

What is the role of the meson cloud in the $\Sigma^{*0} \rightarrow \gamma\Lambda$ and $\Sigma^{*+} \rightarrow \gamma\Sigma$ decays?

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We study the effect of the meson cloud dressing in the octet baryon to decuplet baryon electromagnetic transitions. Combining the valence-quark contributions from the covariant spectator quark model with those of the meson cloud estimated based on the flavor SU(3) cloudy bag model, we calculate the transition magnetic form factors at $Q^2 = 0$ ($Q^2 = -q^2$ and q the four-momentum transfer), and also the decuplet baryon electromagnetic decay widths. The result for the $\gamma^*\Lambda \rightarrow \Sigma^{*0}$ decay width is in complete agreement with the data, while that for the $\gamma^*\Sigma^+ \rightarrow \Sigma^{*+}$ is underestimated by 1.4 standard deviations. This achievement may be regarded as a significant advance in the present theoretical situation.

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I. INTRODUCTION

One of the most interesting challenges in hadronic physics is to study the internal structure of baryons and mesons. A microscopic understanding of the transition between the hadronic states is also very important. Although it is generally accepted that the internal structure of hadrons and the dynamics of quarks and gluons are described by QCD, one has to rely on some effective degrees of freedom in the nonperturbative low- Q^2 region such as constituent quarks which form baryon cores with meson cloud excitations [1,2]. Although there exist some works which attempted to treat the meson cloud explicitly as the $q\bar{q}$ excitations in the so-called *unquenched quark models* [3–6], most of the phenomenological models treat the meson cloud using pointlike meson excitations.

Particular examples of very interesting studies may be the electromagnetic transitions between an octet baryon B (spin-1/2) and a decuplet baryon B' (spin-3/2), $\gamma^*B \rightarrow B'$, and the B' electromagnetic decay reactions, $B' \rightarrow \gamma B$. There are theoretical predictions for the $\gamma^*B \rightarrow B'$ transition magnetic moments based on quark models [7–13], including quark models with meson-cloud dressing [14–17], Skyrme and soliton models [18,19], the large- N_c limit [20], QCD sum rules [21,22], and chiral perturbation theory [23]. There are also some results from lattice QCD [24,25]. One of the strong motivations to study the $\gamma^*B \rightarrow B'$ reactions is to clarify the role of the meson cloud dressing, which is of fundamental importance, as was demonstrated by the $\gamma^*N \rightarrow \Delta$ reaction [1,26–30]. The data, except for the $\gamma^*N \rightarrow \Delta$ reaction, namely the $\Sigma^{*0} \rightarrow \gamma\Lambda$ and $\Sigma^{*+} \rightarrow \gamma\Sigma^+$ decay widths, have become available only recently [31–34]. In general, most of the model predictions significantly underestimate the data, particularly those for the $\Sigma^{*+} \rightarrow \gamma\Sigma^+$ decay width (see Ref. [35] for a more detailed discussion).

In our previous work [35] we studied the $\gamma^*B \rightarrow B'$ reactions using a covariant constituent quark model, complemented by the pion-cloud effects extrapolated by the

$\gamma^*N \rightarrow \Delta$ reaction based on an SU(3) symmetry. The pion-cloud effects were included in the leading order, namely, they included only the processes with the direct photon coupling to the pion. The electromagnetic transition form factors calculated were decomposed into the valence-quark and pion-cloud contributions. We concluded that the pion-cloud effects could help to explain satisfactorily the $\gamma^*N \rightarrow \Delta$ data, but only partially help to explain the data for the $\gamma^*\Lambda \rightarrow \Sigma^{*0}$ and $\gamma^*\Sigma^+ \rightarrow \Sigma^{*+}$ reactions. Therefore, the other effects—such as the contributions from the heavier mesons like the kaon, and alternative higher-order processes involving the meson cloud—may be relevant to explain the experimental decay widths. The next-order processes to be included in this study are the processes which one photon couples to the intermediate baryon states while one meson is in the air. In addition, the heavier mesons to be taken into account are the kaon and eta meson, the next lighter mesons to the pion.

In a model with pointlike quarks the photon coupling with the intermediate baryon states is not expected to be important for the meson-cloud contributions, since the octet-to-decuplet electromagnetic transitions are dominated by the magnetic interactions, and the quark anomalous magnetic moments vanish for the pointlike quarks. However, in a constituent quark model like the one we use in this study—the covariant spectator quark model—the octet-decuplet electromagnetic transitions are dominated by the mechanisms with a quark spin flip (magnetic-type interactions). Therefore, the valence-quark contributions may be very important when the quark anomalous magnetic moments are significant. Furthermore, we can expect important meson-cloud contributions from the intermediate octet-decuplet baryon electromagnetic transitions while one meson is in the air (photon-vertex correction), since these mechanisms also depend on the quark anomalous magnetic moments.

In this work we improve the calculation of the meson-cloud contributions for the $\gamma^*B \rightarrow B'$ transitions made in the previous work [35], and predict the corresponding

electromagnetic decay widths (determined at $Q^2 = 0$). The improvements are the following: i) inclusion of the photon coupling to the intermediate baryon states; ii) inclusion of the effects of the heavier meson clouds—the kaon and eta meson—besides the pion. As in the previous work, the transition form factors can be decomposed into the valence-quark and meson-cloud contributions. The processes included as the meson-cloud contributions in this work are depicted in Fig. 1, in terms of the meson and baryon degrees of freedom.

We will conclude that the effects of the intermediate baryon states combined with the kaon cloud improve the agreement of our model with the experimental data.

The valence-quark contributions are estimated based on the covariant spectator quark model [36–40] as in the previous work. Thus, the baryons are described as three-quark systems. The valence-quark contributions for the transition form factors are calculated using the octet and decuplet baryon wave functions and the quark electromagnetic current of the model, determined in the previous works.

To describe the meson-cloud contributions for the octet-decuplet baryon electromagnetic transitions we need a microscopic model to describe the virtual meson-baryon states.

Contrary to the valence-quark contributions that dominate in the large- Q^2 region, the meson-cloud effects are long-range processes, and are known to be of crucial to explain the transition helicity amplitudes and form factors in the low- Q^2 region [1,2]. To incorporate the meson-cloud effects, we use the cloudy bag model (CBM) [15,41–44] which treats the mesons as pointlike particles to describe the meson cloud dressing in the static approximation for the baryons. All such approximations have been practiced well in the past within the CBM, and may be regarded as under control.

Although the CBM framework differs from the covariant spectator quark model for the treatment of the valence quarks, the CBM can be used as an effective description of the long-range physics of the meson-cloud dressing. The possible conflict between the two models e.g., the lack of

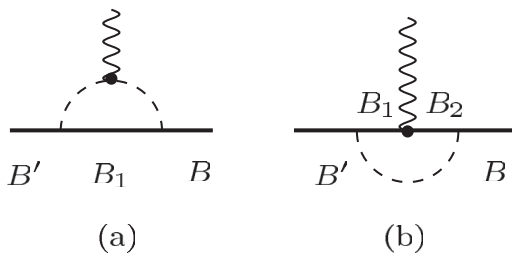


FIG. 1. Meson-cloud contributions for the electromagnetic transition form factors. Between the initial octet (B) and final decuplet (B') baryon states, there are several possible intermediate baryon states: (a) B_1 ; (b) B_1 and B_2 . Depending on the meson M , intermediate meson-baryon states ($M-B_1$ and $M-B_2$) may arise, where B_1 and B_2 are the octet and decuplet baryon states in this study.

the explicit covariance and the limitation of the applicability for the large- Q^2 region in the CBM, can be overcome in a proper manner, since one can define a covariant extension of the model based on covariant parametrizations for the meson-cloud contributions that are equivalent with the CBM result at $Q^2 = 0$; see for instance Refs. [45–47], where meson-cloud contributions were estimated in different reactions. In addition, the merit of using the CBM is that the model is based on SU(3) [SU(6) flavor-spin] symmetry and chiral symmetry.

The explicit calculation of the meson-cloud contributions considered in this study requires two kinds of mechanisms. The first mechanism is the photon coupling with the meson, and for this we use a formalism similar to that applied in Ref. [35]. However, in the present study we take into account the explicit dependence of the baryon and meson (pion, kaon and eta) masses. [Previously, we used SU(3) symmetry for the baryon masses, and only the pion cloud was included.] The second mechanism is the photon coupling with the intermediate baryon states, which is more delicate and model dependent, since this requires an estimate of all the intermediate octet-octet, octet-decuplet, decuplet-octet and decuplet-decuplet transition form factors at $Q^2 = 0$. The corresponding expressions are derived in the CBM framework, but since we describe the valence-quark cores with the covariant spectator quark model, it is necessary to reinterpret the CBM quark magnetic moments in terms of those calculated by the spectator quark model. This will be done using SU(3) symmetry, to be explained in detail later.

Finally, the results from the CBM are normalized by the pion-cloud contribution obtained in the covariant spectator quark model for the $\gamma^*N \rightarrow \Delta$ transition [27], under the assumption that the pion cloud is the dominant meson-cloud contribution. With this procedure—also used in the previous work [35]—we preserve the parametrization of the covariant spectator quark model for the core, and estimate the effects of the meson cloud for the other octet-to-decuplet transitions, as well as the kaon and eta clouds for the $\gamma^*N \rightarrow \Delta$ reaction.

This article is organized as follows. In Sec. II we explain the decomposition of the valence- and meson-cloud contributions for the transition form factors. In Sec. III we review the formalism associated with the valence-quark contributions for the form factors, and express the results in terms of *effective* quark magnetic moments, which are also necessary for the calculation of the meson-cloud effects. In Sec. IV we present the formalism associated with the meson cloud dressing. The results are presented in Sec. V, while the final conclusions are given in Sec. VI.

II. FORMALISM

Next, we discuss briefly the formalism necessary to describe the valence-quark contributions, as well as the mechanism of the meson-cloud dressing.

