## Quark Wigner distributions in a light-cone spectator model

Tianbo Liu<sup>1</sup> and Bo-Qiang Ma<sup>1,2,3,\*</sup>

<sup>1</sup>School of Physics and State Key Laboratory of Nuclear Physics and Technology, Peking University,

Beijing 100871, China

<sup>2</sup>Collaborative Innovation Center of Quantum Matter, Beijing, China

<sup>3</sup>Center for High Energy Physics, Peking University, Beijing 100871, China

(Received 31 March 2014; published 19 February 2015)

We investigate the quark Wigner distributions in a light-cone spectator model. The Wigner distribution, as a quasidistribution function, provides the most general one-parton information in a hadron. Combining the polarization configurations, unpolarized, longitudinal polarized, or transversal polarized, of the quark and the proton, we can define 16 independent Wigner distributions at leading twist. We calculate all these Wigner distributions for the u quark and the d quark, respectively. In our calculation, both the scalar and the axial-vector spectators are included, and the Melosh–Wigner rotation effects for both the quark and the axial-vector spectator are taken into account. The results provide us a very rich picture of the quark structure in the proton.

DOI: 10.1103/PhysRevD.91.034019

PACS numbers: 12.39.Ki, 14.20.Dh, 12.38.-t

## I. INTRODUCTION

One of the main goals of particle physics is to unravel the quark and gluon structure of hadrons, which are fundamentally described by OCD in the framework of the Yang-Mills gauge theory. Because of the nonperturbative nature of the QCD at the hadron scale, it is almost impossible to calculate all the properties of hadrons directly from the QCD at present. Though the Euclidean lattice method provides a first-principle numerical simulation [1], it is challenged by the enormous computational complexity and the unavoidable multihadron thresholds [2]. The Dyson-Schwinger equation method also leads to important insights on the nonperturbative properties of QCD [3]. Besides, the remarkable connection between the quantum gauge field theory in four dimensions and the classical gravity theory in five dimensions, developed in last 15 years and known as the holographic dual, sheds light on the confinement dynamics in OCD [4].

The parton model formulated by Feynman and formalized by Bjorken and Paschos [5,6] has proved successful in explaining the high-energy hadronic scattering experiments. Based on the partonic picture and factorization theorem [7], a general process independent function, referred to as the parton distribution function, is defined to describe the light-cone longitudinal momentum fractions carried by the partons in a nucleon, and it is a powerful tool on data analyses of high-energy hadronic scattering experiments. With the development of the polarization techniques, one is able to obtain spin related partonic information in the nucleon from the helicity and transversity. To describe the three-dimensional structure of the nucleon, the transverse momentum dependent parton distributions (TMDs) [8] and the generalized parton distributions (GPDs) [9] are introduced and proven to be useful tools. The TMD contains three-dimensional momentum distributions of the partons, while the GPD contains longitudinal momentum and two-dimensional transverse coordinates of the partons through the impact parameter dependent densities (IPDs).

As a further generalization, the Wigner distributions are defined as position-momentum joint distributions to understand the partonic structure of the nucleon. The Wigner distribution is a kind of phase-space distribution that contains the most general one-parton information in a nucleon. In fact, it is not a new concept and has been widely applied in many physical areas, such as the quantum information, quantum molecular dynamics, optics, nonlinear dynamics, and so on [10] and is even directly measurable in some experiments [11]. The Wigner distribution was first explored in QCD as a six-dimensional (three positions and three momentums) function [12,13], in which the relativistic effects were neglected. Then, a five-dimensional (two transverse positions and three momentums) Wigner distribution [14] was proposed in the light-cone formalism or in the infinite momentum frame, where the parton language is well defined. The simple QCD vacuum structure in the light-cone form allows an unambiguous definition of the constituents of a hadron [15]. The five-dimensional Wigner distributions will reduce to the TMDs and IPDs, when the transverse coordinates or the transverse momentums are integrated. Thus, possible relations between the TMDs and GPDs may be established via the Wigner distributions. Unlike the TMDs and the IPDs, the Wigner distributions do not have probability interpretations because of the Heisenberg uncertainty principle in quantum theories [16], but at twist 2, they can be expressed in terms of light-cone wave

mabq@pku.edu.cn

functions, and the semiclassical interpretations are still possible [14].

In this paper, we focus on quark Wigner distributions in the proton. With the combination of the polarization configurations (unpolarized or polarized along three spatial directions) of the quark and the proton, 16 Wigner distributions are defined at the leading twist. We investigate all these Wigner distributions in a light-cone spectator model. Both the scalar and the axial-vector spectators are included, and the axial-vector one is necessary for flavor separation. This model is proven successful in calculating the structure functions [17]. With the Melosh–Wigner rotation effect [18] taken into account, this kind of models has been applied to investigate many physical quantities, such as quark unpolarized, helicity, and transversity distributions [19,20]; form factors [21]; TMDs [22-25]; and GPDs [26,27], and the results are comparable with experiments. Therefore, the investigation of the Wigner distributions in such a simple model may still provide us some valuable information to help us understand the structure of the nucleon. The paper is organized as follows. We briefly

review the light-cone spectator model in Sec. II, and then calculate the Wigner distributions in Sec. III. The numerical results are shown in Sec. IV including discussions. A summary is drawn in the last section.

#### **II. LIGHT-CONE SPECTATOR MODEL**

The proton is an eigenstate of the light-cone Hamiltonian  $H_{\rm LC} = P^+P^- - P_{\perp}^2$ , and the eigenvalue is the invariant mass square. In the light-cone (or front) form relativistic dynamics [28], the fields are quantized at fixed light-cone time  $\tau = t + z$  instead of the ordinary time *t* in the instant form. The light-cone longitudinal momentum  $k^+ = k^0 + k^3$  of any massive particle is restricted to be positive definite, and thus the Fock space vacuum state  $|0\rangle$ , i.e., the free vacuum, is exactly the QCD vacuum, if the possibility of the color-singlet states built on massless gluon quanta with zero momentum is ignored [15]. Therefore, one may have an unambiguous definition of the constituents of the hadron. Then, the proton state can be expanded on the complete basis of free multiparticle Fock states as

$$\Psi: P^{+}, \mathbf{P}_{\perp}, S_{z} \rangle = \sum_{n, \{\lambda_{i}\}} \prod_{i=1}^{N} \int \frac{dx_{i} d^{2} \mathbf{k}_{\perp i}}{2\sqrt{x_{i}}(2\pi)^{3}} (16\pi^{3}) \delta\left(1 - \sum_{j=1}^{N} x_{j}\right) \delta^{(2)}\left(\sum_{j=1}^{N} \mathbf{k}_{\perp j}\right) \\ \times \psi_{n}(\{x_{j}\}, \{\mathbf{k}_{\perp j}\}, \{\lambda_{j}\}) | n: \{x_{i} P^{+}\}, \{x_{i} \mathbf{P}_{\perp} + \mathbf{k}_{\perp i}\}, \{\lambda_{i}\} \rangle,$$
(1)

where *N* is the number of the components in the Fock state  $|n\rangle$ ,  $x_i = k_i^+/P^+$  is the light-cone longitudinal momentum fraction carried by the *i*th component, and  $k_{\perp i}$  and  $\lambda_i$  are its intrinsic transverse momentum and light-cone helicity. The projection of the proton state on the Fock state  $|n\rangle$  is the so-called light-cone wave function  $\psi_n$ , which is frame independent.

The proton state is normalized as

$$\begin{split} \langle \Psi : P'^+, \boldsymbol{P}'_{\perp}, S'_z | \Psi : P^+, \boldsymbol{P}_{\perp}, S_z \rangle \\ &= 2P^+ (2\pi)^3 \delta(P^+ - P'^+) \delta^{(2)} (\boldsymbol{P}_{\perp} - \boldsymbol{P}'_{\perp}) \delta_{S_z S'_z}, \qquad (2) \end{split}$$

and the light-cone wave functions are correspondingly normalized as

$$\sum_{n,\{\lambda_i\}} \prod_{i=1}^{N} \int \frac{dx_i d^2 \mathbf{k}_{\perp i}}{2(2\pi)^3} (16\pi^3) \delta\left(1 - \sum_{j=1}^{N} x_j\right) \delta^{(2)} \left(\sum_{j=1}^{N} \mathbf{k}_{\perp j}\right) \\ \times |\psi_n(\{x_j\}, \{\mathbf{k}_{\perp j}\}, \{\lambda_j\})|^2 = 1.$$
(3)

The one-particle Fock state is defined as  $|p\rangle = \sqrt{2p^+}a^{\dagger}(p)|0\rangle$  with the commutation or anticommutation relations,

$$\begin{split} [a(p), a^{\dagger}(p')] &= \{b(p), b^{\dagger}(p')\} \\ &= (2\pi)^3 \delta(p^+ - p'^+) \delta^{(2)}(\pmb{p}_{\perp} - \pmb{p}'_{\perp}), \quad (4) \end{split}$$

where the a(p),  $a^{\dagger}(p)$ , b(p), and  $b^{\dagger}(p)$  are annihilation and creation operators for bosons and fermions, respectively.

