Beyond-proximity-force-approximation Casimir force between two spheres at finite temperature

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A recent experiment [J. L. Garrett, D. A. T. Somers, and J. N. Munday, Phys. Rev. Lett. **120**, 040401 (2018)] measured for the first time the gradient of the Casimir force between two gold spheres at room temperature. The theoretical analysis of the data was carried out using the standard proximity force approximation (PFA). A fit of the data, using a parametrization of the force valid for the sphere-plate geometry, was used by the authors to place a bound on deviations from PFA. Motivated by this work, we compute the Casimir force between two gold spheres at finite temperature. The semianalytic formula for the Casimir force that we construct is valid for all separations, and can be easily used to interpret future experiments in both the sphere-plate and sphere-sphere configurations. We describe the correct parametrization of the corrections to PFA for two spheres that should be used in data analysis.

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I. INTRODUCTION

The Casimir force [1] is the tiny long-range force between (neutral) macroscopic polarizable bodies that originates from the modification of the spectrum of quantum and thermal vacuum fluctuations of the electromagnetic (em) field, caused by the presence of the bodies. This phenomenon represents one of the rare manifestations of the quantum properties of the em field at the macroscopic scale. For reviews see Refs. [2–6].

A characteristic feature of the Casimir force is its *nonadditivity*, reflecting the many-body character of fluctuation forces. This property enormously complicates the computation the Casimir force in nonplanar geometries. We recall that in his original paper Casimir worked out the force in the highly idealized geometry of two perfectly conducting plane-parallel surfaces at zero temperature. Planar systems were the exclusive object of consideration also in the famous paper by Lifshitz [7], who derived a general formula for the Casimir interaction between two plane-parallel slabs, taking into full account realistic material properties of the plates, i.e., their frequency-dependent dielectric permittivity and finite temperature.

Unfortunately, the theoretically simple planar geometry studied by Casimir and Lifshitz has been rarely used in experiments [8,9], because of severe difficulties connected with controlling the parallelism of two macroscopic surfaces separated by a submicron gap. To avoid these problems, the vast majority of Casimir experiments adopt the sphere-plate geometry (see for example [10–20] and references therein), which is obviously immune from parallelism issues.

Very recently, a new experiment [21] has measured for the first time the (gradient of the) Casimir force between two gold-coated spheres. Following the practice of previous experiments, also in this new experiment the Casimir force has been computed using the simple proximity force approximation (PFA) [22], which expresses the Casimir force between two curved surfaces as the average of the plane-parallel force (as given by Lifshitz formula) over the local separation between the surfaces. The results of the experiment have been found to be in good agreement with theoretical predictions based on Lifshitz formula. By using the measurements made with nine sphere-sphere and three sphere-plate systems of different radii, the authors of [21] could also place a bound on the magnitude of deviations from PFA, using the same parametrization of the force that was used in the sphere-plate experiment of the IUPUI group [23].

Motivated by the new experiment, in this paper we compute beyond-PFA curvature corrections to the Casimir force gradient for two gold spheres at room temperature, and we describe the correct parametrization of the force that should be used in the data analysis to measure corrections to PFA in this geometry. The Casimir interaction of two metallic spheres in the limit of large separations was investigated in [24]. We remind the reader that the computation of the Casimir force for nonplanar geometries has been an intractable problem until recently. Only in the early 2000s, an exact scattering formula for the Casimir interaction between dielectric objects of any shape, generalizing early results of Balian and Duplantier [25] and Langbein [26], was worked

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out [27-29]. At finite temperature T, the scattering formula has the form of a sum over so-called Matsubara (imaginary) frequencies $\xi_n = 2\pi n k_B T/\hbar$, n = 0, 1, ... (with k_B Boltzmann constant, and \hbar Planck constant) of functional determinants involving the multipole expansions of the T-operators of the two bodies. Unfortunately, the scattering formula converges very slowly in the characteristic regime of Casimir experiments, in which the (minimum) separation a between the two surfaces is very small compared to their characteristic radius of curvature R. In order to obtain a precise estimate of the Casimir force, as it is needed for a proper interpretation of current precision experiments, it is necessary to push the computation to very high multipole orders, which represents a very challenging task even for present day computers. In [30] simulations of the sphere-plate problem for Drude conductors were done up to multipole order $l_{\text{max}} = 24$, which were later [31] pushed to $l_{\text{max}} = 45$, allowing one to calculate the force for $a/R \ge 0.1$. Very recently [32], a large-scale numerical simulation of the sphere-plate system going up to mutipole order $l_{\text{max}} = 2 \times 10^4$ reached for the first time the experimentally important region $a/R \sim 10^{-3}$.

Large numerical simulations of the scattering formula like that of [32] require sophisticated algorithms for the computation of determinants of hierarchical matrices, which not every researcher may master or be willing to spend time on. This led us to investigate whether the new analytical tools that have been developed recently in the Casimir field could be exploited to construct easy-to-use formulas for the Casimir force having the high degree of precision demanded by current experiments. In Ref. [33] a formula with these features was constructed for the sphere-plate system.

The approach followed in [33] can be described as follows. As we said earlier, the scattering formula has the form of a sum over Matsubara modes ξ_n . The first term of this series, corresponding to n = 0, represents a classical contribution to the Casimir interaction, which becomes dominant in the high temperature limit $a/\lambda_T \gg 1$, where $\lambda_T = \hbar c/2\pi k_B T$ $(\lambda_T = 1.2 \ \mu m$ for room temperature) is the thermal length. In Ref. [34], it was shown that this classical term can be evaluated *exactly* in the geometry of two metallic spheres of any radii, including the sphere-plate case as a special limit. Of course, knowledge of the n = 0 term is not sufficient to compute the full Casimir energy. Unfortunately, the n > 0terms of the scattering formula cannot be computed exactly, but in [33] it was shown that they can be computed very precisely using an asymptotic small-distance formula, which includes corrections to PFA, based on a recently proposed derivative expansion (DE) of the Casimir interaction [35–39]. In [33] it was proved that the semianalytic approximate formula for the sphere-plate Casimir force, resulting from the combination of the exact n = 0 term with the approximate expression of the n > 0 terms, is indeed extremely precise for all separations. The formula derived in [33] is in excellent agreement with the results of the large numerical simulation of [32]. In this paper, we extend the construction of [33] to the sphere-sphere system.

