Collinear velocity relaxation of two spheres in a viscous incompressible fluid

B. U. Felderhof \bullet^* \bullet^*

Institut für Theorie der Statistischen Physik, RWTH Aachen University, Templergraben 55, 52056 Aachen, Germany

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Collinear velocity relaxation of two spheres immersed in a viscous incompressible fluid is studied on the basis of an approximate expression for the retarded hydrodynamic interaction. After a sudden impulse applied to one sphere, the other one instantaneously starts to move as well, with amplitude determined by the added mass effect. The velocities of both spheres eventually decay with a *t*−3/² long-time tail, but the relative velocity decays with a *t*−5/² long-time tail. The three relaxation functions are approximated by simple expressions involving only a small number of poles in the complex square root of frequency plane.

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I. INTRODUCTION

In recent work [\[1\]](#page-6-0) we investigated the retarded hydrodynamic interaction between two spheres immersed in a viscous incompressible fluid. The corresponding frequencydependent mobility matrix relates the translational and rotational velocities of the spheres to the hydrodynamic forces and torques exerted by the fluid. In combination with Newton's equations, the mobility matrix yields the sphere velocities for oscillatory forces and torques applied to the spheres.

By Fourier analysis the mobility matrix also provides information on the behavior in time of the velocity relaxation functions after a sudden impulse or twist applied to one of the spheres. Due to the assumption of incompressibility there is an instantaneous transfer of momentum, and the other sphere starts to move as well. The nature of the motion provides information on the nature of the fluid. The two-sphere system can be used as an investigative tool to analyze the viscoelasticity of the fluid $[2-5]$.

In the following we consider in particular the translational velocities for collinear motions of the two spheres. For motions transverse to the line of centers there is translationrotation coupling, leading to a complicated overall motion. The restriction to collinear motion allows a relatively simple one-dimensional picture, but even for this case the motion is intricate.

The initial values of the sphere velocities after a sudden impulse from a state of rest are determined by added mass effects. We compare the initial values found from the high-frequency behavior of the mobility matrix with those predicted by potential flow theory [\[6\]](#page-6-0). The velocity autocorrelation functions, which describe the time-dependent velocities after the initial impulse, decay with a wide distribution of relaxation times. The low-frequency dependence of the mobility matrix incorporates the long-time decay. It is known that the amplitude of the long-time *t*−3/² decay of the two sphere

velocities is identical to that of a single sphere [\[7\]](#page-6-0). We show that the relative velocity decays with a $t^{-\overline{5}/2}$ long-time tail.

Our analysis is based on the recently derived approximate expression for the frequency-dependent mobility matrix [\[1\]](#page-6-0). In the approximation the hydrodynamic interaction between the two spheres is limited to a single Green function, but finite-size effects are fully taken into account via the exact expression for the primary frequency-dependent Stokes flow, and a Faxén theorem in the calculation of the secondary velocity. The approximation should be accurate, unless the spheres initially are very close.

The mobility function for collinear motion is a complicated function of frequency. We show for a numerical example that it can be well approximated by a much simpler function, involving only a small number of poles in the complex square root of frequency plane. In this manner the relaxation behavior fits a general framework of slow dynamics of linear relaxation systems $[8]$.

The dynamics of two colloidal spheres immersed in a compressible fluid was studied in a computer simulation by Tatsumi and Yamamoto [\[9\]](#page-6-0). In their theoretical analysis these authors used a simple Green function approximation to the mobility function [\[10\]](#page-6-0). In most practical applications it will be sufficient to consider the incompressible limit. Nonetheless it would be of interest to extend the present theory to a compressible fluid.

II. PAIR OF INTERACTING SPHERES

We consider two uniform spheres, labeled *A* and *B*, with radii *a* and *b* and mass densities ρ_A and ρ_B , immersed in a viscous incompressible fluid, and oscillating with frequency ω about positions \mathbf{R}_A and \mathbf{R}_B , as shown in Fig. [1.](#page-1-0) We choose a Cartesian system of coordinates such that the *z* axis is along the vector $\mathbf{R} = \mathbf{R}_B - \mathbf{R}_A$. We may choose the origin at \mathbf{R}_A . The fluid has mass density ρ and shear viscosity η . The fluid flow velocity $v(r, t)$ and pressure $p(r, t)$ satisfy the linearized Navier-Stokes equations:

$$
\rho \frac{\partial \mathbf{v}}{\partial t} = \eta \nabla^2 \mathbf{v} - \nabla p, \qquad \nabla \cdot \mathbf{v} = 0. \tag{2.1}
$$

^{*}ufelder@physik.rwth-aachen.de

FIG. 1. Snapshot of two spheres oscillating along the line of centers.

We can linearize, since the oscillations are assumed to be small, so that the inertial Reynolds term can be omitted on the left-hand side of the first equation. At the surface of the spheres the fluid velocity and pressure are assumed to satisfy mixed-slip boundary conditions with slip coefficients ξ*^A* and ξ_B , respectively [\[11\]](#page-6-0). The value $\xi = 0$ corresponds to no-slip, and $\xi = 1/3$ corresponds to perfect slip [\[12\]](#page-6-0). The fluid fills all space outside the spheres and is at rest at infinity.

