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# Optimal design of calibration signals in space-borne gravitational wave detectors

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Future space-borne gravitational wave detectors will require a precise definition of calibration signals to ensure the achievement of their design sensitivity. The careful design of the test signals plays a key role in the correct understanding and characterization of these instruments. In that sense, methods achieving optimal experiment designs must be considered as complementary to the parameter estimation methods being used to determine the parameters describing the system. The relevance of experiment design is particularly significant for the LISA Pathfinder mission, which will spend most of its operation time performing experiments to characterize key technologies for future space-borne gravitational wave observatories. Here we propose a framework to derive the optimal signals—in terms of minimum parameter uncertainty—to be injected into these instruments during the calibration phase. We compare our results with an alternative numerical algorithm which achieves an optimal input signal by iteratively improving an initial guess. We show agreement of both approaches when applied to the LISA Pathfinder case.

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# I. INTRODUCTION

LISA Pathfinder [1] is an ESA mission with NASA contributions designed to test key technologies for the detection of gravitational waves in space, like the proposed eLISA [2]. The main scientific goal for the mission is expressed in terms of a differential acceleration noise between two test masses in nominally geodesic motion down to a level of  $S_{\Delta a} = 3 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$  at 3 mHz. The relevance of this requirement is not only its demand in terms of noise reduction but also the very low frequency measuring band, which introduces technological difficulties that cannot be addressed by ground-based gravitational wave detectors due to the so-called seismic wall [3].

The LISA Pathfinder mission is currently planned to have a six-month operation period at the Lagrange point L1 that will be split between the two experiments onboard: the European LISA Technology Package (LTP) and the American Disturbance Reduction System (DRS). This leads to a very short operation period of roughly three months for the complete characterization and achievement of the scientific goal for the LTP.

It is worth noticing that, after the demonstration of the technology readiness, a second—yet not less relevant objective of the mission is a detailed characterization of the noise contributions to the main scientific measurement. An extensive list of experiments has been put forward by the scientific team including experiments to characterize the optical metrology [4], the inertial sensor instrument [5], the effects of the thermal [6] and magnetic [7]

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#### MIQUEL NOFRARIAS et al.

environment, and pure free-fall experiments that aim to measure acceleration noise in a configuration that is even more representative of eLISA [8]. All these runs need to be executed via telecommands using a daily eight-hour communication window with the satellite. Internal constraints in preprocessing and validation of telecommands will add a latency from two to three days between the definition of a telecommand sequence and its execution on the spacecraft.

The planning of experiments represents therefore a crucial part of the mission and needs to be optimized accordingly to make sure that the information obtained from each experiment is maximized. As part of this effort, a MATLAB toolbox has been developed with the specific aim to deal with the LTP data during flight operations [9]. Among the different methods and capabilities of this tool, much attention has been paid to the improvement of the methods to obtain precise parameters from the experiments [10–14]. These have been tested with simulated data, taking into account the expected noise performance of the LISA Pathfinder mission, in a series of mock data challenges with data generated using the analysis software's built-in modeling and simulation tools. Agreement between methods was also checked with data generated from an independent spacecraft simulator developed by the prime industrial contractor, as was the case in the LISA Pathfinder operational exercises [15].

These analyses focused on the parameter estimation strategy and the achievement of an optimal precision in the parameters obtained, following the heritage of previous simulated data exercises, like for instance the LISA Mock Data Challenge [16] that focused on the problem of astrophysical parameter estimation from LISA data. Unlike the problem of astrophysical data analysis, in the LISA Pathfinder case, the measured signal is the response of the LISA Pathfinder system to some injected input signal that was specified by the telecommand file. In other words, there exists the opportunity in LISA Pathfinder to design the injected signals so that the measurement of the system parameters is optimized. The operators of ground-based gravitational-wave detectors have a similar opportunity to design signals when characterizing the response of their instruments to various noise sources but, given their easy access to their instruments, not as much emphasis is placed on optimizing signal injections. Instead, for a space-borne gravitational wave observatory, such optimal experiment designs might prove very important for maximizing science return for a given mission duration. LISA Pathfinder thus represents a scenario where careful signal design would produce the most benefit. In the following, we propose a general framework which allows the optimization of the input signals applied to a given system.

Optimal experiment design [17–19] is a long-standing area of research. In general terms, the main objective is to adjust the experiment in such a way that the maximal information is obtained from the data. This general purpose

has of course applicability in a wide variety of areas spanning the study of physical, biological or engineering systems. The reader is referred to reviews covering this extensive field of research for more insight [20–24]. In most cases, experiment design is described as an optimization problem for a given figure of merit, which typically relates to a scalar of the Fisher information matrix. Although the description used here applies to a general case, in the current work we will be mostly interested in the application to the estimation of the main parameters governing the combined dynamics of the test mass and the spacecraft in LISA Pathfinder. Hardware onboard the satellite imposes us a further limitation which is only to consider sinusoidal signals as input signals.

This work is organized as follows. In Sec. II, we introduce the problem of experiment design and the notation used in this work. Section III describes a numerical algorithm to optimize the signal to be injected in a given model and its application to a simple case. In Sec. IV, we introduce the LISA Pathfinder model used for our analysis, and in Sec. V we present our conclusions.

#### **II. FISHER MATRIX ANALYSIS**

#### A. Definitions and notation

In the following, we will describe a given system as

$$\vec{o}(\omega) = \mathbf{H}(\omega; \Theta)\vec{s}(\omega) + \vec{n}(\omega), \tag{1}$$

where  $\vec{o}$  is a vector with the measurements being considered,  $\vec{s}$  is a vector with injection signals that can be applied to test the system and  $\mathbf{H}(\omega; \Theta)$  is the matrix whose components,  $H_{ij}(\omega; \Theta)$ , contain the transfer function describing the dynamics of the system in the frequency domain with a dependence on a set of parameters  $\Theta = \{\theta_1, \dots, \theta_N\}$ .  $\vec{n}(\omega)$ describes the noise contribution of our instrument.

