

# Large field excursions and approximate discrete symmetries from a clockwork axion

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We present a renormalizable theory of scalars in which the low-energy effective theory contains a pseudo-Goldstone boson with a compact field space of  $2\pi F$  and an approximate discrete shift symmetry  $\mathcal{Z}_Q$  with  $Q \gg 1$ , yet the number of fields in the theory goes as  $\log Q$ . Such a model can serve as a UV completion to models of relaxions and is a new source of exponential scale separation in field theory. While the model is local in “theory space,” it appears not to have a continuum generalization (*i.e.*, it cannot be a deconstructed extra dimension). Our framework shows that super-Planckian field excursions can be mimicked while sticking to renormalizable four-dimensional quantum field theory. We show that a supersymmetric extension is straightforwardly obtained, and we illustrate possible UV completions based on a compact extra dimension, where all global symmetries arise accidentally as a consequence of gauge invariance and five-dimensional locality.

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## I. INTRODUCTION

The simplest field theories describing cosmic inflation—the near-classical slow-rolling of a single field—which match current cosmological data require a canonically normalized field to take on values larger than the Planck scale [1]. In addition, a recent proposal for a solution to the electroweak hierarchy problem—the relaxion [2]—requires field excursions in the early universe much larger than the ultraviolet cutoff of the theory. While the range of validity of a low-energy effective field theory is applied to the energy, momenta, and masses involved in any physical process, field values are not constrained. Nevertheless, there is concern that such large field excursions may not be viably generated in a consistent UV completion, especially one encompassing quantum gravity (see for instance [3,4]). Given the implications for inflation and the electroweak hierarchy problem, it is worth looking for existence proofs of field theories and string theories that produce effectively large field ranges in the infrared.

One way to phrase our question is the following. Consider an ordinary renormalizable quantum field theory in which all scalar fields acquire expectation values  $\sim f < M_{\text{Pl}}$  and which contains a pseudo-Goldstone boson with a compact field space of size  $2\pi F$ : under what conditions is it possible to consistently achieve  $F \gg f$ ? For instance how many degrees of freedom, or how much complexity, is needed given a large  $F/f$ ? We address this question by presenting a simple renormalizable field theory with  $N + 1$  scalars, where a global symmetry is broken at a scale  $f$ , and which produces a large field range  $F \equiv 3^N f$  for a pseudo-Goldstone boson in the low-energy theory. The model can include couplings to fermions charged under strong gauge groups, and thus produce axionlike potentials with effective

decay constants that range from  $f$  to  $3^N f$ . The interactions in the theory are of nearest-neighbor type and follow a simple pattern, and thus we hope it helps lead to a more complete ultraviolet theory (*e.g.*, an embedding in string theory). We show how to make our construction more plausible by presenting similar constructions where the global symmetry is an accidental approximate consequence of a higher-dimensional gauge symmetry, and we relate it to multi-axion models already in the literature. We then show a supersymmetric version can trivially be constructed.

## II. MODEL

Our model has  $N + 1$  complex scalars  $\phi_j, j = 0, \dots, N$  and the following renormalizable potential:

$$V(\phi) = \sum_{j=0}^N \left( -m^2 \phi_j^\dagger \phi_j + \frac{\lambda}{4} |\phi_j^\dagger \phi_j|^2 \right) + \sum_{j=0}^{N-1} (\epsilon \phi_j^\dagger \phi_{j+1}^3 + \text{H.c.}) \quad (1)$$

The first sum of terms respects a  $U(1)^{N+1}$  global symmetry, whereas the remainder of the potential breaks this symmetry explicitly to  $U(1)$ . Under the unbroken  $U(1)$ , the fields  $\phi_j, j = 0, 1, 2, 3, \dots, N$ , have charges  $Q = 1, \frac{1}{3}, \frac{1}{9}, \dots, \frac{1}{3^N}$ . Note, we ignored the  $U(1)^{N+1}$  preserving cross terms,  $\phi_j^\dagger \phi_j \phi_k^\dagger \phi_k$ , and the first sum preserves a permutation symmetry, but this is only done for simplicity of analysis. Our conclusions will not rely on these assumptions. Note finally that our Lagrangian is local in field space, in that it involves only interactions involving neighbours  $\phi_j$  and  $\phi_{j+1}$ , and can thus be associated to a discrete extra



