# Charmonium spectrum and electromagnetic transitions with higher multipole contributions

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The charmonium spectrum is calculated with two nonrelativistic quark models, the linear potential model, and the screened potential model. Using the obtained wave functions, we evaluate the electromagnetic transitions of charmonium states up to 4*S* multiplet. The higher multipole contributions are included by a multipole expansion of the electromagnetic interactions. Our results are in reasonable agreement with the measurements. As conventional charmonium states, the radiative decay properties of the newly observed charmoniumlike states, such as X(3823), X(3872), and X(4140/4274), are discussed. The X(3823) as  $\psi_2(1D)$ , its radiative decay properties well agree with the observations. From the radiative decay properties of X(3872), one cannot exclude it as a  $\chi_{c1}(2P)$  dominant state. We also give discussions of possibly observing the missing charmonium states in radiative transitions, which might provide some useful references to look for them in forthcoming experiments. The higher multipole contributions to the electromagnetic transitions are analyzed as well. It is found that the higher contribution from the magnetic part could give notable corrections to some E1 dominant processes by interfering with the E1 amplitudes. Our predictions for the normalized magnetic quadrupole amplitudes  $M_2$  of the  $\chi_{c1,2}(1P) \rightarrow J/\psi\gamma$  processes are in good agreement with the recent CLEO measurements.

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# I. INTRODUCTION

During the past few years, great progress has been achieved in the observation of the charmonia [1-5]. From the review of the Particle Data Group (PDG) [6], one can see that many new charmoniumlike "XYZ" states above open-charm thresholds states have been observed at Belle, BABAR, LHC, BESIII, CLEO, and so on. The observations of these new states not only deepen our understanding of the charmonium physics but also bring us many mysteries in this field to be uncovered [3-5]. If these newly observed XYZ states, such as X(3872), X(3915), X(4140/4274), and Y(4260), are assigned asconventional charmonium states, some properties, such as measured mass and decay modes, may be inconsistent with the predictions. Thus, how to identify these newly observed charmoniumlike XYZ states and how to understand their uncommon nature are great challenges for physicists.

Stimulated by the extensive progress made in the observation of the charmonia, in this work, we study the mass spectrum and electromagnetic (EM) transitions of the charmonium within the widely used linear potential model [7–9] and the screened potential model [10,11]. As we know, the EM decays of a hadron are sensitive to its inner structure. The study of the EM decays is not only

crucial for us to determine the quantum numbers of the newly observed charmonium states but also provides very useful references for our search for the missing charmonium states in experiments. To study the charmonium spectrum and/or their EM decays, besides the widely used potential models [7-18], some other models, such as lattice QCD [19–26], QCD sum rules [27–29], coupled-channel quark models [30], the effective Lagrangian approach [31,32], nonrelativistic effective field theories of QCD [33–36], the relativistic quark model [37], the relativistic Salpeter method [38], the light front quark model [39], the Coulomb gauge approach [40], and the generalized screened potential model [41] have been employed in theory. Recently, the hadronic loop contributions to the radiative decay of charmonium states were also discussed in Refs. [42–44]. Although there are many studies about the EM decays of charmonium states, many properties are not well understood. For example, the predictions for the  $\chi_{cI}(1P) \rightarrow J/\psi\gamma$  and  $\psi(3770) \rightarrow \chi_{cI}(1P)\gamma$  processes are rather different in various models [45]. These differences may come from the wave functions of charmonium states adopted, the higher EM multipole amplitude contributions, the coupled-channel effects, and so on. Thus, to clarify these puzzles, more studies are needed.

In this work, we mainly focus on the following issues: (i) To clearly show the model dependence of the higher charmonium states, we calculate the charmonium spectroscopy within two typical models, i.e., the linear and screened

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potential models. As done in the literature, e.g., Refs. [7–9,18], the spin-dependent potentials are dealt with nonperturbatively so that the corrections of the spin-dependent interactions to the wave functions can be included. (ii) We further analyze the EM transitions between charmonium states. Based on the obtained radiative decay properties and mass spectrum, we discuss the classifications of the newly observed charmoniumlike states, while for the missing excited states, we suggest strategies to find them in radiative transitions. (iii) Finally, we discuss the possible higher EM multipole contributions to an EM transition process.

The paper is organized as follows. In Sec. II, the charmonium spectroscopy is calculated within both the linear and screened potential models. In Sec. III, first, we give an introduction of EM transitions described in the present work. Then, using the wave functions obtained from both the linear and screened potential models, we analyze the EM decays of charmonium states. Finally, a summary is given in Sec. IV.

# **II. MASS SPECTROSCOPY**

#### A. Formalism

In this work, the mass and space wave function of a charmonium state are determined by the Schrödinger equation with a conventional quarkonium potential. The effective potential of spin-independent term V(r) between the quark and antiquark is regarded as the sum of Lorentz vector  $V_V(r)$  and Lorentz scalar  $V_s(r)$  contributions [1], i.e.,

$$V(r) = V_V(r) + V_s(r).$$
 (1)

The Lorentz vector potential  $V_V(r)$  is adopted by the standard color Coulomb form:

$$V_V(r) = -\frac{4}{3} \frac{\alpha_s}{r}.$$
 (2)

The Lorentz scalar  $V_s(r)$  might be taken as

$$V_{s}(r) = \begin{cases} br, & \text{linear potential} \\ \frac{b}{\mu}(1 - e^{-\mu r}), & \text{screened potential}, \end{cases} (3)$$

where *r* is the distance between the quark and antiquark. The linear potential *br* is widely used in the potential models. Considering the screening effect from the vacuum polarization effect of the dynamical light quark might soften the linear potential at large distances [46,47], people suggested a screened potential  $b(1 - e^{-\mu r})/\mu$  in the calculations as well [10,11,15,16,48]. Here,  $\mu$  is the screening factor which makes the long-range scalar potential of  $V_s(r)$  behave like *br* when  $r \ll 1/\mu$  and becomes a constant  $b/\mu$  when  $r \gg 1/\mu$ . The main effects of the screened potential on the spectrum is that the masses of the higher excited states are lowered.

Following the method in Refs. [8,11], we include three spin-dependent potentials in our calculations. For the spin-spin contact hyperfine potential, we take the Gaussian-smeared form [8]

$$H_{SS} = \frac{32\pi\alpha_s}{9m_c^2}\tilde{\delta}_{\sigma}(r)\mathbf{S}_c\cdot\mathbf{S}_{\bar{c}},\qquad(4)$$

where  $\mathbf{S}_c$  and  $\mathbf{S}_{\bar{c}}$  are spin matrices acting on the spins of the quark and antiquark. We take  $\tilde{\delta}_{\sigma}(r) = (\sigma/\sqrt{\pi})^3 e^{-\sigma^2 r^2}$  as in Ref. [8]. The five parameters in the above equations ( $\alpha_s$ , b,  $\mu$ ,  $m_c$ ,  $\sigma$ ) are determined by fitting the spectrum.

For the spin-orbit term and the tensor term, we take the common forms obtained from the leading-order perturbation theory,

$$H_{SL} = \frac{1}{2m_c^2 r} \left( 3 \frac{dV_V}{dr} - \frac{dV_s}{dr} \right) \mathbf{L} \cdot \mathbf{S}, \tag{5}$$

and

$$H_T = \frac{1}{12m_c^2} \left( \frac{1}{r} \frac{dV_V}{dr} - \frac{d^2 V_V}{dr^2} \right) S_T,$$
 (6)

where **L** is the relative orbital angular momentum of *c* and  $\bar{c}$  quarks,  $\mathbf{S} = \mathbf{S}_c + \mathbf{S}_{\bar{c}}$  is the total quark spin, and the spin tensor  $S_T$  is defined by

$$S_T = 6 \frac{\mathbf{S} \cdot \mathbf{r} \mathbf{S} \cdot \mathbf{r}}{r^2} - 2\mathbf{S}^2.$$
(7)

By solving the radial Schrödinger equation  $\frac{d^2u(r)}{dr^2} + 2\mu_R[E - V_{c\bar{c}}(r) - \frac{L(L+1)}{2\mu_R r^2}]u(r) = 0$ , with  $V_{c\bar{c}}(r) \equiv V(r) + H_{SS} + H_{SL} + H_T$  and  $\mu_R \equiv m_c m_{\bar{c}}/(m_c + m_{\bar{c}})$ , we obtain the wave function u(r) and the mass  $M_{c\bar{c}} = 2m_c + E$  for a charmonium state. For simplification, the spin-dependent interactions can be dealt with perturbatively. Although the meson mass obtains perturbative corrections from these spin-dependent potentials, the wave functions obtain no corrections from them. Thus, to reasonably include the corrections from these spin-dependent potentials to both the mass and wave function of a meson state, we deal with the spin-dependent interactions nonperturbatively.

In this work, we solve the radial Schrödinger equation by using the three-point difference central method from the center (r = 0) toward the outside  $(r \to \infty)$  point by point. The details of this method can be found in Ref. [49]. To overcome the singular behavior of  $1/r^3$  in the spindependent potentials, following the method of our previous work [50], we introduce a cutoff distance  $r_c$  in the calculation. Within a small range  $r \in (0, r_c)$ , we let  $1/r^3 = 1/r_c^3$ . It is found that the masses of the  ${}^3P_0$  states are sensitive to the cutoff distance  $r_c$ , which is easily determined by the mass of  $\chi_{c0}(1P)$ .

Considering the progress in the charmonium spectrum in recent years, we do not use the old parameter sets determined in Refs. [8,11]. Combining the new measurements, we slightly adjust the parameter sets of Refs. [8,11] to better describe the data. By fitting the masses of the 12 well-established  $c\bar{c}$  states given in Table I, we obtain the parameter sets for the linear potential model and screened potential model, which are given in Table II.

# B. Results and discussions

Our calculated masses for the nS ( $n \le 5$ ), nP, and nD ( $n \le 3$ ) charmonium states with both the linear and

TABLE I. Charmonium mass spectrum. LP and SP stand for our calculated masses with the linear potential and screened potential models, respectively. For comparison, the measured masses (MeV) from the PDG [6] and the previous predictions with screened potential in Ref. [11] and linear potential in Ref. [8] are also listed.

$n^{2S+1}L_J$	Name	$J^{PC}$	Exp. [6]	[8]	[11]	LP	SP
$1^{3}S_{1}$	$J/\psi$	1	3097 <sup>a</sup>	3090	3097	3097	3097
$1^{1}S_{0}$	$\eta_c(1S)$	$0^{-+}$	2984 <sup>a</sup>	2982	2979	2983	2984
$2^{3}S_{1}$	$\psi(2S)$	1	3686 <sup>a</sup>	3672	3673	3679	3679
$2^{1}S_{0}$	$\eta_c(2S)$	$0^{-+}$	3639 <sup>a</sup>	3630	3623	3635	3637
$3^{3}S_{1}$	$\psi(3S)$	1	$4040^{a}$	4072	4022	4078	4030
$3^{1}S_{0}$	$\eta_c(3S)$	$0^{-+}$		4043	3991	4048	4004
$4^{3}S_{1}$	$\psi(4S)$	1	4415?	4406	4273	4412	4281
$4^{1}S_{0}$	$\eta_c(4S)$	$0^{-+}$		4384	4250	4388	4264
$5^{3}S_{1}$	$\psi(5S)$	1			4463	4711	4472
$5^{1}S_{0}$	$\eta_c(5S)$	$0^{-+}$			4446	4690	4459
$1^{3}P_{2}$	$\chi_{c2}(1P)$	$2^{++}$	3556 <sup>a</sup>	3556	3554	3552	3553
$1^{3}P_{1}$	$\chi_{c1}(1P)$	$1^{++}$	3511 <sup>a</sup>	3505	3510	3516	3521
$1^{3}P_{0}$	$\chi_{c0}(1P)$	$0^{++}$	3415 <sup>a</sup>	3424	3433	3415	3415
$1^{1}P_{1}$	$h_c(1P)$	1+-	3525 <sup>a</sup>	3516	3519	3522	3526
$2^{3}P_{2}$	$\chi_{c2}(2P)$	$2^{++}$	3927 <sup>a</sup>	3972	3937	3967	3937
$2^{3}P_{1}$	$\chi_{c1}(2P)$	$1^{++}$		3925	3901	3937	3914
$2^{3}P_{0}$	$\chi_{c0}(2P)$	$0^{++}$	3918?	3852	3842	3869	3848
$2^{1}P_{1}$	$h_c(2P)$	$1^{+-}$		3934	3908	3940	3916
$3^{3}P_{2}$	$\chi_{c2}(3P)$	$2^{++}$		4317	4208	4310	4211
$3^{3}P_{1}$	$\chi_{c1}(3P)$	$1^{++}$		4271	4178	4284	4192
$3^{3}P_{0}$	$\chi_{c0}(3P)$	$0^{++}$		4202	4131	4230	4146
$3^{1}P_{1}$	$h_c(3P)$	$1^{+-}$		4279	4184	4285	4193
$1^{3}D_{3}$	$\psi_3(1D)$	3		3806	3799	3811	3808
$1^{3}D_{2}$	$\psi_2(1D)$	2	3823 <sup>a</sup>	3800	3798	3807	3807
$1^{3}D_{1}$	$\psi_1(1D)$	1	3778 <sup>a</sup>	3785	3787	3787	3792
$1^{1}D_{2}$	$\eta_{c2}(1D)$	2-+		3799	3796	3806	3805
$2^{3}D_{3}$	$\psi_3(2D)$	3		4167	4103	4172	4112
$2^{3}D_{2}$	$\psi_2(2D)$	2		4158	4100	4165	4109
$2^{3}D_{1}$	$\psi_1(2D)$	1	4191?	4142	4089	4144	4095
$2^{1}D_{2}$	$\eta_{c2}(2D)$	$2^{-+}$		4158	4099	4164	4108
$3^{3}D_{3}$	$\psi_3(3D)$	3		•••	4331	4486	4340
$3^{3}D_{2}$	$\psi_2(3D)$	2		•••	4327	4478	4337
$3^{3}D_{1}$	$\psi_1(3D)$	1		•••	4317	4456	4324
$3^{1}D_{2}$	$\eta_{c2}(3D)$	$2^{-+}$			4326	4478	4336

<sup>a</sup>These masses for the 12 well-established  $c\bar{c}$  states are used as input to determine the model parameters.

TABLE II. Quark model parameters determined by the 12 wellestablished  $c\bar{c}$  states given in Table I.

