Three-loop hard-thermal-loop perturbation theory thermodynamics at finite temperature and finite baryonic and isospin chemical potential

Jens O. Andersen,¹ Najmul Haque,² Munshi G. Mustafa,³ and Michael Strickland²

¹Department of Physics, Norwegian University of Science and Technology, N-7491 Trondheim, Norway

²Department of Physics, Kent State University, Kent, Ohio 44242, USA

³Theory Division, Saha Institute of Nuclear Physics, 1/AF Bidhannagar, Kolkata 700064, India

(Received 7 January 2016; published 29 March 2016)

In a previous paper [N. Haque *et al.*, J. High Energy Phys. 05 (2014) 27], we calculated the three-loop thermodynamic potential of QCD at finite temperature *T* and quark chemical potentials μ_q using the hard-thermal-loop perturbation theory (HTLpt) reorganization of finite temperature and density QCD. The result allows us to study the thermodynamics of QCD at finite temperature and finite baryon, strangeness, and isospin chemical potentials μ_B , μ_S , and μ_I . We calculate the pressure at nonzero μ_B and μ_I with $\mu_S = 0$, and the energy density, the entropy density, the trace anomaly, and the speed of sound at nonzero μ_I with $\mu_B = \mu_S = 0$. The second- and fourth-order isospin susceptibilities are calculated at $\mu_B = \mu_S = \mu_I = 0$. Our results can be directly compared to lattice QCD without Taylor expansions around $\mu_q = 0$ since QCD has no sign problem at $\mu_B = \mu_S = 0$ and finite isospin chemical potential μ_I .

DOI: 10.1103/PhysRevD.93.054045

I. INTRODUCTION

Quantum chromodynamics (QCD) in extreme conditions such as high temperature and high density has been a very active area of research for more than two decades. The interest in QCD at finite temperature has largely been spurred by the experimental programs in heavy-ion collisions at the Relativistic Heavy Ion Collider in Brookhaven and the Large Hadron Collider at CERN. One of the goals of these programs is the creation and study of the quark-gluon plasma-the deconfined phase of QCD. The equation of state of QCD is essential to the phenomenology of the quark-gluon plasma. Lattice gauge theory provides a first-principle method to calculate the thermodynamic functions of QCD at finite temperature and zero baryon chemical potential μ_B . However, at finite μ_B , QCD suffers from the so-called sign problem, namely that the fermion determinant is complex. This prevents one from using standard lattice techniques involving importance sampling to calculate the partition function of QCD. One way to circumvent this problem, at least for small baryon chemical potentials, is to make a Taylor expansion of the thermodynamic functions around μ_{R} . This requires the calculation of the quark-number susceptibilities evaluated at zero quark chemical potentials, $\mu_q = 0.$

Perturbative QCD offers an alternative to lattice gauge theory for the calculations of thermodynamic functions in the deconfined phase. Invoking asymptotic freedom, one might expect that perturbation theory works at sufficiently high temperatures. However, one does not know *a priori* how large T must be in order to obtain a sufficiently good approximation. Using the weakcoupling expansion in the strong coupling constant g, the calculation of the thermodynamic functions has been pushed to order $g^6 \log g$ both at zero [1] and finite chemical potential [2-4]. However, a strict perturbative expansion in q does not converge at temperatures relevant for the heavy-ion collision experiments. It turns out that the convergence is very poor unless the temperature is many orders of magnitude larger than the critical temperature T_c for the deconfinement transition. The source of the poor convergence is the contributions to the thermodynamic functions coming from soft momenta of order gT. The poor convergence of the weak-coupling expansion suggests that one needs to reorganize the perturbative series of thermal QCD. For scalar theories, screened perturbation theory (SPT) has been applied successfully up to four loops [5-8]. SPT is in part inspired by variational perturbation theory [9–14]; see also [15] for a renormalization-group improved reorganization of the perturbative series. In the case of gauge theories, using a local mass term for the gluons breaks gauge invariance, and one needs to generalize SPT. Hard-thermal-loop perturbation theory (HTLpt) represents such a generalization and was developed over a decade ago [16]. Since its invention, HTLpt has been used to calculate thermodynamic functions through three loops at zero chemical potential [17–21] as well as finite chemical potential [22,23]. Depending on the thermodynamic function at hand, the agreement between lattice simulations and the results from HTLpt is very good down to temperatures of approximately $T \simeq 250$ MeV. Application of some HTL-motivated approaches can be found in Refs. [24-35].

While three-color QCD at finite baryon chemical potential has a sign problem, there are a number of other cases

ANDERSEN, HAQUE, MUSTAFA, and STRICKLAND

where the sign problem is absent. This includes QCD in a strong magnetic field B, two-color QCD at finite baryon chemical potential μ_B [36,37], and three-color QCD at finite isospin chemical potential μ_I [38]. In this paper, we will focus on three-color QCD at finite isospin density. There are a few papers on lattice QCD with finite isospin chemical potential [39-46]; however, these mostly focus on the phase transitions themselves and not on the deconfined phase: In addition to the deconfinement transition, there is an additional transition to a Bose condensate of pions at sufficiently low temperature Tand sufficiently large isospin chemical potential μ_I [47]. For T = 0, the critical chemical potential for pion condensation is $\mu_I^c = m_{\pi}$. Moreover, the results of [44] seem to indicate that the first-order deconfinement transition at zero isospin density turns into a crossover at $\mu_I/T \simeq 2.5$. At sufficiently low temperature and high isospin chemical potential, i.e. around the phase boundary, HTLpt is unreliable. Thus, at this point in time, we cannot compare our HTLpt predictions with lattice Monte Carlo at finite μ_I . Therefore, our results should be considered as predictions which can be checked by future lattice simulations. This is in contrast to threecolor QCD at $\mu_B = 0$, where there is a plethora of lattice results [48-66] on the thermodynamics of the deconfined phase.

The paper is organized as follows. In Sec. II, we briefly discuss finite chemical potentials and the sign problem of QCD. In Sec. III, we review hard-thermal-loop perturbation theory and the HTLpt thermodynamic potential through next-next-to-leading order (NNLO). In Sec. IV, we present and discuss our numerical results for the thermodynamic functions. In Sec. V, we summarize and conclude.

II. PARTICLE DENSITIES, CHEMICAL POTENTIALS, AND THE SIGN PROBLEM IN QCD

In massless QCD with N_f flavors there are N_f^2 conserved charges which correspond to the number of generators of the group $SU(N_f) \times U(1)$. For each conserved charge Q_i , we can introduce a nonzero chemical potential μ_i . However, it is possible to specify the expectation values of different charges simultaneously, only if they commute. For $N_f = 2$ and $N_f = 3$, this implies that we can introduce two and three independent chemical potentials, respectively. These can conveniently be chosen as the quark chemical potentials μ_q , which corresponds to the separate conservation of the number of u, d, and squarks. However, any other independent linear combination of μ_q is equivalent, and it is customary to introduce chemical potentials for baryon number n_B , isospin n_I , and strangeness n_S . After having introduced the chemical potentials in the Lagrangian, the partition function as well as all thermodynamic quantities are functions of the temperature and the chemical potentials. For example, the corresponding charge densities n_i are given by

$$n_i = -\frac{\partial \mathcal{F}}{\partial \mu_i},\tag{1}$$

where \mathcal{F} is the free energy density.

