## PHYSICAL REVIEW D

## **Comments and Addenda**

The Comments and Addenda section is for short communications which are not of such urgency as to justify publication in Physical Review Letters and are not appropriate for regular Articles. It includes only the following types of communications: (1) comments on papers previously published in The Physical Review or Physical Review Letters; (2) addenda to papers previously published in The Physical Review or Physical Review Letters, in which the additional information can be presented without the need for writing a complete article. Manuscripts intended for this section should be accompanied by a brief abstract for information-retrieval purposes. Accepted manuscripts will follow the same publication schedule as articles in this journal, and galleys will be sent to authors.

## Bounds on chiral-symmetry breaking and $K_{13}$ slope parameters\*

S. C. Prasad

Department of Physics and Astronomy, University of Rochester, Rochester, New York 14627 (Received 17 December 1973)

Using the rigorous bounds of Li and Pagels and Okubo and the broken  $SU(3) \times SU(3)$  model of Glashow and Weinberg and Gell-Mann, Oakes, and Renner, we derive bounds on the symmetry-breaking parameters which measure the invariance of the Hamiltonian and the vacuum under SU(3). We also obtain bounds on  $\Lambda_0$ , the slope of the spin-0 form factor in  $K_{13}$  decay. Our bounds on  $\Lambda_0$ , though inconsistent with the experimental value obtained by a quadratic fit to the  $K_{13}$  form factors, are consistent with the value obtained by a linear fit. Some theoretical implications of our results are discussed.

Li and Pagels<sup>1</sup> and Okubo<sup>2</sup> have derived some rigorous bounds on the  $K_{13}$  decay parameters:

$$f_{+}^{2}(0) \leq \frac{\Delta(0)}{N} , \qquad (1)$$

$$\left| 4 \left( \frac{m_{\underline{K}} + m_{\pi}}{m_{\pi}} \right)^{2} \Lambda_{0} - 2 + \frac{(4m_{\underline{K}}m_{\pi})^{1/2}}{\left[ (m_{\underline{K}})^{1/2} + (m_{\pi})^{1/2} \right]^{2}} \right|$$

$$\leq \left[ \frac{\Delta(0)}{Nf_{+}^{2}(0)} - 1 \right]^{1/2} , \quad (2)$$

where

$$N = \frac{3}{256\pi} (m_{K} - m_{\pi})^{2} (m_{K} + m_{\pi}) [(m_{K})^{1/2} + (m_{\pi})^{1/2}]^{2} \quad .$$
(3)

Here  $f_+(0)$  is the usual spin-1 form factor,  $\Delta(0)$  is the propagator function for the divergence of the strangeness-changing vector current, and  $\Lambda_0$  is the slope of the spin-0 form factor for the divergence of the current, all quantities evaluated at the zero momentum transfer. Apart from the standard assumptions of analyticity, unitarity, and crossing symmetry, the most crucial assumption under which these bounds are derived is that the spin-0 form factor satisfies an unsubtracted dispersion relation.<sup>1,2</sup> The parameter  $\Lambda_0$  has been measured experimentally, and for a quadratic fit to the form factors the world average value<sup>3</sup> is

$$\Lambda_0 = -0.11 \pm 0.03 \ . \tag{4}$$

On the other hand, a previously reported value<sup>4</sup> obtained by assuming that the form factors depend linearly on momentum transfer is given by

$$\Lambda_0 = -0.024 \pm 0.02 \ . \tag{5}$$

In order to compute the bounds, one needs to know  $\Delta(0)$  or an upper bound to  $\Delta(0)$  in (1) and (2). An estimate has been made for  $\Delta(0)$  using the broken- $SU(3) \times SU(3)$  model of Glashow and Weinberg<sup>5</sup> and Gell-Mann, Oakes, and Renner<sup>5</sup> (GWGMOR) where the symmetry-breaking terms transform as the  $(3,3^*) + (3^*,3)$  representation. One gets  $\Delta(0) \simeq 2.04 m_{\pi}^2 F_{\pi}^2$ , where  $F_{\pi}$  is the pion decay constant,  $F \simeq 94$  MeV. From (1) and (2), one then obtains  $f_+(0) \leq 1.01$  and  $0.011 \leq \Lambda_0 \leq 0.024$ [for  $f_+(0) = 0.85$ ], the bounds for  $\Lambda_0$  being in contradiction with the experimental results (5) and (4). Although the experimental situation at the present time is quite uncertain,<sup>6</sup> it is desirable from the theoretical point of view to reexamine the evaluation of  $\Delta(0)$ . Among various assumptions, the estimate of  $\Delta(0)$  mentioned above is based on the pole-dominance approximation. A priori, it is hard to see how a refinemant in the value of  $\Delta(0)$ , which we do not expect to be drastic, can lead to bounds for  $\Lambda_0$  consistent with the experimental result (4). This is confirmed by the work of  $Okubo^7$ 