The above ansatz for  $a_j(\theta)$  and  $\lambda(\theta)$  guarantees satisfaction of the eigenvalue equation for  $1 \leq j \leq N-1$ . Satisfaction at the boundaries  $j=0$  and  $j=N$  gives instead the constraints

$$A = -\frac{\sin \theta}{q - \cos \theta}$$

$$0 = (1 + q^2) \sin(N+1)\theta - q \sin N\theta - q \sin(N+2)\theta.$$

The second equation admits  $N$  discrete solutions  $0 < \theta_p < 2\pi$  ( $p = 1, \dots, N$ ) corresponding to  $N$  different eigenvalues  $\lambda(\theta_p)$ . Leaving their explicit values aside, these eigenvalues are all located in the band we found in the infinite matrix limit: besides the zero mode, all other masses are finite and of order  $\sqrt{\epsilon}f$ .

Now we see where this is going. In our case,  $q = 3$ , the zero-mode eigenvector is

$$\vec{a}_{(0)}^T = \mathcal{N} \left( 1 \quad \frac{1}{3} \quad \frac{1}{9} \quad \dots \quad \frac{1}{3^N} \right). \quad (13)$$

This mode has exponentially suppressed overlap,  $\sim 3^{-N}$ , with whatever operator  $\phi_N$  couples to in the UV, and can thus produce an effective decay constant which is exponentially larger than  $f$ . For example, if  $\phi_N$  couples to fermions,  $\psi_N$ , charged under a confining gauge group as  $y\phi_N\bar{\psi}_N\psi_N$ , the fermions will pick up a mass  $yf$ , while the zero-mode,  $a$ , at low energies will pick up the standard axion coupling to the gauge bosons:

$$\sim \frac{a}{32\pi^2(3^N f)} H^{\mu\nu} \tilde{H}_{\mu\nu} \quad (14)$$

producing an effective decay constant exponentially larger than  $f$  ( $H^{\mu\nu}$  is the field strength of the new gauge group). The classical rolling of this field over large (even superduper Planckian) field values would correspond to a clockwork of phase-rotations of the  $\phi_j$  fields. With a handful of  $\phi_j$  fields, this Goldstone field can be a single large-field inflaton.

To produce a relaxion-type of potential, we will need to add a periodic potential with period  $\sim f$ . This can be done by adding a second set of fermions and strong group, and the coupling  $y'\phi_0\bar{\psi}_0\psi_0$ , generating the operator  $\sim \frac{\pi}{32\pi^2 f} G\tilde{G}$ . Below the confinement scales of the two strong groups, the Goldstone's potential becomes:

$$V \sim \Lambda_N^4 \cos \pi/F + \Lambda_0^4 \cos \pi/f \quad (15)$$

where  $F \equiv 3^N f$  and  $\Lambda_N$  and  $\Lambda_0$  are roughly the confinement scales of the two strong groups. Corresponding to the relaxion potential, the scale  $\Lambda_N$  should be at or below the ultraviolet cutoff of the standard model (the scale at which loop contributions to the Higgs mass are cut off), while the

second group confines at a scale below the weak scale (and could even be QCD).

### III. ACCIDENTAL GLOBAL SYMMETRY AND THE UV

The structure of the clockwork axion is preserved by an (admittedly bizarre) global  $U(1)_{PQ}$  symmetry, and one can ask if such a symmetry should be preserved (by, say, higher-dimensional operators) in light of the expected violation of global symmetries by quantum gravity. Stated otherwise: one could ask if the symmetry at the basis of our mechanism could arise accidentally in a more fundamental theory where all symmetries are gauged. One way to obtain such accidental symmetries is via locality in extra dimensions.

One example, using extra-dimensional locality [5], as applied to the axion [6], is structured as follows. The strong groups live in the bulk of an extra dimension, and our model's matter, including the scalars and the chiral fermions they couple to, lives on a brane or hypersurface in an extra dimension. A chiral partner of the same matter (a copy theory with opposite PQ charges) lives on a different brane (or, for example, on the other orbifold fixed point of an  $S^1/\mathbb{Z}_2$  compactification). Now, because over the entire space, the PQ symmetry is vectorlike and thus anomaly-free, the  $U(1)$  symmetry can be gauged in the bulk by a  $A^{PQ}$  vector. Anomaly cancellation is insured locally in the five-dimensional theory by including the suitable Chern-Simons terms in the bulk,

$$\frac{C_G}{64\pi^2} A^{PQ} \wedge G^a \wedge G^a + \frac{C_H}{64\pi^2} A^{PQ} \wedge H^\alpha \wedge H^\alpha, \quad (16)$$

with the correct integers,  $C_G$  and  $C_H$ , to cancel the anomalous representations on each orbifold fixed point, and where  $G$  and  $H$  represent the field strengths. While the PQ,  $G$ , and  $H$  four-dimensional vectors are taken to be even under the orbifold, the  $A_5$  components are taken to be odd (have Dirichlet boundary conditions), thus projecting out the zero modes.