The whole system is caused to move by the oscillatory applied forces E_A and E_B and the torques N_A and N_B , acting on the spheres, which force the system to oscillate at frequency ω . With the above assumptions the problem is linear. As a consequence the translational velocity amplitudes U_A and U_B and the rotational velocity amplitudes Ω_A and Ω_B are linear in the amplitudes of the applied forces and torques. They are also linear in the hydrodynamic forces K_A and K_B and the torques *T^A* and *T^B* exerted by the fluid on the spheres. The equations of motion for the spheres in complex shorthand notation are as follows:

$$
-i\omega \mathsf{m}_p \cdot \mathsf{U} = \mathsf{E} + \mathsf{K},
$$

$$
-i\omega \mathsf{I}_p \cdot \Omega = \mathsf{N} + \mathsf{T},
$$
 (2.2)

with six-vectors $V = (V_A, V_B)$. The mass matrix m_p is a diagonal 6×6 matrix incorporating the two masses m_A and m_B , and similarly I_p is a diagonal 6×6 matrix incorporating the two moments of inertia I_A and I_B . The 12×12 mobility matrix μ is defined by the following linear relation:

$$
U = -\mu^{tt} \cdot K - \mu^{tr} \cdot T,
$$

\n
$$
\Omega = -\mu^{rt} \cdot K - \mu^{rr} \cdot T.
$$
 (2.3)

The 12 × 12 friction matrix is the inverse $\zeta = \mu^{-1}$. The *tt* part of this matrix has the property that it varies in proportion to ω at high frequency. The property is called the acceleration reaction by Batchelor [\[13\]](#page-6-0), who discusses it for a single rigid body. The 6×6 added mass matrix m_a is defined by

$$
\zeta^{tt} = -i\omega \mathsf{m}_a + \zeta^{tt'}, \qquad (2.4)
$$

where $\zeta^{tt'}$ is the remaining part, which does not increase in proportion to ω at high frequency. It follows from the recip-rocal theorem that the various matrices are symmetric [\[13\]](#page-6-0). The elements of the mobility matrix define scalar mobility functions. In particular the *tt* part of the matrix takes the form

$$
\mu^{tt} = \begin{pmatrix} \mu_{AA}^{tt} & \mu_{AB}^{tt} \\ \mu_{BA}^{tt} & \mu_{BB}^{tt} \end{pmatrix}, \tag{2.5}
$$

with the self-mobility tensors μ_{AA}^{tt} and μ_{BB}^{tt} and the mutual mobility tensors μ_{AB}^{tt} and μ_{BA}^{tt} given by

$$
\mu_{ij}^{tt} = \alpha_{ij}^{tt}(R,\omega)\hat{R}\hat{R} + \beta_{ij}^{tt}(R,\omega)(I - \hat{R}\hat{R}), \qquad (2.6)
$$

with $\hat{\mathbf{R}} = \mathbf{R}/R$, and the scalar mobility functions $\alpha_{ij}^{tt}(R, \omega)$ and $\beta_{ij}^{tt}(R,\omega)$. The form follows from translational and rotational invariance. Translation-rotation coupling is expressed by the scalar mobility functions $\beta_{ij}^{tr}(R, \omega)$ and $\beta_{ij}^{rt}(R, \omega)$, which are related by the reciprocity relation $\beta_{ij}^{tr}(R, \omega) = -\beta_{ji}^{rt}(R, \omega)$. The tensor μ_{ij}^{tr} takes the form

$$
\mu_{ij}^{tr} = \beta_{ij}^{tr}(R,\omega)\epsilon \cdot \hat{R}, \qquad (2.7)
$$

where ϵ is the Levi-Civita tensor. The tensor μ_{ij}^{rr} takes a form similar to Eq. (2.6) .

We consider first the longitudinal case, where both spheres translate along the *z* axis, parallel to $\hat{\mathbf{R}}$. Then by symmetry there is no translation-rotation coupling, and it suffices to consider the scalar mobility functions $\alpha_{ij}^{tt}(R,\omega)$. Two of these are related by the reciprocity relation $\alpha_{AB}^{tt}(R, \omega) = \alpha_{BA}^{tt}(R, \omega)$, and $\alpha_{BB}^{tt}(R, \omega)$ can be obtained from $\alpha_{AA}^{tt}(R, \omega)$ by *AB* interchange.

