# $CO<sub>2</sub>$  laser absorption measurements of temperature and density in the shock-induced Richimyer-Meshkov mixing zone

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The aim of this paper is to discuss the interest of using spectroscopic diagnostic methods to study shock-induced Richtmyer-Meshkov mixing as a complement to other experimental techniques, numerical simulations, and theoretical studies. Using a  $CO<sub>2</sub>$  laser absorption technique, the average temperatures and densities are experimentally determined through a high velocity gaseous mixing zone. This mixing is originated from a shock-induced acceleration of a plane interface, materialized by a thin plastic membrane, that initially separates two gases of different densities. Two gas combinations have been tested:  $CO_2$ -Ar and  $CO_2$ -He. Temperature and density are determined within the shock accelerated mixing zone from infrared absorption measurements of two characteristic vibrational-rotational lines of the  $CO<sub>2</sub>$  bending mode, using a  $CO<sub>2</sub>$  continuous wave laser as a diagnostic probe. The absorption coefficient is a function of the temperature and density of the  $CO<sub>2</sub>$  for a known linewidth profile. Thus, temperature and density profiles within the mixing region are calculated from two measurements. For this, a CO<sub>2</sub> absorption coefficient model, assuming the Voigt linewidth profile in the mixing region and taking into account the contribution of the hot bands, has been developed. It is shown that the technique is well applicable to this type of study, and that there is a strong correlation between the Atwood number and the shape of the mixing zone.

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# I. INTRODUCTION

#### A. Richtmyer-Meshkov instability

In the context of inertial confinement fusion (ICF) an experimental work is undertaken to analyze the impulsive Rayleigh-Taylor instability. This instability is one of the causes of the generation and development of a turbulent mixing zone, between the shell material and the thermonuclear combustible of an ICF pellet, which contributes to reducing the efficiency of the nuclear reaction by early breakup of the shell and cooling of the combustible.

The absorption method presented in this work has been developed for studying, in the simpler shock tube environment, the evolution of the mixing zone created by the interaction of a high Mach number shock wave and a gaseous interface initially at rest. When a shock wave accelerates the interface, this one is subjected to the impulsive Rayleigh-Taylor [1,2] instability, also called the Richtmyer-Meshkov [3,4] instability. If the membrane is initially planar, which is the case for the experiments presented in this paper, only the random small spatial scales created by the membrane breakup are excited and the turbulent mixing occurs soon.

In most of the experimental studies on the development of mixing regions due to Rayleigh-Taylor or Richtmyer-Meshkov instabilities, results are based on the determination of the mixing zone thickness, because it allows one to determine the turbulent energy developing within the mixing region. But, up to now, few measurement techniques have been developed to determine quantitatively the thermodynamic parameters in the mixing zone, such as temperature or density. Recently, Mikaelian [5] gave calculated turbulent energy profiles within shock accelerated interfaces, but, for this study, he supposes a linear density profile in the turbulent mixing region, and he suggests that more experimental results are needed.

Because of its spectroscopic properties in the infrared band,  $CO<sub>2</sub>$  has been used as one of the two constituting gases of the mixing zone. The other constituent is argon or helium because each presents no infrared emission or absorption in the domain of our experiments. Furthermore, as their densities are, respectively, very different and close to that of  $CO<sub>2</sub>$ , one can study the influence of the gas density gradients on the mixing.

#### B. Laser absorption principle of measurements

The laser absorption technique has been described by Wang [6] to measure temperature and concentration species through a radiating flow. He used an absorption coefficient corresponding to two or more monochromatic lines according to one or more absorbing species present in the burning media. He also shows that it is possible to determine temperature and concentration distributions in nonhomogeneous media with absorption measurements in different directions. Hanson and Falcone [7] use two vibrational-rotational lines of CO, centered on 4.7  $\mu$ m, to measure the temperature of a medium resulting from the combustion of a propane and air mixture at atmospheric pressure. Cheng and Bien [8] study the case of an axisymmetric medium composed of methane and burning carbon. They measure the absorption of two rotationalvibrational lines of CO at different parallel positions. But as the medium studied was strongly nonhomogeneous, they were obliged to realize an Abel inversion to obtain the temperature distribution of the mixing along the radius of the shock tube.