Parameter	Linear potential model	Screened potential model
$m_c$ (GeV)	1.4830	1.4110
$\alpha_s$	0.5461	0.5070
$b (\text{GeV}^2)$	0.1425	0.2100
$\sigma$ (GeV)	1.1384	1.1600
$r_c$ (fm)	0.202	0.180
$\mu$ (GeV)	•••	0.0979

screened potential models have been listed in Table I, respectively. It is found that the mass spectrum calculated from the three-point difference central method is consistent with the previous calculations [8,11]. For the states with a mass of M < 4.1 GeV, both linear and screened potential models give a reasonable description of the mass spectrum compared with the data. However, for the higher resonances with a mass of M > 4.1 GeV, the predictions between these two models are very different. In the linear potential model, the well-established states  $\psi(4160)$  and  $\psi(4415)$ could be assigned as the  $\psi_1(2D)$  and  $\psi(4S)$ , respectively. However, in the screened potential model, the  $\psi(4415)$ might be assigned as  $\psi(5S)$  [11], while for  $\psi(4160)$ , the predicted mass is about 100 MeV less than the measurements. Comparing with the linear potential model, an obvious feature of the screened potential model is that it provides a compressed mass spectrum, which permits many new charmoniumlike XYZ states to be accommodated in the conventional higher charmonium states [11]. Lately, the BESIII Collaboration observed two resonant structures, one with a mass of ~4222 MeV and a width of ~44 MeV and the other with a mass of ~4320 MeV and a width of ~101 MeV, in the cross section for the process  $e^+e^- \rightarrow$  $\pi^+\pi^- J/\psi$  [51], which may correspond to the  $J^{PC} = 1^{--}$ states X(4260) and X(4360) from the PDG [6], respectively. It is found that within the screened potential model X(4260) and X(4360) are good candidates of the  $\psi(4S)$ and  $\psi_1(3D)$ , respectively. Furthermore, we should mention that recently two new charmoniumlike states X(4140) and X(4274) were confirmed by the LHCb Collaboration [52]. Their quantum numbers are determined to be  $J^{PC} = 1^{++}$ . Within the linear potential model, the X(4274) might be identified as the  $\chi_{c1}(3P)$  state, while within the screened potential model, the X(4140) is a good candidate of  $\chi_{c1}(3P)$ . However, neither the linear potential model nor the screened model can give two conventional  $J^{PC} = 1^{++}$ charmonium states with masses around 4.14 and 4.27 GeV at the same time, which may indicate the exotic nature of X(4140) and/or X(4274).

Furthermore, in Table III, we give our predictions of the hyperfine splittings for some *S*-wave states and fine splittings for some *P*-wave states with the linear and screened potentials, respectively. For a comparison, the world average data from the PDG [6] and the previous predictions in

TABLE III. Hyperfine and fine splittings in units of MeV for charmonia. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. The experimental data are taken from the PDG [6]. The theoretical predictions with the previous screened potential model (SNR model) [11], the relativized quark model (GI model), and non-relativistic linear potential model (NR model) [8] are also listed for comparison.

Splitting	LP	SP	SNR [11]	NR [8]	GI [8]	Exp. [6]
$\overline{m(1^3S_1) - m(1^1S_0)}$	114	113	118	108	113	$113.3 \pm 0.7$
$m(2^{3}S_{1}) - m(2^{1}S_{0})$	44	43	50	42	53	$46.7 \pm 1.3$
$m(3^3S_1) - m(3^1S_0)$	30	26	31	29	36	
$m(4^3S_1) - m(4^1S_0)$	22	17		22	25	
$m(5^3S_1) - m(5^1S_0)$	21	13				
$m(1^{3}P_{2}) - m(1^{3}P_{1})$	36	32	44	51	40	$45.5 \pm 0.2$
$m(1^{3}P_{1}) - m(1^{3}P_{0})$	101	106	77	81	65	$95.9 \pm 0.4$
$m(2^{3}P_{2}) - m(2^{3}P_{1})$	30	23	36	47	26	
$m(2^{3}P_{1}) - m(2^{3}P_{0})$	68	66	59	53	37	
$m(3^{3}P_{2}) - m(3^{3}P_{1})$	25	19	30	46	20	
$\frac{m(3^3P_1) - m(3^3P_0)}{m(3^3P_0)}$	51	46	47	69	25	

Refs. [8,11] are listed in the same table as well. It is found that both the linear and screened potential models give comparable results. The predicted splittings are in agreement with the world average data [6]. It should be mentioned that both the linear and screened potential models obtain a similar fine splitting between  $\chi_{c2}(2P)$  and  $\chi_{c0}(2P)$ , i.e.,  $\Delta m \approx 90$  MeV. According to the measured mass of  $\chi_{c2}(2P)$ , one can predict that the mass of  $\chi_{c0}(2P)$  is about 3837 MeV. Thus, assigning the X(3915) as the  $\chi_{c0}(2P)$  state is still problematic, which was also pointed out in Refs. [53–55].

To better understand the properties of the wave functions of the charmonium states, which are important to the decays, we plot the radial probability density as a function of the interquark distance r in Fig. 1. It is found that the spin-dependent potentials have notable corrections to the S- and triplet P-wave states. The spin-spin potential  $H_{SS}$ brings an obvious splitting to the wave functions between  $n^1S_0$  and  $n^3S_1$ , while the tensor potential  $H_T$  brings notable splittings to the wave functions between the triplet *P*-wave states. The spin-dependent potentials only give a tiny correction to wave functions of the higher triplet nD, nF, ... states. On the other hand, comparing the results from the linear potential model with those from the screened potential model, we find that for the low-lying 1S, 2S 1P, and 1D charmonium states the wave functions obtained from both of the models are less different. However, for the higher charmonium states nS  $(n \ge 3)$ ,  $nP, nD... (n \ge 2)$ , the wave functions obtained from these two models show a notable difference.

# III. ELECTROMAGNETIC TRANSITIONS WITH HIGHER MULTIPOLE CONTRIBUTIONS

Using the wave functions obtained from both the linear and screened potential models, we further study the EM transitions between charmonium states with higher multipole contributions. The EM decay properties not only are crucial for us to determine the quantum numbers of the newly observed charmonium states but also provide very useful references for our search for the missing charmonium states in experiments.

# A. Model

The quark-photon EM coupling at the tree level is described by

$$H_e = -\sum_j e_j \bar{\psi}_j \gamma^j_\mu A^\mu(\mathbf{k}, \mathbf{r}) \psi_j, \qquad (8)$$



FIG. 1. Predicted radial probability density  $|u(r)|^2$  for S-, P-, and D-wave charmonium states up to n = 3 shell. The dotted and solid curves stand for the results obtained from the linear and screened potential models, respectively.

where  $\psi_j$  stands for the *j*th quark field in a hadron. The photon has 3-momentum **k**, and the constituent quark  $\psi_j$  carries a charge  $e_j$ .

In this work, the wave functions are calculated nonrelativistically from the potential models. To match the nonrelativistic wave functions of hadrons, we should adopt the nonrelativistic form of Eq. (8) in the calculations. Including the effects of the binding potential between quarks [56], the nonrelativistic expansion of  $H_e$  may be written as [57–59]

$$h_e \simeq \sum_j \left[ e_j \mathbf{r}_j \cdot \boldsymbol{\epsilon} - \frac{e_j}{2m_j} \boldsymbol{\sigma}_j \cdot (\boldsymbol{\epsilon} \times \hat{\mathbf{k}}) \right] e^{-i\mathbf{k} \cdot \mathbf{r}_j}, \qquad (9)$$

where  $m_j$ ,  $\sigma_j$ , and  $\mathbf{r}_j$  stand for the constituent mass, Pauli spin vector, and coordinate for the *j*th quark, respectively. The vector  $\boldsymbol{\epsilon}$  is the polarization vector of the photon. It is found that the first and second terms in Eq. (9) are responsible for the electric and magnetic transitions, respectively. The second term  $\frac{e_j}{2m_j}\sigma_j \cdot (\boldsymbol{\epsilon} \times \hat{\mathbf{k}})$  in Eq. (9) is the same as that used in Ref. [7], while the first term in Eq. (9) differs from  $(1/m_j)\mathbf{p}_j \cdot \boldsymbol{\epsilon}$  used in Ref. [7] for the effects of the binding potential is included in the transition operator. This nonrelativistic EM transition operator has been widely applied to meson photoproduction reactions [59–71].

Finally, the standard helicity transition amplitude  $A_{\lambda}$  between the initial state  $|J\lambda\rangle$  and final state  $|J'\lambda'\rangle$  can be calculated by

$$\mathcal{A}_{\lambda} = -i\sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | h_e | J\lambda \rangle, \qquad (10)$$

where  $\omega_{\gamma}$  is the photon energy. It is easily found that the helicity amplitudes for the electric and magnetic operators are

$$\mathcal{A}_{\lambda}^{E} = -i\sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} e_{j} \mathbf{r}_{j} \cdot \boldsymbol{\epsilon} e^{-i\mathbf{k}\cdot\mathbf{r}_{j}} | J\lambda \rangle, \qquad (11)$$

$$\mathcal{A}^{M}_{\lambda} = +i\sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} \frac{e_{j}}{2m_{j}} \boldsymbol{\sigma}_{j} \cdot (\boldsymbol{\epsilon} \times \hat{\mathbf{k}}) e^{-i\mathbf{k}\cdot\mathbf{r}_{j}} | J\lambda \rangle.$$
(12)

In the initial-hadron-rest system for the radiative decay process, the momentum of the initial hadron is  $\mathbf{P}_i = 0$ , and that of the final hadron state is  $\mathbf{P}_f = -\mathbf{k}$ . Without loss of generality, we select the photon momentum along the *z* axial ( $\mathbf{k} = k\hat{\mathbf{z}}$ ) and take the polarization vector of the photon with the right-hand form, i.e.,  $\boldsymbol{\epsilon} = -\frac{1}{\sqrt{2}}(1, i, 0)$ , in our calculations. To easily work out the EM transition matrix elements, we use the multipole expansion of the plane wave,

$$e^{-i\mathbf{k}\cdot\mathbf{r}_{j}} = e^{-ikz_{j}}$$
  
=  $\sum_{l} \sqrt{4\pi(2l+1)}(-i)^{l}j_{l}(kr_{j})Y_{l0}(\Omega),$  (13)

where  $j_l(x)$  is the Bessel function and  $Y_{lm}(\Omega)$  are the wellknown spherical harmonics. Then, we obtain the matrix element for the electric multipole transitions with angular momentum l (El transitions) [72],

$$\begin{split} \mathcal{A}_{\lambda}^{El} &= \sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} (-i)^{l} B_{l} e_{j} j_{l+1}(kr_{j}) r_{j} Y_{l1} | J\lambda \rangle \\ &+ \sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} (-i)^{l} B_{l} e_{j} j_{l-1}(kr_{j}) r_{j} Y_{l1} | J\lambda \rangle, \quad (14) \end{split}$$

where  $B_l \equiv \sqrt{\frac{2\pi l(l+1)}{2l+1}}$ . We also obtain the matrix element from the magnetic part with angular momentum *l* (M*l* transitions),

$$\begin{aligned} \mathcal{A}_{\lambda}^{\mathrm{M}l} &= \sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} (-i)^{l} C_{l} \frac{e_{j}}{2m_{j}} j_{l-1}(kr_{j}) \sigma_{j}^{+} Y_{l-10} | J\lambda \rangle \\ &= \sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} (-i)^{l} C_{l} \frac{e_{j}}{2m_{j}} j_{l-1}(kr_{j}) \\ &\times [\sigma_{j}^{+} \otimes Y_{l-10}]_{1}^{l} | J\lambda \rangle \\ &+ \sqrt{\frac{\omega_{\gamma}}{2}} \langle J'\lambda' | \sum_{j} (-i)^{l} C_{l} \frac{e_{j}}{2m_{j}} j_{l-1}(kr_{j}) \\ &\times [\sigma_{j}^{+} \otimes Y_{l-10}]_{1}^{l-1} | J\lambda \rangle, \end{aligned}$$
(15)

where  $C_l \equiv i\sqrt{8\pi(2l-1)}$  and  $\sigma^+ = \frac{1}{2}(\sigma_x + i\sigma_y)$  is the spin shift operator. Obviously, the *El* transitions satisfy the parity selection rule,  $\pi_i \pi_f = (-1)^l$ , while the *Ml* transitions satisfy the parity selection rule,  $\pi_i \pi_f = (-1)^{l+1}$ , where  $\pi_i$ and  $\pi_f$  stand for the parities of the initial and final hadron states, respectively. Finally, using the parity selection rules, one can express the EM helicity amplitude  $\mathcal{A}$  with the matrix elements of EM multipole transitions in a unified form:

$$\mathcal{A}_{\lambda} = \sum_{l} \left\{ \frac{1 + (-1)^{\pi_{i}\pi_{f} + l}}{2} \mathcal{A}_{\lambda}^{\mathrm{E}l} + \frac{1 - (-1)^{\pi_{i}\pi_{f} + l}}{2} \mathcal{A}_{\lambda}^{\mathrm{M}l} \right\}.$$
(16)

Combining the parity selection rules, we easily know the possible EM multipole contributions to an EM transition considered in the present work, which are listed in Table IV.

It should be pointed out that the second term of Eq. (15) from the magnetic part is included in the electric

TABLE IV. Possible EM multipole contributions to an EM transition between two charmonium states.

Process	Multipole contribution
$n^3S_1 \leftrightarrow m^1S_0$	M1
$n^3 P_J \leftrightarrow m^3 S_1$	E1, M2
$n^1P_1 \leftrightarrow m^1S_0$	E1
$n^3 D_J \leftrightarrow m^3 P_J$	E1, E3, M2, M4
$n^1D_1 \leftrightarrow m^1P_1$	E1, E3
$n^3 P_J \leftrightarrow m^1 P_1$	M1, M3

part by the most general decomposition of the helicity amplitudes [73–75],

$$\mathcal{A}_{\lambda} = \sum_{k \ge 1} (-1)^{k+1} \sqrt{\frac{2k+1}{2J+1}} a_k \langle k-1; J'\lambda + 1 | J\lambda \rangle, \quad (17)$$

with  $a_k$  corresponding to the multipole amplitude of the EM tensor operators with a rank k. The second term of Eq. (15) is called an "extra" electric-multipole term,  $E_R$ , by Close *et al.* [76]. Specifically, for  ${}^{3}S_1 \leftrightarrow {}^{3}P_1$ ,

$$a_{1} = E_{1} + E_{R} = -\frac{\sqrt{2}}{2}(\mathcal{A}_{0} + \mathcal{A}_{-1}),$$
  
$$a_{2} = M_{2} = -\frac{\sqrt{2}}{2}(\mathcal{A}_{0} - \mathcal{A}_{-1});$$
 (18)

for  ${}^{3}P_2 \rightarrow {}^{3}S_1$ ,

$$a_{1} = E_{1} + E_{R} = \frac{\sqrt{10}}{2} (\sqrt{3}A_{-1} - A_{0}),$$
  
$$a_{2} = M_{2} = \frac{\sqrt{6}}{2} (\sqrt{3}A_{0} - A_{-1}); \qquad (19)$$

and for  ${}^3S_1 \rightarrow {}^3P_2$ ,

$$a_{1} = E_{1} + E_{R} = \frac{\sqrt{10}}{2} (\sqrt{3}A_{0} - A_{-1}),$$
  
$$a_{2} = M_{2} = \frac{\sqrt{6}}{2} (A_{0} - \sqrt{3}A_{-1}).$$
 (20)

Here,  $E_1$  is the leading electric-dipole term determined by Eq. (14), and  $M_2$  is the magnetic-quadrupole term related to the first term of Eq. (15). It should be mentioned that we have  $a_3 = 0$  for the above transitions in the present work.