Elsewhere we derived an approximate expression for the mutual mobility function $\alpha_{AB}^{tt}(R,\omega)$ based on a onepropagator approximation, which takes account of only a single Green function between the two spheres [\[1\]](#page-6-0). The primary flow is a Stokes flow at frequency ω generated by sphere *A* as if it were moving by itself in infinite fluid. The velocity of sphere B as it moves in this flow with zero force is calculated from a Faxén theorem $[14]$. The resulting mutual mobility function reads [\[1\]](#page-6-0) as

$$
\alpha_{BA}^{tt}(R,\omega) = \frac{B_0(\alpha a, \xi_A)B_0(\alpha b, \xi_B) - (1 + \alpha R)e^{\alpha(a+b-R)}}{2\pi \eta \alpha^2 R^3 A_0(\alpha a, \xi_A)A_0(\alpha b, \xi_B)},
$$
\n(2.8)

with the abbreviations

$$
A_0(\lambda, \xi) = (1 - \xi) \frac{1 + \lambda}{1 + \xi \lambda} + \frac{1}{9} \lambda^2,
$$

\n
$$
B_0(\lambda, \xi) = (1 - \xi) \frac{1 + \lambda}{1 + \xi \lambda} + \frac{1}{3} \lambda^2,
$$
 (2.9)

and $\alpha = \sqrt{-i\omega\rho/\eta}$, Re $\alpha > 0$. The approximate self-mobility function $\alpha_{AA}^{tt}(R,\omega)$ is more complicated [\[1\]](#page-6-0). It is calculated from a single reflection from sphere *B*, which is freely moving. The effect of the correction to the single-particle mobility $\mu_A^t = 1/\zeta_A^t$ on the velocity relaxation function is numerically small for not too close distances. In our analysis we use the approximation $\alpha_{AA}^{tt}(R,\omega) \approx 1/\zeta_A^t(\omega,\xi_A)$ with the single sphere friction coefficient

$$
\xi_A^t(\omega, \xi_A) = 6\pi \eta a A_0(\alpha a, \xi_A). \tag{2.10}
$$

We can estimate the validity of the approximation by considering the added mass matrix which follows from these expressions. We find that the part corresponding to longitudinal motions is given by

$$
\mathsf{m}_{a}^{\parallel} = \frac{2\pi\,\rho}{3} \bigg[1 + O\bigg(\frac{a^3 b^3}{R^6}\bigg) \bigg] \bigg(\frac{a^3}{-3\frac{a^3 b^3}{R^3}} - \frac{-3\frac{a^3 b^3}{R^3}}{b^3} \bigg). \tag{2.11}
$$

The terms omitted in $\alpha_{AA}^{tt}(R,\omega)$ contribute to the $O(a^3b^3/R^6)$ term. The diagonal terms follow from the added mass of each individual sphere [\[13,14\]](#page-6-0).

In the same one-propagator approximation we found the following for the mutual mobility function for motions normal to the line of centers:

$$
\beta_{BA}^{tt}(R,\omega) = \frac{-(1+\xi_A\alpha a)(1+\xi_B\alpha b)B_0(\alpha a,\xi_A)B_0(\alpha b,\xi_B) + (1-\xi_A)(1-\xi_B)A_1(\alpha R)e^{\alpha(a+b-R)}}{4\pi \eta \alpha^2 R^3 (1+\xi_A\alpha a)(1+\xi_B\alpha b)A_0(\alpha a,\xi_A)A_0(\alpha b,\xi_B)},
$$
(2.12)

with the function

$$
A_1(\lambda) = 1 + \lambda + \lambda^2. \tag{2.13}
$$

At long range it is not necessary to consider the translationrotation coupling, since at finite frequency the corresponding mobility functions decay exponentially. In the same way as above, for the part of the added mass matrix corresponding to transverse motions, we find

$$
\mathsf{m}_{a}^{\perp} = \frac{2\pi\,\rho}{3} \bigg[1 + O\bigg(\frac{a^{3}b^{3}}{R^{6}} \bigg) \bigg] \bigg(\frac{a^{3}}{3a^{3}b^{3}} - \frac{3a^{3}b^{3}}{2R^{3}} \bigg). \tag{2.14}
$$

Note that the added mass matrices in Eqs. (2.11) and (2.14) are independent of the two slip coefficients.

The expressions for the added mass matrices are identical to the ones derived by Lamb $[15]$ in potential flow theory from the kinetic energy of the flow pattern. The results also follow from a linear response theory [\[6\]](#page-6-0) based on the expression for the force on a sphere subjected to an incident potential flow, as derived by Landau and Lifshitz [\[16\]](#page-6-0) and by Batchelor [\[13\]](#page-6-0). In dipole approximation this leads to an expression for the added mass matrix given by [\[17\]](#page-6-0)

$$
\mathsf{m}_a = -\mathsf{m}_f + 4\pi \rho \mathcal{A},\tag{2.15}
$$

where the matrix \mathbf{m}_f is diagonal with the elements m_{fA} = $4\pi \rho a^3/3$ and $m_{fB} = 4\pi \rho b^3/3$ corresponding to the displaced mass of each sphere, and $\mathcal A$ is the inverse of a matrix involving the interactions between induced dipoles. The evaluation of Eq. (2.15) for two spheres $[18]$ yields results consistent with Eqs. (2.11) and (2.14) .