The same technique has been used by Rea and Hanson [9] to determine temperature in the postflame region corresponding to a hydrocarbon-air combustion. The authors give temperature profiles analogous to those that we have obtained within the mixing zone. We have adapted this technique for our study. However, the mixing region in our case is composed of  $CO<sub>2</sub>$ -He or  $CO<sub>2</sub>$ -Ar, which renders the problem easier because there is only one absorbing species. We have neglected the effects of the boundary layer, which is colder than the main flow, on the crossing of the laser beam. Furthermore, we have shown in Ref. [10] that when the initial interface is plane, the deformation of the mixing zone front is the lowest. Thus, the temperature and density obtained by integration over the laser beam trajectory are average values.

In this paper we present results obtained with a  $CO<sub>2</sub>$ laser absorption technique which has already been presented in Ref. [11] in its first approach. Now, a more rigorous absorption coefficient model is developed.

### II. EXPERIMENTAL SETUP AND INITIAL CONDITIONS

#### A. Experimental setup

Figure <sup>1</sup> represents the experimental setup for the absorption measurements. Experiments are carried out in a double diaphragm shock tube, of  $8.5 \times 8.5$  cm<sup>2</sup> cross section and length about 9 m. The test section is about 1.5 m long downstream from the second diaphragm. The gases, carbon dioxide upstream and helium or argon downstream, are initially separated by a thin plastic film  $(1.5 \mu m)$  which constitutes the second diaphragm. The driver gas is hydrogen. Before introducing the test gases, a vacuum of about  $10^{-3}$  Torr is obtained on both sides of the thin membrane, using a rotary pump followed by a diffusion pump. During the phases of pumping and introducing gases, the maximum amplitude of the initial deformation of the membrane is less than <sup>1</sup> mm. These two phases are the most delicate operations in the preparation of the experiments. Wall heat transfer gauges are used to determine the velocity of the shock wave, with an error on the Mach number less than 3%, while the velocity of the mixing zone is obtained, with the same accuracy, from two simultaneous  $CO<sub>2</sub>$  emission measurements at different abscissas, taken during the same absorption run. The trigger signal is given by the shock wave passage in front of one of the wall gauges, and then suitably delayed.

We use a  $CO<sub>2</sub>$  continuous wave laser beam, Model 81-5500-T, from California Laser, <sup>1</sup> mm diameter, 8.8 mrad aperture angle and 2 mm-diam spot at 40 cm distance, 7 W maximum power, 10.4  $\mu$ m centered wavelength, and 425±25 MHz width under standard operating conditions. Out of these standard operating conditions, which require 1-2 <sup>h</sup> continuous running in a stabilized temperature room before performing the experiments, one can witness an important broadening rate, up to 200 MHz/'C, and, as a consequence, an important decrease of the laser power for a given wavelength. This laser operates on the rotational-vibrational transition 00°1-10°0 of the CO<sub>2</sub> molecule and for different characteristic lines of this transition. Therefore, for each experiment, the laser emission line is controlled with both a powermeter and a spectrum analyzer.

Infrared absorption of  $CO<sub>2</sub>$  is collected with a liquid nitrogen cooled photovoltaic  $Cd_xH_{1-x}$ Te detector. The response time of the detector is less than  $1 \mu s$  and its passband is centered at 10.4  $\mu$ m. As indicated in Fig. 1, the initial laser beam intensity is strongly reduced by the ZnSe and Ge attenuators so that its power is less than 10 mW when it arrives on the sensitive infrared detector cell.

# B. Initial conditions

Measurements are essentially based on infrared absorption of shock heated  $CO<sub>2</sub>$ . The infrared technique requires strong shocks — Mach 4 in  $CO<sub>2</sub>$ , 3.7 and 2 in Ar and He, respectively —and therefore the velocities behind the shock, 700 m/s for the  $CO_2$ -Ar interface and 1060  $m/s$  for the  $CO<sub>2</sub>$ -He one, are into the compressible range. The initial pressure of the gases on both sides of the plastic membrane is the same, about 2000 Pa, in all the experiments, in order to avoid any large scale initial bulge [10]. The use of two gas pairs,  $CO_2$ -He and  $CO_2$ -Ar, allows to analyze the influence of the Atwood number  $(|A_t|_{\text{CO}_2 \to \text{He}} \approx 0.80$  and  $|A_t|_{\text{CO}_2 \to \text{Ar}} \approx 0.26$  on the subsequent turbulent mixing. The Atwood number is defined by

$$
A_t = \frac{\rho_{\text{(Ar or He)}} - \rho_{\text{CO}_2}}{\rho_{\text{(Ar or He)}} + \rho_{\text{CO}_2}}
$$

where  $\rho_i$  is the density of the gas i just after the shock passage. Average temperature and  $CO<sub>2</sub>$  density profiles have been obtained at the abscissa of 940 mm from the initial position of the interface, which corresponds to a time of 1340  $\mu$ s for the CO<sub>2</sub>-Ar mixing zone and 886  $\mu$ s for the  $CO<sub>2</sub>$ -He mixing zone.