Then, the partial decay widths of the EM transitions are given by

$$\Gamma = \frac{|\mathbf{k}|^2}{\pi} \frac{2}{2J_i + 1} \frac{M_f}{M_i} \sum_{\lambda} |\mathcal{A}_{\lambda}|^2$$
$$= \frac{|\mathbf{k}|^2}{\pi} \frac{2}{2J_i + 1} \frac{M_f}{M_i} \sum_{k} |a_k|^2, \qquad (21)$$

where  $J_i$  is the total angular momenta of the initial mesons and  $J_{fz}$  and  $J_{iz}$  are the components of the total angular momentum along the *z* axis of initial and final mesons, respectively. To take into account the relativistic effects, following the idea of Ref. [8], we introduce an overall relativistic phase space factor  $E_f/M_i$  in our predictions of the widths, which is usually not far from unity.  $M_f$  and  $E_f$ stand for the mass and total energy of the final charmonium state, respectively.  $M_i$  is the mass of the initial charmonium state.

Finally, we should mention that in the most general decomposition of the helicity amplitudes [73–75] the  $a_k$  is considered as the magnetic or electric multipole amplitude; thus, in the total decay width, the electric and magnetic multipole amplitudes cannot interfere with each other. However, the extra electric-multipole term comes from the magnetic part, and thus, in this sense, the electric-magnetic interference term appears in the total decay width [76–78].

# **B.** Results and discussions

# 1. Lighter states

First, we calculate the M1 transitions of the low-lying 1S, 2S, and 3S states. Our results compared with experimental data and other model predictions have been listed in Table V. Both our linear and screened potential model calculations obtain a compatible prediction. Our predictions are consistent with those of the NR and GI models [8]. It should be pointed out that our predictions together with those in the framework of the GI and NR potential models [8] give a very large partial width for the  $\psi(2S) \rightarrow \eta_c(1S)\gamma$ process, which is about an order of magnitude larger than the world average data from the PDG [6] and the prediction of  $\Gamma[\psi(2S) \rightarrow \eta_c(1S)\gamma] \simeq 0.4(8)$  keV from lattice QCD [20]. Although our prediction of  $\Gamma[J/\psi \rightarrow \eta_c(1S)\gamma]$  is obviously larger than the PDG average data [6], it is in agreement with the recent measurement  $\Gamma[J/\psi \rightarrow$  $\eta_c(1S)\gamma] \simeq 2.98 \pm 0.18^{+0.15}_{-0.33}$  keV at KEDR [79]. As a whole, strong model dependence exists in the predictions of the M1 transitions, thus more studies are needed in both theory and experiments.

Then, we calculate the E1 dominant radiative decays of the 1P and 2S states. Our results compared with experimental data and other model predictions have been listed in Table VI. Both our linear and screened potential model calculations obtain a compatible prediction because the wave functions and masses for the low-lying states from

TABLE V. Partial widths (keV) of the M1 radiative transitions for some low-lying *S*-wave charmonium states. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. For comparison, the predictions from the relativistic quark model [37] and NR and GI models [8] are listed in the table as well. The experimental average data are taken from the PDG [6].

		$E_{\gamma}$ (MeV) $\Gamma_{\rm M1}$ (keV)							$\Gamma_{\rm M1}~({\rm keV})$		
Initial state	Final state	[37]	NR [8]	GI [8]	Ours	[37]	NR [8]	GI [8]	LP	SP	Exp.
$J/\psi$	$\eta_c(1S)$	115	116	115	111	1.05	2.9	2.4	2.39	2.44	$1.58 \pm 0.37$
$\psi(2S)$	$\eta_c(2S)$	32	48	48	47	0.043	0.21	0.17	0.19	0.19	$0.21\pm0.15$
	$\eta_c(1S)$	639	639	638	635	0.95	4.6	9.6	8.08	7.80	$1.24 \pm 0.29$
$\eta_c(2S)$	$J/\psi$	514	501	501	502	1.53	7.9	5.6	2.64	2.29	
$\psi(3S)$	$\eta_c(3S)$		29	35	30/36		0.046	0.067	0.051	0.088	
	$\eta_c(2S)$		382	436	381		0.61	2.6	1.65	1.78	
	$\eta_c(1S)$		922	967	918		3.5	9.0	6.66	6.76	•••

these two models have fewer differences. Our predictions are in reasonable agreement with the data. The predictions from different models are consistent with each other in a magnitude, although there are differences more or less.

# 2. $\psi(3770)$ , X(3823) and the missing 1D states

The  $\psi(3770)$  resonance is primarily a  $\psi_1(1D)$  state with small admixtures of  $\psi(2S)$  [1]. It can decay into  $\chi_{cJ}(1P)\gamma$ . These decay processes are dominated by the E1 transition. The radiative decays of  $\psi(3770)$  are still not well understood. For example, the predictions of  $\Gamma[\psi(3770) \rightarrow \chi_{c0}(1P)\gamma]$  vary in a very large range (200–500) keV [45]. Considering  $\psi(3770)$  as a pure  $\psi_1(1D)$  state, we calculate the radiative decay widths of  $\Gamma[\psi(3770) \rightarrow \chi_{cJ}(1P)\gamma]$  with the wave functions obtained from the linear and screened potential models, respectively. Our results are listed in Table VII. From the table, we can see that both models give very similar predictions for the partial decay widths. Considering the leading E1 decays only, our predictions are in agreement with the world average data within their uncertainties [6]. However, including the magnetic part, the partial decay widths predicted by us are about a factor of 1.5 larger than the world average data [6] and the recent measurements from BESIII [45,80]. It is unclear whether these discrepancies are caused by our model limitations or come from the experimental uncertainties. It should be mentioned that, although some predictions from the models with a relativistic assumption [8,11] or a coupled-channel correction [30] seem to better agree quantitatively with the experimental data, the corrections of the magnetic part are not included in their calculations. To better understand the radiative decay properties of  $\psi(3770)$ , more studies are needed in both theory and experiments.

TABLE VI. Partial widths  $\Gamma$  (keV) and branching ratios Br for the radiative transitions (E1 dominant) between the low-lying charmonium states. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. For comparison, the predictions from the relativistic quark model [37], NR and GI models [8], and SNR model [11] are listed in the table as well. The experimental average data are taken from the PDG.  $\Gamma_{E1}$  and  $\Gamma_{EM}$  stand for the E1 and EM transition widths, respectively.

Initial	Final	E., (MeV)		$\Gamma_{\rm E1}~({\rm keV})$					$\Gamma_{\rm EM}$	(keV)		Br (%)	
state	state	Ours	[37]	NR/GI [8]	SNR <sub>0/1</sub> [11]	LP	SP	LP	SP	Exp.	LP	SP	Exp.
$\overline{\psi(2S)}$	$\chi_{c2}(1P)$	128	18.2	38/24	43/34	36	44	38	46	$25.2 \pm 2.9$	13.3	15.7	9.1 ± 0.3
,	$\chi_{c1}(1P)$	171	22.9	54/29	62/36	45	48	42	45	$25.5\pm2.8$	14.7	15.7	$9.6\pm0.3$
	$\chi_{c0}(1P)$	261	26.3	63/26	74/25	27	26	22	22	$26.3\pm2.6$	7.7	7.7	$10.0\pm0.3$
$\eta_c(2S)$	$h_c(1P)$	112	41	49/36	146/104	49	52	49	52		0.43	0.46	
$\chi_{c2}(1P)$	$J/\psi$	429	327	424/313	473/309	327	338	284	292	$371\pm34$	14.6	15.0	$19.2\pm0.7$
$\chi_{c1}(1P)$		390	265	314/239	354/244	269	278	306	319	$285\pm14$	34.8	36.3	$33.9\pm1.2$
$\chi_{c0}(1P)$		303	121	152/114	167/117	141	146	172	179	$133\pm8$	1.6	1.7	$1.3\pm0.1$
$h_c(1P)$	$\eta_c(1S)$	499	560	498/352	764/323	361	373	361	373	$357\pm280$	51.6	51.0	$51.0\pm6.0$
$\psi_1(1D)$	$\chi_{c2}(1P)$	215	6.9	4.9/3.3	5.8/4.6	5.4	5.7	7.1	8.1	<24.8	$4.8 \times 10^{-2}$	$5.3  imes 10^{-2}$	$< 9.0 \times 10^{-2}$
	$\chi_{c1}(1P)$	258	135	125/77	150/93	115	111	138	135	$81\pm27$	0.55	0.58	$0.29\pm0.06$
	$\chi_{c0}(1P)$	346	355	403/213	486/197	243	232	272	261	$202\pm42$	0.99	0.95	$0.73\pm0.09$
$\psi_2(1D)$	$\chi_{c2}(1P)$	258	59	64/66	70/55	79	82	91	96		13.3	14.1	
	$\chi_{c1}(1P)$	299	215	307/268	342/208	281	291	285	296	•••	41.9	43.5	•••

TABLE VII. Partial widths  $\Gamma$  (keV) and branching ratios Br for the radiative transitions (E1 dominant) of the higher *D*-wave states. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. For comparison, the predictions from the relativistic quark model [37], NR and GI models [8], and SNR model [11] are listed in the table as well.

Initial	Final		$E_{\gamma}$	(MeV)			$\Gamma_1$	E1 (keV)			$\Gamma_{\rm EM}$	(keV)	F	3r
state	state	[37]	NR/GI [8]	SNR [11]	LP/SP	[37]	NR/GI [8]	SNR <sub>0/1</sub> [11]	LP	SP	LP	SP	LP	SP
$\overline{\psi_3(1D)}$	$\chi_{c2}(1P)$	250	242/282	236	264/264	156	272/296	284/223	377	393	350	364	12%	12%
$\eta_{c2}(1D)$	$h_c(1P)$	275	264/307	260	284/284	245	339/344	575/375	362	376	362	376	72%	75%
$\psi_3(2D)$	$\chi_{c2}(1P)$		566/609		571/518		29/16		83	78	72	67	$4.9 \times 10^{-4}$	$4.5 \times 10^{-4}$
148 <sup>a</sup>	$\chi_{c2}(2P)$		190/231		238/181		239/272		457	256	427	243	$2.9 \times 10^{-3}$	$1.6 \times 10^{-3}$
$\psi_2(2D)$	$\chi_{c2}(1P)$		558/602		564/516		7.1/0.62		16	16	20	20	$1.7 \times 10^{-4}$	$2.2 \times 10^{-4}$
92 <sup>a</sup>	$\chi_{c1}(1P)$		597/640		603/554		26/23		64	64	68	68	$7.4 \times 10^{-4}$	$7.4 \times 10^{-4}$
	$\chi_{c2}(2P)$		182/223		231/178		52/65		101	57	115	64	$1.3 \times 10^{-3}$	$7.0 \times 10^{-4}$
	$\chi_{c1}(2P)$		226/247		222/204		298/225		220	186	223	188	$2.4 \times 10^{-3}$	$2.0 \times 10^{-3}$
$\psi_1(2D)$	$\chi_{c2}(1P)$		559/590		587		0.79/0.027		16	16	17	20	$2.3 \times 10^{-4}$	$2.7 \times 10^{-4}$
74 <sup>a</sup>	$\chi_{c1}(1P)$		598/628		625		14/3.4		25	42	37	63	$5.0 \times 10^{-4}$	$8.5 \times 10^{-4}$
	$\chi_{c0}(1P)$		677/707		704		27/35		120	149	150	189	$2.0 \times 10^{-3}$	$2.6 \times 10^{-3}$
	$\chi_{c2}(2P)$		183/210		256		5.9/6.3		18	21	24	29	$3.2 \times 10^{-4}$	$3.9 \times 10^{-4}$
	$\chi_{c1}(2P)$		227/234		281/281		168/114		253	280	309	347	$4.2 \times 10^{-3}$	$4.7 \times 10^{-3}$
	$\chi_{c0}(2P)$		296/269		312/329		483/191		299	321	332	360	$4.5 \times 10^{-3}$	$4.9 \times 10^{-3}$
$\eta_{c2}(2D)$	$h_c(1P)$		585/634		590/542		40/25		96	92	96	92	$1.3 \times 10^{-3}$	$1.2 \times 10^{-3}$
111 <sup>a</sup>	$h_c(2P)$		218/244	•••	256/203	•••	336/296	•••	438	271	438	271	$3.9 \times 10^{-3}$	$2.4 \times 10^{-3}$

<sup>a</sup>Predicted width (MeV) from Ref. [8].

Recently, X(3823) as a good candidate of  $\psi_2(1D)$  was observed by the Belle Collaboration in the  $B \rightarrow \chi_{c1}\gamma K$ decay with a statistical significance of 3.8 $\sigma$  [81]. Lately, this state was confirmed by the BESIII Collaboration in the process  $e^+e^- \rightarrow \pi^+\pi^-X(3823) \rightarrow \pi^+\pi^-\chi_{c1}\gamma$  with a statistical significance of 6.2 $\sigma$  [82]. Assigning X(3823) as the  $\psi_2(1D)$  state, we predict the radiative decay widths of  $\Gamma[X(3823) \rightarrow \chi_{cJ}(1P)\gamma]$ . Both the linear and screened potential models give quite similar predictions,

$$\Gamma[X(3823) \to \chi_{c1}(1P)\gamma] \simeq 300 \text{ keV}, \qquad (22)$$

$$\Gamma[X(3823) \to \chi_{c2}(1P)\gamma] \simeq 90 \text{ keV}.$$
 (23)

Our prediction of  $\Gamma[X(3823) \rightarrow \chi_{c1}(1P)\gamma]$  is close to the predictions in Refs. [8,11,37,83], while our prediction for  $\Gamma[X(3823) \rightarrow \chi_{c2}(1P)\gamma]$  is about a factor of  $1.4 \sim 1.8$  larger than the predictions in these works. Furthermore, our predicted partial width ratio,

$$\frac{\Gamma[X(3823) \to \chi_{c2}(1P)\gamma]}{\Gamma[X(3823) \to \chi_{c1}(1P)\gamma]} \simeq 30\%,$$
(24)

is consistent with the observations <42% [82]. The X(3823) state mainly decays into the  $\chi_{c1,2}(1P)\gamma$ ,  $J/\psi\pi\pi$ , and ggg channels. The predicted partial widths for the  $J/\psi\pi\pi$  and ggg channels are about (210 ± 110) and 80 keV, respectively [14]. Thus, the total width of the X(3823) might be  $\Gamma_{tot} \simeq 680 \pm 110$  keV, from which we obtain large branching ratios:

$$Br[X(3823) \to \chi_{c1}(1P)\gamma] \simeq 42\%,$$
 (25)

$$Br[X(3823) \to \chi_{c2}(1P)\gamma] \simeq 13\%.$$
 (26)

The large branching fraction  $Br[X(3823) \rightarrow \chi_{c1}(1P)\gamma]$  can explain why the X(3823) was first observed in the  $\chi_{c1}\gamma$  channel.

