

## Erratum: Causality-violating Higgs singlets at the LHC [Phys. Rev. D **87**, 045004 (2013)]

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In our original paper, we present a simple 5D metric that allows closed timelike curves (CTCs) as viewed by an observer confined to our 4D brane. The metric mixes time  $t$  and the fifth dimension  $u$ , as

$$d\tau^2 = dt^2 + 2g(u)dtdu - h(u)du^2 - d\vec{r}^2. \quad (1)$$

For definiteness, we assume a “constraint” equation,  $|\bar{g}| < 1$ , where  $\bar{g}$  is the value of the metric element  $g$  when averaged over the extra dimension. In a comment on our paper, Gielen [1] has shown that our assumed relation is not consistent with the timelike condition of the energy-momentum vector in the 5D space. Using “diagonalized” coordinates  $\tilde{t}$  and  $\tilde{u}$ , defined in our paper, and his, Gielen shows, in a simple way, that  $|\bar{g}| > \bar{D}$  must hold for massless particles, where  $D$  is the square root of the metric’s determinant  $\text{Det} = g^2 + h$  and  $\bar{D}$  is its averaged value. In our work, we set  $\text{Det}$  equal to unity, its Minkowski value on the brane. Thus, Gielen’s condition for us becomes  $|\bar{g}| \geq 1$ .

In this erratum, we rederive Gielen’s constraint equation using standard  $t$  and  $u$  coordinates. Then, with this correct constraint in hand, we correct the few parts of our original paper that require change. These parts are the examples presented in Figs. 1 and 2, the parameter region that allows CTCs presented in Fig. 5, and the piece of Sec. VIIB in which we compute the energy dispersion relation for the time-traveling Higgs singlet particles. The main conclusions of our original paper are not changed.

We begin with the proof that  $|\bar{g}| \geq 1$  must hold for particle geodesics enabling CTCs or prearrivals. In our published paper, we identified three constants of geodesic motion, given in Eqs. (3.1), (3.2), and (3.10) as  $\dot{i} + g(u)\dot{u}$ ,  $\dot{r}$ , and  $\dot{u}$ , respectively, where the overdot denotes a derivative with respect to the particle’s proper time  $\tau$ . The first two constants are the inevitable results of the metric depending only on the coordinate  $u$ , while the latter constant occurs when our ansatz  $\text{Det} = 1$  is implemented. Equivalent to any one of the geodesic constants of the motion is the “first integral” constructed by dividing the line element in Eq. (1) by  $d\tau^2$ :

$$\zeta = \dot{i}^2 + 2g(u)\dot{i}\dot{u} - h(u)\dot{u}^2 - \dot{r}^2, \quad (2)$$

where  $\zeta = 1$  for matter and  $\zeta = 0$  for photons. Substituting in the first two constants of motion, and rearranging the right-hand terms a bit, one gets

$$\zeta = (\dot{i} + g(u)\dot{u})^2 - (\text{Det}(u)\dot{u}^2 + \dot{r}^2). \quad (3)$$

Then, substituting Eq. (3.6) from our original article, we arrive at

$$\zeta = \left( \dot{i} + \left( \frac{g(u)}{\sqrt{\text{Det}}} \dot{u}_0 \right) \right)^2 - (\dot{u}_0^2 + \dot{r}^2). \quad (4)$$

Rearranging terms again and then taking the square root gives

$$\left| \dot{i} + \left( \frac{g(u)}{\sqrt{\text{Det}}} \dot{u}_0 \right) \right| = \sqrt{\zeta + \dot{u}_0^2 + \dot{r}^2}. \quad (5)$$

Next, we take the average over the extra-dimensional transit to get

$$\left| \langle \dot{i} \rangle + \left\langle \frac{g}{\sqrt{\text{Det}}} \right\rangle \dot{u}_0 \right| = \sqrt{\zeta + \dot{u}_0^2 + \dot{r}^2}. \quad (6)$$

In our published paper we showed that CTCs and prearrivals require that  $\bar{g} \equiv \langle g(u) \rangle$  and  $\dot{u}_0$  have the same sign, which we conventionally choose to be positive (so “corotating” particles may experience CTCs and prearrivals, whereas “counterrotating” particles may not). We note that  $\langle \dot{i} \rangle = 0$  for a CTC and  $\langle \dot{i} \rangle < 0$  for a prearriving particle, so writing  $\dot{u}_0$  as  $(du/dt)_0(dt/d\tau)_0 \equiv \beta_0\gamma_0$ , and  $\dot{r}_0$  as  $(d\vec{r}/dt)_0(dt/d\tau)_0 \equiv \vec{v}_0\gamma_0$ , we arrive at the final expression for the necessary and sufficient condition for a CTC or a prearrival to occur:

