Detecting new physics in rare top decays at the LHC

Pratishruti Saha,^{1,*} Ken Kiers,^{2,†} David London,^{1,‡} and Alejandro Szynkman^{3,§}

¹*Physique des Particules, Université de Montréal,*

C.P. 6128, succursale centre-ville, Montréal, Quebec H3C 3J7, Canada

²Physics and Engineering Department, Taylor University,

236 West Reade Avenue Upland, Indiana 46989, USA

³IFLP, CONICET–Departamento de Física, Universidad Nacional de La Plata,

C.C. 67, 1900 La Plata, Argentina

(Received 11 July 2014; published 14 November 2014)

In the companion paper it is shown that there are six observables in $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ that can be used to reveal the presence of new physics (NP) in $t \rightarrow b\bar{b}c$. In the present paper we examine the prospects for detecting and identifying such NP at the LHC, in both the short term and long term. To this end, we develop an algorithm for extracting the NP parameters from measurements of the observables. In the short term, depending on what measurements have been made, there are several different ways of detecting the presence of NP. It may even be possible to approximately determine the values of certain NP parameters. In the long term, it is expected that all six observables will be measured. The values of the NP parameters can then be determined reasonably precisely from a fit to these measurements, which will provide good information about the type of NP present in $t \rightarrow b\bar{b}c$.

DOI: 10.1103/PhysRevD.90.094016

PACS numbers: 14.65.Ha

I. INTRODUCTION

Top physics provides a fertile ground for new-physics (NP) searches. With a mass close to the electroweak scale, the top quark may well be sensitive to interactions that do not affect other fermions. In Ref. [1], the companion paper, top decay is investigated for the presence of NP. It is noted that, given the good agreement between the experimental measurement of Γ_t and its theoretical prediction [2], significant NP contributions to top decay can only be present in decay modes that are suppressed in the standard model (SM). One example is $t \rightarrow b\bar{b}c$, whose amplitude involves the small element V_{cb} ($\simeq 0.04$) of the Cabibbo-Kobayashi-Maskawa quark mixing matrix. Reference [1] focuses on this decay at the LHC, where top production occurs predominantly via gluon fusion: $gg \rightarrow t\bar{t}$. The goal is to find observables in the channel $qq \rightarrow t\bar{t}$ with $t \rightarrow b\bar{b}c$ and $\bar{t} \to \bar{b}\ell\bar{\nu}$ that can reveal the presence of NP.

NP contributions to the decay $t \rightarrow b\bar{b}c$ can be parametrized in terms of higher-dimensional operators. If one restricts to dimension-6 operators, then this is realized in the form of ten operators that span all possible Lorentz structures. In Ref. [1], two types of observables are identified that can then be used to get a handle on this NP. The first consists of invariant mass-squared distributions involving the $\{b, c\}$, $\{\bar{b}, c\}$, or $\{b, \bar{b}\}$ quark pairs coming from $t \rightarrow b\bar{b}c$. As for the second type, we note that, in $gg \rightarrow t\bar{t}$, the spins of the t and \bar{t} are correlated. The spin-correlation coefficient $(\kappa_{t\bar{t}})$ depends only on the production process. However, since the *t* quark has an extremely short lifetime, this quantity has to be inferred by measuring the angular correlation between the decay products of the *t* and those of the \bar{t} . If there are NP contributions to the decay, the inferred value of $\kappa_{t\bar{t}}$ is necessarily altered from that of the SM. It is this feature that provides information about the NP. Therefore, the second type of observable consists of these angular correlations. They are taken between the ℓ^- coming from the \bar{t} decay and one of \bar{b} , *b* or *c* coming from the *t* decay.

The NP operators not only change the top branching fraction of this decay, but also modify the shapes of these distributions. It is shown in Ref. [1] that the NP contributions to all of the above observables can be written in terms of certain combinations of the NP couplings denoted as \hat{A}_i^{σ} . Furthermore, the observables are found to be practically unaffected by parton densities, etc., so that they provide direct access to the values of these \hat{A}_i^{σ} 's.

Now, the observables described above involve the \bar{b} quark coming from the decay of the *t*. However, there is also a \bar{b} produced in the \bar{t} decay. A realistic analysis must deal with the question of how to distinguish the two \bar{b} 's. In addition, while the focus in Ref. [1] was entirely on $t\bar{t}$ production from gluon fusion, there is also a contribution from $q\bar{q} \rightarrow t\bar{t}$ which must be considered.

In the present paper we address these issues. In Ref. [1], the analytical expressions for the observables are compared with the results of a numerical simulation of the LHC using MADGRAPH 5 [3]. Here we extend our MADGRAPH 5 simulations to examine different strategies for extracting the NP parameters. In so doing, we include a method for

pratishruti.saha@umontreal.ca

knkiers@taylor.edu

london@lps.umontreal.ca

[§]szynkman@fisica.unlp.edu.ar

distinguishing the two \bar{b} 's. We also take the contribution from $q\bar{q} \rightarrow t\bar{t}$ into account, examining its effect on the aforementioned observables and their sensitivity to the \hat{A}_i^{σ} 's. In our simulations we consider numbers of total events representative of LHC measurements in both the short and long terms. While the long-term results obviously have smaller errors, it is still possible in the short term to detect and partially identify NP in $t \rightarrow b\bar{b}c$.

We begin in Sec. II by summarizing the results of Ref. [1]. We present the NP operators that contribute to $t \rightarrow b\bar{b}c$, as well as the $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ observables that can reveal the presence of the NP. In Sec. III we develop the algorithm to extract the NP parameters from the observables. We discuss the MADGRAPH 5 simulations in Sec. IV, and apply the algorithm. Here we show that the measurement of the observables at the LHC can lead to the detection of the NP, and possibly even its identification.¹ We conclude in Sec. V.

II. NEW PHYSICS IN TOP DECAY

In this section, we summarize the main results of Ref. [1].

A. $t \rightarrow b\bar{b}c$: Effective Lagrangian

In the SM, the decay $t \to b\bar{b}c$ proceeds through $t \to W^+ b$, followed by $W^+ \to \bar{b}c$. The NP contributions to this can be parametrized by the effective Lagrangian $\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{eff}}^V + \mathcal{L}_{\text{eff}}^S + \mathcal{L}_{\text{eff}}^T$, with

$$\mathcal{L}_{\text{eff}}^{V} = 4\sqrt{2}G_{F}V_{cb}V_{tb}\{X_{LL}^{V}\bar{b}\gamma_{\mu}P_{L}t\bar{c}\gamma^{\mu}P_{L}b + X_{LR}^{V}\bar{b}\gamma_{\mu}P_{L}t\bar{c}\gamma^{\mu}P_{R}b + X_{RL}^{V}\bar{b}\gamma_{\mu}P_{R}t\bar{c}\gamma^{\mu}P_{L}b + X_{RR}^{V}\bar{b}\gamma_{\mu}P_{R}t\bar{c}\gamma^{\mu}P_{R}b\} + \text{H.c.},$$
(1)

$$\mathcal{L}_{\text{eff}}^{S} = 4\sqrt{2}G_{F}V_{cb}V_{tb}\{X_{LL}^{S}\bar{b}P_{L}t\bar{c}P_{L}b + X_{LR}^{S}\bar{b}P_{L}t\bar{c}P_{R}b + X_{RL}^{S}\bar{b}P_{R}t\bar{c}P_{L}b + X_{RR}^{S}\bar{b}P_{R}t\bar{c}P_{R}b\} + \text{H.c.}, \quad (2)$$

$$\mathcal{L}_{\text{eff}}^{T} = 4\sqrt{2}G_{F}V_{cb}V_{tb}\{X_{LL}^{T}\bar{b}\sigma^{\mu\nu}P_{L}t\bar{c}\sigma_{\mu\nu}P_{L}b + X_{RR}^{T}\bar{b}\sigma^{\mu\nu}P_{R}t\bar{c}\sigma_{\mu\nu}P_{R}b\} + \text{H.c.}, \qquad (3)$$

Here the color indices are assumed to contract in the same way as in the SM (i.e., the fields \overline{b} with *t* and \overline{c} with *b*). Color-mismatched terms, in which the indices contract in the opposite way, may occur in certain models and can be incorporated in a straightforward manner [4].