In the spectator model, the proton is viewed as a struck quark and a spectator that contains the remaining part of the proton. Considering the fact that the leading term in the Fock states expansion for the proton is the valence threequark state  $|uud\rangle$ , we approximately regard the proton state as a valence quark and a spectator that carries the diquark quantum numbers, and some nonperturbative effects between the quarks and gluons in the spectator part from higher Fock states are effectively absorbed into the mass of the spectator. This kind of picture has been applied to study many physical observables and is supported by some lightfront holographic model [29]. However, this picture is actually simplistic, and some effects cannot be realistically absorbed into the spectator mass [30]. More realistic analyses will provide better descriptions of the nucleon [31,32]. Here, we only adopt this crude model to have some qualitative results.

Constrained by the quantum numbers of the proton and the quark, the spectator can only be either a scalar or an axial vector, and the latter one is necessary for flavor separation. Therefore, we express the proton state as

$$|\Psi\rangle = \sin\theta\phi_S |qS\rangle + \cos\theta\phi_V |qV\rangle, \tag{5}$$

where *S* and *V* stand for the scalar and the axial-vector spectators, respectively;  $\phi_S$  and  $\phi_V$  are the momentum space light-cone wave functions; and  $\theta$  is an angle to describe the *SU*(6) spin-flavor symmetry breaking. In this paper, we choose the isospin symmetric case  $\theta = \pi/4$  in our calculations. This choice is also adopted in the calculations of helicity and transversity distributions [19,20], form factors [21], and single-spin asymmetries [33], and the results therein are comparable with the data. One might alternatively regard it as a parameter to fit data.

The spin space wave function of the quark-spectator state is acquired from the SU(6) quark model in the instant form [19]:

$$|qS\rangle^{\uparrow/\downarrow} = u_T^{\uparrow/\downarrow}S(ud),$$
  

$$|qV\rangle^{\uparrow/\downarrow} = \pm \frac{1}{3} [u_T^{\uparrow/\downarrow}V_T^0(ud) - \sqrt{2}u_T^{\downarrow/\uparrow}V_T^{\pm 1}(ud) - \sqrt{2}d_T^{\uparrow/\downarrow}V_T^0(uu) + 2d_T^{\downarrow/\uparrow}V_T^{\pm 1}(uu)].$$
(6)

Then, we transform the instant form spinors to the light-cone form spinors through the Melosh–Wigner rotation [18],

$$\chi_T^{\uparrow} = w[(k^+ + m)\chi_F^{\uparrow} - (k^1 + ik^2)\chi_F^{\downarrow}],$$
  
$$\chi_T^{\downarrow} = w[(k^+ + m)\chi_F^{\downarrow} + (k^1 - ik^2)\chi_F^{\uparrow}],$$
 (7)

which plays an important role in understanding the "proton spin puzzle" [34]. The factor  $w = 1/\sqrt{2k^+(k^0 + m)}$  and the subscripts *T* and *F* stand for the instant form and the lightcone form spinors, respectively. This transformation procedure is consistent with the results directly derived from the light-cone field theory [35]. For the spectator, the scalar one does not transform since it has spin zero, and the axial-vector one transforms as [36]

$$V_T^{+1} = w^2 [(k^+ + m)^2 V_F^{+1} - \sqrt{2}(k^+ + m)(k^1 + ik^2) V_F^0 + (k^1 + ik^2)^2 V_F^{-1}],$$

$$V_T^0 = w^2 [\sqrt{2}(k^+ + m)(k^1 - ik^2) V_F^{+1} + 2(k^+(k^0 + m) - (k^1 - ik^2)(k^1 + ik^2)) V_F^0 - \sqrt{2}(k^+ + m)(k^1 + ik^2) V_F^{-1}],$$

$$V_T^{-1} = w^2 [(k^1 - ik^2)^2 V_F^{+1} + \sqrt{2}(k^+ + m)(k^1 - ik^2) V_F^0 + (k^+ + m)^2 V_F^{-1}].$$
(8)

For the momentum space light-cone wave function, we assume the Brodsky–Huang–Lepage (BHL) prescription [37],

$$\phi_D(x, \mathbf{k}_\perp) = A_D \exp\left\{-\frac{1}{8\beta_D^2} \left[\frac{m_q^2 + \mathbf{k}_\perp^2}{x} + \frac{m_D^2 + \mathbf{k}_\perp^2}{1 - x}\right]\right\},\tag{9}$$

where the subscript D represents the type of the spectator with S for the scalar one and V for the axial-vector one,  $m_q$ and  $m_D$  are the masses of the quark and the spectator,  $\beta_D$  is the harmonic oscillator scale parameter, and  $A_D$  is the normalization factor. We simply adopt m = 330 MeV,  $\beta_D = 330$  MeV as often adopted in the literature [38–40]. In principle, the spectator masses have a spectrum that might be extracted from experiments, and some modern understandings on diquark correlations in the nucleon are explored in Ref. [41]. Here, for simplicity, we regard them as parameters and choose the values as  $m_S = 600$  MeV and  $m_V = 800$  MeV. Apart from the BHL prescription, some other forms were proposed in the literature, such as the Terentev-Karmanov prescription [42], Chung-Coaster-Polyzou prescription [43], and Vega-Schmidt-Gutsche-Lyubovitskij prescription [44]. The choice of the momentum space wave function only affects the results quantitatively, and the qualitative properties are essentially determined by the spin structures.

## III. QUARK WIGNER DISTRIBUTIONS IN THE PROTON

The Wigner distribution is a quantum phase-space distribution first introduced by Wigner in quantum mechanics [45]. It is in general not positive definite and has no probability interpretations because of the Heisenberg uncertainty principle in quantum theories. Whereas, at certain situation, semiclassical interpretations are still possible, and thus it can be viewed as a quasidistribution function. Similar to the quark-quark correlation operator, the Hermitian Wigner operator for quarks at a fixed light-cone time is defined as [14]

$$\hat{W}^{[\Gamma]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} \int \frac{dz_{\perp} d^2 \boldsymbol{z}_{\perp}}{(2\pi)^3} e^{i\boldsymbol{k}\cdot\boldsymbol{z}} \bar{\psi}\left(\boldsymbol{y} - \frac{\boldsymbol{z}}{2}\right) \Gamma \mathcal{W} \psi\left(\boldsymbol{y} + \frac{\boldsymbol{z}}{2}\right) \Big|_{\boldsymbol{z}^+ = 0}, \quad (10)$$

where  $y = (0, 0, \mathbf{b}_{\perp})$  and  $\mathcal{W}$  is the gauge link Wilson line connecting the points y - z/2 and y + z/2 to ensure the SU(3) color gauge invariance of the Wigner operator. The  $\Gamma$  represents a Dirac  $\gamma$  matrix, and at twist-2, it is  $\gamma^+, \gamma^+\gamma_5$ , or  $i\sigma^{j+}\gamma_5$  with j = 1 or 2, corresponding to unpolarized, longitudinal-polarized, or *j*-direction transverse-polarized quarks, respectively. The  $\mathbf{b}_{\perp}$  and  $\mathbf{k}_{\perp}$  are the intrinsic transverse position and transverse momentum of the quark. They are not Fourier conjugate variables as demonstrated in Ref. [14], but the  $b_j$  and  $k_j$  along the same direction are still protected by the uncertainty principle, since they are not commutative with each other. Interpolating the Wigner operator (10) between the initial and final proton states with a transverse momentum  $\Delta_{\perp}$  transferred,

$$\rho^{[\Gamma]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, \boldsymbol{x}, \boldsymbol{S}) = \int \frac{d^2 \boldsymbol{\Delta}_{\perp}}{(2\pi)^2} \left\langle P^+, \frac{\boldsymbol{\Delta}_{\perp}}{2}, \boldsymbol{S} \right| \hat{W}^{[\Gamma]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, \boldsymbol{x}) \left| P^+, -\frac{\boldsymbol{\Delta}_{\perp}}{2}, \boldsymbol{S} \right\rangle$$
(11)

where *S* is the spin of the proton state, the five-dimensional Wigner distribution is defined. Here, the Drell–Yan–West frame  $(\Delta^+ = 0)$  [46], which is widely used in the calculation of form factors [47], is adopted, and with this choice, one may have semiclassical probability interpretations [14]. Including a longitudinal momentum transfer, one can define a six-dimensional Wigner distribution  $\rho^{[\Gamma]}(\boldsymbol{b}_{\perp}, \boldsymbol{\xi}, \boldsymbol{k}_{\perp}, x, \boldsymbol{S})$  but will lose the quasiprobability interpretation. Combining the polarization configurations, unpolarized (U), longitudinal polarized (L), and transverse polarized (T), of the proton and the quark, one can define 16 independent twist-2 quark Wigner distributions. They are the unpolarized Wigner distribution

$$\rho_{\rm UU}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} [\rho^{[\gamma^+]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_z) + \rho^{[\gamma^+]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_z)],$$
(12)

the unpolarized-longitudinal Wigner distribution

$$\rho_{\rm UL}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} [\rho^{[\gamma^+ \gamma_5]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_z) + \rho^{[\gamma^+ \gamma_5]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_z)], \quad (13)$$

