

Chiral symmetry breaking, color superconductivity, and equation of state for magnetized strange quark matter

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We investigate the vacuum structure of dense quark matter in strong magnetic fields in a three-flavor Nambu Jona Lasinio (NJL) model including the Kobayashi-Maskawa-t'Hooft (KMT) determinant term using a variational method. The method uses an explicit construct for the “ground” state in terms of quark-antiquark condensates as well as diquark condensates in the background of a constant magnetic field. The coupled mass gap equations and the superconducting gap equation are solved self-consistently and are used to compute the thermodynamic potential along with charge neutrality conditions imposed for bulk matter. Within the model, we observe inverse magnetic catalysis for chiral symmetry breaking for moderate magnetic fields. Further, we observe gapless modes in the presence of the magnetic field when charge neutrality conditions are imposed. The equation of state for charge neutral magnetized strange quark matter is derived, and found to be stiffer compared to the vanishing magnetic field counterpart. This could be relevant for gross structural properties of neutron stars.

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I. INTRODUCTION

The structure of vacuum in quantum chromodynamics (QCD) and its modification under extreme environment has been a major theoretical and experimental challenge in current physics [1]. In particular, it is interesting to study the modification of the structure of ground state at high temperature and/or high baryon densities as related to the nonperturbative aspects of QCD. This is important not only from a theoretical point of view, but also for many applications to problems of quark-gluon plasma (QGP) that could be copiously produced in relativistic heavy ion collisions as well as for the ultradense cold nuclear/quark matter which could be present in the interior of compact stellar objects like neutron stars. In addition to hot and dense QCD, the effect of strong magnetic field on QCD vacuum structure has attracted recent attention. This is motivated by the possibility of creating ultrastrong magnetic fields in noncentral collisions at RHIC and LHC. The strengths of the magnetic fields are estimated to be of hadronic scale [2,3] of the order of $eB \sim 2 m_\pi^2$ ($m_\pi^2 \simeq 10^{18}$ Gauss) at RHIC, to about $eB \sim 15 m_\pi^2$ at LHC [3]. There have been recent calculations

both analytic as well as with lattice simulations, which indicate that the QCD phase diagram is affected by strong magnetic fields [4–6].

In the context of cold dense matter, compact stars can be strongly magnetized. Neutron star observations indicate the magnetic field to be of the order of 10^{12} – 10^{13} Gauss at the surface of ordinary pulsars [7]. Further, the magnetars which are strongly magnetized neutron stars, may have even stronger magnetic fields of the order of 10^{15} – 10^{16} Gauss [8–14]. The physical upper limit on the magnetic field in a gravitationally bound star is 10^{18} Gauss which is obtained by comparing the magnetic and gravitational energies using virial theorem [7]. This limit could be higher for self-bound objects like quark stars [15]. Since the magnetic field strengths are of the order of QCD scale, this can affect both the thermodynamic as well as the hydrodynamics of such magnetized matter [16]. The phase structure of dense matter in the presence of the magnetic field along with a nonzero chiral density has been investigated for two-flavor Polyakov Loop extended Nambu-Jona Lasinio model (PNJL) model for high temperatures relevant for RHIC and LHC [17]. There have also been many investigations to look into the vacuum structure of QCD and it has been recognized that the strong magnetic field acts as a catalyzer of chiral symmetry breaking [18–22]. The effects of magnetic field on the equation of state have been recently studied in the Nambu Jona Lasinio model at zero temperature for three flavors and the equation of state has been computed for the cold quark matter [23,24] taking into account chiral condensate structure with the quark-antiquark pair for the ground state.

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On the other hand, color superconductivity is now an accepted conjectured state of cold and dense quark matter describing Cooper pairing of quarks of different colors and different flavors [25,26]. One can have a rigorous treatment of the phenomenon of such pairing using asymptotic freedom of QCD at very high densities. In its simplest form, when masses of the three quarks can be neglected compared to the chemical potential one can have the color flavor locked (CFL) phase [25,26]. However, to apply it to neutron star matter, the situation is more complicated as for the densities expected in the interior of neutron star, the masses of strange quarks cannot be neglected. Further, many nontrivial complications arise when beta equilibrium and charge neutrality conditions are imposed in such systems [27]. Since the well known sign problem prevents the first principle lattice simulations at finite chemical potentials, one has to rely on effective models at this regime of moderate densities. One model that has been extensively studied in this context has been the Nambu Jona Lasinio (NJL) model with contact interactions [28,29].

Of late, there has been a lot of attention on the investigation of color superconductivity in the presence of the magnetic field [18,19,30–32]. Essentially, this is due to its possible application in the astrophysical situations as the densities in compact star cores are large enough to have a possible superconducting phase as well as such compact stars can have a strong magnetic field as mentioned above. Let us also mention here that although such systems can be color superconductors, these phases can be penetrated by a “rotated” long range magnetic field. The corresponding rotated gauge field is a linear combination of vacuum photon field and the eighth gluon field [33,34]. These rotated magnetic fields are not subjected to the Meissner effect. While the Cooper pair is neutral with respect to the magnetic field, the quark quasiparticles have well-defined charges. Therefore, the pairing phenomenon is affected by the presence of the magnetic field. Initially, the effect of the magnetic field on superconducting phase has been studied for the CFL phase [30] where all three quarks take part in the pairing dynamics. However, for realistic densities, such symmetric pairing is disfavored due to large strange quark mass that leads to large mismatch in the Fermi surface. The condition of charge neutrality further complicates the pairing mechanism leading to gapless modes for homogeneous diquark pairing [35,36]. Superconductivity for the two-flavor quark matter in the presence of the magnetic field has been studied in Refs. [19,37,38] within the NJL model. The effect of charge neutrality along with the interplay of chiral and superconducting condensates has been analyzed in Refs. [37,38] in this model. A complete three-flavor analysis of magnetized dense quark matter including superconductivity has not been attempted so far. In the present investigation we include the effects of strange quarks that take part in chiral condensation but not in the diquark channel in the magnetized quark matter. As we

shall see, the strange quarks, similar to the vanishing magnetic field case, play an important role for charge neutral matter and the resulting equation of state. Moreover, with the inclusion of a flavor mixing interaction term, the strange quark scalar condensate not only affects the light quark condensates but also the diquark condensates.

We had earlier considered a variational approach to study chiral symmetry breaking as well as color superconductivity in hot and dense matter with an explicit structure for the “ground state” [36,39–41] with quark-antiquark condensate. The calculations were carried out within NJL model with minimization of free energy density to decide which condensate will exist at what density and/or temperature. A nice feature of the approach is that the four component quark field operator in the chiral symmetry broken phase gets determined from the vacuum structure. In the present work, we aim to investigate how the vacuum structure in the context of chiral symmetry breaking and color superconductivity gets modified in the presence of a magnetic field. In the context of chiral symmetry breaking, it was seen that, since the vacuum contains quark-antiquark pairs, the Dirac vacuum gets corrections due to the effective magnetic field apart from the modification of the medium or the Fermi sea of quarks. In our analysis we also keep these contributions to the equation of state.

We organize the paper as follows. In Sec. II, we discuss an ansatz state with quark-antiquark pairs related to chiral symmetry breaking, diquark and diantiquark pairs for the light flavors related to color superconductivity in the presence of a magnetic field. We then generalize such a state to include the effects of temperature and density. In Sec. III, we consider the three-flavor NJL model along with the so-called the Kobayashi-Maskawa-t’Hooft (KMT) term—the six fermion determinant interaction term which breaks $U(1)$ axial symmetry as in QCD. We use this Hamiltonian and calculate its expectation value with respect to the ansatz state to compute the energy density as well the thermodynamic potential for this system. We minimize the thermodynamic potential to determine the ansatz functions and the resulting mass gap equations. These coupled mass and superconducting gap equations are solved and we discuss the results in Sec. IV. We discuss here the results with and without constraints of charge neutrality. Finally we summarize and conclude in Sec. V. In the Appendix we give some details of the derivation of the evaluation of expectation values of the order parameters.

II. THE ANSATZ FOR THE GROUND STATE

Let us first consider the ground state structure relevant for chiral symmetry breaking in the presence of a strong magnetic field [24]. We shall then modify the same relevant for color superconductivity. To make the notations clear, we first write down the field operator expansion for quarks with a current quark mass m and charge q in the momentum space in the presence of a constant magnetic field \mathbf{B} .

We take the field direction to be along the z axis. We choose the gauge such that the electromagnetic vector potential is given as $A_\mu(\mathbf{x}) = (0, 0, Bx, 0)$. The quark field operator expansion in the presence of a constant magnetic field is given as [24,42]

$$\psi(\mathbf{x}) = \sum_n \sum_r \int \frac{d\mathbf{p}_\perp}{2\pi} [q_r^0(n, \mathbf{p}_\perp) U_r^0(x, \mathbf{p}_\perp, n) + \tilde{q}_r^0(n, -\mathbf{p}_\perp) V_r^0(x, -\mathbf{p}_\perp, n)] e^{i\mathbf{p}_\perp \cdot \mathbf{x}_\perp}. \quad (1)$$

Here n is the Landau level and the sum over it runs from 0 to infinity. In the above, $\mathbf{p}_\perp \equiv (p_y, p_x)$, and $r = \pm 1$ denotes the up and down spins. We have suppressed the color and flavor indices of the quark field operators. The quark annihilation and antiquark creation operators, q_r^0 and \tilde{q}_r^0 , respectively, satisfy the quantum algebra

$$\{q_r^0(n, \mathbf{p}_\perp), q_{r'}^{0\dagger}(n', \mathbf{p}'_\perp)\} = \{\tilde{q}_r^0(n, \mathbf{p}_\perp), \tilde{q}_{r'}^{0\dagger}(n', \mathbf{p}'_\perp)\} = \delta_{rr'} \delta_{nn'} \delta(\mathbf{p}_\perp - \mathbf{p}'_\perp). \quad (2)$$

In the above, U_r and V_r are the four component spinors for the quarks and antiquarks respectively. The explicit forms of the spinors for the fermions with mass m and electric charge q are given by

$$U_\uparrow^0(x, \mathbf{p}_\perp, n) = \begin{pmatrix} \cos \frac{\phi_0}{2} (\theta(q) I_n + \theta(-q) I_{n-1}) \\ 0 \\ \hat{p}_z \sin \frac{\phi_0}{2} (\theta(q) I_n + \theta(-q) I_{n-1}) \\ -i \hat{p}_\perp \sin \frac{\phi_0}{2} (\theta(q) I_{n-1} - \theta(-q) I_n) \end{pmatrix} \quad (3a)$$

$$U_\downarrow^0(x, \mathbf{p}_\perp, n) = \begin{pmatrix} 0 \\ \cos \frac{\phi_0}{2} (\theta(q) I_{n-1} + \theta(-q) I_n) \\ i \hat{p}_\perp \sin \frac{\phi_0}{2} (\theta(q) I_n - \theta(-q) I_{n-1}) \\ -\hat{p}_z \sin \frac{\phi_0}{2} (\theta(q) I_{n-1} + \theta(-q) I_n) \end{pmatrix} \quad (3b)$$

$$V_\uparrow^0(x, -\mathbf{p}_\perp, n) = \begin{pmatrix} \hat{p}_\perp \sin \frac{\phi_0}{2} (\theta(q) I_n - \theta(-q) I_{n-1}) \\ i \hat{p}_z \sin \frac{\phi_0}{2} (\theta(q) I_{n-1} + \theta(-q) I_n) \\ 0 \\ i \cos \frac{\phi_0}{2} (\theta(q) I_{n-1} + \theta(-q) I_n) \end{pmatrix} \quad (3c)$$

$$V_\downarrow^0(x, -\mathbf{p}_\perp, n) = \begin{pmatrix} i \hat{p}_z \sin \frac{\phi_0}{2} (\theta(q) I_n + \theta(-q) I_{n-1}) \\ \hat{p}_\perp \sin \frac{\phi_0}{2} (\theta(q) I_{n-1} - \theta(-q) I_n) \\ -i \cos \frac{\phi_0}{2} (\theta(q) I_n + \theta(-q) I_{n-1}) \\ 0 \end{pmatrix}. \quad (3d)$$

Here $\theta(x)$ is the Heaviside theta function. In the above, the energy of the n th Landau level is given as $\epsilon_n = \sqrt{m^2 + p_z^2 + 2n|q|B} \equiv \sqrt{m^2 + |p^2|}$ with $p^2 = p_z^2 + p_\perp^2$

so that $p_\perp^2 = 2n|q|B$, $\hat{p}_z = p_z/|p|$, $\hat{p}_\perp = \sqrt{2n|q|B}/|p|$. In Eqs. (3), $\cot \phi_0 = m/|p|$. Clearly, for vanishing masses $\phi_0 = \pi/2$. The functions I_n 's (with $n \geq 0$) are functions of $\xi = \sqrt{|qB|}(x - p_y/|qB|)$ and are given as

$$I_n(\xi) = c_n \exp\left(-\frac{\xi^2}{2}\right) H_n(\xi), \quad (4)$$

where $H_n(\xi)$ is the Hermite polynomial of the n th order and $I_{-1} = 0$. The normalization constant c_n is given by

$$c_n = \sqrt{\frac{\sqrt{|q|B}}{n! 2^n \sqrt{\pi}}}.$$

The functions $I_n(\xi)$ satisfy the orthonormality condition

$$\int d\xi I_n(\xi) I_m(\xi) = \sqrt{|q|B} \delta_{n,m}, \quad (5)$$

so that the spinors are properly normalized. The detailed derivation of these spinors and some of their properties are presented in the Appendix of Ref. [24].

With the field operators now defined in terms of the annihilation and the creation operators in the presence of a constant magnetic field, one can write down an ansatz for the ground state as in Ref. [24]. The ground state is taken as a squeezed coherent state involving quark and antiquarks pairs. Explicitly [24,36,39,41],

$$|\Omega\rangle = \mathcal{U}_Q |0\rangle. \quad (6)$$

Here, \mathcal{U}_Q is a unitary operator which creates quark-antiquark pairs from the vacuum $|0\rangle$ which is annihilated by the quark/antiquark annihilation operators given in Eq. (1). Explicitly, the operator \mathcal{U}_Q is given as [24]

$$\mathcal{U}_Q = \exp\left(\sum_{n=0}^{\infty} \int d\mathbf{p}_\perp q_r^{0i\dagger}(n, \mathbf{p}_\perp) a_{r,s}^i(n, p_z) h^i(n, p_z) \times \tilde{q}_s^{0i}(n, -\mathbf{p}_\perp) - \text{H.c.}\right). \quad (7)$$

In the above ansatz for the ground state, the function $h^i(n, p_z)$ is a real function describing the quark-antiquark condensates related to the vacuum realignment for chiral symmetry breaking to be obtained from a minimization of the thermodynamic potential. In the above equation, the spin dependent structure $a_{r,s}^i$ is given by

$$a_{r,s}^i = \frac{1}{|\mathbf{p}_i|} [-\sqrt{2n|q_i|B} \delta_{r,s} - i p_z \delta_{r,-s}], \quad (8)$$

with $|\mathbf{p}_i| = \sqrt{p_z^2 + 2n|q_i|B}$ denoting the magnitude of the three momentum of the quark/antiquark of i th flavor (with electric charge q_i) in the presence of a magnetic field.

Summation over three colors and three flavors is understood in the exponent of \mathcal{U}_Q in Eq. (7). Clearly, a nontrivial $h_i(n, p_z)$ breaks the chiral symmetry.

It is easy to show that the transformation of the ground state as in Eq. (6) is a Bogoliubov transformation. With the ground state transforming as Eq. (6), any operator O^0 in the $|0\rangle$ basis transforms as

$$O = \mathcal{U}_Q O^0 \mathcal{U}_Q^\dagger, \quad (9)$$

and, in particular, one can transform the creation and annihilation operators of Eq. (1) to define the transformed operators as above satisfying the same anticommutation relations as in Eq. (2):

$$\begin{aligned} \psi(\mathbf{x}) = & \sum_n \sum_r \frac{1}{2\pi} \int d\mathbf{p}_\chi [q_r(n, \mathbf{p}_\chi) U_r(x, n, \mathbf{p}_\chi) \\ & + \tilde{q}_r(n, -\mathbf{p}_\chi) V_r(x, n, -\mathbf{p}_\chi)] e^{i\mathbf{p}_\chi \cdot \mathbf{x}}, \end{aligned} \quad (10)$$

with $q_r|\Omega\rangle = 0 = \tilde{q}_r^\dagger|\Omega\rangle$. In the above, we have suppressed the flavor and color indices. It is easy to see that the form of U, V spinors is exactly similar to the form of the spinors U_0, V_0 as in Eq. (3) but with the shift of the function $\phi_0 \rightarrow \phi = \phi_0 - 2h$ with the function $h(\mathbf{k})$ to be determined by a minimization of free energy. As we shall see later, it is more convenient to vary $\phi(\mathbf{k})$ rather than $h(\mathbf{k})$. Let us note that with Eq. (10), the four component quark field operator gets defined in terms of the vacuum structure for chiral symmetry breaking given through Eq. (6) and Eq. (7) in the presence of a magnetic field [43].

The chiral order parameter in the condensate vacuum $|\Omega\rangle$ can be evaluated explicitly using the field operator expansion given in Eq. (10) and is given by [24] (for i th flavor)

$$I_s^i = \langle \Omega | \bar{\psi}^i \psi^i | \Omega \rangle = - \sum_n N_c \alpha_n \frac{|q_i B|}{(2\pi)^2} \int dp_z \cos \phi^i. \quad (11)$$

This expression for the quark-antiquark condensate is exactly the same form as derived earlier in the absence of the magnetic field [39,40] once one realizes that in the presence of a quantizing magnetic field with discrete Landau levels, one has for the phase space integration [37]

$$\int \frac{d\mathbf{p}}{(2\pi)^3} \rightarrow \sum_{n=0}^{\infty} \alpha_n \frac{|qB|}{(2\pi)^2} \int dp_z.$$

Next, we would like to generalize the ansatz of Eq. (6) with quark-antiquark pairs in the presence of a magnetic field, to include quark-quark pairs for the description of the ground state as relevant for color superconductivity. However, few comments in this context are in order. It is known that in the presence of color superconductivity, the diquark is electromagnetically charged and the usual

magnetic field will have a Meissner effect. However, a linear combination of the photon field and the gluon field given by $\tilde{A}_\mu = \cos \alpha A_\mu - \sin \alpha G_\mu^8$ still remains massless and is unscreened. For two-flavor color superconductivity, $\cos \alpha = g/\sqrt{g^2 + e^2/3} \sim 1/20$ [33]. The electron couples to this rotated gauge field by the coupling $\tilde{e} = e \cos(\alpha)$. The quark field couples to the rotated gauge field through its rotated charge \tilde{Q} . In units of \tilde{e} , the rotated charge matrix in the flavor-color space is given by

$$\tilde{Q} = Q_f \otimes \mathbf{1}_c - \mathbf{1}_f \otimes \frac{T_c^8}{2\sqrt{3}}. \quad (12)$$

Thus, the \tilde{e} charges of red and green u quarks is $1/2$; red and green down and strange quarks is $-1/2$. The blue u quark has \tilde{Q} charge as $+1$, while the blue d and s quarks are \tilde{Q} chargeless. We shall take the rotated U(1) magnetic field along the z axis and spatially constant as before without the absence of superconductivity. The ansatz for the ground state with quark-antiquark condensate is now taken as, with i being the flavor index,

$$|\Omega\rangle_\chi = \exp \sum_{flav} (B_i^\dagger - B_i) |0\rangle. \quad (13)$$

The flavor dependent quark-antiquark pair creation operator for u quark ($i = 1$) is given as, with $a = 1, 2, 3$ being the color indices for red, blue and green respectively,

$$B_u^\dagger = \sum_{a=1}^3 \sum_{n=0}^{\infty} \int d\mathbf{p}_\chi q_r^{1a}(n, \mathbf{p}_\chi)^\dagger a_{r,s}^1(n, p_z) f^{1a}(n, \mathbf{p}_\chi) \tilde{q}_s^{1a}(n, -\mathbf{p}_\chi), \quad (14)$$

while, for the down and strange quarks ($i = 2, 3$) the same is given as

$$\begin{aligned} B_i^\dagger = & \sum_{a=1}^2 \sum_{n=0}^{\infty} \int d\mathbf{p}_\chi q_r^{ia}(n, \mathbf{p}_\chi)^\dagger a_{r,s}^i(n, p_z) h^{ia}(n, \mathbf{p}_\chi) \tilde{q}_s^{ia}(n, -\mathbf{p}_\chi) \\ & + \int d\mathbf{p} q_r^{i3}(\mathbf{p})^\dagger (\sigma \cdot \hat{\mathbf{p}})_{rs} h^i(\mathbf{p}) \tilde{q}_s^{i3}(-\mathbf{p}). \end{aligned} \quad (15)$$

The difference between the pair creation operator in Eqs. (14) and (15) lies on the contribution of the blue color. While the up blue quark has \tilde{Q} charge, the blue quarks of down and strange quark are \tilde{Q} neutral.