The baryon, isospin, and strangeness densities n_B , n_I , and n_S can be expressed in terms of the quark densities n_f as

$$n_B = \frac{1}{3}(n_u + n_d + n_s),$$
(2)

$$n_I = n_u - n_d, \tag{3}$$

$$n_S = -n_s. \tag{4}$$

Equations (2)–(4) can be used to derive relations between the corresponding chemical potentials μ_B , μ_I , and μ_S and the quark chemical potentials μ_q . Equations (1) and (3) give

$$n_{I} = -\frac{\partial \mathcal{F}}{\partial \mu_{I}}$$
$$= -\left(\frac{\partial \mathcal{F}}{\partial \mu_{u}} - \frac{\partial \mathcal{F}}{\partial \mu_{d}}\right) = -\left(\frac{\partial \mu_{u}}{\partial \mu_{I}}\frac{\partial \mathcal{F}}{\partial \mu_{u}} + \frac{\partial \mu_{d}}{\partial \mu_{I}}\frac{\partial \mathcal{F}}{\partial \mu_{d}}\right). \quad (5)$$

Comparing the second and third lines in Eq. (5), we infer that

$$\frac{\partial \mu_u}{\partial \mu_I} = -\frac{\partial \mu_d}{\partial \mu_I} = 1.$$
(6)

In the same manner, one can show that $\frac{\partial \mu_u}{\partial \mu_B} = \frac{\partial \mu_d}{\partial \mu_B} = \frac{\partial \mu_s}{\partial \mu_B} = \frac{1}{3}$, $\frac{\partial \mu_u}{\partial \mu_S} = \frac{\partial \mu_d}{\partial \mu_S} = 0$, and $\frac{\partial \mu_s}{\partial \mu_S} = -1$. This gives the following relations between the chemical potentials μ_B , μ_I , and μ_S and the quark chemical potentials μ_q :

$$\mu_u = \frac{1}{3}\mu_B + \mu_I,\tag{7}$$

$$\mu_d = \frac{1}{3}\mu_B - \mu_I,\tag{8}$$

$$\mu_s = \frac{1}{3}\mu_B - \mu_S. \tag{9}$$

In the chiral (Weyl) representation, we can write the Dirac operator $(D + m - \mu_a \gamma_0)$ for three flavors as

$$\begin{pmatrix} m & iX - \frac{1}{3}\mu_B - \mu_I & 0 & 0 & 0 & 0 \\ iX^{\dagger} - \frac{1}{3}\mu_B - \mu_I & m & 0 & 0 & 0 \\ 0 & 0 & m & iX - \frac{1}{3}\mu_B + \mu_I & 0 & 0 \\ 0 & 0 & iX^{\dagger} - \frac{1}{3}\mu_B + \mu_I & m & 0 & 0 \\ 0 & 0 & 0 & 0 & m & iX - \frac{1}{3}\mu_B + \mu_S \\ 0 & 0 & 0 & 0 & iX^{\dagger} - \frac{1}{3}\mu_B + \mu_S & m \end{pmatrix},$$
(10)

where $iX = D_0 + i\boldsymbol{\sigma} \cdot \mathbf{D}$. The fermion determinant then becomes

$$det(\mathcal{D} + m - \mu_q \gamma_0) = det \left[\left(X^{\dagger} + \frac{1}{3} i\mu_B + i\mu_I \right) \left(X + \frac{1}{3} i\mu_B + i\mu_I \right) + m^2 \right]$$

$$\times det \left[\left(X^{\dagger} + \frac{1}{3} i\mu_B - i\mu_I \right) \left(X + \frac{1}{3} i\mu_B - i\mu_I \right) + m^2 \right]$$

$$\times det \left[\left(X^{\dagger} + \frac{1}{3} i\mu_B - i\mu_S \right) \left(X + \frac{1}{3} i\mu_B - i\mu_S \right) + m^2 \right].$$
(11)

The terms proportional to μ_B and μ_S appear in the same way in combination with X^{\dagger} and X. Consequently, the fermion determinant is real only for $\mu_B = \mu_S = 0$. Using Eqs. (7)–(9), this yields the constraints

$$\mu_u + \mu_d = 0, \tag{12}$$

$$\mu_s = 0. \tag{13}$$

Given the two constraints, there is only one independent chemical potential, for example, the isospin chemical potential $\mu_I = \frac{1}{2}(\mu_u - \mu_d)$. The fermion determinant reduces to

$$det(\vec{D} + m - \mu_q \gamma_0) = det[(X^{\dagger} + i\mu_I)(X + i\mu_I) + m^2]$$
$$\times det[(X^{\dagger} - i\mu_I)(X - i\mu_I) + m^2]$$
$$\times det[X^{\dagger}X + m^2].$$
(14)

We conclude that the fermion determinant is real even for nonzero isospin chemical potential, and this proves that there is no sign problem for $\mu_I \neq 0$.

III. HARD-THERMAL-LOOP PERTURBATION THEORY

In this section, we briefly review hard-thermal-loop perturbation theory. For a detailed discussion, see, for example, Ref. [23]. Hard-thermal-loop perturbation theory is a reorganization of perturbation theory for thermal QCD. The HTLpt Lagrangian density is written as

$$\mathcal{L} = (\mathcal{L}_{\text{QCD}} + \mathcal{L}_{\text{HTL}})|_{q \to \sqrt{\delta}q} + \Delta \mathcal{L}_{\text{HTL}}, \qquad (15)$$

where the HTL improvement term is [67]

$$\mathcal{L}_{\text{HTL}} = (1 - \delta) i m_q^2 \bar{\psi} \gamma^{\mu} \left\langle \frac{y_{\mu}}{y \cdot D} \right\rangle_{\hat{y}} \psi$$
$$- \frac{1}{2} (1 - \delta) m_D^2 \text{Tr} \left(G_{\mu\alpha} \left\langle \frac{y^{\alpha} y_{\beta}}{(y \cdot D)^2} \right\rangle_{\hat{y}} G^{\mu\beta} \right), \quad (16)$$

and $\Delta \mathcal{L}_{HTL}$ contains additional HTLpt counterterms. Here $y^{\mu} = (1, \hat{\mathbf{y}})$ is a lightlike four-vector with $\hat{\mathbf{y}}$ being a threedimensional unit vector, and the angular bracket indicates an average over the direction of $\hat{\mathbf{y}}$. The two parameters m_D and m_q can be identified with the Debye screening mass and the thermal quark mass, respectively, and account for screening effects. HTLpt is defined by treating δ as a formal expansion parameter. The HTLpt Lagrangian (15) reduces to the QCD Lagrangian if we set $\delta = 1$. Physical observables are calculated in HTLpt by expanding in powers of δ , truncating at some specified order, and setting $\delta = 1$ in the end. This defines a reorganization of the perturbative series in which the effects of m_D^2 and m_q^2 terms in (16) are included with leading order but then are systematically subtracted out at higher orders. Note that HTLpt is gauge invariant order-byorder in the δ expansion, and, consequently, the results obtained are independent of the gauge-fixing parameter ξ (in the class of covariant gauges we are using). To zeroth order in δ , HTLpt describes a gas of massive gluonic and quark quasiparticles. Thus, HTLpt systematically shifts the perturbative expansion from being around an ideal gas of massless particles to being around a gas of massive quasiparticles which are the appropriate physical degrees of freedom at high temperature and/or chemical potential.

Higher orders in δ describe the interaction among these quasiparticles and involve standard QCD Feynman diagrams as well as new diagrams generated by the HTL improvement term. If the expansion in δ could be calculated to all orders, the final result would not depend on m_D and m_q when we set

 $\delta = 1$. However, any truncation of the expansion in δ produces results that depend on m_D and m_q . As a consequence, a prescription is required to determine m_D and m_q as a function of T, μ_q , and α_s . Several prescriptions were discussed in [20] at zero chemical potential and generalized to finite chemical potential in [23]. We return to this issue below.

