

**Kinematic cusps with two missing particles. II. Cascade decay topology**Tao Han,<sup>1</sup> Ian-Woo Kim,<sup>2</sup> and Jeonghyeon Song<sup>3</sup><sup>1</sup>*Pittsburgh Particle Physics, Astrophysics, and Cosmology Center, Department of Physics and Astronomy, University of Pittsburgh, 3941 O'Hara Street, Pittsburgh, Pennsylvania 15260, USA*<sup>2</sup>*Department of Physics, University of Michigan, Ann Arbor, Michigan 48109, USA*<sup>3</sup>*Division of Quantum Phases and Devices, School of Physics, Konkuk University, Seoul 143-701, Korea*

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Three-step cascade decays into two invisible particles and two visible particles via two intermediate on-shell particles develop cusped peak structures in several kinematic distributions. We study their basic properties and demonstrate that the masses of the missing particles and the intermediate particles can be determined by the cusp and endpoint positions. Effects from realistic considerations such as finite decay widths, the longitudinal boost of the mother particle, the initial state radiation, and spin correlations are shown to be under control for the processes illustrated.

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**I. INTRODUCTION**

At the energy frontier, the LHC experiments are taking us to an unprecedented territory of the terascale physics beyond the Standard Model (SM). At the cosmo frontier, we have entered an era of precision cosmology. With much progress made in the two frontiers, we have to admit that our understanding of the Universe is still far from being complete. According to the precise measurements of the cosmic microwave background fluctuations, such as WMAP [1], about 95% of the components of the current Universe have never been directly observed in the laboratory. The dominant component ( $\approx 72\%$ ) is dark energy that is responsible for the accelerating expansion of the Universe [2]. The second dominant ( $\approx 23\%$ ) is cold dark matter (CDM), which is assumed to be in a form of non-relativistic matter but cannot be explained within the SM. Even in the realm of particle physics, the SM is regarded as an effective theory below a certain scale, albeit its extraordinary success in explaining current experimental data with incredibly high precision. For example, theoretical unnaturalness of the SM, dubbed as the gauge hierarchy problem, suggests new physics beyond the SM at the TeV scale. Therefore, there is a very intriguing possibility that such CDM components appear in new particle physics models.

Indeed particle physics has theoretical answers for the astrophysical question about CDM. One of the most popular scenarios is a thermal production of weakly interacting massive particles [3]. In this scenario, a stable particle  $X$  had been once in thermal equilibrium in the early history of the Universe but later got frozen out as its reaction rate became slower than the expansion rate of the Universe. The stability of the CDM particle over cosmic time is often attributed to an unbroken parity symmetry or a discrete symmetry. Under such a symmetry, the SM particle fields are in the trivial representation while new particle fields are in some nontrivial representation. The decay of the lightest new particle into SM particles is prohibited. The current

observation highly suggests that the CDM particle has its mass at the electroweak scale and its couplings with a size of weak interaction. Some popular models with weakly interacting massive particles are supersymmetric models with  $R$  parity [4], the universal extra dimension (UED) model with Kaluza-Klein (KK) parity [5], and the littlest Higgs model with  $T$  parity [6].

This weakly interacting massive particle is likely to be produced at the LHC and to be identified by missing transverse energy. The measurement of its mass is of crucial importance to reveal the identity of the CDM, but a very challenging task at the LHC because such invisible particles are produced in pairs. With large errors especially in jet energy measurements, the combinatoric complications disentangle the indirect information on the missing particle mass. In the literature, many new ideas to measure the CDM mass have been proposed [7], such as endpoint methods [8], polynomial methods [9,10],  $M_{T2}$  methods [11–14], and combined methods [15].

Recently, we have proposed a new approach to measure the missing particle mass by using the singular structures in the kinematic distributions of the antler decay [16,17]. The antler decay is a resonant decay of a parity-even particle  $D$  into a pair of the missing particles ( $X_1$  and  $X_2$ ) and a pair of SM visible particles ( $a_1$  and  $a_2$ ) through two on-shell parity-odd intermediate particles ( $B_1$  and  $B_2$ ), as depicted in Fig. 1. We have studied two kinds of singular structures, a cusp and an endpoint. The positions of cusps and endpoints determine the masses of the missing particle as well as the intermediate particle, if the mother particle mass  $m_D$  is known from other decay channels, directly into two SM particles.<sup>1</sup>

Following are a few merits of this method: (i) the positions of the cusps and endpoints are stable under the spin correlation effects since they are purely determined by the phase space; (ii) a cusp as a sharp and nonsmooth peak is

<sup>1</sup>This is possible since the particle  $D$  has even parity.

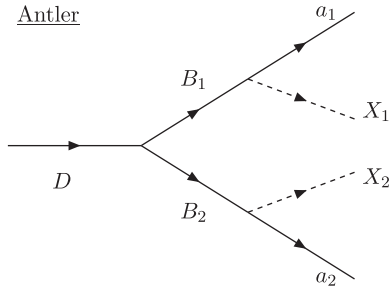


FIG. 1. The antler decay topology of a parity-even particle  $D$  into two missing particles ( $X_1$  and  $X_2$ ) and two visible particles ( $a_1$  and  $a_2$ ).

statistically more advantageous to search than an endpoint and more identifiable to observe than a kink; (iii) the simple configuration of outgoing particles can reduce combinatoric complication which is commonly troublesome in many missing particle mass measurement methods; (iv) the derived analytic functions of some kinematic distributions allow us to reconstruct the mass parameters by best fitting.

As a companion of Ref. [17], this paper focuses on another decay topology of a parity-even particle  $D$  into two visible particles and two missing particles, the cascade decay shown in Fig. 2. In this process, the mother particle  $D$  decays through three steps in series, finally ended up with a missing particle  $X_2$ . There are two nontrivial types of this decay, Type I and Type II, according to at which step the first missing particle  $X_1$  is produced. Unlike the antler decay case with one kind of intermediate particle, the cascade decay involves two different intermediate particles. We thus need to fix one more unknown mass, which requires more independent observables. The study of the basic properties of cusps and endpoints to determine all of the unknown masses is our main purpose. The cusp in the invariant mass distribution of the Type I cascade decay has been discussed in the context of new physics models with the CDM particle stabilized by  $Z_3$  symmetry [18].

The rest of the paper is organized as follows. In Sec. II, we categorize all possible kinematic variables from the four-momenta of the two visible particles. Section III deals with the Type I cascade decay. We present the expressions of cusps and endpoints of various kinematic distributions in a common case where  $m_{a_1} = m_{a_2} = 0$  and  $m_{X_1} = m_{X_2}$ . The functional form of the invariant mass distribution is also given. The general mass case is to be discussed in the

Appendix. In Sec. IV, we present the corresponding results for the Type II cascade decay. Section V is devoted to realistic considerations such as the finite widths of the intermediate particles, the longitudinal boost of the mother particle  $D$ , the initial state radiation (ISR), and the spin correlation. We then conclude in Sec. VI.

## II. KINEMATICS OF CASCADE DECAY TOPOLOGY WITH TWO MISSING PARTICLES

We consider the four-body cascade decay of a heavy particle  $D$  through three sequential steps. The cascade decay into a single missing particle and three visible particles has been extensively studied in the literature. In a supersymmetry model, a good example is the process of  $\bar{q} \rightarrow q\tilde{\chi}_2^0 \rightarrow q\ell_n\tilde{\ell} \rightarrow q\ell_n\ell_f\tilde{\chi}_1^0$ . In the UED model, we have  $Q^{(1)} \rightarrow Z^{(1)}q \rightarrow L^{(1)}\ell_nq \rightarrow B^{(1)}\ell_f\ell_nq$ . Here  $\ell_n(\ell_f)$  denotes the near (far) lepton with respect to the mother particle. In principle, three observable particles provide enough information to determine all of the unknown mass parameters [7,8]. However, there are some difficulties in extracting proper information, especially because of combinatoric complications. It is hard to distinguish  $\ell_n$  from  $\ell_f$ . Furthermore, the mother particle  $D$  is to be pair produced due to its odd parity, yielding another decay chain in the event.

Here we consider the three-step cascade decay with two missing particles. The mother particle  $D$  is of even parity and thus its single production is allowed. The final states are simply two visible particles ( $a_1$  and  $a_2$ ) with missing transverse energy. There is no combinatoric complication when forming the invariant mass of two visible particles. In addition, if the rest frame of  $D$  in the transverse direction can be determined, we can use the transverse momenta of  $a_1$  and  $a_2$  as additional information. As shall be shown below, some transverse momentum variable distributions accommodate cusps.

The cascade decays of  $D \rightarrow a_1a_2X_1X_2$  can be classified according to at which step the first missing particle, say  $X_1$ , is produced. We fix that the other missing particle  $X_2$  is from the last step. If  $X_1$  is also from the last step, the final intermediate particle  $B$  is just missing and this decay is indistinguishable from a two step cascade decay. We do not consider this case. Then, there are two nontrivial three-step cascade decays, as depicted in Fig. 2. In the Type I decay,

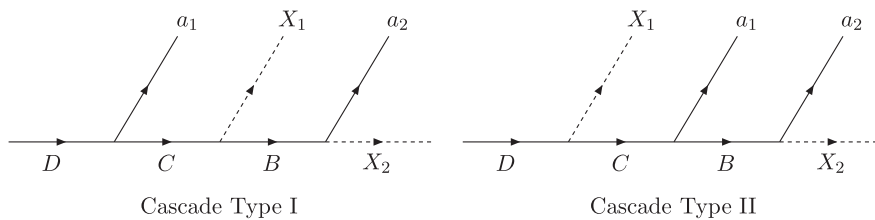


FIG. 2. The cascade decay topology of a parity-even particle  $D$  into two missing particles ( $X_1$  and  $X_2$ ) and two visible particles ( $a_1$  and  $a_2$ ).

$X_1$  is from the second step. The mother particle  $D$  decays into a visible particle  $a_1$  and a new particle  $C$ , followed by the decay of  $C$  into a missing particle  $X_1$  and a new particle  $B$ . Finally  $B$  decays into a visible particle  $a_2$  and a missing particle  $X_2$ . In the Type II decay,  $X_1$  is from the first step:  $D$  decays into  $CX_1$ , followed by  $C \rightarrow a_1B$ , and finally  $B \rightarrow a_2X_2$ . In the viewpoint of two observable particles  $a_1$  and  $a_2$ , this Type II decay is a two-step cascade decay of a new heavy particle  $C$ . As shall be shown, there is no cusp structure in Lorentz-invariant distributions.

It is useful to describe the kinematics in terms of the rapidity of a massive particle  $i$ ,

$$\eta_i^{(k)} = \frac{E_i^{(k)}}{m_i}, \quad (1)$$

where  $E_i^{(k)}$  and  $m_i$  are the energy and mass of the particle  $i$  in the rest frame of a particle (system)  $k$ . To avoid confusion, we adopt the following rapidity notations for the Type I and Type II decays:

|                   | Type I Cascade | Type II Cascade |
|-------------------|----------------|-----------------|
| rapidity notation | $\xi_i$        | $\zeta_i$       |

For the sake of simplicity, we omit the superscript specifying the reference frame when the rapidity is defined in the rest frame of its mother particle.

With the four-momenta  $k_1$  and  $k_2$  of the two observable particles  $a_1$  and  $a_2$  in the lab frame, respectively, we consider the following observables in three categories:

- (1) Lorentz invariant observables: the invariant mass of  $a_1$  and  $a_2$ ,

$$m = \sqrt{(k_1 + k_2)^2}. \quad (3)$$

- (2) Longitudinal-boost invariant observables:

- (a) the magnitude of the transverse momentum of a visible particle  $a_i$ ,

$$p_{Ti} = |\mathbf{k}_i^T|, \quad (4)$$

- (b) the magnitude of the transverse momentum of the  $a_1$ - $a_2$  system,

$$p_T = |\mathbf{k}_1^T + \mathbf{k}_2^T|, \quad (5)$$

- (c) the transverse mass of the  $a_1$ - $a_2$  system,

$$m_T = \sqrt{p_T^2 + m^2}. \quad (6)$$

- (3) Noninvariant observables:

- (a) cosine of  $\Theta_i$ , the angle of the visible particle  $a_i$  in the c.m. frame of  $a_1$  and  $a_2$ , with respect to their c.m. moving direction,

$$\cos\Theta = \frac{\mathbf{k}_1^{(aa)} \cdot \mathbf{k}^{(D)}}{|\mathbf{k}_1^{(aa)}| |\mathbf{k}^{(D)}|}. \quad (7)$$

Here the bold-faced letter denotes the three-vector momentum,  $k = k_1 + k_2$ , and the superscript  $(D)$  and  $(aa)$  denote the  $D$ -rest frame and the c.m. frame of  $a_1$  and  $a_2$ , respectively. We note that the  $\cos\Theta$  variable is observable only if the rest frame of the mother particle  $D$  is reconstructed.