The likelihood function is the probability to observe a measurement for a given set of parameters describing that system. Assuming that the data are Gaussian distributed, the likelihood for our system will be

$$p(\vec{o}|\Theta) = [2\pi\Sigma]^{-1/2} \\ \times \exp\left[-\frac{1}{2}(\vec{o} - \mathbf{H}(\Theta) \cdot \vec{s})^T \Sigma^{-1}(\vec{o} - \mathbf{H}(\Theta) \cdot \vec{s})\right]$$
(2)

where  $\Sigma$  is the noise covariance matrix. Experiment design is based on the analysis of the Fisher matrix, whose elements are defined as

$$F_{ij} = \left\langle \left( \frac{\partial \log(p(\vec{o}|\Theta))}{\partial \theta_i} \right)^T \left( \frac{\partial \log(p(\vec{o}|\Theta))}{\partial \theta_j} \right) \right\rangle \bigg|_{\theta_0}$$
(3)

which can be used to set limit for an expected covariance matrix of the parameters known as the Crámer-Rao bound [25],

OPTIMAL DESIGN OF CALIBRATION SIGNALS IN ...

$$\operatorname{cov}[\theta_i, \theta_i] \ge \mathbf{F}^{-1}. \tag{4}$$

The decomposition of the Fisher matrix into eigenvalues and eigenvectors will prove to be very useful in the following sections. Given a N × N Fisher matrix **F**, defined by a set of N parameters, the eigenvectors  $\vec{u}$  and eigenvalues  $\lambda$  always fulfill

$$\mathbf{F}\vec{u} = \lambda\vec{u}.\tag{5}$$

The eigenvectors can be used to diagonalize the Fisher matrix according to the following property,

$$\mathbf{F} = \mathbf{R}^{\mathrm{T}} \Lambda \mathbf{R},\tag{6}$$

where the columns of the matrix **R** are the (normalized) eigenvectors of **F**, and  $\Lambda$  is a diagonal matrix with the eigenvalues in the diagonal. Notice that **R** can be understood as a rotation matrix that can be used to express the vector of our initial parameters,  $\vec{\Theta}$ , in the new diagonal basis  $\vec{u}$ ,

$$\vec{\zeta} = \mathbf{R}\vec{\Theta},\tag{7}$$

from where we obtain our new set of parameters in the diagonal basis,  $\vec{\zeta}$ .

#### **B.** Fisher matrix tomography

To compute the Fisher matrix, we need to follow Eq. (3). We notice though that even for this simplified problem the straightforward application of this expression leads to a long expression that makes difficult a further analytical treatment. To avoid cumbersome expressions as much as possible, we expand the Fisher matrix in its different composing terms. Following the notation in Eq. (1), we consider an experiment with M inputs,  $\vec{s}$ , and N outputs,  $\vec{o}$ . In such a case, we may write the elements of our Fisher matrix as

$$F_{ij} = \sum_{n,q=1}^{M} \sum_{m,p=1}^{N} F_{mnpq,ij},$$
(8)

where

$$F_{mnpq,ij} = \{\Sigma^{-1}\}_{mp} [\partial_{\theta_i} H_{mn}(\Theta)]^T [\partial_{\theta_j} H_{pq}(\Theta)] s_n s_q.$$
(9)

The definition of the Fisher matrix allows us to combine the information of different experiments by adding their Fisher matrices. However, in this case, we use this same property in the opposite direction: to split a single experiment as the combination of simpler independent experiments. This tomography will be particularly useful to interpret the Fisher matrix since we can split each experiment into the contribution of each transfer function and study them independently. The  $F_{mnpq,ij}$  term can be understood as the mp component of a Fisher matrix corresponding to an experiment which only considers a sinusoidal input applied to the nq channels. We notice here that if the noise covariance matrix,  $\{\Sigma^{-1}\}_{mp}$ , were diagonal, we could consider each  $F_{mnpn,ij}$  as the contribution corresponding to a given transfer function  $H_{mn}(\Theta)$ . However, cross-couplings between our channels imply a mixing of the different transfer function contributions.

# **III. DESIGN OF INPUT SIGNALS**

The experiment design problem can be stated as how to choose an input signal that allows the optimization of a given figure of merit, provided some constraints on our particular experiment. In the literature [23], there are several options for a scalar figure of merit to use as a minimization criteria including (i) the minimization of the trace of the covariance matrix (A-optimality), thus minimizing the average variance of the parameters, (ii) the minimization of the largest eigenvalue of the covariance matrix (E-optimality), which implies minimizing the major axis of the uncertainty ellipsoid in the parameter space or (iii) minimizing the determinant of the covariance matrix (D-optimality), which is the equivalent to minimizing the uncertainty ellipsoid in the parameter space. In the following, we will stick to the latter criterion since, among other advantages, it remains invariant under scale changes in the parameters.

An analytical solution to the problem, as the one proposed in the previous section, has a limited application and becomes unfeasible for complex systems. The usual strategy is to describe the problem as a numerical minimization problem as we show below. For computational simplicity, the inverse of the Fisher matrix is used as an approximation of the covariance matrix. Since we are working in a high SNR regime, it is also a good approximation.