Next, we write down the ansatz state for having quark-quark condensates which is given by

$$|\Omega\rangle = U_d |\Omega\rangle_\chi \equiv \exp(B_d^\dagger - B_d) |\Omega\rangle_\chi. \quad (16)$$

In the above, B_d^\dagger is the diquark (and di-antiquark) creation operator given as

$$B_d^\dagger = \sum_n \int dp_\chi [q_r^{ia}(n, p_\chi)^\dagger r f(n, p_z) q_{-r}^{jb}(n, -p_\chi, p_z) + i \tilde{q}_r^{ia}(n, p_\chi)^\dagger r f_1(n, p_z) \tilde{q}_{-r}^{jb}(n, p_\chi)^\dagger] e^{ij3} e^{3ab}. \quad (17)$$

In the above, i, j are the flavor indices, a, b are the color indices and $r = \pm 1/2$ are the spin indices. The Levi-Civita tensor ensures that the operator is antisymmetric in color and flavor space along with the fact that only u, d quarks with red and green colors take part in diquark condensation. The blue u, d quarks as well as the strange quarks (all three colors) do not take part in the diquark condensation. The functions $f(n, p_z)$ and $f_1(n, p_z)$ are condensate functions associated with quark-quark and antiquark-antiquark condensates respectively. These functions are assumed to be independent of color and flavor indices. We shall give a *post facto* justification for this that these functions depend upon the average energy and average chemical potentials of the quarks that condense.

To include the effects of temperature and density we next write down the state at finite temperature and density $|\Omega(\beta, \mu)\rangle$ through a thermal Bogoliubov transformation over the state $|\Omega\rangle$ using the thermofield dynamics (TFD) method as described in Refs. [24,44,45]. This is particularly useful while dealing with operators and expectation values. We write the thermal state as

$$|\Omega(\beta, \mu)\rangle = \mathcal{U}_{\beta, \mu} |\Omega\rangle = \mathcal{U}_{\beta, \mu} \mathcal{U}_Q |0\rangle, \quad (18)$$

where $\mathcal{U}_{\beta, \mu}$ is given as

$$\mathcal{U}_{\beta, \mu} = e^{\mathcal{B}^\dagger(\beta, \mu) - \mathcal{B}(\beta, \mu)},$$

with

$$\mathcal{B}^\dagger(\beta, \mu) = \sum_{n=0}^{\infty} \int [dk_\chi q_r^{ia}(n, k_\chi)^\dagger \theta_{-}^{ia}(k_z, n, \beta, \mu) \underline{q}_r^{ia}(n, k_\chi)^\dagger + \tilde{q}_r^{ia}(n, k_\chi) \theta_{+}^{ia}(k_z, n, \beta, \mu) \tilde{\underline{q}}_r^{ia}(n, k_\chi)]. \quad (19)$$

In Eq. (19), the underlined operators are the operators in the extended Hilbert space associated with thermal doubling in the TFD method, and the color flavor dependent ansatz functions $\theta_{\pm}^{ia}(n, k_z, \beta, \mu)$ are related to quark and antiquark distributions as can be seen through the minimization of the thermodynamic potential.

All the functions in the ansatz in Eq. (18) are to be obtained by minimizing the thermodynamic potential. We shall carry out this minimization in the next section. However, before carrying out the minimization procedure, let us focus our attention to the expectation values of some known operators to show that with the above variational ansatz for the ground state given in Eq. (18) these reduce to the already known expressions in the appropriate limits.

Let us first consider the expectation value of the chiral order parameter. The expectation value for chiral order parameter for the i th flavor is given as

$$I_s^i = \langle \Omega(\beta, \mu) | \bar{\psi}_i \psi_i | \Omega(\beta, \mu) \rangle = \sum_{a=1}^3 I_s^{ia}. \quad (20)$$

These expectation values can be evaluated easily once we realize that the state $|\Omega(\beta, \mu)\rangle$ as in Eq. (18) is obtained through successive Bogoliubov transformations on the state $|0\rangle$ as in Eqs. (13) and (16). The details of evaluation for the different order parameters is relegated to the Appendix. Explicitly, for the quarks that take part in superconductivity

$$I_s^{ia} = - \sum_n \alpha_n \frac{|q^{ia} B|}{(2\pi)^2} \times \int dp_z \cos \phi^{ia} (1 - F^{ia} - F_1^{ia}), \quad (i, a = 1, 2), \quad (21)$$

where $\alpha_n = (2 - \delta_{n,0})$ is the degeneracy factor of the n th Landau level (all levels are doubly degenerate except the lowest Landau level). Further,

$$F^{ia} = \sin^2 \theta_{-}^{ia} + \sin^2 f (1 - \sin^2 \theta_{-}^{ia} - |e^{ij}| e^{ab} |\sin^2 \theta_{-}^{jb}|), \quad (22)$$

arising from the quarks which condense and

$$F_1^{ia} = \sin^2 \theta_{+}^{ia} + \sin^2 f_1 (1 - \sin^2 \theta_{+}^{ia} - |e^{ij}| e^{ab} |\sin^2 \theta_{+}^{jb}|), \quad (23)$$

arising from antiquarks which condense. Thus, the scalar condensates arising from quarks that take part in superconductivity depend both on the condensate functions in quark-antiquark channel (ϕ^i) as well as in quark-quark channel (f, f_1). Further, the thermal functions $\sin^2 \theta_{\pm}^{ia}$, as we shall see later, will be related to the number density distribution functions.

Next, for the nonsuperconducting blue up quarks, the contribution to the scalar condensate is given by

$$I_s^{13} = - \sum_n \alpha_n \frac{|q^{13} B|}{(2\pi)^2} \int dp_z \cos \phi^{13} (1 - \sin^2 \theta_{-}^{13} - \sin^2 \theta_{+}^{13}). \quad (24)$$

Let us note that in the limit of vanishing of the color superconducting condensate functions ($f, f_1 \rightarrow 0$), the contributions given in Eq. (21) reduce to Eq. (24) as they should [24].

Similarly, scalar condensate contribution from the charged strange quarks (red, green) is given by

$$I_s^{3a} = - \sum_n \alpha_n \frac{|q^{3a} B|}{(2\pi)^2} \times \int dp_z \cos \phi^{3a} (1 - \sin^2 \theta_{-}^{3a} - \sin^2 \theta_{+}^{3a}) \quad (a = 1, 2). \quad (25)$$

Finally, for the uncharged quarks, i.e., blue down and blue strange quarks, the contributions to the scalar condensates are given by, for flavor i ($i = 2, 3$),

$$I_s^{i3} = -\frac{2}{(2\pi)^3} \int d\mathbf{k} \cos \phi^i (1 - \sin^2 \theta_-^{i3} - \sin^2 \theta_+^{i3}). \quad (26)$$

Next, we write down the condensate in the superconducting channel which is given as

$$\begin{aligned} I_D &= \langle \bar{\psi}_c^{ia} \gamma^5 \psi^{jb} \rangle \epsilon^{ij} \epsilon^{3ab} \\ &= \sum_n \alpha_n \frac{|q_i B|}{(2\pi)^2} \int dp_z \cos\left(\frac{\phi_1 - \phi_2}{2}\right) \\ &\quad \times [\sin 2f(1 - \sin^2 \theta_-^1 - \sin^2 \theta_-^2) \\ &\quad + \sin 2f_1(1 - \sin^2 \theta_+^1 - \sin^2 \theta_+^2)]. \end{aligned} \quad (27)$$

Let us note that the superconducting condensate also depends upon the chiral condensate functions $\phi(p_z)$ through the function $\cos(\frac{\phi_1 - \phi_2}{2})$ apart from the thermal distribution functions $\sin^2 \theta_{\pm}^{ia}$. Further, this dependence vanishes when the u and d quark scalar condensates or equivalently the corresponding masses of the quarks are equal.

The other quantity that we wish to investigate is the axial fermion current density that is induced at finite chemical potential including the effect of temperature. The expectation value of the axial current density is given by

$$\langle j_5^3 \rangle \equiv \langle \bar{\psi}_i^a \gamma^3 \gamma^5 \psi_j^a \rangle.$$

Using the field operator expansion Eq. (10) and Eq. (3) for the explicit forms for the spinors, we have for the i th flavor

$$\langle j_5^3 \rangle = \sum_n \frac{N_c}{(2\pi)^2} \int dp_x (I_n^2 - I_{n-1}^2) (\sin^2 \theta_-^i - \sin^2 \theta_+^i). \quad (28)$$

Integrating over dp_y using the orthonormal condition of Eq. (5), all the terms in the above sum for the Landau levels cancel out except for the zeroth Landau level so that

$$\langle j_5^3 \rangle = \frac{N_c |q_i B|}{(2\pi)^2} \int dp_z [\sin^2 \theta_-^{i0} - \sin^2 \theta_+^{i0}], \quad (29)$$

which is identical to that in Ref. [46] once we identify the functions $\sin^2 \theta_{\pm}^{i0}$ as the particle and the antiparticle distribution functions for the zero modes [see e.g., Eq. (55) in the next section]. In the chiral limit at zero temperature and without superconductivity, one gets the following as the axial current after summing over all three flavors:

$$\langle j_5^0 \rangle = \frac{3eB}{2\pi^2} \left[\mu + \frac{1}{3} \sqrt{\mu^2 - m_s^2} \right]. \quad (30)$$

III. EVALUATION OF THERMODYNAMIC POTENTIAL AND GAP EQUATIONS

As has already been mentioned, we shall consider in the present investigation, the three-flavor Nambu Jona Lasinio model including the Kobayashi-Maskawa-t-Hooft (KMT) determinant interaction. The corresponding Hamiltonian density is given as [24,28,36,47]

$$\begin{aligned} \mathcal{H} &= \psi^\dagger (-i\boldsymbol{\alpha} \cdot \boldsymbol{\Pi}) + \gamma^0 \hat{m} \psi \\ &\quad - G_S \sum_{A=0}^8 [(\bar{\psi} \lambda^A \psi)^2 - (\bar{\psi} \gamma^5 \lambda^A \psi)^2] \\ &\quad + K [\det_f [\bar{\psi} (1 + \gamma_5) \psi] + \det_f [\bar{\psi} (1 - \gamma_5) \psi]] \\ &\quad - G_D [(\bar{\psi} \gamma^5 \epsilon \epsilon_c \psi^C) (\bar{\psi}^C \gamma^5 \epsilon \epsilon_c \psi)], \end{aligned} \quad (31)$$

where $\psi^{i,a}$ denotes a quark field with color “ a ” ($a = r, g, b$), and flavor “ i ” ($i = u, d, s$), indices. $\Pi = -i(\nabla - i\tilde{e}\tilde{A}\tilde{Q})$ is the canonical momentum in the presence of the rotated U(1) gauge field \tilde{A}_μ . ϵ is the Levi-Civita tensor in flavor space while ϵ_c is the Levi-Civita tensor in color space. $\psi^C = i\gamma^1 \gamma^2 \psi$ is the charge conjugate spinor. When there is no superconductivity $A_\mu = \tilde{A}_\mu$ which is the usual massless photon field with the coupling to the quark field being given the electromagnetic charge eQ_f , where Q_f is diagonal matrix $(2/3, -1/3, -1/3)$. As mentioned in the previous section, when the superconducting gap is nonvanishing, the massless gauge field is given by $\tilde{A}_\mu = \cos \alpha A_\mu - \sin \alpha G_\mu^8$, where $\cos \alpha = g/\sqrt{g^2 + e^2/3}$. We have taken here the standard convention of $SU(3)_c$ generators in the adjoint representation [33]. The \tilde{Q} charges of the quarks are given in Table I. It may also be relevant here to mention that, while we are taking into account combination of the photon and gluon field which is massless, the other orthogonal massive component, is either Meissner screened or nucleated into vortices [48].

The matrix of current quark masses is given by $\hat{m} = \text{diag}_f(m_u, m_d, m_s)$ in the flavor space. We shall assume in the present investigation, isospin symmetry with $m_u = m_d$.

TABLE I. Table: List of quarks and their electromagnetic and rotated charges.

Quark	e-charge	\tilde{e} -charge
u-red	$\frac{2}{3}$	$\frac{1}{2}$
u-green	$\frac{2}{3}$	$\frac{1}{2}$
u-blue	$\frac{2}{3}$	1
d-red	$-\frac{1}{3}$	$-\frac{1}{2}$
d-green	$-\frac{1}{3}$	$-\frac{1}{2}$
d-blue	$-\frac{1}{3}$	0
s-red	$-\frac{1}{3}$	$-\frac{1}{2}$
s-green	$-\frac{1}{3}$	$-\frac{1}{2}$
s-blue	$-\frac{1}{3}$	0

In Eq. (31), λ^A , $A = 1, \dots, 8$ denote the Gellmann matrices acting in the flavor space and $\lambda^0 = \sqrt{\frac{2}{3}}\mathbf{1}_f$, $\mathbf{1}_f$ as the unit matrix in the flavor space. The four point interaction term $\sim G_S$ is symmetric in $SU(3)_V \times SU(3)_A \times U(1)_V \times U(1)_A$. In contrast, the determinant term $\sim K$ which for the case of three flavors generates a six point interaction which breaks $U(1)_A$ symmetry. In the absence of magnetic field, if the mass term is neglected, the overall symmetry is $SU(3)_V \times SU(3)_A \times U(1)_V$. This spontaneously breaks to $SU(3)_V \times U(1)_V$ implying the conservation of the baryon number and the flavor number. The current quark mass term introduces additional explicit breaking of chiral symmetry leading to partial conservation of the axial current. The last term in Eq. (31) describe a scalar diquark interaction in the color antitriplet and flavor antitriplet channel. Such a form of four point interaction can arise e.g., by Fierz transformation of a four point vector current-current interaction having quantum numbers of a single gluon exchange. In that case the diquark coupling G_D is related to the scalar coupling as $G_D = 0.75G_S$.

Next we evaluate the expectation value of the kinetic term in Eq. (31) which is given as

$$T = \langle \Omega(\beta, \mu) | \psi^{ia\dagger} (-i\boldsymbol{\alpha} \cdot \nabla - \tilde{q}^{ia} B x \alpha_2) \psi^{ia} | \Omega(\beta, \mu) \rangle \\ \equiv \sum_{ia} T^{ia}. \quad (32)$$

In the above the sum over the colors a and flavors i is understood. The color flavor dependent charges

\tilde{q}^{ia} for the quasiparticles is given in Table I. To evaluate this, for nonvanishing \tilde{q} charges, we use Eq. (10) and the results of spatial derivatives on the functions $I_n(\xi)$ ($\xi = \sqrt{|q_i|B}(x - p_y/(|q_i|B))$):

$$\frac{\partial I_n}{\partial x} = \sqrt{|q^{ia}|B} [-\xi I_n + \sqrt{2n} I_{n-1}], \\ \frac{\partial I_{n-1}}{\partial x} = \sqrt{|\tilde{q}^{ia}|B} [-\xi I_{n-1} + \sqrt{2(n-1)} I_{n-2}]. \quad (33)$$

Using the above, a straightforward but somewhat tedious manipulations lead to the contribution arising from the quarks that take part in superconductivity, i.e., for color, flavor indices $i, a = 1, 2$,

$$T^{ia} = - \sum_{n=0}^{\infty} \alpha_n \frac{|\tilde{e}B|}{2(2\pi)^2} \int dp_z (m_i \cos \phi_i + |p_i| \sin \phi_i) \\ \times (1 - F^{ia} - F_1^{ia}), \quad (i, a = 1, 2), \quad (34)$$

where we have defined $|p_i|^2 = p_z^2 + 2n|\tilde{q}B|$, ($\tilde{q} = \tilde{e}/2$). Here, the quark-antiquark condensate effects are encoded in the function ϕ_i while diquark and di-antiquark condensate effects are encoded in the functions F^{ia} and F_1^{ia} respectively as given in Eqs. (22) and (23).

For the blue u quark, which is charged but does not take part in diquark condensation the corresponding contribution to the kinetic term is given by

$$T^{13} = - \sum_{n=0}^{\infty} \alpha_n \frac{|\tilde{e}B|}{(2\pi)^2} \int dp_z (m_1 \cos \phi_1 + |p_1| \sin \phi_1) (1 - \sin^2 \theta_-^{13} - \sin^2 \theta_+^{13}). \quad (35)$$

The contribution of the charged strange quarks (with charges $\tilde{e}/2$) to the kinetic energy is given by, with $a = 1, 2$,

$$T^{3a} = - \sum_{n=0}^{\infty} \alpha_n \frac{|\tilde{e}B|}{2(2\pi)^2} \int dp_z (m_3 \cos \phi_3 + |p_3| \sin \phi_3) (1 - \sin^2 \theta_-^{3a} - \sin^2 \theta_+^{3a}). \quad (36)$$

Finally, the contribution from the \tilde{e} -charge neutral quarks (blue d and blue s) is given as

$$T^{i3} = - \int \frac{d\mathbf{p}}{(2\pi)^3} (m_i \cos \phi_i + p \sin \phi_i) (1 - \sin^2 \theta_-^{i3} - \sin^2 \theta_+^{i3}) \quad (i = 2, 3). \quad (37)$$

The contribution to the energy density from the quartic interaction term in Eq. (31), using Eq. (20) turns out to be

$$V_S \equiv -G_S \langle \Omega(\beta, \mu) | \sum_{A=0}^8 [(\bar{\psi} \lambda^A \psi)^2 - (\bar{\psi} \gamma^5 \lambda^A \psi)^2] | \Omega(\beta, \mu) \rangle \\ = -2G_S \sum_{i=1,3} I_i^2, \quad (38)$$

where $I_i^i = \langle \bar{\psi}_i \psi_i \rangle$ is the scalar quark-antiquark condensate given in Eq. (20). Further, in the above,

we have used the properties of the Gellman matrices $\sum_{A=0}^8 \lambda_{ij}^A \lambda_{kl}^A = 2\delta_{il}\delta_{jk}$.