As shall be shown,  $p_{Ti}$  and  $m_T$  distributions show cusp structures if the mother particle  $D$  is produced at rest in the transverse direction. These additional cusp structures are very valuable to determine all of the unknown masses. At a hadron collider, however, reconstructing the transverse rest frame of  $D$  is not feasible since strong QCD interactions always yield sizable ISR, which causes transverse kick to the mother particle  $D$ . The cusps and endpoints in the  $D$  rest frame get affected. In Sec. V, we study the ISR effects on the  $m_T$  and  $p_{Ti}$  cusps and endpoints.

The final comment in this section is on the simplifying assumption about mass parameters. In general, the involved seven particles ( $D$ ,  $C$ ,  $B$ ,  $a_1$ ,  $a_2$ ,  $X_1$ ,  $X_2$ ) may have different masses. However, the cascade decay processes in many new physics models have massless visible particles and the same kind of invisible particles. In the main text, therefore, we consider only the following case:

$$m_{a_1} = m_{a_2} = 0, \quad m_{X_1} = m_{X_2}. \quad (8)$$

The results for the most general case with seven different masses are presented in the Appendix.

### III. TYPE I CASCADE DECAY

As illustrated in Fig. 2, the Type I cascade decay is the decay of a parity-even particle  $D$  into two missing particles  $X_1$  and  $X_2$  and two visible particles  $a_1$  and  $a_2$  through

$$\begin{aligned} D(P) &\rightarrow C + a_1(k_1), & C &\rightarrow B + X_1, \\ & & B &\rightarrow a_2(k_2) + X_2. \end{aligned} \quad (9)$$

Here  $D$ ,  $C$ ,  $a_1$ , and  $a_2$  are parity even while  $B$ ,  $X_1$ , and  $X_2$  are parity odd. In order to accommodate the Type I cascade decay, we need at least two heavy parity-even particles.

One good example is found in the UED model [19]. It is based on a single flat extra dimension of size  $R$ , compactified on an  $S_1/Z_2$  orbifold. All of the SM fields propagate freely in the whole five-dimensional spacetime. Each field has an infinite number of KK excited states. Since the KK parity is conserved, the lightest KK particle with odd KK parity is stable and becomes a good candidate for the CDM. Usually the first KK mode of the  $U(1)_Y$  gauge boson  $B^{(1)}$  is the lightest KK particle [19,20]. All of the second KK states of the SM particles have even KK parity and mass of  $\sim 2/R$ . Lower limit of  $1/R \gtrsim 400$  GeV is set based on the combination of the constraints from the  $\rho$  parameter [21], the electroweak precision tests [22], the muon  $g - 2$  measurement [23], the flavor-changing neutral

currents [24], the direct searches by the D0 collaborations at the Tevatron [25], and the ATLAS and CMS collaborations [26]. The second KK modes are within the reach of the LHC. Possible Type I cascade decays are

$$Z^{(2)} \rightarrow \ell_n + L^{(2)} \rightarrow \ell_n + B^{(1)}L^{(1)} \rightarrow \ell_n + B^{(1)} + \ell_f B^{(1)}, \quad (10)$$

$$g^{(2)} \rightarrow q_n + q^{(2)} \rightarrow q_n + B^{(1)}q^{(1)} \rightarrow q_n + B^{(1)} + q_f B^{(1)}. \quad (11)$$

Now we present the cusps and endpoints of  $m$ ,  $m_T$ ,  $p_T$ ,  $p_{Ti}$ , and  $\cos\Theta$  distributions in terms of the masses. For the simple case in Eq. (8), there are two independent rapidities,  $\xi_B$  and  $\xi_C$ , given by

$$\begin{aligned} \cosh\xi_B &= \frac{m_C}{2m_B} \left( 1 + \frac{m_B^2}{m_C^2} - \frac{m_X^2}{m_C^2} \right), \\ \cosh\xi_C &= \frac{m_D}{2m_C} \left( 1 + \frac{m_C^2}{m_D^2} \right). \end{aligned} \quad (12)$$

We will also use  $E_n$  and  $E_f$ , the energy of the near visible particle  $a_1$  and the far visible particle  $a_2$  in its mother's rest frame, respectively:

$$E_n = \frac{m_D}{2} \left( 1 - \frac{m_C^2}{m_D^2} \right), \quad E_f = \frac{m_B}{2} \left( 1 - \frac{m_X^2}{m_B^2} \right). \quad (13)$$

For illustration, we take three sets for the mass parameters in Table I. The Mass-A<sub>1</sub> is motivated by the  $Z^{(2)}$  decay in Eq. (10). The KK masses are determined by the UED model parameters of  $\Lambda R = 20$  and  $1/R = 500$  GeV, where  $\Lambda$  is the cutoff scale [19]. Almost equal spacing in the KK mass spectrum leads to very degenerate masses, i.e.,  $m_D \approx m_C \approx 2m_B \approx 2m_X$ . The Mass-B<sub>1</sub> case has substantial mass gaps for each pair of adjacent masses. Finally the Mass-C<sub>1</sub> case has a sizable mass gap between  $m_C$  and  $m_B + m_X$ .

For the precise mass measurement using the singularities, it is necessary to have visible cusp and endpoint. A sharp cusp is considered as a visible one. The visibility of an endpoint is determined by the functional behavior near the endpoint, either fast dropping or long-tailed. We take the former as a visible endpoint.

TABLE I. Test mass spectrum sets for the Type I cascade decay. All of the masses are in units of GeV.

|                     | $m_D$  | $m_C$ | $m_B$ | $m_X$ | $\xi_B$ |
|---------------------|--------|-------|-------|-------|---------|
| Mass-A <sub>1</sub> | 1045.7 | 1023  | 514.2 | 500.9 | 0.12    |
| Mass-B <sub>1</sub> | 600    | 400   | 200   | 100   | 0.60    |
| Mass-C <sub>1</sub> | 600    | 500   | 150   | 100   | 1.16    |

## A. Invariant mass $m$ distribution

We first discuss the distribution of the invariant mass  $m$  of two visible particles. The differential decay rate  $d\Gamma/dm$  is

$$\frac{d\Gamma}{dm} \propto \begin{cases} 2\xi_B m, & \text{for } 0 < m < m_{\text{cas1}}^{\text{cusp}}, \\ m \ln \frac{(m_{\text{cas1}}^{\text{max}})^2}{m^2}, & \text{for } m_{\text{cas1}}^{\text{cusp}} < m < m_{\text{cas1}}^{\text{max}}, \end{cases} \quad (14)$$

where the cusp and endpoint are

$$(m_{\text{cas1}}^{\text{cusp}})^2 = 4E_n E_f e^{\xi_C - \xi_B}, \quad (m_{\text{cas1}}^{\text{max}})^2 = 4E_n E_f e^{\xi_C + \xi_B}. \quad (15)$$

Note that the functional behavior of  $d\Gamma/dm$  is the same as that of the antler decay [17]. The general case with seven different masses is discussed in the Appendix.

The degree of the sharpness of the  $m$  cusp is deduced from Eq. (14). The  $d\Gamma/dm$  function is linear in  $m$  for  $m < m_{\text{cas1}}^{\text{cusp}}$  and a concave function for  $m_{\text{cas1}}^{\text{cusp}} < m < m_{\text{cas1}}^{\text{max}}$ . At  $m = m_{\text{cas1}}^{\text{max}}/e$ , the concave function reaches its maximum. If  $m_{\text{cas1}}^{\text{max}}/e < m_{\text{cas1}}^{\text{cusp}}$ , which is equivalent to  $\xi_B < 1$ , the cusp can be considered to be pronounced.

In Fig. 3, we show the normalized differential decay rate  $d\Gamma/dm$  for the three mass parameter sets in Table I. In order to compare the shapes of cusps only, we present it as a function of  $m/m_{\text{cas1}}^{\text{max}}$ . The vertical lines denote the positions of  $m_{\text{cas1}}^{\text{cusp}}$  in units of  $m_{\text{cas1}}^{\text{max}}$ . The Mass-A<sub>1</sub> case with  $\xi_B = 0.12$  has a very sharp  $m$  cusp. The Mass-B<sub>1</sub> case with  $\xi_B = 0.60$  shows a triangular shape with a cusped peak. However, the Mass-C<sub>1</sub> case with  $\xi_B = 1.16$  has a dull cusp. The endpoints for all of three cases are fast dropping, as suggested by the concave function in Eq. (14).

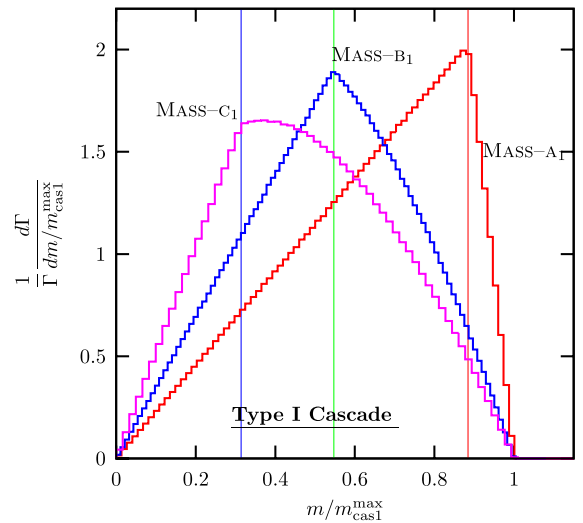


FIG. 3 (color online). The normalized differential decay rate of the invariant mass of two visible particles,  $\frac{d\Gamma}{dm}$  for the Type I cascade decay. The masses are in Table I.

### B. Transverse mass $m_T$ distribution

Figure 4 shows the normalized differential decay rate  $d\Gamma/dm_T$ , which is defined in the rest frame of  $D$ . For all of three cases, the  $m_T$  distributions show visible cusp structures. Even the Mass-C<sub>1</sub> case, which has a dull cusp in the  $m$  distribution, has a quite sharp cusp. This is contrasted to the antler decay case where there is no cusp in the  $m_T$  distribution [17]. As shall be shown in the next section, the Type II cascade decay also has a cusp in the  $m_T$  distribution. Therefore the presence of the  $m_T$  cusp signals the cascade decay topology. The cusp and maximum positions in terms of the masses are

$$\begin{aligned} (m_T)_{\text{cas1}}^{\text{cusp}} &= E_n + E_f e^{\xi_c - \xi_B}, \\ (m_T)_{\text{cas1}}^{\text{max}} &= E_n + E_f e^{\xi_c + \xi_B}, \end{aligned} \quad (16)$$

where  $E_n$  and  $E_f$  are in Eq. (13).

### C. The system $p_T$ distribution

Figure 5 shows the normalized distribution of the transverse momentum  $p_T$  of two visible particle system. For all of the three mass spectra in Table I, the  $p_T$  distribution has smooth peak without a cusp structure. The endpoint of  $p_T$  distribution is

$$(p_T)_{\text{cas1}}^{\text{max}} = E_n + E_f e^{-\xi_c + \xi_B}. \quad (17)$$

Only the Mass-A<sub>1</sub> case has the endpoint of fast-dropping shape, which is attributed to very small momentum transfer to the visible particles. More general cases of Mass-B<sub>1</sub> and Mass-C<sub>1</sub> have long-tailed endpoints. The  $p_T$  distribution is not useful for the mass measurement.

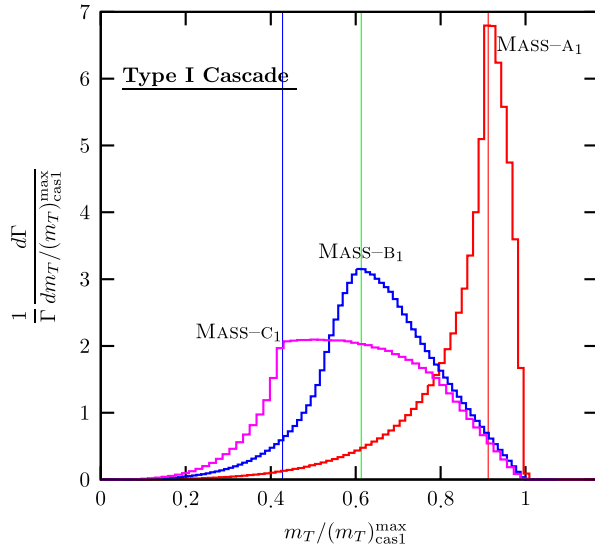


FIG. 4 (color online). The normalized differential decay rate of the transverse mass of two visible particles,  $\frac{d\Gamma}{dm_T}$  for the Type I cascade decay. The masses are in Table I.