In the covariant spectator quark model, baryons are treated as three-quark systems [36–40]. The electromagnetic interactions with the baryons are described by the photon coupling with the constituent quarks in the relativistic impulse approximation, and the quark electromagnetic structure is represented in terms of the quark form factors parameterized by a vector-meson dominance mechanism [37,40]. The parametrization of the quark current, calibrated previously in the studies of the nucleon form factors [37] and by the lattice QCD data for the decuplet baryons [40], encodes effectively the gluon and quark-antiquark substructure of the constituent quarks. The baryon wave functions derived from the $SU(6) \otimes O(3)$ structure, are written in terms of an off-shell quark that is free to interact with the photon fields, and two on-shell quarks. Integrating over the quark-pair degrees of freedom, we reduce the three-quark baryon state to a quark-diquark state, where the diquark can be represented as an on-shell spectator particle with an effective mass of m_D [37,38,40].

Under the assumption that each baryon system can be described by the wave function with an S -state configuration for the quark-diquark system in the first approximation, we calculated the valence-quark contributions for the magnetic form factors G_M^B in the previous work [35], where the upper index B labels the contributions from the quark core (bare), using the wave functions from Refs. [40,45]. (See Ref. [35] for more details.) Contributions from the electric and Coulomb quadrupole form factors appear only beyond the S -state approximation for the decuplet baryon wave functions. However, their contributions are expected to be small (small orbital angular momentum admixtures) [28,30], and thus they are neglected in this work.

As mentioned already, the constituent quarks considered in this work have internal structure, and the structure is encoded in a vector-meson dominance parametrization [37,39,40] that includes effectively, among other effects, the meson-cloud dressing of the quarks. However, it should be emphasized that there are meson-cloud effects that cannot be included in the constituent quark structure, such as the process of meson exchange between the different quarks inside the baryon, which cannot be reduced to a simple diagram of a quark dressing. Processes of this kind have to be represented at the hadronic level (meson and baryon states), as in the diagrams shown in Fig. 1. Thus, the meson cloud in this study is regarded as a process of one meson exchange between the different quarks inside the baryon [39]. Since the meson-cloud dressing can appear in two independent mechanisms (self-dressing of the quarks and the others) there is no double counting. In summary, besides the contributions from the valence-quark core calculated using the quark electromagnetic form factors, there are meson-cloud effects that have to be taken into account in the electromagnetic transitions between the baryon states. These meson-cloud effects are the main focus of the present work.

From the discussions made previously, we conclude that the magnetic transition form factors (G_M) can be represented as the sum of the valence-quark (G_M^B) and meson-cloud (G_M^{MC}) contributions in the present approach: $G_M = G_M^B + G_M^{\text{MC}}$. In particular, for the study of the baryon decuplet decay widths, we need to consider only the case $Q^2 = 0$. Thus, we can write

$$G_M(0) = G_M^B(0) + G_M^{\text{MC}}(0). \quad (2.1)$$

As mentioned already, the meson-cloud contribution for the octet to decuplet transition can be decomposed in the two processes displayed in Fig. 1, for the first and second order, classified by the number of the baryon propagators. Figure 1(a) represents the direct coupling of a photon with the intermediate-state meson (first order, one baryon propagator). Figure 1(b) represents the direct coupling of a photon with the intermediate-state baryons (second order, two baryon propagators). One can then decompose $G_M^{\text{MC}}(0)$ into the contributions from Figs. 1(a) and 1(b),

$$G_M^{\text{MC}}(0) = G_M^{\text{MCa}}(0) + G_M^{\text{MCb}}(0). \quad (2.2)$$

The meson-cloud contributions corresponding to Figs. 1(a) and 1(b) can be further decomposed into the pion, kaon, and eta cloud contributions.

As in the previous work [35], in order to keep the parametrization of the covariant spectator quark model, we regularize the results for the pion-cloud contribution by that from the covariant spectator quark model for the $\gamma^*N \rightarrow \Delta$ reaction,

$$G_M^{\text{MC}\pi}(0) = 3\lambda_\pi, \quad (2.3)$$

where $\lambda_\pi = 0.441$ defines the strength of the pion-cloud effect [28]. In this procedure we assume that the pion cloud is the dominant meson-cloud effect in the $\gamma^*N \rightarrow \Delta$ reaction. Later we will see that this assumption is indeed justified. For the other octet-to-decuplet transition reactions, and also for all the meson clouds considered in the present work, we use the relation

$$G_M^{\text{MC}}(0) = f_{BB'}(3\lambda_\pi), \quad (2.4)$$

where the factor $f_{BB'}$ contains the pion, kaon and eta meson-cloud contributions from the both Figs. 1(a) and 1(b). The calculation of the coefficient $f_{BB'}$ will be explained in Sec. V.

Figure 1(b) includes in the intermediate states the octet-octet, octet-decuplet, decuplet-octet and decuplet-decuplet baryon electromagnetic transitions. Therefore, to estimate the contributions from the possible intermediate baryon-state transitions, we need to calculate all the corresponding transition magnetic form factors at $Q^2 = 0$. One can in principle calculate them in the covariant spectator quark models for this purpose; however, the explicit estimates corresponding to all the intermediate baryon-state transitions would be complex and tedious. Therefore, for Fig. 1(b) we use the estimate made in the CBM/ $SU(6)$

framework, where all the intermediate-state contributions can be related with the valence-quark magnetic moments. In order to relate the CBM/SU(6) quark magnetic moments with the anomalous magnetic moments in the covariant spectator quark model, we will start by reviewing the expressions used for the valence-quark contributions for the octet-to-decuplet electromagnetic transition form factors.

III. VALENCE-QUARK CONTRIBUTIONS

In the covariant spectator quark model the valence-quark contributions for the form factors are calculated using the octet and decuplet baryon wave functions, and the constituent quark current. The quark current has the general form [37,40]

$$j_q^\mu(Q^2) = j_1(Q^2)\gamma^\mu + j_2(Q^2)\frac{i\sigma^{\mu\nu}q_\nu}{2M_N}, \quad (3.1)$$

where j_i ($i = 1, 2$) are the quark form factors that can be parametrized in terms of a vector-dominance mechanism. The form factors j_i can also be decomposed in the quark-isoscalar, quark-isovector and strange-quark components. The details can be found in Refs. [35,37,39,40], but are not important for the present discussion, since we are considering the $Q^2 = 0$ case, where $j_1(0) = e_q$ and $j_2(0) = e_q\kappa_q$. The last equation defines the quark anomalous moment (κ_q) in the covariant spectator quark model formalism.

To calculate the transition form factors, we project the operator j_i on the mixed antisymmetric ($|M_A\rangle$) and mixed symmetric ($|M_S\rangle$) components of the octet and (fully symmetric) decuplet $|B'\rangle$ flavor states,

$$j_i^A = 3\langle B'|j_i|M_A\rangle, \quad (3.2)$$

$$j_i^S = 3\langle B'|j_i|M_S\rangle. \quad (3.3)$$

More details can be found in Refs. [35,39,40]. Note that, for the octet-to-decuplet baryon transitions only the components j_i^S (isovector) are relevant. Finally, the magnetic form factor can be written [35] as

$$G_M^B = \frac{2\sqrt{2}}{3}\sqrt{\frac{2}{3}}\bar{f}_v I, \quad (3.4)$$

with

$$\bar{f}_v = \frac{2M_B}{M_{B'} + M_B} \left\{ \frac{j_1^S}{\sqrt{2}} + \frac{M_{B'} + M_B}{2M_N} \frac{j_2^S}{\sqrt{2}} \right\}, \quad (3.5)$$

where the coefficients $\frac{1}{\sqrt{2}}j_i^S$ can be found in Ref. [35], and I is the overlap integral between the octet (ψ_B) and decuplet ($\psi_{B'}$) radial wave functions (see details also in Ref. [35]). The radial wave functions ψ_B and $\psi_{B'}$ are scalar functions of the baryon and diquark momenta [35,39,40,45]. In Eqs. (3.4) and (3.5) G_M^B , \bar{f}_v and I are exclusive functions of Q^2 .

Now we focus again on the $Q^2 = 0$ case. In this case we can write the factor \bar{f}_v in a more compact form, defining the *effective* quark magnetic moment of the transition $\gamma^*B \rightarrow B'$ as

$$\hat{\mu}_q = \frac{2M_B}{M_{B'} + M_B} + \frac{M_B}{M_N}\kappa_q. \quad (3.6)$$

Note that the expression for $\hat{\mu}_q$ is reduced to the usual form, $\mu_q = (1 + \kappa_q)$, in the limit $M_{B'} = M_B = M_N$. However, $\hat{\mu}_q$ now depends on the masses of the ‘‘submultiplets’’ ($M_{B'}$ and M_B in the $\gamma^*B \rightarrow B'$ transition). We keep this dependence in mind, but suppress the indices B and B' in $\hat{\mu}_q$ for simplicity.

The explicit expressions for \bar{f}_v in terms of $\hat{\mu}_q$ are presented in Table I. In particular, we can express the reactions involving the Σ and Ξ as

$$\bar{f}_v = \frac{1}{6}(2\hat{\mu}_u - \hat{\mu}_d + 2\hat{\mu}_s) + \frac{1}{6}(2\hat{\mu}_u + \hat{\mu}_d)t_3, \quad (3.7)$$

where $t_3 = J_3$ for Σ and $t_3 = \tau_3$ for Ξ . The matrices $J_3 = \text{diag}(1, 0, -1)$ and $\tau_3 = \text{diag}(1, -1)$ are, respectively, isospin-1 and isospin-1/2 operators that act on the isospin states of the baryons B and B' .

TABLE I. Coefficients j_i^S ($i = 1, 2$), \bar{f}_v and valence-quark contributions for the G_M form factors. In the expressions for G_M^B , one has $A = \frac{2\sqrt{2}}{3}$.

