Electrical and thermal conductivities of hot and dense hadronic matter

Guruprasad Kadam,[1,*] Hiranmaya Mishra,[2,†] and Lata Thakur[2,‡]

[1]*Department of Physics, Shivaji University, Kolhapur, Maharashtra 416004, India*
[2]*Theory Division, Physical Research Laboratory, Navarangpura, Ahmedabad 380009, India*

We estimate the electrical and thermal conductivities of hot and dense hadronic matter in the relaxation time approximation of the Boltzmann equation. We estimate the thermodynamical quantities of hot and dense hadronic matter within the ambit of the excluded volume hadron resonance gas model. The relaxation time for all the hadrons is estimated assuming the constant cross section with uniform as well as mass dependent hard-core radius. We compare our results with various existing results. Finally we give an estimate of electrical and thermal conductivities in the context of heavy ion collision experiments.

I. INTRODUCTION

Transport coefficients of hot and dense matter are one of the challenging contemporary research interests, particularly in the field of strong interaction physics. These are interesting quantities for several reasons. For many physical systems, through their dependences on system parameters like temperature, chemical potential can reveal the location of the phase transition in the phase diagram. In the context of heavy ion collisions (HICs), the matter produced in the fireball after a collision, with quarks and gluons degrees of freedom, behaves like a strongly interacting liquid, with a small shear viscosity it expands, cools, and undergoes a crossover transition to hadronic degrees of freedom, which finally free stream to the detector. One of the successful descriptions of such an evolution is through dissipative relativistic hydrodynamics [1–9] and transport simulations [10–17]. Finite but small shear viscosity (η) to entropy (s) ratio is necessary to explain the flow data [18,19]. The smallness of this ratio $\frac{\eta}{s}$ and its connection to the conjectured Kovtun-Son-Starinets bound of $\frac{\eta}{s} = \frac{1}{4\pi}$ obtained using AdS/CFT correspondence [20] has motivated many theoretical investigations of this ratio to understand and derive rigorously from a microscopic theory [21–29]. The other viscosity coefficient ζ has also been realized to be important to be included the dissipative hydrodynamics. During the expansion of the fireball, when

the temperature approaches the critical temperature, ζ can be large and give rise to different interesting phenomena, like cavitation when the pressure vanishes, and hydrodynamic description breaks down [30,31]. The effect of bulk viscosity on the particle spectra and flow coefficients has been investigated [32–34], while the interplay of shear and bulk viscosity coefficients have been studied in Refs. [35–37]. The coefficient of bulk viscosity has been estimated for both the hadronic and partonic systems [38–50]. In the case of noncentral and asymmetric HICs, a large magnetic field as well as electric field is expected to be produced [51,52]. The event by event analysis for the relativistic heavy ion collision system indicates the generation of the magnetic field of the order of $eB \simeq m_\pi^2$ as well as the electric field $eE \simeq m_\pi^2$. The strong magnetic field so produced has exciting possibilities of observing CP violating effects, known as the chiral magnetic and chiral vortical effects. Apart from these, there have been other dynamical manifestations of such strong fields on other observables, like an increase in the elliptical flow coefficient. However, all these interesting and important effects in off central heavy ion collisions require that a reasonably strong magnetic field survives for at least several Fermi proper time. Initially it was thought that the magnetic field decays rapidly after the collision [53]. It was later pointed out that the rapid decrease in the magnetic field leads to induced electric current that slows down the decrease of the magnetic field and satisfies a diffusion equation [54,55]. The crucial parameter that goes in to the estimation of the timescale of this diffusion is the electrical conductivity of the medium σ_{ele}. The time evolution of the magnetic field in relativistic heavy ion collisions is still an open question. This requires a proper estimate of the electrical conductivity of the medium as well as solutions of magneto-hydrodynamic equations, which need further investigation [54,56]. This apart, σ_{ele} also enters in the hydrodynamic

[*]guruprasadkadam18@gmail.com
[†]hm@prl.res.in
[‡]latathakur@prl.res.in

evolution, where charge relaxation also plays an important role. This coefficient influences significantly the soft photon production [57] as well as low mass dilepton enhancement [58].

Several groups have studied the electrical conductivity, including the chiral perturbation theory [59], the numerical solution of the Boltzmann equation [60,61], holography [62], transport models [63,64], Dyson Schwinger calculations [65], a dynamical quasiparticle model [66,67], a quasiparticle model [29,68], the effective fugacity quasiparticle model [69], and lattice gauge theory [70–76]. All these studies aim at the value of σ_{el} in the quark gluon plasma (QGP) phase, but some of these do extend below the transition temperature towards the hadron gas. Despite the importance of electrical conductivity, it has rarely been studied in the literature for the hadronic phase. Recently, σ_{el} has been investigated for a pion gas [77] and a hot hadron gas [78–80]. It has also been studied in the framework of the Polyakov-Nambu-Jona-Lasinio model [81] and the Polyakov-Quark-Meson (PQM) model [82].

The transport coefficient that plays an important role in the hydrodynamic evolution at finite baryon densities is the coefficient of thermal conductivity (κ). The effects of thermal conductivity in the relativistic hydrodynamics has been recently emphasized in Refs. [83,84]. The thermal conduction, which involves relative flow of energy and baryon number, vanishes at zero baryon density. However, for situations, where, e.g., the pion number is conserved, particularly at low temperatures, heat conductivity can be sustained by pions which themselves have zero baryon number [21]. Recently, thermal conductivity has been studied for pionic medium by different groups [21,22, 39,85–87]. The heat conductivity was also obtained using the Kubo formula [41,88,89] and Nambu-Jona-Lasinio (NJL) model [67,90]. Heat conductivity has been investigated recently in a transport model [16] and a PQM coupling model [91].

We might note here that it is of practical as well as fundamental importance to estimate the transport coefficients also in the hadronic phase to distinguish the signatures of QGP matter and hadronic matter. These coefficients can be estimated directly within QCD using Kubo formulation. However, as QCD is strongly coupled for the energies accessible in heavy ion collision experiments, the task is very nontrivial. First principle calculations like lattice QCD simulation are also challenging and are limited to equilibrium properties at small chemical potentials. These coefficients therefore have been estimated within various effective models for strong interaction as well as various approximations in the estimation.

In the present work we intend to estimate the coefficients of electrical conductivity and thermal conductivity for the hadronic phase within the ambit of a hadron resonance gas model (HRGM). The HRG model, which successfully describes the hadronic phase with the multiplicities of

particle abundances of various hadrons in heavy ion collisions [92–94], is assumed to be a free gas of all observed hadrons and their resonances treated as point particles. As shown in Ref. [95], this is a reasonable way to include attractive interaction among hadrons. Apart from hadronic multiplicities, this model has been used to estimate viscosity coefficients [43,96–99] as well as the study of fluctuations in conserved charges in HIC experiments [100,101]. However, the simple HRGM misses the repulsive interactions among hadrons, the existence of which is already known from nucleon-nucleon scattering experiments. Such repulsive interactions can be implemented via an excluded volume approximation, whereby the volume available for the hadrons to move is reduced by the volume they occupy [102–104]. This HRGM with excluded volume (EHRGM) [96,99,105–107] corrections has been found to be in good agreement with lattice QCD results up to temperature $T \sim 140$ MeV. The model has also been used to estimate the viscosity coefficients using relaxation time approximation for solving the relativistic Boltzmann kinetic equation [99]. We use here a similar approximation to estimate the electrical and thermal conductivities of hadronic matter.

We organize the paper as follows. In Sec. II, we recapitulate the excluded volume hadron resonance gas model. In Sec. III, we compute the electrical and thermal conductivities using the relativistic Boltzmann equation in relaxation time approximation relevant for multicomponent hadronic medium. In Sec. IV, we calculate the relaxation time in the limit of isotropic constant scattering cross section for the hadrons. In Sec. V, we discuss our results and finally, in Sec. VI, we summarize findings of the present investigation.