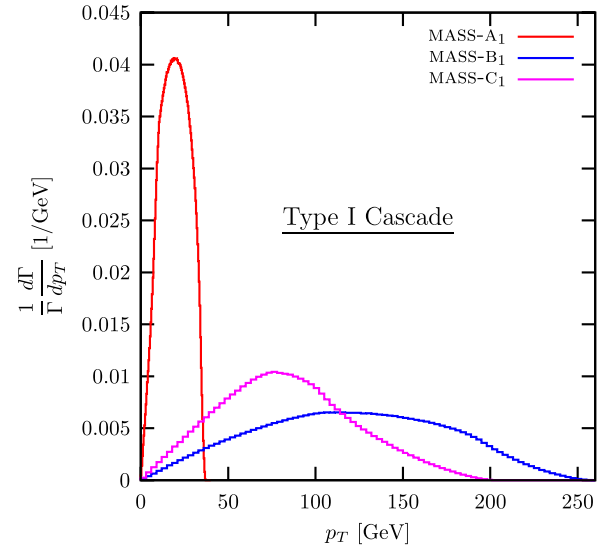


FIG. 5 (color online). The normalized differential decay rate of the transverse momentum of two visible particles,  $\frac{d\Gamma}{dp_T}$  for the Type I cascade decay for the masses in Table I.

### D. Single particle $p_{Ti}$ distribution

We show the individual transverse momentum  $p_{Ti}$  distributions in Fig. 6. The thin solid line labeled by “near” (“far”) is the  $p_{Ti}$  distribution of the near visible particle  $a_1$  (the far visible particle  $a_2$ ). The  $p_{Tf}$  distribution has both the cusp and the endpoint structures. On the contrary, the  $p_{Tn}$  distribution has only an endpoint of a suddenly ending shape, which holds true for all of the mass cases. The positions of the cusp and the endpoint in the  $p_{Tf}$  distribution are given by

$$(p_{Tf})_{\text{cas1}}^{\text{cusp}} = E_f e^{\xi_c - \xi_B}, \quad (p_{Tf})_{\text{cas1}}^{\text{max}} = E_f e^{\xi_c + \xi_B}, \quad (18)$$

and the endpoint in the  $p_{Tn}$  distribution is located at

$$(p_{Tn})_{\text{cas1}}^{\text{max}} = E_n. \quad (19)$$

However, we cannot distinguish  $a_1$  from  $a_2$  event by event. Therefore we consider a more practical observable, the sum of two  $p_{Ti}$  distributions. The thick lines in Fig. 6 represent the sum. Depending on whether  $(p_{Tn})_{\text{cas1}}^{\text{max}} > (p_{Tf})_{\text{cas1}}^{\text{max}}$  or not, the summed distribution shows very different shape. For the former case as in Mass-A<sub>1</sub>, the spiky  $(p_{Tn})_{\text{cas1}}^{\text{max}}$  stands outside the  $p_{Tf}$  distribution, which reveals the  $p_{Tf}$  cusp and endpoint. For the later case as in Mass-B<sub>1</sub> and Mass-C<sub>1</sub>, the  $(p_{Tn})_{\text{cas1}}^{\text{max}}$  peak lies in the middle of  $p_{Tf}$  distribution. The  $p_{Tf}$  cusp gets distorted and thus barely visible.

### E. $\cos\Theta$ distribution

The variable  $\cos\Theta$  in Eq. (7) is defined by the angle of one visible particle. We have two  $\cos\Theta$  distributions for  $a_1$  and  $a_2$ , which cannot be distinguished. In Fig. 7, therefore,

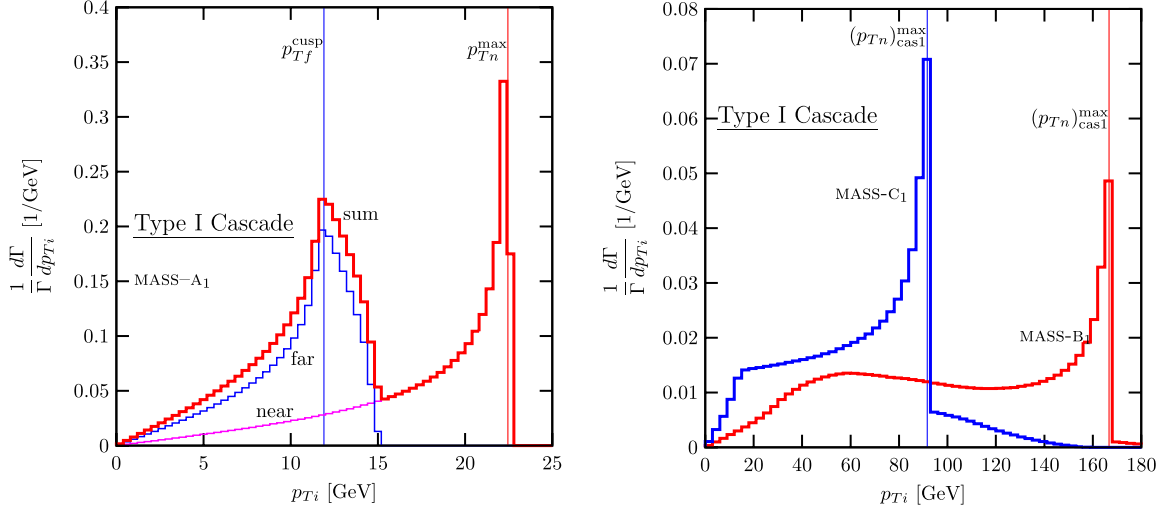


FIG. 6 (color online). The normalized differential decay rate of the transverse momentum of one visible particle,  $\frac{d\Gamma}{\Gamma dp_{Ti}}$  for the Type I cascade decay. In the Mass-A<sub>1</sub> case, the line labeled by “near” (“far”) denotes the  $p_{Ti}$  distribution of  $a_1$  ( $a_2$ ). Thick lines are the summed distributions of  $p_{Ti}$ .

we present the summation of two  $\cos\Theta_i$  distributions in the rest frame of  $D$  for Mass-A<sub>1</sub> and Mass-B<sub>1</sub> cases. It is symmetric about  $\cos\Theta = 0$  and has two cusp structures,  $\cos\Theta_{cas1}^{cusp1}$  and  $\cos\Theta_{cas1}^{cusp2}$ , marked by the vertical arrows. In terms of masses, they are

$$\begin{aligned} \cos\Theta_{cas1}^{cusp1} &= \frac{E_n - E_f \exp(\xi_B - \xi_C)}{E_n + E_f \exp(\xi_B - \xi_C)}, \\ \cos\Theta_{cas1}^{cusp2} &= \frac{E_n - E_f \exp(-\xi_B - \xi_C)}{E_n + E_f \exp(-\xi_B - \xi_C)}. \end{aligned} \quad (20)$$

In the Mass-A<sub>1</sub> case,  $\cos\Theta_{cas1}^{cusp1}$  stands on a steep slope, which is difficult to probe. The Mass-B<sub>1</sub> case shows two pronounced cusps.

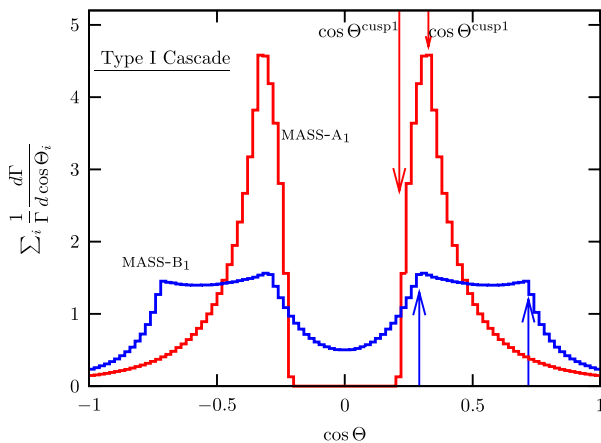


FIG. 7 (color online). The summed distributions of  $\cos\Theta_i$  in the Type I cascade decay for the masses in Table I.

#### IV. TYPE II CASCADE DECAY

Type II cascade decay is a chain decay of

$$\begin{aligned} D(P) &\rightarrow C + X_1, & C &\rightarrow B + a_1(k_1), \\ B &\rightarrow a_2(k_2) + X_2. \end{aligned} \quad (21)$$

A good example can be found in the minimal supersymmetric standard model,

$$H/A \rightarrow \tilde{\chi}_1^0 + \tilde{\chi}_2^0, \quad \tilde{\chi}_2^0 \rightarrow \ell_n + \tilde{\ell}, \quad \tilde{\ell} \rightarrow \ell_f + \tilde{\chi}_1^0. \quad (22)$$

As in the Type I cascade decay, we restrict ourselves to the realistic cascade decay with  $m_{a_1} = m_{a_2} = 0$  and  $m_{X_1} = m_{X_2}$ . Then there are two independent rapidities,  $\zeta_C$  and  $\zeta_B$ ,

$$\begin{aligned} \cosh\zeta_B &= \frac{m_C}{2m_B} \left( 1 + \frac{m_B^2}{m_C^2} \right), \\ \cosh\zeta_C &= \frac{m_D}{2m_C} \left( 1 + \frac{m_C^2}{m_D^2} - \frac{m_X^2}{m_D^2} \right). \end{aligned} \quad (23)$$

For illustration, we consider three mass sets for the Type II cascade decay in Table II.

##### A. Invariant mass $m$ distribution

We first study the distribution of the invariant mass of  $a_1$  and  $a_2$ . Note that in the viewpoint of  $a_1$  and  $a_2$ , this Type II cascade decay is a three-body decay of the mother particle  $C$ . The presence of the invisible  $X_1$  decayed from  $D$  does not change any Lorentz invariant result. The  $m$  distribution is the same as that of, i.e.,  $m_{\ell\ell}$  of the decay  $\tilde{\chi}_2^0 \rightarrow \ell_n \tilde{\ell} \rightarrow \ell_n \ell_f \tilde{\chi}_1^0$  in the minimal supersymmetric standard model. This  $m_{\ell\ell}$  distribution is well known to have no cusp structure. The endpoint is [27],

TABLE II. Test mass spectrum sets for the Type II cascade decay. All of the masses are in units of GeV.

|                     | $m_D$ | $m_C$ | $m_B$ | $m_X$ | $m^{\max}$ |
|---------------------|-------|-------|-------|-------|------------|
| Mass-A <sub>2</sub> | 614   | 299   | 222   | 161   | 138.0      |
| Mass-B <sub>2</sub> | 600   | 300   | 200   | 100   | 193.6      |
| Mass-C <sub>2</sub> | 400   | 250   | 150   | 120   | 120.0      |

$$(m_{\text{cas2}}^{\max})^2 = m_C^2 \left(1 - \frac{m_B^2}{m_C^2}\right) \left(1 - \frac{m_X^2}{m_B^2}\right). \quad (24)$$

In Fig. 8, we show the  $m$  distribution for three sets of the mass parameters in Table II, all of which have right-angled triangle shapes without a cusp.

The absence of a cusp in a two-step cascade decay can be understood by a simple kinematic configuration. For the antler decay ( $D \rightarrow B_1 + B_2 \rightarrow a_1 X_1 + a_2 X_2$ ) in the massless visible particle case ( $m_{a_1} = m_{a_2} = 0$ ), the following four critical points correspond to a kinematic singular structure [17]:

| 1D configuration   | $m_{a_1 a_2}$  |      |
|--|--|------|
| $\begin{array}{c} \leftarrow a_2 \\ \leftarrow B_2 \\ \leftarrow D \\ \leftarrow B_1 \\ \leftarrow a_1 \end{array}$      | $\begin{array}{c} \bullet D \\ \bullet B_1 \\ \bullet a_1 \end{array}$ | max  |
| $\begin{array}{c} \Rightarrow a_2 \\ \Rightarrow B_2 \\ \Rightarrow D \\ \Rightarrow B_1 \\ \Rightarrow a_1 \end{array}$ | $\begin{array}{c} \bullet D \\ \bullet B_1 \\ \bullet a_1 \end{array}$ | cusp |
| $\begin{array}{c} \leftarrow a_2 \\ \leftarrow B_2 \\ \leftarrow D \\ \rightarrow B_1 \\ \rightarrow a_1 \end{array}$    | $\begin{array}{c} \bullet D \\ \bullet B_1 \\ \bullet a_1 \end{array}$ | min  |
| $\begin{array}{c} \rightarrow a_2 \\ \rightarrow B_2 \\ \rightarrow D \\ \rightarrow B_1 \\ \rightarrow a_1 \end{array}$ | $\begin{array}{c} \bullet D \\ \bullet B_1 \\ \bullet a_1 \end{array}$ | min  |

(25)

Here we simplify the picture as a one-dimensional case. It is clear to see that  $m_{a_1 a_2}^{\min}$  happens when two observable particles move in the same direction, while one of two kinematic configurations of back-to-back moving visible particles corresponds to either  $m_{a_1 a_2}^{\max}$  or  $m_{a_1 a_2}^{\text{cusp}}$ . For a

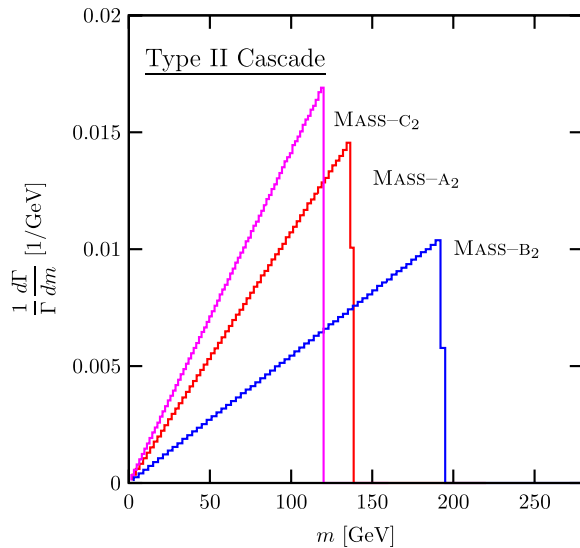


FIG. 8 (color online). The normalized differential decay rate of the invariant mass of two visible particles,  $\frac{d\Gamma}{\Gamma dm}$  for the Type II cascade decay. The mass spectrum sets are described in Table II.

two-step cascade decay ( $C \rightarrow a_1 + B \rightarrow a_1 + a_2 X_2$ ),  $a_1$  and  $a_2$  in one-dimensional space have only two independent kinematic configurations, moving in the same direction and moving in the opposite direction. The former corresponds to the minimum  $m$ , while the latter to the maximum  $m$ . There is no critical point left for the cusp.