The paper is organized as follows: in Sec. II we review the PFA for two spheres, the scattering formula, and display the exact solution for the classical Casimir energy of two Drude spheres discovered in [34]. In Sec. III we compute the contribution of the positive Matsubara modes using the DE. In Sec. IV we display our complete formula for the Casimir energy of two spheres at finite temperature, and use it to compute deviations from PFA. In Sec. IV we also review the recent two-sphere experiment [21] and describe the correct parametrization of corrections to PFA that should be used in the data analysis of experiments with two spheres. In Sec. V we present our conclusions. Finally, in Appendix A.1 we review the DE and in Appendix A.2 we use the DE to compute the leading curvature correction to the force gradient between two spheres.

II. CASIMIR INTERACTION OF TWO SPHERES

We consider a system composed by two spheres of respective radii R_1 and R_2 placed in a vacuum and separated by a gap of width *a* (see Fig. 1).



FIG. 1. The sphere-sphere Casimir setup. The sphere-sphere geometry is characterized by the effective radius $\tilde{R} = R_1 R_2 / (R_1 + R_2)$ and the dimensionless parameter $u = \tilde{R}^2 / R_1 R_2$.

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As it was explained in the introduction, until recently there were no tools to exactly compute the Casimir force in nonplanar geometries, and so one had to resort to the oldfashioned PFA. In the case of two spheres [3] the PFA formula for the Casimir force is

$$F^{(\text{PFA})}(a, R_1, R_2) = 2\pi R \mathcal{F}^{(\text{pp})}(a), \qquad (1)$$

where \tilde{R} is the effective radius of the two spheres,

$$\tilde{R} = \frac{R_1 R_2}{R_1 + R_2},\tag{2}$$

and $\mathcal{F}^{(\text{pp})}(a)$ is the unit-area Casimir free energy for two plane-parallel slabs respectively made of the same materials as the sphere and plate, whose expression was derived long ago by Lifshitz [7],

$$\mathcal{F}^{(\mathrm{pp})}(a,T) = \frac{k_B T}{2\pi} \sum_{n\geq 0}^{\prime} \int_0^\infty k_\perp dk_\perp$$
$$\times \sum_{\alpha=\mathrm{TE},\mathrm{TM}} \ln\left[1 - r_\alpha^2(\mathrm{i}\xi_n, k_\perp)e^{-2aq_n}\right], \quad (3)$$

where the prime in the sum over *n* indicates that the n = 0 term is taken with a weight 1/2, *T* is the temperature of the plates, k_{\perp} is the in-plane momentum, $r_{\alpha}(i\xi_n, k_{\perp})$ denotes the Fresnel reflection coefficient for polarization $\alpha =$ TE, TM of a thick slab, evaluated for the imaginary frequency $\omega = i\xi_n$ and $q_n = \sqrt{\xi_n^2/c^2 + k_{\perp}^2}$. The PFA for the force gradient $F' \equiv \partial F/\partial a$ (here and in what follows a prime denotes a derivative with respect to the separation) easily follows from Eq. (1),

$$F'^{(\text{PFA})}(a, R_1, R_2) = -2\pi \tilde{R} F^{(\text{pp})}(a), \qquad (4)$$

where $F^{(\text{pp})}(a) = -\partial \mathcal{F}^{(\text{pp})}/\partial a$ is the unit-area Casimir force for two parallel slabs. The PFA force and forcegradient for a sphere of radius *R* opposite a plane is recovered from Eqs. (1) and (4), respectively, by taking the radius of one of the two spheres to infinity, i.e., substituting \tilde{R} by *R*. It is important to remark that within the PFA, both *F* and *F'* depend on the radii of the two spheres only via the effective radius \tilde{R} .

For a proper interpretation of current precision Casimir experiments it has become important to estimate curvature corrections beyond PFA. This has been impossible until recently, when an exact scattering formula providing the Casimir energy of two compact dielectric bodies was worked out [27–29]. The general structure of the scattering formula is

$$\mathcal{F} = k_B T \sum_{n \ge 0}^{\prime} \operatorname{Tr} \ln[1 - \hat{M}(\mathrm{i}\xi_n)], \qquad (5)$$

where the prime sign in the sum indicates again that the n = 0 term is taken with weight 1/2. The trace Tr in this

equation is over both spherical multipoles indices (l, m)and polarization indices $\alpha = TE, TM$,

$$Tr = \sum_{m=-\infty}^{\infty} \sum_{l=|m|}^{\infty} tr,$$
(6)

where tr denotes the trace over α . The matrix elements $M_{lm\alpha,m'l'\alpha'}$ of \hat{M} are not reported here for brevity. Their explicit expressions can be found for example in Refs. [40,41]. We just recall that the matrix $M_{lm\alpha m'l'\alpha'}$ involves a product of the T-matrices for the two bodies (i.e., the Mie scattering coefficients in the case of two spheres), both evaluated for the imaginary Matsubara frequencies $i\xi_n$, intertwined with translation matrices that serve to convert the multipole basis relative to either body into the multipole basis relative to the other body (see Refs. [40,41] for details). The expressions for the Casimir force $F = -\mathcal{F}'$ and its gradient F' are obtained by taking derivatives of Eq. (5) with respect to the separation a. Using the scattering formula it has been possible to prove eventually that the PFA formula is indeed asymptotically exact for small separation in the sphereplate and cylinder-plate geometries [42].

Having at our disposal the exact representation Eq. (5), it is natural to ask whether it can be used efficiently to accurately compute the Casimir force in concrete experimental situations. Unfortunately, this is not easy at all. Consider as an example the geometry of a sphere of radius *R* at a minimum distance *a* from a plate. The problem is that to obtain a precise estimate of the Casimir force for experimentally relevant sphere-plate separations (typical aspect ratios $a/R \sim 10^{-3}$) it is necessary to include a huge number of multipoles in the computation. Previous works [28,30–33] found that the multipole order l_{max} for which convergence is achieved scales as $l_{\text{max}} \sim R/a$. To date, the largest numerical simulation of the sphere-plate scattering formula reached up to $l_{\text{max}} = 2 \times 10^4$ [32], which allowed the authors of [32] to probe the Casimir force in the experimentally relevant region $a/R \sim 10^{-3}$. Managing such a large number of multipoles on a computer is not easy at all, and sophisticated algorithms are needed to handle the problem.

At this point we turn to our objective of deriving a simple and very accurate formula for the Casimir force between two spheres. To do this, we go back to the general scattering formula Eq. (5). As we see, it has the form of a sum of terms $\mathcal{F} = \sum_{n\geq 0}' \mathcal{F}_n$ over the Matsubara frequencies ξ_n , $n = 0, 1, \ldots$ It is convenient for our purposes to separate the first term $\mathcal{F}_{n=0}$ of the series from the remaining terms with n > 0. We accordingly decompose Eq. (5) as

$$\mathcal{F} = \mathcal{F}_{n=0} + \mathcal{F}_{n>0},\tag{7}$$

where we set $\mathcal{F}_{n>0} = \sum_{n>0} \mathcal{F}_n$. Consider first $\mathcal{F}_{n=0}$. This term represents a classical contribution to the Casimir

energy, which provides the dominant contribution to the full Casimir energy \mathcal{F} of the system in the limit of large separations $a \gg \lambda_T$.