	\bar{f}_v	G_M^B	G_M^B (CBM)
$\gamma^*p \rightarrow \Delta^+$	$\frac{1}{3}(2\hat{\mu}_u + \hat{\mu}_d)$	$A\frac{1}{3}(2\mu_u + \mu_d)$	$A\bar{\mu}_u$
$\gamma^*n \rightarrow \Delta^0$	$\frac{1}{3}(2\hat{\mu}_u + \hat{\mu}_d)$	$A\frac{1}{3}(2\mu_u + \mu_d)$	$A\bar{\mu}_u$
$\gamma^*\Lambda \rightarrow \Sigma^{*0}$	$\sqrt{\frac{3}{4}}\frac{1}{3}(2\hat{\mu}_u + \hat{\mu}_d)$	$\sqrt{\frac{3}{4}}A\frac{1}{3}(2\mu_u + \mu_d)$	$\sqrt{\frac{3}{4}}A\bar{\mu}_u$
$\gamma^*\Sigma^+ \rightarrow \Sigma^{*+}$	$\frac{1}{3}(2\hat{\mu}_u + \hat{\mu}_s)$	$A\frac{1}{3}(2\mu_u + \mu_s)$	$A\frac{1}{3}(2\bar{\mu}_u + \mu_s)$
$\gamma^*\Sigma^0 \rightarrow \Sigma^{*0}$	$\frac{1}{6}(2\hat{\mu}_u - \hat{\mu}_d + 2\hat{\mu}_s)$	$A\frac{1}{6}(2\mu_u - \mu_d + 2\mu_s)$	$A\frac{1}{3}(\bar{\mu}_u + 2\mu_s)$
$\gamma^*\Sigma^- \rightarrow \Sigma^{*-}$	$\frac{1}{3}(-\hat{\mu}_d + \hat{\mu}_s)$	$A\frac{1}{3}(-\mu_d + \mu_s)$	$A\frac{1}{3}(-\bar{\mu}_u + \mu_s)$
$\gamma^*\Xi^0 \rightarrow \Xi^{*0}$	$\frac{1}{3}(2\hat{\mu}_u + \hat{\mu}_s)$	$A\frac{1}{3}(2\mu_u + \mu_s)$	$A\frac{1}{3}(2\bar{\mu}_u + \mu_s)$
$\gamma^*\Xi^- \rightarrow \Xi^{*-}$	$\frac{1}{3}(-\hat{\mu}_d + \hat{\mu}_s)$	$A\frac{1}{3}(-\mu_d + \mu_s)$	$A\frac{1}{3}(-\bar{\mu}_u + \mu_s)$

The contributions from the valence quarks for the form factors can also be estimated by the SU(6) quark model in terms of the quark magnetic moments μ_q . The results are expressed in terms of the u , d and s quark magnetic moments, μ_u , μ_d and μ_s , respectively. The CBM uses the spin-flavor SU(6) wave functions, but calculates the values of μ_q using the CBM (MIT bag) formalism. Since usually $\mu_u \equiv \mu_d$ in the CBM (reflecting the quark masses used, $m_u = m_d$), we use $\bar{\mu}_u$ to represent either μ_u or μ_d .

Note that the definitions of the quark magnetic moments discussed here do not include the quark charges, contrarily to the convention used for instance in naive quark models [48,49].

The results for G_M^B (quark-core contributions) from the CBM are presented in the last column in Table I. For the $\gamma^*N \rightarrow \Delta$ reaction the result is [8,26]

$$G_M^B(0) = \frac{2\sqrt{2}}{3} \bar{\mu}_u \quad [\text{SU}(6)], \quad (3.8)$$

where $\bar{\mu}_u = \mu_p$, the proton magnetic moment in the SU(2) limit.¹ In this limit also $\mu_n = -\frac{2}{3} \bar{\mu}_u$.

In order to compare the results of the covariant spectator quark model with those of the CBM, we need to relate the spectator-model quark magnetic moments with those of the CBM/SU(6). Motivated by Eq. (3.8), and taking into account that the structure of G_M^B in the covariant spectator quark model, given by Eqs. (3.4) and (3.5), we define $\mu_q \equiv \sqrt{\frac{2}{3}} \hat{\mu}_q I(0)$, or

$$\mu_q = \sqrt{\frac{2}{3}} \left\{ \frac{2M_B}{M_{B'} + M_B} + \frac{M_B}{M_N} \kappa_q \right\} I(0), \quad (3.9)$$

for the covariant spectator quark model.

With the above identification of the quark magnetic moments, we can write the $\gamma^*N \rightarrow \Delta$ magnetic form factor in the covariant spectator quark model as

$$G_M^B(0) = \frac{2\sqrt{2}}{3} \frac{1}{3} (2\mu_u + \mu_d) \quad (\text{Spectator}). \quad (3.10)$$

Note the similarity between Eqs. (3.8) and (3.10). If we replace $\frac{1}{3}(2\mu_u + \mu_d) \rightarrow \bar{\mu}_u$, as in the SU(2) symmetric case, the two equations are equivalent. The expressions for the other octet-to-decuplet transitions are presented in Table I, with $A = \frac{2\sqrt{2}}{3}$. Also, for the other reactions the expressions from the covariant spectator quark model and CBM are equivalent in the SU(2) symmetric limit, although in the case of the covariant spectator quark model μ_q varies from reaction to reaction.

¹For simplicity we ignore the factor $\sqrt{\frac{M_N}{M_\Delta}}$ that transforms the magnetic moment $\mu_{N\Delta}$ into the corresponding form factor $G_M(0)$. This simplification has no consequence in the present work, since to identify the results of the CBM and those of the spectator quark model global factors are not important.

Thus, we can relate the CBM/SU(6) results of the bare core (valence-quark contributions) with those of the covariant spectator quark model defining the quark magnetic moments by Eq. (3.9). The quark magnetic moments of the covariant spectator quark model generalize the *usual* magnetic moments by the inclusion of the octet and decuplet baryon-mass dependence (M_B and $M_{B'}$). Therefore, the effective quark magnetic moment μ_q , defined by Eq. (3.9), differs from transition to transition in the covariant spectator quark model. However, as mentioned already, the *familiar* expression is recovered in the limit $M_{B'} = M_B = M_N$, apart from some constants.

Another interesting point is the dependence of μ_q on the overlap integral $I(0)$, which is a consequence of the difference between the octet and decuplet radial wave functions of the transition. In a naive picture with $M_B = M_{B'}$, the octet and decuplet radial wave functions can be approximated by the same radial wave function (defined in the same frame) and the overlap integral would be $I(0) = 1$. In the present case, as discussed in Ref. [35], $I(0)$ is about 0.8–0.9, depending on the transitions.

A note is in order about the SU(6) result for $G_M^B(0)$, given by Eq. (3.8) for the $\gamma^*N \rightarrow \Delta$ reaction. The numerical result using the experimental value for μ_p is $G_M^B(0) = 2.3$ (including the effect of the nucleon and Δ masses; see footnote 1). This result overestimates the relativistic calculations. Some relativistic calculations take into account the differences between the nucleon and Δ masses and also the nonzero momentum of the nucleon at $Q^2 = 0$ in the Δ rest frame. The Sato-Lee model [50] for instance gives $G_M^B(0) = 2.05$. As for the covariant spectator quark model, we recall that the model predicts the upper limit of $G_M^B(0) = 2.07$ [27,35], but in practice this value is reduced by the overlap of the nucleon and Δ radial wave functions, $I(0)$, which is always smaller than unity, as already mentioned. See Appendix B in Ref. [35] for details. For the present study it is not important even if our expressions differ from those of the SU(6) by a factor. For example, the factor $\sqrt{\frac{2}{3}}$ may be a consequence of the relativistic calculation. Also, the overlap integral $I(0)$ does not appear in the simple SU(6) quark-model expressions due to the static approximation [$I(0) \rightarrow 1$]. The important point is to establish the correspondence between the analytical expressions in the SU(6) quark model and those of the spectator formalism consistently.

Next, we comment on the renormalization of the baryon wave functions. In the present calculation we use the decuplet baryon radial wave functions from Ref. [40] and those of the octet baryons from Ref. [45]. In these cases the decuplet baryon wave functions were determined assuming that they have no meson-cloud dressing, while the octet baryon wave functions were determined assuming a small pion-cloud dressing. As discussed already in Ref. [35], the correction due to the renormalization of the octet baryon wave functions (due to the pion-cloud dressing) is small,

and can be neglected in a first approximation (less than 4% effect). In the present work we include kaon and eta clouds in addition to the pion cloud. Although we cannot calculate the renormalization effects due to these mesons for the baryon wave functions in the covariant spectator quark model framework, we will assume—as was already done for the pion cloud—that the meson-cloud effects are small and can be neglected in the normalization of the wave functions in a first approximation. Later we will discuss the renormalization effect due to the meson-cloud effects, since the meson-cloud contributions also depend on the wave functions.

IV. MESON-CLOUD CONTRIBUTIONS

To estimate the meson-cloud contributions for the processes shown by Figs. 1(a) and 1(b), we apply the CBM [41]. As usually practiced in the CBM, we use the static approximation and neglect the momentum of the baryons in the initial, intermediate and final baryon states by replacing the respective energies by their masses [15,41–43]. The same approximation is also used in the heavy baryon chiral perturbation theory [22,51–53]. In addition, we ignore the possible center-of-mass correction for the three-quark composite baryon (core) systems, keeping in mind that this correction reduces the bare core transition amplitudes for the $\gamma^*N \rightarrow \Delta$ reaction by 5% to 10% in the region $Q^2 \lesssim 0.5 \text{ GeV}^2$ [15]. The effects are expected to be even smaller for the remaining reactions since the corresponding baryons are heavier.

Although the approximations discussed above break the Lorentz covariance, the phenomenological successes and practices in describing the physics in the low- Q^2 region [15,41–43] suggest that the approximations may be well under control in the present study, particularly at $Q^2 = 0$ (small kinematic corrections).

To carry out the calculations of the meson-cloud contributions, all the intermediate states are summed over utilizing the standard angular momentum algebra in flavor and spin spaces combined with the Wigner-Eckart theorem [41–43]. Thus, the summation is made based on the SU(6) symmetry at the flavor-spin wave-function level. The SU(3)-breaking effects are partially included by using the physical baryon and meson masses, and via the quark masses, $m_u = m_d \neq m_s$. Note that in the covariant spectator quark model the SU(3) symmetry is explicitly broken in the octet and decuplet baryon wave functions.

The equations derived in the CBM for $Q^2 = 0$ depend only on one-dimensional integrals. In some cases the CBM integrals have singularities in the integrand functions (poles associated with physical baryons or mesons in the intermediate states). These poles yield imaginary parts for the calculated integrals. For simplicity we evaluate these integrals using the *principal value integral*. Based on the results from the CBM [15] for the $\gamma^*N \rightarrow \Delta$ reaction, we

may expect the imaginary part to be about 15%–20% of the real part near $Q^2 = 0$. Since the decay width depends on $|G_M(0)|^2$, this approximation has only a small effect on the final results [a 20% imaginary part of the real part in $G_M(0)$ leads to a 4% correction for $|G_M(0)|^2$].