For mathematical convenience, our description of the system under study will be in frequency domain. Hence, recalling Eq. (1), the input to our system will be described as

$$\chi(\omega) = (|s(\omega_1)|^2 \dots |s(\omega_F)|^2)$$
(10)

with bounded energy,

$$\sum_{k=1}^{F} |s(\omega_k)|^2 = 1$$
(11)

where  $s(\omega_k)$  is the frequency domain representation of a given input at frequency  $\omega_k$ . Following the literature, we will call  $\chi(\omega)$  our design. Different conditions can be set on the design in order to achieve D-optimality. Indeed, it can be shown that a design maximizing the determinant of the Fisher matrix will minimize the maximum of the quantity [18]

$$\nu(\omega) = \operatorname{tr}[\mathbf{F}^{-1}(\chi)\mathbf{F}(\omega)] \tag{12}$$

where  $\mathbf{F}(\chi)$  is the information matrix from the design  $\chi(\omega)$ and  $\mathbf{F}(\omega)$  is the information matrix from a single frequency input with normalized power spectrum  $|X(\omega)|^2 = 1$ . The quantity  $\nu(\omega)$ , known as a dispersion function or response dispersion, can be understood as the ratio of the variance of the system transfer function to the noise power.

Based on its mathematical properties, the dispersion function has been proposed as a tool for input design optimization. The underline idea is to select a frequency grid where the power of the input signal is initially uniformally distributed among the selected frequencies. The dispersion function is then computed for each frequency in the grid and the power of the signal distributed proportionally to the value of this function. The optimal design is achieved by repeating this procedure iteratively. More precisely, the algorithm steps are [26,27]

- (1) Select a set of frequencies  $\{\omega_1, ..., \omega_F\}$  within the frequency band of interest and distribute the power equally over these frequencies. This constitutes the initial design.
- (2) Compute the dispersion function for the F frequencies.
- (3) Create a new design according to

$$\chi_{i+1}(\omega_k) = \chi_i(\omega_k) \times \nu_i(\omega_k) / N_\theta \qquad (13)$$

(4) If max(ν(χ<sub>i</sub>, ω<sub>k</sub>)) − N<sub>θ</sub>) < ε for a sufficient small ε, then the optimum design is found. If not, we return to step 2.</li>

It can be shown that the algorithm converges to a D-optimal design. [26].

In order to prove the efficiency of the previous numerical design method, we test it in the case of a harmonic oscillator. We can analytically compute the Fisher matrix for this problem to obtain an expression which, as expected, shows a maximum of the spectrum at the natural frequency of the oscillator,  $\omega_0$ . This value is, therefore, the one that minimizes the volume of the error ellipsoid in the parameter space and, hence, the one that the numerical method described in the previous section should retrieve.

In order to check the validity of our methodology, we generated a time series of 10,000 seconds of white noise with variance  $\sigma = 10^{-5}$  that we consider as our initial input design. We choose white noise in order to weight all frequencies equally. We consider a harmonic oscillator with damping ratio  $\xi = 0.01$  and natural frequency  $\omega_0 = 0.07$ , and then we run the algorithm as described above. The result is shown in Fig. 1 where we show the evolution of the input signal as proposed by the algorithm. As shown, two





FIG. 1. Evolution of the algorithm to optimize the input signal for the harmonic oscillator case. The algorithm promotes the natural frequency of the oscillator  $\omega_0 = 0.07$ .

iterations are enough for the algorithm to promote the natural frequency of the oscillator  $\omega_0$  among the others.

### **IV. LISA PATHFINDER MODEL**

In order to apply this methodology to LISA Pathfinder, we will need first to define a model for the experiment. In the following, we introduce the notation to describe the combined dynamics of the two test masses and the satellite required for the analysis. The same description with small variations can also be found in [10,11,13].

#### A. Equation of motion

The measurement onboard the satellite is usually expressed as

$$\vec{o} = (\mathbf{D} \cdot \mathbf{S}^{-1} + \mathbf{C})^{-1} (-\mathbf{C}\vec{o}_i + \vec{g}_n + \mathbf{D} \cdot \mathbf{S}^{-1}\vec{o}_n), \quad (14)$$

where **D** is the dynamical matrix, **C** is the controller, and **S** stands for the sensing matrix, which translates the physical position of the test masses into the interferometer readout,  $\vec{o}$ . Subindex *n* stands for noise quantities, either sensing noise  $(\vec{o}_n)$  or force noise  $(\vec{g}_n)$ , and subindex *i* stands for the injected signals  $(\vec{o}_i)$ . Restricting ourselves to linear motion along the axis between the two test masses (the degree of freedom that is measured by the interferometer), each of the dynamical variables in Eq. (14) can be expressed as two-dimensional vectors with components referring to the  $x_1$  and  $x_{\Delta}$  channels, respectively,

$$\vec{o} = \begin{pmatrix} o_1 \\ o_{\Delta} \end{pmatrix}, \qquad \vec{o}_i = \begin{pmatrix} o_{i1} \\ o_{i\Delta} \end{pmatrix},$$
$$\vec{o}_n = \begin{pmatrix} o_{n1} \\ o_{n\Delta} \end{pmatrix}, \qquad \vec{g}_n = \begin{pmatrix} g_{n1} - g_N \\ g_{n2} - g_{n1} \end{pmatrix}.$$

## OPTIMAL DESIGN OF CALIBRATION SIGNALS IN ...

where subindices 1 and 2 refer to the first and second test mass, subindex  $\Delta$  refers to differences between the first and second test mass, and capitalized subindices (such as force noise on the spacecraft,  $g_N$ ) refer to the spacecraft. The last equation in Eq. (15) shows how  $g_N$  is only measured in the first channel. On the other hand, the differential channel is sensitive to any differential force noises applied to the first and the second test masses.