In Ref. [34], it was shown that this classical term can be evaluated exactly in the following two cases. The first one is that of a scalar field obeying Dirichlet (D) boundary conditions (bc) on the surfaces of two spheres of arbitrary radii, including the sphere-plate geometry as a special case. The second case is that of a scalar field obeying so-called Drude bc on the surfaces of a sphere opposite a plate [43] The latter bc is identical to D bc, apart from the fact that in the Drude case charge monopoles (corresponding to index l = 0) are excluded from the scattering formula. Both sets of bc can be used to describe ohmic conductors, depending on the electric configuration of the system. Drude bc describe isolated conductors, whose total charge is fixed, while D bc describe conductors whose voltages are fixed. The latter is the experimentally important situation, since in all Casimir experiments (including the experiment [21]) one plate is grounded, while the other is connected with a voltage generator that serves to apply a bias potential, to compensate for unavoidable potential differences between the plates, resulting from differences in the respective work functions. The classical sphere-sphere Casimir energies implied by D and Drude bc are undistinguishable for separations much smaller than the spheres radii, while the two models lead to distinct asymptotic behaviors in the limit of large separations $a \gg (R_1, R_2)$, since $\mathcal{F}_{n=0}^{(D)} \sim -k_B T R_1 R_2 / a^2$, while $\mathcal{F}_{n=0}^{(Dr)} \sim -k_B T R_1^3 R_2^3 / a^6$. For two metallic spheres obeying D bc the exact solution

worked out in [34] is

$$\mathcal{F}_{n=0}^{(\text{ex})} = \frac{k_B T}{2} \sum_{l=0}^{\infty} (2l+1) \ln[1-Z^{2l+1}], \qquad (8)$$

where Z is

$$Z = [1 + x + x^2 u/2 + \sqrt{(x + x^2 u/2)(2 + x + x^2 u/2)}]^{-1}.$$
(9)

In the above equation, $x = a/\tilde{R}$, and $u = \tilde{R}^2/R_1R_2$. The parameter *u* depends only on the ratio between R_1 and R_2 and takes values in the interval [0, 1/4], the upper bound u = 1/4 corresponding to two equal spheres $R_1 = R_2$, while the lower bound u = 0 is approached as either of the two radii becomes infinite. The special case of a sphere of radius R opposite a plate is thus recovered by taking u = 0into Eq. (9), and setting $\tilde{R} = R$. The range of the variable Z in Eq. (9) is the interval [0, 1], the upper (lower) bound Z = 1 (Z = 0) corresponding to the limit of vanishing (infinite) separation $x \to 0$ ($x \to \infty$). The corresponding expression for the Casimir force $F_{n=0}^{(ex)} = -\mathcal{F}_{n=0}^{\prime(ex)}$ and its derivative $F_{n=0}^{\prime(ex)}$ are easily obtained by deriving Eq. (8) with respect to a. As we see from the expression of Z in Eq. (9),

the exact classical energy $\mathcal{F}_{n=0}^{(\mathrm{ex})}$ depends not only on the effective radius \tilde{R} , but also on the ratio among the radii via the variable *u*. This feature marks an important difference with respect to the PFA formula.

In [34] the small distance expansion of $\mathcal{F}_{n=0}^{(\mathrm{ex})}$ for the sphere-plate system was worked out, by setting Z = $\exp(-\mu)$ and then taking the small- μ asymptotic expansion of the series on the rhs of Eq. (8). Using the formulas of Ref. [34], it is easy to verify that for small separations the sphere-sphere force gradient has the expansion

$$F'_{n=0}^{(\text{ex})} = k_B T \frac{\zeta(3)\tilde{R}}{4a^3} \left(1 + \frac{1}{12\zeta(3)} \frac{a}{\tilde{R}} + o(a/\tilde{R}) \right), \quad (10)$$

where $\zeta(x)$ is the Riemann zeta function. The leading term coincides with the PFA Eq. (4), since from the Lifshitz formula we find $F_{n=0}^{(pp)} = -k_B T \zeta(3)/(8\pi a^3)$ for two Drude-metal plates, while the next term provides the leading correction to PFA. Interestingly, like the PFA, also the latter correction is independent of the parameter *u*. The correction to PFA in Eq. (10) is consistent with the DE [see Eq. (A17)].

At this point we need to consider the contribution $\mathcal{F}_{n>0}$ of the positive Matsubara modes n > 0 to the free energy. Unfortunately, differently from the classical term $\mathcal{F}_{n=0}$, the quantity $\mathcal{F}_{n>0}$ cannot be computed exactly. Of course $\mathcal{F}_{n>0}$ can be computed numerically, using the scattering formula truncated to a finite multipole order l_{max} . As we explained earlier, such a computation is however very challenging, because the multipole order l_{max} that is necessary is very large for experimentally relevant values of the radii and separation. Below we obtain a very precise and simple analytical formula for $\mathcal{F}_{n>0}$, by using the so-called derivative expansion [35-37,39].

III. DERIVATIVE EXPANSION OF $\mathcal{F}_{n>0}$

The DE [35–39] is an analytical technique to compute curvature corrections to proximity forces between two surfaces of small slope. For the benefit of the reader, we provide in Appendix A.1 a short review to the DE and a guide to the relevant references.

In this section we use the DE to estimate the contribution $\mathcal{F}_{n>0}$ of the nonzero Matsubara modes to the Casimir energy. The DE is particularly well suited to this task, as we now explain. By it very nature, the DE is expected to be very precise in situations in which the slope of the surfaces is small within the interaction region. A little reasoning shows that this condition is met in the problem at hand for all separations a between the spheres, just provided (as it always is the case in current experiments) that the radii of the spheres are both large compared to the thermal length λ_T ($\lambda_T = 1.2 \ \mu m$ at room temperature).

It is a well-known fact [3,4] that the Casimir interaction between two surfaces, with a characteristic radius of curvature R, is localized inside a disk of radius $\rho \sim \sqrt{aR}$

around the point of either surface which is closest to the other. Thus, one is led to expect that the DE is applicable in general only for separations a such that $\rho/R = \sqrt{a/R} \ll 1$. This condition is usually well satisfied in Casimir experiments, for which typically a/R < 0.01. A closer inspection reveals however that the DE of $\mathcal{F}_{n>0}$ is in fact valid for *all* separations, provided that R_1 and R_2 are both larger than λ_T . The point to consider is that positive Matsubara modes of (imaginary) frequency ξ_n can only propagate across a distance of order $\ell_n = c/\xi_n = \lambda_T/n \le \lambda_T$. Because of this constraint, the true size of the interaction region is actually not \sqrt{aR} , but instead $\rho = \min(\sqrt{aR}, \sqrt{\lambda_T R})$. This implies that the DE for $\mathcal{F}_{n>0}$ is actually valid for separations such that $\rho/R = \min(\sqrt{a/R}, \sqrt{\lambda_T/R}) \ll 1$. The latter condition is clearly satisfied for all separations, provided that R_1 and R_2 are both much larger than λ_T .