II. EXCLUDED VOLUME HADRON RESONANCE GAS MODEL

As we have discussed in the Introduction, hadrons cannot be considered as point particles. The repulsive interactions can be taken into account between hadrons via an excluded volume approximation or van der Waals treatment. The thermodynamic pressure is related to the partition function as

$$P_{id} = T \lim_{V \to \infty} \frac{\ln Z_{id}(T, \mu, V)}{V}, \quad (1)$$

where T is temperature, μ is chemical potential, and V is volume of the system. In thermodynamically consistent excluded volume formulation, one can obtain the transcendental equation for the pressure as [107,108]

$$P^{EV}(T, \mu) = P^{id}(T, \tilde{\mu}), \quad (2)$$

where $\tilde{\mu} = \mu - vP^{EV}(T, \mu)$ is an effective chemical potential with v as the parameter corresponding to the proper volume of the particle. At high temperature and low

densities this prescription is equivalent to multiplying a suppression factor of $\exp(-vP^{EV}/T)$ to the pressure in the Boltzmann approximation. Therefore, the pressure in excluded volume hadron resonance gas model becomes

$$P^{EV}(T,\mu) = e^{\frac{-vP^{EV}(T,\mu)}{T}}P^{\mathrm{id}}(T,\mu), \qquad (3)$$

where P^{id} in the Boltzmann approximation can be written as

$$P^{\mathrm{id}}(T,\mu) = \sum_a \frac{g_a}{2\pi^2} m_a^2 T^2 K_2\left(\frac{m_a}{T}\right)\cosh\left(\frac{\mu}{T}\right), \quad (4)$$

where g_a is the degeneracy of ath hadron species. Other thermodynamical quantities can be readily obtained from Eq. (2) by taking appropriate derivatives. The number density, energy density, and entropy density, respectively, can be written as [107]

$$n^{EV}(T,\mu) = \sum_a \frac{n_a^{\mathrm{id}}(T,\tilde{\mu})}{1+\sum_a v_a n_a^{\mathrm{id}}(T,\tilde{\mu})}, \qquad (5)$$

$$\epsilon^{EV}(T,\mu) = \sum_a \frac{\epsilon_a^{\mathrm{id}}(T,\tilde{\mu})}{1+\sum_a v_a n_a^{\mathrm{id}}(T,\tilde{\mu})}, \qquad (6)$$

$$s^{EV}(T,\mu) = \sum_a \frac{s_a^{\mathrm{id}}(T,\tilde{\mu})}{1+\sum_a v_a n_a^{\mathrm{id}}(T,\tilde{\mu})}. \qquad (7)$$

Again in the Boltzmann approximation all the thermodynamical quantities are multiplied by the factor $\exp(-vP^{EV}/T)$. But unlike pressure there is an additional factor $\frac{1}{1+\sum_a v_a n_a(T,\tilde{\mu})}$, which suppresses the thermodynamical quantities at high temperature as compared to their ideal gas counterpart. Once the thermodynamic quantities are estimated, we can calculate the electrical and thermal conductivities using the EHRGM model.

III. TRANSPORT COEFFICIENTS IN RELAXATION TIME APPROXIMATION

A. Electrical conductivity

The electric conductivity (σ_{ele}) represents the response of the system to an applied electric field,

$$\boldsymbol{j} = \sigma_{\mathrm{ele}}\boldsymbol{E}. \qquad (8)$$

We start our calculation from the relativistic Boltzmann transport (RBT) equation. In the presence of an external field, the RBT equation can be written as [109,110]

$$k^\mu \partial_\mu f_a(x,k) + q_a F^{\alpha\beta} k_\beta \frac{\partial}{\partial k^\alpha} f_a(x,k) = \mathcal{C}_a[f_a], \quad (9)$$

where $F^{\alpha\beta}$ is the electromagnetic field strength tensor and $\mathcal{C}_a[f_a]$ is the collision integral. Here we have introduced the

index a on the distribution function for the hadronic species. The relaxation time approximation (RTA) is the simplest scheme to approximate the collision term $\mathcal{C}_a[f_a]$, which is given by

$$\mathcal{C}_a[f_a] \simeq -\frac{k^\mu u_\mu}{\tau_a}\delta f_a, \qquad (10)$$

where $u_\mu = (1,\mathbf{0})$ is the fluid four velocity in the local rest frame and τ_a is the relaxation time, which estimates the timescale for the system to relax towards the equilibrium state. $\delta f_a = f_a - f_a^0$, where we assume that the distribution function f_a is very close to the equilibrium distribution f_a^0 and can be written for deviation in linear order as [61]

$$f_a(x,\mathbf{k}) = f_a^0(x,\mathbf{k})(1+\varphi(x,\mathbf{k})) = f_a^0 + \delta f_a, \qquad (11)$$

where φ ($|\varphi| \ll 1$) is the perturbation. The equilibrium particle distribution function is

$$f_a^0(x,\mathbf{k}) = \frac{1}{e^{(E_a-\mu_a)/T}\pm 1}, \qquad E_a = \sqrt{\mathbf{k}^2+m_a^2}, \qquad (12)$$

where $\pm$ in the denominator corresponds to fermion and boson, respectively; $\mu_a = t_a\mu$, t_a being the baryonic charge of the hadronic species a. For constant electric field $\mathbf{E}$, Eq. (9) becomes

$$q_a\left(k_0\mathbf{E}\cdot\frac{\partial f_a^0}{\partial\mathbf{k}} + \mathbf{E}\cdot\mathbf{k}\frac{\partial f_a^0}{\partial k^0}\right) = -\frac{k^0}{\tau_a}\delta f_a. \qquad (13)$$

After solving, one can get δf_a for the case when $\varphi \ll f_a^0$ as

$$\delta f_a = \frac{q_a\tau_a}{T}\mathbf{E}\cdot\frac{\mathbf{k}}{k^0}f_a^0(1\pm f_a^0). \qquad (14)$$

The electric four current (j^μ) can be written as

$$j^\mu = \sum_a q_a g_a \int \frac{d^3k}{(2\pi)^3 E_a} k^\mu f_a(x,k), \qquad (15)$$

where $q_a(\bar{q}_a)$ and $f_a(x,k)(\bar{f}_a(x,k))$ are the charge and distribution functions for particles (antiparticles) a. After applying an external disturbance, $j^\mu = j_0^\mu + \Delta j^\mu$, four current becomes

$$\Delta j^\mu = \sum_a q_a g_a \int \frac{d^3k}{(2\pi)^3 E_a} k^\mu \delta f_a. \qquad (16)$$

Considering the definition of electrical conductivity and substituting δf_a into that, we get

$$\sigma_{\mathrm{ele}} = \frac{1}{3T}\sum_a g_a q_a^2 \int \frac{d^3k}{(2\pi)^3}\frac{k^2}{E_a^2}\tau_a \times f_a^0(1\pm f_a^0). \qquad (17)$$

In the Boltzmann approximation, the above equation can be written as

$$\sigma_{\mathrm{ele}} = \frac{1}{3T} \sum_a g_a q_a^2 \int \frac{d^3 k}{(2\pi)^3} \frac{k^2}{E_a^2} \tau_a \times f_a^0. \qquad (18)$$

B. Thermal conductivity

Thermal conductivity κ is interesting to study, as it describes the heat flow in interacting systems [111,112]. Recently it has reattained interest in the context of relativistic HICs [16,113]. We will start our calculations from the RBT equation. In the absence of external field, Eq. (9) can be written as [114]

$$k^\mu \partial_\mu f_a(x, k) = -\frac{k^\mu u_\mu}{\tau_a} \delta f_a. \qquad (19)$$

We start our calculation from the energy momentum tensor ($T^{\mu\nu}$) and the baryonic four current (j^μ), which are, respectively, given by [21,114]

$$T^{\mu\nu} = \sum_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} k^\mu k^\nu f_a(x, k) \qquad (20)$$

and

$$J^\mu = \sum_a g_a t_a \int \frac{d^3 k}{(2\pi)^3 E_a} k^\mu f_a(x, k), \qquad (21)$$

where, as before, t_a and g_a are, respectively, the charge and the degeneracy of hadronic species a. In the presence of a small disturbance from the equilibrium distribution function, the change in energy momentum tensor $\Delta T^{\mu\nu}$ can be written as

$$\Delta T^{\mu\nu} = \sum_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} k^\mu k^\nu \delta f_a(x, k). \qquad (22)$$