III. VELOCITY RELAXATION

Added mass effects influence the response of the system to a sudden impulse applied to one of the spheres. Let the fluid and the spheres be at rest for $t < 0$, and consider small applied forces of the form $E(t) = S\delta(t)$ with the impulse vector $S =$ (S_A, S_B) . For $S_B = 0$, both sphere *A* and sphere *B* start to move at $t = 0 +$. More generally we have

$$
\mathsf{U}(0+) = \mathsf{m}^{-1} \cdot \mathsf{S},\tag{3.1}
$$

with the six-dimensional mass matrix $m = m_p + m_a$. The equation generalizes the known acceleration reaction for a single sphere $[13]$ to two spheres. The sudden push on both spheres creates an irrotational flow pattern with boundary values corresponding to the two sphere velocities [\[14\]](#page-6-0). The linear relation between sphere velocities and the imposed impulses follows from the pressure exerted on each sphere and defines the mass matrix [\[6\]](#page-6-0).

At later times

$$
\mathsf{U}(t) = \mathsf{R}(t) \cdot \mathsf{S}, \quad t > 0,\tag{3.2}
$$

with a relaxation matrix, $R(t)$, which has the one-sided Fourier transform

$$
\hat{\mathsf{R}}(\omega) = \int_0^\infty e^{i\omega t} \mathsf{R}(t) \, dt. \tag{3.3}
$$

We can identify

$$
\hat{\mathsf{R}}(\omega) = \mathsf{Y}^{tt}(\omega), \quad \mathsf{Y}(\omega) = [-i\omega \mathsf{M}_p + \zeta(\omega)]^{-1}, \qquad (3.4)
$$

with the 12×12 generalized mass matrix M_p which follows from Eq. [\(2.2\)](#page-1-0). The matrix $Y(\omega)$ is called the admittance matrix [\[19\]](#page-6-0).

The situation is simplest for longitudinal motions. If sphere *A* gets a sudden push in the direction of the line of centers, then also sphere *B* starts to move in the same direction. The reaction is instantaneous due to the assumption of incompressibility. In a compressible fluid the reaction would take some time due to the finite velocity of sound, as seen in the computer simulation of Tatsumi and Yamamoto [\[9\]](#page-6-0). From Eq. (3.4) we can evaluate how the velocity of each sphere relaxes after the initial push. It follows from a general theorem derived by Cichocki and Felderhof [\[7\]](#page-6-0) that at long times the velocity of each sphere decays with a *t*−3/² long-time tail, with an amplitude which is the same as if each sphere were by itself. The initial values of the two velocities are determined by the effective mass matrix, including the added mass terms which come from the high-frequency behavior of the friction matrix $\zeta^{tt}(\omega)$. Neither the initial values nor the amplitude of the long-time tails depend on the slip coefficients (ξ_A, ξ_B) . Since the friction coefficient of a single sphere decreases from the Stokes value $6\pi\eta a$ for no-slip to $6\pi\eta(1-\xi_A)$ with increasing slip coefficient ξ_A , we expect that the mean relaxation time will increase when the slip coefficients (ξ_A, ξ_B) increase.

For two Brownian spheres in thermal equilibrium the velocities are not correlated, and the thermal average $\langle U_A U_B \rangle$ vanishes. From Eq. (2.11) we see that after a very short time the longitudinal velocity components are correlated positively, whereas Eq. (2.14) shows that then the transverse components are correlated negatively. At later times the velocity relaxation function $C_{AB}(t) = \langle U_A(t)U_B(0) \rangle$ is related to the relaxation matrix by the fluctuation-dissipation theorem $C_{AB}(t)$ = $k_B T R_{AB}(t)$.

It is convenient to take the *z* axis along the line of centers. Then by symmetry the six-dimensional matrix $R(t)$ decomposes into a two-dimensional matrix $R^{||}(t)$ corresponding to longitudinal motions in the *z* direction and two identical two-dimensional matrices $\mathsf{R}^{\perp}(t)$ corresponding to transverse motions in the *x* and *y* directions.

We consider scalar autocorrelation functions of the form

$$
C(t) = (\psi | \mathbf{R}(t) | \psi), \tag{3.5}
$$

where $|\psi\rangle$ is a chosen six-dimensional vector selecting a linear combination of translational velocity components. We define the one-sided Fourier transform as

$$
\hat{C}(\omega) = \int_0^\infty e^{i\omega t} C(t) \, dt. \tag{3.6}
$$

This is given by

$$
\hat{C}(\omega) = (\psi | Y^{tt}(\omega) | \psi).
$$
 (3.7)

The initial value of the autocorrelation function is

$$
C(0+) = (\psi|m^{-1}|\psi).
$$
 (3.8)