The matrices describing the dynamics of the LISA Pathfinder system are

$$\mathbf{D} = \begin{pmatrix} s^{2} + \omega_{1}^{2} + \frac{m_{1}}{m_{SC}} \omega_{1}^{2} + \frac{m_{2}}{m_{SC}} \omega_{2}^{2} & \frac{m_{2}}{m_{SC}} \omega_{2}^{2} \\ \omega_{2}^{2} - \omega_{1}^{2} & s^{2} + \omega_{2}^{2} \end{pmatrix},$$
$$\mathbf{C} = \begin{pmatrix} H_{df} & 0 \\ 0 & H_{sus} \end{pmatrix}, \qquad \mathbf{S} = \begin{pmatrix} S_{11} & S_{12} \\ S_{21} & S_{22} \end{pmatrix}, \qquad (15)$$

where  $\omega_1$  and  $\omega_2$  are the stiffnesses—the steady force gradient across the test mass housing per unit mass [28] coupling the motion of each test mass to the motion of the spacecraft;  $H_{df}$  and  $H_{sus}$  are the drag-free and suspension loops controllers, respectively. For the remainder of this work, it is assumed that  $H_{df}$  and  $H_{sus}$  are known since they represent software control loops already tested on ground.

For our current analysis we will assume some approximations in these expressions in order to keep the main scientific information and, at the same time, keep the expressions as simple as possible. For that reason, in the following we will eliminate the back reactions terms,  $m_1 = m_2 \ll m_{SC}$ , consider that the sensing matrix crosscouplings are zero  $S_{12} = S_{21} = 0$ . For convenience, we will take the calibrations  $S_{11} = S_{22} = 1$ . Taking into account these assumptions we can derive expressions for the transfer functions describing the system. We consider input signals injected at the guidance input port which we expressed as  $o_i$  in Eq. (14); hence, the transfer function is defined by

$$\mathbf{H} = (\mathbf{D} \cdot \mathbf{S}^{-1} + \mathbf{C})^{-1} (-\mathbf{C}) \cdot \vec{o}_i$$
$$= \begin{pmatrix} H_{11}(\Theta) & H_{12}(\Theta) \\ H_{21}(\Theta) & H_{22}(\Theta) \end{pmatrix} \begin{pmatrix} o_{i1} \\ o_{i\Delta} \end{pmatrix}, \quad (16)$$

where the transfer functions are given by

$$H_{11} = \frac{H_{\rm df}}{\omega^2 - \omega_1^2 + H_{\rm df}}$$
(17)

$$H_{12} = 0$$
 (18)

$$H_{21} = \frac{H_{\rm df}(\omega_2^2 - \omega_1^2)}{(\omega^2 - \omega_1^2 + H_{\rm df})(\omega^2 - \omega_2^2 + H_{\rm lfs})}$$
(19)

$$H_{22} = \frac{H_{\rm lfs}}{\omega^2 - \omega_2^2 + H_{\rm lfs}},$$
 (20)

TABLE I. LPF noise model parameters.  $p_2$  and  $p_4$  parameters correspond to frequencies in [Hz] and  $p_1$  to amplitude spectral densities in  $[m/\sqrt{Hz}]$  and  $[N/\sqrt{Hz}]$  for read-out noise and force noise, respectively.

Noise parameters			
Parameter	$o_{n1}/o_{n\Delta}$	$g_{n1}/g_{n2}$	$g_N$
$p_1$	$3.6 \times 10^{-12}$	$7 \times 10^{-15}$	$2.5 \times 10^{-10}$
$p_2$	$10 \times 10^{-3}$	$5 \times 10^{-3}$	$12 \times 10^{-3}$
$p_3$	4.2	3	3.8
$p_4$	$1.8 \times 10^{-3}$	$4 \times 10^{-4}$	$1 \times 10^{-3}$
$p_5$	8	8	8

where we realize that  $H_{12}$  is zero because this is proportional to the parameter  $S_{12}$ , which is considered to be zero. At the same time, we see that the cross-coupling from drag-free to differential channel,  $H_{21}$ , is proportional to the differential stiffness,  $\omega_2^2 - \omega_1^2$ .

## **B.** Noise model

Our study of the injection scheme in LISA Pathfinder relies on the Fisher matrix which, in turn, depends on the noise model used for those noise sources identified in Eq. (14). These are interferometer read-out noise for both channels— $o_{n1}$  and  $o_{n\Delta}$ —force noise applied to the test masses— $g_{n1}$  and  $g_{n2}$ —and force noise applied to the spacecraft— $g_N$ . Following [13], we will characterize each of these with the five parameters  $p_{1...5}$  in the expression

$$S_n(\omega) = p_1 \left( 1 + \frac{1}{(\frac{\omega}{2\pi p_2})^{p_3}} + \frac{1}{(\frac{\omega}{2\pi p_4})^{p_5}} \right)^{1/2}.$$
 (21)

Applying the parameters in Table I, we obtain the models in Fig. 2 for the noise spectra of the two main



FIG. 2. Comparison of the noise spectra for the two main interferometer channels for an analytical simple model (dashed line) and a noise data stream generated via a LPF state-space model (solid line).



FIG. 3. Left: evaluation of the eigenvalue  $\lambda_2$  of the  $F_{2121}$  term for a single frequency injection. Right: Output of the numerical algorithm for the optimization of input signals based on the dispersion function applied to a LISA Pathfinder state-space model after 25 iterations. In this case, white noise was injected into the drag-free channel. In both cases (analytical and numerical), the analysis is repeated by rescaling by a factor of 25 and 50 the value of the second test mass stiffness, which is originally considered to be  $\omega_0^2 = -22 \times 10^{-6} \text{ s}^{-2}$ .

interferometer channels. We can compare the predictions from this simplified model to simulations coming from a detailed state-space simulator containing a much elaborate model of the instrument, for instance delays, actuators, and component noise models [29]. As seen in Fig. 2, our simple parametric model agrees well with the noise obtained from the state-space model.