The NP couplings (the X_{AB}^{I} in the above equations) contain weak phases, but the strong phases are negligible [5]. In addition, the X_{AB}^{I} may be assumed, quite generally,

to be O(1). The sizes of the SM and NP contributions to $t \rightarrow b\bar{b}c$ would then be roughly equal. This shows that it is important to include both the SM-NP and NP-NP interference pieces when computing the effect of NP on a particular observable.

B. Observables in $gg \to t\bar{t} \to (b\bar{b}c)(\bar{b}\ell\bar{\nu})$

The kinematics of $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ is represented in Fig. 1. The six-body phase space is decomposed into five solid angles $d\Omega_1^{**}$, $d\Omega_2^*$, $d\Omega_4^{**}$, $d\Omega_5^*$ and $d\Omega_t$, and two invariant masses M_2 and M_5 . The * and ** superscripts on the solid angles indicate that these angles are defined in reference frames that are, respectively, one and two boosts away from the $t\bar{t}$ rest frame. M_2 and M_5 are defined by $M_2^2 = (p_1 + p_2)^2$ and $M_5^2 = (p_4 + p_5)^2$. Note that p_1 , p_2 and p_3 are the momenta of the b, \bar{b} and c quarks in $t \rightarrow b\bar{b}c$, but all permutations are allowed. The observables use several of these possibilities.

The differential cross section for $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ is computed in Ref. [1]. It is a function of the finalstate momenta p_i (i = 1, 2, ..., 6) and SM and NP couplings, and is defined with respect to $dM_2^2 dM_5^2 d\Omega_1^{**} d\Omega_2^* d\Omega_4^{**} d\Omega_5^* d\Omega_t$. The cross section is then integrated over M_5^2 and over all angles except for θ_2^* and θ_ℓ^* . The observables are obtained by (i) assigning the p_i to specific final-state particles, and (ii) integrating further over θ_2^* and θ_ℓ^* , or M_2^2 .

There are three possibilities for the particle assignments: (i) $p_1 = p_c$, $p_2 = p_b$, $p_3 = p_{\bar{b}_1}$, (ii) $p_1 = p_c$, $p_2 = p_{\bar{b}_1}$, $p_3 = p_b$, (iii) $p_1 = p_b$, $p_2 = p_{\bar{b}_1}$, $p_3 = p_c$. Here $p_{\bar{b}_1}$ refers to the \bar{b} coming from the *t*. Also, $p_6 = p_{\ell^-}$. For each case there are two observables.



FIG. 1. Kinematics for the process $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ [6]. Ω_1^{**} denotes the direction of \vec{p}_1^{**} in the rest frame of M_2 , relative to the direction of $\vec{p}_1^* + \vec{p}_2^*$, where $M_2^2 = (p_1 + p_2)^2$. Similarly, Ω_2^* denotes the direction of $(\vec{p}_1^* + \vec{p}_2^*)$ in the *t* rest frame, relative to the direction of \vec{p}_1 in the $t\bar{t}$ rest frame. Ω_t denotes the direction of \vec{p}_t relative to \vec{q}_1 , also in the $t\bar{t}$ rest frame. The solid angles Ω_4^{**} and Ω_5^* are defined analogously to Ω_1^{**} and Ω_2^* , respectively, and $M_5^2 = (p_4 + p_5)^2$.

¹When we refer to the "identification of NP," what is implied is the measurement of the various \hat{A}_i^{σ} 's and $\text{Re}(X_{LL}^V)$.

(1) Invariant mass-squared distribution:

$$\frac{d\sigma}{d\zeta_{12}^2} = \sigma_{\rm SM} \bigg\{ F_{12} \frac{6h_{\rm SM}^{12}(\zeta_{12}^2)}{(1-\zeta_W^2)^2(1+2\zeta_W^2)} + \frac{3G_F m_i^2}{\sqrt{2}\pi^2(1-\zeta_W^2)^2(1+2\zeta_W^2)} \sum_{i,\sigma} \hat{A}_i^{\sigma} h_i^{12}(\zeta_{12}^2) \bigg\}.$$
(4)

(2) Angular correlation:

$$\frac{d\sigma}{d\cos\theta_{3}^{*}d\cos\theta_{\ell}^{*}} = \frac{\sigma_{\rm SM}}{4} \left\{ \left[1 + \rho_{3}(\zeta_{W}^{2})\kappa(r)\cos\theta_{3}^{*}\cos\theta_{\ell}^{*} \right] + \frac{3G_{F}m_{t}^{2}}{4\sqrt{2}\pi^{2}(1-\zeta_{W}^{2})^{2}(1+2\zeta_{W}^{2})} \left[\left(\sum_{i,\sigma}\hat{A}_{i}^{\sigma}\right) + \left(\hat{A}_{3}^{+} - \hat{A}_{3}^{-} - \frac{1}{3}(\hat{A}_{1}^{+} - \hat{A}_{1}^{-} + \hat{A}_{2}^{+} - \hat{A}_{2}^{-}) \right) \kappa(r)\cos\theta_{3}^{*}\cos\theta_{\ell}^{*} \right] \right\}.$$
(5)

The numerical subscripts and superscripts correspond to the particle assignments in each of the three cases. That is, in case (i), the subscript 12 corresponds to bc (b is particle 1, c is particle 2), and similarly for cases (ii) $(12 = \bar{b}c)$ and (iii) $(12 = b\bar{b})$. In the summations, $\sigma = +, -$ and $i = b, \bar{b}, c$.

 $\sigma_{\rm SM}$ is given in Eq. (60) of the Appendix of Ref. [1], $\zeta_{12}^2 \equiv (p_1 + p_2)^2 / m_t^2$, $\zeta_W \equiv m_W / m_t$, and $\kappa(r)$ is defined as

$$\kappa(r) = \frac{(-31r^4 + 37r^2 - 66)r - 2(r^6 - 17r^4 + 33r^2 - 33)\tanh^{-1}(r)}{r^2[(31r^2 - 59)r + 2(r^4 - 18r^2 + 33)\tanh^{-1}(r)]},\tag{6}$$

In addition,

where

$$r \equiv \sqrt{1 - 4m_t^2/Q^2}, \qquad Q \equiv p_t + p_{\bar{t}}.$$
 (7)

The functions h_i^{mn} $(mn = bc, \bar{b}c, b\bar{b}; i = b, \bar{b}, c)$ are defined in Table I, and

 $h_{\rm SM}^{b\bar{b}}(\zeta_{b\bar{b}}^2) = (1-\zeta_W^2-\zeta_{b\bar{b}}^2)(\zeta_W^2+\zeta_{b\bar{b}}^2)\theta(1-\zeta_W^2-\zeta_{b\bar{b}}^2).$

 $h_{\rm SM}^{bc}(\zeta_{bc}^2) = (1 - \zeta_{bc}^2)\zeta_{bc}^2\theta(1 - \zeta_W^2 - \zeta_{bc}^2)$

$$\begin{array}{ll} \text{The functions } h_i^{mn} \ (mn = bc, \bar{b}c, b\bar{b}; \ i = b, \bar{b}, c) \ \text{are defined in Table I, and} \\ h_{\text{SM}}^{bc}(\zeta_{\bar{b}c}^2) = (1 - \zeta_{\bar{b}c}^2)\zeta_{\bar{b}c}^2\theta(1 - \zeta_W^2 - \zeta_{\bar{b}c}^2), \\ h_{\text{SM}}^{\bar{b}c}(\zeta_{\bar{b}c}^2) = \left(\frac{\zeta_W\gamma_W}{6\pi}\right) \frac{(1 - \zeta_{\bar{b}c}^2)^2(1 + 2\zeta_{\bar{b}c}^2)}{(\zeta_{\bar{b}c}^2 - \zeta_W^2)^2 + (\zeta_W\gamma_W)^2}, \end{array}$$

Note that we have neglected some mild dependence on
$$\operatorname{Re}(X_{LL}^V)$$
 in the *bc* and $b\bar{b}$ distributions. This dependence is, however, properly taken into account in the numerical work below.