Using the RTA, $\Delta T^{\mu\nu}$ becomes [114]

$$\Delta T^{\mu\nu} = -\sum_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} \frac{k^\mu k^\nu}{k.u} \tau_a k^\alpha \partial_\alpha f_a(x, k) \qquad (23)$$

and the change in four current Δj^μ becomes

$$\Delta j^\mu = \sum_a g_a t_a \int \frac{d^3 k}{(2\pi)^3 E_a} \frac{k^\mu}{k.u} \tau_a k^\alpha \partial_\alpha f_a(x, k), \qquad (24)$$

where $\partial_\mu = u_\mu D + \nabla_\mu$, and the convective derivatives ($DT, D\mu, Du^\mu$) can be eliminated by using the relation

$$(\varepsilon + P) Du^\mu - \nabla^\mu P = 0, \qquad (25)$$

$$Dn + n\nabla_\mu u^\mu = 0. \qquad (26)$$

After using the above relations, one can obtain [114]

$$\Delta T^{\mu\nu} = \sum_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} \frac{k^\mu k^\nu}{k.u} \frac{1}{T} \left[\tau_a f_a^0 (1 - f_a^0) \right.$$
$$\times \left\{ k.u \left(\frac{\partial k}{\partial \varepsilon} \right)_n \nabla_\alpha u^\alpha + k^\alpha X_\alpha + \frac{k^\alpha k^\beta}{k.u} \nabla_\alpha u_\beta \right.$$
$$\left. + \left(\frac{\partial k}{\partial n} \right)_\varepsilon \nabla_\alpha u^\alpha - \frac{\varepsilon + P}{n} \frac{k^\alpha}{k.u} X_\alpha \right\} \right] \qquad (27)$$

and

$$\Delta j^\mu = \sum_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} \frac{k^\mu}{k.u} \frac{1}{T} \left[\tau_a f_a^0 (1 - f_a^0) \right.$$
$$\times \left\{ k.u \left(\frac{\partial k}{\partial \varepsilon} \right)_n \nabla_\alpha u^\alpha + k^\alpha X_\alpha + \frac{k^\alpha k^\beta}{k.u} \nabla_\alpha u_\beta \right.$$
$$\left. + \left(\frac{\partial k}{\partial n} \right)_\varepsilon \nabla_\alpha u^\alpha - \frac{\varepsilon + P}{n} \frac{k^\alpha}{k.u} X_\alpha \right\} \right], \qquad (28)$$

where

$$X_\alpha = \frac{\nabla_\alpha P}{\varepsilon + P} - \frac{\nabla_\alpha T}{T}, \qquad (29)$$

and $u_\mu = (1, \mathbf{0})$. ε and n are the energy density and baryon number density. The momentum conservation shows that $\nabla P = 0$ [where $\nabla P = (\varepsilon + P) \partial u / \partial t$] in the steady state. Thermal conduction, which involves the relative flow of energy, arises when energy flows relative to the baryonic enthalpy. The T^{0i} component is the energy flux and with the Eckart condition, $T^{0i} = \Delta T^{0i} - \frac{(\varepsilon + P)}{n} \Delta j^i \equiv I^i$, where I^i is the heat current with

$$\Delta T^{0i} = \sum_a g_a \int \frac{d^3 k}{(2\pi)^3} \frac{\mathbf{k}^2}{3T} \tau_a f_a^0 (1 - f_a^0) \left\{ 1 - \frac{\varepsilon + P}{n E_a} \right\} X_i \qquad (30)$$

and

$$\Delta j^i = \sum_a t_a g_a \int \frac{d^3 k}{(2\pi)^3 E_a} \frac{\mathbf{k}^2}{3T} \tau_a f_a^0 (1 - f_a^0) \left\{ 1 - \frac{\varepsilon + P}{n E_a} \right\} X_i. \qquad (31)$$

Using either the Eckart or Landau-Lifshitz condition, one can define the heat conductivity as [114]

$$I^i = -\kappa [\partial_i T - T \partial_i P / (\varepsilon + P)] = \kappa T X_i. \qquad (32)$$

Using Eqs. (30) and (31), one can obtain the expression for thermal conductivity as

$$\kappa = \frac{1}{3T^2} \sum_a g_a \tau_a \int \frac{d^3 k}{(2\pi)^3} \frac{\mathbf{k}^2}{E_a^2} f_a^0 (1 - f_a^0) \left(E_a - \frac{t_a \omega}{n} \right)^2, \qquad (33)$$

where $\omega = \varepsilon + P$ is the enthalpy and t_a is the baryonic charge of hadronic species a. For the baryonic matter and low temperature, the antiparticle contribution in the above sum can be neglected, as the temperatures are much smaller compared to the masses of the baryons. Since we will work in the Boltzmann approximation, the expression for thermal conductivity can be written as

$$\kappa = \frac{1}{3T^2} \sum_a g_a \tau_a \int \frac{d^3 k}{(2\pi)^3} \frac{\mathbf{k}^2}{E_a^2} f_a^0 \left(E_a - \frac{t_a \omega}{n} \right)^2. \quad (34)$$

IV. RELAXATION TIME

The relaxation time τ_a is defined by the expression [99]

$$\tau^{-1}(E_a) = \sum_{bcd} \int \frac{d^3 p_b}{(2\pi)^3} \frac{d^3 p_c}{(2\pi)^3} \frac{d^3 p_d}{(2\pi)^3} W(a, b \to c, d) f_b^0, \quad (35)$$

where $W(a, b \to c, d)$ is the transition rate,

$$W(a, b \to c, d) = \frac{(2\pi)^4 \delta(p_a + p_b - p_c - p_d)}{2E_a 2E_b 2E_c 2E_d} |\mathcal{M}|^2, \quad (36)$$

and $|\mathcal{M}|$ is the transition amplitude. Equation (35) can be simplified in the center of mass frame as

$$\tau^{-1}(E_a) = \sum_b \int \frac{d^3 p_b}{(2\pi)^3} \sigma_{ab} v_{ab} f_b^0, \quad (37)$$

where σ_{ab} is the total scattering cross section for the process, $a(p_a) + b(p_b) \to a(p_c) + b(p_d)$, and v_{ab} is relativistic relative velocity. One can obtain the averaged partial relaxation time by averaging the relaxation time over f_a^0, which is rather a good approximation [115]. Thus, the averaged relaxation time can be written as

$$\tilde{\tau}_a^{-1} = \sum_b n_b \langle \sigma_{ab} v_{ab} \rangle. \quad (38)$$

In the above, v_{ab} is the "relative velocity" defined by

$$v_{ab} = \frac{\sqrt{(p_a \cdot p_b)^2 - m_a^2 m_b^2}}{E_a E_b},$$

with p_a, E_a being the four momentum and energy of particle "a" and, with g_b being the degeneracy of species "b,"

$$n_b = \frac{g_b}{(2\pi)^3} \int \frac{d^3 p_b}{(2\pi)^3} f_b^0$$

is the equilibrium number density of bth hadronic species. Here we use the equilibrium Maxwell-Boltzmann distribution

$$f_a^0 = \exp\left(-\frac{E_a - \mu_a}{T} \right). \quad (39)$$

The thermal average of total cross section times relative velocity, i.e., $\langle \sigma v \rangle$ for the scattering of hard sphere particles of the same species at a given T and μ (having constant cross section, σ), can be calculated as outlined in Refs. [116,117]. The thermal average $\langle \sigma v \rangle$ for the process $a(p_a) + a(p_b) \to a(p_c) + a(p_d)$ can be written as

$$\langle \sigma_{ab} v_{ab} \rangle = \frac{\sigma \int d^3 p_a d^3 p_b v_{ab} e^{-E_a/T} e^{-E_b/T}}{\int d^3 p_a d^3 p_b e^{-E_a/T} e^{-E_b/T}}. \quad (40)$$