# V. CALIBRATION SIGNALS FOR LISA PATHFINDER DYNAMICS

During operations, LISA Pathfinder will run an exhaustive characterization campaign with the objective of calibrating the instrument and identifying the main noise contributions. Here we consider one set of experiments targeting the calibration of the dynamical parameters governing the combined motion of the two test masses and the satellite. For these particular set of experiments, the calibration procedure consists of the injection of a sequence of sinusoids—the only available waveform in the flight software—at different frequencies at a number of input ports. For this work, we will focus on injection in one of the two main interferometer channels. However, the methodology can be easily applied to the remaining degrees of freedom.

# A. The $F_{2121}$ term

In order to demonstrate our method, we consider an injection applied to the drag-free channel. In our framework, this experiment would be completely described by the sum

$$F_{ij} = \sum_{m,p} F_{m1p1,ij},$$
 (22)

where the indices *i* and *j* run over the parameters. The most general case (7 degrees of freedom) correspond to 49 terms. This is not approachable analytically so we focus our attention on one term with particular relevance, the  $F_{2121}$ , which can be expressed as

$$F_{2121,ij} = \{\Sigma^{-1}\}_{22} \times [\partial_{\theta_i} H_{21}(\omega)]^T [\partial_{\theta_j} H_{21}(\omega)] |o_1(\omega)|^2.$$
(23)

This term quantifies the effect of the injection in the first channel as measured by the highly sensitive differential channel. Under the assumptions discussed in sec. IV A, the only parameters that impact this term are the two test mass stiffnesses, which enter through the term in Eq. (19). Due to this simplification, we can describe this problem in analytical terms. Equation (23) turns into a  $2 \times 2$  matrix that we can easily decompose in the related eigenvectors,

$$\vec{u_1} = \begin{bmatrix} \frac{\alpha(\omega_1)\beta(\omega_1)}{\alpha(\omega_2)\beta(\omega_2)} \\ 1 \end{bmatrix} \vec{u_2} = \begin{bmatrix} -\frac{\alpha(\omega_2)\beta(\omega_2)}{\alpha(\omega_1)\beta(\omega_1)} \\ 1 \end{bmatrix}, \quad (24)$$

where

0

$$\alpha(x) = H_{\rm df}(\omega) - x^2 + \omega^2 \tag{25}$$

$$\beta(x) = H_{\rm lfs}(\omega) - x^2 + \omega^2 \tag{26}$$

and associated eigenvalues,

$$\lambda_{1} = 0$$
  

$$\lambda_{2} = \{\Sigma^{-1}\}_{22} H_{df}^{2}(\omega) \frac{\alpha^{2}(\omega_{1})\beta^{2}(\omega_{1})\alpha^{2}(\omega_{2})\beta^{2}(\omega_{2})}{\alpha^{4}(\omega_{1})\beta^{4}(\omega_{2})}.$$
 (27)



FIG. 4. Determinant of the  $F_{2121}$  term for an injected signal with two independent frequencies. The determinant is evaluated for the case  $\omega_2^2 = -22 \times 10^{-7} \text{ s}^{-2}$  (left) and  $\omega_2^2 = 50 \times (-22 \times 10^{-7}) \text{ s}^{-2}$  (right). The values on the legend are the log(det[F<sub>2121</sub>]).

Since eigenvalues are directly related to the inverse of the expected uncertainty on the associated parameter, we conclude that this measurement can only constrain parameters in the  $\vec{u}_2$  direction, while the direction  $\vec{u}_1$  has an associated uncertainty that tends to infinity.

It is important to notice at this point that the analysis we perform in the frequency domain implicitly assumes a unique frequency, i.e. an input signal which is a sinusoid at a given frequency. In the following section, we explore which information can be obtained in such a case.

## 1. Single tone input: Undetermined solution

If we assume that the content of our input signal is a sinusoid with a fixed frequency, we know from the previous eigendecomposition that we will not be able to solve the problem since we have only one valid eigenvector for a two-dimensional problem. Nonetheless, we explore the single frequency solution in order to determine how much information can we get from the system in such a case.

We proceed to diagonalize the Fisher matrix as in Eq. (6),

$$\mathbf{F} = \mathbf{R}^{\mathrm{T}} \Lambda \mathbf{R},$$

from which we obtain a diagonal system with a unique eigenvalue,  $\lambda_2$ , given by Eq. (27). In Fig. 3, we explore this expression as a function of the frequency of the injection. We see that the eigenvalue has a peak when the input is injected at a frequency around f = 1.25 mHz. This becomes more evident if we increase the value of  $\omega_2^2$ , as shown in the figure.

The value of f = 1.25 mHz is therefore the best frequency for a signal composed with a unique frequency component for the experiment under study. Indeed, by maximizing the Fisher matrix we are reducing the error on

the parameter space. However, it must be noted that this is not necessarily an optimal solution since we are dealing with a single frequency injection scheme that leads to singular Fisher matrix.

A second consideration to take into account is that when diagonalizing our system, our parameters are expressed in a new basis which corresponds to applying a rotation matrix **R** to the original vector of parameters  $\vec{\Theta} = \{\omega_1^2, \omega_2^2\}$ . In doing so, we obtain a new set of parameters  $\vec{\zeta} = \mathbf{R} \cdot \vec{\Theta}$ . For the configuration under study, the combination of parameters corresponding to the nonzero eigenvalue is proportional to the sum of stiffness, i.e.

$$\zeta_2 \propto \omega_1^2 + \omega_2^2, \tag{28}$$

confirming that a single frequency signal is not able to break the degeneracy between the two parameters in our system.

We are now prepared to compare the results obtained analytically with the prediction of the numerical algorithm based on the dispersion function, Eq. (12). To do so, we inject a white noise data stream to the input channel under consideration, which for the analysis of the  $F_{2121}$ term is the drag-free channel. In this particular case, we consider as our initial input a white noise time series of  $10^5$  s and  $\sigma = 10^{-6}$  m<sup>2</sup>.