 $\rho_{\bar{b}}(\zeta_W^2) = 1, \qquad \rho_b(\zeta_W^2) = -\left(\frac{1 - 2\zeta_W^2}{1 + 2\zeta_W^2}\right),$

(9)

(10)

TABLE I. Definitions of the h_i^{mn} ($mn = bc, \bar{b}c, b\bar{b}$) functions. The columns correspond to $i = b, \bar{b}, c$.

	b	\bar{b}	С
$ \frac{h_i^{bc}(\zeta^2)}{h_i^{\bar{b}c}(\zeta^2)} $ $ h^{b\bar{b}}(\zeta^2) $	$\frac{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}{3(1-\zeta^2)^2\zeta^2}$ $\frac{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}{2\zeta^2}$	$\frac{3(1-\zeta^2)^2\zeta^2}{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}\\\frac{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}$	$\frac{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}{\frac{1}{2}(1-\zeta^2)^2(1+2\zeta^2)}$ $\frac{3}{3}(1-\zeta^2)^2\zeta^2$

(8)

The NP parameters appear in the observables in the \hat{A}_i^{σ} 's:

$$\begin{split} \hat{A}_{\bar{b}}^{+} &= 4|X_{LL}^{V}|^{2} - 8\text{Re}(X_{LL}^{T}X_{LL}^{S*}) + 32|X_{LL}^{T}|^{2}, \\ \hat{A}_{\bar{b}}^{-} &= 4|X_{RR}^{V}|^{2} - 8\text{Re}(X_{RR}^{T}X_{RR}^{S*}) + 32|X_{RR}^{T}|^{2}, \\ \hat{A}_{\bar{b}}^{+} &= |X_{LL}^{S}|^{2} + |X_{LR}^{S}|^{2} - 16|X_{LL}^{T}|^{2}, \\ \hat{A}_{\bar{b}}^{-} &= |X_{RR}^{S}|^{2} + |X_{RL}^{S}|^{2} - 16|X_{RR}^{T}|^{2}, \\ \hat{A}_{c}^{+} &= 4|X_{LR}^{V}|^{2} + 8\text{Re}(X_{LL}^{T}X_{LL}^{S*}) + 32|X_{LL}^{T}|^{2}, \\ \hat{A}_{c}^{-} &= 4|X_{RL}^{V}|^{2} + 8\text{Re}(X_{RR}^{T}X_{RR}^{S*}) + 32|X_{RR}^{T}|^{2}. \end{split}$$
(11)

As pointed out in the Introduction, the six observables have different functional dependencies on the \hat{A}_i^{σ} 's.

III. EXTRACTING NEW PHYSICS FROM OBSERVABLES

In the companion paper [1], we showed that the chosen observables, namely the three invariant mass-squared distributions $(d\sigma/d\zeta_{12}^2)$ and the three angular correlations $(d\sigma/d\cos\theta_3^*\cos\theta_\ell^*)$, are sensitive to certain combinations of the new-physics parameters. This dependence is represented in Eqs. (4) and (5) as combinations of the various \hat{A}_i^{σ} 's, which in turn can be defined [Eq. (11)] in terms of the NP coefficients X_{AB}^{I} that appear in the effective Lagrangian. In addition, the NP operator proportional to X_{II}^V has the same Lorentz structure as the corresponding SM operator, so that these two interfere, leading to an explicit dependence on X_{LL}^V itself. As discussed in Ref. [1], the shapes of the observables are largely insensitive to effects due to parton distribution functions (PDFs). It is only natural then to expect that it should be possible to extract the various combinations of NP parameters by fitting these distributions. Further, one expects that, by combining information from all six observables, it should be possible to extract the values of the individual \hat{A}_i^{σ} 's and $\operatorname{Re}(X_{LL}^V)$ as well. In this section we develop the procedure to carry out this extraction.

A. Algorithm

Consider first the conventional fitting method, which involves individual observables. If the underlying theory has N theoretical unknowns, their values can be determined only if at least N observables are measured. There are theoretical expressions for these observables in terms of the N unknowns. Using these expressions, the best-fit values of the unknowns are those for which the measured values of the observables are best reproduced.

In our case, the observables are distributions and correlations, specifically $d\sigma/d\zeta_{12}^2$ [Eq. (4)]and $d\sigma/d\cos\theta_3^*\cos\theta_\ell^*$ [Eq. (5)]. Each distribution/correlation contains many measurements [at different values of ζ_{12}^2 or $(\cos \theta_3^*, \cos \theta_\ell^*)$]. Equations (4) and (5) show that each of the ζ_{12}^2 distributions and the angular correlations can be written as a linear combination of the SM piece and several NP pieces [six \hat{A}_i^{σ} 's, Re (X_{LL}^V)]. With this in mind, we use MADGRAPH 5 in conjunction with FEYNRULES [7] to generate eight templates for each distribution/correlation. The templates are nothing but the said distributions/ correlations generated with the input values of the parameters chosen such that certain specific contributions to the observables are retained, while all others are set to zero. The objective is to isolate the contributions coming from the SM, the individual \hat{A}_i^{σ} 's and Re (X_{LL}^V) . Table II gives the parameter choices made for each template (labeled TM-i) and the contributions that they represent.

Once we have these templates, a ζ_{12}^2 distribution or an angular correlation arising from a generic choice of NP parameters can be represented as a linear combination of the corresponding templates with appropriate coefficients. Extracting these coefficients allows one to determine the values of the NP parameters involved.

B. Testing the algorithm

We use "pseudodata" generated in Monte Carlo simulations to test our fitting procedure. Once again, we use MADGRAPH 5 to generate these samples. In Ref. [1] we presented plots of the normalized distributions and correlations. Here we use the unnormalized distributions for the

Template	X^{I}_{AB}	\hat{A}_i^{σ}	Description
TM-0	All $X_{AB}^I = 0$	All $\hat{A}_i^{\sigma} = 0$	SM contribution
TM-1	$X_{LL}^S, X_{LL}^{\overline{T}} \neq 0$	$\hat{A}^+_{\bar{h}} \neq 0$; all other $\hat{A}^{\sigma}_i = 0$	Contribution $\propto \hat{A}_{\vec{b}}^+$
TM-2	$X_{RR}^V \neq 0$	$\hat{A}_{\bar{h}}^{\sigma} \neq 0$; all other $\hat{A}_{i}^{\sigma} = 0$	Contribution $\propto \hat{A}_{\bar{b}}^-$
TM-3	$X_{LL}^S \neq 0$	$\hat{A}_{h}^{+} \neq 0$; all other $\hat{A}_{i}^{\sigma} = 0$	Contribution $\propto \hat{A}_b^+$
TM-4	$X_{RR}^{S} \neq 0$	$\hat{A}_{h}^{\sigma} \neq 0$; all other $\hat{A}_{i}^{\sigma} = 0$	Contribution $\propto \hat{A}_b^-$
TM-5	$X_{LR}^V \neq 0$	$\hat{A}_{c}^{+} \neq 0$; all other $\hat{A}_{i}^{\sigma} = 0$	Contribution $\propto \hat{A}_c^+$
TM-6	$X_{RL}^V \neq 0$	$\hat{A}_{c}^{-} \neq 0$; all other $\hat{A}_{i}^{\sigma} = 0$	Contribution $\propto \hat{A}_c^-$
TM-7	$X_{LL}^{V} \neq 0$	$\hat{A}^+_{\bar{h}} \neq 0$; all other $\hat{A}^{\sigma}_i = 0$	Contributions $\propto \operatorname{Re}(X_{LL}^V)$ and $\hat{A}_{\bar{b}}^+$
TM-8	1	TM-7 – TM-1	Contribution $\propto \operatorname{Re}(X_{LL}^V)$

TABLE II. NP parameter choices for each of the templates TM-i.