Note that, in the above, the chemical potential dependences gets canceled from the numerator and the denominator, which is a consequence of Boltzmann approximations for the equilibrium thermal distribution function. After changing the integration variable as discussed in detail in Ref. [99], the numerator and denominator in Eq. (40) respectively become

$$\int d^3 p_a d^3 p_b v_{ab} e^{-E_a/T} e^{-E_b/T}$$
$$= 2\pi^2 T \int ds \sqrt{s}(s - 4m^2) K_1(\sqrt{s}/T) \quad (41)$$

and

$$\int d^3 p_a d^3 p_b e^{-E_a/T} e^{-E_b/T} = [4\pi m^2 T K_2(m/T)]^2. \quad (42)$$

Therefore, the thermal average $\langle \sigma_{ab} v_{ab} \rangle$ can be written as [99]

$$\langle \sigma_{ab} v_{ab} \rangle = \frac{\sigma}{8m^4 T K_2^2(m/T)} \int_{4m^2}^{\infty} ds \sqrt{s}(s - 4m^2) K_1(\sqrt{s}/T), \quad (43)$$

where $\sqrt{s}$ is the center of mass energy and K_1 (K_2) is the modified Bessel function of order 1 (2). For the case of scattering between different species of the particles, Eq. (43) becomes

$$\langle \sigma_{ab} v_{ab} \rangle = \frac{\sigma}{8T m_a^2 m_b^2 K_2(\frac{m_a}{T}) K_2(\frac{m_b}{T})}$$
$$\times \int_{m_a+m_b}^{\infty} ds \frac{[s - (m_a - m_b)^2]}{\sqrt{s}}$$
$$\times [s - (m_a + m_b)^2] K_1(\sqrt{s}/T). \quad (44)$$

After evaluating the thermal averaged cross section, we can relate it to the relaxation time in Eq. (38).

V. RESULTS AND DISCUSSION

In the hadron resonance gas model, it is customary to include all the hadrons and resonances up to certain cutoff Λ. We choose cutoff $\Lambda = 2.25$ GeV and include all the mesons and baryons listed in Ref. [118]. The only parameter that remains in our model is hard-core radius
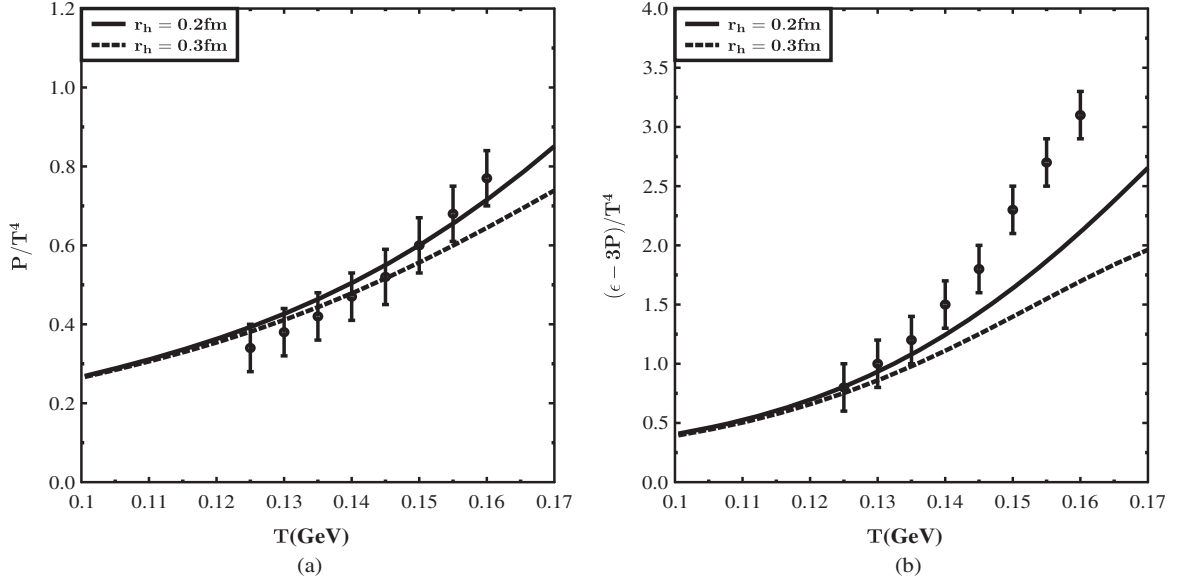
FIG. 1. Thermodynamical functions, pressure (a) and trace anomaly (b), at zero chemical potential.

r_h or the excluded volume parameter v. We choose two different parametrization schemes, viz., uniform excluded volume parameter ($v = \frac{16}{3}\pi r_h^3$) [96] and mass dependent excluded volume parameter ($v = \frac{M_h}{\epsilon_0}$) [106]. Here ϵ_0 is the parameter that we fix to 2 GeV fm^{-3}. Based on the nucleon-nucleon scattering analysis [119], we choose uniform hard-core radius $r_h = 0.3$ fm.

Figure 1 shows the thermodynamical properties of hadron gas estimated within the ambit of the EHRGM. Figure 1(a) shows scaled pressure P/T^4 at zero baryon chemical potential for two different choices of hadron hard-core radius $r_h = 0.2$ and $r_h = 0.3$ fm. We note that the EHRGM estimates deviate from the lattice data at higher temperature. The deviation is large for larger hard-core radius. This is essentially due to the suppression factor $(1 + vn^{EV})^{-1}$, which is large for higher r_h. Figure 1(b) shows the scaled interaction measure $(\epsilon - 3p)/T^4$. Again the EHRGM estimates strongly deviate from the lattice data at higher temperature. The rapid rise in the trace anomaly cannot be explained within the EHRGM alone. But it has been shown in Ref. [120] that, by including the Hagedorn mass spectrum along with the discrete hadron spectrum in the HRG model, the resulting excluded volume model reproduces the lattice data up to 160 MeV at $\mu = 0$. Similar studies extended to include finite baryon chemical potential confirm this result [121]. Note that we will not include the Hagedorn states in our calculations since their quantum numbers, especially the electric charges, are not known experimentally.

In Fig. 2, we show the dimensionless electrical conductivity (σ_{ele}/T) as a function of temperature at zero chemical potential. We have compared our results with the various results that exist in the literature. The red dashed line shows the results of the conformal Super-Yang-Mills

(SYM) plasma [122]. The red open circles represent the data from lattice QCD calculation [123]. However, the hadronic interactions are missing in the lattice calculation. The violet dotted line represents the nonconformal holographic model [62,124]. The cyan dashed line represents chiral perturbation theory (CPT) results [77]. The blue solid line shows the kinetic theory results [79]. The magenta curve shows our results for the uniform excluded
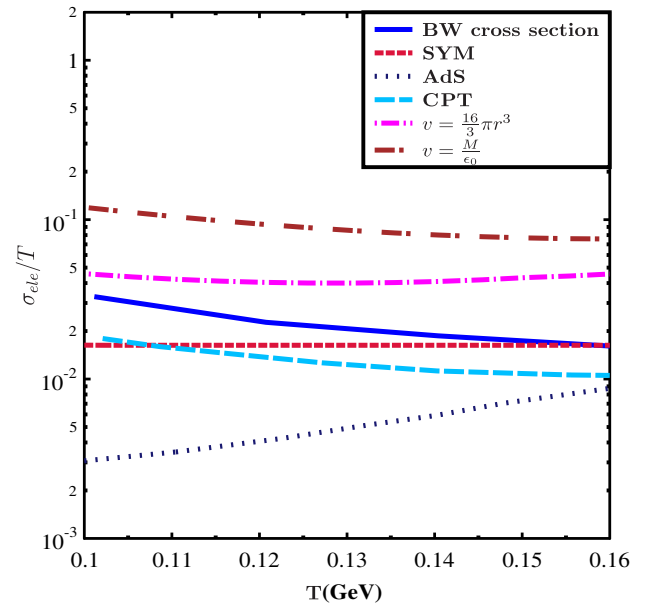