A. Direct coupling with the meson

We first consider the contributions from the processes represented by Fig. 1(a). In the following the upper index M stands for the meson ($M = \pi, K$). The CBM loop integral functions [41,42] for the initial (B) and final (B') baryons that also depend on the intermediate baryon B_1 states for the pion- and kaon-cloud diagrams will be denoted by $H_{BB'}^\pi(B_1)$ and $H_{BB'}^K(B_1)$, respectively.

The contributions from Fig. 1(a) for the $\gamma^*B \rightarrow B'$ can be written as

$$G_M^{\text{MC}a} = \sum_{M,B_1} C_{BB';B_1}^M H_{BB'}^M(B_1), \quad (4.1)$$

where $C_{BB';B_1}^M$ are the coefficients calculated in the CBM framework, and are presented in the Appendix.

The explicit expression for $H_{BB'}^M(B_1)$ is

$$H_{BB'}^M(B_1) = \frac{1}{12\pi^2} \left(\frac{f_{\pi NN}}{m_\pi} \right)^2 \times \int_0^\infty dk \left\{ \frac{k^4 [j_0(kR) + j_2(kR)]^2}{\omega_k [4\omega_k^2 - (M_{B_1} - M_B)^2]} \times \frac{4\omega_k + 2M_{B_1} - M_{B'} - M_B}{(M_{B_1} - M_B + \omega_k)(M_{B_1} - M_{B'} + \omega_k)} \right\}, \quad (4.2)$$

where R is the bag radius, j_l ($l = 0, 2$) are the spherical Bessel functions arising from the CBM form factor, $\omega_k = \sqrt{m_M^2 + k^2}$ for $M = \pi, K$ is the meson energy, and $f_{\pi NN}$ is the pion-nucleon coupling constant. As already mentioned the integral symbol should be read as the principal value integral.

In the present work we take a typical, successful value for the bag radius, $R = 1 \text{ fm}$ [43]. The dependence of the calculated quantities on the values of the bag radius chosen can be found in Refs. [15,41,42].

The factor $\left(\frac{f_{\pi NN}}{m_\pi}\right)^2$ is included in the loop integral definition for all baryon and meson cases, since all the couplings are redefined in terms of $f_{\pi NN}$. For discussions about the renormalized $f_{\pi NN}$ value used in the CBM see Refs. [15,42].

B. Coupling with intermediate baryon states

Next, we consider the contributions from Fig. 1(b) due to the clouds of the pion, kaon and η meson. Since the processes depend on the intermediate baryon states B_1 and B_2 , the respective contributions generally depend on the intermediate-state transition form factors between B_1 and B_2 , and these can, in the SU(6) quark model, be

represented by the combinations of the quark magnetic moments μ_q .

In order to obtain a simple estimate for the meson-cloud contributions without explicitly summing over a huge number of the intermediate states, we use a technique developed and used in the CBM framework [43] with the exact isospin symmetry, $\mu_u = \mu_d$. The use of the isospin symmetry simplifies the calculation drastically by reducing the number of terms to be considered, and can be justified when the difference between μ_u and μ_d is small.

In the following calculations of the meson-cloud effects we will replace the CBM quark magnetic moments by those of the covariant spectator quark model as defined by Eq. (3.9). In order to keep the isospin symmetry in these calculations we replace μ_u and μ_d by an average $\bar{\mu}_u$ to be defined later.

Then, the contributions from Fig. 1(b) for the $\gamma^*B \rightarrow B'$ transition can be written as

$$G_M^{\text{MBb}} = \sum_{M, B_1, B_2} D_{BB'; B_1 B_2}^M H_{BB'}^{2M}(B_1, B_2), \quad (4.3)$$

where the CBM-based integral is represented by $H_{BB'}^{2M}$ to be defined next, and $D_{BB'; B_1 B_2}^M$ are the coefficients which depend on the effective magnetic moments μ_q . The expressions for $D_{BB'; B_1 B_2}^M$ are given in the Appendix.

The integral $H_{BB'}^{2M}$ is defined by

$$\begin{aligned} H_{BB'}^{2M}(B_1, B_2) &= \frac{1}{12\pi^2} \left(\frac{f_{\pi NN}}{m_\pi} \right)^2 \\ &\times \int_0^\infty dk \left[\frac{k^4 [j_0(kR) + j_2(kR)]^2}{\omega_k (M_{B_1} - M_B + \omega_k) (M_{B_2} - M_{B'} + \omega_k)} \right]. \end{aligned} \quad (4.4)$$

Again, the principal value integration should be understood. In Eq. (4.4) B_1 and B_2 are the baryons in the intermediate states with masses M_{B_1} and M_{B_2} , respectively, and the upper index $2M$ indicates that there are two baryon propagators while a meson M is in the air. Besides that, the functions are obtained with a static approximation—the same as for the $H_{BB'}^M(B_1)$ case—but we also now have $\omega_k = \sqrt{m_\eta^2 + k^2}$ when $M = \eta$.

To be consistent with the SU(2) symmetry in the calculation of the function $D_{BB'; B_1 B_2}^M$ for the meson-cloud contributions, we replace in the expressions for the magnetic moments, μ_u and μ_d , by the average

$$\bar{\mu}_u \equiv \frac{1}{3} (2\mu_u + \mu_d). \quad (4.5)$$

With this definition, the results for the core contributions are the same for the CBM and the covariant spectator quark model for the reactions $\gamma^*N \rightarrow \Delta$ and $\gamma^*\Lambda \rightarrow \Sigma^{*0}$. The same expression will be used for the calculation of the meson-cloud contributions represented by Fig. 1(b).

As for the reactions involving the Σ and Ξ in the calculation of the meson-cloud effects, we use the replacement suggested by Eq. (3.7),

$$\begin{aligned} &\frac{1}{6} (2\mu_u - \mu_d + 2\mu_s) + \frac{1}{6} (2\mu_u + \mu_d) t_3 \\ &\rightarrow \frac{1}{6} (\bar{\mu}_u + 2\mu_s) + \frac{1}{2} \bar{\mu}_u t_3, \end{aligned} \quad (4.6)$$

where in the last line we have replaced μ_u and μ_d by $\bar{\mu}_u$. In practice the difference between μ_u , μ_d and $\bar{\mu}_u$ is smaller than 10%.

The effect of the baryon wave-function renormalization due to the meson cloud represented by the Figs. 1(a) and 1(b) can be absorbed in the renormalized coupling constant $f_{\pi NN}$ used in the CBM [15].

V. RESULTS

Before presenting the results, we recall that the contributions from the valence-quark core are given by the covariant spectator quark model as discussed in Sec. III. The formalism was discussed in detail in Ref. [35]. The important point to recall is that for $Q^2 = 0$ the valence-quark contributions G_M^B depend only on the quark anomalous moments κ_u , κ_d , κ_s and the octet/decuplet radial wave functions through $I(0)$. The corresponding parameters were fixed in the previous works.

We start to present the results by discussing the pion-cloud contributions for the $\gamma^*N \rightarrow \Delta$ reaction, and explain how the meson-cloud contributions are calibrated by this reaction. Next, we will present the results for all meson-cloud contributions for the $\gamma^*B \rightarrow B'$ reactions, and discuss the final results of the form factor G_M and decay widths Γ . Finally, we will compare our results with those existing in the literature.

A. Pion-cloud contributions for the $\gamma^*N \rightarrow \Delta$ reaction

The results for the pion-cloud contribution arising from the Figs. 1(a) and 1(b), from the CBM, are presented in Table II (see entry CBM). As we can see in Table II, the final contribution from the pion cloud in this case is 1.634, which is larger than the estimate made by the covariant spectator quark model of 1.323 by about 33% (see entry Spectator). This is not surprising, since the CBM tends to overestimate the effect of the pion cloud for the $\gamma^*N \rightarrow \Delta$

TABLE II. Pion-cloud contributions for the $\gamma^*N \rightarrow \Delta$ reaction. The quantities with the superscript π refer only to the pion cloud. The first entry includes the results from the CBM, while the second includes the corresponding quantities from the covariant spectator quark model.

	$G_M^B(0)$	$G_M^{\text{MC}\pi}(0)$	$G_M^{\text{MCb}\pi}(0)$	$G_M^{\text{MC}\pi}(0)$	$G_M(0)$
CBM [$\tilde{G}_M(0)$]	1.633	0.883	0.754	1.634	3.270
Spectator [$G_M(0)$]	1.633	0.713	0.610	1.323	2.956

reaction. Indeed, in Ref. [15] the pion cloud gives a contribution of about 66% of the total, a contribution substantially larger than in the other calculations [1,27,50].

On the other hand, the pion-cloud contribution in the covariant spectator quark model were determined by a fit to the $\gamma^*N \rightarrow \Delta$ data for $Q^2 \leq 6 \text{ GeV}^2$, combined with the estimate made for the quark-core contribution extracted by the Excited Baryon Analysis Center model [50] by removing the meson-cloud contributions. In addition, the estimate of the quark-core contributions from the covariant spectator quark model was compared successfully with the results of lattice QCD simulations using the pion masses around 350–650 MeV, where the pion-cloud contribution is expected to be suppressed. To compare with the lattice QCD data, the model was generalized to the lattice QCD regime using the vector-meson dominance parametrization for the quark current; see details in Refs. [29,30,39,40]. All these results show that the covariant spectator quark model provides a robust description of both the physical and lattice QCD data, and that it is probably more appropriated than the CBM parametrization for the present study. The small deviation from the result for $Q^2 = 0$ (bare plus pion cloud) given by the covariant quark model, 2.96, compared to the experimental result of 3.02 ± 0.03 [54] is a consequence of the global fit of the covariant spectator quark model for $Q^2 \leq 6 \text{ GeV}^2$, instead of fitting only to the low- Q^2 region data.

In the following we will use \tilde{G}_M to represent the CBM result, and G_M for the present model (covariant spectator quark model). Also, to distinguish between the different $\gamma^*B \rightarrow B'$ reactions, we will use the argument BB' as $G_M(BB')$. Recall that we are only discussing the form factors at $Q^2 = 0$.