We write the autocorrelation function in the form

$$
C(t) = C(0+) \gamma(t/\tau_M), \qquad (3.9)
$$

with initial value $\gamma(0+) = 1$ and the mean relaxation time

$$
\tau_M = \frac{1}{C(0+)} \int_0^\infty C(t) \, dt. \tag{3.10}
$$

From Eqs. (3.7) and (3.8) we find

$$
\tau_M = \frac{(\psi | \mu^{tt}(0) | \psi)}{(\psi | \mathsf{m}^{-1} | \psi)}.
$$
\n(3.11)

We define the variable $z = -i\omega\tau_M$ and the function

$$
\Gamma(z) = \frac{\hat{C}(\omega)}{(\psi|\mu^{tt}(0)|\psi)}.
$$
\n(3.12)

This has the properties

$$
\Gamma(0) = 1, \quad \lim_{z \to \infty} z \Gamma(z) = 1. \tag{3.13}
$$

Since the functions defined in Eqs. (2.8) and (2.12) depend on frequency via the variable α , the dependence of $\Gamma(z)$ on *z* is via $y = \sqrt{z}$. The spectral density $p(u)$ is defined by [\[8\]](#page-6-0)

$$
p(u) = \frac{1}{\pi} \operatorname{Im} \left[\Gamma(y \to -i\sqrt{u}) \right],\tag{3.14}
$$

for positive *u*. The relaxation function $\gamma(\tau)$ in Eq. [\(3.5\)](#page-2-0) is given by the inverse Stieltjes transform [\[20\]](#page-7-0)

$$
\gamma(t/\tau_M) = \int_0^\infty p(u) \exp[-ut/\tau_M] \ du. \tag{3.15}
$$

The function $\Gamma(z)$ has the Stieltjes representation

$$
\Gamma(z) = \int_0^\infty \frac{p(u)}{u+z} du.
$$
 (3.16)

Hence the spectral density has the properties

$$
\int_0^\infty p(u) \, du = 1, \quad \int_0^\infty \frac{p(u)}{u} \, du = 1,\tag{3.17}
$$

corresponding to Eq. (3.13) . The second property in Eqs. (3.13) and (3.17) corresponds to the choice of τ_M as the timescale.

Since we want to compare the spectral densities for different vectors ψ , it is more convenient to write the relaxation function as

$$
\gamma(t/\tau_M) = \int_0^\infty P(s)e^{-st} \, ds,\tag{3.18}
$$

with the rate distribution

$$
P(s) = \tau_M p(\tau_M s), \tag{3.19}
$$

where *s* is the relaxation rate $s = u/\tau_M$.

IV. COLLINEAR MOTION

We consider velocity relaxation along the line of centers in some more detail. The scalar mobility function in Eq. (2.8) is a complicated function of frequency and correspondingly

the various autocorrelation functions are intricate functions of time. We show that, provided the two spheres are of comparable size, a relatively simple approximate description can be found.

It is known that for a single sphere the velocity relaxation function shows an important long-time tail. In that case the spectral density is of the form [\[19\]](#page-6-0)

$$
p_2(u) = \frac{1}{\pi} \frac{\sigma \sqrt{u}}{1 + (\sigma^2 - 2)u + u^2},
$$
 (4.1)

where the parameter σ can be found from the mass and the added mass of the sphere. The long-time tail corresponds to the square root singularity at $u = 0$. The velocity relaxation function is a sum of two *w* functions,

$$
\gamma(\tau) = \frac{1}{\sqrt{\sigma^2 - 4}} [y_+ w(-iy_+ \sqrt{\tau}) - y_- w(-iy_- \sqrt{\tau})], \quad (4.2)
$$

with

$$
w(z) = e^{-z^2} \operatorname{erfc}(-i\sqrt{z})
$$
 (4.3)

and values y_{\pm} which correspond as $y_{\pm} = \sqrt{z_{\pm}}$ to the zeros of the denominator of the Laplace transform of the relaxation function,

$$
\Gamma_2(z) = \frac{1}{1 + \sigma\sqrt{z} + z},\tag{4.4}
$$

where $z = -i\omega\tau_M$, with the mean relaxation time $\tau_M =$ $m^*/(6\pi n a)$ for a sphere of radius *a* and mass m_0 with effective mass $m^* = m_0 + m_f/2$ and $m_f = 4\pi \rho a^3/3$. The relaxation function has the long-time behavior

$$
\gamma(\tau) \approx \frac{\sigma}{2\sqrt{\pi}} \tau^{-3/2}
$$
 as $\tau \to \infty$, (4.5)

where $\tau = t/\tau_M$ and $\sigma = \sqrt{9m_f/2m^*}$. The corresponding long-time behavior of the velocity autocorrelation function is

$$
C(t) \approx \frac{1}{12} \sqrt{\rho} (\pi \eta t)^{-3/2} \quad \text{as} \quad t \to \infty. \tag{4.6}
$$

For the pair of spheres in collinear motion it suffices to consider the two-dimensional matrix $\mathsf{R}^{\parallel}(t)$ and corresponding two-vectors $|\psi\rangle$. We consider the three vectors

$$
|\psi\rangle_A = (1, 0), \quad |\psi\rangle_B = (0, 1), \quad |\psi\rangle_d = (1, -1). \quad (4.7)
$$