We have thus established that the DE is a valid method for $\mathcal{F}_{n>0}$. In Appendix A.2 it is shown that the DE leads to a simple general formula Eq. (A17) for the force gradient between two spheres. The formula for the DE expansion of $\mathcal{F}_{n>0}$ can obtained by making in Eq. (A17) the appropriate substitutions,

$$F'_{n>0} = -2\pi \tilde{R} F^{(\text{pp})}_{n>0}(a) \left[1 - (\tilde{\theta}(a) + u\kappa(a)) \frac{a}{\tilde{R}} \right], \quad (11)$$

where the coefficients $\tilde{\theta}(a)$ and $\kappa(a)$ are [see Eqs. (A18) and (A19)]

$$\tilde{\theta} = \frac{\mathcal{F}_{n>0}^{(\text{pp})}(a) - 2\alpha_{n>0}(a)}{aF_{n>0}^{(\text{pp})}(a)},$$
(12)

$$\kappa(a) = 1 - 2 \frac{\mathcal{F}_{n>0}^{(\text{pp})}(a)}{aF_{n>0}^{(\text{pp})}(a)}.$$
(13)

In the above equations, $\mathcal{F}_{n>0}^{(\mathrm{pp})}$ denotes the contribution of the n > 0 modes to Lifshitz formula Eq. (3), and $F_{n>0}^{(\mathrm{pp})} = -\partial \mathcal{F}_{n>0}^{(\mathrm{pp})}/\partial a$ is the corresponding (unit-area) force.

An important ingredient of Eqs. (11)–(13) is the coefficient $\alpha_{n>0}(a)$ which enters into the expression of $\tilde{\theta}$ in Eq. (12). As it is explained in Appendix A.1, this

coefficient can be extracted from the Green function $\tilde{G}^{(2)}(k; a)$ of the second-order perturbative expansion of $\mathcal{F}_{n>0}$ for a small amplitude deformation of one of the plates around the plane-parallel geometry. The computation of this coefficient for gold plates at room temperature can be carried out following the procedure described in [39], and we address the interested reader to that reference for details. The coefficients $\tilde{\theta}$ and κ for perfect conductors (PC) in the limit of zero temperature are both independent of the separation. Their values can be determined using the formulas listed in [36],

$$\tilde{\theta}_{T=0}^{(\rm PC)} = \frac{20}{3\pi^2} - \frac{1}{9} = 0.564, \qquad \kappa_{T=0}^{(\rm PC)} = \frac{1}{3}. \tag{14}$$

For gold surfaces at room temperature, both coefficients depend on the separation (as well as on the temperature and on the material lengths characterizing the optical properties of gold, in particular, the plasma length). In Table I we list the values of $\tilde{\theta}(a)$ and $\kappa(a)$ for gold at room temperature that we computed using tabulated [44] optical data [45], for several values of the separation in the range from 100 nm to 2 micron. Using these values of $\tilde{\theta}(a)$ and $\kappa(a)$ together with Eq. (A17), it is easily possible to compute $F'_{n>0}$ for any combinations of sphere radii.

In view of later applications, it is important to note that while in the proximity approximation $F'_{n>0}$ depends only on the effective radius \tilde{R} [see Eq. (4)], the more accurate expression of $F'_{n>0}$ in Eq. (11), which includes curvature corrections to PFA, depends also on the parameter *u*, i.e., on the ratio of the radii of the two spheres.

IV. BEYOND-PFA CORRECTIONS

Combining Eq. (8) with Eq. (11) we obtain the following formula for the force gradient between two gold spheres:

$$F' = F_{n=0}^{\prime(\text{ex})} - 2\pi \tilde{R} F_{n>0}^{(\text{pp})}(a) \left[1 - (\tilde{\theta}(a) + u\kappa(a)) \frac{a}{\tilde{R}} \right], \quad (15)$$

which constitutes the main result of the present work. A nice feature of the formula above is that, by construction, it is exact in both limits $a/\tilde{R} \rightarrow 0$ and $a/\lambda_T \rightarrow \infty$. This is so because, on one hand, Eq. (15) is exact for $a/\tilde{R} \rightarrow 0$, since

TABLE I. Values of the coefficients $\tilde{\theta}$ and κ for Au at room temperature.

$a(\mu m)$	0.10	0.15	0.2	0.25	0.3	0.35	0.4	0.45	0.5	0.55	0.6	0.65
<i>θ</i> к	0.456 0.245	0.4715 0.270	0.470 0.289	0.463 0.305	0.454 0.319	0.4445 0.331	0.435 0.342	0.425 0.353	0.415 0.362	0.4055 0.371	0.396 0.380	0.387 0.389
$\overline{a(\mu m)}$	0.70	0.75	0.8	0.85	0.9	0.95	1	1.2	1.4	1.6	1.8	2
θ κ	0.379 0.397	0.370 0.405	0.362 0.413	0.3545 0.421	0.347 0.429	0.3395 0.437	0.332 0.444	0.306 0.474	0.282 0.502	0.261 0.529	0.242 0.554	0.225 0.578

in this limit the DE, which we used to compute the contribution of the positive Matsubara modes, is asymptotically exact. On the other hand, Eq. (15) is exact also for $a/\lambda_T \rightarrow \infty$, because for separations larger than the thermal length $a \gg \lambda_T$ the relative contribution of the positive Matsubara modes vanishes exponentially fast, and then Eq. (15) reduces to the exact n = 0 mode. Comparison with high precision numerical computations of the sphere-plate scattering formula in [33] revealed that Eq. (15) is in fact very accurate also for all intermediate separations. For a gold sphere as small as 8 micron, the maximum error made by Eq. (15) was only of 0.1%, and this was for the large aspect ratio a/R = 0.12. The error is expected to be far smaller in the conditions of the experiment [21], which used spheres with radii larger than 29.8 micron, and probed distances corresponding to aspect ratios smaller than 0.017.