In the right panel of Fig. 3, we show the resulting normalized power spectrum of the input signal as retrieved after 25 iterations of the numerical optimization algorithm. The algorithm promotes the same frequencies that maximized the eigenvalue of the Fisher matrix  $F_{2121}$  as can be seen in the left-hand figure. Moreover, we performed the analysis by rescaling the  $\omega_2^2$  value as in the study of the eigenvalues. Here we observe again how the numerical algorithm selects the f = 1.25 mHz frequency when



FIG. 5. Expected error on parameters for an injection in the drag-free channel considering  $\omega_1^2$  and  $\omega_2^2$  as the only relevant parameters. Black corresponds to the initial proposal of a white noise input, blue represents the expected error for the input signal as obtained with the proposed numerical algorithm after 25 iterations. Histograms are computed based on 5000 samples of a multivariate Gaussian distribution with the covariance matrix obtained by the numerical algorithm.

approaching the case where  $\omega_2^2$  is rescaled by a factor of 50, proving the consistency between the analytical and the numerical approach. The former is derived from Eq. (14), while the latter has its roots in the numerical computation of the dispersion function Eq. (12) using a state-space representation of LISA Pathfinder.

#### 2. Two-tone input: Full-rank solution

Here we take advantage of the analytical solution to go one step further and explore the case of an input signal composed by two sinusoids. In order to combine the information of more than one sinusoid frequency in the input signal, we add the Fisher matrices corresponding to each frequency. Our experiment will therefore be described by

$$F_{2121,ij} = \sum_{k}^{N=2} \bar{F}_{2121,ij}(\omega_k), \qquad (29)$$

where each  $\bar{F}_{2121,ij}(\omega_k)$  corresponds to the contribution of a single sinusoid injection to the final experiment's Fisher matrix.

We first explore the rank of the  $F_{2121}$  matrix when evaluated for different combinations of these two input frequencies. Given that  $F_{2121}$  depends on two parameters, results show that most combinations of frequencies are able to reach the condition rank $(F_{2121}) = 2$ . In fact, only

when the two frequencies are equal-and we come back to our previous case-will we not be in a full-rank situation. This allows us to go one step further and explore which combination of frequencies are optimal, in the sense of maximizing the Fisher matrix, i.e. minimizing the ellipsoid error volume in the parameters space. Figure 4 shows the value for the determinant of the  $F_{2121}$  term as a function of the two injection frequencies. We explore the determinant for two different configurations of the experiment: the standard with  $\omega_2^2 = -22 \times 10^{-7} \text{ s}^{-2}$  and, as before, rescaling  $\omega_2^2 = 50 \times (-22 \times 10^{-7}) \text{ s}^{-2}$ . As expected, the determinant shows symmetry since the two injection frequencies in Eq. (29) can be interchanged producing the same output. The determinant drops to zero at the diagonal since, as commented above, an injection with two equalfrequency sinusoids does not lead to a full-rank solution. It is interesting to see that when we set  $50 \times \omega_2^2$ , a notch appears at the frequency f = 1.25 mHz that we found as a maximum in the single injection case.

In the standard configuration, the maximum of the  $F_{2121}$  determinant appears for frequencies in the very low frequency regime (f < 1 mHz). If, for practical reasons, we set one of the two injections to be  $f_1 = 0.1$  mHz, the maximum of the function displayed in Fig. 4 appears for a second injection at  $f_2 = 0.3$  mHz. With these two values, we can proceed to estimate the expected errors on the parameters by evaluating the Fisher matrix in Eq. (29). By

assuming two sinusoid injections with two cycles each at the obtained frequencies  $f_1 = 0.1$  and  $f_1 = 0.3$  mHz with and amplitude of  $10^{-7}$  m, we can evaluate our expression for the Fisher matrix term  $F_{2121}$ , obtaining a  $7 \times 10^{-3}$ % and  $6 \times 10^{-3}$ % relative error estimate for the two stiffness parameters  $\omega_1^2$  and  $\omega_2^2$ , respectively. It is worth noting here that these are optimal errors representing the contribution of the  $F_{2121}$  term of the Fisher matrix to the overall experiment. We consider it as a useful example to show the capability of the framework here proposed to disentangle the different contributions to the experiment. However, the precise determination of the expected error for a given parameter requires the evaluation of the full Fisher matrix, which is composed in the analytical description of 49 components for the drag-free injection experiment. Hence, analysis considering the whole system is, in most cases, more suited to a numerical approach.

In order to evaluate the improvement on the estimate of the parameters, we run the analysis using the numerical algorithm introduced in Sec. III assuming an injection in the drag-free channel and considering only the two stiffness  $\omega_1^2$ and  $\omega_2^2$  as relevant parameters. As described above, the algorithm evaluates the Fisher matrix at each step so we can trace how the expected errors for each parameter improve by modifying the input signal. The improvement in the error, as given by the Fisher matrix, is shown in Fig. 5, where we compare the expected error on the parameters at the 1st and at the 25th iteration. The input signals associated with these two cases correspond to a white noise injection for the first iteration that turns into a signal focusing all the power at f = 1.25 mHz after 25 iterations. The results show a clear improvement in the expected error on the parameters which decreases roughly by an order of magnitude.

#### VI. ROBUSTNESS ANALYSIS

A last consideration that has to be taken into account in the design of experiments is that of the robustness of the analysis. A key issue of the experiment design framework is that it relies on the evaluation of the Fisher matrix, which depends on the true values of the system under study, precisely the unknowns that the experiment aims to identify. In most cases, some reasonable estimates for the expected values exist and therefore the experiment is designed on the basis of this *a priori* knowledge. The assumption is thus that the design obtained will not show a strong dependence on the values being considered. However, this raises the question if the experiment being defined in such a way is merely reinforcing our previous knowledge about the system [30]. This difficulty has been recognized in the literature, and several approaches have been proposed in order to achieve a robust optimal design scheme [31], although methods that are robust with respect to uncertainties in the system parameters are a wide-open research field [22].