Next, let us discuss the contribution from the six quark determinant interaction term to the energy expectation value. There will be six terms in the expansion of the determinant, each involving three pairs of quark operators of different flavors. These are to be ‘‘contracted’’ in all possible manner while taking the expectation value. This means in the present context of having quark-antiquark and diquark condensates, one can contract a ψ with a $\bar{\psi}$

or ψ with a $\bar{\psi}$. The former leads to condensates having quark-antiquark condensates $I_s^{(i)}$ while the latter leads to diquark condensates I_D . Further, for the case of quark-antiquark condensate contributions, the contracting ψ and $\bar{\psi}$ having the same color will lead to the dominant contribution while contracting similar operators with different colors will lead to a N_c suppressed contribution. Next, regarding the contributions arising from the diquark, terms which are proportional to strange quark-antiquark condensate $\langle \bar{s}s \rangle$ will be dominant. These will have the contractions of strange quark-antiquarks having the same color. The rest of the four terms will be suppressed at least by a factor N_c . Explicitly these two terms are given by $\sim \sum_h \bar{s} O^h_s [\bar{u} \hat{O}^h u \times (\bar{d} \hat{O}^h d) - \bar{u} \hat{O}^h d \times (\bar{d} \hat{O}^h u)]$, where $h = \pm$ and $\hat{O}^\pm = (1 \pm \gamma_5)$. When contracted diquark wise, both terms give identical contributions, except that the contribution of the second term will be of opposite sign as compared to the first term. This is a consequence of the flavor antisymmetric nature of the diquark condensates. This leads to

$$\begin{aligned} V_{\text{det}} &= +K \langle \det_f [\bar{\psi}(1 + \gamma_5)\psi] + \det_f [\bar{\psi}(1 - \gamma_5)\psi] \rangle \\ &= \frac{1}{3} |\epsilon_{ijk}| I_s^{(i)} I_s^{(j)} I_s^{(k)} + \frac{K}{4} I_s^{(3)} I_D^2. \end{aligned}$$

Next, the contribution from the diquark interaction is given by

$$V_D = -\langle G_D [(\bar{\psi} \gamma^5 \epsilon \epsilon_c \psi^C)(\bar{\psi}^C \gamma^5 \epsilon \epsilon_c \psi)] \rangle = -G_D I_D^2, \quad (39)$$

where the diquark condensate I_D is already defined in Eq. (27).

To calculate the thermodynamic potential (negative of the pressure), we also have to specify the chemical potentials relevant for the system. Here, we shall be interested in the form of quark matter that might be present in compact stars that are older than a few minutes so that chemical equilibration for weak interaction is satisfied. The relevant chemical potentials in such a case are the baryon chemical potential $\mu_B = 3\mu_q$, the chemical potential μ_E associated with the electromagnetic charge, and, the color potentials μ_3 and μ_8 . The chemical potential is a matrix that is diagonal in color and flavor space and is given by

$$\mu_{ij,ab} = (\mu \delta_{ij} + Q_{ij} \mu_E) \delta_{ab} + (T_{ab}^3 \mu_3 + T_{ab}^8 \mu_8) \delta_{ij}. \quad (40)$$

Since red and green color of a given flavor of quark is degenerate and the diquark is in the blue direction in the color space, we can assume $\mu_3 = 0$.

The thermodynamic potential, Ω , is then given by using Eqs. (32), (38), and (39), with s being the entropy density,

$$\Omega = T + V_S + V_{\text{det}} + V_D - \langle \mu N \rangle - \frac{1}{\beta} s, \quad (41)$$

where we have introduced

$$\langle \mu N \rangle = \langle \psi^{ia\dagger} \mu_{ij,ab} \psi^{jb} \rangle = \sum_{i,a} \mu^{ia} \rho^{ia}, \quad (42)$$

where ρ^{ia} is the vector density $\rho^{ia} = \langle \psi^{ia\dagger} \psi^{ia} \rangle$. For the superconducting quarks this is given by

$$\rho^{ia} = \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z (F^{ia} - F_1^{ia}), \quad (i, a = 1, 2), \quad (43)$$

while for the blue u quark, the same is given by

$$\rho^{13} = \sum_n \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \int dp_z (\sin^2 \theta_-^{13} - \sin^2 \theta_+^{13}). \quad (44)$$

For the charged strange quarks, this density is given by

$$\rho^{3a} = \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z (\sin^2 \theta_-^{3a} - \sin^2 \theta_+^{3a}), \quad (a = 1, 2). \quad (45)$$

For the \tilde{e} -uncharged quarks (blue down and blue strange), the vector density is given by

$$I_v^{i3} = \frac{2}{(2\pi)^3} \int d\mathbf{p} (\sin^2 \theta_-^{i3} - \sin^2 \theta_+^{i3}), \quad (i = 2, 3). \quad (46)$$

Finally, the entropy density is given by $s = \sum_{i,a} s^{ia}$, where s^{ia} is the entropy density for quarks of flavor i and color a . For the quarks with charge \tilde{q}^{ia} , the phase space is Landau quantized and we have the entropy density given as [44]

$$s^{ia} = -\sum_n \alpha_n \frac{|q^{ia}| B}{(2\pi)^2} \int dp_z \{ (\sin^2 \theta_-^{ia} \ln \sin^2 \theta_-^{ia} + \cos^2 \theta_-^{ia} \ln \cos^2 \theta_-^{ia}) + (- \rightarrow +) \}. \quad (47)$$

On the other hand, for the uncharged (blue down and blue strange) quarks, the entropy density is given by

$$s^{i3} = -\frac{2}{(2\pi)^3} \int d\mathbf{p} \{ (\sin^2 \theta_-^{i3} \ln \sin^2 \theta_-^{i3} + \cos^2 \theta_-^{i3} \ln \cos^2 \theta_-^{i3}) + (- \rightarrow +) \}, \quad (i = 2, 3). \quad (48)$$

Thus, the thermodynamic potential is now completely defined in terms of the condensate functions ϕ^i , $f(k)$ and the thermal distribution functions θ_{\pm}^{ia} which will be determined through a functional extremization of the thermodynamic potential. Minimizing the thermodynamic potential with respect to the quark-antiquark condensate function $\phi_i(p)$, i.e., $\delta\Omega/\delta\phi_i = 0$, leads to

$$\cot\phi^{ia} = \frac{(m_i - 4G_S I_s^i + K\epsilon^{ijk} I_s^j I_s^k + \delta_{i3} \frac{k}{4} I_D^2)}{|p_{ia}|} \equiv \frac{M_i}{|p_{ia}|}, \quad (49)$$

where, as earlier, we have defined $|p_{ia}| = \sqrt{p_z^2 + 2n|q_{ia}|B}$ and we have defined the constituent quark mass $M_i = m_i - 4G_S I_s^i + K|\epsilon_{ijk}| I_s^{(i)} I_s^{(j)} I_s^{(k)} + \delta^{i3} I_D^2 \frac{k}{4}$. These expressions are actually self-consistent equations for the constituent quark masses as scalar condensate $I_s^{(i)}$ as given in Eq. (20) involve M_i through their dependence on ϕ_i . Explicitly, these mass gap equations are given as

$$M^u = m^u - 4G_S I_s^{(u)} + 2K I_s^{(d)} I_s^{(s)}, \quad (50)$$

$$M^d = m^d - 4G_S I_s^{(d)} + 2K I_s^{(u)} I_s^{(s)}, \quad (51)$$

$$M^s = m^s - 4G_S I_s^{(s)} + 2K I_s^{(d)} I_s^{(u)} + \frac{K}{4} I_D^2. \quad (52)$$

Let us note that while the color and flavor dependence on the quark-antiquark condensate functions ϕ^{ia} arises *only* from the momentum $|p_{ia}| = \sqrt{p_z^2 + 2n|\tilde{q}_{ia}|B}$ through the color flavor dependent \tilde{q} charges, the constituent quark masses are color singlets and are given by the solutions of the self-consistent equations (50)–(52). Further, the flavor mixing determinant interaction makes the masses of quark of a given flavor dependent upon the condensates of the other flavor quarks. This apart, the strange quark mass explicitly depends upon the diquark condensates through this determinant interaction. Note that for the two flavor superconductivity as considered here, the strange quark mass is affected explicitly by the superconducting gap given by the last term on the right-hand side Eq. (52). Of course, there is implicit dependence on the superconducting gap in the second term through the functions F and F_1 [given in Eqs. (22) and (23)]. Further, when chiral symmetry is restored for the light quarks, i.e., when the scalar condensates for the nonstrange quarks vanish, still, the determinant term gives rise to a density dependent dynamical strange quark mass arising from diquark condensates of the light quarks [47]. Such a mass generation is very different from the typical mechanism of quark mass generation through quark-antiquark condensates [49].

In a similar manner, minimizing the thermodynamic potential with respect to the diquark function $f(k)$ and di-antiquark function $f_1(k)$, i.e., $\frac{\delta\Omega}{\delta f(k)} = 0$ and $\frac{\delta\Omega}{\delta f_1(k)} = 0$, leads to

$$\begin{aligned} \tan 2f(k) &= \frac{2(G_D - \frac{K}{4} I_s^{(3)}) I_D}{\bar{\epsilon}_n - \bar{\mu}} \cos\left(\frac{\phi_1 - \phi_2}{2}\right) \\ &\equiv \frac{\Delta}{\bar{\epsilon}_n - \bar{\mu}} \cos\left(\frac{\phi_1 - \phi_2}{2}\right); \\ \tan 2f_1(k) &= \frac{\Delta}{\bar{\epsilon}_n + \bar{\mu}} \cos\left(\frac{\phi_1 - \phi_2}{2}\right), \end{aligned} \quad (53)$$

where we have defined the superconducting gap Δ as

$$\Delta = 2\left(G_D - \frac{K}{4} I_s^{(3)}\right) I_D, \quad (54)$$

and $\bar{\epsilon} = (\epsilon_n^u + \epsilon_n^d)/2$, $\bar{\mu} = (\mu^{ur} + \mu^{dg})/2 = \mu + 1/6\mu_E + 1/\sqrt{3}\mu_8$, where we have used Eq. (40) for the chemical potentials. Further, ϵ_n^i is the n th Landau level energy for the i th flavor with constituent quark mass M_i given as $\epsilon_n^i = \sqrt{p_z^2 + 2n|q_i|B} + M_i$. It is thus seen that the diquark condensate functions depend upon the *average* energy and the *average* chemical potential of the quarks that condense. We also note here that the diquark condensate functions depend upon the masses of the two quarks which condense through the function $\cos((\phi_1 - \phi_2)/2)$. The function $\cos\phi_i = M_i/\epsilon_n^i$ can be different for u,d quarks, when the charge neutrality condition is imposed. Such a normalization factor is always there when the condensing fermions have different masses as has been noted in Ref. [50] in the context of the CFL phase.

Finally, the minimization of the thermodynamic potential with respect to the thermal functions $\theta_{\pm}^{ia}(\mathbf{k})$ gives

$$\sin^2\theta_{\pm}^{ia} = \frac{1}{\exp(\beta(\omega_{i,a} \pm \mu_{ia})) + 1}. \quad (55)$$

Various ω^{ia} 's ($i, a \equiv$ flavor, color) are explicitly given as

$$\omega_{n\pm}^{11} = \omega_{n\pm}^{12} = \bar{\omega}_{n\pm} + \delta\epsilon_n \pm \delta\mu \equiv \omega_{n\pm}^u, \quad (56)$$

$$\omega_{n\pm}^{21} = \omega_{n\pm}^{22} = \bar{\omega}_{n\pm} - \delta\epsilon_n \mp \delta\mu \equiv \omega_{n\pm}^d, \quad (57)$$

for the quarks participating in condensation. Here, $\bar{\omega}_{n\pm} = \sqrt{(\bar{\epsilon}_n \pm \bar{\mu})^2 + \Delta^2 \cos^2(\phi_1 - \phi_2)/2}$. Further, $\delta\epsilon_n = (\epsilon_n^u - \epsilon_n^d)/2$ is half the energy difference between the quarks which condense in a given Landau level and $\delta\mu = (\mu_{ur} - \mu_{dg})/2 = \mu_E/2$ is half the difference between the chemical potentials of the two condensing quarks. For the charged quarks which do not participate in the superconductivity,

$$\omega_{n\pm}^{ia} = \epsilon_n^i \pm \mu^{ia}. \quad (58)$$

In the above, the upper sign corresponds to antiparticle excitation energies while the lower sign corresponds to the particle excitation energies.

Let us note that when the charge neutrality conditions are not imposed, the masses of u and d quarks will be almost the same but for the effect of the (rotated) magnetic field as the magnitude of the charges for red and green quarks are the same and that of the blue color is different. Since the chemical potentials of all the quarks are the same when charge neutrality is not imposed, all four quasiparticles taking part in diquark condensation will have (almost) the same energy $\bar{\omega}_{n-}$. On the other hand, when charge neutrality condition is imposed, it is clear from the dispersion relations

given in Eqs. (56) and (57) that it is possible to have zero modes, i.e., $\omega^{ia} = 0$ depending upon the values of $\delta\epsilon_n$ and $\delta\mu$. So, although we shall have nonzero order parameter Δ , there will be fermionic zero modes or the gapless superconducting phase [51,52].

Substituting the solutions for the quark-antiquark condensate function ϕ^i of Eq. (49), we have the solutions for the different quark-antiquark condensates, i.e., I_s^{ia} given by, using Eqs. (21), (24), and (25),

$$I_s^{ia} = -\sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z \frac{M_i}{\sqrt{p_z^2 + 2n(\tilde{e}B/2) + M_i^2}} (1 - F^{ia} - F_1^{ia}), \quad (i, a = 1, 2), \quad (59)$$

$$I_s^{13} = -\sum_n \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \int dp_z \frac{M_1}{\sqrt{p_z^2 + 2n(\tilde{e}B) + M_1^2}} (1 - \sin^2\theta_-^{13} - \sin^2\theta_+^{13}), \quad (60)$$

$$I_s^{3a} = -\sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z \frac{M_3}{\sqrt{p_z^2 + 2n(\tilde{e}B/2) + M_3^2}} (1 - \sin^2\theta_-^{3a} - \sin^2\theta_+^{3a}), \quad (a = 1, 2), \quad (61)$$

for the \tilde{e} charged quarks while for the uncharged quarks (blue down and blue strange quarks),

$$I_s^{i3} = -\frac{2}{(2\pi)^3} \int d\mathbf{p} \frac{M_i}{i\sqrt{\mathbf{p}^2 + M_i^2}} (1 - \sin^2\theta_-^{i3} - \sin^2\theta_+^{i3}), \quad (i = 2, 3). \quad (62)$$

Similarly, substituting the solutions for the diquark/di-antiquark condensate functions from Eq. (53) in Eq. (27), we have, with the usual notations, $\bar{\xi}_{n\pm} = \bar{\epsilon}_n \pm \bar{\mu}$ and $\bar{\omega}_{n\pm} = \sqrt{\bar{\xi}_{n\pm}^2 + \Delta^2 \cos^2(\phi_1 - \phi_2)/2}$,

$$I_D = \frac{2}{(2\pi)^2} \sum_n \alpha_n |\tilde{e}B/2| \int dp_z \Delta \cos^2\left(\frac{\phi_1 - \phi_2}{2}\right) \left[\frac{1}{\bar{\omega}_{n-}} (1 - \sin^2\theta_-^1 - \sin^2\theta_-^2) + \frac{1}{\bar{\omega}_{n+}} (1 - \sin^2\theta_+^1 - \sin^2\theta_+^2) \right]. \quad (63)$$

Thus Eqs. (50)–(52) for the mass gaps, Eq. (54) for the superconducting gap and Eqs. (59)–(63) define the self-consistent mass gap equation for the i th quark flavor and the superconducting gap.

Next we discuss the thermodynamic potential. We substitute the solutions for the condensate functions [Eqs. (49) and (53)] in the expression for the thermodynamic potential [Eq. (41)] and use the gap equations [Eqs. (50)–(52) and (54)]. The thermodynamic potential is then given by

$$\Omega_q = \Omega_{1/2}^{sc} + \Omega_{1/2}^s + \Omega_0 + \Omega_1 + 4G_s \sum_i I_s^i - 4KI_s^u I_s^d I_s^s + \frac{\Delta^2}{4G'_D} - \frac{K}{4} I_s^s I_D^2, \quad (64)$$

where we have defined an effective diquark coupling $G'_D = G_D - \frac{K}{4} I_s^s$ in the presence of the determinant term which mixes the flavors. Let us now discuss each of the terms in Eq. (64). The first term is the contribution from the quarks that take part in superconductivity, i.e., the red and blue, u,d quarks. This contribution is given by

$$\begin{aligned} \Omega_{1/2}^{sc} &= -2 \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int (\epsilon_n^u + \epsilon_n^d) dp_z + 2 \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int ((\bar{\xi}_{n-} + \bar{\xi}_{n+}) - (\bar{\omega}_{n-} + \bar{\omega}_{n+})) \\ &\quad - 2 \sum_n \sum_{i=u,d} \alpha_n \frac{\tilde{e}B}{(2\pi)^2 \beta} \int dp_z [\log(1 + \exp(-\beta(\omega_{n-}^i - \mu_{ir}))) + \log(1 + \exp(-\beta(\omega_{n+}^i + \mu_{ir})))] \\ &\equiv \Omega_{1/2,0}^{sc}(T=0, \mu=0) + \Omega_{1/2,med}^{sc}(T, \mu), \end{aligned} \quad (65)$$

where we have separated the contribution of the medium $\Omega_{1/2,med}^{sc}$ from $T=0, \mu=0$ contribution. Similarly, the (\tilde{e}) charged strange quark contribution to the thermodynamic potential is given by

$$\begin{aligned}\Omega_{1/2}^s &= -2 \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z \epsilon_n^s - \sum_n \sum_{a=1,2} \sum_{s=\pm 1} \alpha_n \frac{\tilde{e}B}{2(2\pi)^2 \beta} \int dp_z [\log(1 + \exp(-\beta(\omega_{3a} + s\mu_{ia})))] \\ &\equiv \Omega_{1/2,0}^s + \Omega_{1/2,\text{med}}^s.\end{aligned}\quad (66)$$

The term Ω_1 in Eq. (64) arises from the blue colored u quark with charge \tilde{e} and is given as

$$\Omega_1 = - \sum_n \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \int (\epsilon_n^u) - \sum_n \sum_{s=\pm 1} \alpha_n \frac{\tilde{e}B}{(2\pi)^2 \beta} \int dp_z [\log(1 + \exp(-\beta(\omega_{33} + s\mu_{33})))] \equiv \Omega_{1,0}^u + \Omega_{1,\text{med}}^u.$$

Finally, the \tilde{e} uncharged quarks' contributions to the thermodynamic potential Ω_0 are given by

$$\Omega_0 = -2 \sum_{i=2,3} \int \frac{d\mathbf{p}}{(2\pi^3)} \epsilon^i(\mathbf{p}) - \frac{2}{(2\pi)^3 \beta} \int d\mathbf{p} \sum_{s=\mp 1} [\log(1 + \exp(-\beta(\omega_{23} + s\mu_{33})))] . \quad (67)$$

Now, all the zero temperature and zero chemical potential contributions of the thermodynamic potential in Eqs. (65)–(67) are ultraviolet divergent. This divergence also gets transmitted to the gap equations through the quark-antiquark as well as diquark condensates in Eqs. (59), (60), (61), and (63). For the chargeless case, these can be rendered finite through a regularization with a sharp cutoff in the magnitude of three momentum as is usually done in the NJL models. However, it is also seen that a sharp cutoff in the presence of magnetic field for charged particles suffers from cutoff artifacts since the continuous momentum dependence in two spatial dimensions are replaced by the sum over discrete Landau levels. To avoid this, some calculations use a smooth parametrization for the cutoff as e.g., in Ref. [17]. In the present work however we follow the elegant procedure that was followed in Ref. [23] by adding and subtracting a vacuum (zero field) contribution to the thermodynamic potential which is also divergent. This manipulation makes e.g., the Dirac vacuum contribution in the presence of magnetic field to a physically more appealing form by separating the same to a zero field vacuum contribution and a finite field contribution written in terms of the Riemann-Hurwitz ζ function. The vacuum contribution to the energy density arising from a charged quark can be written as [23,24]

$$\begin{aligned}& - \sum_{n=0}^{\infty} \frac{\alpha_n |q_i B|}{(2\pi)^2} \int dp_z \sqrt{p_z^2 + 2n|q_i B| + M_i^2} \\ &= - \frac{2}{(2\pi)^3} \int d\mathbf{p} \sqrt{\mathbf{p}^2 + M_i^2} \\ & \quad - \frac{|q_i B|^2}{2\pi^2} \left[\zeta'(-1, x_i) - \frac{1}{2}(x_i^2 - x_i) \ln x_i + \frac{x_i^2}{4} \right], \quad (68)\end{aligned}$$

where we have defined the dimensionless quantity, $x_i = \frac{M_i^2}{2|q_i B|}$, i.e., the mass parameter in units of the magnetic field. Further, $\zeta'(-1, x) = d\zeta(z, x)/dz|_{z=-1}$ is the derivative of the Riemann-Hurwitz zeta function [53].