B. Meson-cloud contributions for the $\gamma^*B \rightarrow B'$ reactions

In order to keep the successful features of the pion-cloud contributions estimated in the covariant spectator quark model for the $\gamma^*N \rightarrow \Delta$ transition, we normalize the CBM result for the pion cloud $\tilde{G}_M^{\text{MC}\pi}(0)$ by the result of the covariant spectator quark model $G_M^{\text{MC}\pi}(0) = 3\lambda_\pi = 1.32$, given by

$$G_M^{\text{MC}\pi}(N\Delta) = \mathcal{R}\tilde{G}_M^{\text{MC}\pi}(N\Delta), \quad (5.1)$$

where

$$\mathcal{R} = \frac{3\lambda_\pi}{\tilde{G}_M^{\text{MC}\pi}(N\Delta)}. \quad (5.2)$$

Numerically, it gives $\mathcal{R} \simeq 0.81$.

To estimate the effect of the other meson clouds, the kaon and η meson in the $\gamma^*N \rightarrow \Delta$ reaction, and also for the other octet-to-decuplet transitions, we use a similar relation, since all the couplings are related with the coupling constant $f_{\pi NN}$. Thus, we use in general

$$G_M^{\text{MC}}(BB') = \mathcal{R}\tilde{G}_M^{\text{MC}}(BB'). \quad (5.3)$$

Except for the fact that we now include Fig. 1(b), the procedure is the same as the one used in the previous work [35]. From Eq. (2.4), we get

$$f_{BB'} = \frac{\tilde{G}_M^{\text{MC}}(BB')}{\tilde{G}_M^{\text{MC}\pi}(N\Delta)}. \quad (5.4)$$

The results of the meson-cloud contributions from Figs. 1(a) and 1(b) for the octet-to-decuplet electromagnetic transition form factors at $Q^2 = 0$ are presented in Table III. Since we expect the results for $\gamma^*n \rightarrow \Delta^0$ and

TABLE III. Meson-cloud contributions for the octet-to-decuplet transition magnetic moments. In each group the first line indicates the pion-cloud contributions, the second line the kaon cloud, the third line the eta cloud, and the fourth the sum of all the meson-cloud contributions (boldface). The column $G_M^B(0)$ presents the contributions from the valence-quark core. The column $G_M^{\text{MC}}(0)$ and $G_M(0)$ show, respectively, the total meson-cloud contributions and the final results (both boldface). The results in the first line for $G_M^{\text{MC}}(0)$ and $G_M(0)$ include only the pion-cloud contributions.

	$G_M^B(0)$	$G_M^{\text{MC}a}(0)$	$G_M^{\text{MC}b}(0)$	$G_M^{\text{MC}}(0)$	$G_M(0)$
$\gamma^*N \rightarrow \Delta$	1.633	0.713	0.610	1.323	2.956
		0.017	0.037		
			0.0062		
		0.730	0.652	1.383	3.016
$\gamma^*\Lambda \rightarrow \Sigma^{*0}$	1.683	0.669	0.358	1.027	2.710
		0.068	0.289		
			0.016		
		0.737	0.663	1.400	3.083
$\gamma^*\Sigma^+ \rightarrow \Sigma^{*+}$	2.094	0.149	0.513	0.663	2.757
		0.155	0.269		
			0.043		
		0.304	0.825	1.129	3.224
$\gamma^*\Sigma^0 \rightarrow \Sigma^{*0}$	0.969	0.000	0.270	0.270	1.239
		0.104	0.010		
			0.015		
		0.104	0.387	0.490	1.460
$\gamma^*\Sigma^- \rightarrow \Sigma^{*-}$	-0.156	-0.149	0.026	-0.124	-0.279
		0.052	-0.065		
			-0.012		
		-0.097	-0.052	-0.149	-0.305
$\gamma^*\Xi^0 \rightarrow \Xi^{*0}$	2.191	0.222	0.086	0.308	2.499
		0.187	0.519		
			0.086		
		0.410	0.691	1.101	3.291
$\gamma^*\Xi^- \rightarrow \Xi^{*-}$	-0.168	-0.222	0.084	-0.138	-0.306
		0.038	-0.108		
			-0.0034		
		-0.185	-0.028	-0.213	-0.380

$\gamma^* p \rightarrow \Delta^+$ to be the same in the present approach, we use the label $\gamma^* N \rightarrow \Delta$ to represent both reactions.

In Table III we can see the contributions from each meson: π , K or η . We can conclude that the pion cloud indeed gives the dominant meson-cloud contribution for the $\gamma^* N \rightarrow \Delta$ reaction. However, for the other reactions, the kaon cloud in particular can give important contributions. The magnitude of the kaon plus eta cloud contributions can be obtained by subtracting the result of the first line (only pion-cloud effects) from the last line (bold, total) for $G_M(0)$. We can then conclude that the kaon- and eta-cloud corrections are about 0.4 for $\gamma^* \Lambda \rightarrow \Sigma^{*0}$ and 0.5 for $\gamma^* \Sigma^+ \rightarrow \Sigma^{*+}$.

The kaon-cloud effects in some cases are comparable or larger than those of the pion cloud, particularly for Fig. 1(b). See for instance the reactions $\gamma^* \Lambda \rightarrow \Sigma^{*0}$, $\gamma^* \Sigma^+ \rightarrow \Sigma^{*+}$ and $\gamma^* \Xi^0 \rightarrow \Xi^{*0}$.

Globally, the meson-cloud contributions can be about 45% of the total for the cases with $|G_M(0)| \approx 3$, or even larger for the $\gamma^* \Sigma^{*-} \rightarrow \Sigma^-$ and $\gamma^* \Xi^{*-} \rightarrow \Xi^-$ cases.

Another interesting point is the magnitude of the contributions from Fig. 1(b). They are in most cases similar or larger than the contributions from Fig. 1(a). This is a consequence of two main factors: i) the quark magnetic moments μ_q are significant (about 2–3 nuclear magneton), which enhances the effect; ii) Fig. 1(b) has a large number of intermediate states to be summed over (see the Appendix).

Note that this feature contradicts the assumption made in the previous work [35], namely that Fig. 1(a) is expected to give the leading-order contribution. However, since in the previous study the meson-cloud contributions were normalized by the total pion-cloud contribution of the $\gamma^* N \rightarrow \Delta$ transition, the difference between the previous and the new results for all the reactions is not drastic. Based on the present results for the $\gamma^* N \rightarrow \Delta$ transition, where roughly 50% of the meson cloud comes from Figs. 1(a) and 1(b), we may regard the previous result as a consequence of the assumption that both diagrams have the same effect (50%) for all the reactions. Recall that only the pion was considered in the previous work.

In the present study, we also normalize the meson-cloud contributions by the pion-cloud contribution for the $\gamma^* N \rightarrow \Delta$ transition, but we leave the contributions from Figs. 1(a) and 1(b) independent, as can be seen in Table III.

Then, we can conclude that the explicit inclusion of the contributions from Fig. 1(b), increases the contribution of the meson cloud, and improves the description of the $\Sigma^{*0} \rightarrow \gamma \Lambda$ and $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ data, as will be discussed next. For further discussion, we recall that the $\Sigma^{*0} \rightarrow \gamma \Lambda$ and $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ decay widths given by the Particle Data Group (PDG) [31] were underestimated by, respectively, 1.2 and 2.4 standard deviations in the previous work [35].

The final results for $G_M(0)$ are also presented in Table IV, in comparison with the estimates extracted from the experimental decay widths [35]. The estimates were made assuming the dominance of $G_M(0)$ compared to the quadrupole electric form factor $G_E(0)$. From Table IV, one can see that the present model can describe well the data for the $\Sigma^{*0} \rightarrow \gamma \Lambda$ (less than one standard deviation), and underestimates the $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ data by 1.5 standard deviations. These features may be regarded as a significant improvement compared to our previous result and other theoretical estimates (see discussion in the next section).

Using the model results obtained for $G_M(0)$, we calculate the decuplet electromagnetic decay widths, assuming the dominance of G_M ,

$$\Gamma_{B' \rightarrow \gamma B} = \frac{\alpha}{16} \frac{(M_{B'}^2 - M_B^2)^3}{M_{B'}^3 M_B^2} |G_M(0)|^2, \quad (5.5)$$

where $\alpha = \frac{e^2}{4\pi} \approx \frac{1}{137}$ is the electromagnetic fine structure constant. The G_M dominance is a good approximation according to theoretical estimates and the experimental results for the $\gamma^* N \rightarrow \Delta$. (See Ref. [35] for a more detailed discussion.)

Our predictions for the decay width, $\Gamma \equiv \Gamma_{B' \rightarrow \gamma B}$, are also presented in Table IV. For the cases of $\Delta \rightarrow \gamma N$ and $\Sigma^{*0} \rightarrow \gamma \Lambda$, we also present the results from PDG [31]. In addition, we present the results for $\Sigma^{*0} \rightarrow \gamma \Lambda$ and $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ from Refs. [33,34]. Our model results deviate from the data only for the $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ reaction by 1.4 standards

TABLE IV. Results for $G_M(0)$ corresponding to the $B' \rightarrow \gamma B$ decays. The values for $|G_M(0)|_{\text{exp}}$ are estimated by Eq. (5.5) using the experimental values of $\Gamma_{B' \rightarrow \gamma B}$.

	$G_M(0)$	$ G_M(0) _{\text{exp}}$	Γ (keV)	Γ_{exp} (keV)
$\Delta \rightarrow \gamma N$	3.02	3.04 ± 0.11 [31]	648	660 ± 47 [31]
$\Sigma^{*0} \rightarrow \gamma \Lambda$	3.08	3.35 ± 0.57 [31] 3.26 ± 0.37 [33,34]	399	470 ± 160 [31] 445 ± 102 [33,34]
$\Sigma^{*+} \rightarrow \gamma \Sigma^+$	3.22	4.10 ± 0.57 [33]	154	250 ± 70 [33]
$\Sigma^{*0} \rightarrow \gamma \Sigma^0$	1.46		32	
$\Sigma^{*-} \rightarrow \gamma \Sigma^-$	-0.31	<0.8 [55]	1.4	<9.5 [55]
$\Xi^{*0} \rightarrow \gamma \Xi^0$	3.29		182	
$\Xi^{*-} \rightarrow \gamma \Xi^-$	-0.38		2.4	

TABLE V. Results for the $B' \rightarrow \gamma B$ decay widths (in keV) given by several works.

