# Scaling properties of the action in the Riemann-Liouville fractional standard map

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The Riemann-Liouville fractional standard map (RL-fSM) is a two-dimensional nonlinear map with memory given in action-angle variables  $(I, \theta)$ . The RL-fSM is parameterized by K and  $\alpha \in (1, 2]$ , which control the strength of nonlinearity and the fractional order of the Riemann-Liouville derivative, respectively. In this work we present a scaling study of the average squared action  $\langle I^2 \rangle$  of the RL-fSM along strongly chaotic orbits, i.e., for  $K \gg 1$ . We observe two scenarios depending on the initial action  $I_0, I_0 \ll K$  or  $I_0 \gg K$ . However, we can show that  $\langle I^2 \rangle / I_0^2$  is a universal function of the scaled discrete time  $nK^2/I_0^2$  (*n* being the *n*th iteration of the RL-fSM). In addition, we note that  $\langle I^2 \rangle$  is independent of  $\alpha$  for  $K \gg 1$ . Analytical estimations support our numerical results.

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## I. PRELIMINARIES

The kicked rotor represents a free rotating stick in an inhomogeneous field that is periodically switched on in instantaneous pulses; see, e.g., Ref. [1]. It is described by the second-order differential equation

$$\ddot{\theta} + K\sin(\theta) \sum_{j=0}^{\infty} \delta\left(\frac{t}{T} - j\right) = 0.$$
 (1)

Here,  $\theta \in [0, 2\pi]$  is the angular position of the stick, *K* is the kicking strength, *T* is the kicking period (that we set to one from now on), and  $\delta$  is Dirac  $\delta$  function. By replacing the second-order derivative in the equation of motion of the kicked rotor with a Riemann-Liouville (RL) derivative of fractional order  $\alpha$  [2,3], the RL fractional kicked rotor (fKR) is obtained [4,5]:

$${}_{0}D_{t}^{\alpha}\theta + K\sin(\theta)\sum_{j=0}^{\infty}\delta(t-j) = 0, \quad 1 < \alpha \leq 2.$$
 (2)

Above [2,3],

т

$${}_{0}D_{t}^{\alpha}\theta(t) = D_{t}^{m}{}_{0}\mathcal{I}_{t}^{m-\alpha}\theta(t)$$
$$= \frac{1}{\Gamma(m-\alpha)}\frac{d^{m}}{dt^{m}}\int_{0}^{t}\frac{\theta^{\tau}d\tau}{(t-\tau)^{\alpha-m+1}},$$
$$-1 < \alpha \leq m.$$

with  $D_t^m = d^m/dt^m$ ,  ${}_0\mathcal{I}_t^m f(t)$  is a fractional integral given by

$${}_0\mathcal{I}_t^m f(t) = \frac{1}{\Gamma(m)} \int_0^t (t-\tau)^{\alpha-1} f(\tau) d\tau,$$

and  $\Gamma$  is the Gamma function.

The RL-fKR has a stroboscopic version, a two-dimensional nonlinear map with memory, which is well known as the RL fractional standard map (RL-fSM) [5]:

$$I_{n+1} = I_n - K \sin(\theta_n),$$
  

$$\theta_{n+1} = \frac{1}{\Gamma(\alpha)} \sum_{i=0}^n I_{i+1} V_{\alpha}^1(n-i+1), \quad \text{mod}\,(2\pi), \quad (3)$$

where  $I(t) \equiv {}_{0}D_{t}^{\alpha-1}\theta(t)$ , *n* is the discrete time, and  $V_{\alpha}^{k}(m) = m^{\alpha-k} - (m-1)^{\alpha-k}$ . Then, the RL-fSM, given in action-angle variables  $(I, \theta)$ , is parameterized by *K* and  $\alpha \in (1, 2]$ , which control the strength of nonlinearity and the fractional order of the RL derivative, respectively. In fact, for  $\alpha = 2$ , the RL-fSM reproduces the celebrated Chirikov standard map (CSM) [6].

Compared with the CSM, which presents the generic transition to chaos (in the context of the Kolmogorov-Arnold-Moser theorem), depending on the parameter pair (K,  $\alpha$ ), the RL-fSM shows richer dynamics: It generates attractors (fixed points, asymptotically stable periodic trajectories, slow converging and slow diverging trajectories, ballistic trajectories, and fractal-like structures) and/or chaotic trajectories [5,7–10]. Moreover, trajectories may intersect, and attractors may overlap [7].