2167

9

and others' who have made estimates of  $\Delta(0)$  under different approximations. Thus, if relatively large negative values like (4) are confirmed by further experiments, one may have to give up one or more fundamental premises underlying the derivation of (1) and (2) or conclude that the GWGMOR model is incorrect.

The purpose of the present note is twofold. First, using (1) in conjunction with the GWGMOR model we derive rigorous bounds on the symmetrybreaking parameters. It turns out that these bounds allow both the GMOR-type solution<sup>5</sup> as well as the type advocated by Brandt and Preparata (BP).<sup>8</sup> We find, however, that in the case of the BP-type solution where the Hamiltonian is approximately SU(3)-invariant the vacuum cannot be approximately SU(3)-invariant. Secondly, on the basis of the GWGMOR model we obtain an upper bound for  $\Delta(0)$  by taking into consideration the departures from the pole-dominance approximation, to see if small negative values such as (5) are consistent with the theoretical bounds in (1) and (2).

We start with the Hamiltonian density for the GWGMOR model

$$H(x) = H_0(x) + \epsilon_0 u^0(x) + \epsilon_8 u^8(x) , \qquad (6)$$

where  $H_0(x)$  is invariant under SU(3)×SU(3) and  $u^i(x)$  (i=0, 1, ..., 8) are the scalar densities which belong to the  $(3, 3^*) + (3^*, 3)$  representation of SU(3)×SU(3). We define the following two-point functions:

$$I_{\alpha\beta} = i \int d^4 x \langle 0 | T(\partial_{\mu} A^{\alpha}_{\mu}(x) \partial_{\nu} A^{\beta}_{\nu}(0)) | 0 \rangle ,$$
  

$$K_{\alpha\beta} = i \int d^4 x \langle 0 | T(\partial_{\mu} V^{\alpha}_{\mu}(x) \partial_{\nu} V^{\beta}_{\nu}(0)) | 0 \rangle ,$$
(7)

where  $A^{\alpha}_{\mu}(x)$  and  $V^{\alpha}_{\mu}(x)$  ( $\alpha = 1, 2, ..., 8$ ) are the usual octet of axial-vector and vector currents. As noted elsewhere,<sup>9</sup> from Eqs. (6) and (7) the following exact relations can be derived:

$$I_{33} = \gamma (1 + a)(1 + b) ,$$

$$I_{44} = \gamma (1 - \frac{1}{2}a)(1 - \frac{1}{2}b) ,$$

$$I_{88} = \gamma (1 - a - b + 3ab) ,$$

$$K_{44} \equiv \Delta (0) = \frac{9}{4} \gamma ab ,$$
(8)

where

$$a = \frac{\epsilon_8}{\sqrt{2}\epsilon_0}, \quad b = \frac{\xi_8}{\sqrt{2}\xi_0}, \quad \gamma = -\frac{2}{3}\epsilon_0\xi_0, \quad (9)$$

with  $\xi_0 = \langle 0 | u^0(0) | 0 \rangle$  and  $\xi_8 = \langle 0 | u^8(0) | 0 \rangle$ . We also note that the GMOR solution for the symmetry-breaking parameter *a* is given by

$$a = \frac{2(m_{\pi}^2 - m_K^2)}{m_{\pi}^2 + 2m_K^2} \simeq -0.89 .$$
 (10)