Using Eq. (68), the quark-antiquark condensate of (\tilde{q}) charged quarks can be written as

$$\begin{aligned}\langle \bar{\psi}^{ia} \psi^{ia} \rangle &= - \frac{2}{(2\pi)^3} \int d\mathbf{p} \frac{M_i}{\sqrt{\mathbf{p}^2 + M_i^2}} \\ & \quad - \frac{M_i |q_i B|}{2\pi^2} \left[x_i (1 - \ln x_i) + \ln \Gamma(x_i) + \frac{1}{2} \ln \left(\frac{x_i}{2\pi} \right) \right] \\ & \quad + I_{s\text{med}}^{ia} \\ & \equiv I_{s\text{vac}}^{ia} + I_{s\text{field}}^{ia} + I_{s\text{med}}^{ia}.\end{aligned}\quad (69)$$

The first term, $I_{s\text{vac}}^{ia}$ can be explicitly evaluated with a cutoff Λ as

$$I_{s\text{vac}}^{ia} = \frac{M_i}{2\pi^2} \left[\Lambda \sqrt{\Lambda^2 + M_i^2} - M_i^2 \log \left(\frac{\Lambda + \sqrt{\Lambda^2 + M_i^2}}{M_i} \right) \right]. \quad (70)$$

The medium contribution to the scalar condensate from the superconducting part is

$$I_{s\text{med}}^{ia} = \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z \frac{M_i}{\epsilon_n^i} (F^{ia} - F_1^{ia}), \quad (71)$$

while, for the nonsuperconducting blue u quarks,

$$I_{s\text{med}}^{13} = \sum_n \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \int dp_z \frac{M_1}{\epsilon_n^1} (\sin^2 \theta_-^{13} - \sin^2 \theta_+^{13}). \quad (72)$$

Similarly, the contribution of the medium to the (\tilde{q}) charged strange quark-antiquark condensate is

$$\begin{aligned}I_{s\text{med}}^{3a} &= \sum_n \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \\ & \quad \times \int dp_z \frac{M_3}{\epsilon_n^3} (\sin^2 \theta_-^{3a} - \sin^2 \theta_+^{3a}), \quad (a=1,2).\end{aligned}\quad (73)$$

In what follows, we shall focus our attention to zero temperature calculations. Using the relation $\lim_{\beta \rightarrow \infty} \frac{1}{\beta} \ln(1 + \exp(-\beta\omega)) = -\omega\theta(-\omega)$ and using Eqs. (65) and (68), we have the zero temperature thermodynamic potential for the color superconducting quarks given as

$$\Omega_{1/2}^{sc}(T=0, \mu, B) = \Omega_{1/2,0}^{sc}(T=0, \mu=0) + \Omega_{1/2,med}^{sc}(T=0, \mu), \quad (74)$$

with

$$\Omega_{1/2,0}^{sc}(T=0, \mu=0) = -2 \times 2 \sum_{i=u,d} G(\Lambda, M_i) - 2 \sum_{i=u,d} F(x_i, B), \quad (75)$$

where we have defined the function $G(\Lambda, M)$ as

$$\begin{aligned} G(\Lambda, M) &= \frac{1}{(2\pi)^3} \int \sqrt{\mathbf{p}^2 + M^2} d\mathbf{p} \\ &= \frac{1}{16\pi^2} \left[\Lambda \sqrt{\Lambda^2 + M^2} (2\Lambda^2 + M^2) \right. \\ &\quad \left. - M^4 \log \left(\frac{\Lambda + \sqrt{\Lambda^2 + M^2}}{M} \right) \right]. \end{aligned} \quad (76)$$

The prefactors in the first term correspond to color and spin degeneracy factors while the same in the second term corresponds to the color degeneracy factor. The magnetic field dependent function, $F(x_i, B)$ with $x_i = M_i^2/|q_i B|$,

$$F(x_i, B) = \frac{|q_i B|^2}{2\pi^2} \left[\zeta'(-1, x_i) - \frac{1}{2} (x_i^2 - x_i) \ln x_i + \frac{x_i^2}{4} \right]. \quad (77)$$

The medium contribution from the superconducting quarks is given as

$$\begin{aligned} \Omega_{1/2,med}^{sc}(T=0, \mu) &= 2 \sum_{n=0}^{n_{\max}} \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int_0^{p_{z,n}^{\max}} dp_z [\bar{\xi}_{n-} + \bar{\xi}_{n+} - (\bar{\omega}_{n-} + \bar{\omega}_{n+})] \\ &\quad + 2 \sum_{n=0}^{n_{\max}} \sum_{i=u,d} \alpha_n \frac{\tilde{e}B}{2(2\pi)^2} \int_0^{p_{z,n}^{\max}} dp_z i [\omega_{n-}^i \theta(-\omega_{n-}^i) + \omega_{n+}^i \theta(-\omega_{n+}^i)]. \end{aligned} \quad (78)$$

The three momentum cutoff Λ for the magnitude of momentum in the absence of magnetic field leads to the sum over the Landau level up to $n_{\max} = \frac{\Lambda^2}{2\tilde{e}B}$. Further, the positivity of the magnitude of p_z restricts the cutoff in $|p_z|$ as $p_{z,n}^{\max} = \sqrt{\Lambda^2 - n\tilde{e}B}$ for a given value of n of the Landau level.

The contribution of the blue up quark to the thermodynamic potential $\Omega_1 = \Omega_{1,0} + \Omega_{1,med}$ with

$$\Omega_{1,0}(T=0, \mu=0) = -2G(\Lambda, M_u) - F(x_u, B), \quad (79)$$

and

$$\Omega_{1,med}(T=0, \mu) = \sum_{n=0}^{n_{\max}^u} \alpha_n \frac{\tilde{e}B}{(2\pi^2)} \left[\mu_{ub} \sqrt{\mu_{ub}^2 - M_{nu}^2} + M_{nu}^2 \log \left(\frac{\mu_{ub} + \sqrt{\mu_{ub}^2 - M_{nu}^2}}{M_{nu}} \right) \right], \quad (80)$$

where $M_{nu} = \sqrt{M_u^2 + 2n\tilde{e}B}$ is the n th Landau level mass for up quark and $n_{\max}^u = \text{Int} \left[\frac{\mu_{ub}^2 - M_u^2}{2\tilde{e}B} \right]$ is the maximum number of Landau level consistent with the zero temperature distribution function.

The \tilde{e} charged strange quark contribution to the thermodynamic potential $\Omega_{1/2}^s = \Omega_{1/2,0}^s + \Omega_{1/2,med}^s$, with

$$\Omega_{1/2,0}^s(T=0, \mu=0) = -2 \times 2G(\Lambda, M_s) - 2F(x_s, B), \quad (81)$$

and

$$\Omega_{1/2,med}^s(T=0, \mu) = 2 \sum_{n=0}^{n_{\max}^s} \alpha_n \frac{\tilde{e}B}{2(2\pi^2)} \left[\mu_{sr} \sqrt{\mu_{sr}^2 - M_{ns}^2} + M_{ns}^2 \log \left(\frac{\mu_{sr} + \sqrt{\mu_{sr}^2 - M_{ns}^2}}{M_{ns}} \right) \right], \quad (82)$$

where $M_{ns} = \sqrt{M_s^2 + 2n\tilde{e}B}$ is the n th Landau level mass for the s quarks. Further, the sum over the Landau levels is restricted to $n_{\max}^s = \text{Int} \left[\frac{\mu_{sr}^2 - M_s^2}{2\tilde{e}B} \right]$ arising from the distribution function at zero temperature $\theta(\mu - \epsilon_n)$.

For the uncharged quarks, i.e., blue down and strange quarks, we have $\Omega_0 = \Omega_{0,0} + \Omega_{0,\text{med}}$ with

$$\Omega_{0,0}(T=0, \mu=0) = -2 \sum_{i=d,s} G(\Lambda, M_i), \quad (83)$$

and for the medium part, with $p_{fi} = \sqrt{\mu_i^2 - M_i^2}$,

$$\Omega_{0,\text{med}}(T=0, \mu) = 2 \sum_{i=d,s} H_i(\mu_{i3}, p_{fi}). \quad (84)$$

In the above H_i is the medium contribution from a single chargeless flavor given as

$$H_i(\mu, p_f) = \frac{1}{16\pi^2} \left[p_{fi} \mu_i (p_{fi}^2 + \mu_i^2) - M_i^4 \log \left(\frac{\mu_i^2 + p_{fi}^2}{M_i^2} \right) \right]. \quad (85)$$

Next, we write down the expressions for the condensates at zero temperature, which are needed to compute the thermodynamic potential in Eq. (64). This is already given by Eq. (69). Here, we write down explicitly the zero temperature limit for the same. The scalar condensate for, say, u quarks is given as

$$I_s^u = I_{s,\text{vac}}^u + I_{s,\text{med}}^{ur} + I_{s,\text{med}}^{ug} + I_{s,\text{med}}^{ub} + \sum_{a=1}^3 I_s^{\text{field}-u}(x_{ua}). \quad (86)$$

The vacuum contribution $I_{s,\text{vac}}^u$ is already given in Eq. (70).

The scalar condensate medium contribution from the superconducting red up and green up quarks is given as

$$I_{s,\text{med}}^{ur} = I_{s,\text{med}}^{ug} = - \sum_{n=0}^{n_{\text{max}}} \alpha_n \frac{(\tilde{\epsilon}B)}{2(2\pi)^2} \int dp_z \frac{M_u}{e_n^u} (F^{ur} + F_1^{ur}). \quad (87)$$

The expressions for the distribution functions F^{ia} and F_1^{ia} are already given in Eqs. (22) and (23) in terms of the diquark condensate functions and the thermal distribution functions. In the zero temperature limit, the distribution functions for e.g., u quarks become

$$F^{ur} = \frac{1}{2} \left(1 - \frac{\tilde{\xi}_{n-}}{\tilde{\omega}_{n-}} \right) (1 - \theta(-\omega^d)), \quad (88)$$

and

$$F_1^{ur} = \frac{1}{2} \left(1 - \frac{\tilde{\xi}_{n+}}{\tilde{\omega}_{n+}} \right). \quad (89)$$

The blue up quark contribution to the scalar condensate is given by

$$I_{s,\text{med}}^{ub} = - \sum_{n=0}^{n_{\text{max}}} 2M\alpha_n \frac{\tilde{\epsilon}B}{(2\pi)^2} \log \left(\frac{p_z^{\text{max}} + \sqrt{p_z^{\text{max}2} + M_{nu}^2}}{M_{nu}} \right). \quad (90)$$

As in Eq. (80) here we have defined the n th Landau level mass for the blue up quark as $M_{nu}^2 = M_u^2 + 2n|\tilde{\epsilon}B|$. The magnetic field contribution to the scalar condensate for the up quarks of a given color “a” is given by

$$I_s^{\text{field}-u}(x_{ua}) = -M_u \frac{|q_a B|}{2\pi^2} \left[x_a (1 - \ln x_a) + \ln \Gamma(x_a) + \frac{1}{2} \frac{x_a}{22\pi} \right], \quad (91)$$

where $x_a = M_u^2/2|q_a B|$ and $q_a = \tilde{\epsilon}/2$ for red and green colors and $q_a = \tilde{\epsilon}$ for blue color up quarks.

In an identical manner, the scalar condensates for the down and strange quarks I_s^d, I_s^s can be written down with appropriate changes for the charges and the masses. The diquark condensate $4I_D$ is given in Eq. (63) where the zero temperature limit can be taken by replacing the distribution functions $\sin^2 \theta^i = \theta(-\omega^i)$, ($i = u, d$). Thus the thermodynamic potential, Ω_q given in Eq. (64) gets completely defined for the quark matter in the presence of a magnetic field.

In the context of neutron star matter, the quark phase that could be present in the interior consists of the u,d,s quarks as well as electrons, in weak equilibrium,

$$d \rightarrow u + e^- + \bar{\nu}_{e^-}, \quad (92a)$$

$$s \rightarrow u + e^- + \bar{\nu}_{e^-}, \quad (92b)$$

and

$$s + u \rightarrow d + u, \quad (92c)$$

leading to the relations between the chemical potentials $\mu_u, \mu_d, \mu_s, \mu_E$ as

$$\mu_s = \mu_d = \mu_u + \mu_E. \quad (93)$$

The neutrino chemical potentials are taken to be zero as they can diffuse out of the star. So there are *two* independent chemical potentials needed to describe the matter in the neutron star interior which we take to be the quark chemical potential μ_q and the electric charge chemical potential μ_e in terms of which the chemical potentials are given by $\mu_s = \mu_q - \frac{1}{3}\mu_e = \mu_d$, $\mu_u = \mu_q + \frac{2}{3}\mu_e$ and $\mu_E = -\mu_e$. In addition, for a description of the charge neutral matter, there is a further constraint for the chemical potentials through the following relation for the particle densities given by

$$Q_E = \frac{2}{3}\rho_u - \frac{1}{3}\rho_d - \frac{1}{3}\rho_s - \rho_E = 0. \quad (94)$$

The color neutrality condition corresponds to

$$Q_8 = \frac{1}{\sqrt{3}} \sum_{i=u,d,s} (\rho^{i1} + \rho^{i2} - 2\rho^{i3}) = 0. \quad (95)$$

In the above, ρ^{ia} is the number density for quarks of flavor i and color a . In particular, the number densities of the condensing quarks are given as

$$\rho^{ia} = \sum_n \frac{\tilde{e}B}{2(2\pi)^2} \int dp_z (F^{ia} - F_1^{ia}), \quad (i, a = 1, 2), \quad (96)$$

where F^{ia}, F_1^{ia} are defined in Eqs. (22) and (23) respectively in terms of the condensate functions and e.g., for zero temperature is given explicitly in Eq. (88) for up red quarks. For the blue colored quarks, the same for the up blue quarks is given by

$$\rho^{ub} = \sum_{n=0}^{n_{\max}^u} \alpha_n \frac{\tilde{e}B}{2\pi^2} \sqrt{\mu_{ub}^2 - M_u^2 - 2n\tilde{e}B}, \quad (97)$$

while for the \tilde{e} uncharged d quarks

$$\rho^{db} = \frac{(\mu_{db}^2 - M_d^2)^{3/2}}{3\pi^2}. \quad (98)$$

For the charged strange quarks the number densities are given by

$$\rho^{sr} = \rho^{sg} = \sum_{n=0}^{n_{\max}^s} \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \sqrt{\mu_{sr}^2 - M_s^2 - n\tilde{e}B}, \quad (99)$$

while for the \tilde{e} uncharged blue strange quarks

$$\rho^{sb} = \frac{(\mu_{sb}^2 - M_s^2)^{3/2}}{3\pi^2}. \quad (100)$$

The electron number density is given by

$$\rho_E = \sum_n \alpha_n \frac{\tilde{e}B}{\pi^2} \left(\sqrt{\mu_E^2 - 2n\tilde{e}B} \right). \quad (101)$$

To discuss the pressure in the context of matter in the core of the neutron star, one also has to add the contribution of the electrons to the thermodynamic potential. Since we shall describe the system as a function of $\tilde{e}B$, we shall take the approximations $\tilde{e} \sim e, A_\mu \sim \tilde{A}_\mu$ to a good approximation as the mixing angle is small. The corresponding thermodynamic potential for the electrons is given by

$$\Omega_e = \sum_{n=0}^{n_{\max}^e} \alpha_n \frac{eB}{(2\pi)^2} \left[\mu_E \sqrt{\mu_E^2 - 2neB} - 2neB \log \left(\frac{\mu_E + \sqrt{(\mu_E^2 - 2neB)}}{\sqrt{2neB}} \right) \right], \quad (102)$$

where $n_{\max}^e = \frac{\mu_E^2}{2|\tilde{e}B|}$. Clearly in Eqs. (101) and (102) we have neglected the electron mass ($m_E \sim 0.5$ MeV), which is small compared to μ_E which is few tens of MeV. Thus the total thermodynamic potential or the negative of the pressure is given as, with Ω_q given in Eq. (64)

$$\Omega = \Omega_q + \Omega_e. \quad (103)$$

The thermodynamic potential [Eq. (103)], the mass and superconducting gap equations [Eqs. (50)–(52) and (54)], along with the charge neutrality conditions [Eq. (94) and (95)] are the basis for our numerical calculations for various physical situations that we shall discuss in detail in the following section.