FIG. 2. Normalized conductivity in the present EHRGM with two different parametrizations for the excluded volume compared with other model estimations. The magenta dot-dashed curve refers to the EHRGM with a uniform excluded volume parameter, while the dashed double-dot curve corresponds to the EHRGM with a mass dependent excluded volume calculation.
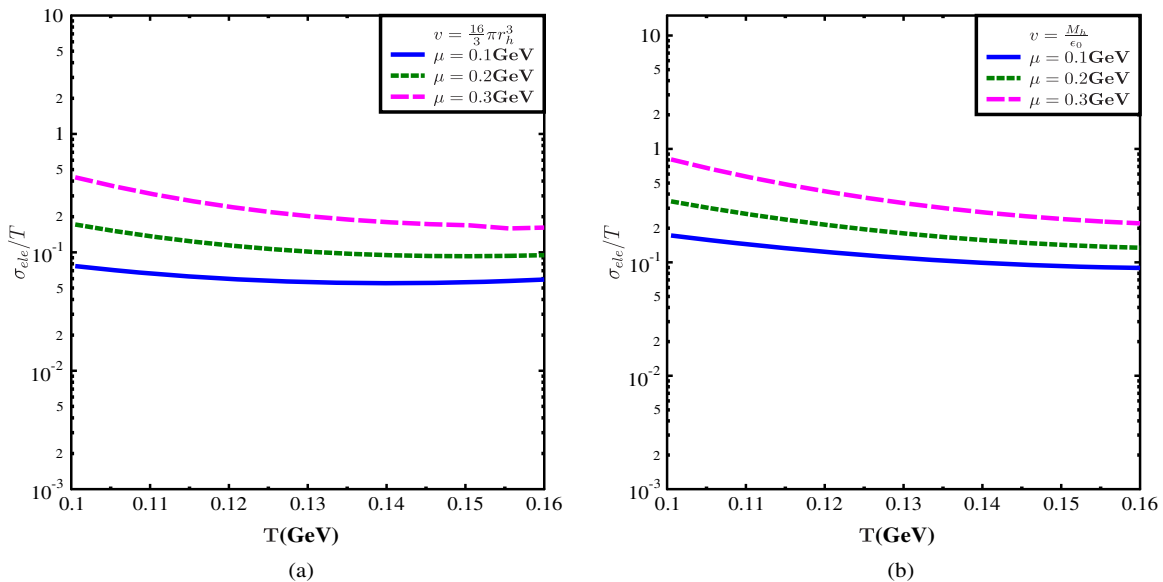
FIG. 3.   The scaled electrical conductivity as a function of temperature for different values for the baryon chemical potential. (a) The variation in the EHRGM with a uniform hard-core excluded volume. (b) Corresponds to the EHRGM model with a mass dependent excluded volume.

volume parameter $(v = \frac{16}{3}\pi r_h^3)$, while the maroon curve corresponds to the mass dependent excluded volume parameter $(v = \frac{M_h}{\epsilon_0})$. The behavior of σ_{ele}/T with temperature from the CPT and kinetic theory results are similar to our results, although there is a difference in magnitude of electrical conductivity. The magnitude of electrical conductivity is higher in the model as compared to other results, especially the kinetic theory estimations of Ref. [79] with Breit-Wigner (BW) crosssections. However, this is not so surprising. The basic reason behind higher conductivity in our model is the smaller cross section. In the case of the uniform excluded volume parameter, the cross sections ~10 mb for all the hadronic species, while in Ref. [79] different cross sections are assumed for different species and the values of the cross section are relatively large. Since the conductivity is inversely proportional to the cross section (through relaxation time τ), its estimation turns out to be large in our model. However, it may be noted that assigning hard-core radius to all the hadrons may not be the correct way to account for the repulsive interactions within the noninteracting HRG model. One possible improvement one can do to this model is to assign repulsive interactions only between baryons and antibaryons, while mesons are kept noninteracting [125]. Estimating the transport coefficients within this model is under progress and will appear elsewhere.

In Fig. 3 we show the variation of electrical conductivity with temperature for different chemical potentials, $\mu = 0.1$, 0.2, and 0.3 GeV. Figure 3(a) corresponds to the uniform excluded volume parameter, while Fig. 3(b) corresponds to the mass dependent excluded volume parameter. We note

that the electrical conductivity increases with the increase in chemical potential, although the general behavior as a function of temperature does not change. This behavior is not hard to understand. In Eq. (40), while the cross section is independent of both μ and T, the thermally averaged cross section times the relative velocity $\langle\sigma v\rangle$ is, in general, dependent on both T and μ arising from the distribution functions. However, in the Boltzmann approximation, the μ dependence gets canceled from the numerator and the denominator. On the other hand, for the thermal averaged cross section times the relative velocity or the inverse of scattering length, given by $\langle\sigma v\rangle_{ab}n_b$ [Eq. (38)], will be an increasing function of μ if species b is a baryon. So this will lead to the relaxation time being a decreasing function of μ. In the expression for the σ_{ele}, τ_a is multiplied by a distribution function, which again is an increasing function of μ when species a is a baryon. Thus, the contribution to σ_{ele} from, say, a baryon will depend upon which of the two parts τ_a (a decreasing function of μ) and the distribution function f_0 (which is an increasing function of μ) dominate the variation with μ. It turns out that, for baryons, the dominant contribution to the relaxation time arises from the baryon scattering with mesons and in that case the corresponding average cross section or, equivalently, the relaxation time is independent of μ. Therefore, the μ variation of the corresponding contribution of the baryon to σ_{ele} is an increasing function of μ.

From Fig. 3 it is clear that the σ_{ele}/T ratio is higher for the mass dependent excluded volume parameter case as compared to the uniform excluded volume parameter case for the different values of chemical potential, although the general behavior of the ratio is similar as a function of
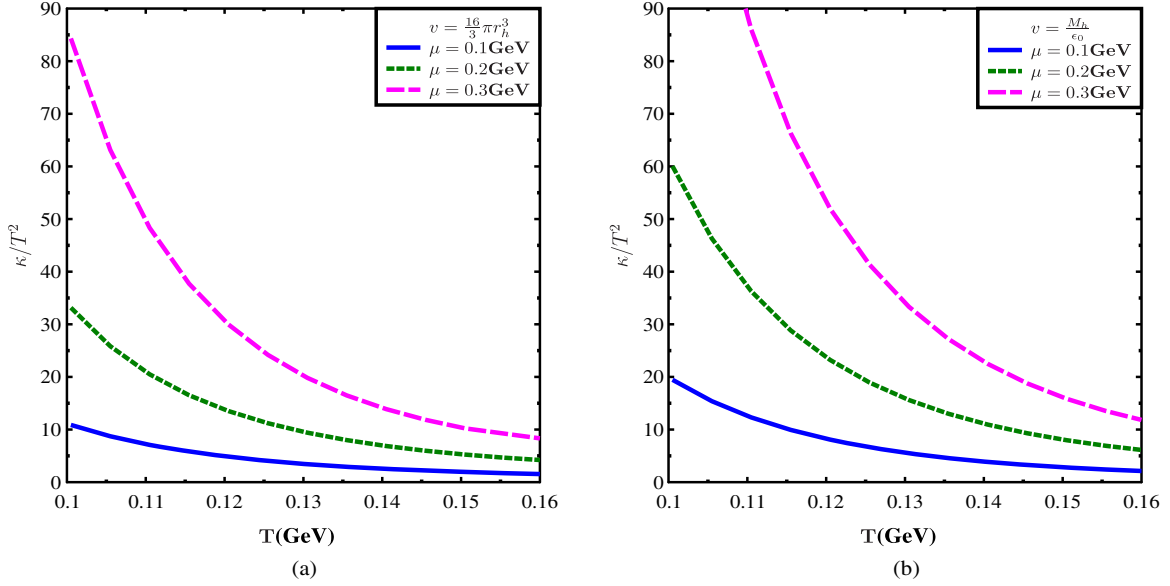
FIG. 4.   Scaled thermal conductivity as a function of temperature in the EHRGM with uniform excluded volume for all hadrons (a) and a mass dependent excluded volume (b).

temperature. This behavior may be the reflection of the fact that, while the cross section in the case of the uniform hard-core excluded volume parameter is ∼10 mb, the same in mass dependent parametrization varies from 3 mb for pions to 10 mb for protons. Thus, the smaller cross section leads to a larger relaxation time and therefore results in a larger electrical conductivity.