Another two 1*D*-wave states,  $\psi_3(1D)$  and  $\eta_{c2}(1D)$ , have not been observed in experiments. According to the theoretical predictions, their masses are very similar to that of  $\psi_2(1D)$ . If X(3823) corresponds to the  $\psi_2(1D)$  state indeed, the masses of the  $\psi_3(1D)$  and  $\eta_{c2}(1D)$  resonances should be around 3.82 GeV. For the singlet 1*D* state  $\eta_{c2}(1D)$ , its main radiative transition is  $\eta_{c2}(1D) \rightarrow$  $h_c(1P)\gamma$ . This process is governed by the E1 transition, and the effects from the E3 transition are negligibly small. Taking the mass of  $\eta_{c2}(1D)$  with M = 3820 MeV, with the wave functions calculated from the linear potential model, we predict that

$$\Gamma[\eta_{c2}(1D) \to h_c(1P)\gamma] \simeq 362 \text{ keV}, \qquad (27)$$

which is consistent with that of the screened potential model. Our results are close to the previous predictions in Refs. [8,11] (see Table VII). Combined with the predicted partial widths of the other two main decay modes gg and  $\eta_c \pi \pi$  [14], the total width of  $\eta_{c2}(1D)$  is estimated to be  $\Gamma_{tot} \approx 760$  keV. Then, we can obtain a large branching ratio:

$$\operatorname{Br}[\eta_{c2}(1D) \to h_c(1P)\gamma] \simeq 48\%.$$
(28)

Combining the measured branching ratios of  $Br[h_c(1P) \rightarrow \eta_c \gamma] \simeq 51\%$  and  $Br[\eta_c \rightarrow K\bar{K}\pi] \simeq 7.3\%$  [6], we obtain

$$\operatorname{Br}[\eta_{c2}(1D) \to h_c(1P)\gamma \to \eta_c\gamma\gamma \to K\bar{K}\pi\gamma\gamma] \simeq 1.8\%. \quad (29)$$

It should be mentioned that the  $\eta_{c2}(1D)$  state could be produced via the  $B \to \eta_{c2}(1D)K$  process as suggested in Refs. [13,84,85]. The expectations are to accumulate  $10^{10}$  $\Upsilon(4S) B\bar{B}$  events at Belle [2,9]. If the branching fraction Br $[B \to \eta_{c2}(1D)K]$  is  $\mathcal{O}(10^{-5})$  as predicted in Ref. [84], the  $10^{10} B\bar{B}$  events could let us observe  $\mathcal{O}(1000) \eta_{c2}(1D)$ events via the two-photon cascade  $\eta_{c2}(1D) \to h_c(1P)\gamma \to \eta_c\gamma\gamma$  in the  $\gamma\gamma K\bar{K}\pi$  final states.

While for the triplet 1*D* state  $\psi_3(1D)$ , its common radiative transition is  $\psi_3(1D) \rightarrow \chi_{c2}(1P)\gamma$ . Taking the mass of  $\psi_3(1D)$  with M = 3830 MeV, we calculate the partial decay widths  $\Gamma[\psi_3(1D) \rightarrow \chi_{c2}(1P)\gamma]$  with both the linear and screened potential models. Both of the models give a very similar result:

$$\Gamma[\psi_3(1D) \to \chi_{c2}(1P)\gamma] \simeq 350 \text{ keV.}$$
(30)

The magnitude of the partial decay width of  $\Gamma[\psi_3(1D) \rightarrow \chi_{c2}(1P)\gamma]$  predicted by us is compatible with that in Refs. [8,11,37]. Combining the predicted total width  $\Gamma_{\text{tot}} \simeq 3$  MeV for  $\psi_3(1D)$  [14], we estimate the branching ratio

$$\operatorname{Br}[\psi_3(1D) \to \chi_{c2}(1P)\gamma] \simeq 12\%. \tag{31}$$

The missing  $\psi_3(1D)$  might be produced via the  $B \rightarrow \psi_3(1D)K$  process at Belle [13,84,86] and reconstructed in the  $\chi_{c2}(1P)\gamma$  decay mode with  $\chi_{c2}(1P) \rightarrow J/\psi\gamma$  and  $J/\psi \rightarrow \mu^+\mu^-/e^+e^-$ . If the branching fraction Br[ $B \rightarrow \psi_3(1D)K$ ] is  $\mathcal{O}(10^{-5})$  [84,86], based on the  $10^{10} B\bar{B}$  data to be accumulated at Belle II, we expect that  $\mathcal{O}(100) \psi_3(1D)$  events could reconstructed in the  $\chi_{c2}(1P)\gamma$  channel with  $\chi_{c2}(1P) \rightarrow J/\psi\gamma$  and  $J/\psi \rightarrow \mu^+\mu^-/e^+e^-$ .

### 3. X(3872, 3915) and the 2P states

In the 2*P*-wave states, only  $\chi_{c2}(2P)$  has been established experimentally. This state was observed by both Belle [87] and *BABAR* [88] in the two-photon fusion process  $\gamma\gamma \rightarrow D\bar{D}$  with a mass  $M \approx 3927$  MeV and a narrow width  $\Gamma \approx 24$  MeV [6]. We analyze its radiative transitions to  $\psi(1D)\gamma$ ,  $J/\psi\gamma$ , and  $\psi(2S)\gamma$ . Both the linear and screened potentials give very similar predictions:

$$\Gamma[\chi_{c2}(2P) \to \psi(3770)\gamma] \simeq 0.4 \text{ keV}, \qquad (32)$$

$$\Gamma[\chi_{c2}(2P) \to \psi_2(1D)\gamma] \simeq 3.2 \text{ keV}, \qquad (33)$$

$$\Gamma[\chi_{c2}(2P) \to \psi_3(1D)\gamma] \simeq 20 \text{ keV.}$$
(34)

Our predictions are notably different from those of the NR potential model [8] (see Table VIII). With the measured width, we further predicted the branching ratios:

$$Br[\chi_{c2}(2P) \to \psi(3770)\gamma] \simeq 1.7 \times 10^{-5},$$
 (35)

$$Br[\chi_{c2}(2P) \to \psi_2(1D)\gamma] \simeq 1.3 \times 10^{-4},$$
 (36)

$$Br[\chi_{c2}(2P) \to \psi_3(1D)\gamma] \simeq 1.5 \times 10^{-3}.$$
 (37)

Combining these ratios with the decay properties of  $\psi(1D)$ and  $\chi_c(1P)$  states, we estimate the combined branching ratios for the decay chains  $\chi_{c2}(2P) \rightarrow \psi(1D)\gamma \rightarrow \chi_c(1P)\gamma\gamma \rightarrow J/\psi\gamma\gamma\gamma$ ; our results are listed in Table X. It is found that the most important decay chains involving  $\psi_2(1D)$  and  $\psi_3(1D)$  are  $\chi_{c2}(2P) \rightarrow \psi_2(1D)\gamma \rightarrow \chi_{c1}(1P)\gamma\gamma \rightarrow J/\psi\gamma\gamma\gamma$  (Br  $\approx 1.9 \times 10^{-5}$ ) and  $\chi_{c2}(2P) \rightarrow \psi_3(1D)\gamma \rightarrow \chi_{c2}(1P)\gamma\gamma \rightarrow J/\psi\gamma\gamma\gamma$  (Br  $\approx 1.4 \times 10^{-5}$ ). These decay chains might be difficult to observe at present because of the very small production cross section of  $\chi_{c2}(2P)$ .

The  $\chi_{c2}(2P)$  state has relatively larger radiative decay rates into  $J/\psi\gamma$  and  $\psi(2S)\gamma$ . With the linear potential model, we obtain

$$\Gamma[\chi_{c2}(2P) \to J/\psi\gamma] \simeq 93 \text{ keV},$$
 (38)

$$\Gamma[\chi_{c2}(2P) \to \psi(2S)\gamma] \simeq 135 \text{ keV},$$
 (39)

which are consistent with those of the screened potential model. Combined with the measured width, the branching ratios are predicted to be

$$Br[\chi_{c2}(2P) \to J/\psi\gamma] \simeq 3.9 \times 10^{-3}, \tag{40}$$

$$Br[\chi_{c2}(2P) \to \psi(2S)\gamma] \simeq 5.6 \times 10^{-3}.$$
 (41)

It might be a challenge to observe  $\chi_{c2}(2P)$  in the  $J/\psi\gamma$  and  $\psi(2S)\gamma$  channels with  $J/\psi/\psi(2S) \rightarrow \mu^+\mu^-$  at BESIII. For example, we produce  $\chi_{c2}(2P)$  via the process  $e^+e^- \rightarrow$  $\gamma \chi_{c2}(2P)$  at BESIII. Based on the cross section ~0.2 pb predicted in Ref. [89], we estimate that, even if BESIII can collect a 10 fb<sup>-1</sup> data sample above the open charm threshold, we only accumulate about 2000  $e^+e^- \rightarrow$  $\gamma \chi_{c2}(2P)$  events. Combining it with our predicted branching ratio  $Br[\chi_{c2}(2P) \rightarrow J/\psi\gamma \rightarrow \gamma\mu^+\mu^-] \sim \mathcal{O}(10^{-4})$ , we find there is less hope for observing the radiative decay modes of  $\chi_{c2}(2P)$  at BESIII. It should be mentioned that the decay chain  $\chi_{c2}(2P) \rightarrow J/\psi\gamma, \psi(2S)\gamma \rightarrow \gamma\mu^+\mu^-$  might be observed at Belle II or LHCb via the  $B \rightarrow \chi_{c2}(2P)X$ decay. The expectations are to accumulate  $10^{10} \Upsilon(4S) B\bar{B}$ events at Belle [2,9]. If the branching ratio of  $Br[B \to \chi_{c2}(2P)X] \sim 10^{-5}$ , we may observe  $\mathcal{O}(10)$  $\chi_{c2}(2P)$ 's in the decay chain  $\chi_{c2}(2P) \rightarrow J/\psi\gamma \rightarrow \gamma\mu^+\mu^-$ .

The  $\chi_{c1}(2P)$  state is still not established in experiments. According to the fine splitting between  $\chi_{c2}(2P)$  and  $\chi_{c1}(2P)$ , we estimate the mass of  $\chi_{c1}(2P)$  to be around M = 3900 MeV. With this mass, we calculate the transitions of  $\chi_{c1}(2P)$  into  $\psi(2S)\gamma$ ,  $J/\psi\gamma$ ,  $\psi_1(1D)\gamma$ , and  $\psi_2(1D)\gamma$ . Our results are listed in Table VIII. With the wave functions obtained from the linear potential model, it is found that the partial widths

TABLE VIII. Partial widths  $\Gamma$  (keV) and branching ratios Br for the radiative transitions (E1 dominant) of the higher 2*P* and 3*P* states. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. For comparison, the predictions from the NR and GI models [8] and SNR model [11] are listed in the table as well.