Test case	X^{I}_{AB}	\hat{A}^{σ}_{i}	$\sigma/\sigma_{ m SM}$
EX-1	$X_{LL}^T = 1; X_{RR}^T = 1$	$\hat{A}^+_{ar{b}}=32;\hat{A}^+_{b}=-16;\hat{A}^+_{c}=32;$	3.1
		$\hat{A}_{\bar{b}}^{-} = 32; \hat{A}_{\bar{b}}^{-} = -16; \hat{A}_{\bar{c}}^{-} = 32$	
EX-2	$X_{LR}^S = 5$	$\hat{A}_{h}^{+} = 25$; all other \hat{A}_{i}^{σ} 's = 0	1.6
EX-3	$X_{LR}^S = 3; X_{RL}^S = 4$	$\hat{A}_{b}^{+} = 9; \hat{A}_{b}^{-} = 16;$ all other \hat{A}_{i}^{σ} 's = 0	1.6
EX-4	$X_{LL}^V = 3; X_{LL}^S = 5$	$\hat{A}_{\bar{b}}^{+} = 36; \hat{A}_{b}^{+} = 25; \text{all other } \hat{A}_{i}^{\sigma}\text{'s} = 0$	2.4

TABLE III. Input values of the NP parameters for the four test cases EX-i. The last column illustrates how the total cross section σ is affected in each of the test cases.

fitting. Since both the templates and the "data sets" are obtained with MADGRAPH 5 with the same choices of PDFs, scale, etc., the overall normalization is automatically accounted for.

Our procedure is as follows. We generate pseudodata using MADGRAPH 5, in conjunction with FEYNRULES, for certain chosen values of the NP parameters (i.e., the X_{AB}^I). This gives us three ζ_{12}^2 distributions and three angular correlations. We divide each of these into 25 bins (using a 5×5 array for the angular correlations). We then perform a single χ^2 minimization involving all six histograms in order to determine the coefficients for the templates that result in the best fit for all six observables simultaneously. For this purpose, we use standard, publicly available routines [8]. Finally, we examine to what extent the values of the NP parameters extracted from the fit agree with their input values.

We consider four different test cases of pseudodata, which we label EX-1, EX-2, EX-3 and EX-4. The values of the input NP parameters for these test cases are listed in Table III, along with the size of the cross section relative to the SM prediction. The data sets EX-i have been generated for the process $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}e^-\bar{\nu}_e)$, taking a benchmark luminosity that corresponds to 10^5 SM events. The uncertainties incorporated in the fitting procedure are statistical only, and are estimated by considering the number of events in each bin in the histograms to be Poisson distributed. The templates TM-i have been generated for the same process but with $O(10^6)$ events, so that uncertainties from these can be neglected in the fit.

1. Fit 1

As detailed above, the templates are generated by assuming there is only a single contribution at a time to the distributions/correlations for the process $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}e^{-}\bar{\nu}_{e})$. The observables are represented by the analytical expressions in Eqs. (4) and (5). However, these expressions have been derived [1] under the (unrealistic) assumption that the two \bar{b} 's in the final state are distinguishable. In Fit 1, as a first test of the algorithm, we retain this assumption.

Table IV shows the values of the \hat{A}_i^{σ} 's and $\operatorname{Re}(X_{LL}^V)$ extracted from the fit for the four different test cases of

pseudodata in Table III. A comparison of the two tables shows that most of the values of the parameters extracted from the fit agree with their input values within $\pm 1\sigma$. This demonstrates that the fundamental idea of the algorithm, namely fitting using the templates, is sound.

The worst-fitted parameter is $\operatorname{Re}(X_{LL}^V)$ in the case where $\hat{A}_{\bar{b}}^+$ is nonzero but $\operatorname{Re}(X_{LL}^V) = 0$. This poor fit is an artifact of the somewhat simple-minded fitting procedure that we adopt: $\hat{A}_{\bar{b}}^+$ and $\operatorname{Re}(X_{LL}^V)$ are treated as independent parameters in the fit, despite the fact that they are correlated [see Eq. (11)]. Note that the contribution proportional exclusively to $\operatorname{Re}(X_{LL}^V)$ appears primarily in $d\sigma/d\zeta_{\bar{b}c}^2$.² Even so, the fit performs rather well when $\operatorname{Re}(X_{LL}^V)$ is, in fact, nonzero. On the other hand, not considering $\operatorname{Re}(X_{LL}^V)$ as a fit parameter leads to an overall worsening of the fits. For this reason we retain it in our fitting algorithm, while taking care to avoid drawing any strong conclusions from the extracted value of this parameter.

2. Fit 2

In Fit 2 we drop the assumption that the two final-state \bar{b} 's are distinguishable. The Monte Carlo pseudodata (as well as the templates) for the process $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}e^-\bar{\nu}_e)$ now include the amplitudes in which the momenta of the two \bar{b} 's in the final state are exchanged. However, in order to construct the above observables, we necessarily need to identify the \bar{b} emerging from the top decay. Hence we must restrict our analysis to regions of phase space where the two \bar{b} 's can effectively be considered to be distinguishable. To do this, we construct the two quantities $m_1^2 = (p_b + p_c + p_{\bar{b}_1})^2$ and $m_2^2 = (p_b + p_c + p_{\bar{b}_2})^2$. If both m_1 and m_2 lie within the range $m_t \pm 15\Gamma_t$, the event is discarded. Otherwise, it is assumed that the \bar{b}_i that yields the smaller value of $|m_i - m_t|$ comes from the *t*-quark decay. This leads to a loss of about 20% of

²Based on our theoretical analysis, we expect the $\zeta_{\bar{b}c}^2$ distribution to have the most sensitivity to $\operatorname{Re}(X_{LL}^V)$. This expectation is confirmed by an examination of the templates. Having said this, the dependence on $\operatorname{Re}(X_{LL}^V)$ is not completely negligible for the other distributions and correlations, and in our numerical work we include the corresponding template (TM-8) in the fits for all distributions and correlations.

Test Case	Fit Results		$\chi^2/d.o.f.$
EX-1	SM coefficient = 1.005 ± 0.003 $\hat{A}_{b}^{+} = 33 \pm 2$ $\hat{A}_{b}^{+} = -16 \pm 2$ $\hat{A}_{c}^{+} = 33 \pm 2$ Re $(X_{U}^{V}) = 0.40 \pm 0.02$	$\hat{A}_{\bar{b}}^{-} = 30 \pm 2$ $\hat{A}_{\bar{b}}^{-} = -15 \pm 2$ $\hat{A}_{c}^{-} = 31 \pm 2$	1.30
EX-2	SM coefficient = 1.000 ± 0.002 $\hat{A}_{b}^{+} = 0 \pm 1$ $\hat{A}_{b}^{+} = 24 \pm 1$ $\hat{A}_{c}^{+} = 1 \pm 1$ $\text{Re}(X_{U}^{V}) = 0.01 \pm 0.02$	$egin{array}{lll} \hat{A}_{ar{b}}^{-} = 0 \pm 1 \ \hat{A}_{ar{b}}^{-} = 1 \pm 1 \ \hat{A}_{c}^{-} = 0 \pm 1 \end{array}$	1.21
EX-3	SM coefficient = 0.994 ± 0.002 $\hat{A}_{\bar{b}}^{+} = 1 \pm 1$ $\hat{A}_{\bar{b}}^{+} = 10 \pm 1$ $\hat{A}_{c}^{+} = 0 \pm 1$ $\text{Re}(X_{LL}^{V}) = 0.02 \pm 0.02$	$\hat{A}_{ar{b}}^- = -1 \pm 1$ $\hat{A}_{b}^- = 15 \pm 1$ $\hat{A}_{c}^- = 0 \pm 1$	1.22
EX-4	SM coefficient = 1.003 ± 0.003 $\hat{A}_{\bar{b}}^{+} = 36 \pm 1$ $\hat{A}_{\bar{b}}^{+} = 25 \pm 1$ $\hat{A}_{c}^{+} = 1 \pm 1$ Re $(X_{LL}^{V}) = 3.03 \pm 0.01$	$\hat{A}_{ar{b}}^{-} = 0 \pm 1$ $\hat{A}_{ar{b}}^{-} = -1 \pm 1$ $\hat{A}_{c}^{-} = -1 \pm 1$	1.43

TABLE IV. Values of the NP parameters extracted from the four test cases EX-i using Fit 1.

the events. This cut also distorts the angular correlation such that it no longer conforms to the familiar $(a_1 + a_2 \cos \theta_3^* \cos \theta_{\ell}^*)$ form, even for the SM. Despite the distortion, the fit can be performed using the same algorithm as long as the same method of event selection is applied to the pseudodata as well as the templates.

The results of the fit are presented in Table V for the four test cases. The agreement between the values of the fitted

TABLE V. Values of the NP parameters extracted from the four test cases EX-i using Fit 2.