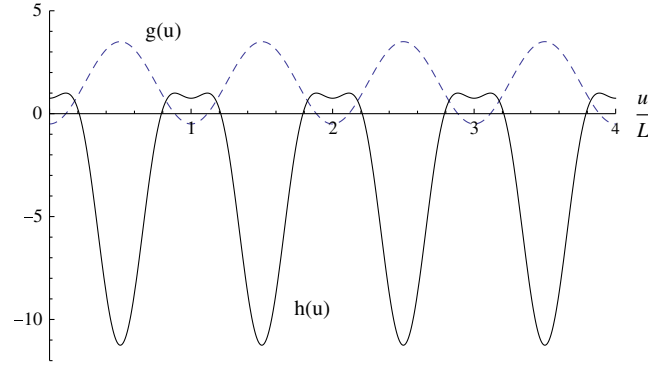


FIG. 1 (color online).  $g(u)$  (dashed line) and  $h(u) = 1 - g^2(u)$  (solid line) versus  $u/L$ , for parameter choices  $g_0 = -0.5$ ,  $a_1 = A = 2$ , and  $a_{n \neq 1} = b_n = 0$ . (Refer to our original paper for the definition of the Fourier mode notation.)

$$\left\langle \frac{g}{\sqrt{\text{Det}}} \right\rangle \geq \sqrt{1 + \frac{(\zeta/\gamma_0^2) + \tilde{v}_0^2}{\beta_0^2}}. \quad (7)$$

(We do not consider here the alternate mathematical solution with very large, negative  $\langle i \rangle$ ; this solution does not connect continuously to the  $\langle i \rangle = 0$  CTC condition.)

The particle's boost factor  $\gamma_0$  may greatly exceed unity, and the particle's velocity  $\tilde{v}_0$  along the brane may be zero. Thus, we may write the necessary but insufficient condition for a CTC or prearrival as simply

$$\left\langle \frac{g}{\sqrt{\text{Det}}} \right\rangle > 1. \quad (8)$$

For a photon (with  $\zeta = 0$ ) but not for massive particles such as Higgs singlet Kaluza-Klein modes (with  $\zeta = 1$ ) of interest to us, the inequality becomes “ $\geq$ .” Thus, with Det taken to be unity, we arrive at the constraint  $\bar{g} \geq 1$  for photons, in agreement with Gielen's result.

The corrected Figs. 1 and 2 that accommodate the constraint  $|\bar{g}| = |g_0 + A| > 1$  are given here [ $g_0 \equiv g(0)$  and  $A \equiv \bar{g} - g_0$ ]. One can easily see from the example in Fig. 2 that Gielen's constraint  $|\bar{g}| > 1$  allows negative travel time.

Next, we correct the energy dispersion relation for the viable time-traveling particles. In Sec. VIIB of [1], the dispersion relation was (correctly) obtained from the equation of motion for the free  $\phi$  field:

$$G^{AB} \partial_A \partial_B \phi + m^2 \phi = 0 \quad [5\text{D Klein-Gordon (KG) equation}]. \quad (9)$$

The general solution to this equation for the  $n$ th energy eigenfunction takes the form

$$\phi_n^{(\text{KG})} = e^{-iE_n[t + \int_0^u g(u) du]} e^{i\vec{p} \cdot \vec{x}} e^{i\xi u}, \quad (10)$$

where  $E_n$  is the energy of the  $n$ th mode (at fixed  $\vec{p}$ ) and  $\vec{p}$  is the standard three-momentum along the brane direction. Since the extra dimension is compactified, we require  $\phi_n(u + L) = \phi_n(u)$ , which in turn requires that