Initial	Final		$E_{\gamma}$ (MeV)			$\Gamma_{\rm E1}~({\rm keV})$			$\Gamma_{\rm EM}$	(keV)	E	Br
state	state	NR/GI [8]	SNR [11]	LP/SP	NR/GI [8]	SNR <sub>0/1</sub> [11]	LP	SP	LP	SP	LP	SP
$\chi_{c2}(2P)$	$\psi_3(1D)$	163/128		96/96	88/29		20	24	20	24	$8.3 \times 10^{-4}$	$1.0 \times 10^{-3}$
$24\pm6^{a}$	$\psi_2(1D)$	168/139		103	17/5.6		3.3	4.1	3.2	4.0	$1.3 \times 10^{-4}$	$1.7 \times 10^{-4}$
	$\psi_1(1D)$	197/204		146	1.9/1.0		0.47	0.62	0.36	0.46	$1.5 \times 10^{-5}$	$1.9 \times 10^{-5}$
	$\psi(2S)$	276/282	235	234	304/207	225/100	146	163	135	150	$5.6 \times 10^{-3}$	$6.3 \times 10^{-3}$
	$J/\psi$	779/784	744	742	81/53	101/109	118	119	93	93	$3.9 \times 10^{-3}$	$3.9 \times 10^{-3}$
$\chi_{c1}(2P)$	$\psi_2(1D)$	123/113		76/76	35/18		2.8	3.4	2.9	3.5	$1.8 \times 10^{-5}$	$2.1 \times 10^{-5}$
165 <sup>b</sup>	$\psi_1(1D)$	152/179		120/120	22/21		8.6	10.8	7.9	9.8	$4.9 \times 10^{-5}$	$5.9 \times 10^{-5}$
	$\psi(2S)$	232/258	182	208/208	183/183	103/60	129	145	139	155	$8.4 \times 10^{-4}$	$9.4 \times 10^{-4}$
	$J/\psi$	741/763	697	720/720	71/14	83/45	64	68	81	88	$4.9 \times 10^{-4}$	$5.3 \times 10^{-4}$
$\chi_{c0}(2P)$	$\psi_1(1D)$	81/143		90/69	13/51		21	12	20	12	$6.7  imes 10^{-4}$	$4.0 \times 10^{-4}$
30 <sup>b</sup>	$\psi(2S)$	162/223	152	179/159	64/135	61/44	108	89	121	99	$4.0 \times 10^{-3}$	$3.3 \times 10^{-3}$
	$J/\psi$	681/733	672	695/678	56/1.3	74/9.3	4.0	1.5	6.1	2.3	$2.0  imes 10^{-4}$	$7.7 \times 10^{-5}$
$h_c(2P)$	$\eta_{c2}(1D)$	133/117		100/100	60/27		25	25	25	25	$2.9  imes 10^{-4}$	$2.9  imes 10^{-4}$
87 <sup>b</sup>	$\eta_c(2S)$	285/305	261	252/252	280/218	309/108	160	176	160	176	$1.8 \times 10^{-3}$	$2.0 \times 10^{-3}$
	$\eta_c(1S)$	839/856	818	808/808	140/85	134/250	135	134	135	134	$1.6 \times 10^{-3}$	$1.6 \times 10^{-3}$
$\chi_{c2}(3P)$	$\psi_3(2D)$	147/118		136/98	148/51		116	64	121	66	$1.8 \times 10^{-3}$	$1.0 \times 10^{-3}$
66 <sup>b</sup>	$\psi_2(2D)$	156/127		143/101	31/10		18	10	18	10	$2.7 \times 10^{-4}$	$1.5 \times 10^{-4}$
	$\psi_1(2D)$	155/141		117/20	2.1/0.77		0.55	0.004	0.44	0.004	$6.7 \times 10^{-6}$	$6.0 \times 10^{-8}$
	$\psi_3(1D)$	481/461		453/364	0.049/6.8		15	10	17	12	$1.1 \times 10^{-4}$	$1.8 \times 10^{-4}$
	$\psi_2(1D)$	486/470		459/370	0.01/0.13		4.6	2.5	4.6	2.4	$7.0  imes 10^{-5}$	$3.6 \times 10^{-5}$
	$\psi_1(1D)$	512/530		495/411	0.00/0.00		1.9	1.0	1.5	0.79	$2.2 \times 10^{-5}$	$1.2 \times 10^{-5}$
	$\psi(3S)$	268/231		261/168	509/199		306	121	281	114	$4.3 \times 10^{-3}$	$1.7 \times 10^{-3}$
	$\psi(2S)$	585/602		574/492	55/30		116	90	97	76	$1.5 \times 10^{-3}$	$1.2 \times 10^{-3}$
	$J/\psi$	1048/1063		1042/967	34/19		83	69	61	51	$9.2 \times 10^{-4}$	$7.7 \times 10^{-4}$
$\chi_{c1}(3P)$	$\psi_2(2D)$	112/108		117/82	58/35		22	11	23	11	$5.9  imes 10^{-4}$	$2.8  imes 10^{-4}$
39 <sup>b</sup>	$\psi_1(2D)$	111/121	•••	92/1	19/15		8.6	0	8.1	0	$2.1 \times 10^{-4}$	0
	$\psi_2(1D)$	445/452		436/353	0.035/4.6		0.13	0.09	0.12	0.09	$3.1 \times 10^{-6}$	$2.3 \times 10^{-6}$
	$\psi_1(1D)$	472/512		476/394	0.014/0.39		4.4	2.7	3.2	2.0	$6.1\times10^{-5}$	$4.1 \times 10^{-5}$
	$\psi(3S)$	225/212		237/149	303/181		305	111	331	117	$8.5 \times 10^{-3}$	$3.0 \times 10^{-3}$
	$\psi(2S)$	545/585		556/475	45/8.9		78	63	94	74	$2.4 \times 10^{-3}$	$1.9 \times 10^{-3}$
	$J/\psi$	1013/1048		1023/952	31/2.2		36	33	50	45	$1.3 \times 10^{-3}$	$1.2 \times 10^{-3}$
$\chi_{c0}(3P)$	$\psi_1(2D)$	43/97		39/45	4.4/35		3.8	9.3	3.8	9.1	$7.5 \times 10^{-5}$	$1.8 \times 10^{-4}$
51 <sup>b</sup>	$\psi_1(1D)$	410/490		427/352	0.037/9.7		0.31	0.44	0.27	0.39	$5.3 \times 10^{-6}$	$7.6 \times 10^{-6}$
	$\psi(3S)$	159/188		186/105	109/145		214	56	241	61	$4.7 \times 10^{-3}$	$1.2 \times 10^{-3}$
	$\psi(2S)$	484/563		509/434	32/0.045		13	6.9	17	9.1	$3.3 \times 10^{-4}$	$1.8 \times 10^{-4}$
	$J/\psi$	960/1029		981/916	27/1.5		0.14	0.08	0.24	0.13	$4.7 \times 10^{-6}$	$2.5 \times 10^{-6}$
$h_c(3P)$	$\eta_{c2}(2D)$	119/109		120/84	99/48		93	47	93	47	$1.2 \times 10^{-3}$	$6.3 \times 10^{-4}$
75 <sup>°</sup>	$\eta_{c2}(1D)$	453/454		453/370	0.16/5.7		15	8.7	15	8.7	$2.0  imes 10^{-4}$	$1.2 \times 10^{-4}$
	$\eta_c(3S)$	229/246		238/185	276/208		237	146	237	146	$3.2 \times 10^{-3}$	$1.9 \times 10^{-3}$
	$\eta_c(2S)$	593/627		602/517	75/43		124	96	124	96	$1.7 \times 10^{-3}$	$1.3 \times 10^{-3}$
	$\eta_c(1S)$	1103/1131		1104/1035	72/38		90	77	90	77	$1.2 \times 10^{-4}$	$1.0 \times 10^{-3}$

<sup>a</sup>Width (MeV) from the PDG [6].

<sup>b</sup>Predicted width (MeV) from Ref. [8].

TABLE IX. Partial widths  $\Gamma$  (keV) and branching ratios Br for the radiative transitions of the higher *S*-wave states. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively. For comparison, the predictions from the NR and GI models [8] are listed in the table as well.

			$E_{\gamma}$ (Me	V)			Γ <sub>E1</sub> (k	eV)		$\Gamma_{\rm EM}$ (	(keV)	E	Br
Initial state	Final state	NR [8]	GI [8]	LP	SP	NR [8]	GI [8]	LP	SP	LP	SP	LP	SP
$\psi(3S)$	$\chi_{c2}(2P)$	67	119	111	111	14	48	65	79	67	82	$8.4 \times 10^{-4}$	$1.0 \times 10^{-3}$
$80\pm10^{\mathrm{a}}$	$\chi_{c1}(2P)$	113	145	138	138	39	43	58	71	55	67	$6.9 \times 10^{-4}$	$8.4 \times 10^{-4}$
	$\chi_{c0}(2P)$	184	180	167	187	54	22	21	31	19	27	$2.4 \times 10^{-4}$	$3.4 \times 10^{-4}$
	$\chi_{c2}(1P)$	455	508	455	455	0.7	13	0.21	2.1	0.25	2.5	$3.1 \times 10^{-6}$	$3.1 \times 10^{-5}$
	$\chi_{c1}(1P)$	494	547	494	494	0.53	0.85	4.8	8.0	4.0	6.7	$5.0 \times 10^{-5}$	$8.4 \times 10^{-5}$
	$\chi_{c0}(1P)$	577	628	577	577	0.27	0.63	9.1	10.6	5.9	6.7	$7.4 \times 10^{-5}$	$8.4 \times 10^{-5}$
$\eta_c(3S)$	$h_c(2P)$	108	108	108	108	105	64	104	128	104	128	$1.3 \times 10^{-3}$	$1.6 \times 10^{-3}$
80 <sup>b</sup>	$h_c(1P)$	485	511	456	456	9.1	28	0.045	1.4	0.045	1.4	$5.6 \times 10^{-7}$	$1.8 \times 10^{-5}$
$\psi(4S)$	$\chi_{c2}(1P)$	775	804	773	664	0.61	5.2	0.13	0.66	0.17	0.84	$2.2 \times 10^{-6}$	$1.1 \times 10^{-5}$
78 <sup>b</sup>	$\chi_{c1}(1P)$	811	841	809	701	0.41	0.53	3.8	3.9	2.9	3.0	$3.7 \times 10^{-5}$	$3.8 \times 10^{-5}$
	$\chi_{c0}(1P)$	887	915	884	778	0.18	0.13	7.5	6.2	3.7	2.7	$4.7 \times 10^{-5}$	$4.2 \times 10^{-5}$
	$\chi_{c2}(2P)$	421	446	458	339	0.62	15	11	4.7	13	5.3	$1.7 \times 10^{-4}$	$6.8 \times 10^{-5}$
	$\chi_{c1}(2P)$	423	469	482	364	0.49	0.92	24	12	20	11	$2.6  imes 10^{-4}$	$1.4 \times 10^{-4}$
	$\chi_{c0}(2P)$	527	502	510	411	0.24	0.39	17	12	12	8.7	$1.5  imes 10^{-4}$	$1.1 \times 10^{-4}$
	$\chi_{c2}(3P)$	97	112	101	69	68	66	80	39	82	40	$1.1 \times 10^{-3}$	$5.1 \times 10^{-4}$
	$\chi_{c1}(3P)$	142	131	126	88	126	54	74	38	71	37	$9.1  imes 10^{-4}$	$4.7 \times 10^{-4}$
	$\chi_{c0}(3P)$	208	155	178	133	0.003	25	40	23	36	21	$4.6  imes 10^{-4}$	$2.7 \times 10^{-4}$
$\eta_c(4S)$	$h_c(1P)$	782	808	778	675	5.2	9.6	0.29	0.63	0.29	0.63	$4.8 \times 10^{-6}$	$1.0 \times 10^{-5}$
61 <sup>b</sup>	$h_c(2P)$	427	444	461	348	10.1	31.3	20	7.9	20	7.9	$3.3 \times 10^{-4}$	$1.3 \times 10^{-4}$
	$h_c(3P)$	104	106	142	70	159	101	102	70	102	70	$1.7  imes 10^{-3}$	$1.1 \times 10^{-3}$

<sup>a</sup>Width (MeV) from the PDG [6].

<sup>b</sup>Predicted width (MeV) from Ref. [8].

$$\Gamma[\chi_{c1}(2P) \to J/\psi\gamma] \simeq 81 \text{ keV},$$
 (42)

$$\Gamma[\chi_{c1}(2P) \to \psi(2S)\gamma] \simeq 139 \text{ keV}$$
 (43)

are slightly smaller than those from the screened potential model. Combined with the predicted width in Ref. [8], the branching ratios might be

$$Br[\chi_{c1}(2P) \to J/\psi\gamma] \simeq 4.9 \times 10^{-4}, \qquad (44)$$

$$Br[\chi_{c1}(2P) \to \psi(2S)\gamma] \simeq 8.4 \times 10^{-4}.$$
 (45)

The partial widths of  $\Gamma[\chi_{c1}(2P) \rightarrow \psi_{1,2}(1D)\gamma]$  are about several keV, and their branching ratios are  $\mathcal{O}(10^{-5})$ .

The X(3872) resonance has the same quantum numbers as  $\chi_{c1}(2P)$  (i.e.,  $J^{PC} = 1^{++}$ ) and a similar mass to the predicted value of  $\chi_{c1}(2P)$ . However, its exotic properties cannot be well understood with a pure  $\chi_{c1}(2P)$  state [4,90]. To understand the nature of X(3872), measurements of the radiative decays of X(3872) have been carried out by the *BABAR* [91], Belle [92], and LHCb [93] collaborations, respectively. Obvious evidence of  $X(3872) \rightarrow J/\psi\gamma$  was observed by these collaborations. Furthermore, the *BABAR* and LHCb collaborations also observed evidence of  $X(3872) \rightarrow \psi(2S)\gamma$ . The branching fraction ratio

$$R_{\psi'\gamma/\psi\gamma}^{\exp} = \frac{\Gamma[X(3872) \to \psi(2S)\gamma]}{\Gamma[X(3872) \to J/\psi\gamma]} \simeq 3.4 \pm 1.4, \quad (46)$$

obtained by the *BABAR* Collaboration [91] is consistent with the recent measurement  $R_{\psi'\gamma/\psi\gamma}^{exp} = 2.46 \pm 0.93$  of the LHCb Collaboration [93].

Considering X(3872) as a pure  $\chi_{c1}(2P)$  state, we calculate the radiative decays  $X(3872) \rightarrow J/\psi\gamma, \psi(2S)\gamma$ . With the linear potential model, we predict that

$$\Gamma[X(3872) \rightarrow J/\psi\gamma) \simeq 72 \text{ keV},$$
 (47)

$$\Gamma[X(3872) \to \psi(2S)\gamma] \simeq 94 \text{ keV.}$$
(48)

With these predicted partial widths, we can easily obtain the ratio

$$R_{\psi'\gamma/\psi\gamma}^{\text{th}} = \frac{\Gamma[X(3872) \to \psi(2S)\gamma]}{\Gamma[X(3872) \to J/\psi\gamma]} \simeq 1.3, \qquad (49)$$

which is slightly smaller than the lower limit of the measurements from *BABAR* [91] and LHCb [93]. Our predictions from the screened potential model are consistent with those from the linear potential model. Thus, from

TABLE X. Three-photon decay chains of  $2^{3}P_{2}$ . The branching fractions are  $Br_{1} = Br[2^{3}P_{2} \rightarrow 1^{3}D_{J}\gamma]$ ,  $Br_{2} = Br[1^{3}D_{J} \rightarrow 1^{3}P_{J}\gamma]$ ,  $Br_{3} = Br[1^{3}P_{J} \rightarrow J/\psi\gamma]$ , and  $Br = Br_{1} \times Br_{2} \times Br_{3}$ . The branching fractions are predicted with the linear potential model.

Decay chain	$Br_1$	Br <sub>2</sub>	Br <sub>3</sub>	Br
$\frac{1}{2^3 P_2 \rightarrow 1^3 D_1 \rightarrow 1^3 P_0 \rightarrow J/\psi}$	$1.5 \times 10^{-5}$	$9.9 \times 10^{-3}$	1.6%	$2.4 \times 10^{-9}$
$2^{3}P_{2} \rightarrow 1^{3}D_{1} \rightarrow 1^{3}P_{1} \rightarrow J/\psi$	$1.5 \times 10^{-5}$	$5.5 \times 10^{-3}$	34.8%	$2.9 \times 10^{-8}$
$2^{3}P_{2} \rightarrow 1^{3}D_{1} \rightarrow 1^{3}P_{2} \rightarrow J/\psi$	$1.5 \times 10^{-5}$	$4.8 \times 10^{-4}$	14.6%	$1.0 \times 10^{-8}$
$2^{3}P_{2} \rightarrow 1^{3}D_{2} \rightarrow 1^{3}P_{1} \rightarrow J/\psi$	$1.3 \times 10^{-4}$	42%	34.8%	$1.9 \times 10^{-5}$
$2^{3}P_{2} \rightarrow 1^{3}D_{2} \rightarrow 1^{3}P_{2} \rightarrow J/\psi$	$1.3 \times 10^{-4}$	13%	14.6%	$2.5 \times 10^{-6}$
$2^{3}P_{2} \rightarrow 1^{3}D_{3} \rightarrow 1^{3}P_{2} \rightarrow J/\psi$	$8.3 \times 10^{-4}$	12%	14.6%	$1.4 \times 10^{-5}$

the view of branching fraction ratio  $R_{\psi'\gamma/\psi\gamma}$ , we cannot exclude the *X*(3872) as a candidate of  $\chi_{c1}(2P)$ .

