

Z_p scalar dark matter from multi-Higgs-doublet modelsI. P. Ivanov^{1,2} and V. Keus¹¹*IFPA, Université de Liège, Allée du 6 Août 17, bâtiment B5a, 4000 Liège, Belgium*²*Sobolev Institute of Mathematics, Koptyug avenue 4, 630090, Novosibirsk, Russia*

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In many models, stability of dark-matter particles is protected by a conserved \mathbb{Z}_2 quantum number. However, dark matter can be stabilized by other discrete symmetry groups, and examples of such models with custom-tailored field content have been proposed. Here, we show that electroweak symmetry-breaking models with N Higgs doublets can readily accommodate scalar dark matter candidates stabilized by groups \mathbb{Z}_p with any $p \leq 2^{N-1}$, leading to a variety of kinds of microscopic dynamics in the dark sector. We give examples in which semiannihilation or multiple semiannihilation processes are allowed or forbidden, which can be especially interesting in the case of asymmetric dark matter.

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I. INTRODUCTION

Despite compelling astronomical evidence for the existence of dark matter [1], there is still no direct experimental clue of which particle can play the role of dark-matter candidate. It is only known that dark matter cannot be satisfactorily explained by the standard model (SM) particle content. In this situation, one can focus on exploring dark-matter candidates arising in various models beyond the SM, especially if these models can simultaneously address other particle physics issues such as electroweak symmetry breaking and small neutrino masses.

Dark-matter particles must be (almost) stable on cosmological time scales. In many models, this stability is provided by imposing by hand a conserved \mathbb{Z}_2 quantum number on the Lagrangian. This quantum number generically called parity appears in different contexts; examples include the R -parity in supersymmetric models, the matter parity [2], the sign flip in the Inert doublet model, [3], which revived the old idea of Deshpande and Ma [4], or in the minimal singlet model [5]. When assigning the parity quantum number of the SM and new fields, different options are available. The simplest is to assume that all the SM particles including the SM-like Higgs boson have positive parity, while the dark sector particles are of negative parity. Another possibility is to exploit statistics by assigning positive \mathbb{Z}_2 parity to all SM boson fields and negative \mathbb{Z}_2 parity to all SM fermions as well as to the new scalar. This situation is realized, for example, in models exploiting the matter parity [2] to stabilize the dark matter. In both cases, the lightest non-SM state, which we generically denote d , cannot couple to SM final states without violating the \mathbb{Z}_2 parity. Therefore, in a \mathbb{Z}_2 -conserving model, this state is stable and represents the dark-matter candidate.

It is natural to ask whether dark matter can be stabilized by a conserving discrete quantum number taking values in a group other than \mathbb{Z}_2 . This idea is not new and was

explored in a descent number of works, in which both Abelian [6–8] and non-Abelian [9] finite groups were used.¹ Requiring that d transforms nontrivially under the discrete group implies that it cannot be a truly neutral particle. It also prohibits direct two-particle annihilation $dd \rightarrow X_{\text{SM}}$, where X_{SM} is any set of SM particles. However, more complicated processes involving several d 's can take place. This certainly changes the kinetics of dark-matter evolution in the early Universe and its relic abundance after the freeze-out. If, in addition, one assumes that an asymmetry between d and its antiparticle d^* is generated at a high-energy scale in a way similar to the baryon asymmetry of the usual matter [11], then the present-day behavior of this asymmetric dark matter is not dominated anymore by dd^* annihilation and can lead to characteristic observational signatures.

One particular class of groups used to stabilize dark matter are cyclic groups \mathbb{Z}_p , see Refs. [6,7]. With this choice, all fields are characterized by a conserved quantum number which we will call the \mathbb{Z}_p charge q and which is additive modulo p . The usual assignment is that all the SM fields including the SM-like Higgs boson have $q = 0$, while the dark-matter candidates have nonzero \mathbb{Z}_p charges. This opens up the possibility of novel two-particle processes such as $dd \rightarrow d^* X_{\text{SM}}$, which was called semiannihilation in Ref. [12], or even multiparticle versions of it, “multiple semiannihilation.” Besides, if the model allows for the existence of several dark-matter candidates d_i with different charges q_i , then inelastic two-particle processes $d_i d_j \rightarrow d_k X_{\text{SM}}$ are also possible. Such processes, too, have impact on the kinetics of the dark-matter abundances in the early Universe [13].