### B. Transverse mass $m_T$ distribution

Unlike the invariant mass distribution, the  $m_T$  distribution contains the information about the transverse momentum of the first missing particle  $X_1$ . As shown in Fig. 9, there is a cusp here. We stress once again that this  $m_T$  cusp appears only when  $D$  is produced at rest in the transverse direction.

Another interesting feature is that the position of the  $m_T$  cusp is nothing but the  $m$  maximum,

$$(m_T)_{\text{cas2}}^{\text{cusp}} = m_{\text{cas2}}^{\max}. \quad (26)$$

This nontrivial equality is a unique feature of the Type II cascade decay.

### C. System $p_T$ distribution

The total  $p_T$  distributions for the Type II cascade decay are shown in Fig. 10. All of the three mass sets have smooth  $p_T$  distributions, and their endpoints are all long tailed. This feature is common for the antler, Type I, and Type II cascade decay topology.

### D. Single particle $p_{T_i}$ distribution

Figure 11 shows the distribution of the individual transverse momentum of the near  $a_1$  and the far  $a_2$ . The near  $p_{T_n}$  distribution has a sharp cusp and a fast dropping

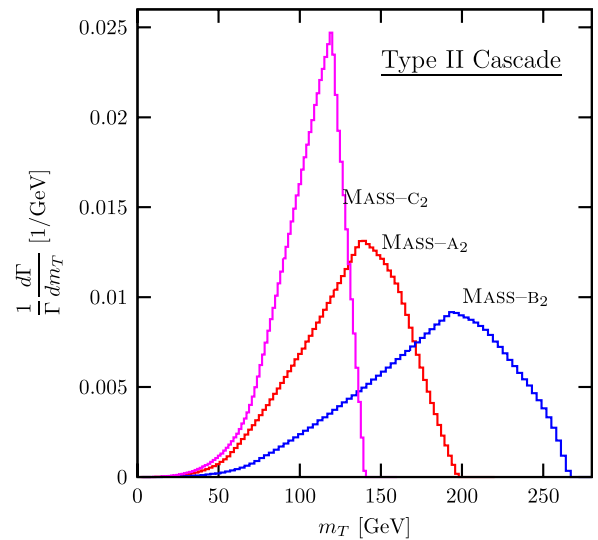


FIG. 9 (color online). The normalized differential decay rate of the transverse mass of two visible particles,  $\frac{d\Gamma}{\Gamma dm_T}$  for the Type II cascade decay. The mass spectrum sets are described in Table II.

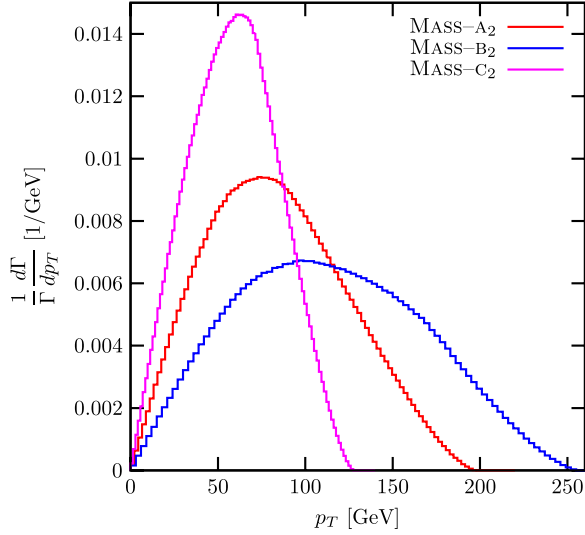


FIG. 10 (color online). The normalized differential decay rate of the transverse momentum of two visible particles,  $\frac{d\Gamma}{\Gamma dp_T}$  for the Type II cascade decay. The mass spectrum sets are in Table II.

endpoint. The  $p_{Tf}$  distribution has a long-tailed endpoint without any cusp. In terms of masses, they are simply

$$\begin{aligned} (p_{Tn})_{\text{cas2}}^{\text{cusp}} &= \frac{m_C}{2} \left(1 - \frac{m_B^2}{m_C^2}\right) e^{-\zeta_C}, \\ (p_{Tn})_{\text{cas2}}^{\text{max}} &= \frac{m_C}{2} \left(1 - \frac{m_B^2}{m_C^2}\right) e^{\zeta_C}. \end{aligned} \quad (27)$$

Note that the product of  $(p_{Tn})_{\text{cas2}}^{\text{cusp}}$  and  $(p_{Tn})_{\text{cas2}}^{\text{max}}$  removes the  $\zeta_C$  dependence, which depends on the intermediate masses  $m_C$  and  $m_B$ . In addition the ratio  $(p_{Tn})_{\text{cas2}}^{\text{cusp}}/(p_{Tn})_{\text{cas2}}^{\text{max}}$  depends only on the rapidity  $\zeta_C$ .

As discussed before, the individual  $p_{Ti}$  distribution cannot be constructed. Instead we show the sum of two

distributions in Fig. 12. For the Mass-A<sub>2</sub> case, the cusp in the  $p_{Tn}$  distribution and the smooth peak of the  $p_{Tf}$  distribution are located nearby. In their sum, the  $p_{Tn}$  cusp survives over the relatively round  $p_{Tf}$  peak and the fast dropping  $p_{Tn}$  endpoint is also measurable. For the Mass-C<sub>2</sub> case, however, the  $p_{Tn}$  cusp and the  $p_{Tf}$  peak are separated so that the summed distribution shows both. With finite number of data, it would be difficult to distinguish the  $p_{Tn}$  cusp from the  $p_{Tf}$  peak.

### E. The $\cos\Theta$ distribution

We consider the  $\cos\Theta$  distribution for the Type II cascade decay. In Fig. 13 we show the normalized differential decay rate  $d\Gamma/d\cos\Theta$  for the near and far visible particles (denoted by thin lines) as well as their sum (thick lines) for the Mass-A<sub>2</sub> and Mass-C<sub>2</sub>. In both cases, the summed distribution of  $\cos\Theta_i$  is symmetric about  $\cos\Theta = 0$  and has one sharp cusp denoted by vertical lines in Fig. 13. The  $\cos\Theta$  cusp position in terms of the mass parameters is

$$\cos\Theta_{\text{cas2}}^{\text{cusp}} = \frac{m_C \left(1 - \frac{m_B^2}{m_C^2}\right) - m_B \left(1 - \frac{m_X^2}{m_B^2}\right) e^{-\zeta_B}}{m_C \left(1 - \frac{m_B^2}{m_C^2}\right) + m_B \left(1 - \frac{m_X^2}{m_B^2}\right) e^{-\zeta_B}}. \quad (28)$$

### F. Mass determination from the cusps and endpoints

Unlike the antler decay with one kind of intermediate particles, the cascade decay has two different intermediate particles. In addition, the Type II decay has fewer independent observables of cusps and endpoints: there is no  $m$  cusp structure, and the  $m_T$  cusp position is the same as the  $m$  endpoint. A concern arises whether we have enough information to determine all of the masses, especially at the LHC where the  $\cos\Theta$  cusp cannot be used. We show that

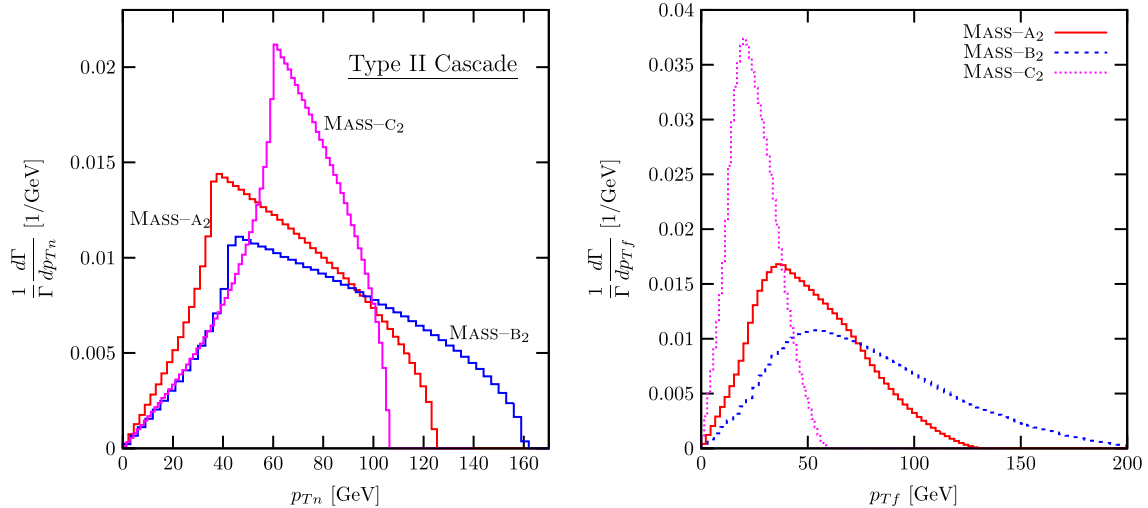


FIG. 11 (color online). The normalized differential decay rate of the transverse mass of one visible particle,  $\frac{d\Gamma}{\Gamma dp_{Ti}}$  for the Type II cascade decay. The left figure is for the near visible particle  $a_1$ , and the right one is for the far visible particle  $a_2$ .



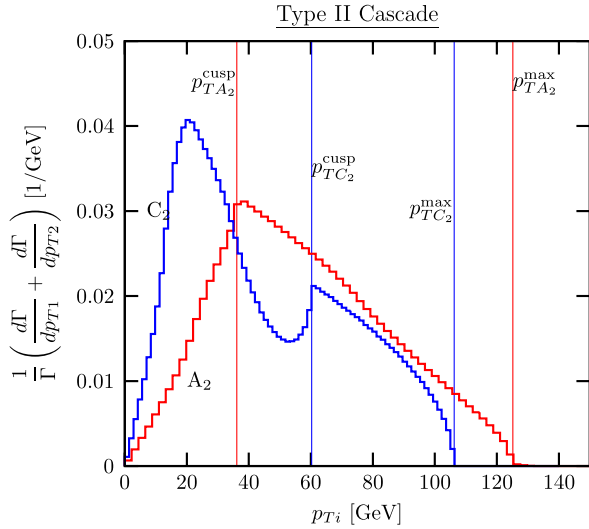


FIG. 12 (color online). The sum of two normalized differential decay rate with respect to the individual transverse momenta of the near and far visible particles.