On the other hand, if X(3872) corresponds to  $\chi_{c1}(2P)$ , with the measured width (about several MeV) [6], we can estimate

$$Br[X(3872) \to J/\psi\gamma] \sim \mathcal{O}(10^{-2}),$$
 (50)

$$Br[X(3872) \to \psi(2S)\gamma] \sim \mathcal{O}(10^{-2}). \tag{51}$$

Combining the branching ratio  $\operatorname{Br}[B \to \chi_{c1}(2P)K] \sim O(10^{-4})$  predicted in Ref. [94], we can further estimate that  $\operatorname{Br}[B \to \chi_{c1}(2P)K] \times \operatorname{Br}[X(3872) \to J/\psi\gamma/\psi(2S)\gamma] \sim O(10^{-6})$ , which is also consistent with the Belle measurements [92].

The  $\chi_{c0}(2P)$  state is still not well established, although X(3915) was recommended as the  $\chi_{c0}(2P)$  state in Ref. [95] and also assigned as the  $\chi_{c0}(2P)$  state by the PDG recently [6]. Assigning X(3915) as the  $\chi_{c0}(2P)$  state will face several serious problems [53,54]. Recently, Zhou *et al.* carried out a combined amplitude analysis of the  $\gamma\gamma \rightarrow D\bar{D}, \omega J/\psi$  data [55]. They demonstrated that X(3915) and X(3930) can be regarded as the same state with  $J^{PC} = 2^{++}$  (i.e.,  $\chi_{c2}(2P)$ ). With the screened and linear potential models, our predicted masses for the  $\chi_{c0}(2P)$  state are ~3848 and ~3869 MeV, respectively, which are consistent with the previous predictions in Refs. [8,11] and the mass extracted by Guo and Meissner by refitting the *BABAR* and

Belle data of  $\gamma\gamma \rightarrow D\bar{D}$  separately [53]. The  $\chi_{c0}(2P)$  state can decay via the radiative transitions  $\chi_{c0}(2P) \rightarrow \psi(3770)\gamma, \psi(2S)\gamma, J/\psi\gamma$ . With the wave functions obtained from both the screened and linear potential models, we calculate the decay rates of these radiative transitions. Our results are listed in Table VIII. From the table, it is found that both of the models give similar predictions. The partial width for  $\chi_{c0}(2P) \rightarrow \psi(2S)\gamma$  is

$$\Gamma[\chi_{c0}(2P) \to \psi(2S)\gamma] \simeq 110 \pm 10 \text{ keV.}$$
 (52)

The  $\chi_{c0}(2P)$  might be very broad with a width of ~200 MeV extracted from experimental data [53], which is about an order of magnitude larger than that predicted in Ref. [8]. With the broad width, the branching ratio is predicted to be

$$Br[\chi_{c0}(2P) \to \psi(2S)\gamma] \simeq 5.5 \times 10^{-4}.$$
 (53)

It should be mentioned that there is less chance of producing the  $\chi_{c0}(2P)$  state via the radiative decay chains  $\psi(4040, 4160, 4415) \rightarrow \chi_{c0}(2P)\gamma\chi_{c0}(2P) \rightarrow \psi(1S, 2S)\gamma\gamma \rightarrow \gamma\gamma\mu^+\mu^-$  (see Tables XI,XII,XIII).

There is no information on  $h_c(2P)$  from experiments. According to our predictions, the mass splitting between  $\chi_{c2}(2P)$  and  $h_c(2P)$  is about  $M_{\chi_{c2}(2P)} - M_{h_c(2P)} =$  $(26 \pm 4)$  MeV. Thus, the mass of  $h_c(2P)$  is most likely  $M_{h_c(2P)} \approx 3900$  MeV. The typical radiative decay channels of  $h_c(2P)$  are  $\eta_c(1S, 2S)\gamma$  and  $\eta_{c2}(1D)\gamma$ . With the wave

TABLE XI. Two-photon decay chains of  $3^{3}S_{1}$ . The branching fractions are  $Br_{1} = Br[3^{3}S_{1} \rightarrow 2^{3}P_{J}\gamma]$ ,  $Br_{2} = Br[2^{3}P_{J} \rightarrow 2^{3}S_{1}\gamma, J/\psi\gamma]$ ,  $Br_{3} = Br[2^{3}S_{1}, J/\psi \rightarrow \mu^{+}\mu^{-}]$  (obtained from PDG [6]), and  $Br = Br_{1} \times Br_{2} \times Br_{3}$ . The theoretical branching fractions are predicted with the linear potential model. The estimated events are based on producing  $5 \times 10^{7} \psi(3S)$  events at BESIII in the coming years as described in the text.

Decay chain	$Br_1(10^{-4})$	$Br_2(10^{-4})$	Br <sub>3</sub> (%)	$Br(10^{-8})$	Events
$\overline{3^3S_1 \rightarrow 2^3P_2 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-}$	8.4	56	0.79	3.7	2
$3^3S_1 \rightarrow 2^3P_1 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	6.9	8.4	0.79	0.46	0.2
$3^3S_1 \rightarrow 2^3P_0 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	2.4	40	0.79	0.76	0.4
$3^3S_1 \rightarrow 2^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	8.4	39	5.9	19	10
$3^3S_1 \rightarrow 2^3P_1 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	6.9	4.9	5.9	2.0	1
$3^3S_1 \rightarrow 2^3P_0 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	2.4	2.0	5.9	0.28	0.1

TABLE XII. Two-photon decay chains of 2*D*-wave states. The branching fractions are  $Br_1 = Br[2D \rightarrow nP\gamma]$ ,  $Br_2 = Br[nP \rightarrow 2^3S_1\gamma, 1S\gamma]$ ,  $Br_3 = Br[2^3S_1, J/\psi \rightarrow \mu^+\mu^-]$  (obtained from PDG [6]), or  $Br_3 = Br[\eta_c(1S) \rightarrow K\bar{K}\pi]$ (obtained from PDG [6]), and  $Br = Br_1 \times Br_2 \times Br_3$ . The theoretical branching fractions are predicted with the linear potential model. The estimated events are based on producing of  $2.4 \times 10^7 \psi(4160)$ 's at BESIII as described in the text.

Decay chain	$Br_1(10^{-3})$	$Br_2(10^{-4})$	Br <sub>3</sub> (%) [6]	$Br(10^{-7})$	Events
$\frac{1}{2^3D_1 \rightarrow 2^3P_2 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-}$	0.32	56	0.79	0.14	0.3
$2^{3}D_{1} \rightarrow 2^{3}P_{1} \rightarrow 2^{3}S_{1} \rightarrow \mu^{+}\mu^{-}$	4.2	8.4	0.79	0.28	0.6
$2^3D_1 \rightarrow 2^3P_0 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	4.5	40	0.79	1.4	3
$2^3D_1 \rightarrow 2^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	0.32	39	5.9	0.74	2
$2^3 D_1 \rightarrow 2^3 P_1 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	4.2	4.9	5.9	1.2	3
$2^3 D_1 \rightarrow 2^3 P_0 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	4.5	2.0	5.9	0.53	1
$2^3D_1 \rightarrow 1^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	0.23	1460	5.9	19.8	47
$2^{3}D_{1} \rightarrow 1^{3}P_{1} \rightarrow J/\psi \rightarrow \mu^{+}\mu^{-}$	0.50	3480	5.9	102	244
$2^3D_1 \rightarrow 1^3P_0 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	2.0	160	5.9	18.8	45
$2^3D_2 \rightarrow 2^3P_2 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	1.3	56	0.79	0.57	
$2^{3}D_{2}^{-} \rightarrow 2^{3}P_{1}^{-} \rightarrow 2^{3}S_{1}^{-} \rightarrow \mu^{+}\mu^{-}$	2.4	8.4	0.79	0.16	
$2^3 D_2 \rightarrow 2^3 P_2 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	1.3	39	5.9	3.0	
$2^3 D_2 \rightarrow 2^3 P_1 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	2.4	4.9	5.9	6.9	
$2^3 D_2 \rightarrow 1^3 P_2 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	0.17	1460	5.9	14.6	
$2^3 D_2 \rightarrow 1^3 P_1 \rightarrow J/\psi \rightarrow \mu^+ \mu^-$	0.74	3480	5.9	152	
$2^3 D_3 \rightarrow 2^3 P_2 \rightarrow 2^3 S_1 \rightarrow \mu^+ \mu^-$	2.9	56	0.79	1.3	
$2^3D_3 \rightarrow 2^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	2.9	39	5.9	6.7	
$2^3D_3 \rightarrow 1^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	0.49	1460	5.9	42	
$2^1 D_2 \rightarrow 2^1 P_1 \rightarrow \eta_c(1S) \rightarrow K \bar{K} \pi$	3.9	16	7.3	4.6	
$2^1 D_2 \to 1^1 P_1 \to \eta_c(1S) \to K\bar{K}\pi$	1.3	5100	7.3	483	

functions obtained from the linear and screened potentials, we further calculate these radiative decays. It is found that the radiative transition rates of the  $h_c(2P) \rightarrow \eta_{c2}(1D)\gamma$ ,  $\eta_c(1S)\gamma$ , and  $\eta_c(2S)\gamma$  channels are fairly large. Both the linear and screened potential models give very similar predictions:

$$\Gamma(h_c(2P) \to \eta_c(1S)\gamma) \simeq 135 \text{ keV},$$
 (54)

$$\Gamma(h_c(2P) \to \eta_c(2S)\gamma) \simeq 160 \text{ keV},$$
 (55)

$$\Gamma(h_c(2P) \to \eta_{c2}(1D)\gamma) \simeq 25 \text{ keV}.$$
 (56)

The rather sizeable partial widths for  $h_c(2P) \rightarrow \eta_c(1S, 2S)\gamma$  are also obtained in the previous potential model calculations [8,11,18]. Combined with the theoretical width  $\Gamma \simeq 87$  MeV from Ref. [8], the branching ratios are predicted to be

$$Br[h_c(2P) \to \eta_c(1S)\gamma] \simeq 1.6 \times 10^{-3},$$
 (57)

$$Br[h_c(2P) \to \eta_c(2S)\gamma] \simeq 1.8 \times 10^{-3},$$
 (58)

$$Br[h_c(2P) \to \eta_{c2}(1D)\gamma] \simeq 2.8 \times 10^{-4}.$$
 (59)

The missing  $h_c(2P)$  state might be produced via the  $B \rightarrow h_c(2P)K$  process and reconstructed in the  $\eta_c(1S, 2S)\gamma$ 

decay modes with  $\eta_c(1S, 2S) \rightarrow K\bar{K}\pi$  at Belle II and LHCb.

#### 4. $\psi(4040)$ and the missing $\eta_c(3S)$ state

The  $\psi(4040)$  resonance is commonly identified with the  $\psi(3S)$  state [1]. This state can decay into  $\chi_{cJ}(1P)\gamma$  and  $\chi_{cI}(2P)\gamma$  via the radiative transitions. We have calculated these processes with both the linear and screened potential model. Our results have been listed in Table IX. In our calculations, we find that the radiative transition rates of  $\psi(4040) \rightarrow \chi_{cJ}(1P)\gamma$  are relatively weak. Using the PDG value for the total width  $\Gamma \simeq 80$  MeV [6], we obtain the branching ratios  $Br[\psi(4040) \rightarrow \chi_{cJ}(1P)\gamma] \sim \mathcal{O}(10^{-5}),$ which are consistent with the measurements  $Br[\psi(4040) \rightarrow$  $\chi_{c1,2}(1P)\gamma < 2\%$  [6]. Interestingly, it is found that the radiative transition rates of  $\psi(4040) \rightarrow \chi_{cJ}(2P)\gamma$  are rather sizeable. The decay rates into the  $\chi_{cJ}(2P)\gamma$  channels are about 1 order of magnitude larger than those into the  $\chi_{cI}(1P)\gamma$  channels. With the screened potential model, we obtain that

$$\Gamma[\psi(4040) \to \chi_{c2}(2P)\gamma] \simeq 82 \text{ keV}, \qquad (60)$$

$$\Gamma[\psi(4040) \to \chi_{c1}(2P)\gamma] \simeq 67 \text{ keV}, \tag{61}$$

$$\Gamma[\psi(4040) \to \chi_{c0}(2P)\gamma] \simeq 27 \text{ keV}, \tag{62}$$

TABLE XIII. Two-photon decay chains of  $4^{3}S_{1}$ . The branching fractions are  $Br_{1} = Br[4^{3}S_{1} \rightarrow 2^{3}P_{J}\gamma, 3^{3}P_{J}\gamma]$ ,  $Br_{2} = Br[n^{3}P_{J} \rightarrow m^{3}S_{1}\gamma]$ ,  $Br_{3} = Br[2^{3}S_{1}, J/\psi \rightarrow \mu^{+}\mu^{-}]$  (obtained from PDG [6]), and  $Br = Br_{1} \times Br_{2} \times Br_{3}$ . The theoretical branching fractions are predicted with the linear potential model. The estimated events are based on producing  $2 \times 10^{7} \psi(4S)$ 's at BESIII in the coming years as described in the text.

Decay chain	$Br_1(10^{-4})$	$Br_2(10^{-4})$	Br <sub>3</sub> (%) [6]	$Br(10^{-9})$	Events
$\overline{4^3S_1 \rightarrow 2^3P_2 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-}$	1.7	56	0.79	7.5	0.15
$4^3S_1 \rightarrow 2^3P_1 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	2.6	8.4	0.79	1.7	0.03
$4^3S_1 \rightarrow 2^3P_0 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	1.5	40	0.79	4.7	0.09
$4^3S_1 \rightarrow 2^3P_2 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	1.7	39	5.9	39	0.78
$4^3S_1 \rightarrow 2^3P_1 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	2.6	4.9	5.9	7.5	0.15
$4^3S_1 \rightarrow 2^3P_0 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	1.5	2.0	5.9	1.8	0.03
$4^3S_1 \rightarrow 3^3P_2 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	11	15	0.79	13	0.26
$4^3S_1 \rightarrow 3^3P_1 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	9.1	24	0.79	17	0.34
$4^3S_1 \rightarrow 3^3P_0 \rightarrow 2^3S_1 \rightarrow \mu^+\mu^-$	4.6	33	0.79	12	0.24
$4^{3}S_{1} \rightarrow 3^{3}P_{2} \rightarrow J/\psi \rightarrow \mu^{+}\mu^{-}$	1.7	9.2	5.9	9.2	0.18
$4^3S_1 \rightarrow 3^3P_1 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	2.6	13	5.9	20	0.40
$4^3S_1 \rightarrow 3^3P_0 \rightarrow J/\psi \rightarrow \mu^+\mu^-$	1.5	0.047	5.9	0.041	0

which are about 15% larger than our linear potential model predictions. Relatively large partial decay widths for  $\psi(4040) \rightarrow \chi_{cJ}(2P)\gamma$  were also found in the previous studies [8]. With the measured width  $\Gamma \simeq 80$  MeV from the PDG [6], we estimate the branching ratios

$$Br[\psi(4040) \to \chi_{c2}(2P)\gamma] \simeq 1.0 \times 10^{-3},$$
 (63)

$$Br[\psi(4040) \to \chi_{c1}(2P)\gamma] \simeq 0.8 \times 10^{-3}, \qquad (64)$$

$$Br[\psi(4040) \to \chi_{c0}(2P)\gamma] \simeq 3.3 \times 10^{-4}.$$
 (65)

It should be pointed out that BESIII plans to collect  $5 \sim 10 \text{ fb}^{-1} \psi(4040)$  in the coming years [96]. Using the cross section of ~10 nb based on BES and CLEO measurements [97–99], we expect to accumulate  $(0.5 \sim 1.0) \times 10^8 \psi(4040)$  events. The  $\psi(4040)$  might provide us a source to produce  $\chi_{cJ}(2P)$  states via the radiative transitions. Thus, we further estimate the number of events of the two-photon cascades involving the  $\chi_{cJ}(2P)$  states. The results are listed in Table XI. From the table, we can see that about  $\mathcal{O}(10) \chi_{c2}(2P)$  events should be observed at BESIII via the radiative transition chain  $\psi(4040) \rightarrow \chi_{c2}(2P)\gamma \rightarrow J/\psi\gamma\gamma \rightarrow \gamma\gamma\mu^+\mu^-$ .