the unpolarized-transverse Wigner distribution

$$\rho_{\mathrm{UT}}^{j}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x) = \frac{1}{2} [\rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x,\hat{\boldsymbol{e}}_{z}) + \rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x,-\hat{\boldsymbol{e}}_{z})], \quad (14)$$

the longitudinal-unpolarized Wigner distribution

$$\rho_{\mathrm{LU}}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x}) = \frac{1}{2} [\rho^{[\gamma^+]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x},\hat{\boldsymbol{e}}_z) - \rho^{[\gamma^+]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x},-\hat{\boldsymbol{e}}_z)],$$
(15)

the longitudinal Wigner distribution

$$\rho_{\rm LL}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} [\rho^{[\gamma^+ \gamma_5]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_z) - \rho^{[\gamma^+ \gamma_5]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_z)], \quad (16)$$

the longitudinal-transverse Wigner distribution

$$\rho_{\mathrm{LT}}^{j}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} [\rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_{z}) - \rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_{z})],$$

$$(17)$$

the transverse-unpolarized Wigner distribution

$$\rho_{\mathrm{TU}}^{i}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} [\rho^{[\gamma^{+}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_{i}) - \rho^{[\gamma^{+}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_{i})], \quad (18)$$

the transverse-longitudinal Wigner distribution

$$\rho_{\mathrm{TL}}^{i}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x) = \frac{1}{2} [\rho^{[\gamma^{+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x,\hat{\boldsymbol{e}}_{i}) - \rho^{[\gamma^{+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},x,-\hat{\boldsymbol{e}}_{i})], \quad (19)$$

the transverse Wigner distribution

$$\rho_{\mathrm{TT}}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} \delta_{ij} [\rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_{i}) - \rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_{i})],$$
(20)

and the pretzelous Wigner distribution

$$\rho_{\mathrm{TT}}^{\perp}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x) = \frac{1}{2} \epsilon_{ij} [\rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \hat{\boldsymbol{e}}_{i}) - \rho^{[i\sigma^{j+}\gamma_{5}]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, -\hat{\boldsymbol{e}}_{i})].$$

$$(21)$$

The first subscript stands for the proton polarization, and the second one stands for the quark polarization. The names are given by considering the polarization configuration of the quark, and then a prefix is added to describe the proton polarization, unless it is parallel polarized with the quark. The pretzelous Wigner distribution is named after the pretzelocity TMD to describe the case that both the quark and the proton are transversely polarized but along two orthogonal directions.

The Wigner distributions have direct connection with the generalized parton correlation functions (GPCFs) [48], which are the fully unintegrated off-diagonal quark-quark correlator for a nucleon. The generalized transverse momentum dependent parton distributions (GTMDs) can be defined by integrating the GPCFs over the light-cone energy. Then, the Wigner distributions can be regarded as the transverse Fourier transformation of the GTMDs. Since all the TMDs and GPDs can be obtained from the GTMDs, the GTMD is viewed as the so-called mother function. Therefore, the Wigner distribution is a connection between the TMDs and IPDs. Since the TMDs and IPDs are in general independent functions, the relations between them, if they exist, should be established at the level of Wigner distributions. Hence, the Wigner distribution is a useful tool to study the nucleon structure, although there are at present no clear methods to extract it from experiments.

The Wilson line W that connects the quark fields at two points y - z/2 and y + z/2 plays an important role in understanding the naive time-reversal-odd (T-odd) TMDs

[49]. The path is process dependent, and a sign change of quark Sivers and Boer–Mulders functions in the semiinclusive deeply inelastic scattering and Drell–Yan processes is predicted on the basis of this understanding. Here, we choose a staplelike path for the Wilson line as [48]

$$y - \frac{z}{2} = \left(0, -\frac{z}{2}, \boldsymbol{b}_{\perp} - \frac{\boldsymbol{z}_{\perp}}{2}\right) \rightarrow \left(0, \infty, \boldsymbol{b}_{\perp} - \frac{\boldsymbol{z}_{\perp}}{2}\right)$$
$$\rightarrow \left(0, \infty, \boldsymbol{b}_{\perp} + \frac{\boldsymbol{z}_{\perp}}{2}\right) \rightarrow \left(0, \frac{z}{2}, \boldsymbol{b}_{\perp} + \frac{\boldsymbol{z}_{\perp}}{2}\right)$$
$$= y + \frac{z}{2}$$

in order to obtain the appropriate Wilson line when taking the TMD and IPD limits. Although the two segments along

PHYSICAL REVIEW D 91, 034019 (2015)

the light-cone direction vanish in the light-cone gauge  $A^+ = 0$ , the transverse segment at the light-cone infinity is still nontrivial [50]. Here, we take a crude truncation of the Wilson line at the first order for simplicity, and it reduces to the unit operator 1. At the leading twist, there are 16 independent GTMDs, which are in general complex-valued functions, while the Wigner distributions are always real valued. If separating the real part and the imaginary part of the GTMDs, one can define 32 real-valued GTMDs, and 16 of them are T even, while the other 16 are T odd. Our truncation of the Wilson line essentially means to neglect the T-odd ones, i.e., the contribution from the imaginary part of the GTMDs.

Quantized at the fixed light-cone time, the quark field operator in Eq. (10) is expressed as

$$\psi\left(y+\frac{z}{2}\right) = \sum_{\lambda} \int \frac{d\ell^{+}}{\sqrt{2\ell^{+}}} \frac{d^{2}\ell_{\perp}}{(2\pi)^{3}} [b_{\lambda}(\ell)u(\ell,\lambda)e^{-i\ell\cdot(y+\frac{z}{2})} + d_{\lambda}^{\dagger}(\ell)v(\ell,\lambda)e^{i\ell\cdot(y+\frac{z}{2})}], \tag{22}$$

and similarly for  $\bar{\psi}(y-z/2)$ . The twist-2 Wigner operators are expressed as

$$\hat{W}^{[\gamma^{+}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x}) = \frac{1}{2} \int \frac{d\ell'^{+} d^{2}\ell'_{\perp}}{(2\pi)^{3}} \int \frac{d\ell'^{+} d^{2}\ell'_{\perp}}{(2\pi)^{3}} \\ \times \left\{ \left[ b^{\dagger}_{\uparrow}(\ell') b_{\uparrow}(\ell) + b^{\dagger}_{\downarrow}(\ell') b_{\downarrow}(\ell) \right] e^{-i(\ell'_{\perp} - \ell_{\perp}) \cdot \boldsymbol{b}_{\perp}} \delta\left(k^{+} - \frac{\ell'^{+} + \ell^{+}}{2}\right) \delta^{(2)}\left(\boldsymbol{k}_{\perp} - \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \right. \\ \left. - \left[ d^{\dagger}_{\uparrow}(\ell) d_{\uparrow}(\ell') + d^{\dagger}_{\downarrow}(\ell) d_{\downarrow}(\ell') \right] e^{i(\ell'_{\perp} - \ell_{\perp}) \cdot \boldsymbol{b}_{\perp}} \delta\left(k^{+} + \frac{\ell'^{+} + \ell^{+}}{2}\right) \delta^{(2)}\left(\boldsymbol{k}_{\perp} + \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \right\},$$
(23)

$$\hat{W}^{[\gamma^{+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x}) = \frac{1}{2} \int \frac{d\ell'^{+}d^{2}\ell'_{\perp}}{(2\pi)^{3}} \int \frac{d\ell'^{+}d^{2}\ell_{\perp}}{(2\pi)^{3}} \\ \times \left\{ [b_{\uparrow}^{\dagger}(\ell')b_{\uparrow}(\ell) - b_{\downarrow}^{\dagger}(\ell')b_{\downarrow}(\ell)]e^{-i(\ell'_{\perp}-\ell_{\perp})\cdot\boldsymbol{b}_{\perp}}\delta\left(k^{+} - \frac{\ell'^{+} + \ell^{+}}{2}\right)\delta^{(2)}\left(\boldsymbol{k}_{\perp} - \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \\ + [d_{\uparrow}^{\dagger}(\ell)d_{\uparrow}(\ell') - d_{\downarrow}^{\dagger}(\ell)d_{\downarrow}(\ell')]e^{i(\ell'_{\perp}-\ell_{\perp})\cdot\boldsymbol{b}_{\perp}}\delta\left(k^{+} + \frac{\ell'^{+} + \ell^{+}}{2}\right)\delta^{(2)}\left(\boldsymbol{k}_{\perp} + \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right)\right\},$$
(24)