A. NNLO HTLpt thermodynamic potential

The QCD free energy to three-loop order in HTLpt for the case that each quark f has a separate quark chemical potential μ_f was calculated in [23]. The final result is

$$\begin{split} & \underline{\Omega}_{\mathbf{N}\mathbf{n}\mathbf{O}} = \frac{7\,d_r}{4d_A} \frac{1}{N_f} \sum_{f} \left(1 + \frac{120}{7} \hat{\mu}_f^2 + \frac{240}{7} \hat{\mu}_f^4 \right) - \frac{s_F\alpha_s}{\pi} \frac{1}{N_f} \sum_{f} \left[\frac{8}{8} (1 + 12\hat{\mu}_f^2) (5 + 12\hat{\mu}_f^2) - \frac{15}{2} (1 + 12\hat{\mu}_f^2) \hat{m}_D \right] \\ & - \frac{15}{2} \left(2\ln\frac{\hat{\lambda}}{2} - 1 - \aleph(z_f) \right) \hat{m}_D^3 + 90 \hat{m}_q^2 \hat{m}_D \right] + \frac{s_{2F}}{N_f} \left(\frac{a_s}{\pi} \right)^2 \sum_{f} \left[\frac{15}{64} \left\{ 35 - 32(1 - 12\hat{\mu}_f^2) \frac{\xi''(-1)}{\zeta(-1)} + 472\hat{\mu}_f^2 + 1328\hat{\mu}_f^4 \right. \\ & + 64(-36\hat{\mu}_f \aleph(2, z_f) + 6(1 + 8\hat{\mu}_f^2) \aleph(1, z_f) + 3\hat{\mu}_f (1 + 4\hat{\mu}_f^2) \aleph(0, z_f)) \right\} - \frac{45}{2} \hat{m}_D (1 + 12\hat{\mu}_f^2) \right] \\ & + \left(\frac{s_F\alpha_s}{\pi} \right)^2 \frac{1}{N_f} \sum_{f} \frac{5}{16} \left[96(1 + 12\hat{\mu}_f^2) \frac{\hat{m}_q^2}{\hat{m}_D} + \frac{4}{3} (1 + 12\hat{\mu}_f^2) (5 + 12\hat{\mu}_f^2) \ln \frac{\hat{\lambda}}{2} \right] \\ & + \left(\frac{s_F\alpha_s}{\pi} \right)^2 \frac{1}{N_f} \sum_{f} \frac{5}{16} \left[96(1 + 12\hat{\mu}_f^2) \frac{\hat{m}_q^2}{\hat{m}_D} + \frac{4}{3} (1 + 12\hat{\mu}_f^2) (5 + 12\hat{\mu}_f^2) \ln \frac{\hat{\lambda}}{2} \right] \\ & + \frac{1}{3} + 4\gamma_E + 8(7 + 12\gamma_E) \hat{\mu}_f^2 + 112\hat{\mu}_f^4 - \frac{64\zeta'(-3)}{15\zeta(-3)} - \frac{32}{3} (1 + 12\hat{\mu}_f^2) \frac{\zeta'(-1)}{\zeta(-1)} \\ & - 96\{8\aleph(3, z_f) + 12\hat{\mu}_f \aleph(2, z_f) - 2(1 + 2\hat{\mu}_f^2) \aleph(1, z_f) - i\hat{\mu}_f \aleph(0, z_f) \} \right] \\ & + \left(\frac{s_F\alpha_s}{\pi} \right)^2 \frac{1}{N_f^2} \sum_{f} \left[\frac{5}{4\hat{m}_D} (1 + 12\hat{\mu}_f^2) (1 + 12\hat{\mu}_g^2) + 90\{2(1 + \gamma_E)\hat{\mu}_f^2\hat{\mu}_g^2} \\ & - \{\aleph(3, z_f + z_g) + \aleph(3, z_f + z_g) - (\hat{\mu}_f - \hat{\mu}_g)^2 \aleph(1, z_f + z_g) - 4\hat{\mu}_f \hat{\mu}_g^2 \Re(0, z_f) \} \right] - \frac{4}{2} \frac{1}{2} (1 + 12\hat{\mu}_f^2) \left(2\ln\frac{\hat{\lambda}}{2} - 1 - \aleph(z_g) \right) \hat{m}_D \right] \\ & + \left(\frac{c_A\alpha_s}{3\pi} \right) \left(\frac{s_F\alpha_s}{\pi N_f} \right) \sum_{f} \left[\frac{15}{2\hat{m}_D} (1 + 12\hat{\mu}_f^2) - \frac{235}{16} \left\{ \left(1 + \frac{792}{4\hat{\mu}} \hat{\mu}_f^2 + \frac{1584}{4\hat{\eta}} \hat{\mu}_f^4 \right) \ln \hat{\lambda}_2 \right] \\ & - \frac{144}{47} (1 + 12\hat{\mu}_f^2) \ln \hat{m}_D + \frac{319}{340} \left(1 + \frac{2040}{319} \hat{\mu}_f^2 + \frac{38640}{319} \hat{\mu}_f^4 \right) - \frac{24\gamma_E}{47} (1 + 12\hat{\mu}_f^2) \left(\ln \frac{1}{2} \right) \\ & - \frac{44}{47} \left(1 + \frac{156}{11} \hat{\mu}_f^2 \right) \frac{\xi''(-1)}{235\zeta(-3)} - \frac{72}{247} \left[4\hat{\mu}_f \Re(0, z_f) + (5 - 92\hat{\mu}_f^2) \Re(1, z_f) \right] \\ & + 144\hat{\mu}_f \Re(2, z_f) + 52\aleph(3, z_f) \right] \right\} + 0 \hat{\frac{m}_B}{\hat{m}_D} + \frac{314}{41} \left\{ \left(1 + \frac{132}{7} \hat{\mu}_f^2 \right) \ln \hat{\lambda}_2 \right] \\ & + \frac{14}{7} (1 + 12\hat{\mu}_f^$$

where $\Omega_0 = -\frac{d_A \pi^2 T^4}{45}$, $\hat{\mu}_f = \mu_f / 2\pi T$, $\hat{\Lambda} = \Lambda / 2\pi T$, and $\hat{m}_D = m_D / 2\pi T$. The QCD Casimir numbers are $c_A = N_c$, $d_A = N_c^2 - 1$, $s_F = N_f / 2$, $d_F = N_c N_f$, and $s_{2F} = C_F s_F$ with $C_F = (N_c^2 - 1) / 2N_c$. The sums over f and g include all quark flavors, $z_f = 1/2 - i\hat{\mu}_f$, and $\Omega_{\text{NNLO}}^{\text{YM}}$ is the pure-glue contribution

$$\frac{\Omega_{\rm NNLO}^{\rm YM}}{\Omega_0} = 1 - \frac{15}{4} \hat{m}_D^3 + \frac{c_A \alpha_s}{3\pi} \left[-\frac{15}{4} + \frac{45}{2} \hat{m}_D - \frac{135}{2} \hat{m}_D^2 - \frac{495}{4} \left(\ln \frac{\hat{\Lambda}_g}{2} + \frac{5}{22} + \gamma_E \right) \hat{m}_D^3 \right] \\
+ \left(\frac{c_A \alpha_s}{3\pi} \right)^2 \left[\frac{45}{4 \hat{m}_D} - \frac{165}{8} \left(\ln \frac{\hat{\Lambda}_g}{2} - \frac{72}{11} \ln \hat{m}_D - \frac{84}{55} - \frac{6}{11} \gamma_E - \frac{74}{11} \frac{\zeta'(-1)}{\zeta(-1)} + \frac{19}{11} \frac{\zeta'(-3)}{\zeta(-3)} \right) \\
+ \frac{1485}{4} \left(\ln \frac{\hat{\Lambda}_g}{2} - \frac{79}{44} + \gamma_E + \ln 2 - \frac{\pi^2}{11} \right) \hat{m}_D \right].$$
(18)

THREE-LOOP HARD-THERMAL-LOOP PERTURBATION ...

In Eq. (17), the functions $\aleph(z)$ and $\aleph(n, z)$ appear. These are defined as

$$\aleph(z) = \Psi(z) + \Psi(z^*), \tag{19}$$

where

$$\zeta'(x, y) = \partial_x \zeta(x, y), \tag{21}$$

$$\Psi(z) = \frac{\Gamma'(z)}{\Gamma(z)}.$$
(22)

Here $\zeta(x, y)$ is the Riemann zeta function and $\Gamma(z)$ is the digamma function.

