metric equations.<sup>3</sup>

is noncausal.

## Demonstration of noncausality for the Rarita-Schwinger equation\*

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Using the Rarita-Schwinger equation for a spin-3/2 particle in a constant magnetic field, we explicitly calculate the propagator in 2 + 1 dimensions. From the behavior of the propagator, it is seen that the propagation is noncausal.

## I. INTRODUCTION

in Ref. 1 that the solutions of the Rarita-Schwinger (RS) equation are noncausal, using the method of characteristics. There are two main criticisms of the Velo and Zwanziger result. First, they are applying the method of characteristics outside the domain of the existing proofs.<sup>2</sup> Secondly, since

the metric used in the RS equation is not positive

definite, even the existence of the solutions is not certain. The positivity of the metric is used in an essential manner for the existence proofs of sym-

There has been further work on this subject, and the existence of solutions for the RS equation has been shown quite recently.<sup>4</sup> However, we think it is still worthwhile to communicate this work, which was done some time ago, which might be complementary to Ref. 1 in some respect. Here,

without treating in pure mathematical rigor the

problems of domains of definition, regularity of the solutions, etc., we explicitly construct the

propagator for the RS equation in 2+1 dimensions. We use Velo and Zwanziger's techniques<sup>1</sup> to re-

duce the RS equation into the modified form where

we have a genuine equation of motion, which is

equivalent to the original equation if we impose certain conditions at time  $t_0$ . For any electromag-

netic potential, we do not know any solutions of

this equation in 3+1 dimensions. In 2+1 dimen-

sions, however, this equation can be easily solved

for a constant magnetic field. From the solutions,

we construct the propagator and explicitly see that

the propagation of the RS field in 2+1 dimensions

Recently, there has been some interest on higher-spin equations. Velo and Zwanziger showed

Using the subsidiary conditions

$$\gamma^{\mu}\psi_{\mu} = \frac{ie}{2m} \left( -F_{\nu\mu}\gamma^{\mu}\psi^{\nu} + \frac{1}{2}F_{\nu\lambda}\gamma^{\nu}\gamma^{\lambda}\gamma_{\mu}\psi^{\mu} \right)$$
(2)

one gets the modified RS equation

$$(\not\!\!D + m)\psi^{\kappa} - (\gamma^{\kappa}m + D^{\kappa})ge(F_{\nu\mu}\gamma^{\mu}\psi^{\nu} - \frac{1}{2}F_{\nu\lambda}\gamma^{\nu}\gamma^{\lambda}\gamma_{\mu}\psi^{\mu}) = 0$$

$$g = -\frac{i}{2m^2} \quad . \tag{3}$$

Equation (1) has two subsidiary conditions: Eq. (2) and the zeroth component of  $L^{\nu}$ . Take

$$\chi \equiv L^{0} = (\gamma^{l} D_{l} - m) \gamma_{k} \psi^{k} + D_{l} \psi^{l} = 0 ,$$
  
$$l, k = 1, 2 \text{ for } t = t_{0} \quad (4)$$

$$\phi = \gamma_{\mu}\psi^{\mu} - ge(F_{\nu\mu}\gamma^{\mu}\psi^{\nu} - \frac{1}{2}F_{\nu\lambda}\gamma^{\nu}\gamma^{\lambda}\gamma_{\mu}\psi^{\mu}) = 0 ,$$
  
for  $t = t_0$ . (5)

By straightforward computation<sup>5</sup> one can show that  $\chi$  and  $\phi$  satisfy the equations of motion

$$(-\not D + m)\phi + 2(D_k\gamma^k - m)\phi - 2\chi = 0 , \qquad (6)$$

$$(\gamma_0 D_0 - \gamma^l D_l - m)\chi = \frac{1}{2}eF_{lk}\gamma^l\gamma^k\phi , \qquad (7)$$

which gives the wave equation for  $\phi$ :

$$(D_{\mu}D^{\mu} + m^{2})\phi + eF_{\mu\nu}\gamma^{\mu}\gamma^{\nu}\phi = 0.$$
 (8)

So, if we impose  $\phi = 0$ ,  $\chi = 0$ ,  $\partial_k \phi = 0$  at  $t = t_0$ , Eqs. (6) and (7) give  $\chi = 0 = \phi$  for all time. This shows that Eqs. (2)-(4) are equivalent to Eq. (1), i.e., if we impose the subsidiary conditions  $L^0 = 0$  and Eq. (2) on the solutions of Eq. (3) at  $t = t_0$ , we get the solutions of Eq. (1). One can also get to Eq. (3) using the methods described in Ref. 6.

Now we study Eq. (3) with a constant magnetic field,

$$A_1 = Bx_2, \quad A_2 = A_3 = 0$$
.

Then Eq. (3) reads

$$(\not D + m)\psi^{\nu} - (\gamma^{\nu} m + D^{\nu})egB\gamma^{1}\gamma^{2}\gamma_{0}\psi^{0} = 0 .$$
 (9)

For  $\nu = 0$ ,  $\psi^0$  totally decouples from  $\psi_1$  and  $\psi_2$ .