The first corresponds to the motion of sphere *A*, the second to the motion of sphere *B*, and the third to the relative motion. We then have

$$
C_{AA}^{\parallel}(t) = R_{AA}^{\parallel}(t), \quad C_{BB}^{\parallel}(t) = R_{BB}^{\parallel}(t),
$$

\n
$$
C_{dd}^{\parallel}(t) = R_{AA}^{\parallel}(t) - 2R_{AB}^{\parallel}(t) + R_{BB}^{\parallel}(t). \tag{4.8}
$$

This shows that conversely the three elements $R_{ij}^{\parallel}(t)$ can be found from the three autocorrelation functions:

$$
R_{AA}^{\parallel}(t) = C_{AA}^{\parallel}(t), \quad R_{BB}^{\parallel}(t) = C_{BB}^{\parallel}(t),
$$

\n
$$
R_{AB}^{\parallel}(t) = \frac{1}{2} \Big[C_{AA}^{\parallel}(t) - C_{dd}^{\parallel}(t) + C_{BB}^{\parallel}(t) \Big].
$$
\n(4.9)

It follows from a general theorem [\[7\]](#page-6-0) that the relaxation functions $R_{AA}^{\parallel}(t)$ and $R_{BB}^{\parallel}(t)$ have exactly the same long-time behavior as in Eq. (4.6). At long times the two spheres move

collectively. We show below that the relative velocity decays with a *t*−5/² long-time tail.

We express the autocorrelation functions as in Eq. (3.17) . This yields the rate distributions $P_{AA}^{\parallel}(s)$, $P_{dd}^{\parallel}(s)$, and $P_{BB}^{\parallel}(s)$ as functions of the rate *s*. Hence we find the time-dependent autocorrelation functions $C_{AA}^{\parallel}(t)$, $C_{dd}^{\parallel}(t)$, and $C_{BB}^{\parallel}(t)$ by numerical integration.

The initial values of the autocorrelation functions, as given by Eq. [\(3.8\)](#page-3-0), are

$$
C_{AA}^{\parallel}(0+) = \frac{3}{2\pi a^3} \frac{R^6(\rho + 2\rho_B) - 18a^3b^3\rho_B}{R^6(\rho + 2\rho_A)(\rho + 2\rho_B) - 36a^3b^3\rho_A\rho_B},
$$

\n
$$
C_{dd}^{\parallel}(0+) = \frac{3}{2\pi a^3b^3} \frac{a^3R^6(\rho + 2\rho_A) + b^3R^6(\rho + 2\rho_B) - 6a^3b^3R^3\rho - 18a^6b^3\rho_A - 18a^3b^6\rho_B}{R^6(\rho + 2\rho_A)(\rho + 2\rho_B) - 36a^3b^3\rho_A\rho_B},
$$
\n(4.10)

with $C_{BB}^{\parallel}(0+)$ found by *AB*- interchange.

The three mean relaxation times τ_{MAA} , τ_{Mdd} , and τ_{MBB} are given by Eq. (3.11) as

$$
\tau_{MAA}^{\parallel} = \frac{a^2}{9\eta} \frac{R^6(\rho + 2\rho_A) - 18a^3b^3\rho_A}{R^6 - 9a^3b^3},
$$

$$
\tau_{Mdd}^{\parallel} = \frac{a^3b + ab^3 - 3abR^2 + aR^3 + bR^3}{6\pi\eta abR^3C_{dd}(0+)},
$$
 (4.11)

with τ_{MBB}^{\parallel} being found by *AB* interchange.

We note that the two-dimensional admittance matrix has the following low-frequency expansion [\[7\]](#page-6-0):

$$
\mathsf{Y}^{tt\|}(\omega) = \pmb{\mu}^{tt\|}(0) - \frac{\alpha}{6\pi\eta} \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} + O(\alpha^2 a^2). \tag{4.12}
$$

This corresponds to the universal nature of the collective longtime motion, mentioned above.

We show below for a numerical example with two spheres of comparable size that the three relaxation functions are well approximated by a sum of w functions similar to Eq. (4.2) with either two or three terms. The simple approximate description corresponds to a small number of poles of the admittance in the complex \sqrt{z} plane.

V. NUMERICAL EXAMPLE

As an example we consider two neutrally buoyant spheres of radii $a, b = 2a$, at the center-to-center distance $R = 5a$, with mass densities $\rho_A = \rho$ and $\rho_B = \rho$, with no-slip boundary conditions, corresponding to slip coefficients $\xi_A = \xi_B =$ 0. We consider the motion along the line of centers. The explicit expression for the admittance matrix $Y^{tt}(\omega)$ is quite complicated, but it is straightforward to obtain numerical results. We compare with the simplified description discussed above.