Separating out the single pion and kaon contributions to  $I_{33}$  and  $I_{44}$ , respectively, we write

$$I_{33} = F_{\pi}^{2} m_{\pi}^{2} (1 + \delta_{\pi}) , \qquad (11)$$

$$I_{44} = F_K^2 m_K^2 (1 + \delta_K) , \qquad (12)$$

where  $\delta_{\pi}$  and  $\delta_{K}$  are related to the multiparticle intermediate-state contributions, and  $F_K$  is the kaon decay constant. Since the contribution of any intermediate state to  $I_{33}$  and  $I_{44}$  is positive-definite, we conclude that  $\delta_{\pi} \ge 0$ ,  $\delta_{K} \ge 0$ . Now from wellknown arguments suggesting dominance by nearby singularities, and the success of the hypothesis of partially conserved axial-vector currents (PCAC), we believe, especially for  $I_{33}$ , that pole dominance is perhaps a reasonable assumption. To be generous, however, we shall adopt in the following the weaker assumption that the continuum contribution does not overwhelm the contribution from the single-particle states. In particular, we assume that  $\delta_{\pi} \leq 1$ ,  $\delta_{K} \leq 1$ . We emphasize that this is a rather weak assumption, and in practice one does not expect  $\delta_{\pi}$  and  $\delta_{\kappa}$  to be much more than 10-20%. With considerable reasonableness, we then expect

$$1 \ge \delta_{\pi} \ge 0$$
,  $1 \ge \delta_{K} \ge 0$ . (13)

We now proceed to derive bounds on the symmetry-breaking parameters a and b. Eliminating  $\gamma$  and b from the relations for  $I_{33}$ ,  $I_{44}$ , and  $\Delta(0)$ , we get

$$\Delta(0) = \frac{3a}{2} \left( \frac{I_{33}}{1+a} - \frac{2I_{44}}{2-a} \right) \quad . \tag{14}$$

Now for the physical solution we expect  $-1 \le a \le 0$ . Equations (14) and (1) then give the following relation:

$$a^{2}[3(2I_{44} + I_{33}) - 2Nf_{+}^{2}(0)] + 2a[3(I_{44} - I_{33}) + Nf_{+}^{2}(0)] + 4Nf_{+}^{2}(0) \le 0.$$
(15)

Using Eqs. (11) and (12) we obtain from (15) the following exact bounds for a:

$$-\frac{D+B}{2A} \le a \le \frac{D-B}{2A} \quad , \tag{16}$$

where

$$A = \frac{3}{2} [2m^2 \alpha^2 f_+^2(0)(1+\delta_K) + 1 + \delta_\pi] - Rf_+^2(0) ,$$
  
$$B = 3 [m^2 \alpha^2 f_+^2(0)(1+\delta_K) - 1 - \delta_\pi] + Rf_+^2(0) ,$$

and

$$D = 3\{ [m^2 \alpha^2 f_+^2(0)(1+\delta_K) + 1 + \delta_\pi - R f_+^2(0)]^2 -4(1+\delta_\pi)(1+\delta_K)m^2 \alpha^2 f_+^2(0) \}^{1/2} ,$$

(17)

with

$$m = \frac{m_K}{m_{\pi}}, \quad R = \frac{N}{F_{\pi}^2 m_{\pi}^2}, \quad \alpha = \frac{F_K}{F_{\pi} f_+(0)}.$$
 (18)

Using the experimental value  $\alpha = 1.28$ , m = 3.659and the fact that the Ademollo-Gatto theorem<sup>10</sup> requires  $f_+(0)$  to be close to unity, in the exact poledominance approximation ( $\delta_K = \delta_{\pi} = 0$ ) we get the following bounds for *a*:

$$-0.925 \le a \le -0.071$$
, (19)

where we have taken  $f_+(0) = 1.0$ . However, if we use instead the weaker assumption (13) for  $\delta_{\pi}$  and  $\delta_{\kappa}$ , we get the absolute bounds

$$-0.964 \le a \le -0.034$$
 . (20)

Note that both the GMOR and BP solutions, where a = -0.89 and a = -0.17, respectively, are consistent with (19) and (20).