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In 2+1 dimensions, the Rarita-Schwinger equa-

**II. THE CALCULATION AND DISCUSSION** 

$$\begin{split} L^{\nu} &= (\not\!\!\!D + m)\psi^{\nu} - (\gamma^{\nu}D_{\mu} + D^{\nu}\gamma_{\mu})\psi^{\mu} - \gamma^{\nu}(\not\!\!\!D - m)\gamma_{\mu}\psi^{\nu} \\ &= 0 , \qquad \qquad \nu = 0, 1, 2 , \quad (1) \\ D^{\mu} &= \partial^{\mu} - ieA^{\mu} . \end{split}$$

$$(\not\!\!\!\!/ + m)\psi^0 - (\gamma^0 m + \partial^0) eg B \gamma^1 \gamma^2 \gamma^0 \psi^0 - i e \gamma_1 A^1 \psi_0 = 0 .$$
(10)

One can solve the  $\nu = 0$  equation for  $\psi^0$ , then use  $\psi^0$  as a source term in the equations for  $\psi^1$  and  $\psi^2$ . This simplification of the problem is our main reason to go to 2+1 dimensions.

We have three wave equations for  $\psi^0$ ,  $\psi^1$ , and  $\psi^2$ , and two subsidiary conditions, Eqs. (2) and (4). We can solve Eq. (2) for  $\psi^2$  in terms of  $\psi^1$  and  $\psi^0$ . Then Eq. (4) gives a differential equation for  $\psi^1$ , where  $\psi^0$  is treated as a source term, and where only space derivatives appear. We do not have any restrictions on  $\psi^0$  and  $\partial_0 \psi^0$  at  $t = t_0$ ; fixing them determines  $\psi^1$  and  $\psi^2$ . So, we have a well-defined Cauchy problem for  $\psi^0$ .

In 2+1 dimensions  $\gamma^{\mu}$  has the representation of the Pauli  $\tau$  matrices

$$\gamma_0 = \begin{pmatrix} i & 0 \\ 0 & -i \end{pmatrix}, \quad \gamma_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \gamma_2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}.$$

After eliminating the first-order time derivative by taking  $\psi_0 = e^{i\alpha t}\phi$ , where

$$\alpha = \frac{eB}{m(1 - e^2 B^2 / 4m^4)} , \qquad (11)$$

Eq. (10) goes into

$$\left\{-1\left(1-\frac{e^2B^2}{4m^4}\right)(\partial^0)^2+(\partial_1-iex_2B)^2+\partial_2^2+\left[i\sigma_0eB-\frac{e^2B^2}{m(1-e^2B^2/4m^4)}-m^2\left(1-\frac{e^2B^2}{4m^4}\right)\right]\right\}\phi=0,$$
(12)

where

$$-i\sigma_0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} .$$

Note that if  $e^2B^2 > 4m^4$ , the equation above is not hyperbolic. A similar situation occurs in 3+1 dimensions as shown in Ref. 1.

For the sake of simplicity we change to variables

$$x = (eB)^{1/2} x_1, \quad y = (eB)^{1/2} x_2, \quad t = (eB)^{1/2} x_0.$$

Equation (12) has a stationary solution,

$$\phi_n = \frac{e^{-iB_n t}}{2\pi (n!)^{1/2}} \frac{H_n((y+p_1)/\sqrt{2})}{\pi^{1/4}} e^{-(y+p_1)^2/2} e^{-ip_1 x} ,$$
(13)

where

$$E_{n} = \left(\frac{2n+1+M^{2}}{1-e^{2}B^{2}/4m^{4}}\right)^{1/2},$$

$$M^{2} = \frac{eB}{m(1-e^{2}B^{2}/4m^{4})} + \frac{m^{2}}{eB}\left(1-\frac{e^{2}B^{2}}{4m^{4}}\right) - i\sigma_{0}.$$
(14)

 $H_n$  are the Hermite polynomials of degree n.

Here we want to study the propagation properties of

$$\psi^{0}(x) = \int D(x, x') \overline{\partial}_{0} \psi_{0}(x') d^{3}x' \quad . \tag{15}$$

Imposing Eqs. (2) and (4) at time  $t = t_0$ , we get a solution for Eq. (1). The propagator D(x, x') propagates that solution to time t. Therefore, it is sufficient to establish the noncausality for D(x, x') to show the noncausality for the solutions of the RS equation in 2+1 dimensions.