Test case	Fit results		$\chi^2/d.o.f.$
EX-1	SM coefficient = 1.005 ± 0.004 $\hat{A}_{\bar{b}}^{+} = 29 \pm 2$ $\hat{A}_{\bar{b}}^{+} = -15 \pm 2$ $\hat{A}_{c}^{+} = 33 \pm 2$ $\text{Re}(X_{II}^{V}) = 0.40 \pm 0.03$	$\hat{A}_{\bar{b}}^{-} = 33 \pm 2$ $\hat{A}_{\bar{b}}^{-} = -16 \pm 2$ $\hat{A}_{c}^{-} = 31 \pm 2$	1.09
EX-2	SM coefficient = 0.999 ± 0.003 $\hat{A}_{\bar{b}}^{\pm} = 1 \pm 1$ $\hat{A}_{\bar{b}}^{+} = 26 \pm 1$ $\hat{A}_{c}^{+} = 1 \pm 2$ $\operatorname{Re}(X_{LL}^{V}) = -0.01 \pm 0.02$	$\hat{A}_{ar{b}}^{-} = -1 \pm 1$ $\hat{A}_{b}^{-} = 0 \pm 1$ $\hat{A}_{c}^{-} = -2 \pm 2$	1.05
EX-3	SM coefficient = 1.005 ± 0.003 $\hat{A}_{\bar{b}}^{+} = -1 \pm 1$ $\hat{A}_{\bar{b}}^{+} = 10 \pm 1$ $\hat{A}_{c}^{+} = 0 \pm 2$ Re $(X_{LL}^{V}) = -0.01 \pm 0.02$	$\hat{A}_{ar{b}}^{-} = 1 \pm 1$ $\hat{A}_{ar{b}}^{-} = 15 \pm 1$ $\hat{A}_{c}^{-} = 0 \pm 2$	0.89
EX-4	SM coefficient = 0.997 ± 0.003 $\hat{A}_{b}^{+} = 38 \pm 2$ $\hat{A}_{b}^{+} = 23 \pm 2$ $\hat{A}_{c}^{+} = 1 \pm 2$ $\text{Re}(X_{LL}^{V}) = 2.98 \pm 0.02$	$\hat{A}_{ar{b}}^{-} = -1 \pm 2$ $\hat{A}_{ar{b}}^{-} = 2 \pm 2$ $\hat{A}_{c}^{-} = -1 \pm 2$	1.29

Test Case	Fit Results		$\chi^2/d.o.f.$
EX-1	SM coefficient = 1.002 ± 0.004 ; $\hat{A}_{b}^{+} = 32 \pm 3$; $\hat{A}_{b}^{+} = -16 \pm 2$; $\hat{A}_{c}^{+} = 32 \pm 3$; Re $(X_{VL}^{V}) = 0.42 \pm 0.03$;	$\hat{A}_{\overline{b}} = 31 \pm 3$ $\hat{A}_{\overline{b}} = -15 \pm 2$ $\hat{A}_{\overline{c}} = 33 \pm 3$	1.01
EX-2	SM coefficient = 0.998 ± 0.003; $\hat{A}_{\bar{b}}^{+} = -1 \pm 2;$ $\hat{A}_{\bar{b}}^{+} = 24 \pm 2;$ $\hat{A}_{c}^{+} = 1 \pm 2;$ Re $(X_{II}^{V}) = -0.01 \pm 0.02;$	$egin{array}{lll} \hat{A}_{ar{b}} &= 0 \pm 2 \ \hat{A}_{ar{b}} &= 1 \pm 2 \ \hat{A}_{ar{c}} &= -1 \pm 2 \end{array} \end{array}$	1.00
EX-3	SM coefficient = 1.001 ± 0.003 ; $\hat{A}_{b}^{+} = 1 \pm 2$; $\hat{A}_{b}^{+} = 9 \pm 2$; $\hat{A}_{c}^{+} = 1 \pm 2$; Re $(X_{LL}^{V}) = -0.01 \pm 0.02$;	$\hat{A}_{\overline{b}}^- = -1 \pm 2$ $\hat{A}_{\overline{b}}^- = 16 \pm 2$ $\hat{A}_{\overline{c}}^- = -1 \pm 2$	1.08
EX-4	SM coefficient = 0.999 ± 0.003 ; $\hat{A}_{b}^{+} = 38 \pm 2$; $\hat{A}_{b}^{+} = 24 \pm 2$; $\hat{A}_{c}^{+} = 1 \pm 3$; Re $(X_{LL}^{V}) = 2.97 \pm 0.02$;	$\hat{A}_{ar{b}}^{-} = -2 \pm 2$ $\hat{A}_{ar{b}}^{-} = 1 \pm 2$ $\hat{A}_{c}^{-} = -1 \pm 2$	1.01

TABLE VI. Values of the NP parameters extracted from the four test cases EX-i using Fit 3.

parameters and their input values is almost as good as in the idealized case (Fit 1): apart from $\text{Re}(X_{LL}^V)$ in EX-1, all values agree within $\pm 1.5\sigma$. We find that, with the event selection discussed above, despite the resulting loss of statistics, one obtains a slight improvement in the goodness of fit, as can be seen from the smaller values of $\chi^2/\text{d.o.f.}$ All of this demonstrates that our algorithm continues to hold, even when one imposes a "cut" to distinguish the two \bar{b} 's in the final state.

3. Fit 3

Finally, at the LHC, there is a small ($\approx 10\%$ -15%) contribution to $t\bar{t}$ production from $q\bar{q}$ annihilation. In Fit 3 we consider the impact of this additional contribution.

It must be said that we do not expect a significant effect. Since the NP couplings play a role only in top decay, the structure of Eqs. (4) and (5) remains largely unchanged. The change in Eq. (4) is the analytical form of the factor $\sigma_{\rm SM}$; in Eq. (5), the changes appear in the expressions for $\sigma_{\rm SM}$ and $\kappa(r)$. The decomposition of the NP contribution in terms of a linear combination of \hat{A}_i^{σ} 's and $\operatorname{Re}(X_{LL}^V)$ therefore remains valid for the purposes of the fit. Moreover, in the $\sqrt{\hat{s}}$ range that is sampled,³ $t\bar{t}$ production is overwhelmingly dominated by gg fusion, simply because the gluon density is large at low values of momentum fractions (the well-known Bjorken x_1 and x_2). This means that the corrections due to $q\bar{q} \rightarrow t\bar{t}$ are small in magnitude over the entire region of phase space that can be probed.

The results of Fit 3 are presented in Table VI. As expected, the fitting procedure described above proves just as effective for the full process $pp \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}e^{-}\bar{\nu}_{e})$.

Although the essential structure of our statistical analysis is based on the analytical expressions obtained in Ref. [1], where several simplifying assumptions were made, through the above series of fits we have obtained a reliable algorithm that includes a procedure to distinguish the two final-state \bar{b} 's, and works well even in the presence of the contribution from $q\bar{q} \rightarrow t\bar{t}$. We now use this algorithm to examine the prospects for obtaining information about NP in the decay $t \rightarrow b\bar{b}c$ at the LHC, in both the short and long terms. This is discussed in the next section.

IV. DETECTING NEW PHYSICS IN TOP DECAY

Above, we have established a method for the extraction of NP parameters involved in the decay $t \rightarrow b\bar{b}c$. However, should a sizeable NP contribution exist, it is likely that it would first be detected simply by measuring the total cross section in this channel. It is only afterwards that the ζ_{12}^2 distributions and the angular correlations discussed in the preceding sections would be used to indicate the presence of NP. While this is true, it should also be pointed out that the overall normalization of the cross section suffers from inherent theoretical uncertainties such as the choice of

³At a 14 TeV pp collider, with $\mathcal{O}(10^5)$ events, this range is approximately 350 to 1200 GeV.



FIG. 2 (color online). Normalized $d\sigma/d\zeta_{12}^2$ distributions: (a) $d\sigma/d\zeta_{bc}^2$, (b) $d\sigma/d\zeta_{b\bar{b}}^2$, (c) $d\sigma/d\zeta_{\bar{b}c}^2$, (d) $d\sigma/d\zeta_{\bar{b}c}^2$ on a semilog scale.