FIG. 6. Normalized input spectrum for a Monte Carlo analysis running the experiment design algorithm, considering two input channels and five parameters. The solid line represents the mean, for each channel, of the 1000 runs while the grey shadow is the corresponding standard variation.

In the particular case of LISA Pathfinder, we focus on the sequential design scheme [27,32,33] which proposes to overcome the circular reasoning above by iteratively switching from parameter estimation to experiment design using the most recent parameter estimates from the previous step to define the next experiment. Indeed, this experiment design strategy fits particularly for our mission scenario since it includes in a natural way the process of improving the knowledge of our system parameters that will occur during the mission. Moreover, the sequential design can be already tested against the mission simulators that we have previously introduced, as we show in the following.

Let us consider a LISA Pathfinder state-space model, where instead of the 2 degrees of freedom considered before, we increase now the complexity of the problem to include five unknown parameters in the system, corresponding to the parameter set  $\{G_{df}, G_{lfs}, \omega_1^2, \omega_2^2, S_{21}\}$  in Eq. (14). That way, we allow our system to explore a wider parameter space.

In order to quantify the robustness of our experiment design strategy, we run a Monte Carlo analysis allowing the parameters of our model to be uniformly distributed as

$$G_{\rm df} \sim G_{\rm lfs} \sim \mathcal{U}[0.9, 0.2],$$
  

$$\omega_1^2 \sim \omega_1^2 \sim \mathcal{U}[1.9 \times 10^{-6}, 1.1 \times 10^{-6}],$$
  

$$S_{21} \sim \mathcal{U}[1.55 \times 10^{-4}, 1.45 \times 10^{-4}].$$
(30)

It is worth noticing here that the allowed range of discrepancy in the parameters is orders of magnitude higher than the expected uncertainty in the parameters. Indeed, previous studies [10–14] have shown that the order of magnitude of expected error on the parameters considered in our analysis is  $\sigma_{G_{\rm df}} \simeq \sigma_{G_{\rm lfs}} \simeq 10^{-5}$  for the control gains,  $\sigma_{\omega_1^2} \simeq \sigma_{\omega_2^2} \simeq 10^{-10} \text{ s}^{-2}$  for the stiffnesses and  $\sigma_{S_{21}} \simeq 10^{-8}$  for the sensing cross-coupling.

We run 1000 analyses of the algorithm described in Sec. III. Each of the individual runs computes the dispersion function in a 100-frequency bin grid spanning 0.1 mHz to 1 Hz. The input spectrum is built iteratively using the dispersion function as a figure of merit. Each individual input spectrum of the Monte Carlo analysis was computed based on 50 iterations of the algorithm. The result of the Monte Carlo run is shown in Fig. 6. The solid line represents, for each channel, the mean of all the input spectrum while the shadow area stands for the standard deviation from all the runs. As clearly shown, in the LISA Pathfinder mission framework and considering a scenario of high uncertainty in the parameters, the methodology described is still robust. Results show a clear convergence around an input power spectrum for both channels and guarantee that in a mission-realistic scenario-with unknown parameter values-the methodology described here is a valid protocol to define the calibration signals that will help us get the maximum information from the experiments and, hence, to optimize the mission time line.

#### **VII. CONCLUSIONS**

LISA Pathfinder and future space-borne gravitational wave detectors will require precise calibration of their dynamical systems in order to operate at their design sensitivities. Given the operational constraints for such missions, the design of injection signals used for calibration is a key aspect for efficient characterization of the instrument.

We have introduced a methodology to design experiments for these instruments based on the minimization of the uncertainty ellipsoid in parameter space. This methodology allows one to decompose the Fisher information matrix in its different contributions, each related to a unique physical coupling—or transfer function—of the experiment. By studying these contributions, we can evaluate the expected error for a given spectrum of the injected test signal.

We have compared this with a numerical algorithm capable of generating an optimal input signal by iteratively improving a proposed input spectrum. The algorithm uses the dispersion function of the system to promote those frequencies which minimize the error on the parameters under study. We have applied both techniques to one example of LISA Pathfinder injection experiments, obtaining agreement in the injection signals obtained with both approaches.

As an example, we have considered the contributions to the expected error for a given term of the Fisher matrix decomposition: the  $F_{2121}$ , which describes the coupling of the  $x_1$  (the drag-free channel) and the  $x_{12}$  (the differential channel) for the case when a signal is injected in the former. The methodology proposed here is, however, general and can be equally applied to other instruments requiring an accurate calibration in terms of parameter uncertainties, such as ground-based gravitational wave detectors.

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- [1] F. Antonucci *et al.*, From laboratory experiments to LISA pathfinder: Achieving LISA geodesic motion, Classical Quantum Gravity **28**, 094002 (2011).
- [2] The eLISA Consortium *et al.*, The Gravitational Universe, arXiv:1305.5720.
- [3] J. Abadie *et al.*, All-sky search for gravitational-wave bursts in the first joint ligo-geo-virgo run, Phys. Rev. D 81, 102001 (2010).
- [4] H. Audley *et al.*, The LISA pathfinder interferometry hardware and system testing, Classical Quantum Gravity 28, 094003 (2011).
- [5] R. Dolesi, D. Bortoluzzi, P. Bosetti, L. Carbone, A. Cavalleri, I. Cristofolini, M. DaLio, G. Fontana, V. Fontanari, B. Foulon, C. D. Hoyle, M. Hueller, F. Nappo, P. Sarra,

D. N. A. Shaul, T. Sumner, W. J. Weber, and S. Vitale, Gravitational sensor for LISA and its technology demonstration mission, Classical Quantum Gravity **20**, S99 (2003).