To obtain bounds on b we eliminate  $\gamma$  between  $I_{33}$  and  $\Delta(0)$  and use (1) and (11); we get

$$\frac{9ab}{4(1+a)(1+b)} \ge \frac{Rf_{+}^{2}(0)}{1+\delta_{\pi}} .$$
 (21)

The physical solution for b is expected to lie between 0 and -1; we can therefore write (21) as

$$b \leq \frac{4(1+a)Rf_{+}^{2}(0)}{9a(1+\delta_{\pi})-4(1+a)Rf_{+}^{2}(0)} \quad .$$
 (22)

Given  $\delta_{\pi}$ ,  $f_{+}(0)$ , and *a* the bounds on *b* can be calculated from (22). We have calculated the bounds on *b* by taking  $f_{+}(0) = 1.0$ , taking two different values for  $\delta_{\pi}$  (0 and 0.1, respectively), and varying *a* between 0 and -1. The allowed values for *b* are given by the shaded area in Fig. 1. The shaded area increases as we increase  $\delta_{\pi}$  from 0 to 0.1. In particular, for the GMOR model (a = -0.89) we get



FIG. 1. The physically allowed values of a and b are indicated by the shaded region. The allowed region increases as we increase  $\delta_{\pi}$  from 0 to 0.1.

$$b \le -0.104, \quad \delta_{\pi} = 0$$
  
 $b \le -0.095, \quad \delta_{\pi} = 0.1$  (23)

and for the BP model (a = -0.17) we get

$$b \le -0.823, \quad \delta_{\pi} = 0$$
 (24)  
 $b \le -0.809, \quad \delta_{\pi} = 0.1$ .

We note that in the BP model where the Hamiltonian is approximately SU(3)-invariant the vacuum *cannot* be approximately SU(3)-invariant.

We next turn to the derivation of bounds for  $\Lambda_0$ , the slope of the spin-0 form factor. For this purpose we need an upper bound for  $\Delta(0)$  in (2). Using Eqs. (11), (12), and (14) we get

$$\Delta(0) = \frac{F_{\pi}^{2} m_{\pi}^{2} [3a(2-a)(1+\delta_{\pi}) - 6a(1+a)m^{2}\alpha^{2}f_{+}^{2}(0)(1+\delta_{K})]}{2(1+a)(2-a)}$$
(25)

Now for the physical solution, since  $a \le 0$ , the maximum value of  $\Delta(0)$  is obtained when  $\delta_K$  is maximum and  $\delta_{\pi}$  is minimum. Using the bounds from (13) with  $\delta_K = 1$  and  $\delta_{\pi} = 0$ , we get

$$\Delta_{\max}(0) = \frac{F_{\pi}^{2} m_{\pi}^{2} [3a(2-a) - 12a(1+a)m^{2}\alpha^{2}f_{+}^{2}(0)]}{2(1+a)(2-a)}$$
(26)

Numerically for the GMOR solution (a = -0.89) we obtain

$$\Delta_{\max}(0) = 26.051 F_{\pi}^2 m_{\pi}^2, \quad f_+(0) = 1.0 \tag{27}$$

$$\Delta_{\max}(0) = 18.749 F_{\pi}^{2} m_{\pi}^{2}, \quad f_{+}(0) = 0.9 \quad . \tag{28}$$

In order to obtain the bounds on  $\Lambda_{_{0}}$  we note that we can write (2) as

$$\left| 4(m+1)^2 \Lambda_0 + 2 \left[ \frac{\sqrt{m}}{(\sqrt{m}+1)^2} - 1 \right] \right| \leq \left[ \frac{\Delta_{\max}(0)}{N f_+^2(0)} - 1 \right]^{1/2}.$$
(29)

The relation (29) implies the following bounds for  $\Lambda_0$ :

- 1/0

$$\left|\frac{\Delta_{\max}(0)}{Nf_{+}^{2}(0)} - 1\right|^{1/2} \ge 4(m+1)^{2}\Lambda_{0} + 2\left[\frac{\sqrt{m}}{(\sqrt{m}+1)^{2}} - 1\right]$$
$$\ge -\left[\frac{\Delta_{\max}(0)}{Nf_{+}^{2}(0)} - 1\right]^{1/2}$$
(30)

Using the estimates of  $\Delta_{max}(0)$  from (27) and (28), we get from (30) the following bounds for  $\Lambda_0$ :

$$0.058 \ge \Lambda_0 \ge -0.022, \quad f_+(0) = 1.0$$
 (31)

$$0.057 \ge \Lambda_0 \ge -0.020, \quad f_+(0) = 0.9$$
 (32)