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FIG. 1. Average squared action  $\langle I_n^2 \rangle_{\text{int}}$  as a function of the discrete time *n* for (a)  $\alpha = 1.1$ , (b)  $\alpha = 1.5$ , and (c)  $\alpha = 1.9$ . Open symbols (full symbols) correspond to  $I_0 \ll K$  ( $I_0 \gg K$ ). The blue dashed lines, plotted to guide the eye, are proportional to *n*. The average is over 100 orbits with initial random phases in the interval  $0 < \theta_0 < 2\pi$ . Red full lines are Eq. (9).

Among the available studies on the RL-fKR, e.g., Refs. [5,7–10], the analysis of strongly chaotic orbits has been left unexplored. Therefore, here we undertake this task and characterize the dynamics of the RL-fSM by computing the squared average action  $\langle I_n^2 \rangle$  when  $K \gg 1$ .

#### **II. NUMERICAL RESULTS**

For the numerical study we wrote a program in FORTRAN90 to compute the orbits of map (3) by straightforward iteration. Moreover, since the memory property in map (3) forbids the use of a large number of orbit realizations, instead of investigating  $\langle I_n^2 \rangle$  directly, to smooth the curves  $\langle I_n^2 \rangle$  vs *n*, we compute its cumulative-normalized value:

$$\langle I_n^2 \rangle_{\rm int} = \frac{1}{n} \int_{n_0=0}^n \langle I_{n'}^2 \rangle dn'.$$
 (4)

Specifically, we compute  $\langle I_n^2 \rangle_{\text{int}}$  for map (3) following two steps: First we calculate the average squared action over the orbit associated with the initial condition j as  $\langle I_{n,j}^2 \rangle =$  $(n+1)^{-1} \sum_{i=0}^{n} I_{i,j}^2$ , where *i* refers to the *i*th iteration of the map. Then,  $\langle I_n^2 \rangle_{\text{int}}$  is defined as the average over M = 100independent realizations of the map (by randomly choosing values of  $\theta_0$  in the interval  $0 < \theta_0 < 2\pi$ ):  $\langle I_n^2 \rangle_{\text{int}}(I_0, K, \alpha) =$  $M^{-1} \sum_{j=1}^{M} \langle I_{n,j}^2 \rangle$ .

In Fig. 1 we plot  $\langle I_n^2 \rangle_{int}$  as a function of the discrete time *n* for representative values of  $\alpha$  in the interval (1,2]: (a)  $\alpha = 1.1$ , (b)  $\alpha = 1.5$ , and (c)  $\alpha = 1.9$ . Several combinations of parameter pairs ( $I_0, K$ ) are considered, as indicated in the right-hand side of the figure. From this figure we observe two scenarios, depending on the initial action  $I_0$  as compared with  $K: I_0 \ll K$  (open symbols) or  $I_0 \gg K$  (full symbols). Specifically, when  $I_0 \ll K, \langle I_n^2 \rangle_{int} \propto n$  for all *n* (see the blue dashed lines), while for  $I_0 \gg K$ , first,  $\langle I_n^2 \rangle_{int}$  remains approximately constant and proportional to  $I_0^2$  up to a crossover time  $n^*$ , after which  $\langle I_n^2 \rangle_{int}$  grows proportional to *n*.

From Fig. 1 it can also be seen that the crossover time  $n^*$  depends on both  $I_0$  and K, while the dependence of  $n^*$  with  $\alpha$  is not evident. Then, to look for the dependence of  $n^*$  on the map parameters, in Fig. 2(a) [Fig. 2(b)] we plot  $\langle I_n^2 \rangle_{int}$  vs *n* for several values of  $I_0$  [K] and fixed K [ $I_0$ ]. In both figures we

use  $\alpha = 1.1$ . We numerically extract  $n^*$  as the crossing point between the functions  $\langle I_n^2 \rangle_{\text{int}} = I_0^2$  and  $\langle I_n^2 \rangle_{\text{int}} = Cn$  (which is the fitting to the data in the growing regime); as examples,