# III. ANALYTICAL MODEL OF THE CO<sub>2</sub> ABSORPTION COEFFICIENT IN THE MIXING ZONE

The knowledge of the absorption coefficient of an optical active medium allows us to know the thermodynamic state of this medium. In the context of our work, we have to obtain some information on the thermodynamic state of a turbulent gaseous mixing zone  $(CO_2-He$  or  $CO<sub>2</sub>-Ar$ , moving at high velocities in a shock tube. For this, we need to know the absorption coefficient of the  $CO<sub>2</sub>$  molecule, in the gaseous mixings  $CO<sub>2</sub>$ -He or  $CO<sub>2</sub>$ -Ar for the transition  $\Sigma$  (00°1-10°0) when these media are traversed by a low power continuous  $CO<sub>2</sub>$  laser beam [12,13].

We use the semiclassical theory of wave-matter interaction. The absorption coefficient is the result of the contributions of the laser induced absorption and emission as well as the spontaneous emission. We have taken into account the spectral width of the laser line and the spectral width of the  $CO<sub>2</sub>$  absorption line in the test section. This  $CO<sub>2</sub>$  absorption line is a function of the Doppler width due to the thermal molecular motion, the collisional width dependent on the gas pressure and temperature in the considered medium, and its composition: that is, a  $CO<sub>2</sub>$  Voigt profile [14]. We also have taken into



Absorption signal of  $CO<sub>2</sub>$  -Ar mixing; Mach number 4



account the hot bands of the  $\Sigma$  (00°1-100°) transition which can give a non-negligible contribution to the absorption coefficient [15]. Thus, we have realized a theoretical model of the absorption coefficient of the  $CO<sub>2</sub>$ molecule in gaseous mixing media,  $CO_2$ -Ar or  $CO_2$ -He, which takes into account all the contributions previously described.

#### A. The absorption coefficient

Let us consider an absorbing medium with an elementary thickness dl, crossed by a low intensity radiation  $I_0$ 

FIG. 1. (a) Sketch of the experimental setup; (b) typical absorption signal record.

with a frequency  $v$ . The intensity of this radiation changes of a quantity  $dI$  which is proportional to the incident intensity as predicted by the Bourguer-Lambert  $[16]$  law:

$$
dI = -\alpha I \, dl \tag{1}
$$

where  $\alpha$  is a coefficient of proportionality called the absorption coefficient and  $I$  the intensity after absorption. The integration of Eq.  $(1)$  along the distance L gives the expression of the absorption coefficient  $\alpha$ :

r

$$
\alpha = \frac{1}{L} \ln \left( \frac{I_0}{I} \right) \,. \tag{2}
$$

Let us now consider a system of particles submitted to a radiation  $I_0$  with the frequency v. Let  $N_l$  be the number of molecules in the fundamental quantic state and  $N_u$ the number of molecules in the quantic excited state. Then,  $\rho_{v/u}$ , the density of the energy of the incident radiation, is [16,17]:

$$
\rho_{\nu l u} = \frac{8\pi h v_{l u}^3}{c^3} \frac{1}{e^{h v_{l u}/KT} - 1} \ . \tag{3}
$$

Assuming thermodynamical equilibrium, there is equality between induced absorption, spontaneous emission, and isotropic induced emission:

$$
N_{u}B_{u-l}\rho_{vlu} + N_{u}A_{u-l} = N_{l}B_{l-u}\rho_{vlu} \tag{4}
$$

With the classical hypothesis of a Boltzmann distribution function for the population of the transition levels [16]:

$$
\frac{N_l}{N_u} = g_l e^{-h v_{lu}/KT}
$$
\n(5)\n
$$
x = \sqrt{\ln 2} \frac{(\nu - \nu_0)}{v_{lu}} \qquad a = \ln 2 \frac{\Delta v_c}{}
$$