	$\Delta \rightarrow \gamma N$	$\Sigma^{*0} \rightarrow \gamma \Lambda$	$\Sigma^{*+} \rightarrow \gamma \Sigma^+$
<i>U</i> -spin [7,35]		292 ± 27	138 ± 13
HB χ PT [23]	670–790	252–540	70–220
Algebraic model [13]	342–344	221.3	140.7
QCD SR [22]	887	409	150
Large N_c [20]	669 ± 42	336 ± 81	149 ± 36
Spectator	648	399	154
Data [31,33,34]	660 ± 47	470 ± 160	250 ± 70
		445 ± 102	

deviations. The upper limit for the $\Sigma^{*-} \rightarrow \gamma \Sigma^-$ reaction [55] is also shown in Table IV.

We call attention to the fact that the present estimate of the meson-cloud contribution is affected by some uncertainties related to the effective quark magnetic moments used to calculate Fig. 1(b). Since some of the intermediate states correspond to elastic transitions [where $I(0) = 1$], we can question the use of the prescription (3.9) with the factor $I(0) \leq 1$ (octet-decuplet radial wave function overlap integral) given by the inelastic $\gamma^* B \rightarrow B'$ reaction. Therefore, an upper limit for the meson-cloud contribution can be obtained by setting $I(0) = 1$ (perfect overlap of the radial wave functions). In this case the final results for the decay width are enhanced by 1%–6%. In particular, the $\Sigma^{*0} \rightarrow \gamma \Lambda$ decay width increases by 4.4% and that for the $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ case by 2.6%. Therefore, in the latter case a possible enhancement due to the intermediate-state baryon wave-function overlap is small, and it does not significantly increase the final result enough to bring the present result closer to the experimental data.

Taking into account the typical uncertainty in the CBM of about 10%, and also assuming that the meson-cloud contribution is about 50% of the total in the CBM, this gives about 5% ambiguity. Combining the two ambiguities—one from the wave function overlap of 1%–6%, and the other from the CBM estimate for meson-cloud contributions of about 5%—we can conclude that our estimate can be affected by a value around 10%.

Another interesting exercise can be to check if the discrepancy between our estimate and the experimental $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ decay width may be a consequence of neglecting the effect of $G_E(0)$ in the calculation of the decay width. In this case, the deviation from the data would be the result of dropping the term $3|G_E(0)|^2$ in the sum with $|G_M(0)|^2$ in the decay-width calculation. Using our result for $|G_M(0)|$, we would be able to reproduce the experimental $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ decay width if $|G_E(0)|$ is about 30% of $|G_M(0)|$.

C. Discussion

In general, most of the existing quark models underestimate the $\Sigma^{*0} \rightarrow \gamma \Lambda$ and $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ decay widths by about 50%. Chiral quark models with mesonic effects are

also included in this category [12,16,17]. In those cases they give a $\Delta \rightarrow \gamma N$ decay width of about 400 keV, smaller than the experimental result of 660 ± 47 keV. A more detailed comparison between the model results and data can be found in the previous work [35].

A better result for the $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ decay width is obtained by an algebraic model of the hadronic structure [13]. However, the Δ decay width is again underestimated (see Table V).

Some calculations give values that are closer to the experimental results, e.g., the heavy baryon chiral perturbation theory [23] as presented in Table V. The windows associated with the results are, however, too broad.

Also, the predictions based on *U*-spin symmetry proposed in the 1970s [7] give a good description of the data, using the updated result for the $\Delta \rightarrow \gamma N$ decay width [35]. As can be seen in Table V, the best result differs, respectively, by 0.8 and 1.4 standard deviations for the Λ and Σ^+ cases. It is worth mentioning that the estimates made in Ref. [33]—also based on a *U*-spin symmetry—have a much better agreement with the data. However, as also discussed in our previous work [35], their *U*-spin symmetry-based estimates did not take into account the effect of the baryon masses in the conversion between the form factors and the helicity amplitudes.

The results from QCD sum rules [22] are close to the $\Sigma^{*+} \rightarrow \gamma \Sigma^+$ and $\Sigma^{*0} \rightarrow \gamma \Lambda$ decay-width data, but overestimate the $\Delta \rightarrow \gamma N$ decay width by about 220 keV.

An excellent description of all the data was also obtained in Ref. [20] using a $1/N_c$ expansion. The unknown coefficients in the expansion are fitted to the known octet and decuplet magnetic moments (n , p , Σ^\pm , $\Xi^{0,-}$, Δ^+ , Ω^- and $\Sigma^0 \rightarrow \gamma \Lambda$ transition), providing a prediction for the remaining cases. Note that the values from Ref. [20] are very close to our own results.

We would like to emphasize that our estimate is a pure prediction, since the parameters involved in the calculations (quarks anomalous magnetic moments and wave functions) were already determined and calibrated in the previous works. The only adjustable ingredient in the present calculation is the magnitude of the pion-cloud contribution for the $\gamma^* N \rightarrow \Delta$ reaction, chosen to match the pion-cloud contribution of the original covariant spectator quark model [28].

VI. CONCLUSIONS

In this work we have studied the decuplet-to-octet electromagnetic decay widths, which are related to the magnetic transition form factors defined at $Q^2 = 0$. To describe the baryon quark core we have used the covariant spectator quark model, the model parameters of which were calibrated in the previous works on the octet and decuplet baryon systems. To estimate the effects of the meson cloud, including those from the pion, kaon and eta meson, we have been guided by the cloudy bag model, improved by the result from the covariant spectator quark model for the $\gamma^*N \rightarrow \Delta$ reaction. The effects included as the meson cloud are the direct photon coupling to the meson [Fig. 1(a)] and the photon coupling to the intermediate baryon states while one meson is in the air [Fig. 1(b)].

We conclude that the inclusion of the contributions from Fig. 1(b), as well as the effects of the kaon cloud [Figs. 1(a) and 1(b)], are both very important. When the meson-cloud contributions are combined with the quark-core contributions calculated by the covariant spectator quark model the present model can reproduce the experimental results well. The inclusion of only the valence-quark contributions leads to significant underestimates of the data. The meson-cloud effects are particularly important for the reactions $\Sigma^{*0} \rightarrow \gamma\Lambda$ and $\Sigma^{*+} \rightarrow \gamma\Sigma^+$. Furthermore, the effect of Fig. 1(b) is also very important for the $\Delta \rightarrow \gamma N$ reaction.

In summary, we are able to describe the $\Sigma^{*0} \rightarrow \gamma\Lambda$ decay width very well, and also obtain a very reasonable result (1.4 standard deviations) for that of the $\Sigma^{*+} \rightarrow \gamma\Sigma^+$. The present approach also describes the $\Delta \rightarrow \gamma N$ decay width rather well. However, in the last case, the agreement is a consequence of the fit made previously by the model, although the explicit inclusion of the extra kaon-cloud effects improves the agreement slightly.

Our predictions for the transition form factors of the other reactions are consistent with the estimates made based on the U -spin symmetry, namely, $G_M(\Sigma^{*+}\Sigma^+) \approx G_M(\Xi^{*0}\Xi^0)$, and $G_M(\Sigma^{*-}\Sigma^-) \approx G_M(\Xi^{*-}\Xi^-)$.

We can, in general, conclude that the meson-cloud effects are of fundamental importance to describe the $\gamma^*B \rightarrow B'$ reactions—especially in the low- Q^2 region—and the decuplet baryon decay widths. To further test the

conclusions of the present study, an accurate experimental determination of the unknown decuplet baryon electromagnetic decay widths is crucial. In addition, precise lattice QCD simulations for several pion-mass values can also help to constrain the contributions from the valence quarks and test our estimates of the quark-core contributions.

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APPENDIX: MESON-CLOUD CONTRIBUTIONS

The meson-cloud contributions calculated by the CBM corresponding to Fig. 1(a) are presented in Table VI. Compared to the results presented in Ref. [35] the present results include a factor $\sqrt{\frac{2}{3}}(2M_B)$ in each transition. The factor is necessary to represent the form factors in their natural units (dimensionless). Note that the factor $(2M_B)$ multiplied by $H_{BB'}^M(B_1)$ gives a dimensionless quantity.

In the exact SU(3) limit, the factors $\sqrt{\frac{2}{3}}(2M_B)$ become the same for all the octet-to-decuplet transitions, and as a consequence the factor can be ignored in the calculation of $f_{BB'}$ since the factors will be canceled out by the normalization, and divided by the $\gamma^*N \rightarrow \Delta$ contribution following the procedure of this work. Taking the limit $H_{BB'}^\pi(B_1) = H_\pi$ (independent of the octet and decuplet baryon masses) and $H_{BB'}^K(B_1) = 0$, we recover the previous results given in Ref. [35] for the contributions from Fig. 1(a).

To calculate the contributions from Fig. 1(b), it is convenient to define the following quantities:

$$\begin{aligned} \mu_S &= \frac{1}{3}(\bar{\mu}_u + 2\mu_s), & \mu_V &= \bar{\mu}_u, \\ \mu_1 &= \frac{1}{3}(2\bar{\mu}_u + \mu_s), & \mu_2 &= \bar{\mu}_u - \mu_s, \\ \mu_3 &= \frac{1}{9}(\bar{\mu}_u + 8\mu_s), & \mu_4 &= \frac{1}{3}(-\bar{\mu}_u + 4\mu_s). \end{aligned} \quad (\text{A1})$$

TABLE VI. Meson-cloud contributions for G_M from Fig. 1(a).