IV. RESULTS AND DISCUSSIONS

We begin the discussions with the parameters of the NJL model. The model parameters are the three current masses of quarks, namely m_u, m_d and m_s and the couplings G_S, G_D and the determinant coupling K . This apart, one additional parameter, the momentum cutoff Λ , is also required to regularize the divergent integrals which are characteristic of the four point interaction of NJL models. Except for the diquark coupling G_D , there are several parameter sets for the couplings derived from fitting of the meson spectrum and chiral condensate [54–56]. The diquark coupling is not known from fitting since one does not have a diquark spectrum to fit with. The Fierz transforming quark-antiquark term from one gluon exchange term gives the relation $G_D = 0.75 G_S$. Although not precise, many other references use this value [47,57,58]. However some other references [59,60] also consider the case of stronger diquark coupling $G_D = G_S$ apart from $G_D = 0.75 G_S$. In the following we shall limit ourselves only to the case of $G_D = 0.75 G_S$. For a nice discussion on this we refer the interested reader to Sec. 4.2.2 of Ref. [61]. The parameters used in our calculations are $m_u = 5.5$ MeV, $m_d = 5.5$ MeV, $m_s = 140.7$ MeV for the current quark masses, the momentum cutoff $\Lambda = 602.5$ MeV and the couplings $G_S, \Lambda^2 = 1.835$ and $K\Lambda^5 = 12.36$ as have been chosen in Ref. [56]. After choosing the light current quark mass $m_u = m_d = 5.5$ MeV, the remaining four parameters are chosen to fit vacuum values of pion decay constant f_π , masses of pion, kaon and η' . With this set of parameters the η meson mass is underestimated by about 6 percent and leads to u and d constituent mass in vacuum to be about 368 MeV. The strange mass is about 549 MeV at zero temperature and density. The determinant interaction is responsible for $U(1)_A$ anomaly and getting the correct eta mass. Further, this interaction also mixes the various gap equations and affects the superconducting gap significantly as we shall see. However, we must point out that there is a large discrepancy in the determination of this six fermion interaction coupling K . For example, in Ref. [54] the parameter $K\Lambda^5$ differs by as

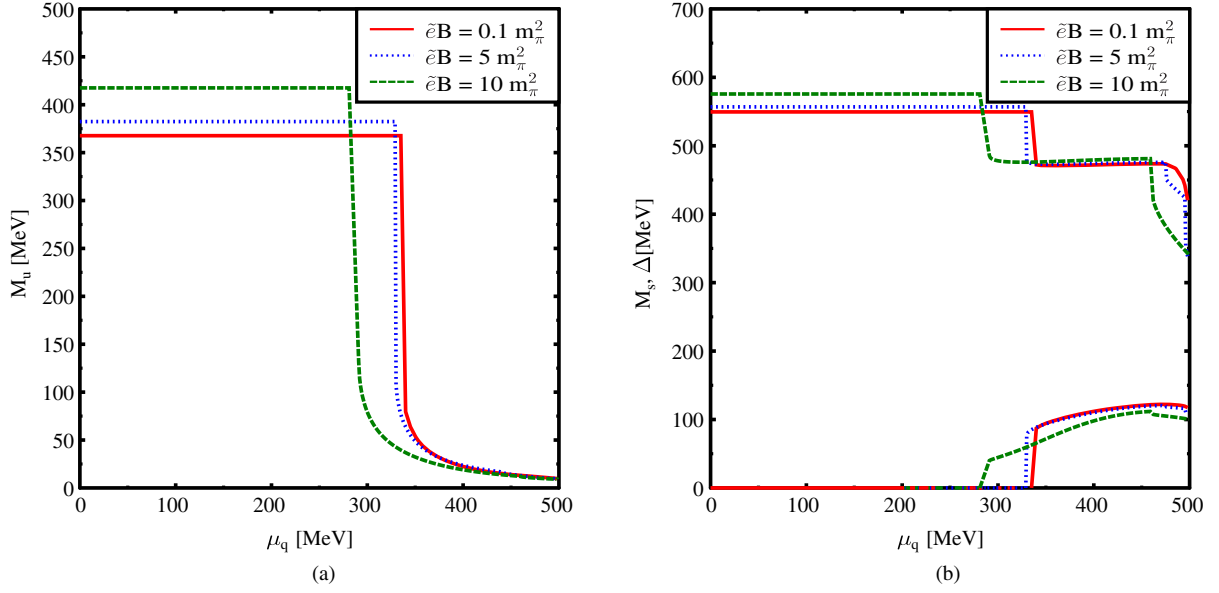


FIG. 1. Constituent quark masses and superconducting gap when charge neutrality conditions are not imposed. Part (a) shows the M_i at zero temperature as a function of quark chemical potential for different values of the magnetic field. Part (b) shows the same for the strange quark mass M_s and the superconducting gap.

large as 30 percent as compared to the value chosen here. This discrepancy is due to the difference in the treatment of η' mesons with a high mass [28]. In fact, this leads to an unphysical imaginary part for the corresponding polarization diagram in the η' meson channel. This is unavoidable because NJL is not confining and is unrealistic in this context. Within the above-mentioned limitations of the model and the uncertainty in the value of the determinant coupling, we proceed with the present parameter set which has already been used for phase diagram of dense matter in Refs. [28,59] and for neutron star matter in Ref. [62].

We begin our discussion for the simpler case where the charge neutrality conditions are not imposed. In this case, the electrical and color charge chemical potential are set to zero so that all the quarks have same potential μ_q . In this case we have to solve four gap equations, three for the constituent masses [Eqs. (50)–(52)] and the fourth for the superconducting gap [Eqs. (54) and (63)]. For given values of quark chemical potential and magnetic field we solve the gap equations self-consistently. A few comments regarding solving these gap equations may be in order. Although the gap equations and the thermodynamic potential has been written down for a given T and μ , we confine our attention to the case of zero temperature only in the present investigation. Second, for nonvanishing magnetic fields, all the Landau levels for the medium part up to a cutoff, $n_{\max} = \frac{\sqrt{\mu^2 - M_i^2}}{2\tilde{e}B}$ for each flavor i , are taken into account. Near the μ_c , the critical chemical potential for chiral transition for light quarks, there can be multiple solutions for the gap equations. We have chosen the solutions which have the lowest thermodynamic potential.

In Fig. 1, we have shown the variation of the masses as a function of quark chemical potential μ_q for three different values of magnetic fields, $\tilde{e}B = 0.1 m_\pi^2$, $5 m_\pi^2$, $10 m_\pi^2$. The results for $\tilde{e}B = 0.1 m_\pi^2$ reproduce the vanishing magnetic field results. As the chemical potential increases, the masses remain constant up to a critical value of quark chemical potential μ_c and the superconducting gap remains zero. At the critical chemical potential there is a first order phase transition and the constituent masses drop sharply from their vacuum values and the superconducting gap becomes nonzero. For vanishing magnetic field, the isospin symmetry for the light quarks is unbroken and the constituent masses of u and d quarks are degenerate. The critical chemical potential, μ_c , is about 340 MeV for (almost) vanishing magnetic field. In this case, the up and the down quark masses decrease from their vacuum values of about 368 MeV to about 80 MeV. The strange mass being coupled to other gaps via determinant interaction also decreases from 549 to 472 MeV when this first order transition happens for the light quarks. However, since this μ_c is still less than the strange mass its density remains zero. The superconducting gap rises from 0 to 88.0 MeV at μ_c . As the chemical potential is increased beyond μ_c , the superconducting gap shows a mild increase reaching a maximum value of 122 MeV at around $\mu_q \sim 475$ MeV beyond which the gap shows a mild decrease with μ .

Such a decrease of the gap with chemical potential could be due to two reasons. First, at higher chemical potentials, beyond $\mu = 475$ MeV, the strange quark mass starts decreasing rapidly. This leads to a decrease of the effective diquark coupling $G'_D = G_D + \frac{K}{4} \langle \bar{s}s \rangle$ resulting in a decrease

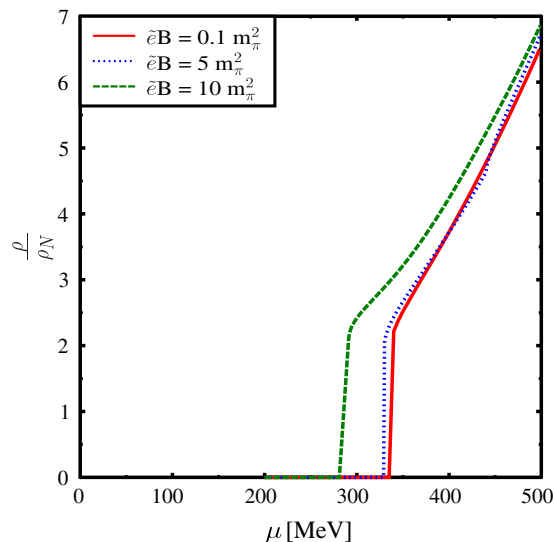


FIG. 2. Baryon number density in units of nuclear matter density as a function of chemical potential for different strengths of magnetic field at zero temperature.

in the superconducting gap with increasing chemical potential. Second, such a behavior of decreasing superconducting gap with μ for large μ could also be a manifestation of a finite cutoff in the momentum integration in e.g., Eq. (78). One may note that the first term in Eq. (78) is the contribution from the medium. Indeed, for $T = 0$, $\mu = 0$ the contributions of the two terms in the integrand here cancel out unlike the term in Eq. (68) which is a genuine vacuum contribution and is divergent. The second term in Eq. (78) in any case gives a contribution from the medium when $\delta\mu \neq 0$. For both terms the upper limit of p_z integration, $p_{z,n}^{\max}$, has a Λ dependence. Therefore, it is expected that there will be a cutoff dependence in the contribution of this term to the thermodynamic potential. This effect of finite cutoff therefore will be more pronounced at large μ . Thus the decrease of the superconducting gap at large μ in Fig. 1(b) could also be a reflection of this effect. Therefore the decreasing behavior of Δ can be both due to the decrease of effective diquark coupling G'_D , and the effects of a finite cutoff.

In Fig. 2, we have plotted the total baryon number density in units of nuclear matter density ($\rho_N = 0.17/\text{fm}^{-3}$) as a function of quark chemical potential. For vanishing magnetic field, at the critical chemical potential $\mu_c \sim 340$ MeV, the baryon density jumps from 0 to 0.38 fm^{-3} which is about 2.2 times the nuclear matter density.

Upon increasing the magnetic field, as seen in Fig. 1, the vacuum constituent quark masses increase due to magnetic catalysis at zero density. It may also be observed here that the μ_c for chiral transition for the light quarks decreases with the magnetic field. Such a phenomenon is known as inverse magnetic catalysis at finite chemical potential [63]. Let us note that in the superconducting phase the \tilde{e} charges

of the u and d quarks are identical in magnitude while that of the unpaired blue quark are different for u and d quarks. This results in the color summed scalar condensate I_s^u and I_s^d to be different in the presence of a magnetic field. This leads to the difference in constituent masses for the light quarks. For $\tilde{e}B = 10 m_\pi^2$ the u mass in the chiral symmetry broken phase increases by about 13.6 percent and strange mass by about 4.7 percent. The critical chemical potential decreases from about 340 MeV to about 291 MeV. As seen in the plot, the superconducting gap decreases and the peak value decreases from 122 to 111 MeV. As may be seen from Eqs. (54) and (63), the superconducting gap depends upon the effective diquark coupling $G'_D = G_D - \frac{\kappa}{4} I_s^s$. With an increase in magnetic field the effective coupling G'_D has a slight increase in magnitude as the strange quark condensate increases with magnetic field. Therefore, one would have expected an increase in Δ with magnetic field. However, the variation in Δ due to the magnetic field is essentially decided by Eq. (63). From here also one would have expected an increase in Δ with magnetic field as $\tilde{e}B$ occurs in the numerator in Eq. (63). In fact, this behavior is actually seen for high magnetic field, where only the lowest Landau level contributes to the integral in Eq. (63). For moderately strong magnetic fields, contributions of the higher Landau levels become relevant for the behavior of gap with magnetic field. As long as the contribution of higher Landau levels are nonvanishing, the gap equation can support the solution for the gap that decreases with magnetic field. We may point out that $\tilde{e}B = 5 m_\pi^2$ and $10 m_\pi^2$ the cutoff for Landau levels n_{\max} equals 3 and 1 respectively. For $\tilde{e}B \geq 20 m_\pi^2$ only the lowest Landau level contributes to the integral in Eq. (63) and the gap increases with magnetic field. One may also note that at higher magnetic fields the charge asymmetry between the u and d quark becomes apparent in their masses as expected. At $10 m_\pi^2$ the difference is about 3.4 percent and at $15 m_\pi^2$ its about 5.7 percent at lower chemical potentials.

One may note that below the critical chemical potential μ_c the u quarks have higher mass compared to d quarks as all the three colors are charged for u quarks while for the d quarks, the blue color is chargeless. However, beyond the critical chemical potential the u quark has a lower mass compared to d quarks. This is because with magnetic field the medium contribution to chiral condensate increases. This increase is the same for the condensing pairs of u and d quarks but different for the blue quarks. The blue up quark has charge $\tilde{e} = 1$ whereas it is zero for the down blue quark. Therefore the medium contribution from the up quark is more than the down quark and it reduces the condensate for the up quark and consequently its mass too. As we shall see later, imposing charge neutrality requires the d quark chemical potential to be much higher compared to u quarks to balance their larger positive charge. This forces the d quark mass to be smaller compared to u quark mass above critical chemical potential. This results in an opposite

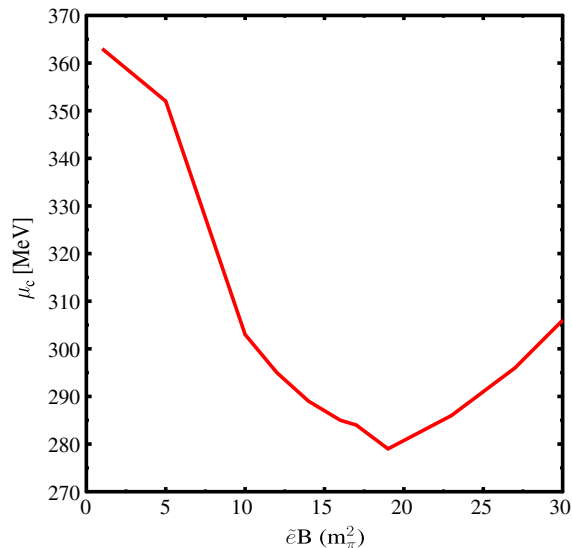


FIG. 3. Critical chemical potential for chiral transition at zero temperature as a function of magnetic field.

behavior for the u and d quark masses with chemical potential, beyond μ_c when the charge neutrality condition is imposed *vis-à-vis* when such a condition is not imposed.

As may be observed from Fig. 2, the baryon number density increases with magnetic field for a given chemical potential. This is because for the magnetic fields considered here, the symmetry is restored for lower chemical potential at higher magnetic field. Thus for a given chemical potential beyond the critical chemical potential the masses become smaller for higher magnetic field leading to larger baryon number density. This is consistent with inverse magnetic catalysis. One may note however that for very large fields, there is magnetic catalysis of chiral symmetry breaking in the sense that critical chemical potential increases with magnetic field. In Fig. 3 we show the behavior of μ_c as a function of magnetic field. It is observed that μ_c is minimum for $\tilde{e}B = 19 m_\pi^2$.

To examine the effect of flavor mixing determinant interaction, we show in Fig. 4 the variation of the masses and the superconducting gap without the determinant interaction. As expected, without the mixing of flavors the strange mass remains unaffected when u and d quark masses decrease. This is significantly different behavior compared to Fig. 1 where the strange mass decreases by about 74 MeV beyond μ_c when there is a first order transition for the light quarks. This also affects the superconducting gap. The superconducting gap is smaller as the effective diquark coupling decreases without the determinant interaction term.

In Fig. 5 we show the variation of the gaps as a function of the magnetic field for $\mu = 200$ MeV and $\mu = 400$ MeV. $\mu = 200$ MeV is less than the critical μ_c for any value of magnetic field considered here. Hence the constituent masses are high and the superconducting gap is zero.

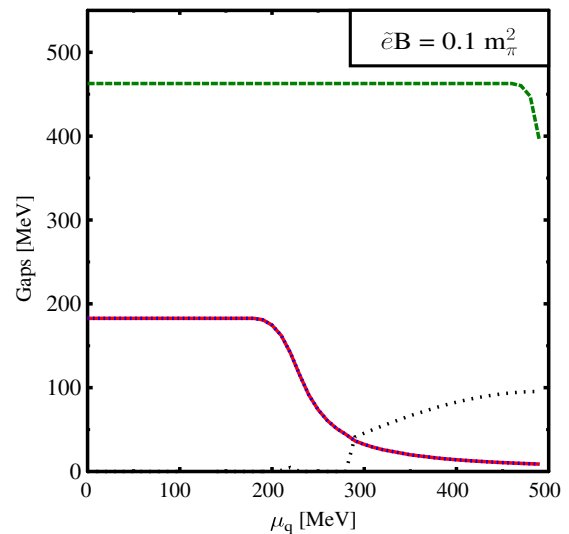


FIG. 4. Gaps without determinant interaction at zero temperature as a function of quark chemical potential. The solid curve refers to masses of u-d quarks, the dashed curve refers to the mass of strange quark and the dotted curve corresponds to the superconducting gap.

We find that the masses increase monotonically with the magnetic field. At $\tilde{e}B = 10 m_\pi^2$, the u mass increases by 14 percent of its zero field value while strange mass increases by 5 percent. Similarly for $\mu = 400$ MeV which is larger than the critical chemical potential for magnetic fields considered here, one also has finite superconducting gap. However, in this case it is observed that the u and d masses decrease slowly and monotonically with magnetic field while strange quark mass remains almost constant. The superconducting gap shows an oscillatory behavior with increase in magnetic field. The oscillatory behavior is associated with the discontinuous changes in the density of states due to Landau quantization and is similar to de Hass van Alphen effects for magnetized condensed matter system.

Finally, in Fig. 6 we have plotted the axial current density normalized to the same for three flavor without any condensates as given in Eq. (30) as a function of baryon density for values of magnetic field 5 and 10 m_π^2 . For smaller chemical potentials but above the chiral transition this ratio is about 0.75 since strange quarks do not contribute as their masses are larger than these values of chemical potential. For μ_q about 480 MeV the strange quarks contribution to the axial current density becomes nonvanishing and the ratio approaches to the value when there are no condensates. Let us note that while quark masses decrease with chemical potential, the superconducting gap increases with chemical potential. This leads to a nearly constant value for this ratio for the range of chemical potential below the strange quark mass. Above $\mu_q = 480$ MeV, the ratio shows a monotonic increase with chemical potential as the strange quark mass starts decreasing.