- [6] P. Canizares *et al.*, The diagnostics subsystem on board LISA PathFinder and LISA, Classical Quantum Gravity 26, 094005 (2009).
- [7] M. Diaz-Aguilo, E. García-Berro, and A. Lobo, Inflight magnetic characterization of the test masses onboard LISA pathfinder, Phys. Rev. D 85, 042004 (2012).
- [8] A. Grynagier, W. Fichter, and S. Vitale, The LISA pathfinder drift mode: implementation solutions for a robust algorithm, Classical Quantum Gravity 26, 094007 (2009).
- [9] M. Hewitson *et al.*, Data analysis for the LISA Technology Package, Classical Quantum Gravity 26, 094003 (2009).

OPTIMAL DESIGN OF CALIBRATION SIGNALS IN ...

- [10] G. Congedo, L. Ferraioli, M. Hueller, F. De Marchi, S. Vitale, M. Armano, M. Hewitson, and M. Nofrarias, Time domain maximum likelihood parameter estimation in LISA pathfinder data analysis, Phys. Rev. D 85, 122004 (2012).
- [11] L. Ferraioli, G. Congedo, M. Hueller, S. Vitale, M. Hewitson, M. Nofrarias, and M. Armano, Quantitative analysis of LISA pathfinder test-mass noise, Phys. Rev. D 84, 122003 (2011).
- [12] N. Karnesis, M. Nofrarias, C. F. Sopuerta, F. Gibert, M. Armano, H. Audley, G. Congedo, I. Diepholz, L. Ferraioli, M. Hewitson, M. Hueller, N. Korsakova, P. W. McNamara, E. Plagnol, and S. Vitale, Bayesian model selection for LISA pathfinder, Phys. Rev. D 89, 062001 (2014).
- [13] M. Nofrarias, C. Röver, M. Hewitson, A. Monsky, G. Heinzel, K. Danzmann, L. Ferraioli, M. Hueller, and S. Vitale, Bayesian parameter estimation in the second LISA pathfinder mock data challenge, Phys. Rev. D 82, 122002 (2010).
- [14] S. Vitale, G. Congedo, R. Dolesi, V. Ferroni, M. Hueller, D. Vetrugno, W. J. Weber, H. Audley, K. Danzmann, I. Diepholz, M. Hewitson, N. Korsakova, L. Ferraioli, F. Gibert, N. Karnesis, M. Nofrarias, H. Inchauspe, E. Plagnol, O. Jennrich, P. W. McNamara, M. Armano, J. I. Thorpe, and P. Wass, Data series subtraction with unknown and unmodeled background noise, Phys. Rev. D 90, 042003 (2014).
- [15] M. Nofrarias, L. Ferraioli, G. Congedo, M. Hueller, M. Armano, M. Diaz-Aguiló, A. Grynagier, M. Hewitson, and S. Vitale, Parameter estimation in LISA pathfinder operational exercises, J. Phys. Conf. Ser. 363, 012053 (2012).
- [16] S. Babak *et al.*, The mock LISA data challenges: From challenge 3 to challenge 4, Classical Quantum Gravity 27, 084009 (2010).
- [17] V. V. Fedorov, *Theory of Optimal Signals* (Academic Press, New York, 1972).
- [18] G. Goodwin and R. Payne, Dynamic System Identification: Experiment Design and Data Analysis (Academic Press, New York, 1977).
- [19] F. Pukelsheim, *Optimal Design of Experiments* (John Wiley, New York, 1993).
- [20] K. Chaloner and I. Verdinelli, Bayesian experimental design: A review, Stat. Sci. 10, 273 (1995).

- [21] M. Gevers, Identification for control: From the early achievements to the revival of experiment design, European Journal of Control 11, 335 (2005).
- [22] H. Hjalmarsson, From experiment design to closed-loop control, Automatica 41, 393 (2005).
- [23] R. Mehra, Optimal input signals for parameter estimation in dynamic systems–survey and new results, IEEE Trans. Autom. Control 19, 753 (1974).
- [24] E. Walter and L. Pronzato, Qualitative and quantitative experiment design for phenomenological models—a survey, Automatica 26, 195 (1990).
- [25] M. Vallisneri, Use and abuse of the fisher information matrix in the assessment of gravitational-wave parameterestimation prospects, Phys. Rev. D 77, 042001 (2008).
- [26] R. Pintelon and J. Schoukens, System Identification: A Frequency Domain Approach (John Wiley, New York, 2004).
- [27] E. Walter and L. Pronzato, Identification of Parametric Models from Experimental Data (Springer, New York, 1997).
- [28] D. Bortoluzzi, L. Carbone, A. Cavalleri, M. D. Lio, R. Dolesi, C. D. Hoyle, M. Hueller, S. Vitale, and W. J. Weber, Measuring random force noise for LISA aboard the LISA pathfinder mission, Classical Quantum Gravity 21, S573 (2004).
- [29] M. Nofrarias et al., State space modelling and data analysis exercises in LISA Pathfinder, in 9th LISA Symposium, Vol. 467 of the Astronomical Society of the Pacific Conference Series, edited by G. Auger, P. Binétruy, and E. Plagnol (Astronomical Society of the Pacific, 2013), p. 161.
- [30] G. C. Goodwin, C. R. Rojas, and J. S. Welsh, Good, bad and optimal experiments for identification, in *Forever Ljung in System Identification-Workshop on the Occasion of Lennart Ljung's 60th Birthday, Studentlitteratur AB*, 2006.
- [31] C. R. Rojas, J. S. Welsh, G. C. Goodwin, and A. Feuer, Robust optimal experiment design for system identification. Automatica 43, 993 (2007).
- [32] I. Ford, D. M. Titterington, and C. F. J. Wu, Inference and sequential design, Biometrika 72, 545 (1985).
- [33] C. Wu, Asymptotic inference from sequential design in a nonlinear situation, Biometrika **72**, 553 (1985).