Taking the compactification scale to be  $\pi R \gg 1/M_*$  (where  $M_*$  is the five-dimensional cutoff), and assuming the only bulk charged fields are of masses  $\sim M_*$ , the contact terms between the two sectors are suppressed by a Yukawa potential  $\sim e^{-\pi R M_*}$ , and produce operators such as

$$\frac{\mathcal{O}_Q^n \tilde{\mathcal{O}}_{-Q}^\ell}{M_*^{\ell+n-4}} e^{-\pi R M_*}, \quad (17)$$

where  $\mathcal{O}$  and  $\tilde{\mathcal{O}}$  are operators from each of the two sectors with dimensions  $n$  and  $\ell$  and PQ charges  $\pm Q$ , respectively. Thus, effectively, there are two symmetries—one linear combination which is gauged, and the other which is an approximate, accidental global symmetry due to locality. A spontaneous breaking of the PQ symmetry on one brane

causes the (exact) Goldstone to be eaten by the PQ gauge field. The spontaneous breaking on the second brane produces a near-exact Goldstone (up to exponentially suppressed effects), and thus the global-symmetry is extremely well preserved.

Another partly UV complete theory and structurally analogous to ours is obtained by extrapolating the bi-axion and tri-axion constructions in Ref. [7] to arbitrary  $N$  axions. The setup involves a fifth dimension again compactified on  $S^1/\mathbb{Z}_2$  of length  $\pi R$ , with a  $U(1)^{N+1} \equiv U(1)_0 \times U(1)_1 \dots \times U(1)_N$  gauge theory living in the bulk. The  $A_\mu^j$  components ( $j = 0, \dots, N$ ) of the gauge fields along the noncompact directions are assumed to satisfy Dirichlet boundary conditions at the boundaries, so that no massless extra photon survives compactification. The fifth components  $A_5^j$  satisfy von Neumann boundary conditions and give rise to  $N + 1$  four-dimensional scalars parametrized by the gauge invariant Wilson loops around  $S_1$ :

$$\pi_j \equiv \frac{1}{g_5 \sqrt{2\pi R}} \oint dx^5 A_5^j(x^\mu, x^5) \quad (18)$$

(where the normalization ensures  $\pi_j$  are canonically normalized when taking a  $-F_{MN}F^{MN}/4g_5^2$  five-dimensional gauge kinetic term). At the classical level, the potential for the  $\pi_j$ 's vanishes exactly. In the presence of five-dimensional charged matter, a potential will be generated by the Casimir effect, which will depend on the  $\pi_j$  via the Aharonov-Bohm phase around the circle. Given a bulk multiplet with charges  $(q_0, \dots, q_N)$  under  $U(1)^{N+1}$  the Aharonov-Bohm phase will be

$$\Phi = \frac{1}{f} \sum_{j=0}^N q_j \pi_j \quad f \equiv \frac{1}{g_5 \sqrt{2\pi R}}, \quad (19)$$

and the Casimir energy will be a function of  $e^{i\Phi}$ . Now, our construction is realized by having  $N - 1$  bulk fields  $\Psi_{j,j+1}$  for  $j = 0, \dots, N - 1$  each with charges  $q_j = 1, q_{j+1} = -q$  and  $q_k = 0$  for  $k \neq j, j + 1$ . The quantum numbers of these fields, either bosons or fermions, or both, form a moose structure connecting the  $U(1)$ 's and representing a discrete compact dimension. The effective potential will be a periodic function of the Aharonov-Bohm phases,

$$\Phi_{j,j+1} = \frac{1}{f} (\pi_j - q\pi_{j+1}), \quad (20)$$

which are precisely the same combinations we encountered in the four-dimensional model of the previous section. For bulk field masses  $m \sim 1/R$ , the effective potential will have the form