$$\hat{W}^{[i\sigma^{1+}\gamma_{5}]}(\boldsymbol{b}_{\perp},\boldsymbol{k}_{\perp},\boldsymbol{x}) = \frac{1}{2} \int \frac{d\ell'^{+}d^{2}\ell'_{\perp}}{(2\pi)^{3}} \int \frac{d\ell'^{+}d^{2}\ell'_{\perp}}{(2\pi)^{3}} \\ \times \left\{ [b^{\dagger}_{\uparrow}(\ell')b_{\downarrow}(\ell) + b^{\dagger}_{\downarrow}(\ell')b_{\uparrow}(\ell)]e^{-i(\ell'_{\perp}-\ell_{\perp})\cdot\boldsymbol{b}_{\perp}}\delta\left(k^{+} - \frac{\ell'^{+} + \ell^{+}}{2}\right)\delta^{(2)}\left(\boldsymbol{k}_{\perp} - \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \\ - [d^{\dagger}_{\uparrow}(\ell)d_{\downarrow}(\ell') + d^{\dagger}_{\downarrow}(\ell)d_{\uparrow}(\ell')]e^{i(\ell'_{\perp}-\ell_{\perp})\cdot\boldsymbol{b}_{\perp}}\delta\left(k^{+} + \frac{\ell'^{+} + \ell^{+}}{2}\right)\delta^{(2)}\left(\boldsymbol{k}_{\perp} + \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right)\right\}, \quad (25)$$

where we use the Lepage–Brodsky conventions for the properties of the light-cone spinors [51]. In the spectator model, the proton state in Eq. (11) is expressed as

$$\left| P^{+}, -\frac{\mathbf{\Delta}_{\perp}}{2}, \mathbf{S} \right\rangle = \sum_{\sigma, s} \int \frac{dx_{q}}{2\sqrt{x_{q}}} \frac{d^{2}\mathbf{k}_{q\perp}}{(2\pi)^{3}} \int \frac{dx_{D}}{2\sqrt{x_{D}}} \frac{d^{2}\mathbf{k}_{D\perp}}{(2\pi)^{3}} 16\pi^{3}\delta(1 - x_{q} - x_{D})\delta^{(2)}(\mathbf{k}_{q\perp} + \mathbf{k}_{D\perp}) \times \psi^{S}_{\sigma s}(x_{q}, \mathbf{k}_{q\perp}) \sqrt{2k_{q}^{+}} \sqrt{2k_{D}^{+}} b^{\dagger}_{\sigma} \left(x_{q}, \mathbf{k}_{q\perp} - x_{q}\frac{\mathbf{\Delta}_{\perp}}{2}\right) a^{\dagger}_{s} \left(x_{D}, \mathbf{k}_{D\perp} - x_{D}\frac{\mathbf{\Delta}_{\perp}}{2}\right) |0\rangle,$$
(26)

where  $k_{q\perp}$  and  $k_{D\perp}$  are the intrinsic transverse momentum for the quark and the spectator and  $x_q = k_q^+/P^+$  and  $x_D =$  $k_D^+/P^+$  are the light-cone momentum fractions carried by the quark and the spectator. The arguments for the creation (annihilation) operators are the frame dependent transverse momentum in the one-particle Fock state definition, whereas the light-cone wave function  $\psi_{\sigma s}^{S}(x, \mathbf{k}_{q\perp})$ , which describes the amplitude of the quark-spectator state in the proton, is frame independent and only depends on boost invariant variables x and  $k_{q\perp}$ . The light-cone wave function can be factored into a spin space part and a momentum space part. The spin space part is given by the Melosh-Wigner rotations (7) and (8), and the momentum space part is given by the BHL prescription (9). Then, sandwiching the Wigner operators in Eqs. (23)–(25) between the proton states with different polarization configures, we can obtain the expressions of all the 16 quark Wigner distributions in terms of the overlap of light-cone wave functions, as shown in the Appendix.

#### **IV. NUMERICAL RESULTS AND DISCUSSIONS**

The Wigner distribution is a five-dimensional function of  $b_x$ ,  $b_y$ ,  $k_x$ ,  $k_y$ , and x. As a phase-space distribution, what we concern ourselves with most is the behavior in the

transverse coordinate and transverse momentum space. Therefore, we integrate over the light-cone momentum fraction x and display its behavior in the remaining four dimensions, i.e., the so-called transverse Wigner distributions [14]. We plot the Wigner distribution in the transverse coordinate space with a definite transverse momentum and in the transverse momentum space with a definite coordinate. Apart from the TMDs and IPDs, which can all be obtained from the Wigner distributions, one may define the mixing distributions by integrating over a transverse coordinate and a transverse momentum along two perpendicular transverse directions as

$$\tilde{\rho}_{\Lambda\lambda}(b_x, k_y, x) = \int db_y dk_x \rho_{\Lambda\lambda}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x), \quad (27)$$

$$\bar{\rho}_{\Lambda\lambda}(b_y, k_x, x) = \int db_x dk_y \rho_{\Lambda\lambda}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x). \quad (28)$$

Unlike the Wigner distributions, the mixing distributions are real distribution functions since the remaining variables are not protected by the uncertainty principle, and they describe the correlation of quark transverse coordinate and transverse momentum in orthogonal directions. Both Wigner distributions and mixing distributions are



FIG. 1 (color online). Unpolarized Wigner distributions  $\rho_{UU}$  and mixing distributions  $\tilde{\rho}_{UU}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the Wigner distributions in transverse coordinate space with definite transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_{y}$ . The second column is the Wigner distributions in transverse momentum space with definite transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_{y}$ . The third column is the mixing distributions  $\tilde{\rho}_{UU}$ .

QUARK WIGNER DISTRIBUTIONS IN A LIGHT-CONE ... PHYSICAL REVIEW D 91, 034019 (2015)  $\rho_{\rm UL}^{u}$  $\tilde{\rho}_{\text{UL}}^{\mu}$  $\rho_{\rm UL}^{u}$ 0.010 0.010 0.02 0.005 0.005 0.01 0.000 0.00 0.000 -0.005-0.01-0.005 0.25 0.25 -0.010 -0.02 -0.010-0.5  $b_{\rm s}/{\rm fm}$ k./GeV k/GeV -0.5-0.25 -0.5 0.5 0.25 0.25  $b_x/\text{fm}$ k./GeV  $b_x/fm$ 0.5 0.25 0.5 -0.5 -0.5 0.5  $\tilde{\rho}_{\text{UL}}^{d}$  $\rho_{\mathrm{UL}}^d$  $\rho_{\mathrm{UL}}^d$ 0.02 0.005 0.005 0.01 0.000 0.00 0.00 -0.005-0.01 0.25 -0.00 0.25 -0.02-0.5 k./GeV k./GeV  $b_v/fm$ -0.25-0.5 -0.5 0.25 -0.25 -0.5  $k_x/\text{GeV}$  $b_{\rm u}/{\rm fm}$  $b_{\rm r}/{\rm fm}$ 0.25 0.5 -0.5 -0.5 0.5

FIG. 2 (color online). Unpolarized-longitudinal Wigner distributions  $\rho_{UL}$  and mixing distributions  $\tilde{\rho}_{UL}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the Wigner distributions in transverse coordinate space with definite transverse momentum  $k_{\perp} = 0.3 \text{ GeV } \hat{e}_y$ . The second column is the Wigner distributions in transverse momentum space with definite transverse coordinate  $b_{\perp} = 0.4 \text{ fm} \hat{e}_y$ . The third column is the mixing distributions  $\tilde{\rho}_{UL}$ .

dimensionless quantities in natural units. In this section, we display the numerical results of quark Wigner distributions and mixing distributions.

## A. Unpolarized proton

In Fig. 1, we plot the unpolarized Wigner distributions  $\rho_{UU}$  and mixing distributions  $\tilde{\rho}_{UU}$ . The unpolarized Wigner distributions represent the transverse phase-space distribution of the unpolarized quark in an unpolarized proton. Its behaviors in transverse coordinate space and transverse momentum space are plotted with fixed transverse momentum and transverse coordinate separately. We denote the direction of the fixed momentum or the fixed coordinate as the *y* direction, since there is no privileged transverse plane. The same case also occurs for  $\rho_{UL}$ ,  $\rho_{LU}$ , and  $\rho_{LL}$  for which neither the quark nor the proton is transverse polarized.

As general properties, the left  $(-\hat{e}_x)$ -right  $(\hat{e}_x)$  symmetry of the unpolarized Wigner distributions is a direct result of the space symmetry and reflects the fact that the proton viewed from any direction is the same. This property is also observed in the mixing distributions in which one may clearly find the quark has no preference on moving either clockwise or counterclockwise. It is consistent with our topological knowledge that any preference must break the SO(3) symmetry. The up  $(\hat{e}_y)$ -down  $(-\hat{e}_y)$  symmetry of the Wigner distributions reflects the nature that the quarks in the proton is in equilibrium. The two classes of mixing distributions defined in Eqs. (27) and (28) have a simple relation,

$$\tilde{\rho}_{\mathrm{UU}}(b_x, k_y, x) = \bar{\rho}_{\mathrm{UU}}(b_y, -k_x, x), \qquad (29)$$

and this relation also survives for  $\tilde{\rho}_{\rm UL}$ ,  $\tilde{\rho}_{\rm LU}$ , and  $\tilde{\rho}_{\rm LL}$ .

Comparing the behaviors of the u quark and the d quark, we find in our model that the u quark favors concentrating in the center relative to the d quark in the coordinate space, while they have similar spread behaviors in the momentum space. Though it is a common result to find that the uconcentrates in the center in quark models, this property does have dependence on the parameters in the model. The similar behavior in the momentum space is due to the same choice of the  $\beta$  parameter in the scalar and axial-vector cases. Besides, as found in some constituent quark model and chiral quark soliton model [14], the property that the distributions decrease faster in the y direction than the xdirection is not obvious in our model.