	$\tilde{G}_M^{\text{MCa}}(BB')$
$\gamma^*N \rightarrow \Delta$	$\tilde{G}_M^{\text{MCa}}(N\Delta) = \frac{4\sqrt{2}}{9}(2M_N)[\frac{1}{5}H_{N\Delta}^\pi(N) + H_{N\Delta}^\pi(\Delta) + \frac{1}{25}H_{N\Delta}^K(\Sigma) + \frac{1}{5}H_{N\Delta}^K(\Sigma^*)]$
$\gamma^*\Lambda \rightarrow \Sigma^*$	$\tilde{G}_M^{\text{MCa}}(\Lambda\Sigma^*) = \frac{2\sqrt{2}}{15\sqrt{3}}(2M_\Lambda)[\frac{4}{5}H_{\Lambda\Sigma^*}^\pi(\Sigma) + 4H_{\Lambda\Sigma^*}^\pi(\Sigma^*) + \frac{3}{5}H_{\Lambda\Sigma^*}^K(N) - \frac{1}{5}H_{\Lambda\Sigma^*}^K(\Xi) + 2H_{\Lambda\Sigma^*}^K(\Xi^*)]$
$\gamma^*\Sigma \rightarrow \Sigma^*$	$\tilde{G}_M^{\text{MCa}}(\Sigma\Sigma^*) = \frac{\sqrt{2}}{3}(2M_\Sigma)[\frac{2}{75}H_{\Sigma\Sigma^*}^K(N) + \frac{8}{15}H_{\Sigma\Sigma^*}^K(\Delta) + \frac{2}{15}H_{\Sigma\Sigma^*}^K(\Xi) + \frac{4}{15}H_{\Sigma\Sigma^*}^K(\Xi^*)] + \frac{\sqrt{2}}{3}(2M_\Sigma)[\frac{4}{25}H_{\Sigma\Sigma^*}^\pi(\Lambda) - \frac{8}{75}H_{\Sigma\Sigma^*}^\pi(\Sigma) + \frac{4}{15}H_{\Sigma\Sigma^*}^\pi(\Sigma^*) - \frac{2}{75}H_{\Sigma\Sigma^*}^K(N) + \frac{4}{15}H_{\Sigma\Sigma^*}^K(\Delta) + \frac{2}{15}H_{\Sigma\Sigma^*}^K(\Xi) + \frac{4}{15}H_{\Sigma\Sigma^*}^K(\Xi^*)]J_3$
$\gamma^*\Xi \rightarrow \Xi^*$	$\tilde{G}_M^{\text{MCa}}(\Xi\Xi^*) = \frac{\sqrt{2}}{3}(2M_\Xi)[-\frac{1}{25}H_{\Xi\Xi^*}^K(\Lambda) + \frac{1}{5}H_{\Xi\Xi^*}^K(\Sigma) + \frac{2}{5}H_{\Xi\Xi^*}^K(\Sigma^*) + \frac{2}{5}H_{\Xi\Xi^*}^K(\Omega)] + \frac{\sqrt{2}}{3}(2M_\Xi)[\frac{4}{75}H_{\Xi\Xi^*}^\pi(\Xi) + \frac{4}{15}H_{\Xi\Xi^*}^\pi(\Xi^*) + \frac{1}{25}H_{\Xi\Xi^*}^K(\Lambda) + \frac{2}{15}H_{\Xi\Xi^*}^K(\Sigma) + \frac{2}{15}H_{\Xi\Xi^*}^K(\Sigma^*) + \frac{2}{5}H_{\Xi\Xi^*}^K(\Omega)]\tau_3$

Note that the quantities above are dependent on the transitions under consideration [see Eq. (3.9)].

We can now write the meson-cloud contributions corresponding to Fig. 1(b) as

$$\begin{aligned} \tilde{G}_M^{\text{MCb}}(N\Delta) = & \frac{2\sqrt{2}}{3} \mu_V \left\{ \frac{4}{9} H_{N\Delta}^{2\pi}(N, N) + \frac{5}{9} H_{N\Delta}^{2\pi}(N, \Delta) + \frac{8}{225} H_{N\Delta}^{2\pi}(\Delta, N) + \frac{4}{9} H_{N\Delta}^{2\pi}(\Delta, \Delta) + \frac{4}{25} H_{N\Delta}^{2K}(\Lambda, \Sigma) + \frac{1}{5} H_{N\Delta}^{2K}(\Lambda, \Sigma^*) \right. \\ & \left. + \frac{8}{225} H_{N\Delta}^{2K}(\Sigma, \Sigma) + \frac{4}{45} H_{N\Delta}^{2K}(\Sigma^*, \Sigma^*) - \frac{1}{45} H_{N\Delta}^{2K}(\Sigma, \Sigma^*) + \frac{4}{225} H_{N\Delta}^{2K}(\Sigma^*, \Sigma) + \frac{1}{15} H_{N\Delta}^{2\eta}(N, \Delta) \right\}, \end{aligned} \quad (\text{A2})$$

$$\begin{aligned} \tilde{G}_M^{\text{MCb}}(\Lambda\Sigma^{*0}) = & \sqrt{\frac{2}{3}} \frac{M_\Lambda}{M_N} \mu_V \left\{ \frac{8}{75} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma, \Lambda) + \frac{32}{225} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma, \Sigma) + \frac{8}{45} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma, \Sigma^*) + \frac{4}{75} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma^*, \Lambda) - \frac{8}{225} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma^*, \Sigma) \right. \\ & + \frac{16}{45} H_{\Lambda\Sigma^*}^{2\pi}(\Sigma^*, \Sigma^*) + \frac{4}{15} H_{\Lambda\Sigma^*}^{2K}(N, N) + \frac{8}{15} H_{\Lambda\Sigma^*}^{2K}(N, \Delta) + \frac{4}{225} H_{\Lambda\Sigma^*}^{2K}(\Xi, \Xi) + \frac{4}{45} H_{\Lambda\Sigma^*}^{2K}(\Xi, \Xi^*) \\ & \left. + \frac{8}{225} H_{\Lambda\Sigma^*}^{2K}(\Xi^*, \Xi) + \frac{8}{45} H_{\Lambda\Sigma^*}^{2K}(\Xi^*, \Xi^*) + \frac{8}{75} H_{\Lambda\Sigma^*}^{2\eta}(\Lambda, \Sigma) \right\}, \end{aligned} \quad (\text{A3})$$

$$\begin{aligned} \tilde{G}_M^{\text{MCb}}(\Sigma\Sigma^*) = & \frac{\sqrt{2}}{3} \frac{M_\Sigma}{M_N} \left\{ \mu_S \left[\frac{16}{45} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma, \Sigma^*) + \frac{8}{225} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma^*, \Sigma) + \frac{4}{75} H_{\Sigma\Sigma^*}^{2\eta}(\Sigma^*, \Sigma) + \frac{4}{9} H_{\Sigma\Sigma^*}^{2K}(\Xi, \Xi^*) - \frac{8}{225} H_{\Sigma\Sigma^*}^{2K}(\Xi^*, \Xi) \right] \right. \\ & + \mu_1 \left[\frac{32}{225} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma, \Sigma) + \frac{8}{75} H_{\Sigma\Sigma^*}^{2\eta}(\Sigma, \Sigma) \right] + \mu_s \frac{8}{75} H_{\Sigma\Sigma^*}^{2\pi}(\Lambda, \Lambda) - \mu_2 \frac{16}{135} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma^*, \Sigma^*) \\ & + \mu_V \left[\frac{4}{225} H_{\Sigma\Sigma^*}^{2K}(N, N) + \frac{16}{45} H_{\Sigma\Sigma^*}^{2K}(\Delta, \Delta) \right] + \mu_3 \frac{4}{15} H_{\Sigma\Sigma^*}^{2K}(\Xi, \Xi) + \mu_4 \frac{8}{45} H_{\Sigma\Sigma^*}^{2K}(\Xi^*, \Xi^*) \left. \right\} \\ & + J_3 \frac{\sqrt{2}}{3} \frac{M_\Sigma}{M_N} \mu_V \left\{ -\frac{8}{75} H_{\Sigma\Sigma^*}^{2\pi}(\Lambda, \Sigma) + \frac{4}{15} H_{\Sigma\Sigma^*}^{2\pi}(\Lambda, \Sigma^*) + \frac{16}{75} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma, \Lambda) + \frac{32}{225} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma, \Sigma) \right. \\ & + \frac{8}{45} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma, \Sigma^*) - \frac{4}{75} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma^*, \Lambda) + \frac{4}{225} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma^*, \Sigma) - \frac{8}{45} H_{\Sigma\Sigma^*}^{2\pi}(\Sigma^*, \Sigma^*) + \frac{4}{45} H_{\Sigma\Sigma^*}^{2K}(N, N) \\ & - \frac{4}{45} H_{\Sigma\Sigma^*}^{2K}(N, \Delta) + \frac{16}{225} H_{\Sigma\Sigma^*}^{2K}(\Delta, N) + \frac{8}{9} H_{\Sigma\Sigma^*}^{2K}(\Delta, \Delta) + \frac{4}{45} H_{\Sigma\Sigma^*}^{2K}(\Xi, \Xi) + \frac{4}{9} H_{\Sigma\Sigma^*}^{2K}(\Xi, \Xi^*) \\ & \left. - \frac{8}{225} H_{\Sigma\Sigma^*}^{2K}(\Xi^*, \Xi) - \frac{8}{45} H_{\Sigma\Sigma^*}^{2K}(\Xi^*, \Xi^*) + \frac{16}{75} H_{\Sigma\Sigma^*}^{2\eta}(\Sigma, \Sigma) + \frac{4}{75} H_{\Sigma\Sigma^*}^{2\eta}(\Sigma^*, \Sigma) \right\}, \end{aligned} \quad (\text{A4})$$