$$V \sim \frac{1}{16\pi^2} \frac{1}{R^4} F(\{e^{i\Phi_{j,j+1}}\}), \quad (21)$$

and around its minimum will give a mass of order  $1/(R^2 f)$  to  $N$  of the scalar, leaving out the same zero mode of Eq. (11). Notice that, depending on the spin and the mass of the link fields, some of the Wilson lines may acquire an expectation value. The details are beyond the scope of this discussion. However, it is rather clear that such details in the vacuum dynamics do not affect our main point, which concerns the survival of a massless mode with eigenvector given by Eq. (11). Our result for the potential is a direct consequence of the choice of quantum numbers for the light bulk fields. In general, one should expect all sorts of charges under  $U(1)^{N+1}$  corresponding to all possible Aharonov-Bohm phases, and thus no residual massless Goldstone. However, provided the corresponding fields have a mass  $M \gg 1/R$ , their contribution to the effective potential will be suppressed by  $e^{-2\pi MR}$ , with some chance of being negligible.

To make contact with the model of the previous section, we can further assume that the gauge fields of color  $SU(3)$  and of some other non-Abelian gauge group  $H$  propagate in the bulk. We can now consider coupling Abelian to non-Abelian gauge fields directly via five-dimensional Chern-Simons terms. In particular, we can couple  $U(1)_0$  to  $SU(3)$  and  $U(1)_N$  to  $H$  by explicitly adding

$$\frac{C_0}{64\pi^2} A^0 \wedge G^a \wedge G^a + \frac{C_N}{64\pi^2} A^N \wedge H^\alpha \wedge H^\alpha, \quad (22)$$

with  $C_0, C_N$  integers. Notice that, in this case, the orbifold boundary conditions are reversed for the  $U(1)$ s, and their gauge transformations are  $\mathbb{Z}_2$  odd, while  $G$  and  $H$  are even; thus, the above CS terms are not associated to anomaly inflow at the boundary: they imply a coupling to the corresponding Wilson line scalars  $\pi_0$  and  $\pi_N$  in the low-energy effective theory. Focusing just on the surviving axion zero mode  $a$ , we have

$$\frac{a}{f} \frac{\mathcal{N}}{64\pi^2} \left( C_0 G_{\mu\nu}^a \tilde{G}_{\mu\nu}^a + \frac{C_N}{q} H_{\mu\nu}^\alpha \tilde{H}_{\mu\nu}^\alpha \right), \quad (23)$$

so that nonperturbative effects in the two gauge groups give rise to an axion potential modulated over two hierarchically separated periods  $f$  and  $q^N f$ .

One could employ a more elaborate version of this construction to produce an explicit model of the relaxion. For instance, the Higgs boson could be a composite of the confining dynamics of the gauge group factor  $H$ . Then all Higgs-related parameters would be modulated by the relaxion  $a$  but with the slow variation  $q^{-N}/f$ . We leave such detailed constructions for future work.

One final issue in both UV completions above: one could worry about the possible relevance of the contribution to the potential of heavy bulk modes of arbitrary charge. As we said, these effects are controlled by  $e^{-2\pi RM}$ . It is not realistic to make these effects arbitrarily small by taking



$MR$  arbitrarily large. The largest we can hope  $M$  to be is the cutoff of the five-dimensional theory  $M_*$ . In turn,  $M_*$  is bounded from above by the scale at which the bulk non-Abelian gauge interactions,  $G$  and  $H$  become strong. It is an interesting fact<sup>1</sup> that one finds

$$2\pi RM_* \sim \frac{16\pi^2}{g^2}, \quad (24)$$

where  $g$  is the four-dimensional gauge coupling at the compactification scale  $1/R$ . Therefore, the unwanted corrections to the potential from heavy modes are formally similar to the nonperturbative four-dimensional effects, such as those associated with confinement. However, given these effects all appear in exponents, there exists a wide range of possibilities. Overall, it does not seem implausible that the IR four-dimensional effects dominate, but it is not guaranteed. There is, however, no way to ascertain this in the absence of a full UV completion at the five-dimensional cutoff.