FIG. 2. (a), (b) Average squared action  $\langle I_n^2 \rangle_{int}$  as a function of *n* for (a)  $K = 10^3$  and several values of  $I_0$  (2×10<sup>3</sup>, 4×10<sup>3</sup>, 10<sup>4</sup>, 2×10<sup>4</sup>,  $4 \times 10^4$ , 10<sup>5</sup>, and  $2 \times 10^5$ , from bottom to top) and (b)  $I_0 = 10^5$  and several values of K (4×10<sup>2</sup>, 10<sup>3</sup>, 2×10<sup>3</sup>, 4×10<sup>3</sup>, 10<sup>4</sup>, 2×10<sup>4</sup>, and  $4 \times 10^4$ , from bottom to top). As in Fig. 1, the average is taken over 100 orbits with initial random phases in the interval  $0 < \theta_0 < 2\pi$ . All data in (a), (b) correspond to  $\alpha = 1.1$ . The horizontal red dashed lines in (a), (b) indicate  $\langle I_n^2 \rangle_{\text{int}} = I_0^2$  with (a)  $I_0 = 4 \times 10^4$  and (b)  $I_0 = 10^5$ , respectively. The transverse red dashed lines in (a), (b) are fittings of  $\langle I_n^2 \rangle_{\text{int}} = Cn$  to the data represented by asterisks (for  $n \ge 10^4$ ) with fitting constants (a)  $\mathcal{C} = 294\,658$  and (b)  $\mathcal{C} = 3\,984\,348$ . (c), (d) Crossover time  $n^*$  (c) as a function of  $I_0$  for constant K ( $K = 10^3$ ) and (d) as a function of K for constant  $I_0$  ( $I_0 = 10^5$ ). In (c), (d), three values of  $\alpha$  are reported:  $\alpha = 1.1, 1.5, \text{ and } 1.9$ . Red dashed lines in (c), (d) are power-law fittings to the data of the form (a)  $n^* \propto I_0^{\gamma_1}$ with  $\gamma_1 \approx 2$  and (b)  $n^* \propto K^{\gamma_2}$  with  $\gamma_2 \approx -2$ . Blue dot-dashed lines in (c), (d) are Eq. (11).



FIG. 3. Normalized average squared action  $\langle I_n^2 \rangle_{int} / I_0^2$  as a function of the normalized time  $nK^2 / I_0^2$ . Same data sets as in Fig. 1. The red line is Eq. (10).

see the horizontal and transverse dashed lines in Figs. 2(a) and 2(b), respectively. Thus, in Figs. 2(c) and 2(d) we plot the obtained values of  $n^*$  for  $\alpha = 1.1$  but also for  $\alpha = 1.5$  and 1.9. Figures 2(c) and 2(d) reveal the power-law dependence

$$n^* \propto I_0^{\gamma_1} K^{\gamma_2} \tag{5}$$

and the independence of  $n^*$  on  $\alpha$ . Power-law fittings of the data in Figs. 2(c) and 2(d) provide  $\gamma_1 \approx 2$  and  $\gamma_2 \approx -2$ , see red dashed lines.

Equation (5) together with the observation that  $\langle I_n^2 \rangle_{\text{int}} \approx I_0^2$ for  $n < n^*$  allow us to scale the curves  $\langle I_n^2 \rangle_{\text{int}}$  vs *n*. Indeed, in Fig. 3 we plot  $\langle I_n^2 \rangle_{\text{int}} / I_0^2$  as a function of the normalized time  $nK^2/I_0^2$  (i.e.,  $n/n^*$ ) and observe the collapse of all curves on top of a *universal* function.