B. Mass prescription

In order to complete a calculation in HTLpt, we must have a prescription for the mass parameters m_D and m_q appearing in the HTL Lagrangian. A variational prescription seems natural; i.e., one looks for solutions of

$$\frac{\partial}{\partial m_D} \Omega(T, \alpha_s, m_D, m_q, \mu_q, \delta = 1) = 0, \qquad (23)$$

$$\frac{\partial}{\partial m_q} \Omega(T, \alpha_s, m_D, m_q, \mu_q, \delta = 1) = 0.$$
(24)

However, in some cases the resulting gap equations only have complex solutions and one must look for other prescriptions. Inspired by dimensions reduction, one equates the mass parameter m_D with the mass parameter of threedimensional electric QCD in [68]. This mass can be interpreted as the contribution to the Debye mass from the hard scale T and is well defined and gauge invariant order-by-order in perturbation theory. This prescription was used in Ref. [23] and will be used in the remainder of the paper as well. Originally, the two-loop perturbative mass was calculated in Ref. [68] for zero chemical potential; however, Vuorinen has generalized it to finite chemical potential. The resulting expression for \hat{m}_D^2 is [2,3]

$$\hat{m}_{D}^{2} = \frac{\alpha_{s}}{3\pi} \left\{ c_{A} + \frac{c_{A}^{2}\alpha_{s}}{12\pi} \left(5 + 22\gamma_{E} + 22\ln\frac{\hat{\Lambda}_{g}}{2} \right) + \frac{1}{N_{f}} \sum_{f} \left[s_{F}(1 + 12\hat{\mu}_{f}^{2}) + \frac{c_{A}s_{F}\alpha_{s}}{12\pi} \left((9 + 132\hat{\mu}_{f}^{2}) + 22(1 + 12\hat{\mu}_{f}^{2})\gamma_{E} + 2(7 + 132\hat{\mu}_{f}^{2})\ln\frac{\hat{\Lambda}}{2} + 4\aleph(z_{f}) \right) + \frac{s_{F}^{2}\alpha_{s}}{3\pi} (1 + 12\hat{\mu}_{f}^{2}) \left(1 - 2\ln\frac{\hat{\Lambda}}{2} + \aleph(z_{f}) \right) - \frac{3}{2} \frac{s_{2F}\alpha_{s}}{\pi} (1 + 12\hat{\mu}_{f}^{2}) \right] \right\}.$$

$$(25)$$

The effect of the in-medium quark mass parameter m_q in thermodynamic functions is small, and following Ref. [20], we take $m_q = 0$.

IV. NUMERICAL RESULTS

In this section, we present our results for the NNLO HTLpt thermodynamic functions at finite temperature *T* and isospin chemical potential μ_I , and $\mu_B = \mu_S = 0$. We emphasize that all thermodynamic functions can be calculated for nonzero values of the three independent chemical potentials.

A. Running coupling and scales

In Ref. [20], we showed that the renormalization of the three-loop HTLpt free energy is consistent with the standard one-loop running of the strong coupling constant [69,70]. Using a one-loop running is therefore selfconsistent and will be used in the remainder of this paper.¹ In this case, the running coupling $\alpha_s(\Lambda)$ is given by

$$\alpha_s(\Lambda) = \frac{1}{b_0 t},\tag{26}$$

with $t = \ln(\Lambda^2/\Lambda_{\overline{\text{MS}}}^2)$ and $b_0 = (11c_A - 2N_f)/12\pi$. We fix the scale $\Lambda_{\overline{\text{MS}}}$ by requiring that $\alpha_s(1.5 \text{ GeV}) = 0.326$ which is obtained from independent lattice measurements [71]. For one-loop running, this procedure gives $\Lambda_{\overline{\text{MS}}} =$ 176 MeV.

For the renormalization scale we use separate scales, Λ_g and Λ_q , for purely gluonic and fermionic graphs, respectively. We take the central values of these renormalization scales to be $\Lambda_g = 2\pi T$ and $\Lambda = \Lambda_q = 2\pi \sqrt{T^2 + (\mu_B^2 + 2\mu_I^2)/(N_f\pi^2)}$. In all plots, the thick lines indicate the result obtained using these central values, and the light-blue band indicates the variation of the result under the variation of both of these scales by a factor of 2, e.g., $\pi T \leq \Lambda_g \leq 4\pi T$. For all numerical results below we use $c_A = N_c = 3$ and $N_f = 3$.

Since our final result for the thermodynamic potential (17) and the thermodynamic functions that are derived from it are expansions in m_D/T and m_q/T , we cannot push our results to very high values of μ_I ; the Debye mass in Eq. (25) depends on the quark chemical potentials μ_f . An estimate for the reliability of HTLpt is that $m_D \simeq gT$. If $T < \sqrt{3}\mu_f/\pi$, the μ_f -dependent term of m_D just starts to dominate over the *T*-dependent term. Thus we consider $\mu_f \lesssim \pi T$ as reasonable. For temperatures down to 150 MeV, we decide to err on the safe side and use μ_I no larger than 400 MeV.

¹In our previous paper [23], we used one-loop running as well as three-loop running to gauge the sensitivity of our results. Generally, our three-loop HTLpt predictions were rather insensitive to whether we used one-loop or three-loop running.

B. Pressure

The pressure of the quark-gluon plasma can be obtained directly from the thermodynamic potential (17)

$$\mathcal{P}(T,\Lambda,\mu_u,\mu_d,\mu_s) = -\Omega(T,\Lambda,\mu_u,\mu_d,\mu_s), \quad (27)$$

PHYSICAL REVIEW D 93, 054045 (2016)

where Λ includes both Λ_g and Λ_q . The pressure can be obtained using our general expression Eq. (17) for nonzero values of μ_B and μ_I and for $\mu_S = 0$ using Eqs. (7)–(9). For simplicity, we are presenting here the NNLO HTL pressure only at the nonzero value of μ_I and for $\mu_B = \mu_S = 0$ as

$$\begin{aligned} \frac{\mathcal{P}_{\text{NLO}}}{\mathcal{P}_{0}} &= \frac{7\,d_{r}}{4d_{A}\,N_{f}} \left(N_{f} + \frac{240}{7}\,\hat{\mu}_{f}^{2} + \frac{480}{7}\,\hat{\mu}_{f}^{4}\right) - \frac{s_{F}\alpha_{s}}{\pi}\frac{1}{N_{f}} \left[\frac{5}{8}(5N_{f} + 144\hat{\mu}_{f}^{2} + 288\hat{\mu}_{f}^{4}) - \frac{15}{2}(N_{f} + 24\hat{\mu}_{f}^{2})\hat{m}_{D} \\ &\quad -\frac{15}{2}\left(2N_{f}\ln\frac{\hat{\lambda}}{2} - N_{f} + 2(2\log 2 + \gamma_{E}) - 2\aleph(z_{I})\right)\hat{m}_{D}^{3} + 90\hat{m}_{g}^{2}\hat{m}_{D}\right] + \frac{s_{F}}{N_{f}}\left(\frac{\alpha_{s}}{\pi}\right)^{2} \\ &\quad \times \left[\frac{15}{64}\left\{35N_{f} - 32\log 2 - 32(N_{f} - 1 - 24\hat{\mu}_{I}^{2})\frac{\xi'(-1)}{\zeta(-1)} + 944\hat{\mu}_{I}^{2} + 2656\hat{\mu}_{I}^{4} - 384(12i\hat{\mu}_{I}\aleph(2, z_{I}) \right. \\ &\quad -2(1 + 8\hat{\mu}_{I}^{2})\aleph(1, z_{I}) - i\hat{\mu}_{I}(1 + 4\hat{\mu}_{I}^{2})\aleph(0, z_{I}))\right\} - \frac{45}{2}\hat{m}_{D}(N_{f} + 24\hat{\mu}_{I}^{2})\right] \\ &\quad + \left(\frac{s_{F}\alpha_{s}}{\pi}\right)^{2}\frac{1}{N_{f}}\frac{5}{16}\left[96(N_{f} + 24\hat{\mu}_{I}^{2})\frac{\hat{m}_{g}^{2}}{\hat{m}_{D}} + \frac{4}{3}(5N_{f} + 144\hat{\mu}_{I}^{2} + 288\hat{\mu}^{4})\ln\frac{\hat{\lambda}}{2} + \frac{N_{f}}{3}(1 + 12\gamma_{E}) \right. \\ &\quad + 16(7 + 12\gamma_{E})\hat{\mu}_{I}^{2} + 224\hat{\mu}_{I}^{4} - \frac{8}{15N_{f}}(8N_{f}^{2} - 21N_{f} + 54)\frac{\xi'(-3)}{\zeta(-3)} - \frac{16}{N_{f}}(N_{f}(1 + 16\hat{\mu}_{I}^{2}) - 12\hat{\mu}_{I}^{2})\frac{\xi'(-1)}{\zeta(-1)} \\ &\quad -\frac{8}{5N_{f}}(11N_{f} + 120\hat{\mu}_{I}^{2})\log 2 - 192\{8\aleph(3, z_{I}) + 12i\hat{\mu}_{I}\aleph(2, z_{I}) - 2(1 + 2\hat{\mu}_{I}^{2})\aleph(1, z_{I}) - i\hat{\mu}_{I}\aleph(0, z_{I})\}\right] \\ &\quad + \left(\frac{\epsilon_{F}\alpha_{s}}{\pi}\right)^{2}\frac{1}{N_{f}^{2}}\left[\frac{5}{4\hat{m}_{D}}(N_{f} + 24\hat{\mu}_{I}^{2})^{2} + 360\{2(1 + \gamma_{E})\hat{\mu}_{I}^{4} - \aleph(3, 2z_{I}) - 2\aleph(3, z_{I} + z_{0}) \right. \\ &\quad -4i\hat{\mu}_{I}[\aleph(2, 2z_{I}) + \aleph(2, z_{I} + z_{0})] + 4\hat{\mu}_{I}^{2}\Re(1, z_{I}) + 2\hat{\mu}_{I}^{2}(2\aleph(1, 2z_{I}) + \aleph(1, z_{I} + z_{0})) + 4i\hat{\mu}_{I}^{3}\aleph(0, z_{I})\} \right] \\ &\quad -\frac{15}{2}(N_{f} + 24\hat{\mu}_{I}^{2})\left(2N_{f}\ln\frac{\hat{\lambda}}{2} - N_{f} + 2(2\log 2 + \gamma_{E}) - 2\aleph(z_{I})\right)\hat{m}_{D}\right] + \left(\frac{\epsilon_{A}\alpha_{s}}{3\pi}\right)\left(\frac{s_{F}\alpha_{s}}{3\ln}\right)\left(\frac{s_{F}\alpha_{s}}{319}\hat{\mu}_{I}^{4}\right) \\ &\quad -\frac{24\gamma_{E}}{47}(N_{f} + 24\hat{\mu}_{I}^{2})\left(-\frac{2}{47}(22N_{f} + 15 + 624\hat{\mu}_{I}^{2})\frac{\zeta'(-1)}{\zeta(-1)} - \frac{1}{23}(268N_{f} - 273)\frac{\zeta'(-3)}{\zeta(-3)} + \frac{113}{235}\log 2 \right. \\ \\ &\quad -\frac{144}{47}\left[4\hat{\mu}_{I}(\aleph(0, z_{I}) + (5 - 92\hat{\mu}_{I}^{2})\aleph(1, z_{I}) + 144\hat{\mu}_{I}(\aleph(2, z_{I}) + 52\Re(3, z_{I})]\right\right\} + 90\frac{\hat{m}_{R}^{2}}{\frac{315}{4$$