The  $\eta_c(3S)$  state is not established in experiments. According to the model predictions, the hyperfine splitting between  $3^3S_1$  and  $3^1S_0$  is about 30 MeV (see Table III). Thus, the mass of  $\eta_c(3S)$  is most likely to be ~4010 MeV. With this mass, we calculate the radiative transitions  $\eta_c(3S) \rightarrow h_c(1P)\gamma$ ,  $h_c(2P)\gamma$ . Our prediction of the decay rate of  $\eta_c(3S) \rightarrow h_c(1P)\gamma$  is tiny. However, the partial decay widths are rather sizeable, and with the screened potential model, we predict that

$$\Gamma[\eta_c(3S) \to h_c(2P)\gamma] \simeq 130 \text{ keV},$$
 (66)

which is slightly (~20%) larger than our prediction with the linear potential model. Our prediction of  $\Gamma[\eta_c(3S) \rightarrow h_c(2P)\gamma]$  is consistent with the previous calculation in Ref. [8] (see Table IX). Combined with the predicted width  $\Gamma \approx 80$  MeV from Ref. [8], the branching ratio of Br[ $\eta_c(3S) \rightarrow h_c(2P)\gamma$ ] is estimated to be  $1.6 \times 10^{-3}$ .

### 5. $\psi(4160)$ and the missing 2D states

The 1<sup>--</sup> state  $\psi(4160)$  is commonly identified with the  $2^{3}D_{1}$  state. The average experimental mass and width from the PDG are  $M = 4191 \pm 5$  and  $\Gamma = 70 \pm 10$  MeV, respectively [6], which are consistent with linear potential model predictions. However, with a screened potential, the predicted mass for  $\psi_{1}(2D)$  is about 100 MeV smaller than the observation. The  $\psi_{1}(2D)$  resonance can decay into  $\chi_{cJ}(1P)\gamma$  and  $\chi_{cJ}(2P)\gamma$  via the radiative transitions.

Considering  $\psi(4160)$  as a pure  $2^3D_1$  state, and with the linear potential model, we predict that

$$\Gamma[\psi(4160) \to \chi_{c0}(1P)\gamma] \simeq 150 \text{ keV}, \qquad (67)$$

$$\Gamma[\psi(4160) \to \chi_{c1}(1P)\gamma] \simeq 37 \text{ keV}, \qquad (68)$$

$$\Gamma[\psi(4160) \to \chi_{c2}(1P)\gamma] \simeq 17 \text{ keV.}$$
(69)

Similar results are also obtained with the screened potential model. Our predictions of  $\Gamma[\psi(4160) \rightarrow \chi_{c0,1}(1P)\gamma]$  are slightly smaller than those obtained in Ref. [17]; however, our predictions are notably larger than those in Ref. [8] (see Table VII). Combining the measured decay width of  $\psi(4160)$  with our predicted partial widths from the linear potential model, we estimate the branching fractions:

$$Br[\psi(4160) \to \chi_{c0}(1P)\gamma] \simeq 2.1 \times 10^{-3},$$
 (70)

$$Br[\psi(4160) \to \chi_{c1}(1P)\gamma] \simeq 0.5 \times 10^{-3},$$
 (71)

$$Br[\psi(4160) \to \chi_{c2}(1P)\gamma] \simeq 0.2 \times 10^{-3}.$$
 (72)

Our predictions are in the range of the recent measurements  $\operatorname{Br}[\psi(4160) \rightarrow \chi_{c1}(1P)\gamma] < 6.1 \times 10^{-3}$  and  $\operatorname{Br}[\psi(4160) \rightarrow \chi_{c2}(1P)\gamma] < 16.2 \times 10^{-3}$  from the Belle Collaboration [100]. We expect that more accurate observations can be carried out in future experiments.

Furthermore, we calculate the partial decay width of  $\Gamma[\psi(4160) \rightarrow \chi_{cJ}(2P)\gamma]$  with the linear and screened potential models, respectively. Our results are listed in Table VII. Both of the models give a similar result. It is found that the decay rates of  $\psi(4160) \rightarrow \chi_{c0}(2P)\gamma, \chi_{c1}(2P)\gamma$  are rather large; their partial decay widths may be 300–400 keV. Similar results were also obtained in Refs. [8,10]. The estimated branching ratios are

$$Br[\psi(4160) \to \chi_{c2}(2P)\gamma] \simeq 0.3 \times 10^{-3},$$
 (73)

$$Br[\psi(4160) \to \chi_{c1}(2P)\gamma] \simeq 4.4 \times 10^{-3},$$
 (74)

$$Br[\psi(4160) \to \chi_{c0}(2P)\gamma] \simeq 4.4 \times 10^{-3}.$$
 (75)

It should be mentioned that 3 fb<sup>-1</sup> of new data of  $\psi(4160)$  have been collected at BESIII [96]. Using the cross section of ~8 nb based on BES and CLEO measurements [97–99], we estimate that  $2.4 \times 10^7$  events of  $\psi(4160)$  have been accumulated at BESIII. Thus, if  $\psi(4160)$  is the  $2^3D_1$  state indeed, it might provide us a source to look for the missing  $\chi_{c0}(2P)$  and  $\chi_{c1}(2P)$  states via the transition chains  $\psi(4160) \rightarrow \chi_{cJ}(2P)\gamma \rightarrow \psi(1S, 2S)\gamma\gamma$ . The combined branching ratios of these decay chains and the producing events of  $\chi_{cJ}(2P)$  estimated by us have been listed in Table XII. It is found that if the  $\chi_{cJ}(2P)$  is to be observed at BESIII via the transition chains of  $\psi(4160) \rightarrow \chi_{cJ}(2P)\gamma \rightarrow \psi(1S, 2S)\gamma\gamma \rightarrow \gamma\gamma\mu^+\mu^-$  one should accumulate more data samples of  $\psi(4160)$  in the coming years.

The other three 2D-wave states,  $\psi_2(2D)$ ,  $\psi_3(2D)$ , and  $\eta_{c2}(2D)$ , are still not observed in experiments. With the masses and wave functions predicted from the linear and screened potential models, we calculate their radiative decay properties. Our results are listed in Table VII. It is seen that, although the predictions in details from both the linear and screened potential models have a notable difference, both models predict that these 2D-wave states  $\psi_{2,3}(2D)$  and  $\eta_{c2}(2D)$  have relatively large transition rates into the 1P- and 2P-wave states. The partial decay widths for the  $2D \rightarrow 1P\gamma$  processes are about 10s keV, and their branching ratios are estimated to be Br[ $2D \rightarrow 1P\gamma$ ] ~  $\mathcal{O}(10^{-4})$ , while the partial decay widths for the  $2D \rightarrow 2P\gamma$  processes usually reach 100s keV, and their branching

ratios are estimated to be  $\text{Br}[2D \to 2P\gamma] \sim \mathcal{O}(10^{-3})$ . The large decay rates of the  $2D \to 2P\gamma$  processes were also predicted in Ref. [8]. We further estimate the combined branching ratios of the two-photon cascades  $2D \to nP \to mS$ . Our results have been listed in Table XII. In these decay chains, the most prominent two-photon cascades are  $\psi_2(2D) \to \chi_{c1}(1P)\gamma \to J/\psi\gamma\gamma \to \gamma\gamma\mu^+\mu^ (\text{Br} \approx 1.5 \times 10^{-5})$  and  $\eta_{c2}(2D) \to h_c(1P)\gamma \to \eta_c\gamma\gamma \to$  $\gamma\gamma K\bar{K}\pi$  (Br  $\approx 4.8 \times 10^{-5}$ ). In the coming years, Belle II will accumulate a  $10^{10} B\bar{B}$  data sample, which might let us obtain enough events of 2D-wave states via  $B \to \psi_2(2D)X$ and  $B \to \eta_{c2}(2D)X$  decays. If the branching fractions of Br $[B \to \psi_2(2D)X]$  and Br $[B \to \eta_{c2}(2D)X]$  are  $\mathcal{O}(10^{-5})$ , the missing 2D-wave states might be observed in the above two-photon cascades.

### 6. X(4140, 4274) and the 3P states

Until now, no 3P charmonium states have been established in experiments. According to the predicted masses and wave functions of the 3P charmonium states, we estimate their radiative properties decay properties with both the linear and screened potential models, which are listed in Table VIII. From the table, it is seen that most of our results from both models are similar in magnitude. The  $\chi_{c0}(3P)$  state has a large decay rate into the  $\psi(3S)\gamma$ channel; the partial width might be  $10 \text{ s} \sim 100 \text{ s} \text{ keV}$ , which is consistent with the prediction in Ref. [8]. The  $\chi_{c1,2}(3P)/h_c(3P)$  state has a large partial decay width into  $\psi(1S, 2S, 3S)\gamma/\eta_c(1S, 2S, 3S)\gamma$  channels, which are  $10 \text{ s} \sim 100 \text{ s}$  keV as well. Combined with the predicted widths from Ref. [8], the estimated branching ratios of  $\operatorname{Br}[\chi_{c1,2}(3P) \to \psi(1S, 2S, 3S)\gamma]$  and  $\operatorname{Br}[\chi_{c0}(3P) \to \psi(3S)\gamma]$ are  $\mathcal{O}(10^{-3})$ . Using  $B \to \chi_{c1,2}(3P)K/h_c(3P)K$  decays, the forthcoming Belle II and LHC experiments might reconstruct these higher  $\chi_{c1,2}(3P)/h_c(3P)$  states in the  $\psi(1S, 2S)\gamma/\eta_c(2S)\gamma$  decay modes.

Recently, two new charmoniumlike states X(4140)( $\Gamma \approx 16$  MeV) and X(4274) ( $\Gamma \approx 56$  MeV) are confirmed by the LHCb Collaboration [52]. Their quantum numbers are determined to be  $J^{PC} = 1^{++}$ . According to the predicted mass from the linear potential model, the X(4274)might be a good candidate of  $\chi_{c1}(3P)$ . However, within the screened potential model, X(4140) seems to favor the  $\chi_{c1}(3P)$  state. If the X(4140) state is assigned as  $\chi_{c1}(3P)$ , within the screened potential model, the partial radiative decay widths of the dominant channels are predicted to be

$$\Gamma[\psi(4140) \to J/\psi\gamma] \simeq 38 \text{ keV},$$
 (76)

$$\Gamma[\psi(4140) \to \psi(2S)\gamma] \simeq 51 \text{ keV}, \tag{77}$$

$$\Gamma[\psi(4140) \to \psi(3S)\gamma] \simeq 36 \text{ keV.}$$
(78)

Combining the average measured width with the predicted partial radiative decay widths of X(4140), the branching ratios are estimated to be

$$Br[\psi(4140) \to J/\psi\gamma] \simeq 2.4 \times 10^{-3},$$
 (79)

$$Br[\psi(4140) \to \psi(2S)\gamma] \simeq 3.2 \times 10^{-3},$$
 (80)

$$Br[\psi(4140) \to \psi(3S)\gamma] \simeq 2.3 \times 10^{-3},$$
 (81)

while, if the X(4274) state is assigned as  $\chi_{c1}(3P)$ , within the linear potential model, the partial radiative decay widths of the dominant channels are predicted to be

$$\Gamma[\psi(4274) \rightarrow J/\psi\gamma] \simeq 48 \text{ keV},$$
 (82)

$$\Gamma[\psi(4274) \to \psi(2S)\gamma] \simeq 88 \text{ keV}, \tag{83}$$

$$\Gamma[\psi(4274) \to \psi(3S)\gamma] \simeq 297 \text{ keV}, \qquad (84)$$

Combining the measured width with the predicted partial radiative decay widths of X(4274), the branching ratios are estimated to be

$$Br[\psi(4274) \to J/\psi\gamma] \simeq 0.9 \times 10^{-3},$$
 (85)

$$Br[\psi(4274) \to \psi(2S)\gamma] \simeq 1.6 \times 10^{-3},$$
 (86)

Br[
$$\psi(4274) \to \psi(3S)\gamma$$
]  $\simeq 5.3 \times 10^{-3}$ . (87)

The search for X(4274) and X(4140) in the  $\psi(1S, 2S, 3S)\gamma$  channels and the measurements of their partial width ratios might be helpful in uncovering the nature of these two newly observed states.

### 7. 4S states

In the 4S states, the  $\psi(4S)$  resonance seems to favor the  $1^{--}$  state  $\psi(4415)$  according to the linear potential model predictions [8]. However, there are other explanations about  $\psi(4415)$ . For example, in the screened potential model,  $\psi(4415)$  favors  $\psi(5S)$  more than  $\psi(4S)$  [11], while with a coupled-channel method, the  $\psi(4415)$ resonance is suggested to be the  $\psi_1(1D)$  resonance [17]. According to the screened potential model prediction, the  $J^{PC} = 1^{--}$  state X(4260) from the PDG [6] could be a good candidate of the  $\psi(4S)$ . Very recently, BESIII Collaboration observed a new structure Y(4220) with a width of  $\Gamma \simeq 66 \text{ MeV}$  in the  $e^+e^- \rightarrow \pi^+\pi^-h_c$  cross sections [101]. The resonance parameters of Y(4220)are consistent with those of the resonance observed in the  $e^+e^- \rightarrow \omega \chi_{c0}$  [102]. The newly observed Y(4220) might be a candidate of  $\psi(4S)$  as well. To establish the 4S states, more studies are needed.