2169

Both the bounds (31) and (32) are inconsistent with the experimental value for  $\Lambda_0$  in (4) and are barely consistent with that in (5). We could have obtained bounds for  $\Lambda_0$  by choosing the BP value of a instead of the GMOR value. It can be readily checked that the BP estimate of a gives bounds on  $\Lambda_0$  which are in even worse agreement with the experimental results in (4) and (5). Therefore, within our framework the only possibility for a relatively large and negative value for  $\Lambda_0$  is when  $\Delta_{\max}(0)$  is about an order of magnitude larger than in Eqs. (27) and (28). This is very unlikely because this requires the kaon pole dominance of  $I_{44}$ to be very badly violated and  $\delta_K$  to be much larger than 1. We can understand this by putting the experimental values of  $\Lambda_0$  in (29) and then calculating  $\Delta_{\max}(0)$ , which via (25) gives bounds on  $\delta_{\kappa}$ . We

- \*Work supported in part by the U. S. Atomic Energy Commission.
- <sup>1</sup>L.-F. Li and H. Pagels, Phys. Rev. D <u>3</u>, 2191 (1971); <u>4</u>, 255 (1971).
- <sup>2</sup>S. Okubo, Phys. Rev. D <u>3</u>, 2807 (1971); <u>4</u>, 725 (1971).
- <sup>3</sup>L. M. Chounet, J. M. Gaillard, and M. K. Gaillard, Phys. Rep. 4C, 199 (1972).
- <sup>4</sup>M. Chounet and M. K. Gaillard, Phys. Lett. <u>32B</u>, 505 (1970); M. K. Gaillard and M. Chounet, CERN Report No. CERN-TH-70-14, 1970 (unpublished).
- <sup>5</sup>S. Glashow and S. Weinberg, Phys. Rev. Lett. <u>20</u>, 244 (1968); M. Gell-Mann, R. J. Oakes, and B. Renner, Phys. Rev. <u>175</u>, 2196 (1968).
- <sup>6</sup>M. K. Gaillard, in *Proceedings of the XVI International Conference on High Energy Physics, Chicago-Batavia, Ill., 1972,* edited by J. D. Jackson and A. Roberts

find

$$\delta_K \ge 1.12^{-1.02}_{+1.65}$$
,  $\Lambda_0 = -0.024 \pm 0.02$  (33)

and

$$\delta_K \ge 12.77^{-5.41}_{+6.85}$$
,  $\Lambda_0 = -0.11 \pm 0.03$ . (34)

Such large violations of kaon pole dominance as in (34) are unwarranted. Therefore, if future experiments confirm a relatively large and negative value for  $\Lambda_0$  as in (4), we believe that we will have to give up either one or more of the assumptions used in the derivation of the bound on  $\Lambda_0$  in (2), or the GWGMOR model on which our estimate of  $\Delta(0)$  is based.

I am thankful to Professor V. S. Mathur for several discussions and a critical reading of the manuscript.

(NAL, Batavia, Ill., 1973), Vol. 2, p. 269; S. A. Wojcicki, *ibid.*, p. 209; G. Donaldson *et al.*, Phys. Rev. Lett. 31, 337 (1973).

- <sup>7</sup>S. Okubo and I-Fu Shih, Phys. Rev. D <u>4</u>, 2020 (1971);
- I-Fu Shih and S. Okubo, Phys. Rev. D 4, 3519 (1971); 6, 1393 (1972); E. E. Radescu, *ibid*. 5, 135 (1972); G. J. Aubrecht, D. M. Scott, K. Tanaka, and R. Torge-
- son, *ibid.* <u>4</u>, 1423 (1971); K. Tanaka and R. Torgeson, *ibid.* <u>5</u>, 1164 (1972); R. Acharya, *ibid.* <u>5</u>, 768 (1972); C. Bourrely, Nucl. Phys. <u>B43</u>, 434 (1972).
- <sup>8</sup>R. Brandt and G. Preparata, Ann. Phys. (N.Y.) <u>61</u>, 119 (1970); Phys. Rev. Lett. <u>26</u>, 1605 (1971).
- <sup>9</sup>V. S. Mathur and S. Okubo, Phys. Rev. D <u>1</u>, 3468 (1970).
   <sup>10</sup>M. Ademollo and R. Gatto, Phys. Rev. Lett. <u>13</u>, 264 (1965).