First we evaluate the initial values of the relaxation functions, as given by Eqs. (4.9) and (4.10) . This yields

$$
m_A R_{AA}^{t\dagger \parallel}(0+) = 0.6660, \quad m_A R_{BB}^{t\dagger \parallel}(0+) = 0.0832,
$$

$$
m_A R_{AB}^{t\dagger \parallel}(0+) = 0.0053,
$$
 (5.1)

where $m_A = 4\pi \rho a^3/3$. Note that the first value is close to $m_A/m_A^* = 2/3$, where m_A^* is the effective mass of a single sphere of radius *a*. The difference from unity is due to the added mass effect in an incompressible fluid. For a compressible fluid the value would be unity, as explained by Zwanzig and Bixon [\[21\]](#page-7-0). Similarly the second value is close to $m_A/m_B^* = 1/12$. The difference from zero of the third value is also due to the instantaneous transfer of momentum in an incompressible fluid.

Next we evaluate the mean relaxation times, as given by Eq. (4.11) . This yields

$$
\tau_{MAA} = 0.33385 \tau_{vA}, \quad \tau_{MBB} = 1.3354 \tau_{vA}
$$

$$
\tau_{Mdd} = 0.2828 \tau_{vA}, \qquad (5.2)
$$

with the viscous relaxation time $\tau_{vA} = a^2 \rho / \eta$. The first value is close to the single sphere value $m_A^*/(6\pi\eta a) = \tau_{vA}/3$, and the second value is close to the single sphere value *m*^{*}_{*B*}/(6πη*b*) = 4τ_{*vA*}/3.

The transforms $\Gamma_{AA}(z)$, $\Gamma_{BB}(z)$, and $\Gamma_{dd}(z)$ can be found from Eqs. (3.6) , (3.12) , and (4.7) with in each case the appropriate value $z = -i\omega \tau_{Mjj}$, with $j = A$, *B*, and *d*. The corresponding spectral densities $p_{AA}(u)$, $p_{BB}(u)$, and $p_{dd}(u)$ are found from Eq. [\(3.14\)](#page-3-0).

In Fig. 2 we plot the spectral density $p_{AA}(u)$ calculated from Eq. [\(3.14\)](#page-3-0) with $z = -i\omega\tau_{MAA}$. We compare with the two-pole expression $p_{2AA}(u)$ calculated for sphere *A* from Eq. [\(4.1\)](#page-3-0) with the parameter

$$
\sigma_{AA} = \sqrt{\frac{9m_{fA}}{2} C_{AA}^{\parallel}(0+)},
$$
\n(5.3)

FIG. 2. Plot of the spectral density $p_{AA}(u)$ for the two spheres *A* and *B* as specified at the beginning of Sec. V (solid curve). We compare with the two-pole approximation $p_{2AA}(u)$ (dashed curve).

FIG. 3. Plot of the spectral density $p_{BB}(u)$ for the two spheres *A* and *B* as specified at the beginning of Sec. [V](#page-4-0) (solid curve). We compare with the two-pole approximation $p_{2BB}(u)$ (dashed curve).

with $C_{AA}^{\parallel}(0+)$ given by Eq. [\(4.10\)](#page-4-0). In Fig. 3 we plot the spectral density $p_{BB}(u)$ calculated from Eq. [\(3.14\)](#page-3-0) with $z =$ $-i\omega\tau_{MBB}$. We compare with $p_{2BB}(u)$, calculated in the same manner with the parameter σ_{BB} . In both cases the agreement is nearly perfect. The spectral density covers several decades

We cannot expect the spectral density $p_{dd}(u)$ to be well approximated by a two-pole expression, since the expansion of the function $\Gamma_{dd}(z)$ in powers of \sqrt{z} does not have a term linear in \sqrt{z} , as seen from the second term in Eq. [\(4.12\)](#page-4-0) and the expression $|\psi_d\rangle = (1, -1)$ from Eq. [\(4.7\)](#page-3-0). Instead we compare with an expression of the form

$$
\Gamma_3(z) = \frac{1}{1 + z + \frac{Cz}{1 + D\sqrt{z}}},\tag{5.4}
$$

familiar from the theory of viscoelasticity of suspensions [\[22,23\]](#page-7-0), with $z = -i\omega\tau_{Mdd}$. This is a three-pole expression, characterized by three poles in the complex $y = \sqrt{z}$ plane. The parameters *C* and *D* can be found by comparison with the expansion of $\Gamma_{dd}(z)$ in powers of $y = \sqrt{z}$.