In Fig. 1, we show the NNLO pressure obtained using HTLpt as a function of *T* normalized to that of an ideal gas of massless particles for $\mu_I = 200$ MeV, $\mu_B = 0$ (left) and $\mu_I = 200$ MeV, $\mu_B = 400$ MeV (right). The pressure is an increasing function of *T*, but stays well below the ideal-gas value even for the highest temperatures shown.

In Fig. 2, we show the normalized NNLO pressure of HTLpt as a function of *T* for four different values of the isospin chemical potential μ_I . We notice that the pressure is an increasing function of μ_I for fixed temperature and that

the pressure curves converge at a temperature of approximately 800 MeV.

C. Energy density

Once we know the pressure \mathcal{P} , we can calculate the energy density \mathcal{E} by the Legendre transform

$$\mathcal{E} = T \frac{\partial \mathcal{P}}{\partial T} + \mu_q \frac{\partial \mathcal{P}}{\partial \mu_q} - \mathcal{P} = T \frac{\partial \mathcal{P}}{\partial T} + \mu_I \frac{\partial \mathcal{P}}{\partial \mu_I} - \mathcal{P}, \quad (29)$$

THREE-LOOP HARD-THERMAL-LOOP PERTURBATION ...



FIG. 1. The pressure normalized to that of an ideal gas of massless particles as a function of *T*. The left part is for $\mu_I = 200$ MeV, $\mu_B = 0$, and the right part is for $\mu_I = 200$ MeV, $\mu_B = 400$ MeV.



FIG. 2. The pressure normalized to that of an ideal gas of massless particles as a function of *T* for various values of the isospin chemical potential μ_I at $\mu_B = \mu_S = 0$ (left) and $\mu_B = 400$ MeV, $\mu_S = 0$ (right). Here $\Lambda_g = 2\pi T$ and $\Lambda_q = 2\pi \sqrt{T^2 + (\mu_B^2 + 2\mu_I^2)/(3\pi^2)}$ were used.

where we have used that $\mu_I = \frac{1}{2}(\mu_u - \mu_d)$ and $\mu_s = 0$. In Fig. 3, we show the energy density as a function of the temperature for $\mu_I = 0$ (left) and $\mu_I = 200$ MeV (right). As in the case of the pressure, the energy density is an increasing function of *T* and stays well below the ideal-gas value for all temperatures.

In Fig. 4, we show the normalized energy density for four different values of the isospin chemical potential μ_I . For $\mu_I = 0$ or $\mu_I = 200$, the energy density is an increasing function of *T*. Note, however, that there is a minimum for the energy density for low temperatures and higher values of the isospin chemical potential. We would like to mention here that HTLpt probably cannot be trusted at these low temperatures with a large chemical potential, and one cannot attribute any interesting physics to this nonmonotonic behavior. Likewise, the curves converge at high temperatures, here already at approximately T = 600 MeV.

D. Trace anomaly

The trace anomaly or interaction measure \mathcal{I} is defined by the difference

$$\mathcal{I} = \mathcal{E} - 3\mathcal{P}.\tag{30}$$

For an ideal gas of massless particles, the trace anomaly vanishes since $\mathcal{E} = 3\mathcal{P}$. For massless particles and nonzero g, \mathcal{I} is nonzero and is a measure of the interactions in the plasma.² In Fig. 5, we show the interaction measure as a function of the temperature for two different values of the

²For nonzero current quark masses m_0 , $\mathcal{I} \neq 0$ even in the absence of interactions.



FIG. 3. The energy density normalized to that of an ideal gas of massless particles as a function of *T*. The left part is for $\mu_I = 0$, and the right part is for $\mu_I = 200$ MeV. $\mu_B = \mu_S = 0$ in both plots.



FIG. 4. The energy density normalized to that of an ideal gas of massless particles as a function of *T* for four different values of the isospin chemical potential and $\mu_B = \mu_S = 0$. Here $\Lambda_g = 2\pi T$ and $\Lambda_g = 2\pi \sqrt{T^2 + 2\mu_I^2/(3\pi^2)}$ were used.

isospin chemical potential, $\mu_I = 0$ (left) and $\mu_I = 200$ MeV (right). The trace anomaly is a decreasing function of *T*, and it converges to zero for large values of *T* due to asymptotic freedom.

In Fig. 6, we show the normalized interaction measure as a function of the temperature T for four different values of the isospin chemical potential μ_I . As the figure demonstrates, the curves are essentially identical.

E. Speed of sound

The speed of sound c_s is defined by

$$c_s^2 = \frac{\partial \mathcal{P}}{\partial \mathcal{E}}.$$
 (31)

In Fig. 7, we show the speed of sound squared c_s^2 for two different values of the isospin chemical potential, $\mu_I = 0$ (left) and $\mu_I = 200$ MeV (right). The horizontal dotted line is the ideal-gas value $c_s^2 = \frac{1}{3}$. As this figure



FIG. 5. Trace anomaly divided by T^4 as a function of the temperature *T*. The left part is for $\mu_I = 0$, and the right part is for $\mu_I = 200$ MeV. $\mu_B = \mu_S = 0$ in both plots.

demonstrates, the speed of sound is an increasing function of T.

In Fig. 8, we show the speed of sound squared c_s^2 for four different values of the isospin chemical potential μ_I . We notice that the speed of sound is an increasing function of μ_I for fixed *T* and that the curves converge rather quickly, here at approximately T = 400 MeV.