$$\begin{aligned} \tilde{G}_M^{\text{MCb}}(\Xi\Xi^*) = & \frac{\sqrt{2}}{3} \frac{M_\Xi}{M_N} \left\{ \mu_S \left[-\frac{1}{15} H_{\Xi\Xi^*}^{2\pi}(\Xi, \Xi^*) + \frac{4}{75} H_{\Xi\Xi^*}^{2\pi}(\Xi^*, \Xi) + \frac{1}{5} H_{\Xi\Xi^*}^{2\eta}(\Xi, \Xi^*) + \frac{4}{75} H_{\Xi\Xi^*}^{2\eta}(\Xi^*, \Xi) \right] \right. \\ & + \frac{2}{3} H_{\Xi\Xi^*}^{2K}(\Sigma, \Sigma^*) - \frac{4}{75} H_{\Xi\Xi^*}^{2K}(\Sigma^*, \Sigma) \left. \right] + \mu_3 \left[+\frac{2}{25} H_{\Xi\Xi^*}^{2\pi}(\Xi, \Xi) + \frac{6}{25} H_{\Xi\Xi^*}^{2\eta}(\Xi, \Xi) \right] \\ & + \mu_4 \left[\frac{2}{15} H_{\Xi\Xi^*}^{2\pi}(\Xi^*, \Xi^*) - \frac{2}{15} H_{\Xi\Xi^*}^{2\eta}(\Xi^*, \Xi^*) \right] + \mu_1 \frac{4}{15} H_{\Xi\Xi^*}^{2K}(\Sigma, \Sigma) \\ & + \mu_s \left[\frac{4}{75} H_{\Xi\Xi^*}^{2K}(\Lambda, \Lambda) + \frac{8}{15} H_{\Xi\Xi^*}^{2K}(\Omega, \Omega) \right] + \mu_2 \frac{8}{45} H_{\Xi\Xi^*}^{2K}(\Sigma^*, \Sigma^*) \left. \right\} \\ & + \tau_3 \frac{\sqrt{2}}{3} \frac{M_\Xi}{M_N} \mu_V \left\{ -\frac{2}{225} H_{\Xi\Xi^*}^{2\pi}(\Xi, \Xi) + \frac{1}{45} H_{\Xi\Xi^*}^{2\pi}(\Xi, \Xi^*) + \frac{2}{45} H_{\Xi\Xi^*}^{2\pi}(\Xi^*, \Xi^*) - \frac{4}{225} H_{\Xi\Xi^*}^{2\pi}(\Xi^*, \Xi) \right. \\ & - \frac{4}{75} H_{\Xi\Xi^*}^{2K}(\Lambda, \Sigma) + \frac{2}{15} H_{\Xi\Xi^*}^{2K}(\Lambda, \Sigma^*) + \frac{4}{15} H_{\Xi\Xi^*}^{2K}(\Sigma, \Lambda) + \frac{16}{45} H_{\Xi\Xi^*}^{2K}(\Sigma, \Sigma) + \frac{4}{9} H_{\Xi\Xi^*}^{2K}(\Sigma, \Sigma^*) \\ & + \frac{4}{75} H_{\Xi\Xi^*}^{2K}(\Sigma^*, \Lambda) - \frac{8}{225} H_{\Xi\Xi^*}^{2K}(\Sigma^*, \Sigma) + \frac{16}{45} H_{\Xi\Xi^*}^{2K}(\Sigma^*, \Sigma^*) + \frac{2}{25} H_{\Xi\Xi^*}^{2\eta}(\Xi, \Xi) + \frac{1}{5} H_{\Xi\Xi^*}^{2\eta}(\Xi, \Xi^*) \\ & \left. + \frac{4}{75} H_{\Xi\Xi^*}^{2\eta}(\Xi^*, \Xi) + \frac{2}{15} H_{\Xi\Xi^*}^{2\eta}(\Xi^*, \Xi^*) \right\}. \end{aligned} \quad (\text{A5})$$

In the equations above the factor $\frac{M_B}{M_N}$ is a consequence of the factor $\sqrt{\frac{2}{3}}(2M_B)$ combined with $1/(2M_N)$ in units of the quark magnetic moments. The final result is thus dimensionless.

- [1] V. D. Burkert and T. S. H. Lee, *Int. J. Mod. Phys. E* **13**, 1035 (2004).
- [2] I. G. Aznauryan *et al.*, *Int. J. Mod. Phys. E* **22**, 1330015 (2013).
- [3] P. Geiger and N. Isgur, *Phys. Rev. D* **55**, 299 (1997).
- [4] S. Capstick and W. Roberts, *Prog. Part. Nucl. Phys.* **45**, S241 (2000).
- [5] R. Bijker and E. Santopinto, *Phys. Rev. C* **80**, 065210 (2009).
- [6] S. Capstick *et al.*, *Eur. Phys. J. A* **35**, 253 (2008).
- [7] H. J. Lipkin, *Phys. Rev. D* **7**, 846 (1973).
- [8] R. Koniuk and N. Isgur, *Phys. Rev. D* **21**, 1868 (1980); **23**, 818(E) (1981).
- [9] J. W. Darewych, M. Horbatsch, and R. Koniuk, *Phys. Rev. D* **28**, 1125 (1983).
- [10] M. Warns, W. Pfeil, and H. Rollnik, *Phys. Lett. B* **258**, 431 (1991).
- [11] R. K. Sahoo, A. R. Panda, and A. Nath, *Phys. Rev. D* **52**, 4099 (1995).
- [12] G. Wagner, A. J. Buchmann, and A. Faessler, *Phys. Rev. C* **58**, 1745 (1998).
- [13] R. Bijker, F. Iachello, and A. Leviatan, *Ann. Phys. (N.Y.)* **284**, 89 (2000).
- [14] E. Kaxiras, E. J. Moniz, and M. Soyeur, *Phys. Rev. D* **32**, 695 (1985).
- [15] D.-H. Lu, A. W. Thomas, and A. G. Williams, *Phys. Rev. C* **55**, 3108 (1997).
- [16] L. Yu, X.-L. Chen, W.-Z. Deng, and S.-L. Zhu, *Phys. Rev. D* **73**, 114001 (2006).
- [17] N. Sharma, H. Dahiya, P. K. Chatley, and M. Gupta, *Phys. Rev. D* **81**, 073001 (2010).
- [18] C. L. Schat, C. Gobbi, and N. N. Scoccola, *Phys. Lett. B* **356**, 1 (1995).
- [19] T. Haberichter, H. Reinhardt, N. N. Scoccola, and H. Weigel, *Nucl. Phys. A* **615**, 291 (1997).
- [20] R. F. Lebed and R. H. TerBeek, *Phys. Rev. D* **83**, 016009 (2011).
- [21] T. M. Aliev and A. Ozpineci, *Nucl. Phys. B* **732**, 291 (2006).
- [22] L. Wang and F. X. Lee, *Phys. Rev. D* **80**, 034003 (2009).
- [23] M. N. Butler, M. J. Savage, and R. P. Springer, *Nucl. Phys. B* **399**, 69 (1993).
- [24] D. B. Leinweber, T. Draper, and R. M. Woloshyn, *Phys. Rev. D* **48**, 2230 (1993).
- [25] D. Arndt and B. C. Tiburzi, *Phys. Rev. D* **69**, 014501 (2004).
- [26] V. Pascalutsa, M. Vanderhaeghen, and S. N. Yang, *Phys. Rep.* **437**, 125 (2007).
- [27] G. Ramalho, M. T. Peña, and F. Gross, *Eur. Phys. J. A* **36**, 329 (2008).
- [28] G. Ramalho, M. T. Peña, and F. Gross, *Phys. Rev. D* **78**, 114017 (2008).
- [29] G. Ramalho and M. T. Peña, *J. Phys. G* **36**, 115011 (2009).
- [30] G. Ramalho and M. T. Peña, *Phys. Rev. D* **80**, 013008 (2009).
- [31] K. Nakamura *et al.* (Particle Data Group), *J. Phys. G* **37**, 075021 (2010).
- [32] S. Taylor *et al.* (CLAS Collaboration), *Phys. Rev. C* **71**, 054609 (2005); **72**, 039902(E) (2005).
- [33] D. Keller *et al.* (CLAS Collaboration), *Phys. Rev. D* **83**, 072004 (2011).
- [34] D. Keller *et al.* (CLAS Collaboration), *Phys. Rev. D* **85**, 059903 (2012).
- [35] G. Ramalho and K. Tsushima, *Phys. Rev. D* **87**, 093011 (2013).
- [36] G. Ramalho, F. Gross, M. T. Peña, and K. Tsushima, in *Proceedings of the 4th Workshop on Exclusive Reactions at High Momentum Transfer*, edited by A. Radyushkin (World Scientific, Singapore, 2011), p. 287.
- [37] F. Gross, G. Ramalho, and M. T. Peña, *Phys. Rev. C* **77**, 015202 (2008).
- [38] F. Gross, G. Ramalho, and M. T. Peña, *Phys. Rev. D* **85**, 093005 (2012).
- [39] G. Ramalho and K. Tsushima, *Phys. Rev. D* **84**, 054014 (2011).
- [40] G. Ramalho, K. Tsushima, and F. Gross, *Phys. Rev. D* **80**, 033004 (2009).
- [41] A. W. Thomas, *Adv. Nucl. Phys.* **13**, 1 (1984).
- [42] S. Theberge and A. W. Thomas, *Nucl. Phys. A* **393**, 252 (1983).
- [43] Y. Kohyama, K. Oikawa, K. Tsushima, and K. Kubodera, *Phys. Lett. B* **186**, 255 (1987); K. Tsushima, T. Yamaguchi, M. Takizawa, Y. Kohyama, and K. Kubodera, *Phys. Lett. B* **205**, 128 (1988); K. Tsushima, T. Yamaguchi, Y. Kohyama, and K. Kubodera, *Nucl. Phys. A* **489**, 557 (1988); T. Yamaguchi, K. Tsushima, Y. Kohyama, and K. Kubodera, *Nucl. Phys. A* **500**, 429 (1989).
- [44] Y. Umino and F. Myhrer, *Nucl. Phys. A* **554**, 593 (1993).
- [45] G. Ramalho, K. Tsushima, and A. W. Thomas, *J. Phys. G* **40**, 015102 (2013).
- [46] G. Ramalho and K. Tsushima, *Phys. Rev. D* **82**, 073007 (2010).
- [47] G. Ramalho and K. Tsushima, *Phys. Rev. D* **86**, 114030 (2012).
- [48] C. Amsler *et al.* (Particle Data Group), *Phys. Lett. B* **667**, 1 (2008). See page 1023 for information about the naive quark model.
- [49] I. C. Cloet, D. B. Leinweber, and A. W. Thomas, *Phys. Rev. C* **65**, 062201 (2002).
- [50] B. Julia-Diaz, T. S. Lee, T. Sato, and L. C. Smith, *Phys. Rev. C* **75**, 015205 (2007).
- [51] E. E. Jenkins and A. V. Manohar, *Phys. Lett. B* **255**, 558 (1991).
- [52] V. Bernard, N. Kaiser, J. Kambor, and U. G. Meissner, *Nucl. Phys. B* **388**, 315 (1992).
- [53] V. Bernard, *Prog. Part. Nucl. Phys.* **60**, 82 (2008).
- [54] L. Tiator, D. Drechsel, O. Hanstein, S. S. Kamalov, and S. N. Yang, *Nucl. Phys. A* **689**, 205 (2001).
- [55] V. V. Molchanov *et al.* (SELEX Collaboration), *Phys. Lett. B* **590**, 161 (2004).