The expansion of $\Gamma_{dd}(z)$ in powers of *y*,

$$
\Gamma_{dd}(z) = 1 + c_2 y^2 + c_3 y^3 + O(y^4),\tag{5.5}
$$

yields the following for the coefficients *C* and *D*:

$$
C = -1 - c_2, \quad D = \frac{-c_3}{1 + c_2}.
$$
 (5.6)

For the coefficient c_2 , we find

$$
c_2 = \frac{-ab}{36\eta R^3} \frac{n_2}{a^3b + ab^3 - 3abR^2 + (a+b)R^3},
$$
 (5.7)

with the numerator

$$
n_2 = 2(a^2 + b^2)(a^2 + b^2 - 6R^2)(\rho_A a^3 + \rho_B b^3) + [(8\rho_A - 5\rho)a^4 + (8\rho_B - 5\rho)b^4 + 2a^2b^2(4\rho_A + 4\rho_B - 11\rho)]R^3
$$

+18(\rho_A a^3 + \rho_B b^3)R^4 + 6[(7\rho - 4\rho_A)a^2 + (7\rho - 4\rho_B)b^2]R^5 + 8[(\rho_A - 4\rho)a + (\rho_B - 4\rho)b]R^6 + 27\rho R^7, (5.8)

and for the coefficient c_3 , we find

in the dimensionless variable *u*.

$$
c_3 = \left(\frac{\rho}{\eta \tau_{Mdd}}\right)^{3/2} \frac{ab}{45} \frac{4a^5 + 4b^5 + 10a^2b^3 + 10a^3b^2 - 30(a^3 + b^3)R^2 - 9R^5}{a^3b + ab^3 - 3abR^2 + (a + b)R^3}.
$$
\n(5.9)

In Fig. 4 we plot the spectral density $p_{dd}(u)$ calculated from Eq. [\(3.14\)](#page-3-0) with $z = -i\omega\tau_{Mdd}$ and the relaxation time τ_{Mdd} given by Eq. (4.11) . We compare with $p_{3dd}(u)$, calculated in

FIG. 4. Plot of the spectral density $p_{dd}(u)$ for the two spheres *A* and *B* as specified at the beginning of Sec. [V](#page-4-0) (solid curve). We compare with the three-pole approximation $p_{3dd}(u)$ (dashed curve).

the same manner from $\Gamma_3(z)$ with the parameters *C* and *D*. In both cases the agreement is nearly perfect. The spectral density covers several decades in the dimensionless variable *u*. The low-frequency behavior of the transform $\Gamma_{dd}(z)$ corresponds to a $t^{-5/2}$ long-time tail of the function $C_{dd}(t)$.

In Fig. [5](#page-6-0) we plot the rate distributions $P_{AA}^{\parallel}(s)$, $P_{dd}^{\parallel}(s)$, and $P_{BB}^{\parallel}(s)$, multiplied by $C_{jj}^{\parallel}(0+)m_a/\tau_{vA}$ for $j = (A, d, B)$, as functions of $\log_{10}(s\tau_{vA})$, as given by Eq. [\(3.19\)](#page-3-0) for the respective mean relaxation times. It can be seen from the behavior of the spectra for small relaxation rates that the relaxation functions $R_{AA}^{\parallel}(t)$ and $R_{BB}^{\parallel}(t)$ have the same long-time behav-ior. In Fig. [6](#page-6-0) we plot $\log_{10}[m_A R^{\parallel}_{AA}(t)]$ and $\log_{10}[m_A R^{\parallel}_{dd}(t)]$ as functions of $\log_{10}(t/\tau_{vA})$. The first function has a $t^{-3/2}$ long-time tail, and the second one has a *t*−5/² long-time tail.

VI. DISCUSSION

In the above we studied velocity relaxation of two spheres immersed in a viscous incompressible fluid. In the simplest configuration the motion of both spheres is along the line of centers. The transfer of momentum after an initial impulse applied to one of the spheres is instantaneous due to

FIG. 5. Plot of the three rate distributions $P_{AA}^{\parallel}(s)$ (solid curve), $P_{dd}^{\parallel}(s)$ (short-dashed curve), and $P_{BB}^{\parallel}(s)$ (long-dashed curve), multiplied by $C_{jj}^{\parallel}(0+)m_a/\tau_{vA}$ for $j = (A, d, B)$, as functions of $\log_{10}(s\tau_{vA})$.

incompressibility. The initial values of the two velocities in this situation are expressed by an added mass matrix which shows a long-range dependence on the distance between centers. The relaxation at subsequent times is complicated and occurs due to viscous diffusion and interference of flow patterns. The velocities of both spheres decay eventually with the same *t*−3/² long-time tail.

The explicit calculations of the velocity relaxation functions are performed on the basis of an approximate expression

FIG. 6. Plot of the functions $log_{10}[m_A R_{AA}^{\parallel}(t)]$ (solid curve) and $\log_{10}[m_A R_{dd}^{\parallel}(t)]$ (dashed curve) as functions of $\log_{10}(t/\tau_{vA})$.

for the retarded scalar mobility function derived elsewhere [1]. We expect that for not too near distances the approximation provides accurate results which may be compared with experiments and computer simulations.

We showed that for collinear motion the relaxation functions can be described by a relatively small number of elementary modes with parameters which can be evaluated from the sphere properties. The relaxation functions are characterized conveniently by rate distributions. We expect that other aspects of velocity relaxation of two spheres in hydrodynamic interaction can be described in similar manner.

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