_i$  instead of the  $\chi_i$ . We discuss this question below for classical systems within the strictures of the canonical ensemble.

### **V. EQUIPARTITION THEOREM**

In classical statistical mechanics there exists a useful general result concerning the energy *E* of a system expressed as a function of 2*N* generalized coordinates  $\xi_i$  (for instance, *N* coordinates  $r_i$  and *N* momenta  $p_i$ ). Thus,  $E = E(\xi_1, \ldots, \xi_{2N})$ . The result holds in the case of the following (frequent) occurrence.

(i) The energy splits additively into the form  $E = \epsilon_i(\xi_i) + E'(\xi_1, \dots, \xi_{i-1}, \xi_{i+1}, \dots, \xi_{2N})$ , where  $\epsilon_i(\xi_i)$  involves only the

variable ξ<sub>i</sub> and the remaining part E' does not depend on ξ<sub>i</sub>.
(ii) The function ε<sub>i</sub>(ξ<sub>i</sub>) is quadratic in ξ<sub>i</sub>.

In these circumstances  $\langle \epsilon_i \rangle = kT/2$ , with *k* Boltzmann's constant and *T* the temperature. This is the equipartition theorem [9]. The mean value of each independent quadratic term in the energy *E* equals kT/2, where  $\beta = 1/kT$  is the (Shannon-Boltzmann-Gibbs) Lagrange multiplier associated with the mean-energy constraint  $\langle E \rangle = \int d\tau f E$ . Its demonstration assumes that the thermal equilibrium Boltzmann-Gibbs equilibrium probability distribution

$$f = \frac{1}{Z}e^{-\beta E},\tag{11}$$

with  $d\tau$  the phase-space volume element. Setting  $\Gamma = 1/k^2 T_0^2$ , with  $T_0$  an arbitrary but fixed reference temperature, yields a dimensionless Fisher information measure (8) for the canonical ensemble:

$$I = -\frac{1}{k^2 T_0^2} \frac{\partial \langle E \rangle}{\partial \beta}.$$
 (12)

### VI. RECIPROCITY

Since we assume equipartition, we immediately find [9]

$$\langle E \rangle = N\beta^{-1}, \tag{13}$$

implying

$$\frac{\partial \langle E \rangle}{\partial \beta} = -N\beta^{-2} = -\frac{\langle E \rangle^2}{N},\tag{14}$$

entailing that, according to Eq. (10), the Fisher multiplier (defined as  $\gamma = 1/kT_F$ ) is

$$\gamma = \frac{1}{kT_F} = -\frac{1}{k^2 T_0^2} \frac{\partial}{\partial \langle E \rangle} \frac{\partial \langle E \rangle}{\partial \beta} = \frac{2}{k^2 T_0^2 \beta}.$$
 (15)

Since the multipliers are inverse temperatures, we obtain the interesting relationship

$$T_F = \frac{T_0^2}{2T},$$
 (16)

our main result here, which, on reflection, should not surprise anyone since it is a well-known fact that whenever I grows, Shannon's S decreases and vice versa [2]. Note that the Fisher information (12) adopts now the following appearance:

$$I = \frac{\langle E \rangle^2}{Nk^2 T_0^2}, \quad \frac{\partial I}{\partial \langle E \rangle} = \frac{2}{k^2 T_0^2 \beta}, \tag{17}$$

where we have used equipartition  $(\langle E \rangle = N/\beta)$ , leading to

$$\frac{1}{\beta} = \frac{k^2 T_0^2}{2} \frac{\partial I}{\partial \langle E \rangle}, \quad \langle E \rangle \equiv \langle E \rangle_{\beta}. \tag{18}$$

With reference to Eq. (15), let us introduce now the Fisher result  $\gamma = 2/(k^2 T_0^2 \beta)$  as the multiplier entering the canonical probability distribution *f* in Eq. (11) and repeat the preceding

discussion, starting with the relation that takes now the place of Eq. (13), here

$$\langle E \rangle_{\gamma} = N \gamma^{-1}. \tag{19}$$

One has

$$\frac{\partial \langle E \rangle_{\gamma}}{\partial \gamma} = -N\gamma^{-2} = -\frac{\langle E \rangle_{\gamma}^{2}}{N}.$$
 (20)

We ask ourselves what is now the new Fisher multiplier  $\gamma_2$ . The answer is, using Eqs. (19) and (15),

$$\gamma_{2} = -\frac{1}{k^{2}T_{0}^{2}} \frac{\partial}{\partial\langle E\rangle_{\gamma}} \frac{\partial\langle E\rangle_{\gamma}}{\partial\gamma} = \frac{2}{k^{2}T_{0}^{2}} \frac{\langle E\rangle_{\gamma}}{N} = \gamma \beta \frac{\langle E\rangle_{\gamma}}{N}, \quad (21)$$

i.e.,

$$\gamma_2 = \beta, \tag{22}$$

which is indeed consistent with Eq. (18). Here we encounter reciprocity. The "Fisher multiplier"  $\gamma_2$  is the inverse

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Boltzmann-Gibbs-Shannon temperature that verifies [cf. Eq. (17)]

$$\frac{1}{kT} = \beta = \frac{\partial I(\gamma)}{\partial \langle E \rangle_{\gamma}},$$
$$\frac{1}{kT_{F}} = \gamma = \frac{\partial I(\beta)}{\partial \langle E \rangle_{\beta}},$$
(23)

in self-explanatory notation. Equation (17) can be written in either the " $\beta$ " language or in the " $\gamma$ " one, indistinctly.

### VII. CONCLUDING REMARKS

We have in this Brief Report provided two results that we deem important for the Fisher practitioners: namely, (a)  $T_F = T_0^2/2T$ , with  $T_0$  an arbitrary but fixed reference Boltzmann temperature, and (b) the reciprocity relations given by Eq. (23).

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