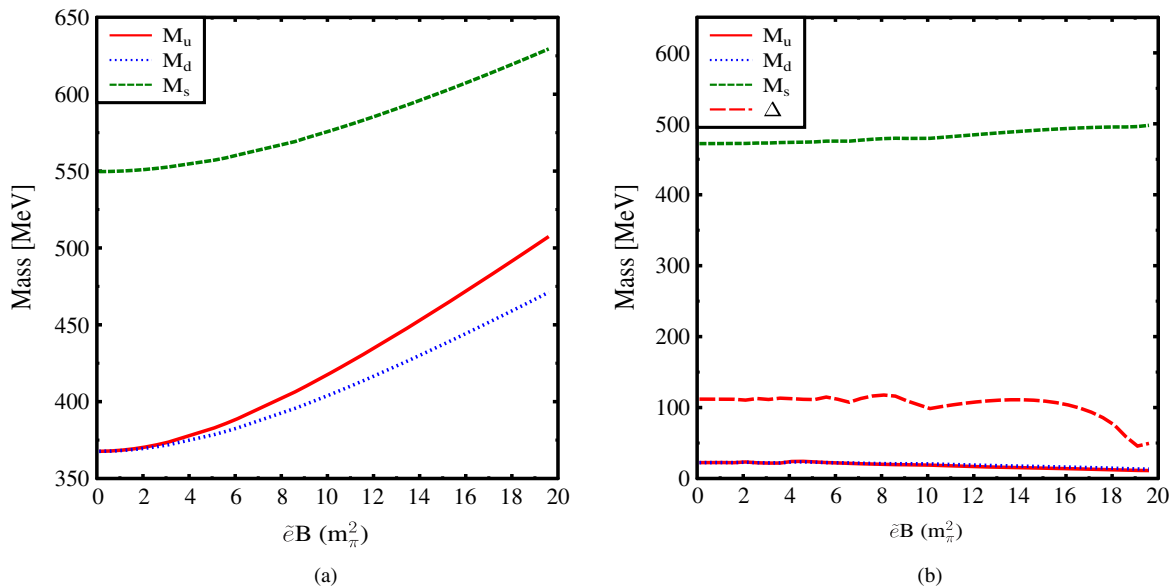


FIG. 5. Constituent quark masses as a function of magnetic field for $T = 0$. Part (a) shows the masses of the three quarks below the chiral transition for $\mu = 200$ MeV. Part (b) shows the same for the masses along with the superconducting gap above the chiral transition for $\mu_q = 400$ MeV.

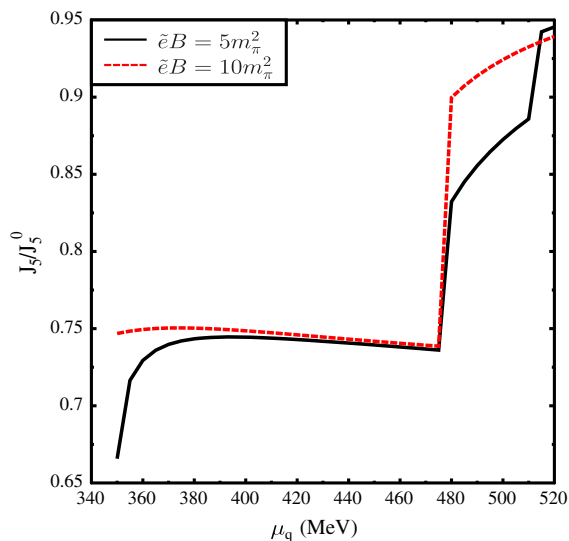


FIG. 6. Axial current density for $\tilde{e}B = 5m_\pi^2$ (black solid line) and $\tilde{e}B = 10m_\pi^2$ (red dashed line).

A. Charge neutral magnetized quark matter

Next we discuss the consequences of imposing charge neutrality conditions ($Q_E = 0, Q_8 = 0$). In Fig. 7 we show the results for the masses and the superconducting gaps for strength of the external magnetic field $\tilde{e}B = 0.1m_\pi^2$ [Fig. 7(a)] and $\tilde{e}B = 10m_\pi^2$ [Fig. 7(b)]. For small magnetic field ($\tilde{e}B = 0.1m_\pi^2$) the masses in the symmetry broken phase are the same as before but the critical chemical potential is now shifted to around $\mu_c = 364$ MeV as compared to $\mu_c = 335$ MeV when the condition is not imposed. At the transition point with neutrality the u-quark

mass decreases from 367 to 111 MeV and the down quark mass from 367 to 87 MeV. Charge neutrality requires d quark number densities to be higher as compared to u quarks. Let us note that near the critical chemical potential there are multiple solutions of the gap equations. The solution which is thermodynamically preferred when the charge neutrality condition is not imposed may no longer be the preferred solution when the constraint of charge neutrality is imposed [36]. The strange quark mass is higher than the chemical potential at the chiral restoration so its density is zero. However due to the determinant interaction the strange mass decreases at the chiral restoration from 549 to 472 MeV. At still higher chemical potential the strange quark density becomes nonzero and strange quark also helps in maintaining charge neutrality.

The critical baryon density when charge neutrality is imposed is however similar to the case when neutrality is not imposed. Specifically $\rho_c \sim 2.25\rho_0$ with charge neutrality while $\rho_c \sim 2.26\rho_0$ without charge neutrality despite the fact that μ_c is higher ($\mu_c = 364$ MeV) for the charge neutral matter compared when such charge neutrality condition is not imposed ($\mu_c = 335$ MeV). This is because the constituent masses at the transition is large ($M_u \sim 111$ MeV and $M_d \sim 87$ MeV) for charge neutral case compared to ($M_u \sim M_d \sim 85$ MeV) without the charge neutrality condition. For $\tilde{e}B = 0.1m_\pi^2$, at the chiral transition $\mu_c = 364$ MeV the superconducting gap increases from zero to 69 MeV. As the chemical potential is further increased the superconducting gap increases to 80 MeV until $\mu = \mu_1 \sim 420$ MeV where it shows a sudden jump to 106 MeV. This happens when the gapless modes cease to exist as explained below. As magnetic field is increased to $\tilde{e}B = 10m_\pi^2$, as

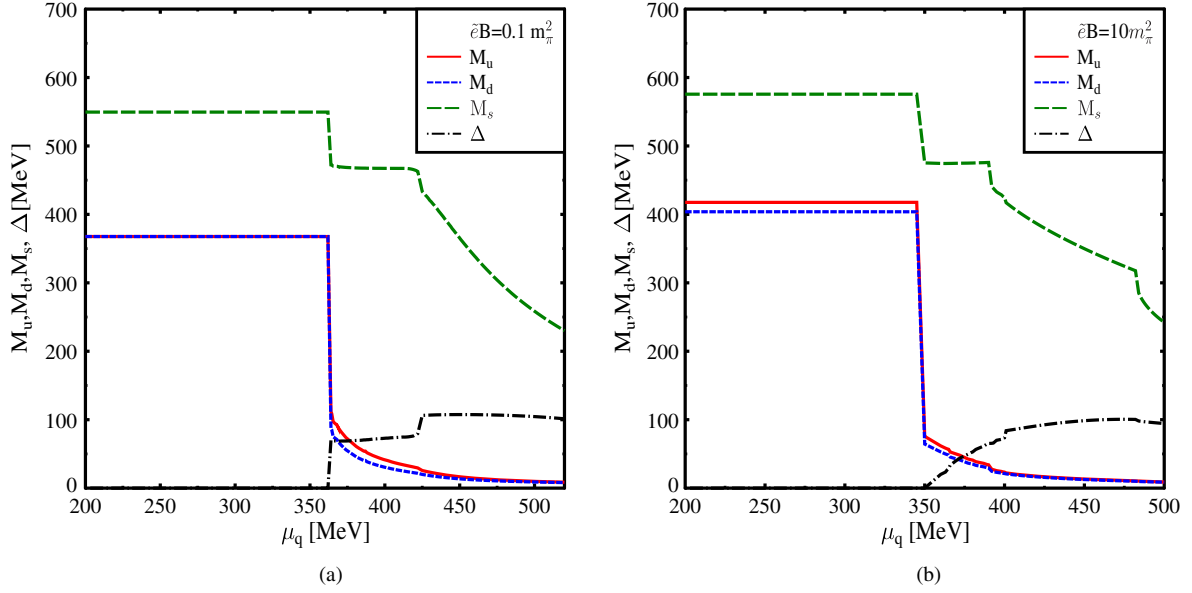


FIG. 7. Constituent quark masses and superconducting gap when charge neutrality conditions are imposed. Part (a) shows the masses and superconducting gap at zero temperature as a function of quark chemical potential for magnetic field $\tilde{B} = 0.1 m_\pi^2$. Part (b) shows the same for $\tilde{B} = 10 m_\pi^2$.

may be observed in Fig. 6(b), the critical chemical potential μ_c for the charge neutral matter decreases to 350 MeV similar to the case without the charge neutrality condition with inverse magnetic catalysis. The superconducting gap on the other hand becomes smaller. One can also observe that unlike the vanishingly small magnetic field case, the superconducting gap increases smoothly with chemical potential from zero initial value to 73 MeV at $\mu = \mu_1 \sim 400$ MeV where it again jumps to a value of 83 MeV.

B. Gapless modes

In the region between μ_c and μ_1 the system shows gapless mode which we discuss now in some detail. Without magnetic field this has earlier been seen for charge neutral matter [35,36,64].

As discussed earlier, from the dispersion relations for Landau levels for the superconducting matter as given in Eqs. (56) and (57), it is possible to have zero modes depending upon the values of $\delta\mu$ and $\delta\epsilon_n$. These quantities are not independent parameters but are dependent dynamically on the charge neutrality condition and the gap equations. For charge neutral matter, near μ_c , the d-quark number density is larger so that $\delta\mu = \mu_E/2$ is negative. This renders $\omega_n^u(p_z) > 0$ for any value of momentum p_z . On the other hand, for $\delta\mu$ negative, ω_n^d can vanish for some values of p_z . This defines the Fermi surfaces for the superconducting d quarks. It is easy to show that the excitation energy of n th Landau level ω_n^d for the condensing d quarks vanishes for momenta $|p_{zn}| = \sqrt{\mu_\pm^2 - 2n\tilde{B}}$. Here $\mu_\pm = (\bar{\mu} \pm \sqrt{\delta\mu^2 - \Delta^2})\theta(\delta\mu - \Delta)$. Thus higher Landau levels can also have gapless modes so long as $\sqrt{\mu_\pm^2 - 2n\tilde{B}}$ is

non-negative. Gapless modes occur when the chemical potential difference $\delta\mu$ is greater than the superconducting gap. In Fig. 8(a), we have plotted the dispersion relation i.e., the excitation energy as a function of momentum for the lowest Landau level for the condensing quarks for $\mu_q = 340$ MeV and magnetic field $\tilde{B} = 10 m_\pi^2$. The superconducting gap turns out to be $\Delta = 35.3$ MeV and $\delta\mu = -74.5$ MeV. The dispersion for the d quarks is given as $\omega_{0-}^d = \bar{\omega}_{0-} - \delta\epsilon + \delta\mu$ while the same for the u quark is given as $\omega_{0-}^u = \bar{\omega}_{0-} + \delta\epsilon - \delta\mu$. The average chemical potential is $\bar{\mu} = 366$ MeV. Far from the pairing region, $|p_z| \sim \bar{\mu} = 366$ MeV the spectrum looks like the usual BCS-type dispersion relation. Of the two excitation energies, ω_{0-}^u shows a minimum at $p_z = \bar{\mu}$ with a value $\omega_{0-}^u(|p_z| = \bar{\mu}) \sim \Delta - \delta\mu = 110$ MeV. On the other hand, ω_{0-}^d vanishes at momenta $|p_z| = \mu_\pm$. In this breached pairing region one has only unpaired d quarks and no u quarks. This can be seen explicitly as below.

The number densities of u quarks participating in condensation is given by

$$\begin{aligned} \rho_{sc}^u &= \rho^{ur} + \rho^{ug} \\ &= \sum_n \frac{\alpha_n \tilde{B}}{(2\pi)^2} \int dp_z \left[\frac{1}{2} \left(1 - \frac{\bar{\zeta}_{n-}}{\bar{\omega}_{n-}} \right) (1 - \theta(-\omega_n^d)) \right. \\ &\quad \left. - \frac{1}{2} \left(1 - \frac{\bar{\zeta}_{n+}}{\bar{\omega}_{n+}} \right) \right], \end{aligned} \quad (104)$$

where $\bar{\zeta}_{n-} = \bar{\epsilon}_n - \bar{\mu}$, $\mu = \frac{\mu_1 + \mu_2}{2}$ and $\bar{\epsilon} = \frac{\epsilon_u + \epsilon_d}{2}$. This is because $\omega_{n-}^u = \bar{\omega}_{n-} - \delta\mu + \delta\epsilon$ is always positive as $\delta\mu = \frac{\mu_- - \mu^d}{2}$ is negative and the theta function $\theta(-\omega_n^d)$ does not

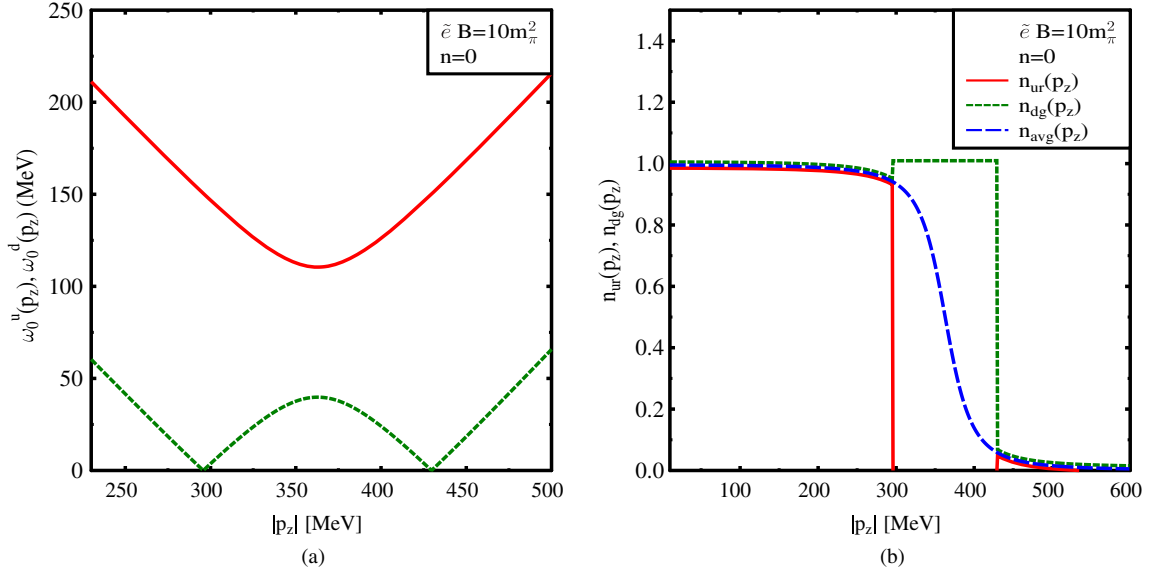


FIG. 8. Dispersion relation and the occupation number for condensing quarks at $T = 0$, $\mu_q = 340$ MeV. Part (a) shows the dispersion relation for the condensing quarks for zeroth Landau level. The upper curve is for u quark and the lower curve corresponds to d quark dispersion relation. Part (b) shows the occupation number as a function of momentum for $\tilde{e}B = 10 m_\pi^2$.

contribute. Similarly the density of d quarks participating in condensation is given by

$$\begin{aligned} \rho_{sc}^d &= \rho^{dr} + \rho^{dg} \\ &= \sum_n \alpha_n \frac{\tilde{e}B}{(2\pi)^2} \int dp_z \left[\theta(-\omega_n^d) + \frac{1}{2} \left(1 - \frac{\tilde{\zeta}_{n-}}{\tilde{\omega}_{n-}} \right) (1 - \theta(-\omega_n^d)) \right. \\ &\quad \left. - \frac{1}{2} \left(1 - \frac{\tilde{\zeta}_{n+}}{\tilde{\omega}_{n+}} \right) \right]. \end{aligned} \quad (105)$$

For positive ω_{n-}^d , the θ -function contributions vanishes and the distribution functions are the BCS distribution function. On the other hand, when $|p_z| \in [P_{n-}, P_{n+}]$, ω_n^d is negative leading to ρ_{sc}^d to vanish but for the antiparticle contribution. In this region of momenta, ρ_{sc}^d is unity. We have plotted in Fig. 8(b) the occupation number of the up and down quarks that take part in condensation as a function of the magnitude of momentum p_z i.e., the integrands of Eqs. (104) and (105) respectively for the lowest Landau level. It is easy to see from Eqs. (104) and (105) e.g., for the lowest Landau level that, except for the interval (μ_-, μ_+) , the distribution function is like the BCS distribution function. This is shown by the blue long-dashed line. The u-quark distribution is shown by the red solid line while the d-quark distribution is shown by the green short dashed line. Indeed, except for the interval (μ_-, μ_+) , all three curves overlap with each other. In the ‘‘gapless’’ momentum region, the u-quark occupation vanishes while d-quark occupation is unity. This leads to the fact that the momentum integrated distribution function for the condensing u and d quarks is not the same for the gapless region unlike the usual BCS phase. We have plotted the

number densities for the u and d quarks in Fig. 9 which shows a fork structure in the gapless region.

Gapless modes have been considered earlier for two flavor quark matter both with [37,38] and without magnetic field [35,36]. However it has been shown [65,66] that in QCD at zero temperature the gapless 2SC phases are unstable. This instability manifests itself in imaginary Meissner mass of some species of the gluons. Finite temperature calculations [67] show that at some critical value of temperature the instability vanishes. This value

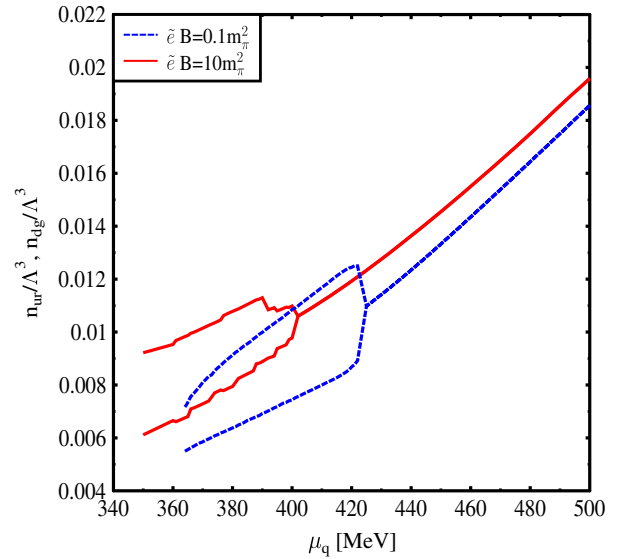


FIG. 9. Number densities of up and down quarks participating in the superconductivity for $\tilde{e}B = 0.1 m_\pi^2$ (dashed line) and $\tilde{e}B = 10 m_\pi^2$ (solid line).

may range from few MeV to tens of MeV. The instability of the gapless phases indicates that there should be other phases of quark matter breaking translational invariance e.g., inhomogeneous phase of quark matter like crystalline color superconductivity [68,69]. One may note that these considerations apply to the case without magnetic field and may change in the presence of a strong magnetic field.