PDFs, the renormalization and factorization scales, etc. On the other hand, compared to the total cross section, the distributions/correlations have additional discriminating power since their shapes also get modified under the influence of NP.

In the following subsections, using the distributions/ correlations, we perform simulations to examine the prospects for detecting NP, for measuring certain combinations of NP parameters, and for partially identifying the NP. The simulations are done using a total number of events consistent with either short-term or long-term measurements at the LHC.

A. Short term

The $t\bar{t}$ cross section at the LHC at a center-of-mass energy of 14 TeV is ~900 pb [9]. Considering the SM branching fractions for $t \rightarrow b\bar{b}c$ and $\bar{t} \rightarrow \bar{b}\ell^-\bar{\nu}_\ell$, the effective cross section in this channel is ~0.1 pb. For the shortterm simulations, we consider an integrated luminosity which, after factoring in the *b*-tagging efficiency,⁴ will lead to 10⁴ events of the type $pp \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell^-\bar{\nu}_\ell)$ from the SM alone.⁵ This is expected to be delivered by 2020–2021 [10].

In the preceding sections, we noted that, in the presence of NP, the shapes of the distributions/correlations can be modified. This suggests that NP can be detected by examining a particular distribution/correlation and seeing a clear difference between the measured shape and its SM prediction. This is explored in Fig. 2 Here all three ζ_{12}^2 distributions are shown for the NP scenario EX-1. Clearly, in the cases of $d\sigma/d\zeta_{bc}^2$ and $d\sigma/d\zeta_{\bar{b}c}^2$, the measurement of the distributions alone would indicate the presence of NP. On the other hand, it would be difficult to draw conclusions from the shape of the corresponding $d\sigma/d\zeta_{b\bar{b}}^2$ distribution.⁶

However, even in the case of $d\sigma/d\zeta_{b\bar{b}}^2$, information about the NP can be obtained. To see this, we use the fitting procedure developed in the previous section and examine what kind of information can be extracted by fitting this

⁴This is assumed to be 70% for each of the three *b* or \bar{b} 's in the final state.

⁵Here $\ell = e, \mu$. In the *CP*-conserving scenario that we consider, there would be an equal number of events in which $\bar{t} \rightarrow \bar{b}b\bar{c}$ and $t \rightarrow b\ell^+\nu_{\ell}$. We assume that the events in which the \bar{t} decays leptonically can be identified by tagging the charge of the lepton and consider only those events in our analysis.

^oIf we were not normalizing the distribution to the total number of events, the difference between the SM and EX-1 cases would be much more apparent.

distribution alone. From Eq. (4) one sees that the $\zeta_{b\bar{b}}^2$ distribution depends on three distinct kinematic structures: $h_{\rm SM}^{b\bar{b}}$, $h_c^{b\bar{b}}$ and $\bar{h}_b^{b\bar{b}}(=h_{\bar{b}}^{b\bar{b}})$. These kinematic structures will in principle be modified by cuts, such as we place on the final state \bar{b} 's. In addition, as described above, there is a mild (but potentially important) dependence on $\operatorname{Re}(X_{LL}^V)$ in the $\zeta_{b\bar{b}}^2$ distribution. We did not include this dependence in the analytical expressions above, but we do retain it here in our numerical work. Thus, a fit using only the $\zeta_{b\bar{b}}^2$ distribution would be sensitive to the relative weights of the SM contribution, $\operatorname{Re}(X_{LL}^V)$, $(\hat{A}_c^+ + \hat{A}_c^-)$ and $(\hat{A}_{\bar{b}}^+ + \hat{A}_{\bar{b}}^- + \hat{A}_{\bar{b}}^+ + \hat{A}_{\bar{b}}^+)$. Accordingly, we modify our fitting procedure: instead of using all eight templates in the fit, we use only four, namely TM-0, TM-1, TM-5 and TM-8. The values that we obtain for the above combinations of NP parameters are presented in Table VII. Once again, $\operatorname{Re}(X_{LL}^V)$ proves to be the weakest link. For the other combinations of NP parameters, the values extracted from the fit agree with their input values within $\pm 3.2\sigma$. However, the key point is this: in each test case, a parameter combination whose input value is nonzero is found from the fit to be nonzero to at least 7σ . So, for the NP scenario EX-1, although one cannot draw any conclusions about NP from a visual examination of the normalized $d\sigma/d\zeta^2_{b\bar{b}}$ distribution, a fit provides statistically significant evidence that NP is present.

Similar fits can be performed with the $d\sigma/d\zeta_{bc}^2$ and $d\sigma/d\zeta_{bc}^2$ distributions to extract other combinations of \hat{A}_i^{σ} 's. If all the $d\sigma/d\zeta_{12}^2$'s are combined in one fit, it is possible

TABLE VII. Values of the combinations of NP parameters extracted from $d\sigma/d\zeta_{b\bar{b}}^2$. The integrated luminosity corresponds to 10^4 SM events.

Test case	Parameter	Input value	Fit result λ	$r^{2}/d.o.f.$
EX-1	SM coefficient	1	1.2 ± 0.2	1.31
	$\hat{A}_{c}^{+} + \hat{A}_{c}^{-}$	64	72 ± 6	
	$\hat{A}^{+}_{\bar{b}} + \hat{A}^{-}_{\bar{b}} + \hat{A}^{+}_{b} + \hat{A}^{-}_{b}$	32	16 ± 5	
	$\operatorname{Re}(X_{LL}^V)$	0	3 ± 1	
EX-2	SM coefficient	1	0.9 ± 0.1	1.24
	$\hat{A}_{c}^{+}+\hat{A}_{c}^{-}$	0	-1 ± 4	
	$\hat{A}^{+}_{\bar{b}} + \hat{A}^{-}_{\bar{b}} + \hat{A}^{+}_{b} + \hat{A}^{-}_{b}$	25	28 ± 3	
	$\operatorname{Re}(X_{LL}^V)$	0	0.0 ± 0.7	
EX-3	SM coefficient	1	1.3 ± 0.1	1.32
	$\hat{A}_{c}^{+}+\hat{A}_{c}^{-}$	0	-10 ± 4	
	$\hat{A}^{+}_{\bar{h}} + \hat{A}^{-}_{\bar{h}} + \hat{A}^{+}_{\bar{h}} + \hat{A}^{-}_{\bar{h}}$	25	22 ± 3	
	$\operatorname{Re}(X_{LL}^V)$	0 -	-1.0 ± 0.7	
EX-4	SM coefficient	1	1.0 ± 0.1	0.96
	$\hat{A}_{c}^{+} + \hat{A}_{c}^{-}$	0	-2 ± 5	
	$\hat{A}^{+}_{\bar{b}} + \hat{A}^{-}_{\bar{b}} + \hat{A}^{+}_{b} + \hat{A}^{-}_{b}$	61	63 ± 4	
	$\operatorname{Re}(X_{LL}^V)$	3	2.3 ± 0.8	

to obtain the combinations $(\hat{A}_{\bar{b}}^+ + \hat{A}_{\bar{b}}^-)$, $(\hat{A}_{\bar{b}}^+ + \hat{A}_{\bar{b}}^-)$ and $(\hat{A}_c^+ + \hat{A}_c^-)$.

Analogous to the above example, one may wish to attempt the fit of a single angular correlation. This, however, is a more complicated task. As argued earlier, the fits are not sensitive to the individual \hat{A}_i^{σ} 's, but rather to the different kinematic structures that are present. In the case of the angular correlations [see Eq. (5)], there are two kinematic structures: a constant term and a term proportional to $\cos \theta_3^* \cos \theta_{\ell}^*$. However, the coefficients of these pieces involve both SM and NP parameters. Therefore only these combinations of SM and NP parameters can be extracted. Furthermore, note that, in the case of the ζ_{12}^2 distributions, each template is proportional to a single kinematic structure. For example, for $d\sigma/d\zeta_{b\bar{b}}^2$, TM-5 is only sensitive to $h_c^{b\bar{b}}$. On the other hand, in the case of the angular correlations, each template contains both the constant piece and the $\cos \theta_3^* \cos \theta_{\ell}^*$ piece.