F. Susceptibilities

Using the thermodynamic potential given by Eq. (28), we can compute the quark-number susceptibilities. In the most general case, we have one quark chemical potential μ_f for each quark flavor f, which we can organize in an N_f -dimensional vector $\boldsymbol{\mu} = (\mu_u, \mu_d, \mu_s, ..., \mu_{N_f})$. The single quark susceptibilities are defined by

$$\chi_{ijk\cdots}(T) = \frac{\partial^{i+j+k+\cdots}\mathcal{P}(T,\boldsymbol{\mu})}{\partial \mu_u^i \partial \mu_d^j \partial \mu_s^k \cdots} \Big|_{\boldsymbol{\mu} = \boldsymbol{\mu}_0}, \quad (32)$$

where μ_0 is a configuration of quark chemical potentials. When computing the derivatives with respect to the chemical potential, we will use $\mu_0 = 0$. We treat Λ_q as being a constant and only put the chemical potential dependence of Λ_q in after the derivatives are taken. We have done this in order to more closely match the procedure used to compute the susceptibilities using resummed dimensional reduction [2] and to ensure that the susceptibilities vanish when $N_f = 0$. In the following, we will use a shorthand notation for the quark susceptibilities by specifying derivatives by a string of quark



FIG. 6. Trace anomaly divided by T^4 as a function of the temperature *T* for four different values of the isospin chemical potential and $\mu_B = \mu_S = 0$. Here $\Lambda_g = 2\pi T$ and $\Lambda_q = 2\pi \sqrt{T^2 + 2\mu_I^2/(3\pi^2)}$ were used.

flavors using superscript. For example, $\chi_2^{uu} = \chi_{200}$, $\chi_2^{ds} = \chi_{011}$, and $\chi_4^{uudd} = \chi_{220}$. For a three-flavor system with (u, d, s) quarks with $\mu_B = \mu_S = 0$, the *n*th-order isospin number susceptibility evaluated at $\mu_I = 0$ is defined by

$$\chi_n^I \equiv \frac{\partial^n \mathcal{P}}{\partial \mu_I^n} \Big|_{\hat{\mu}_I = 0}.$$
 (33)

We can analytically express various order susceptibilities as

$$\begin{split} \chi_{2}^{l} &= \frac{1}{48} T^{2} (32N_{c} - 3d_{A} \hat{m}_{D}^{2} \hat{m}_{D}'') - \frac{d_{A} \alpha_{s} T^{2}}{96\pi} \bigg[16(3 - 6\hat{m}_{D} + 7\hat{m}_{D}^{3} \zeta(3)) - \bigg\{ 2(2c_{A} + N_{f}) - 24c_{A} \hat{m}_{D} \\ &- \bigg(15c_{A} - 36\gamma_{E} + 66c_{A} \gamma_{E} + 6N_{f} - 72\log 2 - 36\ln\frac{\hat{\Lambda}}{2} + 66c_{A} \ln\frac{\hat{\Lambda}}{2} \bigg) \hat{m}_{D}^{2} \bigg\} \hat{m}_{D}'' \bigg] \\ &+ \frac{d_{A} \alpha_{s}^{2} T^{2}}{144\pi^{2}} \bigg[\frac{24(3 + 2c_{A})}{\hat{m}_{D}} + 360 - 91c_{A} + 105c_{F} + 12c_{A} \gamma_{E} - 146N_{f} + 24\gamma_{E} N_{f} \\ &+ (-504 + 144c_{A} + 96c_{F} + 168N_{f}) \frac{\zeta'(-1)}{\zeta(-1)} - (72 + 112c_{A} - 48c_{F} + 8N_{f})\log 2 - (132c_{A} - 24N_{f})\log\frac{\hat{\Lambda}}{2} \\ &+ 288c_{A}\log\hat{m}_{D} - 12\bigg(36c_{F} + 36\gamma_{E} - 6N_{f} + 72\log 2 - 6(11c_{A} - 2N_{f})\log\frac{\hat{\Lambda}}{2} - 21\zeta(3) - c_{A}(33 + 66\gamma_{E} + 14\zeta(3)) \bigg) \hat{m}_{D} \\ &- \bigg\{ \frac{(3 + 2c_{A})^{2}}{4\hat{m}_{D}^{2}} + \frac{6(3 + 2c_{A})}{\hat{m}_{D}} + \frac{3}{4} \bigg(3c_{A}(9 + 14\gamma_{E} - 16\log 2) - c_{A}^{2}(79 - 44\gamma_{E} + 4\pi^{2} - 44\log 2) \\ &- 6(6c_{F} + 6\gamma_{E} - N_{f} + 12\log 2) + 2(21c_{A} + 22c_{A}^{2} - 6N_{f})\ln\frac{\hat{\Lambda}}{2} \bigg) \bigg\} \hat{m}_{D}' \bigg], \end{split}$$

ANDERSEN, HAQUE, MUSTAFA, and STRICKLAND

$$\chi_{4}^{I} = \frac{1}{\pi^{2}} \left(4N_{c} - \frac{d_{A}\hat{m}_{D}}{64} (6\hat{m}_{D}^{\prime\prime2} + \hat{m}_{D}\hat{m}_{D}^{\prime\prime}) \right) - \frac{d_{A}\alpha_{x}}{384\pi^{3}} \left[1152 - 5952\hat{m}_{D}^{3}\zeta(5) - (576 - 2016\hat{m}_{D}^{2}\zeta(3))\hat{m}_{D}^{\prime\prime} \right] \\
+ \left\{ 72c_{A} - \left(216\gamma_{E} - 90c_{A} - 396c_{A}\gamma_{E} - 36N_{f} + 432\log 2 + 216\ln\frac{\hat{\Lambda}}{2} - 396c_{A}\ln\frac{\hat{\Lambda}}{2} \right) \hat{m}_{D} \right\} \hat{m}_{D}^{\prime\prime2} \\
- \left\{ 4c_{A} + 2N_{f} - 24c_{A}\hat{m}_{D} - \left(15c_{A} + 6N_{f} - 72\log 2 - 6(6 - 11c_{A}) \left(\gamma_{E} + \ln\frac{\hat{\Lambda}}{2} \right) \right) \hat{m}_{D}^{2} \right\} \hat{m}_{D}^{\prime\prime\prime} \right] \\
+ \frac{d_{A}\alpha_{x}^{2}}{24\pi^{4}} \left[\frac{144}{\hat{m}_{D}} - 36 + 21c_{F}(5 + 4\zeta(3)) + 72\gamma_{E} + 2N_{f} \left(11 + 14\zeta(3) + 12\ln\frac{\hat{\Lambda}}{2} \right) - 576\log 2 \\
- 11c_{A} \left(17 + 12\gamma_{E} + 24\log 2 - 7\zeta(3) + 12\ln\frac{\hat{\Lambda}}{2} \right) + 6(168\zeta(3) - 93\zeta(5) - 62c_{A}\zeta(5))\hat{m}_{D} \\
+ \left\{ - \frac{6(3 + 2cA)}{\hat{m}_{D}^{2}} + \frac{72c_{A}}{\hat{m}_{D}} + 3(-36c_{F} - 36\gamma_{E} + 6N_{f} - 72\log 2 + 6(11c_{A} - 2N_{f})\ln\frac{\hat{\Lambda}}{2} \right) \\
+ 21\zeta(3) + c_{A}(33 + 66\gamma_{E} + 14\zeta(3)) \right\} \hat{m}_{D}^{\prime\prime} + \frac{(3 + 2c_{A})}{16\hat{m}_{D}^{3}} \left\{ (3 + 2c_{A}) - 12c_{A}\hat{m}_{D} \right\} \hat{m}_{D}^{\prime\prime2} \\
- \left\{ \frac{(3 + 2c_{A})^{2}}{96\hat{m}_{D}^{\prime}} - \frac{c_{A}(3 + 2c_{A})}{4\hat{m}_{D}} - \frac{1}{32} \left(3c_{A}(9 + 14\gamma_{E} - 16\log 2) + c_{A}^{2}(-79 + 44\gamma_{E} - 4\pi^{2} + 44\log 2) \right) \\
- 6(6c_{F} + 6\gamma_{E} - N_{f} + 12\log 2) + 2(21c_{A} + 22c_{A}^{2} - 6N_{f})\ln\frac{\hat{\Lambda}}{2} \right\} \hat{m}_{D}^{\prime\prime} \right\}$$