and

$$
g_l B_{l-u} = g_u B_{u-l} \tag{6}
$$

we obtain

$$
A_{u-l} = \frac{8\pi h v_{lu}^3}{c^3} B_{u-l} \t{,} \t(7)
$$

where  $N$  and  $g$  correspond, respectively, to the population and the statistic weight of each considered level; A and  $B$  are the Einstein coefficients;  $l$  and  $u$  the lower and upper level indices, respectively;  $h$  and  $K$  are the Planck and Boltzmann constants, respectively; and  $c$  the velocity of light. If we suppose that the intensity variation  $dI$  of the incident radiation is due to induced emission and absorption, then we obtain

$$
dI = \frac{h\,\nu}{c} (N_u B_{u-l} - N_l B_{l-u}) I \, dl \, , \qquad (8)
$$

and combining relations (2), (6), (7), and (8),

$$
\alpha = \frac{h v}{c} g_l B_{l-u} \left[ \frac{N_l}{g_l} - \frac{N_u}{g_u} \right]. \tag{9}
$$

In the semiclassical theory framework of the radiationmatter interaction,  $B_{u-l}$  is given by [16–19]

$$
B_{u-l} = \frac{8\pi^3}{3h^2} |R_{12}|^2 |R_{jj'}|^2 H f(\nu - \nu_0) , \qquad (10)
$$

where  $R_{12}$  is the matrix element term of the dipolar electrical moment depending on the change of the vibrational level,  $R_{ii'}$  the matrix element term of the dipolar electrical moment depending on the change of the rotational level given by [18,19]

$$
|R_{jj'}| = \frac{S_{j'l'}}{g_j} \text{ and } S_{j'l'} = \frac{m^2 - l^2}{m}, \qquad (11)
$$

and  $H=1-C_1 m - C_2 m^2$  is the term which takes into account the rotational-vibrational interaction ( $\sim$  1 for CO<sub>2</sub>),  $m=j$  or  $j+1$ ,  $C_1 \sim 10^{-3}$  and  $C_2 \sim (6\pm 1) \times 10^{-5}$  are two constants, and  $f(v-v_0)$  the spectral profile of the incident radiation absorption line.

Thus, Eq. (9) becomes

$$
\alpha = \frac{8\pi^3 \nu}{3hc} g_I |R_{JJ'}|^2 |R_{12}|^2 Hf(\nu - \nu_0) \left| \frac{N_I}{g_I} - \frac{N_u}{g_u} \right| \,. \tag{12}
$$

The spectral profile  $f(v-v_0)$  of the incident radiation is a Voigt profile. In the case of the  $CO_2$ -Ar or  $CO_2$ -He combination of gaseous mixings, it is given by [12,18]

$$
f_{\text{Voigt}}(\nu) = \frac{a}{\pi \Delta \nu_D} \left[ \frac{\ln 2}{\pi} \right]^{1/2} \int_{-\infty}^{+\infty} \frac{e^{-y^2}}{a^2 + (x - y)^2} dy , \qquad (13)
$$

with

$$
x = \sqrt{\ln 2} \frac{(v - v_0)}{\Delta v_D} , \quad a = \ln 2 \frac{\Delta v_c}{\Delta v_D} ,
$$

and where  $\Delta v_c$  is the collisional profile depending on the pressure and temperature of the gaseous mixing,  $\Delta v<sub>D</sub>$  is the Doppler profile which is characteristic of the thermal motion in the medium, and  $\Delta v_c$  and  $\Delta v_D$  are given by [18,19]

$$
\Delta v_c = CP\gamma_{\text{CO}_2\text{-CO}_2} \left[ \psi_{\text{CO}_2} \left( \frac{300}{T} \right)^{n_{\text{CO}_2\text{-CO}_2}} + \sum_i \psi_X \left( \frac{300}{T} \right)^{n_{\text{CO}_2\text{-}X}} \times \left( \frac{\gamma_{\text{CO}_2\text{-}X}^{(j)}}{\gamma_{\text{CO}_2\text{-}CO_2}} \right) \right], \tag{14}
$$

with

$$
C = 1.65 \times 10^{-4} \text{ m}^2 \text{K}^{1/2} \text{kg}^{-1/2}
$$

and

$$
\Delta v_D = 2v_0 \left[ \frac{2 \ln(2)KT}{Mc^2} \right]^{1/2},
$$
 (15)