three unknown masses ( $m_C$ ,  $m_B$ , and  $m_X$ ) are unambiguously determined by three singularities of  $m_{\text{cas}2}^{\text{max}}$ ,  $(p_{Tn})_{\text{cas}2}^{\text{cusp}}$ , and  $(p_{Tn})_{\text{cas}2}^{\text{max}}$ , as

$$m_C = R_\alpha m_D, \quad m_B = \sqrt{1 - \frac{\alpha_1}{R_\alpha} m_C}, \quad (29)$$

$$m_X = \sqrt{1 - \frac{\alpha_2}{R_\alpha} m_B},$$

where  $R_\alpha$  is

$$R_\alpha = \frac{1 + \alpha_1 \alpha_2}{\alpha_3 - \alpha_1 - \alpha_2}, \quad (30)$$

and  $\alpha_{1,2,3}$  are

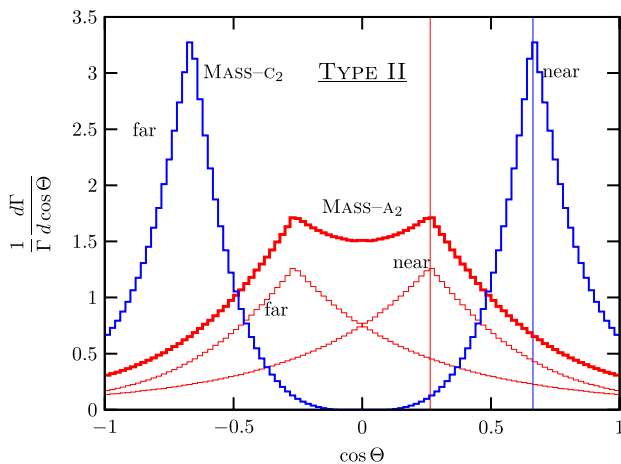


FIG. 13 (color online). The sum of  $d\Gamma/d\cos\Theta_i$  for the Type II cascade decay.

$$\alpha_1 = \frac{(m_{\text{cas}2}^{\text{max}})^2}{2m_D \sqrt{(p_{Tn})_{\text{cas}2}^{\text{max}} / (p_{Tn})_{\text{cas}2}^{\text{cusp}}}},$$

$$\alpha_2 = \frac{2\sqrt{(p_{Tn})_{\text{cas}2}^{\text{max}} / (p_{Tn})_{\text{cas}2}^{\text{cusp}}}}{m_D}, \quad (31)$$

$$\alpha_3 = \sqrt{\frac{(p_{Tn})_{\text{cas}2}^{\text{max}}}{(p_{Tn})_{\text{cas}2}^{\text{cusp}}}} + \sqrt{\frac{(p_{Tn})_{\text{cas}2}^{\text{cusp}}}{(p_{Tn})_{\text{cas}2}^{\text{max}}}}.$$

## V. EFFECTS OF REALISTIC CONSIDERATIONS

All of the previous expressions of the cusps and endpoints have been derived in an idealistic situation: the total decay widths of decaying particles are ignored; the  $D$  rest frame is assumed to be reconstructed; the ISR effects are neglected; and the spin-correlation effects from the full matrix elements are negligible. In this section, we investigate these effects on the position and shape of the kinematic cusp and endpoint.

### A. Finite width effects

Up to now we have applied the narrow width approximation, ignoring the width of decaying particles. Since the effect of finite  $\Gamma_D$  is shown to be very minor in Ref. [16], we focus on the effects of  $\Gamma_B$  and  $\Gamma_C$ .

We find that the mass spectrum is the most crucial factor to determine the stability of the cusp and endpoint structures under the width effects. Out of six cases in Tables I and II, the Mass-A<sub>1</sub> has very vulnerable structures. This case is special because of its degenerate masses: the observable particles have very small momentum transfer and their kinematic phase space is highly limited.

In Figs. 14 and 15 we show, for the Mass-A<sub>1</sub> case, the finite width effects on the  $m$ ,  $m_T$ ,  $\sum p_{Ti}$ , and  $\cos\Theta$  distributions.

We present four cases for  $\Gamma_B$  and  $\Gamma_C$ : on-shell (solid line),  $\Gamma/M = 0.01$  (long dashed line),  $\Gamma/M = 0.1$  (short dashed line), and  $\Gamma/M = 0.5$  (dotted line). Here  $\Gamma/M \equiv \Gamma_B/m_B = \Gamma_C/m_C$  for simplicity. Just one percent of  $\Gamma/M$  destroys all of the sharp cusp structures into smooth peaks. In addition, the positions of the peaks are shifted significantly from the true cusp positions. For  $\Gamma/M \geq 0.1$  the summed  $p_{Ti}$  and  $\cos\Theta_i$  distributions lose their functional characteristics completely, leaving very smooth and featureless distributions. In summary, sizable widths like  $\Gamma/M \geq 0.01$  smash the cusps.

The fast-falling endpoints in the  $m$ ,  $m_T$ , and  $p_{Ti}$  distributions are also smeared out considerably. The degree of its shifting is large even for  $\Gamma/M = 1\%$ . One interesting observation is that two shifted endpoints of the  $m$  and  $m_T$  distributions are the same as  $m_D - 2m_X$ , denoted by vertical arrows. This new endpoint is from the kinematic configuration where two visible particles' momenta span all of the phase space determined by  $m_D$  and  $m_X$ . Even

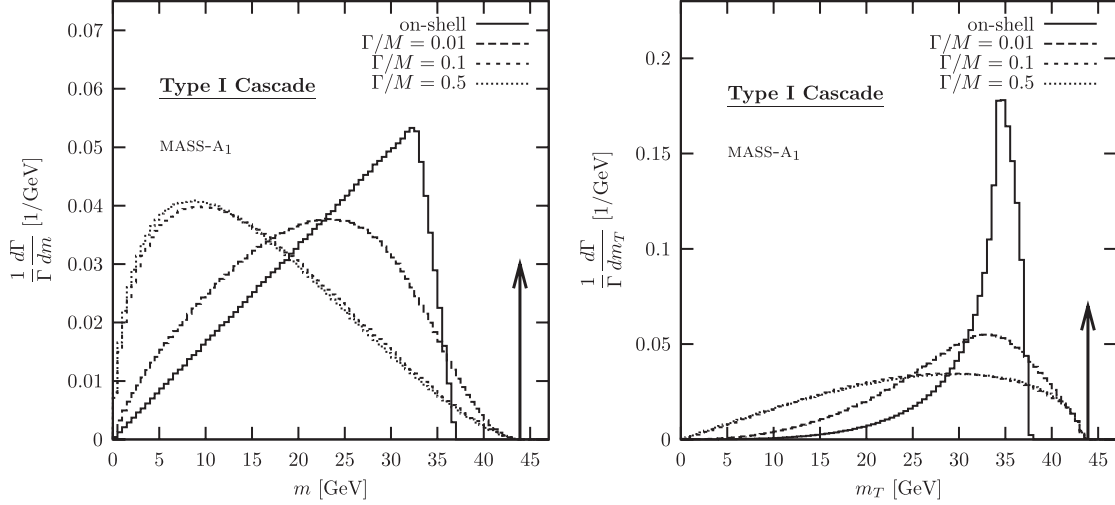


FIG. 14. The finite decay width effects on the  $m$  and  $m_T$  distributions in the Mass- $A_1$  case. Solid lines are for the on-shell decay, the long dashed lines for  $\Gamma/M = 0.01$ , the short dashed lines for  $\Gamma/M = 0.1$ , and the dotted lines for  $\Gamma/M = 0.5$ . Here  $\Gamma/M \equiv \Gamma_B/M_B = \Gamma_C/M_C$ .

though we do not know the intermediate particle masses, the missing particle mass  $m_X$  can be read off. For this information, the  $m_T$  distribution is more advantageous than the  $m$  distribution because of its fast falling shape.

In a realistic new physics process, however, this Mass- $A_1$  case does not allow even one percent of  $\Gamma/M$ . For example, the  $Z^{(2)}$  decay in the minimal UED model has the decay widths of

$$\begin{aligned} \Gamma_D = \Gamma_{Z^{(2)}} &\simeq 270 \text{ MeV}, & \Gamma_C = \Gamma_{L^{(2)}} &\simeq 5 \text{ MeV}, \\ \Gamma_B = \Gamma_{L^{(1)}} &\simeq 1 \text{ MeV}, \end{aligned} \quad (32)$$

which lead to  $\Gamma/M \sim 10^{-5}$ . In summary, the extreme Mass- $A_1$  case has generically negligible width effects. The cusp and endpoint structures are reserved.

We consider more general mass parameters, Mass- $B_1$  for the Type I and Mass- $C_2$  for the Type II cascade decay. First we examine the finite width effects on the invariant mass distributions in Fig. 16. These cases show more stable cusp and endpoint structures from the finite width effects. For  $\Gamma/M = 1\%$ , the  $m$  distributions in both Type I and Type II decays do not change, keeping the same cusp and endpoint structures. For 10% of  $\Gamma/M$ , the  $m$  cusp of the Type I decay retains its position, though losing its sharpness. The  $m$  endpoints in both Type I and Type II decays are shifted into the new position  $m_D - 2m_X$ . If  $\Gamma/M = 50\%$ , the Type I decay does not retain the shape and position of the  $m$  cusp, and the Type II decay loses the right-angled triangle shape of the  $m$  distribution. Both cases have the same new endpoint at  $m_D - 2m_X$ , which

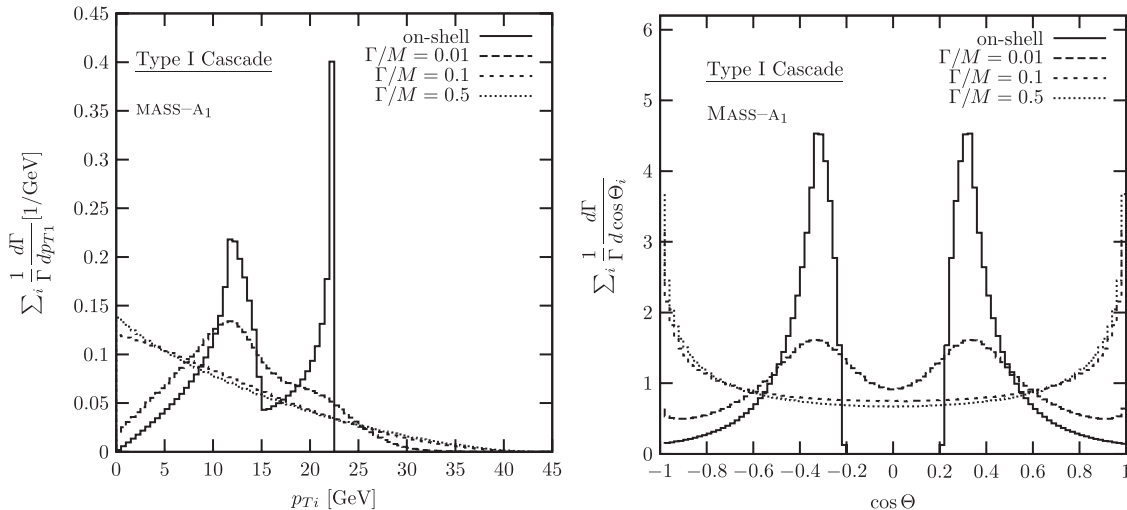


FIG. 15. The finite decay width effects on the summed distributions of  $p_{Ti}$  and  $\cos\Theta_i$  in the Mass- $A_1$  case of the Type I cascade decay. As before, we take  $\Gamma/M = 0, 0.01, 0.1, 0.5$ .

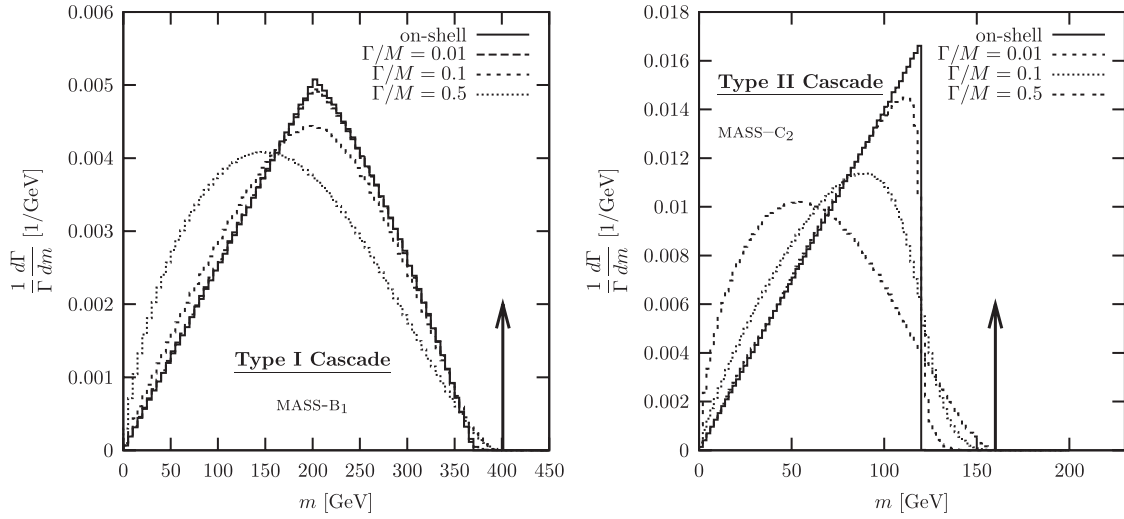


FIG. 16. Finite width effects on the normalized  $m$  distribution. We take the Mass-B<sub>1</sub> case for Type I decay and the Mass-C<sub>2</sub> for Type II decay. As before, we take  $\Gamma/M = 0, 0.01, 0.1, 0.5$ .

is also valuable information for the missing particle mass measurement.