For later use, it is useful to work out the small distance limit of our formula for the sphere-sphere force gradient. This can be easily done substituting $F'_{n=0}^{(ex)}$ on the rhs of Eq. (15) by its small distance expansion Eq. (10). After simple algebraic manipulations, one finds

$$F' = -2\pi \tilde{R} F^{(\mathrm{pp})}(a) \left[1 - (\hat{\theta}(a) + u\hat{\kappa}(a)) \frac{a}{\tilde{R}} + o(a/\tilde{R}) \right].$$
(16)

The coefficients $\hat{\theta}(a)$ and $\hat{\kappa}(a)$ are

$$\hat{\theta}(a) = \frac{F_{n>0}^{(\text{pp})}}{F^{(\text{pp})}} \tilde{\theta} - \frac{F_{n=0}^{(\text{pp})}}{F^{(\text{pp})}} \frac{1}{12\zeta(3)},$$
(17)

$$\hat{\kappa}(a) = \frac{F_{n>0}^{(\text{pp})}}{F^{(\text{pp})}}\tilde{\kappa}.$$
(18)

In Table II, we provide the values of the coefficients $\hat{\theta}$ and $\hat{\kappa}$ for gold at room temperature. We note that the coefficient $\hat{\theta}$ introduced here coincides with the opposite of the coefficient $\hat{\theta}_1$ of [39].

In order to show the deviations of the force gradient from PFA predicted by our formula, in Fig. 2 we plot the quantity $(\tilde{R}/a)(F'/F'_{\rm PFA}-1)$ for a system of two identical spheres of radius $R_1 = R_2 = 30 \ \mu m$ (lower solid line) and $R_1 = R_2 = 100 \ \mu m$ (lower dashed line). The upper solid and dashed lines in Fig. 2 refer to a sphere-plate system, for a sphere of radius $R = 30 \ \mu m$ (solid line) and $R = 100 \ \mu m$ (dashed line). Figure 2 demonstrates that deviations from

PFA of the force gradient are practically independent of the effective radius \tilde{R} , for fixed value of u. However deviations from PFA depend significantly on the ratio among the radii of the spheres via the parameter u. This is clearly seen from Fig. 3, where the quantity $(\tilde{R}/a)(F'/F'_{PFA} - 1)$ is displayed versus the parameter u (for constant \tilde{R} and a). In Fig. 3 solid lines are for $\tilde{R} = 30 \ \mu$ m, while dashed lines are for $\tilde{R} = 100 \ \mu$ m. The four pairs of solid and dashed lines from top to bottom correspond to the four separations $a = 1 \ \mu$ m, 800 nm, 400 nm, and 100 nm, respectively. We recall that the extreme values u = 0 and u = 1/4 correspond, respectively, to a sphere plate and to two spheres of equal radii. Figure 3 shows that the u dependence of the deviations from PFA is linear for the considered separations.

For small values of a/\tilde{R} , the quantity displayed in Figs. 2 and 3 can be basically identified with the parameter β' that was introduced in the sphere-plate experiment [23] as a measure of the deviation of the data from PFA. In [23] starting from the force gradient F', an effective pressure $P^{(\text{eff})}(a, R)$ was defined as

$$P^{(\text{eff})}(a,R) \equiv -\frac{F'}{2\pi R}.$$
(19)

If the PFA Eq. (4) was exact, $P^{(\text{eff})}(a, R) = F^{(\text{pp})}(a)$. However, the PFA is not exact, and so the authors of [23] parametrized deviations from PFA by a coefficient $\beta'(a)$ such that

$$P^{(\text{eff})} = F^{(\text{pp})}(a) \left(1 + \beta' \frac{a}{R} + o(a/R) \right).$$
(20)

It is clear from Eq. (20) that, up to higher order corrections, $\beta'(a)$ coincides with the quantity $(\tilde{R}/a)(F'/F'_{PFA} - 1)$. The parameter β' was determined in [23] by measuring the effective pressure $P^{(eff)}$ for certain fixed sphere plate separations using spheres of different radii, and then fitting $P^{(eff)}$ versus 1/R with a straight line. The experiment [23] placed a bound $|\beta'| < 0.4$ at 95% C.L. in the separation range from 150 to 300 nm. This bound is in substantial agreement with the theoretical prediction (see the upper curves of Fig. 2).

The authors of [21] used the same procedure to study deviations of their data from PFA. In particular, they assumed that the measured forces can be parametrized as

TABLE II. Values of the coefficients $\hat{\theta}$ and $\hat{\kappa}$ for Au at room temperature.

$a(\mu m)$	0.05	0.10	0.15	0.2	0.25	0.3	0.35	0.4	0.45	0.5	0.55	0.6
$\hat{\theta}$	0.378	0.439	0.449	0.443	0.432	0.419	0.405	0.392	0.378	0.365	0.352	0.340
ĥ	0.209	0.237	0.259	0.275	0.288	0.298	0.306	0.313	0.320	0.325	0.330	0.334



FIG. 2. Beyond-PFA corrections for the gradient of the Casimir force between two gold spheres at room temperature are shown as a function of the separation for two identical spheres of radius $R = 30 \ \mu m$ (lower solid line) and $R = 100 \ \mu m$ (lower dashed line). The upper pair of lines is for a sphere-plate system with sphere radius $R = 30 \ \mu m$ (upper solid line) and $R = 100 \ \mu m$ (upper dashed line).

$$\frac{F'}{\tilde{R}} = -2\pi F^{(\text{pp})}(a) \left(1 + \beta' \frac{a}{\tilde{R}} + o(a/\tilde{R}) \right), \quad (21)$$

i.e., by a function of the same form as that of the sphereplate system, apart from the substitution of R by the effective radius \tilde{R} of the two-sphere system. Importantly, they assumed that β' is independent of the radii of the spheres. Based on this assumption, the authors of [21] tried to determine $\beta'(a)$ by doing a linear fit of F'/\tilde{R} versus $1/\tilde{R}$ using for that purpose 12 measurements of F'. Three sets of



FIG. 3. Beyond-PFA corrections for the gradient of the Casimir force between two gold spheres at room temperature are shown as a function of the parameter $u = \tilde{R}^2/(R_1R_2)$ (for constant \tilde{R} and *a*). Solid lines are for $\tilde{R} = 30 \ \mu$ m; dashed lines for $\tilde{R} = 100 \ \mu$ m. The four pairs of solid and dashed lines from top to bottom correspond to the four separations $a = 1 \ \mu$ m, 800 nm, 400 nm, and 100 nm, respectively. The extreme values u = 0 and u = 1/4 correspond, respectively, to a sphere-plate configuration and to two spheres of equal radii.

data were taken in sphere-plate setups (as in the experiment [23]), using three different spheres of radii $R = 40.7 \ \mu m$, 36.1 μm , and 34.2 μm , while the remaining nine data sets were taken with nine different sphere-sphere setups, corresponding to the different combinations of each of three spheres with radii $R_1 = 34.2 \ \mu m$, 36.1 μm , and 40.7 μm , with each of the three spheres of radii $R_2 = 29.8$, 38.0, and 46.9 μm . The measurements and the corresponding fits were repeated for 26 values of the separation, in the interval from 40 to 300 nm. It was found that $\beta' = -6 \pm 27$ was within the 2σ confidence interval of their calculated β' for all considered separations.