where  $\Psi_{\text{CO}_2}$  is the mass concentration of  $\text{CO}_2$  in the mixing CO<sub>2</sub>-X (X=Ar, He),  $\Psi_X$  is the mass concentration of argon or helium in the mixing  $CO<sub>2</sub>$ -X, and

$$
M = \frac{M_{\text{CO}_2} M_x}{M_{\text{CO}_2} + M_x}
$$

### B. General expression of the  $CO<sub>2</sub>$  absorption coefficient taking into account the hot bands

For a given line of the  $CO<sub>2</sub>$  infrared spectrum, one can find adjacent lines in near resonance with it which contribute to the absorption of the  $CO<sub>2</sub>$  in this band in a non-negligible quantity  $( \leq 40\%)$ . These bands are called "hot bands." Therefore, we have determined the bands contributing to the total absorption of the  $CO<sub>2</sub>$  for the transition  $\Sigma$  (00°1-100°) and calculated the percentage of the population of the low level transitions centered on a frequency near 961 cm<sup>-1</sup> (10.4  $\mu$ m). Transitions whose low levels have a population that may be considered as low are neglected, and the others are what we call hot bands. We have considered all the hot bands in resonance with the absorption line likely to contribute to the absorption.

Figure 2(a) schematizes the configurations that we have

used in the present work, which are a laser beam of frequency  $v_0$  and width  $\Delta v_0$  and the two hot bands adjacent to the principal absorption band the spectral range of which is about of the order of the absorption line. Thus, the total absorption coefficient represents the sum of the contribution of the principal band and the contribution of the hot bands:

$$
\alpha = \alpha_0 + \sum_{i=1}^n \alpha_i \tag{16}
$$

where  $\alpha_0$  is the absorption coefficient of the principal band and  $\alpha_i$  the absorption coefficient of the *i*th hot band.

Figure 2(b) represents the profiles of the principal  $CO<sub>2</sub>$ band and the laser. Within the hypothesis of a Voigt profile [relation (13)] the resonant absorption coefficient of the centered line is

$$
\alpha_0 = \frac{1}{\Delta v_0} \int_{v_0 - \Delta v_0/2}^{v_0 + \Delta v_0/2} \alpha_0^0(\nu) f_0(\nu) d\nu , \qquad (17)
$$

where  $\alpha_0^0$  is given by relation [12].

Now, considering  $\Delta v_0/v_0$  as a small quantity, and [18,19]



FIG. 2. (a) Spectral profile of 10.4- $\mu$ m CO<sub>2</sub> laser and CO<sub>2</sub> absorption line and its hot bands; (b) case (1): spectral profile of 10.4  $\mu$ m  $CO<sub>2</sub>$  laser and  $CO<sub>2</sub>$  absorption line without hot bands, and case (2): spectral profile of 10.4  $\mu$ m  $CO<sub>2</sub>$  laser and  $CO<sub>2</sub>$  absorption line with its hot bands.

$$
\frac{a}{\pi} \int_{-\infty}^{+\infty} \frac{e^{-y^2}}{a^2 + (y)^2} dy = \exp(a^2) [1 - \text{erf}(a)] , \qquad (18)
$$

with

$$
\mathrm{erf}(x) = \frac{2}{\sqrt{\pi}} \int_0^x e^{-z^2} dz
$$

relation (17) becomes:

$$
\alpha_0(\nu) = \alpha_0^0(\nu) \left[ \frac{\ln 2}{\pi} \right]^{1/2} \exp(a^2) [1 - \text{erf}(a)] . \tag{19}
$$

The  $CO<sub>2</sub>$  hot band profiles and their positions relative to the center of the laser line are schematized in Fig. 2(b). As previously said, when taking into account the nonadjustment of the laser line and the *i*th hot band, we define a new profile:

$$
g_i(\nu) = f_{\text{Voigt},i}(\nu) f_L(\nu_0 + k_i + \nu_i i) , \qquad (20)
$$

where  $f_{\text{Voigt}, i}(\nu)$  is the Voigt profile for the hot band i and  $f_L(v_0+k_i+v_i)$  is the value of the laser line profile, assumed to be Lorentzian, on the point  $v=v_0+k_i+v$ ,  $v_0$ being the laser line frequency,  $v_i$  the *i*th hot band frequency,  $k_i$  the distance between them, and

$$
f_L(\nu) = \frac{1}{\pi} \frac{1}{1 + 4 \left[ \frac{k_i}{\Delta \nu_c} + \frac{x_i \Delta \nu_D}{\sqrt{\ln 2} \Delta \nu_c} \right]^2} \tag{21}
$$