¹If one takes seriously the argument that quantum-gravitational effects violate any global discrete symmetry, one should require that this quantum number arise as a remnant of a $U(1)$ gauge symmetry spontaneously broken at a high-energy scale [10].

Examples of \mathbb{Z}_p -stabilized dark sectors often involve a variety of new fields which interact via Lagrangians designed specifically to incorporate a given symmetry group, see e.g. Ref. [7]. Even when linked to electroweak physics, these models involve new fields with different electroweak quantum numbers (doublets, singlets, etc.) [13]. Indeed, if one assumes that extra fields come from a hidden sector coupled to the SM fields via “portal” operators [14], then they must be electroweak singlets by construction.

In this paper, we demonstrate that \mathbb{Z}_p -stabilized scalar dark matter can easily arise in multi-Higgs-doublet models. This, perhaps, is not surprising on its own. A less trivial fact is that even with few doublets, one can get \mathbb{Z}_p with a rather large p . To be precise, in models with N Higgs doublets, \mathbb{Z}_p with any $p \leq 2^{N-1}$ is realizable in the scalar sector. We will show that this fact can be instrumental in avoiding semiannihilation processes even with multicomponent dark-matter sectors.

The structure of the paper is the following. In Sec. II, we quickly review what is known about symmetries in the scalar sector of N -Higgs-doublet models (NHDM). Then, in Sec. III, we give a 3-Higgs-doublet model (3HDM) example with dark-matter candidates stabilized by the \mathbb{Z}_3 symmetry group. In Sec. IV, we show what is possible with four doublets and consider in some detail the \mathbb{Z}_7 -symmetric 4-Higgs-doublet model (4HDM), in which one can avoid semiannihilation processes. We end the paper with a discussion and conclusions.

II. SCALAR SECTOR OF NHDM AND ITS SYMMETRIES

The N -Higgs-doublet model is a conceptually simple extension of the SM Higgs mechanism. It is driven by the idea that the Higgs fields, similarly to fermions, can come in several generations. Its simplest version with only two doublets, 2HDM, is of special interest because it mimics the Higgs sector of minimal supersymmetric standard model, and it has been thoroughly studied in the last four decades; see Ref. [15] and references therein. In addition, many particular models employing more than two scalar doublets have also been proposed [16].

In NHDM, one introduces N doublets of complex scalar fields ϕ_i , $i = 1, \dots, N$, with electroweak isospin $Y = 1/2$, and constructs the self-interaction renormalizable Higgs potential

$$V = Y_{ij}(\phi_i^\dagger \phi_j) + Z_{ijkl}(\phi_i^\dagger \phi_j)(\phi_k^\dagger \phi_l), \quad (1)$$

where all indices run from 1 to N . The free parameters of the potential are written as components of tensors Y_{ij} and Z_{ijkl} ; in the most general case, there are $N^2(N^2 + 3)/2$ independent parameters. Once these coefficients and the Yukawa couplings are provided, the model is completely defined, and the entire phenomenology should follow. In practice, however, inferring these consequences directly

from the Lagrangian is impeded by algebraic obstacles at the very first step, namely, the minimization of a sufficiently generic potential. The consequence is that only very few general results are known for $N > 2$ [17–19].

One particular issue which is of much importance and where certain progress has been recently made concerns accidental symmetries which can be encoded in the scalar sector of NHDM. These are transformations which mix several Higgs doublets (also called Higgs-basis transformations) but still leave the potential invariant due to specific patterns in the tensors Y_{ij} and Z_{ijkl} . Although particular models with several Higgs doublets based on various symmetry groups have been proposed and studied over the last decades, no attempt at systematic classification of possible symmetries was made until very recently.