In Fig. 2, we plot the unpolarized-longitudinal Wigner distributions  $\rho_{\rm UL}$  and mixing distributions  $\tilde{\rho}_{\rm UL}$ . They



FIG. 3 (color online). Unpolarized-transverse Wigner distributions  $\rho_{\text{UT}}$  and mixing distributions  $\bar{\rho}_{\text{UT}}$  for *u* quark (upper) and *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the second column is those with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_y$  perpendicular to the quark polarization. The third column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_y$  perpendicular to the quark polarization. The fourth column is the mixing distributions.

describe the longitudinal-polarized quark in an unpolarized proton. This distribution will vanish in either the TMD or IPD limit, and therefore no general relations at twist 2 are found between the TMDs or IPDs and the unpolarizedlongitudinal Wigner distributions. Hence, the information contained in this distribution cannot be observed from the leading-twist TMDs or IPDs, though some effects might be found at subleading twist.

Because of the privileged direction defined by the quark polarization, dipole structures are found in the unpolarizedlongitudinal Wigner distributions, and correspondingly quadrupole structures appear in the mixing distributions. Considering the physical interpretations, these distributions essentially reflect quark spin-orbit correlations. This correlation can be clearly observed from the quadrupole structure of the unpolarized-longitudinal mixing distributions, and no correlation case corresponds to vanishing mixing distributions.

We find in our model that both the u quark and the d quark have negative spin-orbit correlations. This result has little dependence on the choice of momentum space wave functions or the parameters in the model. It is mainly determined by the spin structures of the quark and the spectator. However, it is indeed model dependent, and our result is opposite of that from some constituent quark model [14]. Therefore, realistic analyses in QCD on this distribution may help us clarify this issue.

In Fig. 3, we plot the unpolarized-transverse Wigner distributions  $\rho_{\rm UT}$  and mixing distributions  $\bar{\rho}_{\rm UT}$ . They describe the transverse-polarized quark in an unpolarized proton. The SO(2) symmetry in the transverse space is

explicitly broken by the quark polarization direction, referred to as the *x* direction.

At the TMD limit, the unpolarized-transverse Wigner distribution will reduce to the Boer–Mulders function  $h_1^{\perp}$ , while at the IPD limit, it is related to the  $\tilde{H}_T$  together with some other distributions, whereas the  $h_1^{\perp}$  corresponds to the T-odd part, while the  $\tilde{H}_T$  corresponds to the T-even part. Since the T-odd part is neglected in this paper, the results of  $\rho_{\rm UT}$  here will vanish at the TMD limit.

As a general property determined by the spin structures, the unpolarized-transverse Wigner distributions vanish when the quark intrinsic transverse coordinate is parallel to the polarization. In other words, quark transverse spin has no correlation with its parallel transverse coordinate. Besides, the results displayed here have no dependence on the direction of quark transverse momentum, but this behavior may be changed by a nontrivial Wilson line.

#### **B.** Longitudinal-polarized proton

In Fig. 4, we plot the longitudinal-unpolarized Wigner distributions  $\rho_{LU}$  and mixing distributions  $\tilde{\rho}_{LU}$ . They describe quark phase-space distributions in a longitudinal-polarized proton without any information of quark spins. Thus, they are suitable to study quark orbital angular momentum related issues [52]. Similar to the unpolarized-longitudinal Wigner distribution, no twist-2 TMDs or IPDs are related to the longitudinal-unpolarized Wigner distribution. Therefore, it will vanish at the TMD or IPD limit, and the phase-space behavior in this distribution cannot generally be extracted from leading-twist TMDs or IPDs.

QUARK WIGNER DISTRIBUTIONS IN A LIGHT-CONE ... PHYSICAL REVIEW D 91, 034019 (2015)  $\rho_{LU}^{u}$  $\tilde{\rho}^{\mu}_{LU}$  $\rho_{LU}^{u}$ 0.005 0.010 0.01 0.005 0.000 0.00 0.000 -0.005.25 0.25 -0.00 -0.0 -0.010 -0.5 k./GeV k./GeV -0.5-0.25-0.50.25 0.25  $b_{\rm r}/{\rm fm}$ k<sub>x</sub>/GeV  $b_x/fm$ 0.5 0.25 0.5 -0.5 -0.5 0.5 1  $\rho_{LU}^d$  $\tilde{\rho}_{LU}^d$  $\rho_{\rm LU}^d$ 0.0010 0.0005 0.002 0.0005 0.0000 0.0000 0.000  $-0.000^{4}$ 0.25 0.25 -0.002 -0.0005 -0.0010k./GeV -0.5  $b_{\rm v}/{\rm fm}$ k<sub>v</sub>/GeV -0.5-0.5 -0.250.25 0.25 -0.5  $b_{\rm v}/{\rm fm}$ k./GeV 0.5  $b_{\rm x}/{\rm fm}$ 0.25 -0.5 -0.5 0.5

FIG. 4 (color online). Longitudinal-unpolarized Wigner distributions  $\rho_{LU}$  and mixing distributions  $\tilde{\rho}_{LU}$  for the *u* quark (upper) and *d* quark (lower). The first column is the Wigner distributions in transverse coordinate space with definite transverse momentum  $k_{\perp} = 0.3 \text{ GeV } \hat{e}_y$ . The second column is the Wigner distributions in transverse momentum space with definite transverse coordinate  $b_{\perp} = 0.4 \text{ fm} \hat{e}_y$ . The third column is the mixing distributions  $\tilde{\rho}_{LU}$ .

The longitudinal-unpolarized Wigner distributions have dipole structures, and correspondingly the mixing distributions have quadrupole structures. The preference of quark orbital motions, clockwise or counterclockwise, are clearly observed from them. We find in our calculations that the *u* quark has positive orbital angular momentum, while the d quark has negative orbital angular momentum, but this is model dependent, and the choice of the wave functions may also change this behavior. In addition, the sign change of the d quark distribution in a large coordinate or momentum region is found, as also observed in some constituent quark model. Comparing the longitudinalunpolarized distributions with the unpolarized-longitudinal distributions, we find in this model that quark orbital motions have stronger correlation to quark spins than to nucleon spins. However, this property is not general. Therefore, more careful investigation on this distribution is required.

In Fig. 5, we plot the longitudinal Wigner distributions  $\rho_{LL}$  and mixing distributions  $\tilde{\rho}_{LL}$ . They describe the longitudinal-polarized quark in a longitudinal-polarized proton and correspond to the helicity distributions after integrating over transverse variables.

In the simple model, we find the u quark is positive polarized, while the d quark is negative polarized. This is qualitatively consistent with our knowledge from the axial charge. In the large transverse momentum region, a sign change is clearly observed in d quark mixing distribution, as also found in some constituent quark model. This kind of sign change is also found in the large longitudinal momentum region of d quark helicity distributions in similar models [24]. Besides, it is natural to find that the quark helicity distributions concentrate on the center in the phase space, though it is a result of a simplistic model.

In Fig. 6, we plot the longitudinal-transverse Wigner distributions  $\rho_{LT}$  and mixing distributions  $\bar{\rho}_{LT}$ . They describe the correlation of the phase-space distribution with quark transverse spin and proton longitudinal spin. At the TMD limit, the longitudinal-transverse Wigner distribution will reduce to one worm-gear function, the longitudinal transversity  $h_{1L}^{\perp}$ . At the IPD limit, it is related to the IPDs  $H_T$  and  $\tilde{H}_T$  together with other distributions. Unlike the unpolarized-transverse distributions, both the TMD and the IPDs are related to the T-even part of the longitudinal-transverse Wigner distribution at certain limits. Thus, it is possible to build some relation between them, though no general relations have been found yet.

In the longitudinal-transverse Wigner distribution, the transverse SO(2) symmetry is explicitly broken by the transverse polarization direction, which we refer to as the x direction. With a trivial Wilson line as adopted in this paper, the Wigner distribution as well as the corresponding mixing



FIG. 5 (color online). Longitudinal Wigner distributions  $\rho_{LL}$  and mixing distributions  $\tilde{\rho}_{LL}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the Wigner distributions in transverse coordinate space with definite transverse momentum  $k_{\perp} = 0.3 \text{ GeV } \hat{e}_y$ . The second column is the Wigner distributions in transverse momentum space with definite transverse coordinate  $b_{\perp} = 0.4 \text{ fm} \hat{e}_y$ . The third column is the mixing distributions  $\tilde{\rho}_{LL}$ .



FIG. 6 (color online). Longitudinal-transversal Wigner distributions  $\rho_{LT}$  with fixed  $\mathbf{k}_{\perp}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3$  GeV  $\hat{\mathbf{e}}_x$  parallel to the quark polarization. The second column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4$  fm  $\hat{\mathbf{e}}_x$  parallel to the quark polarization, and the third column is those with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4$  fm  $\hat{\mathbf{e}}_y$  perpendicular to the quark polarization. The fourth column is the mixing distributions  $\bar{\rho}_{LT}$ .