The procedure used in [21] to determine β' is not entirely correct, however, because the parametrization in Eq. (21) misses the dependence of β' on the parameter *u*. Indeed, by comparing Eq. (21) with the small distance expansion of the sphere-sphere force Eq. (16) one finds that β' has the expression

$$\beta' = -(\hat{\theta}(a) + u\hat{\kappa}(a)). \tag{22}$$

This formula shows that in the two-sphere case, contrary to the assumption made in [21], β' does depend on the radii of the spheres via the parameter *u*. Of course, this dependence disappears in the sphere-plate case, for which u = 0. The linear dependence on *u* of deviations from PFA is clearly visible from Fig. 3. Since in [21] the 12 combinations of radii used to determine β' correspond to values of *u* that vary from 0 (for the three sphere-plate setups) to 0.2498 (corresponding to the sphere-sphere setup with $R_1 =$ 36.1 μ m and $R_2 = 38 \ \mu$ m), the dependence of β' on *u* should be considered in the data analysis. Substituting Eq. (22) into Eq. (21), we find that in a two-sphere system F'/\tilde{R} has the expression

$$\frac{F'}{\tilde{R}} = -2\pi F^{(\text{pp})}(a) \left(1 - \frac{a\hat{\theta}}{\tilde{R}} - \frac{a\hat{\kappa}}{R_1 + R_2} + o(a/\tilde{R}) \right), \quad (23)$$

where in the second term between the brackets we used the relation $u/\tilde{R} = 1/(R_1 + R_2)$. This formula shows that the correct procedure to determine the coefficients $\hat{\theta}$ and $\hat{\kappa}$ is to make a joint two-dimensional linear fit of F'/\tilde{R} versus $1/\tilde{R}$ and $1/(R_1 + R_2)$.

The present sensitivity of the experiment [21] is not yet sufficient to detect the small deviations from PFA predicted by Eq. (23). We estimate that an increase in the sensitivity by over 1 order of magnitude would be necessary for that purpose. It is hoped that future improvements of the apparatus will achieve this goal.

V. CONCLUSIONS

Motivated by the recent experiment in [21], we have performed a precise computation of the gradient of the Casimir force between two gold spheres at room temperature. Our computation provides an accurate estimate of beyond-PFA corrections for this system. The semianalytic formula for the Casimir force that we construct is valid for all separations and can be easily used to interpret future experiments in both the sphere-plate and sphere-sphere configurations. We have also described the correct parametrization of the corrections to PFA that should be used to carry out the data analysis in experiments using the sphere-sphere geometry.

In our computations we modeled the gold plates as ohmic conductors (connected to charge reservoirs). In recent years it has been argued by some researchers [4] that a better agreement with Casimir experiments is obtained if metallic bodies are modeled as dissipationless plasmas. The main change introduced by this model is in the classical n = 0 Matsubara term for transverse electric (TE) polarization, which is 0 within the Drude prescription, but different from 0 in the plasma model. A detailed comparison between the Drude and plasma models for the sphere-plate configuration, based on a large-scale numerical simulation of the scattering formula, has been reported in [32], where it was shown that the deviations from PFA engendered by the plasma prescription have the same qualitative behavior as the Drude model, but are slightly larger in magnitude and show a more pronounced dependence on the aspect ration a/R. We plan to study the plasma prescription for the sphere-sphere case in a forthcoming work [47].

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APPENDIX: THE DE EXPANSION

1. General formulae

For the convenience of the reader, in this appendix we briefly review general properties of the DE that are useful for the present work. The DE [35-39] is an asymptotic expansion that allows one to compute curvature corrections of any sufficiently local functional $\hat{\mathcal{F}}$ that describes the interaction between two (nonintersecting) surfaces Σ_1 and Σ_2 . The idea behind the DE is intuitive. One considers that the two surfaces can be described by smooth height profiles $z = H_1(x, y)$ and $z = H_2(x, y)$, where (x, y) are Cartesian coordinates spanning some reference plane Σ and z is a coordinate perpendicular to Σ (see Fig. 1). Since the two surfaces are nonintersecting, it can always be assumed that $H_2(x, y) < H_1(x, y)$. At this point one considers that for surfaces of small slopes $|\nabla H_i| \ll 1, i = 1, 2$ [48] it should be possible to expand $\hat{\mathcal{F}}[H_1, H_2]$ in powers of derivatives of increasing order of the height profiles, at least up to some order. It is rather easy to convince oneself that for a functional $\hat{\mathcal{F}}[H_1, H_2]$ that is invariant under simultaneous rotations and translations of H_1 and H_2 in the reference plane Σ (like the Casimir force between two plates made of a homogeneous and isotropic material) the most general expression of the DE valid to second order in the slopes of the surfaces is of the form

$$\begin{aligned} \hat{\mathcal{F}}[H_1, H_2] &= \int_{\Sigma} d^2 x [\hat{\mathcal{F}}^{(\text{pp})}(H) + \alpha_1(H) (\nabla H_1)^2 \\ &+ \alpha_2(H) (\nabla H_2)^2 + \alpha_\times(H) \nabla H_1 \cdot \nabla H_2 \\ &+ \alpha_-(H) \nabla H_1 \times \nabla H_2] + \rho^{(2)}, \end{aligned} \tag{A1}$$

where we set $H = H_1 - H_2$ and $\rho^{(2)}$ is a remainder that becomes negligible as the local radii of curvature of the surfaces go to infinity for fixed minimum surface-surface distance *a*. Note that invariance of $\hat{\mathcal{F}}$ under translations of Σ in the *z* direction implies that $\hat{\mathcal{F}}^{(\text{pp})}$ and the α 's can depend only on the height difference *H* and not on the individual heights H_1 and H_2 . It is evident that the quantity $\hat{\mathcal{F}}^{(\text{pp})}(a)$ in Eq. (A1) provides the (unit-area) interaction of two plane-parallel surfaces at distance *a*, and thus the first term on the rhs of Eq. (A1) reproduces the Derjaguin approximation (DA) [49] for the functional $\hat{\mathcal{F}}$,

$$\hat{\mathcal{F}}^{(\mathrm{DA})} = \int_{\Sigma} d^2 x \hat{\mathcal{F}}^{(\mathrm{pp})}(H).$$
 (A2)