Then, the absorption coefficient of the *i*th hot band is

$$
\alpha_i = \frac{1}{\Delta v_i} \int_{v_i - \frac{\Delta v_i}{2}}^{v_i + \frac{\Delta v_i}{2}} \alpha_i^0(\nu) g_i(\nu) d\nu , \qquad (22)
$$

where  $\alpha_i^0$  is defined by relation (12) and  $f_i(v)$  by relation (13).

With the hypothesis of a Voit profile, the absorption coefficient for the ith hot band becomes

$$
\alpha_i = \frac{2\Delta v_D \alpha_i^0(\nu)a}{\pi^{3/2} \Delta v_i} \int_{v_i - v_i/2}^{v_i + v_i/2} \int_{-\infty}^{+\infty} \left\{ \frac{e^{-y^2}}{a^2 + (x_i - y)^2} \right\} \frac{1}{1 + 4 \left\{ \frac{k_i}{\Delta v_c} + \frac{x_i \Delta v_D}{\sqrt{\ln 2} \Delta v_c} \right\}^2},
$$
\n(23)

with

$$
x_i = \frac{(\nu - \nu_i)}{\nu_D} \sqrt{\ln 2} , \quad \Delta x_i = \frac{\Delta \nu}{\nu_D} \sqrt{\ln 2}
$$

If we suppose that the normalized width is smaller than 1, we can develop  $\alpha_i$  around zero:

$$
\alpha_i(\Delta x_i) = \alpha_i(0) + (\Delta x_i)\alpha_i(0) + o(\Delta x_i^n) , \qquad (24)
$$

and considering  $\Gamma'(x_i)$ ,

$$
\Gamma'(x_i) = \frac{1}{1 + 4\left(\frac{k_i}{\Delta v_c} + \frac{x_i \Delta v_D}{\sqrt{\ln 2} \Delta v_c}\right)^2},
$$
\n(25)

with  $x_i$  small, the development of  $\Gamma'(x_i)$  near zero gives

$$
\Gamma'(x_i) = \frac{1}{1+4\left[\frac{k_i}{\Delta v_c}\right]^2} - 8\left[\frac{(k_i - v_i)\Delta v_c^2}{\Delta v_c^2 \sqrt{\ln 2}}\right] \times \frac{1}{\left[1+4\left[\frac{k_i}{\Delta v_c}\right]^2\right]^2},
$$

then the corresponding absorption coefficient  $\alpha_i$  can be written  $\Lambda$   $\sim$ 

$$
\alpha_i(\nu) = \alpha_i^0(\nu) \frac{\Delta \nu_i}{\Delta \nu_D} \frac{1}{\left[1 + 4\left[\frac{k_i}{\Delta \nu_c}\right]^2\right]}
$$

$$
\times \left(\frac{\ln 2}{\pi}\right)^{1/2} \exp(a^2) [1 - \text{erf}(a)]. \tag{27}
$$

Finally, the general expression of the absorption coefficient is

$$
\alpha_{\omega}(\rho, T) = \frac{8\pi_{j}^{3}v_{0}}{3KT} \left[ \frac{\ln 2}{\pi} \right]^{1/2} |R_{12}|^{2} |R_{jj'}|^{2} \rho_{\text{CO}_{2}} Q_{\nu}^{-1}
$$

$$
\times \exp(a^{2}) \left[ \frac{\sqrt{\pi/2}}{a + \{[(\pi - 2)a/2]^{q} + p^{2}/4\}^{1/q}} \right]
$$

$$
\times \sum_{i} \alpha(j_{i}, T) , \qquad (28)
$$

$$
(26) \qquad \text{with} \qquad
$$

$$
\alpha(j_i, T) = \left[ b_1 e^{-[b_1 j_i(j_i+1)hc]/KT} - b_2 e^{[-b_2(j_i+j')(j_i+j'+1)hc]/KT} \right] \left[ 1 + \frac{1}{\left[ 1 + 4\left[ \frac{k_i}{\Delta v_c} \right]^2 \right]} \frac{|R_{i,12}^{\prime 2}|}{|R_{12}^2|} \right]
$$

where  $j_i$  is the rotational number of the *i*th hot band,  $j' = 0$  for an R line,  $j' = -1$  for a P line,  $R'_{i,12}$  is the intensity of the ith hot band of the considered transition,  $p \approx 1.902$ , and, finally,  $i = 0$  corresponds to the principal band for which  $k_i = 0$ .