In Fig. 17, we show the width effects on the  $m_T$  distributions. The  $m_T$  cusp structures are more stable than the  $m$  cusps in both Type I and Type II decays. For  $\Gamma/M = 1\%$ , the changes in the distribution are unnoticeable. For  $\Gamma/M \geq 10\%$ , we start to lose the sharpness of the cusps but still keep the positions for the cusp in both cases. If  $\Gamma/M = 50\%$ , the cusped peaks become dull further with relatively stable positions, and the  $m_T$  endpoints are shifted into  $m_D - 2m_X$ .

Figures 18 and 19 show the width effects on the summed distributions of  $p_{Ti}$  and  $\cos\Theta_i$ , respectively. Both distributions preserve the cusp structure for  $\Gamma/M = 1\%$ . If  $\Gamma/M \geq 10\%$ , however, the finite width effects almost smear the cusp and endpoint structures.

## B. Longitudinal boost effect

In hadronic collisions, the longitudinal motion of the particle  $D$  is not determined. This longitudinal ambiguity affects the kinematic variable  $\cos\Theta$ , which is defined in the  $D$  rest frame. In order to see the longitudinal boost effects, we convert the  $\cos\Theta$  distribution in the  $D$  rest frame into the  $pp$  frame at the LHC, by convoluting with the parton distribution functions of a proton. In Fig. 20, we compare the summed distributions of  $\cos\Theta_i$  in the  $D$  rest frame (thin curves) with that in the  $pp$  lab frame at  $\sqrt{s} = 14$  TeV (thick curves). For the parton distribution function, we have used CTEQ6 [28]. We take the Mass-A<sub>1</sub> for Type I and the Mass-C<sub>2</sub> for Type II decay. For simplicity we assume that the heavy particle  $D$  is singly produced through the  $s$ -channel gluon fusion.

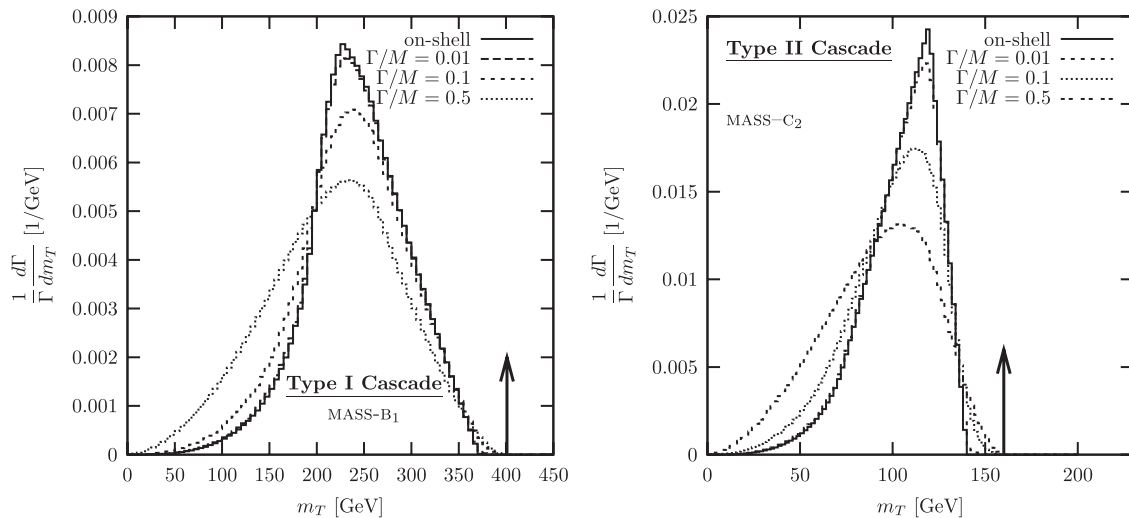


FIG. 17. The width effects on the normalized  $d\Gamma/dm_T$  for the Type I Mass-B<sub>1</sub> and Type II Mass-C<sub>2</sub> cascade decays. As before, we take  $\Gamma/M = 0, 0.01, 0.1, 0.5$ .

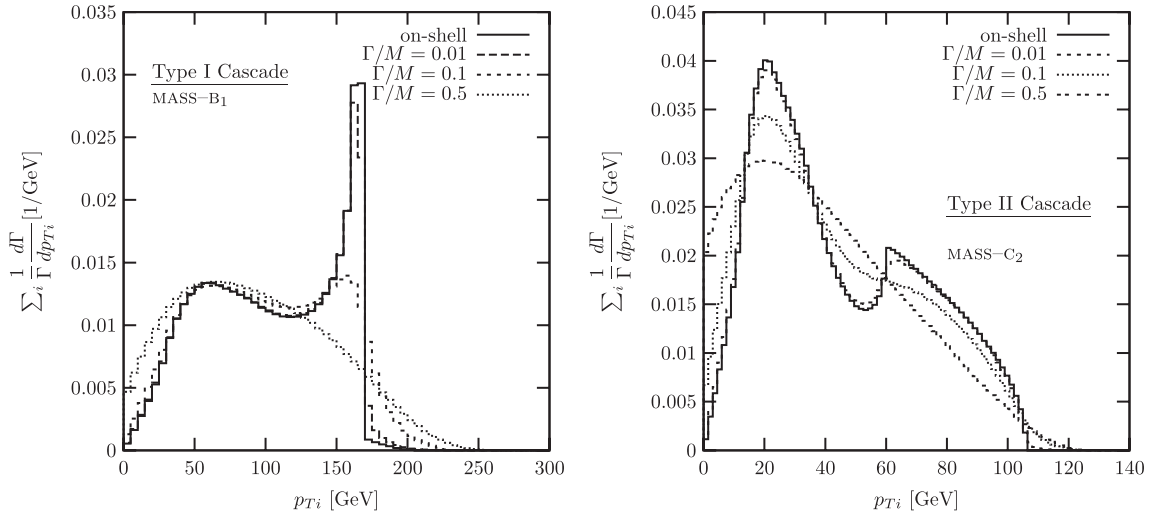


FIG. 18. The width effects on the summed distributions of  $p_{Ti}$  for  $\Gamma/M = 0, 0.01, 0.1, 0.5$ .

Unlike the finite width effects, the longitudinal boost effect does not completely smash the characteristic shape. The sharp cusp structures survive to some extent in both Type I and Type II cascade decays. The shift of the  $\cos\Theta$  cusp position is minor. Moreover, the overall functional shape remains the same even though the  $pp$  frame allows nonvanishing events around  $\cos\Theta = 0$ . The cusp in the  $\cos\Theta$  distribution, though Lorentz noninvariant, is quite useful to draw mass information. Again, we emphasize that the  $e^+e^-$  linear collider does not have this ambiguity.

### C. ISR effects

The cusps and endpoints in the  $m_T$  and  $p_{Ti}$  distributions play a crucial role in the mass measurement of cascade, especially Type II, decays. The results are based on the assumption that the transverse momentum of the mother particle  $D$  is known event by event so that its rest frame

along the transverse direction can be reconstructed. At a hadron collider, even though  $D$  is singly produced at the parton level, the gluon radiation from initial patrons, called the ISR, is inevitable because of strong QCD interaction: the mother particle  $D$  receives transverse kick and the ambiguity arises in its transverse motion [29].

We examine the ISR effects on the cusp and endpoint structures in the  $m_T$  and  $p_{Ti}$  distributions. As in Ref. [17], we adopt the simplifying assumptions that  $D$  is produced by  $q\bar{q}$  initial states, and decays through electroweak interactions so that final state radiation is neglected. We also veto hard ISR gluons if  $p_T^{\text{jet}} > 50$  GeV. We use the parton shower Monte Carlos in PYTHIA for the ISR [30].

Figure 21 compares the  $m_T$  distributions with and without ISR effects, denoted by the thin (red) line and the thick (blue) line, respectively. We take Mass-A<sub>1</sub> for Type I cascade decay, and Mass-C<sub>2</sub> for Type II decay. In both cases, the  $m_T$  cusps remain almost intact from the ISR

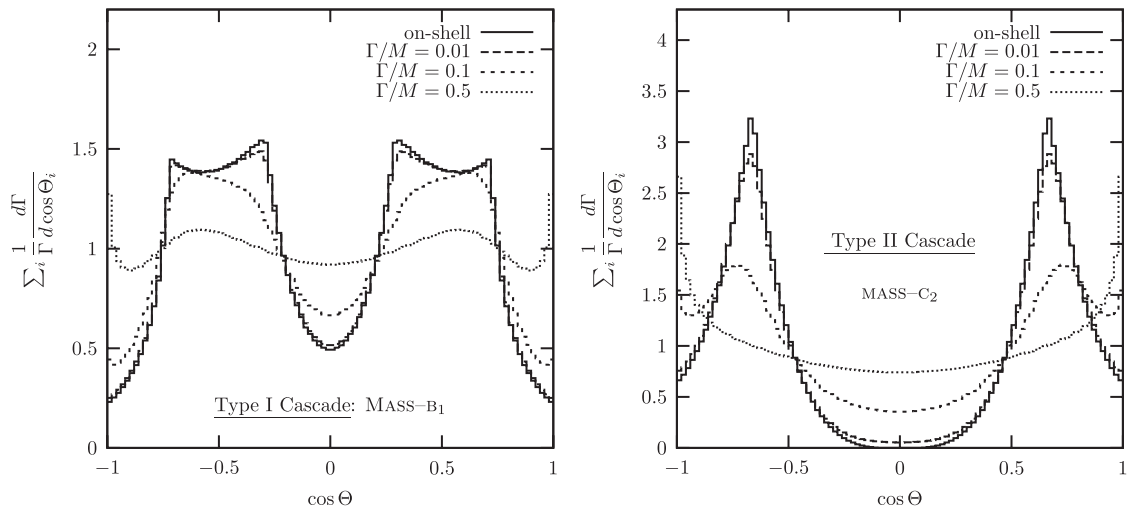


FIG. 19. The summed  $\cos\Theta$  distributions for  $\Gamma/M = 0, 0.01, 0.1, 0.5$ .

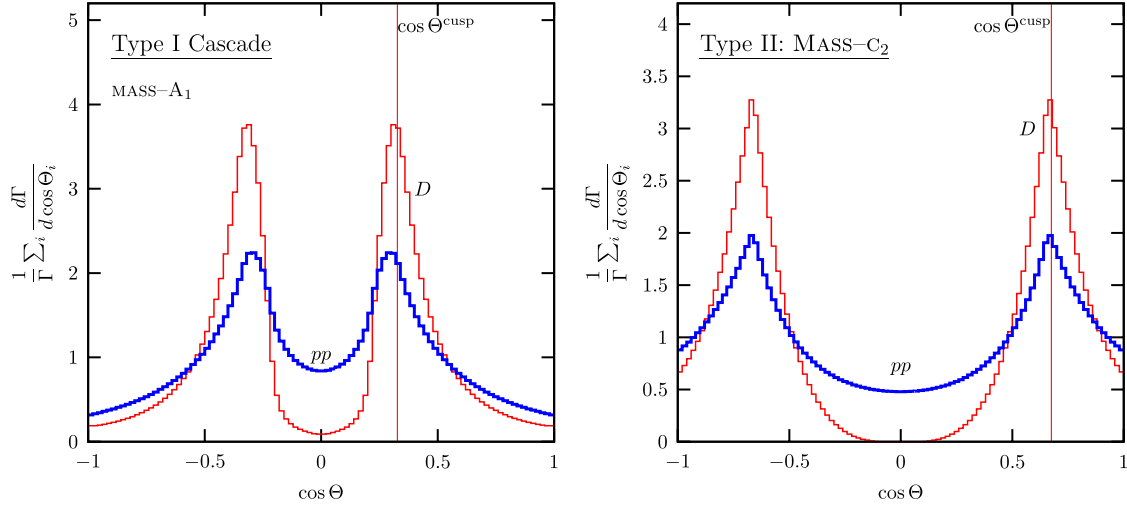


FIG. 20 (color online). Normalized differential decay rates versus  $\cos\Theta$  in the  $D$  rest frame (thin curves) and in the  $pp$  lab frame with  $\sqrt{s} = 14$  TeV (thick curves).

effects. The sharpness and position keep those in the  $D$  rest frame. The endpoints get tailed, but not significantly. Almost linearly dropping behaviors for both cases are preserved until the vicinity of endpoints. The extrapolation of the distribution will help to find the endpoint without the ISR effects.