where

$$\hat{m}_{D}^{\prime\prime} = \frac{\partial^{2} \hat{m}_{D}}{\partial \hat{\mu}_{I}^{2}} \Big|_{\hat{\mu}_{I}=0}^{\hat{\mu}_{I}=0} = \frac{\alpha_{s}}{9\pi^{2} \hat{m}_{D}} \left[36\pi + \alpha_{s} \left\{ 6(11N_{c} - 2N_{f}) \ln \frac{\hat{\Lambda}}{2} - 54c_{F} + N_{f}(6 - 12\gamma_{E} - 24\log 2 + 7\zeta(3)) + N_{c}(33 + 66\gamma_{E} + 14\zeta(3)) \right\} \right],$$
(36)

$$\hat{m}_{D}^{iv} = \frac{\partial^{4} \hat{m}_{D}}{\partial \hat{\mu}_{I}^{4}} \Big|_{\hat{\mu}_{I}=0}^{\mu} = -\frac{\alpha_{s}^{2}}{54\pi^{4} \hat{m}_{D}^{3}} \Big[2 \Big\{ 36\pi + \alpha_{s} \Big(6(11c_{A} - 2N_{f}) \ln \frac{\hat{\Lambda}}{2} - 54c_{F} + N_{f}(6 - 12\gamma_{E} - 24\log 2 + 7\zeta(3)) \\ + c_{A}(33 + 66\gamma_{E} + 14\zeta(3)) \Big) \Big\}^{2} - \alpha_{s} \Big\{ 12\pi(2c_{A} + N_{f}) + \alpha_{s} \Big(2(5 + 22\gamma_{E})c_{A}^{2} + 3c_{A}(9 + 14\gamma_{E} - 16\log 2) \\ - 6(9c_{F} + N_{f}(-1 + 2\gamma_{E} + 4\log 2)) + 2(21c_{A} + 22c_{A}^{2} - 6N_{f}) \ln \frac{\hat{\Lambda}}{2} \Big) \Big\} (84N_{f}\zeta(3) - 31(2c_{A} + N_{f})\zeta(5)) \Big].$$
(37)



FIG. 7. Speed of sound squared as a function of the temperature *T*. The left part is for $\mu_I = 0$, and the right part is for $\mu_I = 200$ MeV. $\mu_B = \mu_S = 0$ in both plots.



FIG. 8. Speed of sound squared as a function of the temperature *T* for different values of the isospin chemical potential μ_I and $\mu_B = \mu_S = 0$. Here $\Lambda_g = 2\pi T$ and $\Lambda_q = 2\pi \sqrt{T^2 + 2\mu_I^2/(3\pi^2)}$ were used.

For a three-flavor system consisting of (u, d, s) quarks, we can express the isospin susceptibilities in terms of the quark susceptibilities as

$$\chi_2^I = [\chi_2^{uu} + \chi_2^{dd} - 2\chi_2^{ud}], \qquad (38)$$

$$\chi_4^I = [\chi_4^{uuuu} + \chi_4^{dddd} - 4\chi_4^{uuud} - 4\chi_4^{dddu} + 6\chi_4^{uudd}].$$
 (39)

The isospin susceptibilities are expressed in terms of diagonal (same flavor on all indices) quark susceptibilities or off-diagonal (different flavor on some or all indices). In HTLpt, there are off-diagonal susceptibilities arising explicitly from some of the three-loop graphs [20,23]. There are also potential off-diagonal contributions coming from all HTL terms since the mass parameter m_D receives contributions from all quark flavors. However, these contributions vanish when we evaluate the susceptibilities at $\mu_f = 0$. In

this case, the HTLpt second- and fourth-order isospin susceptibilities reduce to

$$\chi_2^I = 2\chi_2^{uu},\tag{40}$$

$$\chi_4^I = [2\chi_4^{uuuu} + 6\chi_4^{uudd}]. \tag{41}$$

In Fig. 9, we show the HTLpt predictions for the isospin second- and fourth-order susceptibilities χ_2^I/T^2 and χ_4^I as functions of *T*. The horizontal dotted lines are the corresponding isospin susceptibilities for an ideal gas, indicated by the Stefan-Boltzmann limit. The central line for the second-order susceptibility is almost flat, while the central line for the fourth-order susceptibility is slowly increasing.

V. SUMMARY

In this paper, we presented results for a number of thermodynamic functions of QCD at finite temperature Tand finite isospin chemical potential μ_I using hard-thermalloop perturbation theory. The pressure was also calculated at nonzero baryon chemical potential μ_B . Our results were derived from the three-loop thermodynamic potential, which was computed in Ref. [23] as a function of temperature and quark chemical potentials. Our final results depend on two renormalization scales Λ_q and Λ_q which are expected to be approximately $2\pi T$ and $2\pi\sqrt{T^2 + (\mu_B^2 + 2\mu_I^2)/(3\pi^2)}$. In order to gauge the theoretical uncertainty associated with the scale choice, we varied both Λ_q and Λ_q by a factor of 2 (light-blue bands in some figures). We found that most quantities have a sizable scale variation and, at this moment in time, we do not have a method to reduce the size of the bands. A solution to this problem is suggested by the authors of Ref. [15]. In this approach, dubbed renormalization group optimized perturbation theory, the authors modify standard optimized perturbation theory or SPT. This is done by changing



FIG. 9. Second- and fourth-order susceptibilities as functions of the temperature T normalized to T^2 and one, respectively. $\mu_I = \mu_B = \mu_S = 0$ in both plots.

ANDERSEN, HAQUE, MUSTAFA, and STRICKLAND

the added/subtraction mass term, including a finite vacuum term, and imposing renormalization group invariance on the pressure. In the case of ϕ^4 -theory, the result for the pressure up to two-loop order is very stable and has narrow bands under a scale variation. Note, however, that some quantities, e.g., χ_4^I , have very small scale variations for temperatures $T \gtrsim 400$ MeV, and hence HTLpt provides testable predictions.

Given the relatively good agreement between lattice results and the predictions of NNLO HTLpt at zero and finite baryon chemical potential for $T \gtrsim 250$ MeV, we expect that the lattice results at finite μ_I should fall close to the central (black) lines predicted herein at high temperatures. We are looking forward to lattice measurements of QCD thermodynamics at finite μ_I and high temperatures (with $\mu_B = \mu_S = 0$) in order to test the predictions made herein. Since the necessary lattice measurements can be done without Taylor expansion, they would provide a highprecision test of NNLO HTLpt.

ACKNOWLEDGMENTS

N. H. was supported by an award from the Kent State University Office of Research and Sponsored Programs. M. G. M. was supported by the Indian Department of Atomic Energy under the project "Theoretical Physics across the energy scales." M. S. was supported by the U.S. Department of Energy under Award No. DE-SC0013470.