### IV. TEMPERATURE AND DENSITY CALCULATION

If we suppose that the rotational-vibrational lines of  $CO<sub>2</sub>$  transitional band 0°01 – 1°00 are completely separated and if the gap between two nearby lines is larger than 27 GHz and the spectral width of each absorption line smaller than 2 GHz (for a pressure range less than 5 atm), one can assume that the absorption coefficient for a given line is only a function of its spectral and hot band properties. Furthermore, as the pressure ratio between the  $CO<sub>2</sub>$ in the shock tube and the  $CO<sub>2</sub>$  in the laser cavity is less than 10, we can neglect  $[14]$  the shift between the absorption line of  $CO<sub>2</sub>$  in the shock tube and the laser emission line. On the other hand, emission measurements done with  $CO<sub>2</sub>$  at 10.6  $\mu$ m have shown that we can neglect the contribution of emission in that range of wavelengths.

The temperature and density obtained are average values since the laser beam crosses the shock tube section, and they are then integrated along a cylinder of 8.5 cm long and 2 mm in diameter because we suppose that the mixing has quickly reached its equilibrium temperature. The  $CO<sub>2</sub>$  absorption coefficient given by relation (27) is a function of the temperature and density of  $CO<sub>2</sub>$ within  $CO_2$ -Ar or  $CO_2$ -He mixing zones. Thus, for each line we can write

$$
\alpha = \alpha_{\omega}(T,\rho) \tag{29}
$$

So if we accurately know the  $CO<sub>2</sub>$  absorption coefficient  $\alpha_{\omega}(T,\rho)$  in the CO<sub>2</sub>-Ar or CO<sub>2</sub>-He mixing zones for two given lines, we can access to the values of  $T$  and  $\rho$  within these zones by solving the system of two equations:

$$
\alpha_1 = \alpha_{\omega_1}(T, \rho) ,
$$
  
\n
$$
\alpha_2 = \alpha_{\omega_2}(T, \rho) ,
$$
\n(30)

where  $\omega_1$  and  $\omega_2$  are the frequencies of the first and the second lines, respectively. The accuracy of the method depends on both the choice of the couple of gases which constitute the mixing and the absorption line. For the purpose of the present study, a selection was made of the best couple of frequency bands, as to the signal-to-noise ratio.

Absorption by the test gas  $CO<sub>2</sub>$  in the mixing zone of two continuous  $CO<sub>2</sub>$  laser lines, corresponding to transition between the vibrational levels  $00^{\circ}$ 1  $\rightarrow$  10°0 (about 10.4  $\mu$ m wavelength), yields average profiles of density, temperature, and  $CO<sub>2</sub>$  mass concentration in the mixing zone. Error bars have been estimated to be  $\pm 8\%$  near pure  $CO<sub>2</sub>$  and a little larger, because of the decrease of  $CO<sub>2</sub>$ , but less than  $\pm 12\%$ , near pure argon or helium.

#### V. RESULTS AND DISCUSSION

As was said before, when the shock wave interacts with the 1.5- $\mu$ m thin membrane, a CO<sub>2</sub>-Ar or CO<sub>2</sub>-He mixing appears, because of the local vorticities and instabilities. Initially, the temperature distribution should be caused by heat diffusion in the mixing zone, but quickly, turbulence dominates, leading to a rapid temperature and concentration equilibrium.