In Fig. 22, we show the normalized  $p_{Tn}$  and  $p_{Tf}$  distributions with and without the ISR effects by the thin (red) line and the thick (blue) line, respectively, in the same setup as in Fig. 21. In the Type I cascade decay, the far visible particle accommodates both the cusp and endpoint singularities in its transverse momentum distribution. The ISR effects do not affect them. The near visible particle in the Type I decay has only spiky and suddenly ending endpoint. The ISR smears this into the tailed one. Still

the fast dropping behavior survives to some extent, which can be used to read the endpoint without ISR effects. In the Type II cascade decay, the near particle  $p_{Ti}$  distribution has the cusp and endpoint. The ISR effects do not change the cusp but do smear the endpoint. The  $p_{Ti}$  distribution of the far visible particle, which has only the endpoint, is not affected.

In summary, the inclusion of ISR with veto on hard jets does not change the cusp structures but gets the endpoint tailed.

#### D. Spin-correlation effect

Our main results are based on the kinematics only, ignoring the spin correlation in the full matrix elements. Since this paper is focused on the basic properties of the

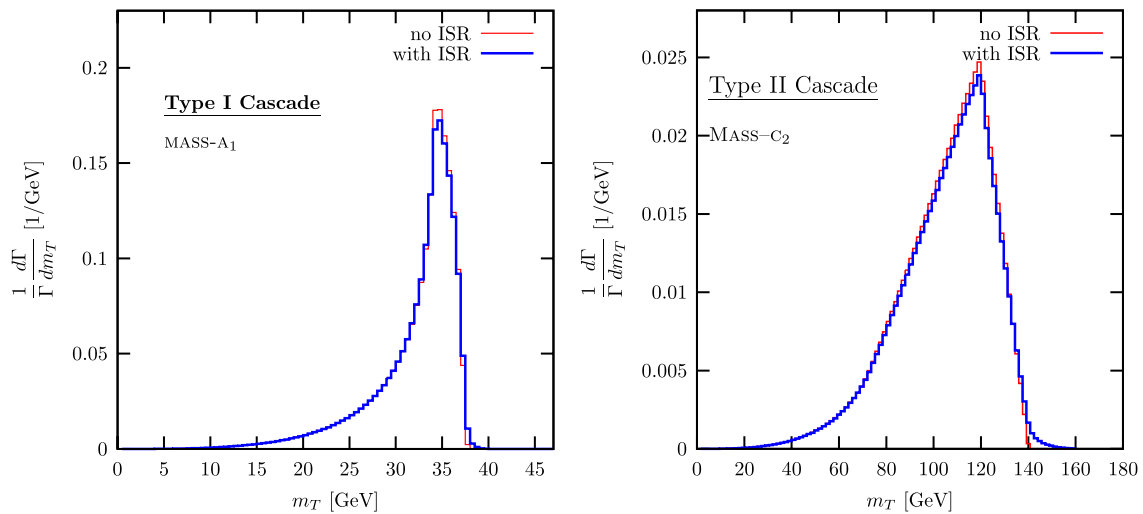


FIG. 21 (color online). The normalized transverse mass distributions with and without the ISR effects by the thin (red) line and the thick (blue) line, respectively. We take Mass- $A_1$  for the Type I cascade decay and Mass- $C_2$  for the Type II. Extra jets with  $p_T > 50$  GeV are vetoed.

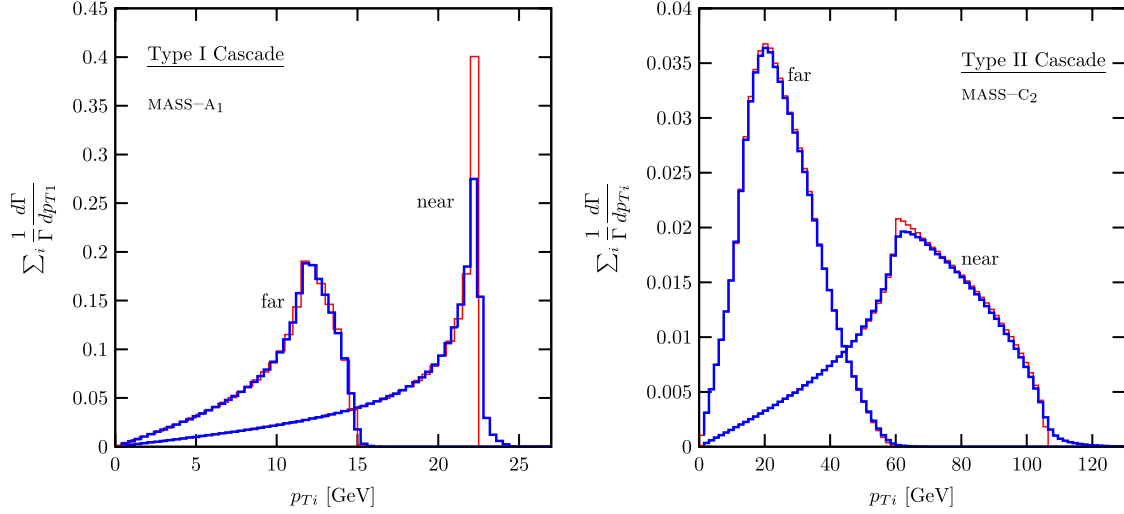


FIG. 22 (color online). The normalized  $p_{T_i}$  and  $p_{T_f}$  distributions with and without the ISR effects by the thin (red) line and the thick (blue) line, respectively. We take Mass-A<sub>1</sub> for the Type I cascade decay and Mass-C<sub>2</sub> for the Type II. Extra jets with  $p_T > 50$  GeV are vetoed.

kinematic singular structures in the cascade decays, full analysis for each new physics process is beyond the scope of this paper. Nevertheless we expect that the algebraic singularity origin of the cusp and endpoint keeps them stable under the spin-correlation effects [31].

In order to demonstrate this, we consider one example, the  $Z^{(2)}$  decay in the the UED model:

$$\begin{aligned} \text{Cascade Type I: } Z^{(2)} &\rightarrow \ell + L^{(2)} \rightarrow \ell + B^{(1)}L^{(1)} \\ &\rightarrow \ell\ell B^{(1)}B^{(1)}. \end{aligned} \quad (33)$$

In Fig. 23, we show their spin-correlation effects. We find that the spin correlations do not change the  $m$  and  $p_{T_i}$

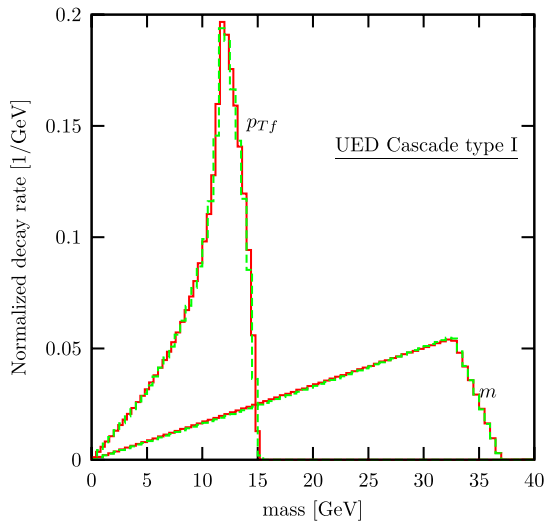


FIG. 23 (color online). The  $d\Gamma/dm$  and  $d\Gamma/dp_{T_i}$  for the process of  $Z^{(2)} \rightarrow \ell + L^{(2)} \rightarrow \ell + B^{(1)}L^{(1)} \rightarrow \ell\ell B^{(1)}B^{(1)}$  with and without spin correlations.

distributions. Two distributions with and without spin-correlation effects are almost identical.

We note that there are other uncertainties such as the SM background and experimental resolutions. The magnitudes of those effects depend on a specific process, which are beyond the scope of this paper. For a benchmark process of  $pp \rightarrow Z' \rightarrow \tilde{\ell}^+ \tilde{\ell}^- \rightarrow \ell^+ \tilde{\chi}_1^0 \ell^- \tilde{\chi}_1^0$  in a supersymmetry model with an extra  $U(1)$  gauge field, we have considered all the realistic effects including the SM backgrounds and detector simulation in Ref. [16]. The total uncertainty in the mass measurement is about 10%. Most of the uncertainty is from huge SM  $t\bar{t}$  backgrounds. Unless the observed decay products are jets, the experimental resolution does not affect the cusp and endpoint structures dominantly. The crucial role was done by best-fitting the  $m$  distribution based on its analytic functional form.

## VI. SUMMARY AND CONCLUSIONS

We have studied the singularity structure, such as cusps and endpoints, in the kinematic distributions of three-step cascade decay of a new parity-even particle  $D$  and the determination of the missing particle mass by using such singularities.

Two nontrivial decay topologies, called the Type I and Type II cascade decays, have been studied. In the Type I decay ( $D \rightarrow a_1 C$ ,  $C \rightarrow X_1 B$ ,  $B \rightarrow a_2 X_2$ ), the distribution of the invariant mass  $m$  of two visible particles,  $a_1$  and  $a_2$ , develops a cusp. Full functional form of the  $m$  distribution for general mass parameters has been derived. If the mother particle  $D$  is produced at rest in the transverse direction, various longitudinal-boost invariant observables accommodate cusp structures. First the transverse mass  $m_T$  distribution has a cusp, which is complementary for the  $m$  cusp since the  $m_T$  cusp is sharp even when the  $m$  cusp is

TABLE III. The presence or absence of the cusp in the kinematic distributions of  $m$ ,  $m_T$ ,  $p_T$ ,  $p_{Ti}$ ,  $\cos\Theta$  of the antler, Type I cascade, Type II cascade decays.

|              | Antler | Cascade |         |
|--------------|--------|---------|---------|
|              |        | Type I  | Type II |
| $M$          | Yes    | Yes     | No      |
| $M_t$        | No     | Yes     | Yes     |
| $P_t$        | No     | No      | No      |
| $P_{tN}$     | Yes    | No      | Yes     |
| $P_{tF}$     | Yes    | Yes     | No      |
| $\cos\Theta$ | Yes    | Yes     | Yes     |

dull. Although the transverse momentum distribution of the two-visible-particle system does not develop a visible cusp structure and a sharp endpoint, we note that the transverse momentum distribution of the far visible particle  $a_2$  has a cusp and that of the near visible particle  $a_1$  has an endpoint in the shape of a steep cliff. We also study the summed distribution of  $\cos\Theta_i$ , which has two independent cusp structures.

In the Type II decay ( $D \rightarrow X_1 C$ ,  $C \rightarrow a_1 B$ ,  $B \rightarrow a_2 X_2$ ), the kinematics of  $a_1$  and  $a_2$  is determined solely by the two-step cascade decay from the first intermediate particle  $C$ . The invariant mass distribution does not have a cusp structure. However, the kinematic distributions of the transverse motion carry the information from the whole three-step cascade decay. Both the  $m_T$  and  $\sum_i p_{Ti}$  distributions have distinctive cusp structures. In the individual transverse momentum distribution, only the near visible particle has both a sharp cusp and a fast-falling endpoint. The  $\cos\Theta$  distribution also shows a cusp as well. Including the antler decay topology, we have summarized the existence of cusp in the kinematic distributions of  $m$ ,  $m_T$ ,  $p_T$ ,  $p_{Ti}$ , and  $\cos\Theta$  in Table III.

We have also considered the effects of the finite decay widths of intermediate particles, the longitudinal boost of the mother particle  $D$ , the ISR, and the spin correlation. The finite width effects are significant if the decay width is sizable like  $\Gamma/M \gtrsim 10\%$ : the sharp cusp gets smeared; the endpoint position gets shifted to  $m_D - 2m_X$ . The longitudinal motion of the mother particle  $D$  affects the distribution of  $\cos\Theta$ . At least for the sample mass parameters, however, the  $\cos\Theta$  cusp remains sharp after convoluting with the parton distribution functions of a proton at the LHC. The ISR effects do not change the cusp structure much if we veto hard jets, but they do smear the endpoints. Spin correlation effects from full  $S$ -matrix elements turn out to be negligible in most cases, which is expected since the singularities are determined by the kinematic relations.

With the companion paper on the antler decay [17], our analysis presents the general properties and useful formulas of the kinematic cusps and endpoints for the decay topologies with two visible particles and two missing particles. By looking at the singularity structures of various

kinematic distributions, the hidden nature of the missing particle can be probed effectively and elegantly. With the outstanding performance of the LHC and detectors, this is an exciting time for such investigation.