#### IV. SUPERSYMMETRIC CLOCKWORK

One can do a supersymmetric version of the clockwork relatively easily. Starting with the following superpotential with  $3(N+1)$  chiral superfields,

$$W = \sum_{j=0}^N \lambda S_j (\phi_j \bar{\phi}_j - f^2) + \epsilon \sum_{j=0}^{N-1} (\bar{\phi}_j \phi_{j+1}^2 + \phi_j \bar{\phi}_{j+1}^2), \quad (25)$$

and taking  $\epsilon$  small, one can see that the first terms will cause a spontaneous breaking of the approximate  $U(1)^{N+1}$  global symmetries (preserving the approximate R symmetry), and that the  $S_j$  fields marry a combination of  $\phi_j$  and  $\bar{\phi}_j$  fields to produce  $2(N+1)$  massive chiral superfields and  $N+1$  massless ones. We parametrize the theory below the mass scale,  $\lambda f$ , by replacing  $\phi_j \rightarrow f e^{\Pi_j/f}$ ,  $\bar{\phi}_j \rightarrow f e^{-\Pi_j/f}$ , where  $\Pi_j$  are the

<sup>1</sup>Which can be explicitly checked using the results in Ref. [8].

(Goldstone) chiral superfields. The low-energy superpotential becomes

$$\rightarrow 2\epsilon f^3 \sum_{j=0}^{N-1} (\cosh(\Pi_j - 2\Pi_{j+1})/f), \quad (26)$$

which produces supersymmetric masses for all but one of the chiral superfields with a mass (not mass-squared) matrix of the form (6) with  $q=2$ . Thus, an effective decay constant of  $F \equiv 2^N f$  can be produced in the low-energy theory.

#### V. COMMENTS

We have shown that a renormalizable field theory with a modest number of fields can produce an effective field range which is exponentially large compared with the input scales in the theory. While our construction is *ad hoc*, it suggests a specific pattern which may be indicative of an even simpler or more complete theory. We also suggested UV completions to render all global symmetries accidental, and it would be interesting to see if similar structures can be made manifest in string constructions and what, if any, the relation is to existing constructions such as axion monodromy [9–12]. The elements we employ are not totally new and are already present in one form or another in the existing literature [7,13–19]. However, we have added a twist, and we believe the combination of all those elements and the simplicity of our construction is a novel feature which could open new pathways in model building.

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[1] A. D. Linde, *Phys. Lett. B* **129**, 177 (1983).  
 [2] P. W. Graham, D. E. Kaplan, and S. Rajendran, *Phys. Rev. Lett.* **115**, 221801 (2015).  
 [3] C. Vafa, *arXiv:hep-th/0509212*.  
 [4] N. Arkani-Hamed, L. Motl, A. Nicolis, and C. Vafa, *J. High Energy Phys.* **06** (2007) 060.  
 [5] N. Arkani-Hamed and S. Dimopoulos, *Phys. Rev. D* **65**, 052003 (2002).  
 [6] H. C. Cheng and D. E. Kaplan, *arXiv:hep-ph/0103346*.

[7] A. de la Fuente, P. Saraswat, and R. Sundrum, *Phys. Rev. Lett.* **114**, 151303 (2015).  
 [8] A. Hebecker and A. Westphal, *Ann. Phys. (Berlin)* **305**, 119 (2003).  
 [9] E. Silverstein and A. Westphal, *Phys. Rev. D* **78**, 106003 (2008).  
 [10] L. McAllister, E. Silverstein, and A. Westphal, *Phys. Rev. D* **82**, 046003 (2010).

- [11] N. Kaloper and L. Sorbo, *Phys. Rev. Lett.* **102**, 121301 (2009).
- [12] N. Kaloper, A. Lawrence, and L. Sorbo, *J. Cosmol. Astropart. Phys.* **03** (2011) 023.
- [13] N. Arkani-Hamed, A. G. Cohen, and H. Georgi, *Phys. Rev. Lett.* **86**, 4757 (2001).
- [14] J. E. Kim, H. P. Nilles, and M. Peloso, *J. Cosmol. Astropart. Phys.* **01** (2005) 005.
- [15] Y. Bai and B. A. Stefanek, *Phys. Rev. D* **91**, 096012 (2015).
- [16] K. Choi, H. Kim, and S. Yun, *Phys. Rev. D* **90**, 023545 (2014).
- [17] T. Higaki and F. Takahashi, *J. High Energy Phys.* **07** (2014) 074.
- [18] R. Kappl, S. Krippendorf, and H. P. Nilles, *Phys. Lett. B* **737**, 124 (2014).
- [19] I. Ben-Dayan, F. G. Pedro, and A. Westphal, *Phys. Rev. Lett.* **113**, 261301 (2014).
- [20] K. Choi and S. H. Im, *J. High Energy Phys.* **01** (2016) 149.