### **III. ANALYTICAL ESTIMATION**

Now, to support and better understand the scaling performed above, we derive an analytical estimation for  $\langle I_n^2 \rangle_{int}$ ; see, e.g., Ref. [11]. From the first line of map (3) we have that  $I_{n+1}^2 = I_n^2 - 2KI_n \sin(\theta_n) + K^2 \sin^2(\theta_n)$ , so we can write

$$\langle I_{n+1}^2 \rangle = \langle I_n^2 \rangle - 2K \langle I_n \rangle \langle \sin(\theta_n) \rangle + K^2 \langle \sin^2(\theta_n) \rangle.$$

Since for chaotic orbits we can assume that  $\langle \sin(\theta_n) \rangle = 0$ , the term  $2K \langle I_n \rangle \langle \sin(\theta_n) \rangle$  can be eliminated. Therefore

$$\left\langle I_{n+1}^2 \right\rangle = \left\langle I_n^2 \right\rangle + \frac{K^2}{2},\tag{6}$$

where we have used  $\langle \sin^2(\theta_n) \rangle = 1/2$ . Then, by noticing that

$$\langle I_{n+1}^2 \rangle - \langle I_n^2 \rangle = \frac{\langle I_{n+1}^2 \rangle - \langle I_n^2 \rangle}{(n+1) - n} \approx \frac{dJ}{dn} ,$$

we rewrite Eq. (6) as the first-order differential equation,

$$\frac{dJ}{dn} = \frac{K^2}{2},\tag{7}$$

where  $J \equiv \langle I_n^2 \rangle$ . Therefore, by solving (7) we can write

$$\langle I_n^2 \rangle = I_0^2 + \frac{K^2}{2}n,$$
 (8)

where we have used  $J_0 = \langle I_0^2 \rangle = I_0^2$  and  $n_0 = 0$ . Finally, by substituting Eq. (8) into Eq. (4), we can also write down an explicit expression for  $\langle I_n^2 \rangle_{\text{int}}$ :

$$\langle I_n^2 \rangle_{\rm int} = I_0^2 + \frac{K^2}{4}n.$$
 (9)

Indeed, Eq. (9) reproduces our numerical data well, as can be seen in Fig. 1 where we have included Eq. (9) as red lines.

### IV. DISCUSSION AND CONCLUSIONS

Given the good correspondence of Eq. (9) and the numerical data, it is clear that it reproduces the scaling laws reported in Sec. II, which can be summarized as

$$\langle I_n^2 \rangle_{\text{int}} = \begin{cases} \propto K^2 n , & \text{when } I_0 \ll K, \\ \approx I_0^2 , & n < n^* \\ \propto K^2 n , & n > n^* \end{cases} \quad \text{when } I_0 \gg K.$$

Moreover, Eq. (9) can also be used to demonstrate that the ratio  $\langle I_n^2 \rangle_{\text{int}} / I_0^2$  is a simple *universal* function of the variable  $\overline{n} = n/n^*$ :

$$\frac{\langle I_n^2 \rangle_{\text{int}}}{I_0^2} = 1 + \overline{n},\tag{10}$$

where the crossover time  $n^*$  is now naturally defined as

$$n^* \equiv 4I_0^2 K^{-2}, \tag{11}$$

in agreement with Eq. (5). Finally, in Figs. 2(c) and 2(d), and 3 we plot Eqs. (11) and (10) (see dot-dashed blue lines and red full lines), respectively, and observe an excellent agreement with the numerical data.

It is relevant to notice that for strongly chaotic orbits,  $K \gg 1$ , the average squared action  $\langle I_n^2 \rangle$  for the RL-fSM does

not depend on the order  $\alpha$  of the fractional derivative. Indeed, the panorama reported here for  $\langle I_n^2 \rangle$  vs *n* is equivalent to that of CSM [11,12] as well as that of the discontinuous standard map [11,13], both with  $K \gg 1$ . This could be understood from the analytical estimation of Sec. III by noticing that to obtain the expression for  $\langle I_n^2 \rangle$  we mainly used the first equation of map (3), which does not contain the parameter  $\alpha$ ; i.e., the property of memory, parametrized by  $\alpha$ , is only present in the equation for  $\theta$ , which is a cyclic variable. So when  $K \gg 1$ ,  $\langle I_n^2 \rangle$  must be independent of  $\alpha$ . That is, to observe effects of  $\alpha$  on the dynamics of the RL-fSM,  $K \sim 1$  should be set, see Refs. [5,7–10].

We stress that similar studies can be carried out for other types of nonlinearity (not just the continuous sine-shaped nonlinearity in the first equation of the RL-fSM) and for other

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types of fractional derivatives. This, in fact, will be the subject of future investigations.

Finally, we want to add that our work falls within the scope of the general fractional dynamics (GFDynamics), a line of research recently introduced in Ref. [14].

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