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## APPENDIX: INVARIANT MASS DISTRIBUTIONS FOR THE GENERAL TYPE I CASE

In this appendix, we present the invariant mass distribution in the general Type I cascade decays,

$$D(P) \rightarrow C + a_1(k_1), \quad C \rightarrow B + X_1, \quad (A1)$$

$$B \rightarrow a_2(k_2) + X_2.$$

As discussed in the main text, the Type II cascade decay is practically a three-body decay in the viewpoint of visible particles. This four-body decay generally has seven different mass parameters. We define the rapidities of six particles as

$$\begin{aligned} \cosh\xi_C &= \frac{m_D^2 + m_C^2 - m_{a_1}^2}{2m_D m_C}, \\ \cosh\xi_{a_1} &= \frac{m_D^2 + m_{a_1}^2 - m_C^2}{2m_D m_{a_1}}, \\ \cosh\xi_B &= \frac{m_C^2 + m_B^2 - m_{X_1}^2}{2m_C m_B}, \\ \cosh\xi_{X_1} &= \frac{m_C^2 + m_{X_1}^2 - m_B^2}{2m_C m_{X_1}}, \\ \cosh\xi_{a_2} &= \frac{m_B^2 + m_{a_2}^2 - m_{X_2}^2}{2m_B m_{a_2}}, \\ \cosh\xi_{X_2} &= \frac{m_B^2 + m_{X_2}^2 - m_{a_2}^2}{2m_B m_{X_2}}. \end{aligned} \quad (A2)$$

A very useful kinematic variable is  $\chi$ , the rapidity of the particle  $a_2$  in the rest frame of  $a_1$ ,

$$\chi \equiv \cosh\xi_{a_2}^{(a_1)} = \frac{m^2 - m_{a_1}^2 - m_{a_2}^2}{2m_{a_1} m_{a_2}}, \quad (A3)$$

where the superscript  $(a_1)$  denotes that the rapidity is defined in the rest frame of  $a_1$ .

The functional expression of  $d\Gamma/dm$  is different according to the mass relations. The derivation of  $d\Gamma/dm$  is similar to that presented in the appendix of Ref. [17]. For simple presentation, we introduce

$$\xi_{++} = \xi_B + \xi_{a_1} + \xi_{a_2} + \xi_C, \quad (\text{A4})$$

$$\xi_{+-} = |\xi_B + \xi_{a_1} - \xi_{a_2} - \xi_C|, \quad (\text{A5})$$

$$\xi_{-+} = |\xi_B - \xi_{a_1} + \xi_{a_2} + \xi_C|, \quad (\text{A6})$$

$$\xi_{--} = |\xi_B - \xi_{a_1} - \xi_{a_2} - \xi_C|. \quad (\text{A7})$$

We order  $\xi_{+-}$ ,  $\xi_{-+}$  and  $\xi_{--}$  and name them  $\xi_1 \leq \xi_2 \leq \xi_3$ . Analytic function forms of  $d\Gamma/d\chi$  are then written as  
 (i) if  $|\xi_B - \xi_{a_2} - \xi_C| \geq \xi_{a_1}$  or  $\xi_B + \xi_{a_2} + \xi_C \leq \xi_{a_1}$ ,

$$\frac{d\Gamma}{d\chi} \propto \begin{cases} -\xi_1 + \cosh^{-1}\chi, & \text{if } \cosh\xi_1 \leq \chi \leq \cosh\xi_2, \\ \xi_2 - \xi_1, & \text{if } \cosh\xi_2 \leq \chi \leq \cosh\xi_3, \\ \xi_{++} - \cosh^{-1}\chi, & \text{if } \cosh\xi_3 \leq \chi \leq \cosh\xi_{++}, \\ 0, & \text{otherwise.} \end{cases} \quad (\text{A8})$$

(ii) if  $|\xi_B - \xi_{a_2} - \xi_C| < \xi_{a_1} < \xi_B + \xi_{a_2} + \xi_C$ ,

$$\frac{d\Gamma}{d\chi} \propto \begin{cases} 2\cosh^{-1}\chi, & \text{if } 1 \leq \chi \leq \cosh\xi_1, \\ \xi_1 + \cosh^{-1}\chi, & \text{if } \cosh\xi_1 \leq \chi \leq \cosh\xi_2, \\ \xi_1 + \xi_2, & \text{if } \cosh\xi_2 \leq \chi \leq \cosh\xi_3, \\ \xi_{++} - \cosh^{-1}\chi, & \text{if } \cosh\xi_3 \leq \chi \leq \cosh\xi_{++}, \\ 0, & \text{otherwise.} \end{cases} \quad (\text{A9})$$

The positions of the minimum, cusp, and maximum of the invariant mass distribution are

$$\begin{aligned} M_{\text{cas1}}^{\text{min}} &= \begin{cases} \sqrt{m_{a_1}^2 + m_{a_2}^2 + 2m_{a_1}m_{a_2}\cosh\xi_1}, & \text{for } \mathcal{R}_{1,\dots,6} \\ m_{a_1} + m_{a_2}, & \text{for } \mathcal{R}_{7,\dots,12} \end{cases} \\ M_{\text{cas1}}^{\text{cusp}} &= \sqrt{m_{a_1}^2 + m_{a_2}^2 + 2m_{a_1}m_{a_2}\cosh\xi_3}, \\ M_{\text{cas1}}^{\text{max}} &= \sqrt{m_{a_1}^2 + m_{a_2}^2 + 2m_{a_1}m_{a_2}\cosh\xi_{++}}. \end{aligned} \quad (\text{A10})$$

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- [1] S. Perlmutter *et al.*, *Astrophys. J.* **517**, 565 (1999); A. G. Riess *et al.*, *Astron. J.* **116**, 1009 (1998); A. G. Riess, R. P. Kirshner, B. P. Schmidt *et al.*, *Astron. J.* **117**, 707 (1999).  
 [2] E. J. Copeland, M. Sami, and S. Tsujikawa, *Int. J. Mod. Phys. D* **15**, 1753 (2006); L. M. Krauss and M. S. Turner, *Gen. Relativ. Gravit.* **27**, 1137 (1995).  
 [3] G. Bertone, D. Hooper, and J. Silk, *Phys. Rep.* **405**, 279 (2005).  
 [4] A. Djouadi, *Phys. Rep.* **459**, 1 (2008); G. Jungman, M. Kamionkowski, and K. Griest, *Phys. Rep.* **267**, 195 (1996).  
 [5] G. Servant and T.M.P. Tait, *Nucl. Phys.* **B650**, 391 (2003); F. Burnell and G.D. Kribs, *Phys. Rev. D* **73**, 015001 (2006); K. Kong and K.T. Matchev, *J. High Energy Phys.* **01** (2006) 038.  
 [6] I. Low, *J. High Energy Phys.* **10** (2004) 067; A. Birkedal, A. Noble, M. Perelstein, and A. Spray, *Phys. Rev. D* **74**, 035002 (2006); J. Hubisz and P. Meade, *Phys. Rev. D* **71**, 035016 (2005); A. Freitas, P. Schwaller, and D. Wyler, *J. High Energy Phys.* **12** (2009) 027; **02** (2011) 032(E).  
 [7] M. Burns, K. Kong, K. T. Matchev, and M. Park, *J. High Energy Phys.* **03** (2009) 143.  
 [8] I. Hinchliffe, F.E. Paige, M. D. Shapiro, J. Soderqvist, and W. Yao, *Phys. Rev. D* **55**, 5520 (1997); H. Bachacou, I. Hinchliffe, and F.E. Paige, *Phys. Rev. D* **62**, 015009



- (2000); B.C. Allanach, C.G. Lester, M.A. Parker, and B.R. Webber, *J. High Energy Phys.* **09** (2000) 004; B.K. Gjelsten, D.J. Miller, and P. Osland, *J. High Energy Phys.* **12** (2004) 003; **06** (2005) 015.
- [9] M.M. Nojiri, G. Polesello, and D.R. Tovey, [arXiv:hep-ph/0312317](https://arxiv.org/abs/hep-ph/0312317); K. Kawagoe, M.M. Nojiri, and G. Polesello, *Phys. Rev. D* **71**, 035008 (2005).
- [10] H.C. Cheng, J.F. Gunion, Z. Han, G. Marandella, and B. McElrath, *J. High Energy Phys.* **12** (2007) 076; M.M. Nojiri and M. Takeuchi, *J. High Energy Phys.* **10** (2008) 025; H.C. Cheng, D. Engelhardt, J.F. Gunion, Z. Han, and B. McElrath, *Phys. Rev. Lett.* **100**, 252001 (2008).
- [11] C.G. Lester and D.J. Summers, *Phys. Lett. B* **463**, 99 (1999).
- [12] A. Barr, C. Lester, and P. Stephens, *J. Phys. G* **29**, 2343 (2003); P. Meade and M. Reece, *Phys. Rev. D* **74**, 015010 (2006); S. Matsumoto, M.M. Nojiri, and D. Nomura, *Phys. Rev. D* **75**, 055006 (2007); C. Lester and A. Barr, *J. High Energy Phys.* **12** (2007) 102.
- [13] W.S. Cho, K. Choi, Y.G. Kim, and C.B. Park, *Phys. Rev. Lett.* **100**, 171801 (2008); B. Gripaios, *J. High Energy Phys.* **02** (2008) 053; A.J. Barr, B. Gripaios, and C.G. Lester, *J. High Energy Phys.* **02** (2008) 014; W.S. Cho, K. Choi, Y.G. Kim, and C.B. Park, *J. High Energy Phys.* **02** (2008) 035; M.M. Nojiri, Y. Shimizu, S. Okada, and K. Kawagoe, *J. High Energy Phys.* **06** (2008) 035.
- [14] M. Serna, *J. High Energy Phys.* **06** (2008) 004; M.M. Nojiri, K. Sakurai, Y. Shimizu, and M. Takeuchi, *J. High Energy Phys.* **10** (2008) 100.
- [15] H.-C. Cheng and Z. Han, *J. High Energy Phys.* **12** (2008) 063; A.J. Barr, B. Gripaios, and C.G. Lester, *J. High Energy Phys.* **11** (2009) 096.
- [16] T. Han, I.W. Kim, and J. Song, *Phys. Lett. B* **693**, 575 (2010).
- [17] T. Han, I.W. Kim, and J. Song, [arXiv:1206.5633](https://arxiv.org/abs/1206.5633).
- [18] K. Agashe, D. Kim, M. Toharia, and D.G.E. Walker, *Phys. Rev. D* **82**, 015007 (2010); W.S. Cho, D. Kim, K.T. Matchev, and M. Park, [arXiv:1206.1546](https://arxiv.org/abs/1206.1546).
- [19] A. Datta, K. Kong, and K.T. Matchev, *Phys. Rev. D* **72**, 096006 (2005); **72**, 119901(E) (2005); H.C. Cheng, K.T. Matchev, and M. Schmaltz, *Phys. Rev. D* **66**, 036005 (2002).
- [20] H.C. Cheng, K.T. Matchev, and M. Schmaltz, *Phys. Rev. D* **66**, 036005 (2002).
- [21] T. Appelquist and H.U. Yee, *Phys. Rev. D* **67**, 055002 (2003).
- [22] T.G. Rizzo and J.D. Wells, *Phys. Rev. D* **61**, 016007 (1999); A. Strumia, *Phys. Lett. B* **466**, 107 (1999); C.D. Carone, *Phys. Rev. D* **61**, 015008 (1999); I. Gogoladze and C. Macesanu, *Phys. Rev. D* **74**, 093012 (2006).
- [23] P. Nath and M. Yamaguchi, *Phys. Rev. D* **60**, 116006 (1999); K. Agashe, N.G. Deshpande, and G.H. Wu, *Phys. Lett. B* **511**, 85 (2001).
- [24] D. Chakraverty, K. Huiti, and A. Kundu, *Phys. Lett. B* **558**, 173 (2003); A.J. Buras, M. Spranger, and A. Weiler, *Nucl. Phys.* **B660**, 225 (2003); A.J. Buras, A. Poschenrieder, M. Spranger, and A. Weiler, *Nucl. Phys.* **B678**, 455 (2004); K. Agashe, N.G. Deshpande, and G.H. Wu, *Phys. Lett. B* **514**, 309 (2001).
- [25] B. Abbott *et al.* (D0 Collaboration), *Phys. Rev. Lett.* **83**, 4937 (1999).
- [26] S. Chang, K.Y. Lee, and J. Song, *Phys. Lett. B* **708**, 144 (2012); *Phys. Rev. D* **85**, 055006 (2012); S. Chang, K.Y. Lee, S.Y. Shim, and J. Song, *Phys. Rev. D* **86**, 117503 (2012).
- [27] I. Hinchliffe, F.E. Paige, M.D. Shapiro, J. Soderqvist, and W. Yao, *Phys. Rev. D* **55**, 5520 (1997); H. Bachacou, I. Hinchliffe, and F.E. Paige, *Phys. Rev. D* **62**, 015009 (2000).
- [28] S. Kretzer, H.L. Lai, F.I. Olness, and W.K. Tung, *Phys. Rev. D* **69**, 114005 (2004).
- [29] Y.-K. Kim, U. Yang, Report No. CDF/PHY/TOP/CDFR/6804; J.M. Campbell, J.W. Huston, and W.J. Stirling, *Rep. Prog. Phys.* **70**, 89 (2007).
- [30] T. Sjostrand, S. Mrenna, and P.Z. Skands, *J. High Energy Phys.* **05** (2006) 026.
- [31] I.W. Kim, *Phys. Rev. Lett.* **104**, 081601 (2010).