First, if we compare the theoretical results calculated from the Rankine-Hugoniot equations to the experimental values obtained in pure gases (on both sides of the mixing region}, we see that our experimental results are very close to the calculated values of temperature and density in pure gases: for the  $CO<sub>2</sub>$ -Ar case we obtain  $T({\rm CO}_2) = 765$  K,  $\rho({\rm CO}_2) = 0.17$  kg/m<sup>3</sup>, and  $T(Ar) = 1240$  K and for the CO<sub>2</sub>/He case we obtain  $T(CO_2)=637$  K,  $\rho(CO_2)=0.08$  kg/m<sup>3</sup>, and  $T(He)=548$ K. Results showing the evolutions to temperature and density within the mixing region are presented in Figs. 3 and 4 for the  $CO_2$ -Ar and Figs. 5 and 6 for  $CO_2$ -He. The most important result that is observed is that the evolution of temperature in the  $CO<sub>2</sub>$ -Ar mixing zone seems to be more regular than in the case of the  $CO<sub>2</sub>$ -He mixing region. One can say that the  $CO<sub>2</sub>$  and the argon diffuse into each other slowly and more regularly. In the  $CO<sub>2</sub>$ -He case, the system progresses in a more rapid way, and the mixing does not occur in a uniform way. The temperature profile is afFected and the mixing seems to be largely composed of helium because its temperature quickly reaches that of this pure gas. Here, we make clear the importance and influence of the Atwood number on the evolution of the mixing. The efFect of the Atwood number is also visible when we compare the thicknesses of the two mixing zones: 36 mm for the  $CO<sub>2</sub>$ -He case with  $|A_t|_{\text{CO}_2\text{-He}} \approx 0.80$  and 26 mm for the CO<sub>2</sub>-Ar case with  $|A_t|_{\text{CO}_2\text{-Ar}} \approx 0.26$ . During the interpenetration of the two gases, all perturbations present at the interface are amplified—the more so, the larger product  $A_{, *}\Delta U$ . The vorticity rise at the interface is advected in the mixing zone, and widens. Then, thermodynamic equilibrium establishes rapidly, due more to turbulence than to molecular difFusion. On the other hand, the shape of the curves (for argon a regular evolution and for helium a sharp decrease followed by a residual slope) might tell something



FIG. 3. Temperature profile within the  $CO<sub>2</sub>$ -Ar incident mixing zone.



FIG. 4. Density profile within the  $CO<sub>2</sub>-Ar$  incident mixing zone.



FIG. 5. Temperature profile within the  $CO<sub>2</sub>$ -He incident mixing zone.



FIG. 6. Density profile within the  $CO<sub>2</sub>$ -He incident mixing zone.

about the topology of the two mixing zones. One expects, as predicted by numerical studies [20], the penetration of  $CO<sub>2</sub>$  into helium to be in a form of long narrow jets because of the high Atwood number but that the interpenetration of  $CO<sub>2</sub>$  into argon is more regular. This leads us to think that the  $CO<sub>2</sub>$ -He mixing zone should be composed of a greater percentage of helium downstream (in the pure helium side) and a strong concentration of  $CO<sub>2</sub>$  in the portion upstream. On the contrary, in the case of the  $CO<sub>2</sub>$ -Ar mixing zone the interpenetration between the two gases occurs in a much more homogeneous way (low difference between the densities), and we obtain more regular profiles. Thus, one can say that the  $CO<sub>2</sub>$ -Ar mixing zone is more homogeneous than the  $CO<sub>2</sub>$ -He one.

#### VI. CONCLUSION

The infrared  $CO<sub>2</sub>$  laser absorption method, like many other optical or spectroscopical techniques, can contribute to the understanding of the shock induced Richtmyer-Meshkov mixing process in a shock tube, when one of the constituents of the mixing has appropriate spectroscopic properties. Even though the absorption measurements are averaged across the shock tube section, the fact that the only absorbent gas is  $CO<sub>2</sub>$  allows us to obtain some information on the dynamic and thermodynamic states of the shocked gaseous mixing zone. In the present study, our objective was not a diagnostic technique innovation, but to show that this preexisting method could be applicable, with some adaptation, to shock tube experiments, and could provide some helpful information about Richtmyer-Meshkov instability. Unfortunately, when the interaction of the shock and the mixing zone is too greatly influenced by threedimensional boundary layer effects, then the signals obtained are hardly exploitable, and the method is no longer valid. From a dynamical point of view, we have emphasized the influence of the vorticity and the turbulent difFusion on the mixing process, as well as that of the Atwood number. We can say that, at the initial instants, the Atwood number plays an important role in the amplification of the instabilities, and, later, turbulent diffusion become predominant. Presently, our team is improving the absorption technique, using a multidirectional laser absorption method. This method could provide for several regions, instead of one, quantitative information about the wall boundary layer and threedimensional efFects which are quite significant in this kind of experiment.

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Absorption signal of CO<sub>2</sub> -Ar mixing; Mach number 4



FIG. 1. (a) Sketch of the experimental setup; (b) typical absorption signal record.