PHYSICAL REVIEW D 91, 034019 (2015)



FIG. 7 (color online). Transverse-unpolarized Wigner distributions  $\rho_{TU}$  and mixing distributions  $\bar{\rho}_{TU}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_x$  parallel to the proton polarization, and the second column is those with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_y$  perpendicular to the proton polarization. The third column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ m} \hat{\mathbf{e}}_y$  perpendicular to the proton polarization. The fourth column is the mixing distributions.

distribution is vanishing when the quark transverse momentum is perpendicular to the transverse polarization. Careful examinations including a nontrivial Wilson line may lead to nonzero values in this situation. In general, this behavior reflects a strong correlation to the direction of the quark intrinsic transverse momentum, and in contrast the correlation to the direction of transverse coordinate is weak in this model.

## C. Transverse-polarized proton

In Fig. 7, we plot the transverse-unpolarized Wigner distributions  $\rho_{TU}$  and mixing distributions  $\bar{\rho}_{TU}$ . They describe the unpolarized quark in a transverse-polarized proton and thus represent the correlation of quark transverse distributions and proton transverse spin. A privileged direction in the transverse plane is determined by the proton polarization, which is referred to as the *x* direction.

At the TMD limit, the transverse-unpolarized Wigner distribution corresponds to a naive T-odd distribution, the Sivers function  $f_{1T}^{\perp}$ . At the IPD limit, it is related to the IPDs *H* and *E* together with other distributions. However, these two limits select two different parts of the Wigner distribution. The Sivers TMD corresponds to the T-odd part, while the IPDs *H* and *E* only depend on the T-even part. Thus, there is no general relation between them if no further assumptions are applied to connect the T-odd and T-even parts. However, some relation between the Sivers function  $f_{1T}^{\perp}$  and the GPD *E* is suggested by some model calculations [26,53] and also implicated in some experiments [54]. Thus, this kind of connection is not impossible.

In our calculation, since the effect from the Wilson line is neglected, the vanishing result is obtained at the TMD limit. Therefore, the nontrivial gauge link plays an important role in establishing the possible relation between the Sivers TMD and the GPDs.

Similar to the situation in the unpolarized-transverse Wigner distribution, the transverse-unpolarized Wigner distribution is vanishing if quark transverse coordinate is parallel to the proton polarization. Together with the behavior of the unpolarized-transverse Wigner distribution, the quark transverse coordinate has no correlations to either the quark parallel transverse spin or proton parallel transverse spin. Besides, its behavior has no dependence on the direction of quark transverse momentum, but a nontrivial Wilson line may change this property.

In Figs. 8, we plot the transverse-longitudinal Wigner distributions  $\rho_{TL}$  and mixing distributions  $\bar{\rho}_{TL}$ . They describe the longitudinal-polarized quark in a transverse-polarized proton. At the TMD limit, the transverse-longitudinal Wigner distribution will reduce to the other worm-gear function, the transverse-helicity  $g_{1L}$ . At the IPD limit, it is related to the IPDs  $\tilde{H}$  and  $\tilde{E}$  together with other distributions. Both of the transverse-helicity TMD  $g_{1L}$  and the IPDs  $\tilde{H}$  and  $\tilde{E}$  only depend on the T-even part of the transverse-longitudinal distribution. Thus, it is possible to establish some relations between them. Besides, the T-odd part, which is neglected in our model calculation, will provide us information beyond TMDs and IPDs.

Similar to the situation of the longitudinal-transverse Wigner distribution, the isotropy in the transverse plane is



FIG. 8 (color online). Transverse-longitudinal Wigner distributions  $\rho_{TL}$  and mixing distributions  $\bar{\rho}_{TL}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_x$  parallel to the proton polarization. The second column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  parallel to the proton polarization, and the third column is those with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  perpendicular to the proton polarization. The fourth column is the mixing distributions.

explicitly violated by the transverse polarization, to which we refer as the x direction, in transverse-longitudinal Wigner distributions. Determined by the combination of the polarization configurations, the phase-space distribution is zero when quark transverse momentum is perpendicular to the proton spin, but a nontrivial Wilson line may lead to nonzero values. This behavior reflects a strong correlation to the direction of quark transverse momentum. Compared to the longitudinal-transverse Wigner distribution, the correlation to the direction of quark transverse coordinate is not weak in this case.

In Figs. 9 and 10, we plot the transverse Wigner distributions  $\rho_{TT}$  and mixing distributions  $\tilde{\rho}_{TT}$  and  $\bar{\rho}_{TT}$ . They describe the phase-space distribution in the case that

![](_page_11_Figure_5.jpeg)

FIG. 9 (color online). Transverse Wigner distributions  $\rho_{\text{TT}}$  for the *u* quark (upper) and the *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the second column is those with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV} \hat{\mathbf{e}}_y$  perpendicular to the quark polarization. The third column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the fourth column is those with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  perpendicular to the quark polarization.

![](_page_12_Figure_1.jpeg)

FIG. 10 (color online). Transverse mixing distributions  $\tilde{\rho}_{TT}$  (left) and  $\bar{\rho}_{TT}$  (right) for the *u* quark (upper) and *d* quark (lower) with the quark polarized along the *x* direction.

both the quark and the proton are transverse polarized. Because of the two degrees of freedom of the choice of the transverse polarization direction, there are two independent combinations of the transverse polarizations of the quark and the proton. For the transverse Wigner distributions, the quark and proton are parallel polarized and referred to as the x direction. The situation that the quark and proton are transverse polarized along two orthogonal directions is

described by the pretzelous Wigner distributions, as plotted in Fig. 11.

The transverse Wigner distributions will reduce to the transversity distributions when the transverse phase space is integrated. Unlike the longitudinal Wigner distributions in which the quark helicity concentrates on the center of the phase space, we find in the model that the maximum of the d quark transversity distribution in the phase space is not on the center point. Although this behavior is not forbidden by any physical principles, it is beyond our intuition that the increasing distance away from the center [55]. Though this is a model dependent result, it provides us the possibility that may be taken into account in realistic analyses.

For the pretzelous Wigner distribution in Fig. 11, a quadrupole structure is observed. This behavior is essentially determined by the spin structures. In this model, we find opposite correlation in the phase space for the u quark and d quark in this situation.

## **V. SUMMARY**

In this paper, we investigate quark Wigner distributions in a light-cone spectator model. The Wigner distribution, as a phase-space distributions, contains full one-parton information in the proton. All TMDs and IPDs can be obtained from the Wigner distributions at certain limits. Therefore, it is a possible bridge to build the relation between the TMDs and GPDs. Besides, the Wigner distributions do contain information that cannot in general be extracted from the

![](_page_12_Figure_10.jpeg)

FIG. 11 (color online). Pretzelous Wigner distributions  $\rho_{TT}^{\perp}$  for the *u* quark (upper) and *d* quark (lower). The first column is the distributions in transverse coordinate space with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV } \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the second column is those with fixed transverse momentum  $\mathbf{k}_{\perp} = 0.3 \text{ GeV } \hat{\mathbf{e}}_y$  parallel to the proton polarization. The third column is the distributions in transverse momentum space with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the fourth column is those with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_x$  parallel to the quark polarization, and the fourth column is those with fixed transverse coordinate  $\mathbf{b}_{\perp} = 0.4 \text{ fm} \hat{\mathbf{e}}_y$  parallel to the proton polarization.

TMDs and GPDs. Thus, the study on Wigner distributions will provide us new knowledge on understanding the nucleon structures.

We perform the calculations of all twist-2 quark Wigner distributions in a spectator model, which has been applied to study many physical observables in high-energy scattering experiments. In our calculation, we include both the scalar and the axial-vector spectators to have flavor separation. For the Wilson line, we take a crude truncation in this paper, and this means the neglect of the naive T-odd part in Wigner distributions. Though the calculations are performed in a simplistic model, we can still find some general properties of quark Wigner distributions from the results. Thus, the plots displayed here can be view as qualitative results. More realistic analyses in QCD are required in the future.

The dipole and quadrupole structures of some Wigner distributions and mixing distributions are determined by the spin structures of the quark and proton and should be general properties. Although the Wilson line is neglected in this study, some possible effects from the nontrivial Wilson line are discussed. Some Wigner distributions that have no corresponding TMDs or IPDs are also interesting, such as the unpolarized-longitudinal distribution  $\rho_{\rm UL}$  and the longitudinal-unpolarized distribution  $\rho_{\rm LU}$ . They are related to the issues on spin-orbit correlation and orbit angular momentum, and variant results can be obtained from different models. Therefore, more careful investigation on these

Wigner distributions is important to clarify the physical pictures. In addition, we have some discussion on the possibility to establish the relations between TMDs and GPDs. Though no general relations have been found, it is suggested by some model calculations and implicated by some experiments to relate the Sivers function with the GPD *E*. Since they are related to the same Wigner distribution but different parts, it is an opportunity to find the relation at Wigner distribution level. Therefore, the investigation on Wigner distributions can improve our understanding on the nucleon structure, and more realistic analyses are deserved.

## ACKNOWLEDGMENTS

This work is supported by National Natural Science Foundation of China (Grants No. 11035003 and No. 11120101004).