In Fig. 10, we have plotted the electric and color chemical potentials μ_E and μ_8 to maintain the electric and color charge neutrality conditions given in Eqs. (94) and (95) as a function of quark chemical potential. For $2 + 1$ flavor matter, strange quarks play an important

role in maintaining charge neutrality. As the quark chemical potential increases, $|\mu_E|$ increases to maintain charge neutrality. When the chemical potential becomes large enough for strange quarks to contribute to densities, they also help in maintaining charge neutrality. This leads to a decrease in electron density or the corresponding chemical potential $|\mu_E|$. This behavior is reflected in Figs. 10(a) and 10(b) as the initial slow rise of the $|\mu_E|$. However, as $|\mu_E|$ increases, the difference $\delta\mu = -\mu_E/2$ also increases and at μ_1 , the condition $\delta\mu > \Delta$ for gapless modes to exist ceases to be satisfied. At the gapless to BCS transition point, the u-quark number density increases while that of d quarks

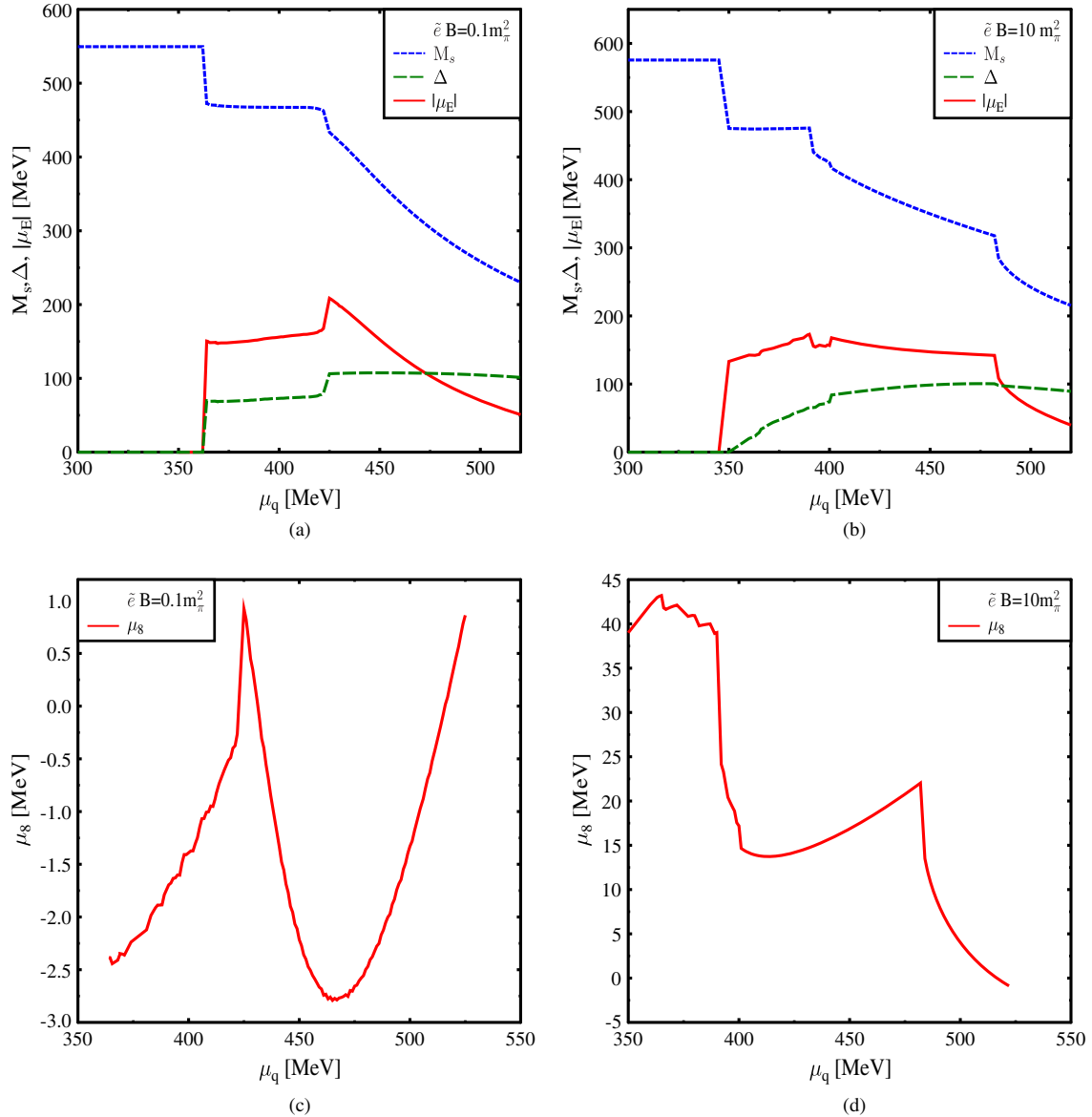


FIG. 10. Chemical potential μ_E and μ_8 for charge neutral quark matter. $|\mu_E|$ is plotted as a function of quark chemical potential μ_q for magnetic field $\tilde{e}B = 0.1 m_\pi^2$ (a) and for $\tilde{e}B = 10 m_\pi^2$ (b). In (a) and (b) we have also plotted the mass of strange quarks and superconducting gap as a function of quark chemical potential to highlight the dependence of charge chemical potential on these two parameters. In the lower two plots, the color chemical potential μ_8 is plotted as a function μ_q for $\tilde{e}B = 0.1 m_\pi^2$ (c) and for $\tilde{e}B = 10 m_\pi^2$ (d).

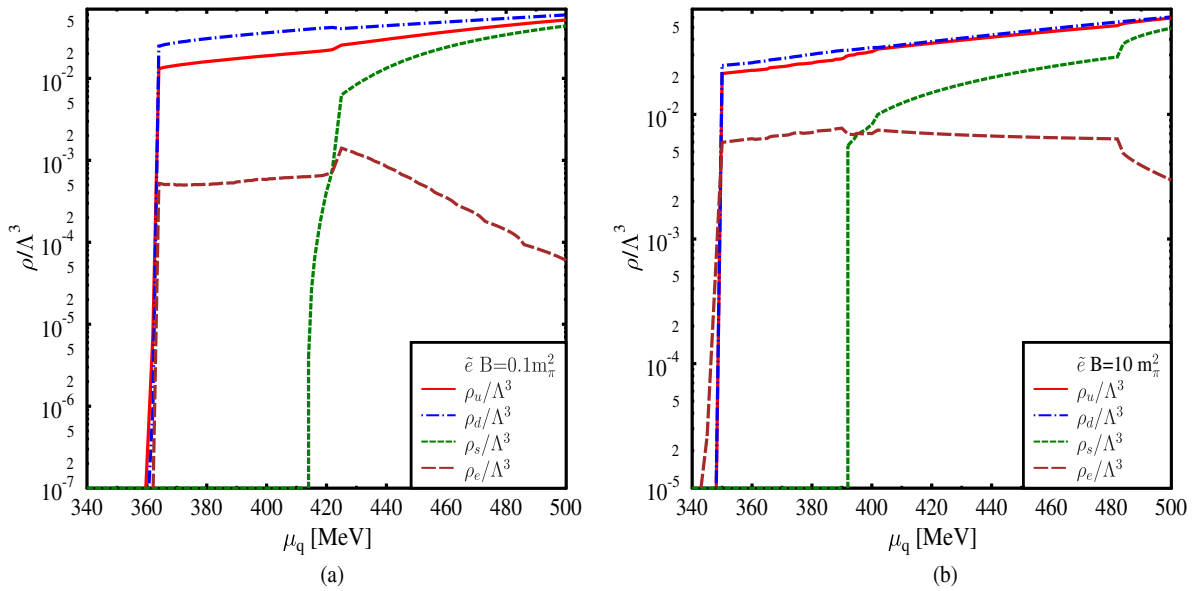


FIG. 11. Population of different species for charge neutral quark matter for $\tilde{e}B = 0.1 m_\pi^2$ (a) and for $\tilde{e}B = 10 m_\pi^2$ (b).

decreases and both become equal as in the usual BCS pairing phase. This leads to an increase in the positive electric charge density. To maintain electrical charge neutrality, the electron density increases at this point. Therefore gapless to BCS transition is accompanied with an increase in $|\mu_E|$. On the other hand, at higher densities when strange quarks start contributing to the density, it is accompanied with a drop in $|\mu_E|$ as strange quarks help in maintaining the charge neutrality along with the electrons. It turns out that for $\tilde{e}B = 0.1 m_\pi^2$, the strange quark densities become nonvanishing after the gapless to BCS transition. This leads to the continuous decrease in the $|\mu_E|$ in the BCS phase as seen in Fig. 10(a). On the other hand, for larger fields, e.g., $\tilde{e}B = 10 m_\pi^2$, chiral transition occurs at a lower μ_c due to magnetic catalysis and the strange quark density starts becoming nonvanishing at lower chemical potential. This leads to a decrease in $|\mu_E|$ at $\mu = 392$ MeV as may be seen in Fig. 10(b). At $\mu = 400$ MeV, there is the transition from the gapless to BCS phase and is accompanied with a rise in $|\mu_E|$ as discussed above. Beyond $\mu = 400$ MeV, $|\mu_E|$ starts decreasing monotonically as strange quark density increase.

In Figs. 10(c) and 10(d), we have plotted the color chemical potential μ_8 . For the weak field case, μ_8 is rather small (few MeVs) compared to both the electric chemical potential as well as the quark chemical potential which are 2 orders of magnitude larger. For the small field, the difference in densities of red and green quarks and the blue quarks essentially arises because of the difference in the distribution functions. This results in a small but finite net color charge. To maintain color neutrality one needs a small μ_8 . On the other hand, at the large magnetic field, the net color charge difference becomes larger as the \tilde{e} charges of red and green quarks and that of blue quarks are

different. This requires a somewhat larger μ_8 to maintain color neutrality as seen in Fig. 10(d). In Fig. 11 we have plotted the number densities of each species for the charge neutral matter for two different magnetic fields. As may be clear from both plots the electron number densities get correlated with the strange quark number densities.

Finally, we discuss the equation of state (EOS) for different magnetic fields. In Fig. 12 we have plotted pressure as a function of energy for $\tilde{e}B = 0.1 m_\pi^2$ and $10 m_\pi^2$. One can observe that the EOSs become stiffer with increase in magnetic field. This can be understood as follows. For $\mu < \mu_c$, the thermodynamic potential contribution from the

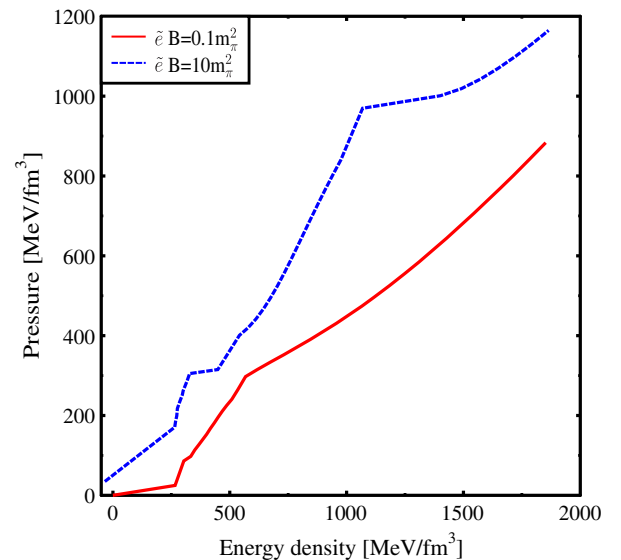


FIG. 12. Equation of state for $\tilde{e}B = 0.1 m_\pi^2$ (dashed line) and $\tilde{e}B = 10 m_\pi^2$ (solid line).

field as in Eqs. (75), (79), and (81) is dominant and decreases with an increase in magnetic field. This leads to a higher pressure for higher magnetic field. As the chemical potential increases, for $\mu > \mu_c$, the medium contribution becomes dominant. As the masses decrease with magnetic field, the medium contribution increases with magnetic field. Moreover, the field contributions also lead to an increase in pressure. Both these effects make the resulting EOS stiffer at higher magnetic field as may be seen in Fig. 12.

V. SUMMARY

We have analyzed here the effect of magnetic field and neutrality conditions on the chiral as well as diquark condensates within the framework of a three-flavor NJL model. This essentially generalizes the results of Ref. [24] to include the u-d superconductivity in the presence of a magnetic field. The methodology uses an explicit variational construct for the ground state in terms of quark-antiquark pairing for all three flavors as well as diquark pairing for the light quarks. A nice feature of the approach is that the four component quark field operator in the presence of a magnetic field could get expressed in terms of the ansatz functions that appear for the description of the ground state. Apart from the methodology being different, we also have new results. Namely, the present investigations have been done in a three-flavor NJL model along with a flavor mixing six quark determinant interaction at finite temperature and density and fields within the same framework. In that sense it generalizes the two flavor color superconductivity in the presence of a magnetic field considered earlier in Refs. [19,37,38]. The gap functions and the thermal distribution functions could be determined self-consistently for given values of the temperature, the quark chemical potential and the strength of magnetic field.

For the charge neutral matter the chiral transition is a first order transition and we observe inverse magnetic catalysis at finite density. The chiral condensate for strange quark affects the u-d superconductivity through the flavor mixing determinant interaction. The effective diquark coupling increases in the presence of strange quark condensates. On the other hand the diquark condensates contribute to the mass of the strange quark through the determinant interaction. Inverse magnetic catalysis is observed for magnetic fields up to $19 m_\pi^2$. Beyond it magnetic catalysis is observed for chiral symmetry breaking [63].

At finite densities, the effects of Landau quantization get manifested in the oscillation of the order parameters similar to the de Hass van Alphen effect for magnetization in metals. However, in the present case of dense quark matter, the order parameters, the masses and the superconducting gap themselves are dependant on the strength of magnetic fields which leads to a nonperiodic oscillation of the order parameter.

Imposition of charge neutrality condition for the quark matter leads to gapless modes even in presence of magnetic

field. The superconducting gaps in gapless modes are smaller compared to the gaps in the BCS phase. The transition from gapless to BCS phase is a sharp transition. The difference in the gap in the two phases at this transition decreases with magnetic field. For charge neutral matter the strange quark plays an important role in maintaining the charge neutrality. This leads to a depletion of electron density at higher chemical potential where strange quarks start to contribute to the densities. The resulting equation of state becomes stiffer with magnetic field.

We have considered here quark-antiquark pairing and diquark pairing in the ansatz for ground state which is homogeneous with zero total momentum. However it is possible that the condensates be spatially inhomogeneous [70] with a net total momentum [71–74]. Indeed, the gapless modes for the charge neutral matter leads to instability arising from imaginary Meissner masses for some of the gluons when $\delta\mu > \Delta$ [66]. This can be suggestive of having inhomogeneous superconducting phases [68,69] which are not considered here. The phase structure here would be nontrivial and interesting in the presence of two vectors, the magnetic field and nonzero momentum of the condensate. Furthermore, the equation of state derived for charge neutral quark matter combined with the same for hadronic matter can be used to study structural properties of neutron star with quark matter core. It will be interesting to see the compatibility of such an equation of state which is constrained by astrophysical observations like GW170817 [75]. Some of these investigations are in progress and will be reported elsewhere.

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APPENDIX: EVALUATION OF OPERATOR EXPECTATION VALUES OF SOME OPERATORS

We give here some details of the evaluation of some operators at finite T, μ and B in the state given in Eq. (18). As the state is obtained from $|0\rangle$, one can calculate the expectation values of different operators, e.g.,

$$\langle q_r^{ia\dagger}(n, k_\chi), q_{r'}^{jb}(n', k'_\chi) \rangle = \delta^{ij} \delta^{ab} \delta_{rr'} \delta_{nn'} \delta(\mathbf{k}_\chi - \mathbf{k}'_\chi) F^{ia}(\mathbf{k}_\chi), \quad (\text{A1})$$

where

$$F^{ia}(\mathbf{k}_\chi) = \sin^2 \theta_-^{ia} + \sin^2 f (1 - \sin^2 \theta_-^{ia} - |e^{ij} e^{ab}| \sin^2 \theta_-^{jb}) \times (1 - \delta^{a3})(1 - \delta^{i3}). \quad (\text{A2})$$

Similarly for the expectation values for the operators involving antiquarks, we have

$$\langle \tilde{q}_r^{ia\dagger}(n, k_\chi), \tilde{q}_r^{jb}(n', k'_\chi) \rangle = \delta^{ij} \delta^{ab} \delta_{rr'} \delta_{nn'} \delta(\mathbf{k}_\chi - \mathbf{k}'_\chi) (1 - F_1^{ia}(\mathbf{k}_\chi)), \quad (\text{A3})$$

where

$$F_1^{ia}(\mathbf{k}_\chi) = \sin^2 \theta_+^{ia} + \sin^2 f_1 (1 - \sin^2 \theta_+^{ia} - |e^{ij} \epsilon^{ab}| \sin^2 \theta_+^{jb}) (1 - \delta^{a3}) (1 - \delta^{i3}). \quad (\text{A4})$$

Using the field operator expansion of Eq. (10) and Eqs. (A1) and (A3), one can evaluate

$$\langle \psi_\alpha^{ia\dagger}(\mathbf{x}) \psi_\beta^{jb}(\mathbf{y}) \rangle = \sum_n \frac{|q_i B|}{(2\pi)^2} \int dk_\chi e^{ik_\chi(\mathbf{x}-\mathbf{y})} \Lambda_-^{ia,jb} \beta_\alpha(n, k_\chi) \quad (\text{A5})$$

with

$$\Lambda_-^{ia,jb} = \delta^{ij} \delta^{ab} [F^{ia}(n, k_z) U_{\beta r}(n, k_\chi) U_{r\alpha}(n, k_\chi)^\dagger + (1 - F_1^{ia}(n, k_z)) V_{\beta r}(n, -k_\chi) V_{r\alpha}(n, -k_\chi)^\dagger]. \quad (\text{A6})$$

Explicitly,

$$\begin{aligned} U_r(n, \mathbf{p}_\chi) U_r^\dagger(n, \mathbf{p}_\chi) &= \frac{1}{2} \begin{pmatrix} (1 + \cos \phi) I_n^2 & 0 & \hat{p}_z \sin \phi I_n^2 & i \hat{p}_\perp \sin \phi I_n I_{n-1} \\ 0 & (1 + \cos \phi) I_{n-1}^2 & -i \hat{p}_\perp \sin \phi I_n I_{n-1} & -\hat{p}_z \sin \phi I_{n-1}^2 \\ \hat{p}_z \sin \phi I_n^2 & i \hat{p}_\perp \sin \phi I_n I_{n-1} & (1 - \cos \phi) I_n^2 & 0 \\ -i \hat{p}_\perp \sin \phi I_n I_{n-1} & -\hat{p}_z \sin \phi I_{n-1}^2 & 0 & (1 - \cos \phi) I_{n-1}^2 \end{pmatrix} \\ &= \frac{1}{2} \left[I_n^2 (1 + \gamma^0 \cos \phi) \Pi^+ + I_{n-1}^2 (1 + \gamma^0 \cos \phi) \Pi^- + \frac{\hat{p}_z}{2} \sin \phi (\gamma_0 \gamma^3 (I_n^2 + I_{n-1}^2) + \gamma^5 (I_n^2 - I_{n-1}^2)) \right. \\ &\quad \left. - \hat{p}_\perp \sin \phi \gamma^2 \gamma^0 \right], \end{aligned} \quad (\text{A7})$$

where we have defined $\Pi^\pm = (1 \pm i\gamma^1 \gamma^2)/2$, $\hat{p}_z = \frac{p_z}{|p|}$, $\hat{p}_\perp = \frac{\sqrt{2nqB}}{|p|}$ with $|p| = \sqrt{p_z^2 + 2nqB}$.