With the screened potential model, we predict that the masses of the 4S states are about 4.28 GeV, while in the linear potential model, their masses are about 4.41 GeV. Taking the predicted masses of  $\psi(4S)$  with 4412 MeV and 4281 GeV from the linear and screened potential models, respectively, we calculate the radiative transitions of the  $\psi(4S)$  state within these two models. Our results are listed in Table IX. It is found that both the linear and screened potential models give comparable predictions of the decay rates for the 4S states in the magnitude, although the details are different. The decay rates of  $4S \rightarrow 2P$ , 3P are sizeable; the partial widths for the transitions  $\psi(4S) \rightarrow \chi_{cI}(2P)\gamma$  are about 10–20 keV; and for the transitions,  $\psi(4S) \rightarrow$  $\chi_{cI}(3P)\gamma$  are about 20–80 keV. Combined with the predicted widths  $\Gamma \simeq 78$  MeV from Ref. [8], the branching ratios for the  $4S \rightarrow 2P, 3P$  transitions are  $\mathcal{O}(10^{-4})$ .

In the coming years, BESIII plans to collect  $5 \sim 10 \text{ fb}^{-1}$  data samples at  $\psi(4S)$  [96]. Using the cross section of ~4 nb based on BES measurements [97,98], we expect to accumulate  $(2 \sim 4) \times 10^7 \psi(4S)$ 's. To know the production possibilities of 2P and 3P states via the radiative decay of  $\psi(4S)$ , we estimate the number of production events from the two-photon cascades  $\psi(4S) \rightarrow \chi_{cJ}(2P, 3P)\gamma \rightarrow \psi(1S, 2S)\gamma\gamma$ , which has been listed in Table XIII. Unfortunately, it is found that the higher 2P and 3P states cannot be produced via these radiative decay chains at BESIII.

#### 8. Higher multipole contributions

In our calculations, we find that the corrections from the magnetic part to some radiative transitions of the *S*-, *P*-, and *D*-wave states are notable (see Tables VI, IX, VIII, and VII). For example, the magnetic part could give a 10% - 30% correction to the radiative partial decay widths of  $\Gamma[\chi_{cJ}(1P) \rightarrow J/\psi\gamma]$ ,  $\Gamma[\psi_1(1D) \rightarrow \chi_{c1,2}(1P)\gamma]$  and  $\Gamma[\psi(3S) \rightarrow \chi_{cJ}(1P)\gamma]$ . This large correction is mainly caused by the interferences between the extra electric-dipole term  $E_R$  from the magnetic part and the leading E1 transitions. About the higher-order EM corrections to the radiative transitions, some discussions can be found in the literature [19,20,74,103–109].

In experiments, the higher-order amplitudes for the transition(s)  $\chi_{c1,2}(1P) \rightarrow J/\psi\gamma$  and/or  $\psi(2S) \rightarrow \chi_{c1,2}(1P)\gamma$  have been measured in different experiments [110–115]. Our predictions with both the linear and screened potential models compared with the data have been listed in Table XIV. From the table, it is seen that both models give comparable results. The predicted ratios between the magnetic quadrupole amplitude and the electric-dipole amplitude,  $a_2/a_1$ , for the  $\chi_{c1,2}(1P) \rightarrow J/\psi\gamma$  processes are in good agreement with the recent measurements from CLEO [114]. The ratios of  $a_2/a_1$  for the  $\psi(2S) \rightarrow \chi_{c1,2}(1P)\gamma$  are small. Their absolute values are comparable to the measurements from CLEO [114]; however, the sign of  $a_2/a_1$  predicted by us seems to be opposite to the

TABLE XIV. The predicted ratios between the magnetic quadrupole amplitude  $a_2$  and the electric-dipole amplitude  $a_1$  compared with the data. The predicted ratios between the extra electric-dipole  $E_R$  and electric-dipole  $a_1$  are also listed. LP and SP stand for our results obtained from the linear potential and screened potential models, respectively.

Process	$\frac{\underline{E}_{R}}{a_{1}}$ LP	$\frac{\underline{E}_{R}}{a_{1}}$ SP	$\frac{\frac{a_2}{a_1}}{\mathbf{SP}}$	$\frac{a_2}{a_1}$ LP	$\frac{a_2}{a_1}$ Lat. [20]	$\frac{a_2}{a_1}$ CLEO [114]	$\frac{\frac{a_2}{a_1}}{\text{BESII [113]}}$	$\frac{\frac{a_2}{a_1}}{\text{Crystal Ball [110]}}$	$\frac{\frac{a_2}{a_1}}{\text{BESIII}} [115]$
$\overline{\chi_{c1}(1P) \to J/\psi\gamma}$	+0.062	+0.065	-0.065	-0.062	-0.09(7)	-0.0626(87)		$-0.002^{+0.008}_{-0.020}$	
$\chi_{c2}(1P) \rightarrow J/\psi\gamma$	-0.078	-0.082	-0.110	-0.105	-0.39(7)	-0.093(19)		$-0.333^{+0.116}_{-0.292}$	
$\psi(2S) \rightarrow \chi_{c1}(1P)\gamma$	-0.030	+0.031	-0.031	-0.030		0.0276(96)		$0.077_{-0.045}^{+0.050}$	
$\psi(2S) \to \chi_{c2}(1P)\gamma$	+0.021	+0.022	-0.030	-0.028	•••	0.010(16)	$-0.051\substack{+0.054\\-0.036}$	$0.132\substack{+0.098\\-0.075}$	0.046(23)

TABLE XV. The predicted ratios  $\frac{a_2}{a_1}$  and  $\frac{E_R}{a_1}$  with the linear potential (LP) and screened potential (SP) models.

process	$\frac{E_R}{a_1}$ (LP)	$\frac{E_R}{a_1}$ (SP)	$\frac{a_2}{a_1}$ (SP)	$\frac{a_2}{a_1}$ (LP)
$\overline{\psi_1(1D) \to \chi_{c1}(1P)\gamma}$	+0.088	+0.092	+0.041	+0.040
$\psi_1(1D) \rightarrow \chi_{c2}(1P)\gamma$	+0.214	+0.224	+0.074	+0.066
$\chi_{c1}(2P) \to J/\psi\gamma$	+0.108	+0.113	-0.113	-0.108
$\chi_{c2}(2P) \rightarrow J/\psi\gamma$	-0.143	-0.151	-0.203	-0.192
$\chi_{c1}(2P) \rightarrow \psi(2S)\gamma$	+0.034	+0.036	-0.036	-0.034
$\chi_{c2}(2P) \rightarrow \psi(2S)\gamma$	-0.041	-0.043	-0.058	-0.055
$\chi_{c1}(3P) \rightarrow J/\psi\gamma$	+0.147	+0.144	-0.144	-0.147
$\chi_{c2}(3P) \rightarrow J/\psi\gamma$	-0.213	-0.207	-0.277	-0.286
$\chi_{c1}(3P) \rightarrow \psi(2S)\gamma$	+0.086	+0.078	-0.078	-0.086
$\chi_{c2}(3P) \rightarrow \psi(2S)\gamma$	-0.107	-0.096	-0.128	-0.144
$\chi_{c1}(3P) \rightarrow \psi(3S)\gamma$	+0.038	+0.026	-0.026	-0.038
$\chi_{c2}(3P) \to \psi(3S)\gamma$	-0.046	-0.031	-0.041	-0.062

measurements. It should be mentioned that our prediction of  $a_2/a_1$  for the  $\psi(2S) \rightarrow \chi_{c2}(1P)\gamma$  is consistent with the previous measurement from BESII [113]. More accurate measurements may clarify the sign problem. The ratios between the extra electric-dipole amplitudes  $E_R$  and the  $a_1$ are also predicted. It is found that  $|\frac{E_R}{a_1}| \approx |\frac{a_2}{a_1}|$ .

Furthermore, we predict the ratios  $E_R/a_1$  and  $a_2/a_1$  for some unmeasured processes  $\psi_1(1D) \rightarrow \chi_{c1,2}(1P)\gamma$  and  $\chi_{cJ}(nP) \rightarrow \psi(mS)\gamma$ , in which the magnetic part plays an important role. Our results have been listed in Table XV. From the table, it is found that most of the ratios are fairly large. Some ratios can reach to ~30%. Since the  $\psi_1(1D)$  and  $\chi_{c2}(2P)$  have been established and the ratios of  $a_2/a_1$  for  $\psi_1(1D) \rightarrow \chi_{c1,2}(1P)\gamma$  and  $\chi_{c2}(2P) \rightarrow J\psi\gamma$  are fairly large, we suggest the experimentalists measure the ratios of  $a_2/a_1$  for these transitions in future experiments.

# **IV. SUMMARY**

In this work, we calculate the charmonium spectrum with two models, the linear potential model and screened potential model. We should emphasize the following:

(i) The hyperfine and fine splittings show less model dependence. The predicted splitting,

 $m(2^{3}P_{2}) - m(2^{3}P_{1}) \approx 90$  MeV, does not support the X(3915) assigned as the  $\chi_{c0}(2P)$  state.

- (ii) In the screened potential model, the states X(4260) and X(4360) with  $J^{PC} = 1^{--}$  may be good candidates of the  $\psi(4S)$  and  $\psi_1(3D)$ , respectively.
- (iii) For the  $J^{PC} = 1^{++}$  states X(4140) and X(4274) newly confirmed by the LHCb, within the linear potential model, the X(4274) might be identified as the  $\chi_{c1}(3P)$  states, while within the screened potential model, the X(4140) is a good candidate of  $\chi_{c1}(3P)$ .

Second, we further evaluate the EM transitions of charmonium states up to the 4S multiplet. The following are found:

- (i) For the EM transitions of the well-established lowlying charmonium states  $J/\psi$ ,  $\psi(2S)$ ,  $\chi_{cJ}(1P)$ ,  $h_c(1P)$ , and  $\psi(3770)$  both linear potential and screened potential models give similar descriptions, which are in reasonable agreement with the measurements.
- (ii) Identifying the newly observed state X(3823) at Belle and BESIII as the  $\psi_2(1D)$ , its EM decay properties of are in good agreement with the measurements.
- (iii) Assigning the X(3872) resonance as the  $\chi_{c1}(2P)$  state, the ratio  $\frac{\Gamma[X(3872) \rightarrow \psi(2S)\gamma]}{\Gamma[X(3872) \rightarrow J/\psi\gamma]} \approx 1.3$  predicted by us is close to the lower limit of the measurements from *BABAR* and LHCb. Thus, the X(3872) as the  $\chi_{c1}(2P)$  cannot be excluded.

Third, we discuss the observations of the missing charmonium states by using radiative transitions:

- (i) The large  $B\bar{B}$  data sample from Belle II should let us have chances to establish the missing  $\eta_{c2}(1D)$  and  $\psi_3(1D)$  states in forthcoming experiments. The  $\eta_{c2}(1D)$  state should be produced via the  $B \rightarrow$  $\eta_{c2}(1D)K$  process and reconstructed in the  $h_c(1P)\gamma$  decay mode with  $h_c(1P) \rightarrow \eta_c\gamma$  and  $\eta_c \rightarrow K\bar{K}\pi$ , while the  $\psi_3(1D)$  state should be produced via the  $B \rightarrow \psi_3(1D)K$  process and reconstructed in the  $\chi_{c2}(1P)\gamma$  decay mode with  $\chi_{c2}(1P) \rightarrow J/\psi\gamma$  and  $J/\psi \rightarrow \mu^+\mu^-/e^+e^-$ .
- (ii) If BESIII can accumulate a  $5 \sim 10 \text{ fb}^{-1} \psi(4040)$ data sample in the coming years, significant

numbers of  $\chi_{c2}(2P)$  are to be produced via the radiative decay of  $\psi(4040)$  and reconstructed in the  $J/\psi\gamma$  decay mode with  $J/\psi \rightarrow \mu^+\mu^-$ .

- (iii) Relatively large data samples of 2*D*-wave states  $\psi_2(2D)$  and  $\eta_{c2}(2D)$  might be collected at Belle II or LHCb via  $B \rightarrow \psi_2(2D)X$  and  $B \rightarrow \eta_{c2}(2D)X$  decays in forthcoming experiments; the two-photon decay chains  $2^3D_2 \rightarrow \chi_{c1}(1P)\gamma \rightarrow J/\psi\gamma\gamma \rightarrow \gamma\gamma\mu^+\mu^-$  (Br  $\approx 1.5 \times 10^{-5}$ ) and  $\eta_{c2}(2D) \rightarrow h_c(1P)\gamma \rightarrow \eta_c(1S)\gamma\gamma \rightarrow K\bar{K}\pi\gamma\gamma$  (Br  $\approx 4.8 \times 10^{-5}$ ) are worth observing.
- (iv) The missing 3*P*-wave states might be observed at LHCb and Belle II in the  $B \rightarrow \chi_{c1,2}(3P)K/h_c(3P)K$  decays and reconstructed in the  $\psi(1S, 2S)\gamma/\eta_c(2S)\gamma$  decay modes with  $\psi(1S, 2S) \rightarrow \mu^+\mu^-/\eta_c(2S) \rightarrow K\bar{K}\pi$ .

Finally, we study the corrections of higher EM multipole amplitudes to the EM transitions. The magnetic part could give about a 10% ~ 30% correction to the radiative partial decay widths of  $\Gamma[\chi_{cJ}(1P) \rightarrow J/\psi\gamma]$ ,  $\Gamma[\psi_1(1D) \rightarrow \chi_{c1,2}(1P)\gamma]$  and  $\Gamma[\psi(3S) \rightarrow \chi_{cJ}(1P)\gamma]$ . This large correction is mainly caused by the interferences between the extra electric-dipole term  $E_R$  from the magnetic part and the leading E1 amplitudes. Our predictions for the normalized magnetic quadrupole amplitude  $M_2$  of the  $\chi_{c1,2}(1P) \rightarrow J/\psi\gamma$  processes are in good agreement with the recent measurements from CLEO [114]. About the normalized magnetic quadrupole amplitude of  $\psi(2S) \rightarrow \chi_{c1,2}(1P)\gamma$ , there may be a sign difference between our predictions and the measurements. The normalized extra electric-dipole amplitudes  $E_R$  are also predicted. It is found that  $|E_R| \approx |M_2|$ . Furthermore, we find that there are fairly large magnetic quadrupole amplitudes  $M_2$  for the  $\chi_{c1,2}(2P, 3P) \rightarrow \psi(1S, 2S)\gamma$  and  $\psi_1(1D) \rightarrow \chi_{c1,2}(1P)$  processes. We suggest the experimentalists measure the higher magnetic quadrupole amplitudes  $M_2$  of the  $\chi_2(2P) \rightarrow \psi(1S, 2S)\gamma$  and  $\psi_1(1D) \rightarrow \chi_{c1,2}(1P)$  processes in future experiments.

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