In Fig. 4 we show the variation of thermal conductivity with temperature for $\mu = 0.1$, 0.2, and 0.3 GeV. We note that the thermal conductivity decreases with increase in temperature. Further, at given a temperature, κ/T^2 is always larger for higher chemical potential. The coefficient of thermal conductivity depends on three factors, viz., the relaxation time $\tilde{\tau}$, the distribution function f_0, and the quantity w/n [see Eq. (34)]. Although the relaxation time decreases with μ, f_0 and w/n increase with increase in chemical potential. It turns out that the latter wins over the former and the overall effect is to increase κ with μ. We further note that the magnitude of κ/T^2 in the uniform excluded volume scheme is smaller than that of the mass dependent one. This observation can again be attributed to the fact that the cross section in the former parametrization is relatively larger than that of the latter.

In order to make the connection with the heavy ion collision experiments, we need the beam energy dependence ($\sqrt{s}$) of the electrical and thermal conductivities. This is extracted from a statistical thermal model description of the particle yield at various $\sqrt{s}$ [126]. $T(\mu)$ is parametrized by $T(\mu) = a - b\mu^2 - c\mu^4$, with $a = 0.166 \pm 0.002$ GeV, $b = 0.139 \pm 0.016$ GeV^{-1}, and $c = 0.053 \pm 0.021$ GeV^{-3}. The energy dependence of the baryon chemical potential is parametrized as $\mu = d/(1 + e\sqrt{s})$, where, $d = 1.308 \pm 0.028$ GeV and $e = 0.273 \pm 0.008$ GeV^{-1} [126]. In Fig. 5(a), we have

shown the variation of electrical (σ_{ele}/T) conductivity with the center of mass energy ($\sqrt{s}$). We note that the electrical conductivity first decreases along the freeze-out line with increasing collision energy and then attains almost constant value at large $\sqrt{s}$ for both the uniform excluded volume parameter ($v = \frac{16}{3}\pi r_h^3$) and mass dependent volume parameter case ($v = \frac{M_h}{\epsilon_0}$). This is reasonable because low $\sqrt{s}$ corresponds to low temperature and high chemical potential along the freeze-out curve, at which electrical conductivity is larger. We can conclude that along the freeze-out line electrical conductivity of the hadron gas does not change.

In Fig. 5(b) we have shown the variation of thermal conductivity (κ/T^2) with $\sqrt{s}$. We observe that thermal conductivity first decreases at small values of $\sqrt{s}$, then decreases slowly, becomes minimum, and finally increases at its larger values for both the uniform excluded volume parameter and mass dependent excluded volume parameter case. This can be understood from the expression for thermal conductivity as given in Eq. (34). In order to discuss Fig. 5(b), we can approximate μ, T to be much smaller than the masses of the baryons. In that case, we have $\omega = \epsilon + p \simeq n(m + T)$, where the baryon number density is given by

$$n = 2g\left(\frac{mT}{2\pi}\right)^{3/2} e^{-\beta m} \sinh(\beta\mu), \qquad (45)$$

so that the factor $(E_a - \frac{\omega}{n})^2$ in Eq. (34) becomes $\simeq (E_a - \frac{(m+T)T}{2\mu})^2$. Therefore, as μ increases, the second term in the parentheses decreases, leading to an increase of the thermal conductivity as seen in Figs. 4 and 5.
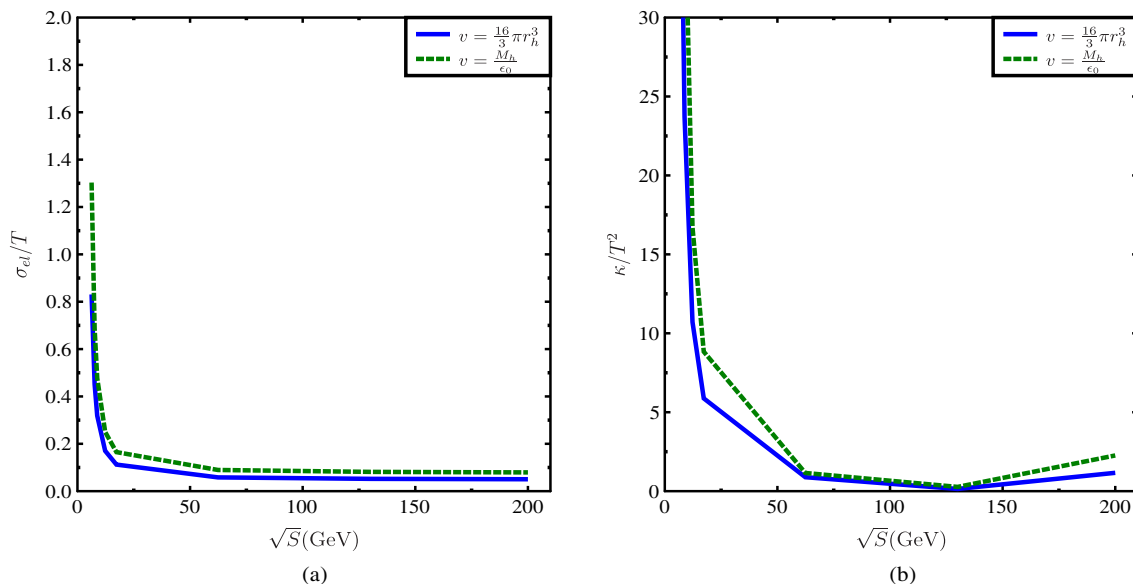
FIG. 5.   Variation of scaled electrical conductivity ($\frac{\sigma_{\mathrm{ele}}}{T}$) (a) and scaled thermal conductivity ($\frac{\kappa}{T^2}$) (b) with the center of mass energy for both uniform excluded volume parameter (blue solid line) and mass dependent volume parameter (green dashed line).

However, as μ becomes vanishingly small, the second term dominates over the first term and diverges for $\mu = 0$. Therefore, κ/T^2 will show a minimum as a function of $\sqrt{s}$ as seen in Fig. 5(b). Similar to the electrical conductivity, the value of κ/T^2 is also more with the mass dependent excluded volume parameter case as compared to the case of the uniform excluded volume parameter for all the values of $\sqrt{s}$ except for its small values.

VI. SUMMARY

We have studied the electrical and thermal conductivity of hot and dense hadron gas by using the Boltzmann equation in the relaxation time approximation. First we have estimated the relaxation time for all the hadrons by assuming the constant cross section. Here we have used the hadron resonance gas model, where the repulsive interactions are parametrized through excluded volume corrections in the ideal hadron resonance gas. We choose the uniform excluded volume and mass dependent excluded volume parametrization scheme. We have included all the hadrons and their resonances with mass cutoff 2.25 GeV. Here we take $r_h = 0.3$ fm for hadrons. We have compared our results for both the mass dependent excluded volume parameter and the uniform excluded volume parameter case with various existing results. We found that the magnitude of electrical conductivity is higher in our case as compared

to these existing results and is more for the mass dependent excluded volume parameter case as compared to the case of the uniform excluded volume parameter. We have shown the behavior of the electrical and thermal conductivity with temperature for different values of the chemical potential. We found that electrical and thermal conductivity increases with increase in chemical potential. The increase in electrical and thermal conductivity is more for the mass dependent excluded volume parameter case as compared to the case of the uniform excluded volume parameter.

Further, we have shown the variation of electrical (σ_{ele}/T) and thermal (κ/T^2) conductivity with the collision energy ($\sqrt{s}$). We found that electrical conductivity first decreases at small values of $\sqrt{s}$ and then remains almost constant at its larger value for both the case of the uniform excluded volume parameter and the mass dependent excluded volume parameter. Thus, we can conclude that electric conductive behavior of hadrons remains same along the freeze-out line. On the other hand, thermal conductivity first decreases with $\sqrt{s}$, attains minimum, and then increases very slowly.

ACKNOWLEDGMENTS

[1] C. Gale, S. Jeon, and B. Schenke, Int. J. Mod. Phys. A **28,** 1340011 (2013).

[2] B. Schenke, J. Phys. G **38,** 124009 (2011).

[3] C. Shen and U. Heinz, Phys. Rev. C **85,** 054902 (2012).

[4] P. F. Kolb and U. W. Heinz, in *Quark Gluon Plasma*, edited by R. C. Hwa and X.-N. Wang (Wold scientific, River edge, New Jersey, 2004), pp. 634–714.