In order to work around these difficulties, we proceed as follows. First, we fix the weight of the SM contribution⁷ to be 1.0. Second, the templates themselves have to be reorganized. For example, consider $d\sigma/d\cos\theta_c^*d\cos\theta_c^*$. Here a template defined as (TM-5 + TM-6) would be proportional to the constant piece, and another defined as (TM-5 – TM-6) would be proportional to the coefficients of these two modified templates would then be expected to yield the values of $(\hat{A}_c^+ + \hat{A}_c^- + \hat{A}_b^+ + \hat{A}_b^- + \hat{A}_b^+ + \hat{A}_b^-)$ and $(\hat{A}_c^+ - \hat{A}_c^- - \hat{A}_b^+ - \hat{A}_b^-)$.

The results of the fit for the different EX-i are presented in Table VIII. For all four EX-i, the agreement between best-fit and input values is very good for $(\hat{A}_c^+ + \hat{A}_c^- + \hat{A}_b^+ + \hat{A}_b^- + \hat{A}_b^+ + \hat{A}_b^-)$. The key point is that, in all cases, this combination of NP parameters is definitely nonzero. For $(\hat{A}_c^+ - \hat{A}_c^- - \frac{1}{3}\{\hat{A}_b^+ - \hat{A}_b^- + \hat{A}_b^+ - \hat{A}_b^-\})$ the error bars are larger: the best-fit and input values agree to within 1–3 σ . Nevertheless, a fit to a single angular correlation can provide statistically significant evidence that NP is present. The measurement of an angular correlation would, however, most likely be more challenging than the measurement of a ζ_{12}^2 distribution, which is essentially an invariant mass-squared distribution. Hence it is very likely that NP, if present, will be discovered first in a ζ_{12}^2 distribution.

The simplest approach towards the fitting of the angular correlations would have been to fit them to the functional form $a_1 + a_2 \cos \theta_3^* \cos \theta_{\ell}^*$, as is the usual

⁷Note that this could also have been done for the fits to the ζ_{12}^2 distributions. However, in the case of an angular correlation it *must* be done.

⁸This holds as long as the values of \hat{A}_c^+ and \hat{A}_c^- used to generate TM-5 and TM-6, respectively, are identical.

Test Case	Parameter	Input Value	Fit Result	$\chi^2/d.o.f.$
EX-1	$\hat{A}^+_c + \hat{A}^c + \hat{A}^+_{ar{b}} + \hat{A}^{ar{b}} + \hat{A}^+_b + \hat{A}^b$	96	97 ± 1	1.19
	$\hat{A}_{c}^{+} - \hat{A}_{c}^{-} - \frac{1}{3}(\hat{A}_{\bar{b}}^{+} - \hat{A}_{\bar{b}}^{-} + \hat{A}_{b}^{+} - \hat{A}_{\bar{b}}^{-})$	0	26 ± 11	
EX-2	$\hat{A}^+_c + \hat{A}^c + \hat{A}^+_{ar{b}} + \hat{A}^{ar{b}} + \hat{A}^+_b + \hat{A}^b$	25	26 ± 1	1.00
	$\hat{A}_{c}^{+} - \hat{A}_{c}^{-} - \frac{1}{3}(\hat{A}_{\bar{b}}^{+} - \hat{A}_{\bar{b}}^{-} + \hat{A}_{b}^{+} - \hat{A}_{\bar{b}}^{-})$	-8.33	-10 ± 7	
EX-3	$\hat{A}^+_c + \hat{A}^c + \hat{A}^+_{ar{b}} + \hat{A}^{ar{b}} + \hat{A}^+_b + \hat{A}^b$	25	26 ± 1	1.01
	$\hat{A}_{c}^{+} - \hat{A}_{c}^{-} - \frac{1}{3}(\hat{A}_{\bar{b}}^{+} - \hat{A}_{\bar{b}}^{-} + \hat{A}_{b}^{+} - \hat{A}_{\bar{b}}^{-})$	2.33	-3 ± 7	
EX-4	$\hat{A}^+_c + \hat{A}^c + \hat{A}^+_{ar{b}} + \hat{A}^{ar{b}} + \hat{A}^+_b + \hat{A}^b$	61	64 ± 1	0.97
	$\hat{A}_{c}^{+} - \hat{A}_{c}^{-} - \frac{1}{3}(\hat{A}_{\bar{b}}^{+} - \hat{A}_{\bar{b}}^{-} + \hat{A}_{b}^{+} - \hat{A}_{\bar{b}}^{-})$	-20.33	-31 ± 9	

TABLE VIII. Values of the combinations of NP parameters extracted from $d\sigma/d\cos\theta_c^*d\cos\theta_c^*$. The integrated luminosity corresponds to 10^4 SM events. The weight of the SM contribution is fixed to be 1.0.

procedure for measuring $\kappa_{t\bar{t}}$. However, this possibility is precluded due to the fact that the event-selection criteria described in Fit 2 distorts the shape of the correlation. We have used a somewhat simple-minded approach to deal with the identical \bar{b} 's in the final state. It is certainly possible that experimentalists will find a better way to deal with this situation (perhaps through the use of some sophisticated multivariate technique, such as neural networks or boosted decision trees) and that such an approach would lead to less distortion of the shape of the correlation.

Finally, we consider the full fit involving all six observables, with statistics corresponding to 10^4 events

for the SM. The results are presented in Table IX. Apart from $\text{Re}(X_{LL}^V)$ in EX-1, the values of all NP parameters agree with their input values within $\pm 1.7\sigma$.

To sum up, the above simulations demonstrate that, even in the short term, it is possible to detect the presence of NP in top decay through the measurement of the invariant mass-squared distributions and/or the angular correlations. This can be done by comparing the measured shapes of the distributions/correlations with the SM predictions. More sensitivity can be obtained by performing fits to extract combinations of NP parameters. If all six distributions and correlations can be measured, a combined fit can be performed to extract all the NP parameters. The

Test Case	Fit Results		$\chi^2/d.o.f.$
EX-1	SM coefficient = 1.00 ± 0.01 $\hat{A}_{\bar{b}}^{+} = 36 \pm 9$ $\hat{A}_{\bar{b}}^{+} = -22 \pm 8$ $\hat{A}_{c}^{+} = 47 \pm 9$ Re(X_{c}^{V}) = 0.31 ± 0.09	$egin{array}{lll} \hat{A}_{ar{b}}^{-} &= 28 \pm 9 \ \hat{A}_{ar{b}}^{-} &= -11 \pm 8 \ \hat{A}_{c}^{-} &= 18 \pm 9 \end{array}$	1.18
EX-2	SM coefficient = 0.988 ± 0.009 $\hat{A}_{b}^{+} = 6 \pm 6$ $\hat{A}_{b}^{+} = 23 \pm 5$ $\hat{A}_{c}^{+} = 0 \pm 6$ Re(X_{L}^{V}) = 0.03 ± 0.05	$\hat{A}_{ar{b}}^- = -3 \pm 6 \ \hat{A}_{ar{b}}^- = 1 \pm 6 \ \hat{A}_{c}^- = -1 \pm 6$	0.92
EX-3	SM coefficient = 1.013 ± 0.009 $\hat{A}_{b}^{+} = 4 \pm 6$ $\hat{A}_{b}^{+} = 16 \pm 5$ $\hat{A}_{c}^{+} = 1 \pm 6$ Re $(X_{LL}^{V}) = 0.03 \pm 0.05$	$\hat{A}_{\overline{b}} = -3 \pm 6$ $\hat{A}_{\overline{b}} = 8 \pm 6$ $\hat{A}_{\overline{c}} = -2 \pm 6$	0.94
EX-4	SM coefficient = 1.00 ± 0.01 $\hat{A}_{\bar{b}}^{+} = 47 \pm 7$ $\hat{A}_{\bar{b}}^{+} = 34 \pm 7$ $\hat{A}_{c}^{+} = 8 \pm 8$ $\operatorname{Re}(X_{LL}^{V}) = 2.91 \pm 0.05$	$\hat{A}_{\bar{b}}^- = -12 \pm 7$ $\hat{A}_{\bar{b}}^- = -9 \pm 7$ $\hat{A}_{\bar{c}}^- = -8 \pm 8$	0.99

TABLE IX. Values of the NP parameters extracted from a fit to all six observables. The integrated luminosity corresponds to 10^4 SM events.

determination of which parameters are nonzero allows for a partial identification of the NP.