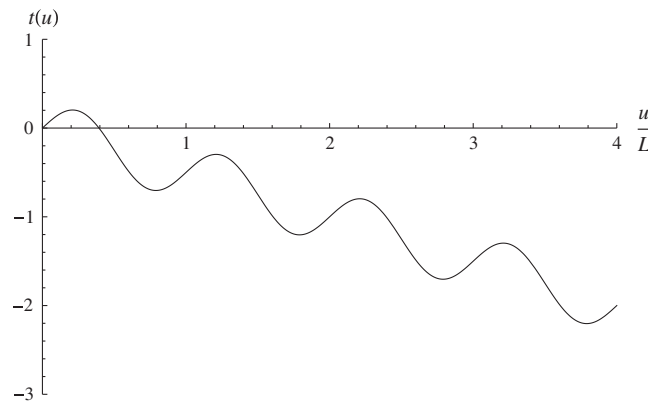


FIG. 2.  $t(u)$  versus  $u/L$ , for the same parameter choices as in Fig. 1, and with  $\beta_0 \equiv (\frac{du}{dt})_0 = 2/3$ .

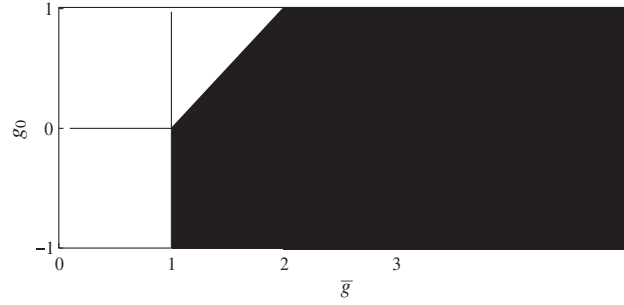


FIG. 3. The shaded region in the  $g_0$ - $\bar{g}$  plane for which CTCs are possible.

$$\xi = \bar{g}E_n - \frac{2\pi n}{L} \quad \text{with} \quad n = 0, \pm 1, \pm 2, \dots \quad (11)$$

Thus, the solution to the KG equation is given by

$$\phi_n^{(\text{KG})} = e^{-iE_n t} e^{i\vec{p} \cdot \vec{x}} e^{-iE_n \int_0^{g-\bar{g}} du} e^{-inu/R}, \quad (12)$$

where we have defined an extra-dimensional “radius”  $R \equiv L/2\pi$  to streamline some notation.

To determine the energy dispersion relation, we simply need to plug Eq. (12) into the 5D KG equation above and solve for  $E_n$ . A bit of algebra yields the quadratic dispersion relation

$$(\bar{g}^2 - 1)E_n^2 - 2\bar{g}E_n \frac{n}{R} + \vec{p}^2 + \frac{n^2}{R^2} + m^2 = 0. \quad (13)$$

Solving for  $E_n$  then gives

$$E_n = \frac{\bar{g} \frac{n}{R} \pm \sqrt{\frac{n^2}{R^2} - (\bar{g}^2 - 1)(\vec{p}^2 + m^2)}}{\bar{g}^2 - 1}. \quad (14)$$

Equation (14) makes it clear that the mode energy  $E_n$  depends on  $\vec{p}$  as well as on  $n$ ; nevertheless, for brevity of notation, we will continue to use the “fixed  $\vec{p}$ ” notation for both  $E_n$  and  $\phi_n$ . In order to ensure that  $E_n$  is real, the condition  $\frac{n^2}{R^2} > (\bar{g}^2 - 1)(\vec{p}^2 + m^2)$  needs to be satisfied. Also, we discard the case with  $\bar{g} < 0$ , which leads to negative  $E_n$ . Both the reality and positive definiteness of  $E_n$  are required to provide the stable modes for the CTC geodesics.

From Eq. (13) we may also derive a lower bound on the time-traveling particle’s energy. The result is

$$E_n > \lim_{\bar{g} \rightarrow 1} E_n = \frac{1}{2} \left( \frac{n}{R} + \frac{\vec{p}^2 + m^2}{(n/R)} \right). \quad (15)$$

The ultimate conditions on the metric which guarantee CTC solutions are simple: (i)  $\bar{g} > 1$  for massive particles (with  $\dot{u}_0$  taken to be positive by convention), (ii)  $|g_0| \leq 1$ , and (iii)  $|A| = |\bar{g} - g_0| > 1$ . The first condition is the result corrected in this erratum, and the second and third conditions are unchanged from our original paper (and derived therein). In Fig. 3 (which replaces Fig. 5 in our original paper), we display the shaded region in the  $g_0$ - $\bar{g}$  plane that allows CTCs. The allowed positive values of  $\bar{g}$  extend to  $+\infty$ .

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[1] S. Gielen, Phys. Rev. D **88**, 068701 (2013).