[5] D. Teaney, J. Lauret, and E. V. Shuryak, Phys. Rev. Lett. **86,** 4783 (2001).

[6] L. Del Zanna, V. Chandra, G. Inghirami, V. Rolando, A. Beraudo, A. De Pace, G. Pagliara, A. Drago, and F. Becattini, Eur. Phys. J. C **73,** 2524 (2013).

[7] I. Karpenko, P. Huovinen, and M. Bleicher, Comput. Phys. Commun. **185,** 3016 (2014).

[8] H. Holopainen, H. Niemi, and K. J. Eskola, J. Phys. G **38,** 124164 (2011).

[9] A. Jaiswal, B. Friman, and K. Redlich, Phys. Lett. B **751,** 548 (2015).

[10] Z. Xu and C. Greiner, Phys. Rev. C **71,** 064901 (2005).

[11] I. Bouras, E. Molnar, H. Niemi, Z. Xu, A. El, O. Fochler, C. Greiner, and D. H. Rischke, Phys. Rev. C **82,** 024910 (2010).

[12] I. Bouras, A. El, O. Fochler, H. Niemi, Z. Xu, and C. Greiner, Phys. Lett. B **710,** 641 (2012).

[13] O. Fochler, Z. Xu, and C. Greiner, Phys. Rev. C **82,** 024907 (2010).

[14] C. Wesp, A. El, F. Reining, Z. Xu, I. Bouras, and C. Greiner, Phys. Rev. C **84,** 054911 (2011).

[15] J. Uphoff, O. Fochler, Z. Xu, and C. Greiner, Phys. Lett. B **717,** 430 (2012).

[16] M. Greif, F. Reining, I. Bouras, G. S. Denicol, Z. Xu, and C. Greiner, Phys. Rev. E **87,** 033019 (2013).

[17] P. Danielewicz and M. Gyulassy, Phys. Rev. D **31,** 53 (1985).

[18] M. Gyulassy and L. McLerran, Nucl. Phys. **A750,** 30 (2005).

[19] L. P. Csernai, J. I. Kapusta, and L. D. McLerran, Phys. Rev. Lett. **97,** 152303 (2006).

[20] P. Kovtun, D. T. Son, and A. O. Starinets, Phys. Rev. Lett. **94,** 111601 (2005).

[21] S. Gavin, Nucl. Phys. **A435,** 826 (1985).

[22] M. Prakash, M. Prakash, R. Venugopalan, and G. Welke, Phys. Rep. **227,** 321 (1993).

[23] A. Dobado and F. J. Llanes-Estrada, Phys. Rev. D **69,** 116004 (2004).

[24] J. W. Chen, Y. H. Li, Y. F. Liu, and E. Nakano, Phys. Rev. D **76,** 114011 (2007).

[25] K. Itakura, O. Morimatsu, and H. Otomo, Phys. Rev. D **77,** 014014 (2008).

[26] A. Dobado, F. J. Llanes-Estrada, and J. M. Torres-Rincon, Phys. Rev. D **80,** 114015 (2009).

[27] N. Demir and S. A. Bass, Phys. Rev. Lett. **102,** 172302 (2009).

[28] A. Puglisi, S. Plumari, and V. Greco, Phys. Lett. B **751,** 326 (2015).

[29] L. Thakur, P. K. Srivastava, G. P. Kadam, M. George, and H. Mishra, Phys. Rev. D **95,** 096009 (2017).

[30] K. Rajagopal and N. Tripuraneni, J. High Energy Phys. 03 (2010) 018.

[31] J. R. Bhatt, H. Mishra, and V. Sreekanth, Phys. Lett. B **704,** 486 (2011); Nucl. Phys. **A875,** 181 (2012); Phys. Rev. C **80,** 054906 (2009).

[32] A. Monnai and T. Hirano, Phys. Rev. C **80,** 054906 (2009).

[33] G. S. Denicol, T. Kodama, T. Koide, and P. Mota, Phys. Rev. C **80,** 064901 (2009).

[34] K. Dusling and T. Schfer, Phys. Rev. C **85,** 044909 (2012).

[35] H. Song and U. W. Heinz, Phys. Rev. C **81,** 024905 (2010).

[36] J. Noronha-Hostler, G. S. Denicol, J. Noronha, R. P. G. Andrade, and F. Grassi, Phys. Rev. C **88,** 044916 (2013).

[37] J. Noronha-Hostler, J. Noronha, and F. Grassi, Phys. Rev. C **90,** 034907 (2014).

[38] A. Dobado and S. N. Santalla, Phys. Rev. D **65,** 096011 (2002).

[39] D. Davesne, Phys. Rev. C **53,** 3069 (1996).

[40] D. Kharzeev and K. Tuchin, J. High Energy Phys. 09 (2008) 093.

[41] D. Fernandez-Fraile and A. Gomez Nicola, Eur. Phys. J. C **62,** 37 (2009).

[42] J. W. Chen and J. Wang, Phys. Rev. C **79,** 044913 (2009).

[43] J. Noronha-Hostler, J. Noronha, and C. Greiner, Phys. Rev. Lett. **103,** 172302 (2009); Phys. Rev. C **86,** 024913 (2012).

[44] C. Sasaki and K. Redlich, Nucl. Phys. **A832,** 62 (2010); Phys. Rev. C **79,** 055207 (2009).

[45] D. Fernandez-Fraile and A. Gomez Nicola, Phys. Rev. Lett. **102,** 121601 (2009).

[46] A. Dobado and J. M. Torres-Rincon, Phys. Rev. D **86,** 074021 (2012).

[47] V. Ozvenchuk, O. Linnyk, M. I. Gorenstein, E. L. Bratkovskaya, and W. Cassing, Phys. Rev. C **87,** 064903 (2013).

[48] U. Gangopadhyaya, S. Ghosh, S. Sarkar, and S. Mitra, Phys. Rev. C **94,** 044914 (2016).

[49] P. Chakraborty and J. I. Kapusta, Phys. Rev. C **83,** 014906 (2011).

[50] C. Sasaki and K. Redlich, Phys. Rev. C **79,** 055207 (2009).

[51] K. Tuchin, Adv. High Energy Phys. **2013,** 490495 (2013).

[52] V. Voronyuk, V. D. Toneev, W. Cassing, E. L. Bratkovskaya, V. P. Konchakovski, and S. A. Voloshin, Phys. Rev. C **83,** 054911 (2011).

[53] D. E. Kharzeev, L. D. Mclerran, and H. J. Warringa, Nucl. Phys. **A803,** 227 (2008); M. Asakawa, A. Mujumder, and B. Mueller, Phys. Rev. C **81,** 064912 (2010).

[54] K. Tuchin, Phys. Rev. C **83,** 017901 (2011); K. Marasinghe and K. Tuchin, Phys. Rev. C **84,** 044908 (2011).

[55] A. Das, S. S. Dave, P. S. Saumia, and A. M. Srivatava, Phys. Rev. C **96,** 034902 (2017).

[56] P. Sushruth Reddy, C. S. Amal Jahan, N. Dhale, A. Mishra, and J. Schaeffener-Bielich, Phys. Rev. C **97,** 065208 (2018).

[57] Y. Yin, Phys. Rev. C **90,** 044903 (2014).

[58] Y. Akamatsu, H. Hamagaki, T. Hatsuda, and T. Hirano, J. Phys. G **38,** 124184 (2011).

[59] K. Fukushima, D. E. Kharzeev, and H. J. Warringa, Phys. Rev. D **78,** 074033 (2008).

[60] M. Greif, I. Bouras, C. Greiner, and Z. Xu, Phys. Rev. D **90,** 094014 (2014).

[61] A. Puglisi, S. Plumari, and V. Greco, Phys. Rev. D **90,** 114009 (2014).

[62] S. I. Finazzo and J. Noronha, Phys. Rev. D **89,** 106008 (2014).

[63] W. Cassing, O. Linnyk, T. Steinert, and V. Ozvenchuk, Phys. Rev. Lett. **110**, 182301 (2013).