# APPENDIX: LIGHT-CONE WAVE FUNCTION OVERLAP REPRESENTATION FOR THE WIGNER DISTRIBUTION

In this section, we take the  $\rho^{[\gamma^+]}(\boldsymbol{b}_{\perp}, \boldsymbol{k}_{\perp}, x, \uparrow)$  as an example to demonstrate how the twist-2 Wigner distributions are expressed in terms of the overlap of light-cone wave functions in the spectator model.

Interpolating the Wigner operator in Eq. (23) between two proton states with a transverse momentum  $\Delta_{\perp}$  transferred, we can obtain the Wigner distribution with Eq. (11) as

$$\begin{split} \rho^{[r^+]}(b_{\perp},k_{\perp},x,\uparrow) \\ &= \int \frac{d^2 \mathbf{\Delta}_{\perp}}{(2\pi)^2} \sum_{\sigma',s'} \int \frac{dx'_q d^2 k'_{q\perp}}{2\sqrt{x'_q}(2\pi)^3} \int \frac{dx'_D d^2 k'_{D\perp}}{2\sqrt{x'_D}(2\pi)^3} 16\pi^3 \delta(1-x'_q-x'_D) \delta^{(2)}(k'_{q\perp}+k'_{D\perp}) \\ &\times \sum_{\sigma,s} \int \frac{dx_q d^2 k_{q\perp}}{2\sqrt{x'_q}(2\pi)^3} \int \frac{dx_D d^2 k_{D\perp}}{2\sqrt{x_D}(2\pi)^3} 16\pi^3 \delta(1-x_q-x_D) \delta^{(2)}(k_{q\perp}+k_{D\perp}) \\ &\times \frac{1}{2} \int \frac{d\ell'' d^2 \ell'_{\perp}}{(2\pi)^3} \int \frac{d\ell' + d^2 \ell_{\perp}}{(2\pi)^3} e^{-i(\ell'_{\perp}-\ell_{\perp})\cdot b_{\perp}} \delta\left(k^+ - \frac{\ell'^+ + \ell^+}{2}\right) \delta^{(2)}\left(k_{\perp} - \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \\ &\times (2P^+)^2 \sqrt{x'_q x'_D} \sqrt{x_q x_D} \psi^{\dagger}_{\sigma's'}(x'_q, k'_{q\perp}) \psi^{\dagger}_{\sigma s}(x_q, k_{q\perp}) \langle 0|a_{s'}\left(x'_D P^+, k'_{D\perp} + x'_D \frac{\Delta_{\perp}}{2}\right) b_{\sigma'}\left(x'_q P^+, k'_{q\perp} + x'_q \frac{\Delta_{\perp}}{2}\right) \\ &\times \sum_{\lambda} b^{\dagger}_{\lambda} (\ell'^+, \ell'_{\perp}) b_{\lambda} (\ell^+, \ell_{\perp}) b^{\dagger}_{\sigma}\left(x_q P^+, k_{q\perp} - x_q \frac{\Delta_{\perp}}{2}\right) a^{\dagger}_{s}\left(x_D P^+, k_{D\perp} - x_D \frac{\Delta_{\perp}}{2}\right) |0\rangle \\ &= \sum_{\sigma',s'} \sum_{\sigma,s} \sum_{\lambda} \frac{1}{2} \int \frac{d^2 \Delta_{\perp}}{(2\pi)^3} \int \frac{dx'_q d^2 k'_{q\perp}}{(2\pi)^3} \int \frac{dx_q d^2 k_{q\perp}}{(2\pi)^3} \int \frac{d\ell'' d^2 \ell'_{\perp}}{(2\pi)^3} \int \frac{d\ell'' d^2 \ell'_{\perp}}{(2\pi)^3} \\ &\times (P^+)^2 \psi^{\uparrow *}_{\sigma's'}(x'_q, k'_{q\perp}) \psi^{\dagger}_{\sigma s}(x_q, k_{q\perp}) e^{-i(\ell'_{\perp} - \ell_{\perp})\cdot b_{\perp}} \delta\left(k^+ - \frac{\ell'^+ + \ell'^+}{2}\right) \delta^{(2)}\left(k_{\perp} - \frac{\ell'_{\perp} + \ell_{\perp}}{2}\right) \\ &\times (2\pi)^3 \delta(x'_q P^+ - x_q P^+) \delta^{(2)}\left(-k'_{q\perp} + k_{q\perp} + \left(1 - \frac{x'_q + x_q}{2}\right)\Delta_{\perp}\right) \delta_{s's} \\ &\times (2\pi)^3 \delta(\ell'' + - x'_q P^+) \delta^{(2)}\left(\ell'_{\perp} - k'_{q\perp} - x'_q \frac{\Delta_{\perp}}{2}\right) \delta_{\sigma'\lambda}(2\pi)^3 \delta(\ell'' + x_q P^+) \delta^{(2)}\left(\ell'_{\perp} - k_q - x_q \frac{\Delta_{\perp}}{2}\right) \delta_{\sigma\lambda} \end{split}$$

$$\begin{split} &= \sum_{\lambda,s} \frac{1}{2} \int \frac{d^2 \mathbf{\Delta}_{\perp}}{(2\pi)^2} \int \frac{dx'_q d^2 \mathbf{k}'_{q\perp}}{(2\pi)^3} \int \frac{dx_q d^2 \mathbf{k}_{q\perp}}{(2\pi)^3} (P^+)^2 \psi_{\lambda s}^{\uparrow *}(x'_q, \mathbf{k}'_{q\perp}) \psi_{\lambda s}^{\uparrow}(x_q, \mathbf{k}_{q\perp}) \\ &\times e^{-i(\mathbf{k}'_{q\perp} - \mathbf{k}_{q\perp} + \frac{x'_q + x_q}{2} \mathbf{\Delta}_{\perp}) \cdot \mathbf{b}_{\perp}} \delta \left( \mathbf{k}^+ - \frac{x'_q + x_q}{2} P^+ \right) \delta^{(2)} \left( \mathbf{k}_{\perp} - \frac{\mathbf{k}'_{q\perp} + \mathbf{k}_{q\perp}}{2} - \frac{x'_z - x_q}{2} \mathbf{\Delta}_{\perp} \right) \\ &\times (2\pi)^3 (x'_x P^+ - x_q P^+) \delta \left( -\mathbf{k}'_{q\perp} + \mathbf{k}_{q\perp} + \left( 1 - \frac{x'_q + x_q}{2} \right) \mathbf{\Delta}_{\perp} \right) \\ &= \sum_{\lambda,s} \int \frac{d^2 \mathbf{\Delta}_{\perp}}{(2\pi)^2} \frac{e^{-i\mathbf{\Delta}_{\perp} \cdot \mathbf{b}_{\perp}}}{16\pi^3} \psi_{\lambda s}^{\uparrow *} \left( x, \mathbf{k}_{\perp} + (1 - x) \frac{\mathbf{\Delta}_{\perp}}{2} \right) \psi_{\lambda s}^{\uparrow} \left( x, \mathbf{k}_{\perp} - (1 - x) \frac{\mathbf{\Delta}_{\perp}}{2} \right). \end{split}$$

This result is consistent with the Drell–Yan–West assignment [46] in the form factor calculations, as demonstrated in the Appendix in Ref. [35]. For the other twist-2 Wigner distributions, similar expressions can be derived but with different combinations of the helicity states. This approach can also be applied to any *N*-particle Fock state.

- [1] K.G. Wilson, Phys. Rev. D 10, 2445 (1974).
- [2] Z. Fodor and C. Hoelbling, Rev. Mod. Phys. 84, 449 (2012).
- [3] J. M. Cornwall, Phys. Rev. D 26, 1453 (1982); C. D. Roberts and A. G. Williams, Prog. Part. Nucl. Phys. 33, 477 (1994); R. Alkofer and L. von Smekal, Phys. Rep. 353, 281 (2001); C. S. Fischer, J. Phys. G 32, R253 (2006); R. J. Holt and C. D. Roberts, Rev. Mod. Phys. 82, 2991 (2010).
- [4] J. M. Maldacena, Int. J. Theor. Phys. 38, 1113 (1999); Adv. Theor. Math. Phys. 2, 231 (1998); S. J. Brodsky and G. F. de Teramond, Phys. Lett. B 582, 211 (2004); G. F. de Teramond and S. J. Brodsky, Phys. Rev. Lett. 94, 201601 (2005); 102, 081601 (2009).
- [5] R. P. Feynman, Phys. Rev. Lett. 23, 1415 (1969).
- [6] J. Bjorken and E. A. Paschos, Phys. Rev. 185, 1975 (1969).
- [7] J. C. Collins and D. E. Soper, Annu. Rev. Nucl. Part. Sci. 37, 383 (1987).
- [8] J. C. Collins and D. E. Soper, Nucl. Phys. B193, 381 (1981);
  B194, 445 (1982); A. Kotzinian, Nucl. Phys. B441, 234 (1995); P. J. Mulders and R. D. Tangerman, Nucl. Phys. B461, 197 (1996); D. W. Sivers, Phys. Rev. D 41, 83 (1990);
  D. Boer and P. J. Mulders, Phys. Rev. D 57, 5780 (1998).
- [9] Müller, D. Robaschik, B. Geyer, F.-M. Dittes, and J. Hořejši, Fortschr. Phys. 42, 101 (1994); K. Goeke, M. V. Polyakov, and M. Vanderhaeghen, Prog. Part. Nucl. Phys. 47, 401 (2001); M. Diehl, Phys. Rep. 388, 41 (2003); X. Ji, Annu. Rev. Nucl. Part. Sci. 54, 413 (2004); A. Belitsky and A. Radyushkin, Phys. Rep. 418, 1 (2005); S. Boffi and B. Pasquini, Riv. Nuovo Cimento 30, 387 (2007).
- [10] N. Balazs and B. Jennings, Phys. Rep. 104, 347 (1984); M. Hillery, R. F. O'Connell, M. O. Scully, and E. P. Wigner, Phys. Rep. 106, 121 (1984); H.-W. Lee, Phys. Rep. 259, 147 (1995).
- [11] K. Vogel and H. Risken, Phys. Rev. A 40, 2847 (1989); D. T.
   Smithey, M. Beck, M. G. Raymer, and A. Faridani, Phys.
   Rev. Lett. 70, 1244 (1993); G. Breitenbach, S. Schiller, and
   J. Mlynek, Nature (London) 387, 471 (1997); K. Banaszek,