B. Long term

As noted above, the effective cross section in $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$ is ~0.1 pb. The LHC is projected to deliver 3000 fb⁻¹ worth of data by the year 2030 [10]. Assuming this integrated luminosity and a *b*-tagging efficiency of 70% for each of the three *b* or \bar{b} 's in the final state, one obtains $\approx 10^5$ events of the type $pp \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell^-\bar{\nu}_\ell)$ from the SM. This is the number of events in our long-term simulations.

By 2030, all six distributions and correlations will, in all likelihood, have been measured. For this reason we consider only the fit to all distributions/correlations with 10^5 SM events. The corresponding results have already been presented in Table VI. Apart from $\text{Re}(X_{LL}^V)$ in EX-1, the best-fit values of all NP parameters differ from their input values by at most 1σ . The errors on the \hat{A}_i^{σ} 's are typically in the range 1.75–2.75. Thus, any \hat{A}_i^{σ} that is $\gtrsim 10$ will be found to be nonzero at a statistically significant level. In this way it will be possible to determine which NP parameters are nonzero, thus producing an identification of the NP.

V. CONCLUSIONS

In Ref. [1], the companion paper, NP in the decay $t \rightarrow b\bar{b}c$ is considered. There, ten dimension-6 NP operators contributing to $t \rightarrow b\bar{b}c$ are delineated, and two types of observables are identified that can be used to search for this NP in the process $gg \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell\bar{\nu})$. They are (i) invariant mass-squared distributions involving the $\{b, c\}, \{\bar{b}, c\}, \text{ or } \{b, \bar{b}\}$ quark pairs coming from $t \rightarrow b\bar{b}c$, and (ii) angular correlations between the $\ell^$ coming from the \bar{t} decay and one of \bar{b} , b or c coming from the t decay. It is further shown that the NP contributions to these observables can be written in terms of certain combinations of the NP couplings denoted as \hat{A}_i^{σ} . In the present paper we examine the prospects for detecting and identifying such NP at the LHC.

The first step is to develop an algorithm to extract the \hat{A}_i^{σ} 's and $\operatorname{Re}(X_{LL}^V)$ from the observables. From the analytical expressions obtained in Ref. [1] [summarized here in Eqs. (4) and (5)], we learn that the NP contribution to the observables can be represented as a linear combination of pieces proportional to the different \hat{A}_i^{σ} 's and $\operatorname{Re}(X_{LL}^V)$. Using this idea, we perform a Monte Carlo simulation using MADGRAPH 5 to compute "templates," which are the contributions of the SM, each \hat{A}_i^{σ} and $\operatorname{Re}(X_{LL}^V)$ to the observables. We also generate Monte Carlo data for four possible NP scenarios. For each of these scenarios, we extract the NP parameters simply by obtaining the weights with which the templates must be combined to reproduce the Monte Carlo data.

Although the fit algorithm is based on a simple premise, there are two issues that must be taken into account. First, the construction of the observables requires distinguishing the decay products of the t from those of the \overline{t} . However, the final state contains two \bar{b} 's, which are indistinguishable, at least in some parts of phase space. We designate the \bar{b} that yields the smaller value of $|m_t - \sqrt{(p_{\bar{b}} + p_b + p_c)^2}|$ as that having come from the t decay. However, if both \bar{b} 's in the event yield sufficiently small values of this quantity (less than $15\Gamma_t$), then we consider them to be indistinguishable and exclude such events from the analysis. Second, the contribution to $t\bar{t}$ production from a $q\bar{q}$ initial state is not included in the analytical expressions. However, this must be taken into account as there are no known algorithms that can efficiently separate $t\bar{t}$ pairs coming from gluon fusion from those occurring due to $q\bar{q}$ annihilation.

In order to examine the prospects for detecting the presence of NP in $t \rightarrow b\bar{b}c$, and for its identification, we perform further simulations of the distributions/ correlations. The simulations are done for either short-term or long-term measurements at the LHC. For the short-term analysis we use 10^4 events of the type $pp \rightarrow t\bar{t} \rightarrow (b\bar{b}c)(\bar{b}\ell^-\bar{\nu}_\ell)$. This is expected to be delivered by 2020–2021. For the long term we use 10^5 events, which is projected by the year 2030.

In the short term not all distributions/correlations may be measured, and what can be learned about the NP depends on what measurements have been made. In the presence of a sufficiently large NP contribution to $t \rightarrow b\bar{b}c$, the shapes of the distributions/correlations can be significantly modified. Thus, NP in $t \rightarrow b\bar{b}c$ may be inferred by observing a clear difference between the shape of a measured distribution and its SM prediction. Even if there is no discernible difference in the shapes, it may still be possible to obtain information about NP contributions. Using the above algorithm with a slight modification, one can perform a fit to a single distribution. In this case, not all the individual \hat{A}_{i}^{σ} 's are extracted, but rather certain combinations of the \hat{A}_i^{σ} 's. We show that, even for a scenario in which the presence of NP does not induce a substantial change in the shape of the distribution, a fit may still yield statistically significant evidence that NP is present. Finally, if all six distributions/correlations are measured, we can use the algorithm to perform a simultaneous fit on all the observables to extract $\operatorname{Re}(X_{LL}^V)$ and all the \hat{A}_i^{σ} 's separately. In the examples studied, we find that the values of all NP parameters agree with their input values within $\pm 1.7\sigma$. Although the errors are large, this provides an approximate determination of the values of the NP parameters. More importantly it allows us to infer that a nonzero NP contribution to $t \rightarrow b\bar{b}c$ exists.

In the long term, it is likely that all six distributions/ correlations will be measured. Furthermore, the availability of larger statistics will lead to an improvement in the quality of the fits. We find that, with 10^5 events, the best-fit values of all NP parameters differ from their input values by at most 1σ . Thus, if NP is present in $t \rightarrow b\bar{b}c$, the fit will allow the determination of its nature.

ACKNOWLEDGMENTS

The authors wish to thank the MADGRAPH and FEYNRULES teams for extensive discussions about MADGRAPH and FEYNRULES, respectively. P. S. would like to thank Georges Azuelos for helpful discussions about

MADGRAPH. The authors also wish to thank S. Judge and J. Melendez for collaboration at an early stage of this work and J. Kiers and R. Rezvani for helpful discussions. This work was financially supported by NSERC of Canada (D. L., P. S.). This work has been partially supported by ANPCyT under Grant No. PICT-PRH 2009-0054 and by CONICET (A. S.). The work of K. K. was supported by the U.S. National Science Foundation under Grant No. PHY–1215785. K. K. also acknowledges sabbatical support from Taylor University.

- K. Kiers, P. Saha, A. Szynkman, D. London, S. Judge, and J. Melendez, Phys. Rev. D 90, 094015 (2014).
- [2] V. M. Abazov *et al.* (D0 Collaboration), Phys. Rev. D 85, 091104 (2012); T. A. Aaltonen *et al.* (CDF Collaboration), Phys. Rev. Lett. 111, 202001 (2013).
- [3] J. Alwall, M. Herquet, F. Maltoni, O. Mattelaer, and T. Stelzer, J. High Energy Phys. 06 (2011) 128; http:// madgraph.hep.uiuc.edu/.
- [4] K. Kiers, T. Knighton, D. London, M. Russell, A. Szynkman, and K. Webster, Phys. Rev. D 84, 074018 (2011).
- [5] A. Datta and D. London, Phys. Lett. B **595**, 453 (2004).

- [6] E. Byckling and K. Kajantie, *Particle Kinematics* (Wiley, New York, 1973).
- [7] N. D. Christensen and C. Duhr, Comput. Phys. Commun. 180, 1614 (2009); A. Alloul, N. D. Christensen, C. Degrande, C. Duhr, and B. Fuks, Comput. Phys. Commun. 185, 2250 (2014); http://feynrules.irmp.ucl.ac.be/.
- [8] W. Press et al., Numerical Recipes in Fortran 77: The Art of Scientific Computing (Cambridge University Press, Cambridge, England, 1992), ISBN-13:.
- [9] N. Kidonakis, arXiv:1304.7775; S. Moch and P. Uwer, Phys. Rev. D 78, 034003 (2008).
- [10] S. Myers, https://indico.cern.ch/event/73513/session/13/ contribution/73.