[64] T. Steinert and W. Cassing, Phys. Rev. C **89**, 035203 (2014).

[65] S. x. Qin, Phys. Lett. B **742**, 358 (2015).

[66] H. Berrehrah, E. Bratkovskaya, W. Cassing, and R. Marty, J. Phys. Conf. Ser. **612**, 012050 (2015).

[67] R. Marty, E. Bratkovskaya, W. Cassing, J. Aichelin, and H. Berrehrah, Phys. Rev. C **88**, 045204 (2013).

[68] P. K. Srivastava, L. Thakur, and B. K. Patra, Phys. Rev. C **91**, 044903 (2015).

[69] S. Mitra and V. Chandra, Phys. Rev. D **94**, 034025 (2016).

[70] S. Gupta, Phys. Lett. B **597**, 57 (2004).

[71] G. Aarts, C. Allton, J. Foley, S. Hands, and S. Kim, Phys. Rev. Lett. **99**, 022002 (2007).

[72] P. V. Buividovich, M. N. Chernodub, D. E. Kharzeev, T. Kalaydzhyan, E. V. Luschevskaya, and M. I. Polikarpov, Phys. Rev. Lett. **105**, 132001 (2010).

[73] H.-T. Ding, A. Francis, O. Kaczmarek, F. Karsch, E. Laermann, and W. Soeldner, Phys. Rev. D **83**, 034504 (2011).

[74] Y. Burnier and M. Laine, Eur. Phys. J. C **72**, 1902 (2012).

[75] B. B. Brandt, A. Francis, H. B. Meyer, and H. Wittig, J. High Energy Phys. 03 (2013) 100.

[76] A. Amato, G. Aarts, C. Allton, P. Giudice, S. Hands, and J. I. Skullerud, Phys. Rev. Lett. **111**, 172001 (2013).

[77] D. Fernández-Fraile and A. Gomez Nicola, Phys. Rev. D **73**, 045025 (2006).

[78] S. Samanta, S. Ghosh, and B. Mohanty, J. Phys. G **45**, 075101 (2018).

[79] M. Greif, C. Greiner, and G. S. Denicol, Phys. Rev. D **93**, 096012 (2016).

[80] S. Ghosh, Phys. Rev. D **95**, 036018 (2017).

[81] K. Saha, S. Ghosh, S. Upadhaya, and S. Maity, Phys. Rev. D **97**, 116020 (2018).

[82] P. Singha, A. Abhishek, G. Kadam, S. Ghosh, and H. Mishra, arXiv:1705.03084.

[83] G. S. Denicol, H. Niemi, I. Bouras, E. Molnar, Z. Xu, D. H. Rischke, and C. Greiner, Phys. Rev. D **89**, 074005 (2014).

[84] J. I. Kapusta and J. M. Torres-Rincon, Phys. Rev. C **86**, 054911 (2012).

[85] S. Mitra and S. Sarkar, Phys. Rev. D **89**, 054013 (2014).

[86] S. Ghosh, Int. J. Mod. Phys. E **24**, 1550058 (2015).

[87] A. Dobado, F. J. Llanes-Estrada, and J. M. Torres Rincon, arXiv:hep-ph/0702130.

[88] D. Fernandez-Fraile and A. Gomez Nicola, Int. J. Mod. Phys. E **16**, 3010 (2007).

[89] Sourav Sarkar, Adv. High Energy Phys. **2013**, 627137 (2013).

[90] P. Deb, G. P. Kadam, and H. Mishra, Phys. Rev. D **94**, 094002 (2016).

[91] A. Abhishek, H. Mishra, and S. Ghosh, Phys. Rev. D **97**, 014005 (2018).

[92] P. Braun-Munzinger, J. Stachel, J. P. Wessels, and N. Xu, Phys. Lett. B **365**, 1 (1996).

[93] G. D. Yen and M. I. Gorenstein, Phys. Rev. C **59**, 2788 (1999).

[94] F. Becattini, J. Cleymans, A. Keranen, E. Suhonen, and K. Redlich, Phys. Rev. C **64**, 024901 (2001).

[95] R. Dashen, S. K. Ma, and H. J. Bernstein, Phys. Rev. **187**, 345 (1969).

[96] M. I. Gorenstein, M. Hauer, and O. N. Moroz, Phys. Rev. C **77**, 024911 (2008).

[97] J. Noronha-Hostler, J. Noronha, and C. Greiner, Phys. Rev. C **86**, 024913 (2012).

[98] S. K. Tiwari, P. K. Srivastava, and C. P. Singh, Phys. Rev. C **85**, 014908 (2012).

[99] G. P. Kadam and H. Mishra, Phys. Rev. C **92**, 035203 (2015).

[100] A. Bhattacharyya, R. Ray, S. Samanta, and S. Sur, Phys. Rev. C **91**, 041901 (2015).

[101] P. Garg, D. K. Mishra, P. K. Netrakanti, B. Mohanty, A. K. Mohanty, B. K. Singh, and N. Xu, Phys. Lett. B **726**, 691 (2013).

[102] R. Hagedorn and J. Rafelski, Phys. Lett. **97B**, 136 (1980).

[103] R. Hagedorn, Z. Phys. C **17**, 265 (1983).

[104] M. I. Gorenstein, V. K. Petrov, and G. M. Zinovev, Phys. Lett. **106B**, 327 (1981).

[105] M. Albright, J. Kapusta, and C. Young, Phys. Rev. C **90**, 024915 (2014).

[106] J. I. Kapusta and K. A. Olive, Nucl. Phys. **A408**, 478 (1983).

[107] D. H. Rischke, M. I. Gorenstein, H. Stoecker, and W. Greiner, Z. Phys. C **51**, 485 (1991).

[108] G. D. Yen, M. I. Gorenstein, W. Greiner, and S. N. Yang, Phys. Rev. C **56**, 2210 (1997).

[109] K. Yagi, T. Hatsuda, and Y. Miake, *Quark-Gluon Plasma: From Big Bang to Little Bang* (Cambridge University Press, Cambridge, 2005).

[110] C. Crecignani and G. M. Kremer, *The Relativistic Boltzmann Equation: Theory and Applications* (Birkhiiuser, Boston, 2002).

[111] W. Israel and J. M. Stewart, Ann. Phys. (N.Y.) **118**, 341 (1979).

[112] S. R. de Groot, W. A. van Leewen, and C. G. van Weert, *Relativistic Kinetic Theory: Principles and Applications* (Elsevier North-Holland, Amsterdam, 1980).

[113] G. S. Denicol, H. Niemi, E. Molnar, and D. H. Rischke, Phys. Rev. D **85**, 114047 (2012).

[114] A. Hosoya and K. Kajantie, Nucl. Phys. **B250**, 666 (1985).

[115] O. Moroz, Ukr. J. Phys. **58**, 1127 (2013).

[116] M. Cannoni, Phys. Rev. D **89**, 103533 (2014).

[117] P. Gondolo and G. Gelmini, Nucl. Phys. **B360**, 145 (1991).

[118] C. Amsler *et al.* (Particle Data Group), Phys. Lett. B **667**, 1 (2008).

[119] A. Bohr and B. Mottelson, *Nuclear Structure* (Benjamin, New York, 1969), Vol. 1, p. 266.

[120] V. Vovchenko, D. V. Anchishkin, and M. I. Gorenstein, Phys. Rev. C **91**, 024905 (2015).

[121] G. Kadam and S. Pawar, arXiv:1802.01942.

[122] S. Caron-Hout, P. Kovtun, G. D. Moore, A. Starinets, and L. G. Yaffe, J. High Energy Phys. 06 (2006) 015.

[123] G. Aarts, C. Allton, A. Amato, P. Giudice, S. Hands, and J.-I. Skullerud, J. High Energy Phys. 02 (2015) 186.

[124] R. Rougemont, J. Noronha, and J. Noronha-Hostler, Phys. Rev. Lett. **115**, 202301 (2015).

[125] V. Vovchenko, M. I. Gorenstein, and H. Stoecker, Phys. Rev. Lett. **118**, 182301 (2017).

[126] J. Cleymans, H. Oeschler, K. Redlich, and S. Wheaton, Phys. Rev. C **73**, 034905 (2006).