The integrals on the rhs of Eq. (A1) that are proportional to α 's represent curvature corrections beyond the DA, and thus we see that the DE provides a systematic way to improve the old-fashioned DA. Arbitrariness in the choice of the reference plane Σ further constrains the three coefficients α in Eq. (A1) [36]. In particular, invariance of $\hat{\mathcal{F}}$ with respect to tilting of Σ (for details, see [36]) implies

$$2(\alpha_1(H) + \alpha_2(H) + \alpha_{\times}(H)) + H \frac{d\hat{\mathcal{F}}^{(\text{pp})}}{dH} - \hat{\mathcal{F}}^{(\text{pp})} = 0,$$

$$\alpha_-(H) = 0.$$

(A3)

The above relations show that, to second order in the gradient expansion, the two-surface problem actually reduces to the simpler problem of a single curved surface opposite a plane, since α_1 and α_2 can be determined in that case, and then α_{\times} follows from the first of Eqs. (A3). We now make the simplifying assumption that the field(s) that mediate the interaction obeys the same boundary conditions on Σ_1 and Σ_2 . Then

$$\alpha_1(H) = \alpha_2(H) \equiv \alpha(H). \tag{A4}$$

Taking advantage of Eqs. (A3) and (A4) the DE can then be recast in the form

$$\begin{split} \hat{\mathcal{F}}[H_1, H_2] &= \hat{\mathcal{F}}^{(\mathrm{DA})} + \int_{\Sigma} d^2 x \alpha(H) (\nabla H)^2 \\ &+ \frac{1}{2} \int_{\Sigma} d^2 x \left(\hat{\mathcal{F}}^{(\mathrm{pp})} - H \frac{d \hat{\mathcal{F}}^{(\mathrm{pp})}}{dH} \right) \nabla H_1 \cdot \nabla H_2 \\ &+ \rho^{(2)}. \end{split} \tag{A5}$$

We thus see that to second order in the slope the interaction $\hat{\mathcal{F}}$ is fully determined by knowledge of the (unit-area) interaction $\hat{\mathcal{F}}^{(\text{pp})}(a)$ of two parallel plates and by the single coefficient $\alpha(H)$.

The latter coefficient can be determined by comparing the DE Eq. (A5) to a perturbative expansion of the functional $\hat{\mathcal{F}}[H, 0]$ around flat plates H = a + h(x, y) to second order in the deformation h(x, y). Note that the latter perturbation requires a deformation of small amplitude $h(x, y)/a \ll 1$, while the DE relies on the condition that the slope of the surface be small. To second order in *h* the perturbative expansion of $\hat{\mathcal{F}}$ reads

$$\begin{aligned} \hat{\mathcal{F}}[a+h(\mathbf{x})] &= A\hat{\mathcal{F}}^{(\mathrm{pp})}(a) + \mu(a)\tilde{h}(\mathbf{0}) \\ &+ \int \frac{d^2\mathbf{k}}{(2\pi)^2} \tilde{G}^{(2)}(k;a) |\tilde{h}(\mathbf{k})|^2 + \tilde{\rho}^{(2)}[h], \end{aligned}$$
(A6)

where A is the surface area, **k** is the in-plane wave vector, $\tilde{h}(\mathbf{k})$ is the Fourier transform of h(x), and $\tilde{\rho}^{(2)}[h]$ refers to higher order corrections. The function $\alpha(H)$ can now be determined if the kernel $\tilde{G}^{(2)}(k;a)$ can be expanded to second order in k. Indeed, matching the expansion

$$\tilde{G}^{(2)}(k;a) = \gamma(a) + \delta(a)k^2 + o(k^2)$$
 (A7)

to Eq. (A5) one finds

$$\hat{\mathcal{F}}^{(\mathrm{pp})\prime}(a) = \mu(a), \qquad \hat{\mathcal{F}}^{(\mathrm{pp})\prime\prime}(a) = 2\gamma(a), \qquad \alpha(a) = \delta(a),$$
(A8)

where a prime denotes a derivative with respect to *a*. The above equation shows that a necessary condition for existence of the second-order DE is existence of the Taylor expansion of the perturbative kernel $\tilde{G}^{(2)}(k;a)$ to second order in the in-plane momentum. Indeed, it can be shown that the DE can be formally recovered by an (infinite) resummation of the perturbative series for small in-plane momenta [38].

Whenever applicable, the DE has been successfully used to compute curvature corrections beyond the PA in various problems involving interactions among gently curved surfaces. In the context of Casimir physics, it was used in [35] to compute curvature corrections to the zero temperature Casimir energy for a scalar field obeying D bc in the sphere-plate and cylinder-plate geometries. The zero temperature Casimir problem for the em field with PC bc, as well as a scalar field obeying Neumann bc, or mixed DN bc (i.e., D bc on one surface and N on the other), was studied in [36] for two spheres and for two inclined cylinders. The curvature corrections obtained in the latter work for the em field with PC bc in the sphere-plate and sphere-sphere geometries were subsequently confirmed in [41,50] by working out a rigorous small distance expansion of the scattering formula. The experimentally important case of the Casimir interaction between gold sphere and plate at finite temperature was instead studied in [39]. Even in this case, the results obtained by the DE were later shown to be in agreement with the small distance expansion of the scattering formula [51]. Curvature corrections obtained by the DE have also been found to be in agreement with the small distance expansion of the rare exact Casimir energies in nonplanar geometries that have been discovered so far, i.e., in the cases of two Drude or D spheres in the classical limit [34], and for two three-spheres with D or PC bc in four Euclidean dimensions [52]. The DE has been also used to study curvature effects in the Casimir-Polder interaction of a particle with a gently curved surface [53,54], and to estimate the shifts of the rotational levels of a diatomic molecule due to its van der Waals interaction with a curved dielectric surface [55]. In a non-Casimir context, the DE hase been also used to compute curvature corrections to the scattering amplitude for an em wave impinging on a curved surface [56] and to the electrostatic interaction among two curved plates [57].

2. Computing the leading curvature correction to the force gradient

The small-slope approximation of the interaction energy $\hat{\mathcal{F}}$ provided by Eq. (A5) still involves a surface integral over Σ of functions depending on the height profiles of the surfaces. As such, Eq. (A5) is not very convenient for a practical use. A better route is to expand Eq. (A5) in powers of the small parameter a/R, where R is the characteristic radius of curvature of the surfaces. The leading order of this expansion reproduces the standard PFA, while in the next order it provides us with the desired curvature correction beyond the PFA. We carry out this expansion not directly for the energy $\hat{\mathcal{F}}$, but rather for the gradient of the force $\hat{F}' = -\hat{\mathcal{F}}''$, which is the quantity that was measured in the experiment [21]. Moreover, we restrict attention to the sphere-sphere system, which is again the geometry used in [21].