- K. Kajantie, M. Laine, K. Rummukainen, and Y. Schroder, Phys. Rev. D 67, 105008 (2003).
- [2] A. Vuorinen, Phys. Rev. D 67, 074032 (2003).
- [3] A. Vuorinen, Phys. Rev. D 68, 054017 (2003).
- [4] A. Ipp, K. Kajantie, A. Rebhan, and A. Vuorinen, Phys. Rev. D 74, 045016 (2006).
- [5] F. Karsch, A. Patkos, and P. Petreczky, Phys. Lett. B 401, 69 (1997).
- [6] S. Chiku and T. Hatsuda, Phys. Rev. D 58, 076001 (1998).
- [7] J. O. Andersen, E. Braaten, and M. Strickland, Phys. Rev. D 63, 105008 (2001).
- [8] J. O. Andersen and L. Kyllingstad, Phys. Rev. D 78, 076008 (2008).
- [9] V. I. Yukalov, Teor. Mat. Fiz. 26, 403 (1976).
- [10] P. M. Stevenson, Phys. Rev. D 23, 2916 (1981).
- [11] A. Duncan and M. Moshe, Phys. Lett. B 215, 352 (1988).
- [12] A. Duncan and H. F. Jones, Phys. Rev. D 47, 2560 (1993).
- [13] A. N. Sisakian, I. L. Solovtsov, and O. Shevchenko, Int. J. Mod. Phys. A 09, 1929 (1994).
- [14] W. Janke and H. Kleinert, Phys. Rev. Lett. 75, 2787 (1995).
- [15] J.-L. Kneur and M. B. Pinto, Phys. Rev. Lett. **116**, 031601 (2016); Phys. Rev. D **92**, 116008 (2015).
- [16] J. O. Andersen, E. Braaten, and M. Strickland, Phys. Rev. Lett. 83, 2139 (1999).
- [17] N. Su, J. O. Andersen, and M. Strickland, Phys. Rev. Lett. 104, 122003 (2010).
- [18] J. O. Andersen, M. Strickland, and N. Su, J. High Energy Phys. 08 (2010) 113.
- [19] J. O. Andersen, L. E. Leganger, M. Strickland, and N. Su, Phys. Lett. B 696, 468 (2011).
- [20] J. O. Andersen, L. E. Leganger, M. Strickland, and N. Su, J. High Energy Phys. 08 (2011) 053.
- [21] J. O. Andersen, L. E. Leganger, M. Strickland, and N. Su, Phys. Rev. D 84, 087703 (2011).
- [22] N. Haque, J. O. Andersen, M. G. Mustafa, M. Strickland, and N. Su, Phys. Rev. D 89, 061701 (2014).

- [23] N. Haque, A. Bandyopadhyay, J. O. Andersen, M. G. Mustafa, M. Strickland, and N. Su, J. High Energy Phys. 05 (2014) 027.
- [24] P. Chakraborty, M. G. Mustafa, and M. H. Thoma, Eur. Phys. J. C 23, 591 (2002).
- [25] P. Chakraborty, M. G. Mustafa, and M. H. Thoma, Phys. Rev. D 67, 114004 (2003).
- [26] P. Chakraborty, M. G. Mustafa, and M. H. Thoma, Phys. Rev. D 68, 085012 (2003).
- [27] N. Haque, M. G. Mustafa, and M. H. Thoma, Phys. Rev. D 84, 054009 (2011).
- [28] N. Haque and M. G. Mustafa, Nucl. Phys. A862-863, 271 (2011).
- [29] N. Haque and M. G. Mustafa, arXiv:1007.2076.
- [30] J. P. Blaizot, E. Iancu, and A. Rebhan, Eur. Phys. J. C 27, 433 (2003).
- [31] J. P. Blaizot, E. Iancu, and A. Rebhan, Phys. Lett. B **523**, 143 (2001).
- [32] J. P. Blaizot, E. Iancu, and A. Rebhan, Nucl. Phys. A698, 404 (2002).
- [33] J. P. Blaizot, E. Iancu, and A. Rebhan, Phys. Rev. Lett. 83, 2906 (1999).
- [34] J. P. Blaizot, E. Iancu, and A. Rebhan, Phys. Lett. B 470, 181 (1999).
- [35] J. P. Blaizot, E. Iancu, and A. Rebhan, Phys. Rev. D 63, 065003 (2001).
- [36] E. Dagotto, F. Karsch, and A. Moreo, Phys. Lett. 169B, 421 (1986); E. Dagotto, A. Moreo, and U. Wolff, Phys. Rev. Lett. 57, 1292 (1986).
- [37] S. Hands, J. B. Kogut, M.-P. Lombardo, and S. E. Morrison, Nucl. Phys. B558, 327 (1999).
- [38] M. G. Alford, A. Kapustin, and F. Wilczek, Phys. Rev. D 59, 054502 (1999).
- [39] J. B. Kogut and D. K. Sinclair, Phys. Rev. D 66, 014508 (2002).
- [40] J. B. Kogut and D. K. Sinclair, Phys. Rev. D 66, 034505 (2002).

- [41] J. B. Kogut and D. K. Sinclair, Phys. Rev. D 70, 094501 (2004).
- [42] C. R. Allton, M. Doring, S. Ejiri, S. J. Hands, O. Kaczmarek, F. Karsch, E. Laermann, and K. Redlich, Phys. Rev. D 71, 054508 (2005).
- [43] S. Ejiri, F. Karsch, and K. Redlich, Phys. Lett. B 633, 275 (2006).
- [44] P. de Forcrand, M. A. Stephanov, and U. Wenger, *Proc. Sci.*, LAT2007 (2007) 237.
- [45] W. Detmold, K. Origonos, and Z. Shi, Phys. Rev. D 86, 054507 (2012).
- [46] G. Endrődi, Phys. Rev. D 90, 094501 (2014).
- [47] D. T. Son and M. A. Stephanov, Phys. At. Nucl. 64, 834 (2001).
- [48] S. Borsanyi, G. Endrodi, Z. Fodor, A. Jakovac, S. D. Katz, S. Krieg, C. Ratti, and K. K. Szabo, J. High Energy Phys. 11 (2010) 077.
- [49] S. Borsanyi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti, and K. Szabo, J. High Energy Phys. 01 (2012) 138.
- [50] S. Borsanyi, S. Durr, Z. Fodor, C. Hoelbling, S. D. Katz, S. Krieg, D. Nogradi, K. K. Szabo, B. C. Toth, and N. Trombitas, J. High Energy Phys. 08 (2012) 126.
- [51] Sz. Borsányi, G. Endrődi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti, and K. K. Szabó, J. High Energy Phys. 08 (2012) 053.
- [52] S. Borsanyi, Nucl. Phys. A904-905, 270c (2013).
- [53] S. Borsanyi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti, and K. K. Szabo, Phys. Rev. Lett. **111**, 062005 (2013).
- [54] S. Sharma, Adv. High Energy Phys. 2013, 452978 (2013).

- [55] F. Karsch, B.-J. Schaefer, M. Wagner, and J. Wambach, Phys. Lett. B 698, 256 (2011).
- [56] A. Bazavov et al., Phys. Rev. Lett. 111, 082301 (2013).
- [57] A. Bazavov, H.-T. Ding, P. Hegde, F. Karsch, C. Miao, S. Mukherjee, P. Petreczky C. Schmidt, and A. Velytsky, Phys. Rev. D 88, 094021 (2013).
- [58] A. Bazavov et al., Phys. Rev. Lett. 109, 192302 (2012).
- [59] C. Bernard, T. Burch, C. DeTar, J. Osborn, S. Gottlieb, E. B. Gregory, D. Toussaint, U. M. Heller, and R. Sugar, Phys. Rev. D 71, 034504 (2005).
- [60] A. Bazavov et al., Phys. Rev. D 80, 014504 (2009).
- [61] A. Bazavov et al., Phys. Rev. D 86, 034509 (2012).
- [62] P. Petreczky, J. Phys. G **39**, 093002 (2012).
- [63] A. Bazavov *et al.* (HotQCD Collaboration), Phys. Rev. D 90, 094503 (2014).
- [64] H.-T. Ding, F. Karsch, and S. Mukherjee, Int. J. Mod. Phys. E 24, 1530007 (2015).
- [65] R. Bellwied, S. Borsanyi, Z. Fodor, S. D. Katz, A. Pasztor, C. Ratti, and K. K. Szabo, Phys. Rev. D 92, 114505 (2015).
- [66] H.-T. Ding, S. Mukherjee, H. Ohno, P. Petreczky, and H.-P. Schadler, Phys. Rev. D 92, 074043 (2015).
- [67] E. Braaten and R. D. Pisarski, Phys. Rev. D 45, R1827 (1992).
- [68] E. Braaten and A. Nieto, Phys. Rev. D 53, 3421 (1996).
- [69] D. J. Gross and F. Wilczek, Phys. Rev. Lett. 30, 1343 (1973).
- [70] H. D. Politzer, Phys. Rev. Lett. 30, 1346 (1973).
- [71] A. Bazavov, N. Brambilla, X. Garcia i Tormo, P. Petreczky, J. Soto, and A. Vairo, Phys. Rev. D 86, 114031 (2012).