C. Radzewicz, K. Wodkiewicz, and J. S. Krasinski, Phys. Rev. A 60, 674 (1999).

- [12] X.-d. Ji, Phys. Rev. Lett. 91, 062001 (2003).
- [13] A. V. Belitsky, X.-D. Ji, and F. Yuan, Phys. Rev. D 69, 074014 (2004).
- [14] C. Lorcé and B. Pasquini, Phys. Rev. D 84, 014015 (2011).
- [15] S. J. Brodsky, H.-C. Pauli, and S. S. Pinsky, Phys. Rep. 301, 299 (1998).
- [16] W. A. Heisenberg, Z. Phys. 43, 172 (1927).
- [17] R. P. Feynman, Photon-Hadron Interactions, *Frontiers in Physics* (Addison-Wesley, Reading, 1972).
- [18] E. P. Wigner, Ann. Math. 40, 149 (1939); H. Melosh, Phys. Rev. D 9, 1095 (1974); see also F. Buccella, C. A. Savoy, and P. Sorba, Lett. Nuovo Cimento 10, 455 (1974).
- [19] B.-Q. Ma, Phys. Lett. B 375, 320 (1996).
- [20] B.-Q. Ma, I. Schmidt, and J. Soffer, Phys. Lett. B **441**, 461 (1998).
- [21] B.-Q. Ma, D. Qing, and I. Schmidt, Phys. Rev. C 65, 035205 (2002); 66, 048201 (2002); T. Liu and B.-Q. Ma, Phys. Rev. C 89, 055202 (2014).
- [22] Z. Lu and B.-Q. Ma, Nucl. Phys. A741, 200 (2004).
- [23] J. She, J. Zhu, and B.-Q. Ma, Phys. Rev. D **79**, 054008 (2009).
- [24] A. Bacchetta, F. Conti, and M. Radici, Phys. Rev. D 78, 074010 (2008).
- [25] Z. Lu and I. Schmidt, Phys. Lett. B 712, 451 (2012).
- [26] M. Burkardt and D. S. Hwang, Phys. Rev. D 69, 074032 (2004).
- [27] D. Chakrabarti and A. Mukherjee, Phys. Rev. D 72, 034013 (2005); D. S. Hwang and D. Mueller, Phys. Lett. B 660, 350 (2008).
- [28] P. A. M. Dirac, Rev. Mod. Phys. 21, 392 (1949).
- [29] G. F. de Teramond, H. G. Dosch, and S. J. Brodsky, Phys. Rev. D 87, 075005 (2013).
- [30] G. D. Cates, C. W. de Jager, S. Riordan, and B. Wojtsekhowski, Phys. Rev. Lett. 106, 252003 (2011).

- [31] C. D. Roberts, M. S. Bhagwat, A. Holl, and S. V. Wright, Eur. Phys. J. Spec. Top. 140, 53 (2007).
- [32] I. C. Cloet, G. Eichmann, B. El-Bennich, T. Klahn, and C. D. Roberts, Few Body Syst. 46, 1 (2009); I. C. Cloet and G. A. Miller, Phys. Rev. C 86, 015208 (2012).
- [33] J. Zhu and B.-Q. Ma, Phys. Rev. D 82, 114022 (2010);
   Phys. Lett. B 696, 246 (2011); Z. Lu, B.-Q. Ma, and J. Zhu,
   Phys. Rev. D 84, 074036 (2011).
- [34] B.-Q. Ma, J. Phys. G 17, L53 (1991); B.-Q. Ma and Q.-R. Zhang, Z. Phys. C 58, 479 (1993).
- [35] B. W. Xiao and B.-Q. Ma, Phys. Rev. D 68, 034020 (2003).
- [36] D. V. Ahluwalia and M. Sawicki, Phys. Rev. D 47, 5161 (1993).
- [37] S. J. Brodsky, T. Huang, and G. P. Lepage, Report No. SLAC-PUB-2540, 1980; Conference Proceedings C810816, 143 (1983); Springer Tracts Mod. Phys. 100, 81 (1982).
- [38] F. Schlumpf, Phys. Rev. D 48, 4478 (1993).
- [39] H. J. Weber, Ann. Phys. (N.Y.) 207, 417 (1991).
- [40] T. Huang, B.-Q. Ma, and Q.-X. Shen, Phys. Rev. D 49, 1490 (1994).
- [41] P. Maris, Few Body Syst. 32, 41 (2002); Few Body Syst. 35, 117 (2004); G. Eichmann, I. C. Cloet, R. Alkofer, A. Krassnigg, and C. D. Roberts, Phys. Rev. C 79, 012202 (2009); G. Eichmann, R. Alkofer, A. Krassnigg, and D. Nicmorus, Phys. Rev. Lett. 104, 201601 (2010); G. Eichmann and C. S. Fischer, Phys. Rev. D 85, 034015 (2012); I. C. Cloet, C. D. Roberts, and A. W. Thomas, Phys. Rev. Lett. 111, 101803 (2013).
- [42] M. V. Terentev, Yad. Fiz. 24, 207 (1976); [Sov. J. Nucl. Phys. 24, 106 (1976)]; V. A. Karmanov, Nucl. Phys. B166, 378 (1980).
- [43] P. L. Chung, F. Coester, and W. N. Polyzou, Phys. Lett. B 205, 545 (1988).

- [44] A. Vega, I. Schmidt, T. Gutsche, and V.E. Lyubovitskij, arXiv:1306.1597.
- [45] E. P. Wigner, Phys. Rev. 40, 749 (1932).
- [46] S. D. Drell and T.-M. Yan, Phys. Rev. Lett. 24, 181 (1970);
   G. B. West, Phys. Rev. Lett. 24, 1206 (1970).
- [47] B.-W. Xiao, X. Qian, and B.-Q. Ma, Eur. Phys. J. A 15, 523 (2002); B.-W. Xiao and B.-Q. Ma, Phys. Rev. D 71, 014034 (2005); J.-h. Yu, B.-W. Xiao, and B.-Q. Ma, J. Phys. G 34, 1845 (2007).
- [48] S. Meissner, A. Metz, and M. Schlegel, J. High Energy Phys. 08 (2009) 056; S. Meissner, A. Metz, M. Schlegel, and K. Goeke, J. High Energy Phys. 08 (2008) 038; C. Lorcé and B. Pasquini, J. High Energy Phys. 09 (2013) 138.
- [49] S. J. Brodsky, D. S. Hwang, and I. Schmidt, Phys. Lett. B 530, 99 (2002); J. C. Collins and A. Metz, Phys. Rev. Lett. 93, 252001 (2004); C. J. Bomhof, P. J. Mulders, and F. Pijlman, Eur. Phys. J. C 47, 147 (2006); J. Collins and J. W. Qiu, Phys. Rev. D 75, 114014 (2007); C. J. Bomhof and P. J. Mulders, Nucl. Phys. B795, 409 (2008); T. C. Rogers and P. J. Mulders, Phys. Rev. D 81, 094006 (2010); J. C. Collins, Phys. Lett. B 536, 43 (2002).
- [50] X.-d. Ji and F. Yuan, Phys. Lett. B 543, 66 (2002).
- [51] G.P. Lepage and S.J. Brodsky, Phys. Rev. D 22, 2157 (1980).
- [52] C. Lorcé, B. Pasquini, X. Xiong, and F. Yuan, Phys. Rev. D 85, 114006 (2012).
- [53] A. Bacchetta and M. Radici, Phys. Rev. Lett. 107, 212001 (2011); Z. Lu and I. Schmidt, Phys. Rev. D 75, 073008 (2007).
- [54] K. Allada *et al.* (Jefferson Lab Hall A Collaboration), Phys. Rev. C 89, 042201 (2014).
- [55] M. Anselmino, M. Boglione, U. D'Alesio, A. Kotzinian, F. Murgia, A. Prokudin, and C. Turk, Phys. Rev. D 75, 054032 (2007).