According to Eq. (A5), the formula for the force gradient \hat{F}' can be split as

$$\hat{F}' = \hat{F}'^{(\mathrm{DA})} + I_2 + I_3,$$
 (A9)

where

and we set

$$I_2 = \frac{1}{2} \int_{\Sigma} d^2 x (H \hat{\mathcal{F}}^{(\mathrm{pp})\prime\prime})' \nabla H_1 \cdot \nabla H_2,$$

$$I_3 = -\int_{\Sigma} d^2 x \alpha'' (H) (\nabla H)^2.$$
 (A11)

We consider proximity forces that decay rapidly with the distance, like the Casimir force. For forces of this nature, the interaction among the surfaces is localized within a small area, typically of radius $\rho \sim \sqrt{a\tilde{R}}$, around the point of closest approach. Under such circumstances, it is legitimate to take the Taylor expansion of the height profiles $H_1(x, y)$ and $H_2(x, y)$ of the two spheres around their tips, which we imagine placed at x = y = 0. Since the position of the reference plane Σ in Fig. 1 is immaterial, we are free to take for Σ the tangent plane to the sphere of radius R_2 , passing through the sphere tip. Then,

$$H_1(x, y) = a + \frac{r^2}{2R_1} + \frac{r^4}{8R_1^3} + \cdots,$$

$$H_2(x, y) = -\frac{r^2}{2R_2} - \frac{r^4}{8R_2^3} + \cdots,$$
 (A12)

where $r^2 = x^2 + y^2$. To evaluate the integrals I_j it is convenient to introduce polar coordinates (r, θ) in the (x, y) plane, and then substitute r by the dimensionless quantity $\xi = r^2/a\tilde{R}$. An essential property of the integrals I_j is that they involve derivatives of certain functions (i.e., $\hat{\mathcal{F}}^{(\text{pp})}$ and α) of the height difference H with respect to the separation a. These derivatives can be converted into derivatives with respect to ξ , using the identity

$$U' = U_{,\xi}(H_{,\xi})^{-1}$$

= $2\frac{U_{,\xi}}{a} \left[1 - \frac{\xi}{2} \left(\frac{a\tilde{R}^2}{R_1^3} + \frac{a\tilde{R}^2}{R_2^3} \right) + o(a/\tilde{R}) \right]$ (A13)

(comas denote derivatives), which holds for any function U of H. We are ready now to take the small distance expansion of \hat{F}' . We start from $\hat{F}'^{(\text{DA})}$. Using Eq. (A13), and omitting corrections of order $o(a/\tilde{R})$ we find

$$\hat{F}'^{(\text{DA})} = 2\pi\tilde{R} \int_0^\infty d\xi \hat{F}^{(\text{DA})}_{,\xi} \left[1 - \frac{\xi}{2} \left(\frac{a\tilde{R}^2}{R_1^3} + \frac{a\tilde{R}^2}{R_2^3} \right) \right] \\ = -2\pi\tilde{R}\hat{F}^{(\text{pp})}(a) - a\pi \left(\frac{\tilde{R}^3}{R_1^3} + \frac{\tilde{R}^3}{R_2^3} \right) \int_0^\infty d\xi \hat{\mathcal{F}}^{(\text{pp})\prime} \\ = -2\pi\tilde{R}\hat{F}^{(\text{pp})}(a) + 2\pi \left(\frac{\tilde{R}^3}{R_1^3} + \frac{\tilde{R}^3}{R_2^3} \right) \hat{\mathcal{F}}^{(\text{pp})}(a). \quad (A14)$$

The first term on the last line of Eq. (A14) coincides with the standard PFA for the force gradient [see Eq. (4)], while its second term represents a curvature correction.

By following an analogous procedure for I_2 , and again omitting higher order terms, we obtain

$$\begin{split} I_{2} &= -\pi \frac{a\tilde{R}^{2}}{R_{1}R_{2}} \int_{0}^{\infty} d\xi (H\hat{\mathcal{F}}^{(\mathrm{pp})\prime\prime})_{,\xi} \xi \\ &= \pi \frac{a\tilde{R}^{2}}{R_{1}R_{2}} \int_{0}^{\infty} d\xi (H\hat{\mathcal{F}}^{(\mathrm{pp})\prime\prime}) = \frac{2\pi\tilde{R}^{2}}{R_{1}R_{2}} \int_{0}^{\infty} d\xi H(\hat{\mathcal{F}}^{(\mathrm{pp})\prime})_{,\xi} \\ &= \frac{2\pi\tilde{R}^{2}}{R_{1}R_{2}} a\hat{F}^{(\mathrm{pp})}(a) - \frac{2\pi\tilde{R}^{2}}{R_{1}R_{2}} \int_{0}^{\infty} d\xi \hat{\mathcal{F}}^{(\mathrm{pp})}_{,\xi} \\ &= \frac{2\pi\tilde{R}^{2}}{R_{1}R_{2}} [a\hat{F}^{(\mathrm{pp})}(a) + \hat{\mathcal{F}}^{(\mathrm{pp})}(a)]. \end{split}$$
(A15)

Finally, for I_3 we obtain

$$I_3 = -4\pi \int_0^\infty d\xi \alpha_{\xi\xi} \xi = -4\pi \alpha(a).$$
(A16)

Upon combining Eqs. (A14)–(A16), after simple algebraic transformations, we obtain the following small distance expansion of the force gradient, correct up to terms of order $o(a/\tilde{R})$:

$$\hat{F}' = -2\pi \tilde{R} \hat{F}^{(\text{pp})}(a) + 2\pi [\hat{\mathcal{F}}^{(\text{pp})}(a) - 2\alpha(a)] + 2\pi u [a \hat{F}^{(\text{pp})}(a) - 2\hat{\mathcal{F}}^{(\text{pp})}(a)] \equiv -2\pi \tilde{R} \hat{F}^{(\text{pp})}(a) \left[1 - (\tilde{\theta}(a) + u\kappa(a)) \frac{a}{\tilde{R}} \right], \quad (A17)$$

where the coefficients $\tilde{\theta}(a)$ and $\kappa(a)$ are

$$\tilde{\theta} = \frac{\hat{\mathcal{F}}^{(\text{pp})}(a) - 2\alpha(a)}{a\hat{F}^{(\text{pp})}(a)},$$
(A18)

$$\kappa(a) = 1 - 2\frac{\hat{\mathcal{F}}^{(\text{pp})}(a)}{a\hat{F}^{(\text{pp})}(a)}.$$
(A19)

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