Similarly for the antiquark spinors

$$\begin{aligned} V_r(n, -\mathbf{p}_\chi) V_r^\dagger(n, -\mathbf{p}_\chi) &= \frac{1}{2} \begin{pmatrix} (1 - \cos \phi) I_n^2 & 0 & -\hat{p}_z \sin \phi I_n^2 & -i \hat{p}_\perp \sin \phi I_n I_{n-1} \\ 0 & (1 - \cos \phi) I_{n-1}^2 & i \hat{p}_\perp \sin \phi I_n I_{n-1} & \hat{p}_z \sin \phi I_{n-1}^2 \\ -\hat{p}_z \sin \phi I_n^2 & -i \hat{p}_\perp \sin \phi I_n I_{n-1} & (1 + \cos \phi) I_n^2 & 0 \\ i \hat{p}_\perp \sin \phi I_n I_{n-1} & \hat{p}_z \sin \phi I_{n-1}^2 & 0 & (1 + \cos \phi) I_{n-1}^2 \end{pmatrix} \\ &= \frac{1}{2} \left[I_n^2 (1 - \gamma^0 \cos \phi) \Pi^+ + I_{n-1}^2 (1 - \gamma^0 \cos \phi) \Pi^- - \frac{\hat{p}_z}{2} \sin \phi (\gamma_0 \gamma^3 (I_n^2 + I_{n-1}^2) + \gamma^5 (I_n^2 - I_{n-1}^2)) \right. \\ &\quad \left. + \hat{p}_\perp \sin \phi \gamma^2 \gamma^0 \right]. \end{aligned} \quad (\text{A8})$$

This leads to, e.g., for the expectation value of chiral condensate for a given flavor as

$$I_s^i = \langle \bar{\psi}^i \psi^i \rangle = -\frac{1}{(2\pi)^2} \sum_n \sum_a \int dp_y dp_z (1 - F^{ia} - F_1^{ia}) \cos \phi_n^i (I_n^2 + I_{n-1}^2). \quad (\text{A9})$$

One can integrate over dp_y to obtain the contribution for the quarks that are charged as

$$I_s^i = \sum_a \sum_n \frac{\alpha_n}{(2\pi)^2} |q_i B| \int dp_z (1 - F^{ia} - F_1^{ia}) \cos \phi_n^i. \quad (\text{A10})$$

On the other hand, the contribution to the scalar condensate from the quarks that are neutral (down blue and strange blue) is given as

$$I_s^i = \frac{2}{(2\pi)^3} \int d\mathbf{p} \cos \phi^i (1 - \sin^2 \theta_-^{i3} - \sin^2 \theta_+^{i3}) \quad (i = 2, 3). \quad (\text{A11})$$

Next, we discuss about the contributions to diquark condensates. Similar to Eq. (A12), we have

$$\begin{aligned} \langle q_r^{ia}(n, k_\chi), q_{r'}^{jb}(n', k'_\chi) \rangle &= r \delta_{r,-r'} \epsilon^{ij} \epsilon^{3ab} \delta_{nn'} \delta(\mathbf{k}_\chi + \mathbf{k}'_\chi) \sin 2f(n, k_z) (1 - \sin^2 \theta_-^{ia} - \sin^2 \theta_-^{jb}) \\ &\equiv r \delta_{r,-r'} \epsilon^{ij} \epsilon^{3ab} \delta_{nn'} \delta(\mathbf{k}_\chi + \mathbf{k}'_\chi) G(k_z, n) \end{aligned} \quad (\text{A12})$$

and, for antiquark operators

$$\begin{aligned} \langle \tilde{q}_r^{ia}(n, k_\chi), \tilde{q}_{r'}^{jb}(n', k'_\chi) \rangle &= r \delta_{r,-r'} \epsilon^{ij} \epsilon^{3ab} \delta_{nn'} \delta(\mathbf{k}_\chi + \mathbf{k}'_\chi) \sin 2f(n, k_z) (1 - \sin^2 \theta_-^{ia} - \sin^2 \theta_-^{jb}) \\ &\equiv r \delta_{r,-r'} \epsilon^{ij} \epsilon^{3ab} \delta_{nn'} \delta(\mathbf{k}_\chi + \mathbf{k}'_\chi) G_1(k_z, n). \end{aligned} \quad (\text{A13})$$

For the diquark condensates we have

$$\langle \psi_\alpha^{ia}(\mathbf{x}) \psi_\beta^{jb}(\mathbf{y}) \rangle = \epsilon^{ij} \epsilon^{3ab} \sum_n \frac{|q_i B|}{(2\pi)^2} \int dk_\chi e^{ik_\chi(\mathbf{x}-\mathbf{y})} [P_u C \gamma^5 G(k_z, n) + P_v C \gamma^5 G_1(k_z, n)]_{\beta\alpha}, \quad (\text{A14})$$

where $P_u C \gamma^5 = \sum_r r U_{ar} U'_{-r\beta}$ and $P_v C \gamma^5 = \sum_r r V_{ar} V'_{-r\beta}$ and the prime on the spinors denotes a spinor with opposite charge and momentum corresponding to the unprimed spinors. Explicitly,

$$P_u = \frac{1}{2} \begin{pmatrix} \cos \frac{\phi}{2} \cos \frac{\phi'}{2} I_n^2 & 0 & \hat{p}_z \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n^2 & i \hat{p}_\perp \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n I_{n-1} \\ 0 & \cos \frac{\phi}{2} \cos \frac{\phi'}{2} I_{n-1}^2 & -i \hat{p}_\perp \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n I_{n-1} & -\hat{p}_z \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_{n-1}^2 \\ \hat{p}_z \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n^2 & i \hat{p}_\perp \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n I_{n-1} & \sin \frac{\phi}{2} \sin \frac{\phi'}{2} I_n^2 & 0 \\ -i \hat{p}_\perp \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_n I_{n-1} & -\hat{p}_z \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_{n-1}^2 & 0 & \sin \frac{\phi}{2} \sin \frac{\phi'}{2} I_{n-1}^2 \end{pmatrix} \quad (\text{A15})$$

and

$$P_v = \frac{1}{2} \begin{pmatrix} -\sin \frac{\phi}{2} \sin \frac{\phi'}{2} I_n^2 & 0 & \hat{p}_z \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_n^2 & i \hat{p}_\perp \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_n I_{n-1} \\ 0 & -\sin \frac{\phi}{2} \sin \frac{\phi'}{2} I_{n-1}^2 & -i \hat{p}_\perp \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_n I_{n-1} & -\hat{p}_z \sin \frac{\phi}{2} \cos \frac{\phi'}{2} I_{n-1}^2 \\ \hat{p}_z \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n^2 & i \hat{p}_\perp \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n I_{n-1} & -\cos \frac{\phi}{2} \cos \frac{\phi'}{2} I_n^2 & 0 \\ -i \hat{p}_\perp \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_n I_{n-1} & -\hat{p}_z \cos \frac{\phi}{2} \sin \frac{\phi'}{2} I_{n-1}^2 & 0 & -\cos \frac{\phi}{2} \cos \frac{\phi'}{2} I_{n-1}^2 \end{pmatrix}. \quad (\text{A16})$$

This leads to, e.g., for expectation value of the diquark condensate as

$$\begin{aligned} I_D &= \langle \bar{\psi}_c^{ia} \gamma^5 \psi^{jb} \rangle \epsilon^{ij} \epsilon^{3ab} \\ &= \frac{2}{(2\pi)^2} \sum_n \alpha_n |q_i B| \int dp_z \cos \left(\frac{\phi_1 - \phi_2}{2} \right) [\sin 2f(1 - \sin^2 \theta_-^1 - \sin^2 \theta_-^2) + \sin 2f_1(1 - \sin^2 \theta_+^1 - \sin^2 \theta_+^2)]. \end{aligned} \quad (\text{A17})$$

[1] For reviews see K. Rajagopal and F. Wilczek, [arXiv:hep-ph/0011333](#); D. K. Hong, *Acta Phys. Pol. B* **32**, 1253 (2001); M. G. Alford, *Annu. Rev. Nucl. Part. Sci.* **51**, 131 (2001);

G. Nardulli, *Riv. Nuovo Cimento* **25**, 1 (2002); S. Reddy, *Acta Phys. Pol. B* **33**, 4101 (2002); T. Schaefer, [arXiv:hep-ph/0304281](#); D. H. Rischke, *Prog. Part. Nucl. Phys.* **52**, 197

- (2004); H. C. Ren, arXiv:hep-ph/0404074; M. Huang, *Int. J. Mod. Phys. E* **E14**, 675 (2005); I. Shovkovy, *Found. Phys.* **35**, 1309 (2005).
- [2] D. Kharzeev, L. McLerran, and H. Warringa, *Nucl. Phys.* **A803**, 227 (2008); K. Fukushima, D. Kharzeev, and H. Warringa, *Phys. Rev. D* **78**, 074033 (2008).
- [3] V. Skokov, A. Illarionov, and V. Toneev, *Int. J. Mod. Phys. A* **24**, 5925 (2009).
- [4] D. E. Kharzeev, J. Liao, S. A. Voloshin, and G. Wang, *Prog. Part. Nucl. Phys.* **88**, 1 (2016); D. Kharzeev, *Ann. Phys. (N.Y.)* **325**, 205 (2010); K. Fukushima, M. Ruggieri, and R. Gatto, *Phys. Rev. D* **81**, 114031 (2010).
- [5] M. D'Elia, S. Mukherjee, and F. Sanfilippo, *Phys. Rev. D* **82**, 051501 (2010).
- [6] A. J. Mizher, M. N. Chenodub, and E. Fraga, *Phys. Rev. D* **82**, 105016 (2010).
- [7] D. Bandyopadhyaya, S. Chakrabarty, and S. Pal, *Phys. Rev. Lett.* **79**, 2176 (1997); S. Chakrabarty and S. Mandal, *Phys. Rev. C* **75**, 015805 (2007).
- [8] R. C. Duncan and C. Thompson, *Astrophys. J.* **392**, L9 (1992).
- [9] C. Thompson and R. C. Duncan, *Astrophys. J.* **408**, 194 (1993).
- [10] C. Thompson and R. C. Duncan, *Mon. Not. R. Astron. Soc.* **275**, 255 (1995).
- [11] C. Thompson and R. C. Duncan, *Astrophys. J.* **473**, 322 (1996).
- [12] C. Y. Cardall, M. Prakash, and J. M. Lattimer, *Astrophys. J.* **554**, 322 (2001).
- [13] A. E. Broderick, M. Prakash, and J. M. Lattimer, *Phys. Lett. B* **531**, 167 (2002).
- [14] D. Lai and S. L. Shapiro, *Astrophys. J.* **383**, 745 (1991).
- [15] E. J. Ferrer, V. Incera, J. P. Keith, and P. Springsteen, *Phys. Rev. C* **82**, 065802 (2010).
- [16] X. G. Huang, M. Huang, D. H. Rischke, and A. Sedrakian, *Phys. Rev. D* **81**, 045015 (2010).
- [17] K. Fukushima, M. Ruggieri, and R. Gatto, *Phys. Rev. D* **81**, 114031 (2010).
- [18] E. V. Gorbar, V. A. Miransky, and I. Shovkovy, *Phys. Rev. C* **80**, 032801(R) (2009); *Phys. Lett. B* **695**, 354 (2011).
- [19] Sh. Fayazbakhsh and N. Sadhooghi, *Phys. Rev. D* **82**, 045010 (2010); **83**, 025026 (2011).
- [20] V. P. Gusynin, V. Miranski, and I. Shovkovy, *Phys. Rev. Lett.* **73**, 3499 (1994); *Phys. Lett. B* **349**, 477 (1995); *Nucl. Phys.* **B462**, 249 (1996); E. J. Ferrer and V. de la Incera, *Phys. Rev. Lett.* **102**, 050402 (2009); *Nucl. Phys.* **B824**, 217 (2010).
- [21] D. Ebert and K. G. Klimenko, *Nucl. Phys.* **A728**, 203 (2003).
- [22] J. K. Boomsma and D. Boer, *Phys. Rev. D* **81**, 074005 (2010).
- [23] D. P. Menezes, M. Benghi Pinto, S. S. Avancini, and C. Providencia, *Phys. Rev. C* **80**, 065805 (2009); D. P. Menezes, M. Benghi Pinto, S. S. Avancini, A. P. Martinez, and C. Providencia, *Phys. Rev. C* **79**, 035807 (2009).
- [24] B. Chatterjee, H. Mishra, and A. Mishra, *Phys. Rev. D* **84**, 014016 (2011).
- [25] M. Alford, K. Rajagopal, and F. Wilczek, *Phys. Lett. B* **422**, 247 (1998); *Nucl. Phys.* **B537**, 443 (1999).
- [26] M. Alford, C. Kouvaris, and K. Rajagopal, *Phys. Rev. Lett.* **92**, 222001 (2004).
- [27] K. Rajagopal and A. Schmitt, *Phys. Rev. D* **73**, 045003 (2006).
- [28] M. Buballa, *Phys. Rep.* **407**, 205 (2005).
- [29] S. P. Klevensky, *Rev. Mod. Phys.* **64**, 649 (1992).
- [30] E. J. Ferrer, V. de la Incera, and C. Manuel, *Phys. Rev. Lett.* **95**, 152002 (2005); E. J. Ferrer and V. de la Incera, *Phys. Rev. Lett.* **97**, 122301 (2006); *Phys. Rev. D* **76**, 114012 (2007).
- [31] J. Noronah and I. Shovkovy, *Phys. Rev. D* **76**, 105030 (2007).
- [32] K. Fukushima and H. J. Warringa, *Phys. Rev. Lett.* **100**, 032007 (2008).
- [33] E. V. Gorbar, *Phys. Rev. D* **62**, 014007 (2000).
- [34] M. G. Alford, J. Berges, and K. Rajagopal, *Nucl. Phys.* **B571**, 269 (2000).
- [35] M. Huang and I. Shovkovy, *Nucl. Phys.* **A729**, 835 (2003).
- [36] A. Mishra and H. Mishra, *Phys. Rev. D* **71**, 074023 (2005).
- [37] T. Mandal, P. Jaikumar, and S. Digal, arXiv:0912.1413; T. Mandal and P. Jaikumar, *Phys. Rev. C* **87**, 045208 (2013); *Phys. Rev. D* **94**, 074016 (2016).
- [38] M. Coppola, P. Allen, A. G. Grunfeld, and K. N. N. Scoccola, *Phys. Rev. D* **96**, 056013 (2017).
- [39] A. Mishra and H. Mishra, *Phys. Rev. D* **69**, 014014 (2004).
- [40] H. Mishra and S. P. Misra, *Phys. Rev. D* **48**, 5376 (1993).
- [41] H. Mishra and J. C. Parikh, *Nucl. Phys.* **A679**, 597 (2001).
- [42] K. Bhattacharya, arXiv:0705.4275; M. de J. Anguiano-Galicia, A. Bashir, and A. Raya, *Phys. Rev. D* **76**, 127702 (2007).
- [43] A. Mishra and S. P. Misra, *Z. Phys. C* **58**, 325 (1993).
- [44] H. Umezawa, H. Matsumoto, and M. Tachiki, *Thermofield Dynamics and Condensed States* (North-Holland, Amsterdam, 1982); P. A. Henning, *Phys. Rep.* **253**, 235 (1995).
- [45] A. Mishra and H. Mishra, *J. Phys. G* **23**, 143 (1997).
- [46] M. A. Metlitsky and A. R. Zhitnitsky, *Phys. Rev. D* **72**, 045011 (2005).
- [47] A. Mishra and H. Mishra, *Phys. Rev. D* **74**, 054024 (2006).
- [48] M. Alford and A. Sedrakian, *J. Phys. G* **37**, 075202 (2010).
- [49] A. W. Steiner, *Phys. Rev. D* **72**, 054024 (2005).
- [50] F. Gastineau, R. Nebauer, and J. Aichelin, *Phys. Rev. C* **65**, 045204 (2002).
- [51] A. A. Abrikosov and L. P. Gorkov, *Sov. Phys. JETP* **12**, 1243 (1961).
- [52] M. G. Alford, J. Berges, and K. Rajagopal, *Phys. Rev. Lett.* **84**, 598 (2000).
- [53] E. Elizalde, *J. Phys. A* **18**, 1637 (1985).
- [54] T. Hatsuda and T. Kunihiro, *Phys. Rep.* **247**, 221 (1994).
- [55] M. Lutz, S. Klimt, and W. Weise, *Nucl. Phys.* **A542**, 521 (1992).
- [56] P. Rehberg, S. P. Klevansky, and J. Huefner, *Phys. Rev. C* **53**, 410 (1996).
- [57] M. Buballa and M. Oertel, *Nucl. Phys.* **A703**, 770 (2002).
- [58] A. W. Steiner, S. Reddy, and M. Prakash, *Phys. Rev. D* **66**, 094007 (2002).
- [59] S. B. Ruester, V. Werth, M. Buballa, I. Shovkovy, and D. H. Rischke, *Phys. Rev. D* **72**, 034004 (2005); S. B. Ruester, I. Shovkovy, and D. H. Rischke, *Nucl. Phys.* **A743**, 127 (2004).
- [60] D. Blaschke, S. Fredriksson, H. Griegorian, A. M. Oztas, and F. Sandin, *Phys. Rev. D* **72**, 065020 (2005).
- [61] M. Buballa, *Phys. Rep.* **407**, 205 (2005).
- [62] K. Schertler, S. Leupold, and J. Schaffner-Bielich, *Phys. Rev. C* **60**, 025801 (1999).

- [63] F. Preis, A. Rebhan, and A. Schmitt, *J. High Energy Phys.* **03** (2011) 033; *Lect. Notes Phys.* **871**, 51 (2013).
- [64] W. V. Liu and F. Wilczek, *Phys. Rev. Lett.* **90**, 047002 (2003); E. Gubankova, W. V. Liu, and F. Wilczek, *Phys. Rev. Lett.* **91**, 032001 (2003).
- [65] I. Giannakis and H.-C. Ren, *Phys. Lett. B* **611**, 137 (2005).
- [66] M. Huang and I. Shovkovy, *Phys. Rev. D* **70**, 094030 (2004).
- [67] K. Fukushima, *Phys. Rev. D* **72**, 074002 (2005).
- [68] K. Rajagopal and R. Sharma, *Phys. Rev. D* **74**, 094019 (2016).
- [69] J. A. Bowers and K. Rajagopal, *Phys. Rev. D* **66**, 065002 (2002).
- [70] S. Sarkar and R. Sharma, *Phys. Rev. D* **96**, 094025 (2017).
- [71] D. Nickel, *Phys. Rev. D* **80**, 074025 (2009).
- [72] G. Baser, G. Dunne, and D. Kharzeev, *Phys. Rev. Lett.* **104**, 232301 (2010).
- [73] H. Abuki, *Phys. Rev. D* **98**, 054006 (2018).
- [74] I. E. Frolov, V. Ch. Zhukovsky, and K. G. Klimenko, *Phys. Rev. D* **82**, 076002 (2010).
- [75] F. J. Fattoyev, J. Piekarewicz, and C. J. Horowitz, *Phys. Rev. Lett.* **120**, 172702 (2018).