Using Ref. 7, we write down the homogeneous propagator D for  $\psi^0$ ,

$$D = \sum_{n} \int dp_{1} e^{ip_{1}(x-x')} e^{-[(y+p_{1})^{2}+(y'+p_{1})^{2}]/2}$$

$$\times e^{i\alpha(t-t')} \frac{\sin E_{n}(t-t')}{E_{n}}$$

$$\times \frac{H_{n}((y+p_{1})/\sqrt{2})H_{n}((y'+p_{1})/\sqrt{2})}{\sqrt{\pi} n!} .$$
(16)

D satisfies

LD=0,

where L is the differential operator of Eq. (10),

$$\frac{\partial D}{\partial x_0}\Big|_{x_0=0} = \delta(\bar{\mathbf{x}} - \bar{\mathbf{x}}'),$$

$$D = D^{adv} - D^{ret}.$$
(18)

Using the completeness property of  $e^{-ip_1x}$  and  $H_n((p_1+y)/\sqrt{2})$ , we perform the integration<sup>8</sup> and get

$$\frac{1}{\sqrt{\pi}} \frac{1}{n!} \int_{-\infty}^{\infty} dp_1 e^{ip_1(x-x')} e^{-[(y+p_1)^2 + (y'+p_1)^2]/2} H_n\left(\frac{y+p_1}{\sqrt{2}}\right) H_n\left(\frac{y'+p_1}{\sqrt{2}}\right) = e^{-(x-x')(y+y')/2} L_n(\sigma') e^{-\sigma'/2},$$
(19)

where  $L_n$  is the Laguerre function and

$$\sigma' = \frac{1}{2} [(x - x')^2 + (y - y')^2] .$$

Using the integral representation for Bessel functions.<sup>9</sup>

$$\frac{\sin E_n(t-t')}{E_n} = (\frac{1}{2}\pi)^{1/2} \frac{1}{2\pi i} (t-t') \int_{c-i\infty}^{c+i\infty} d\beta \frac{\exp\{\frac{1}{2}[\beta - E_n(t-t')/\beta]^2\}}{\beta^{3/2}} , \ c > 0$$
(21)

(20)

and the generating function for the Laguerre polynomials,

$$\sum_{n} u^{n} L_{n}(\sigma') e^{-\sigma'/2} = \frac{\exp[\frac{1}{2}\sigma'(1+u)/(1-u)]}{1-u} , \ u < 1$$
(22)

after changing variables,  $\phi = \beta/(t - t')^2$ , and restoring to the original variables  $x_0, x_1, x_2, x_1^{0}$  we end up with

$$D = \frac{1}{2\pi i} (\frac{1}{2}\pi)^{1/2} \exp\left[\frac{ieB(x_0 - x_0')}{m(1 - e^2B^2/4m^4)}\right] \exp\left[\frac{1}{2}ieB(x_2 + x_2')(x_1 - x_1')\right] \\ \times \int_{c - i\infty}^{c + i\infty} \frac{d\phi}{2\phi^{3/2}} \exp\left(\frac{1}{2}\left\{-\lambda\phi - \frac{1}{\phi}\left[\frac{e^2B^2}{m^2(1 - e^2B^2/4m^4)^2} + m'^2\right]\right\}\right) \frac{\exp\left\{\frac{1}{2}\sigma eB\left[\frac{2\phi}{eB} - \frac{2eB\phi}{4m^4} - \coth\frac{eB}{2\phi(1 - e^2B^2/4m^4)}\right\}}{\sinh\left[\frac{eB}{2\phi(1 - e^2B^2/4m^4)}\right]}\right\},$$
(23)

where

$$-\lambda = 2\sigma \left( 1 - \frac{e^2 B^2}{4m^4} \right) + (x_0 - x_0')^2 , \qquad (24)$$

$$\sigma = \frac{1}{2} [(x_1 - x_1')^2 + (x_2 - x_2')^2] , \qquad (25)$$

$$m'^2 = m^2 - \frac{i\sigma_0 eB}{(1 - e^2 B^2 / 4m^4)}$$
, (26)

c > 0.

For  $\lambda > 0$ , we can close the contour on the righthand side without encountering any singularities, since all the singularities lie on the imaginary axis. Therefore, for  $\lambda > 0$ , D=0. This corresponds to the region where

$$(x_0 - x_0')^2 - \left(1 - \frac{e^2 B^2}{4m^4}\right) \left[(x_1 - x_1')^2 + (x_2 - x_2')^2\right] < 0 .$$

If B=0, we retain the causal behavior, since the propagator is zero for spacelike separations.

When  $B \neq 0$ ,  $\lambda$  does not equal the spacelike distance. We see that the light-cone singularity occurs at  $\lambda = 0$ , which is shifted to the spacelike region.<sup>11</sup> This establishes the noncausal behavior of the  $\psi^0$  component of the RS field in 2 + 1 dimensions. Once we find a solution which is nonzero in a spacelike region; as with the singular solution we found for  $\lambda = 0$ , we have established the noncasuality of the propagator. From the other components of Eq. (9), one easily sees that the other two components are taken together, the RS equation exhibits noncausality in 2 + 1 dimensions.

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- <sup>10</sup>Reference 5. Similar manipulations are described in Refs. 7 and 8.
- $^{11}{\rm For}\ \lambda=0$ , the propagator is reduced to a finite part coming from integrating in a finite region, plus terms like

$$\int_{c}^{\infty} \frac{d\phi}{\phi^{1/2}} + \int_{c}^{\infty} d\phi \ O(\phi^{-3/2}), \quad c \gg 1$$

which diverges.

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<sup>\*</sup>This is part of the author's University of Pittsburgh dissertation.

<sup>&</sup>lt;sup>4</sup>R. Seiler, private communication.

<sup>&</sup>lt;sup>5</sup>M. Hortaçsu, in Ref. 2.