One of the problems here comes from the observation that just imposing a certain symmetry group G on the Higgs potential can often lead to potentials symmetric under a *larger* group, which includes G as a subgroup. This feature was discussed in Ref. [18] where one-parametric symmetry groups were studied for 3HDM and in Ref. [19] where an attempt to understand symmetries in NHDM via geometric constructions in the space of bilinears was made. Therefore, when describing symmetries of the model, one should focus only on the true symmetry groups, the ones which are not automatically extended to larger groups. These groups were called in Ref. [19] “realizable groups.”

Unfortunately, no systematic way to reconstruct the true symmetry group of a given NHDM potential for $N > 2$ is known so far. One even does not know the list of realizable symmetry groups possible for $N = 3$. However, a step forward was recently made in Ref. [20], which completely characterized all *Abelian* groups of Higgs-family transformations and generalized *CP*-transformations which can be realized as symmetry groups of the scalar sector of NHDM. In particular, in what concerns realizable cyclic groups, it was proved there that for a model with N doublets, one can construct the Higgs potential symmetric under the group \mathbb{Z}_p with any $1 < p \leq 2^{N-1}$ (for $N = 3$, this conclusion was known before [18]). One can therefore wonder if these models can be used to construct dark matter sectors stabilized by these symmetry groups.

An electroweak-symmetry-breaking model with \mathbb{Z}_p -stabilized scalar dark matter must satisfy several conditions. First, the entire Lagrangian and not only the Higgs potential must be \mathbb{Z}_p -symmetric. The simplest way to achieve this is to set the \mathbb{Z}_p charges of all the SM particles to zero and to require that only one Higgs doublet (the SM-like doublet) couples to fermions. The \mathbb{Z}_p charge of this doublet must be zero, and it does not matter which doublet is chosen to be SM-like due to the freedom to simultaneously shift the \mathbb{Z}_p charges of all the doublets. Second, the \mathbb{Z}_p symmetry must remain after electroweak symmetry breaking. This is possible when only the SM-like doublet

acquires a nonzero vacuum expectation value (v.e.v.). Third, if we insist on \mathbb{Z}_p stabilization, that is, we require that not only decays but also $2-$, $3-$, \dots , $(p-1)$ -particle annihilation to SM fields are forbidden by quantum numbers, then the dark matter candidates must have \mathbb{Z}_p charge q which is coprime with p .

Below, we show that multi-Higgs-doublet models can easily satisfy these conditions. We start first with the simplest model of this kind, \mathbb{Z}_3 -symmetric 3HDM, and then we outline the plethora of kinds of microscopic dark matter dynamics possible in models with four doublets.

III. \mathbb{Z}_3 -SYMMETRIC 3HDM

A natural way to implement the \mathbb{Z}_3 symmetry in 3HDM would be to construct a potential symmetric under $\phi_1 \rightarrow \phi_2 \rightarrow \phi_3 \rightarrow \phi_1$. However, upon an appropriate Higgs-basis change, this transformation will turn into pure phase rotations of certain doublets. In fact, it can be proven that *any* Abelian subgroup of $SU(N)$ can be mapped onto a group of (possibly correlated) phase rotations of individual doublets, see Ref. [20] for explicit construction. Therefore, we will always use below the phase rotation representations of the cyclic symmetry groups. Also, to keep the notation short, we will describe any such transformation $\phi_i \rightarrow e^{i\alpha_i} \phi_i$ by providing the N -tuple of phases α_i .

A scalar potential invariant under a certain group G of phase rotations can be written as a sum $V = V_0 + V_G$, where V_0 is invariant under any phase rotation, while V_G is a collection of extra terms which realize the chosen symmetry group. The generic phase rotation invariant part has form

$$V_0 = \sum_i [-m_i^2(\phi_i^\dagger \phi_i) + \lambda_{ii}(\phi_i^\dagger \phi_i)^2] + \sum_{ij} [\lambda_{ij}(\phi_i^\dagger \phi_i)(\phi_j^\dagger \phi_j) + \lambda'_{ij}(\phi_i^\dagger \phi_j)(\phi_j^\dagger \phi_i)], \quad (2)$$

while V_G obviously depends on the group. In particular, for the group \mathbb{Z}_3 in 3HDM, we have

$$V_{\mathbb{Z}_3} = \lambda_1(\phi_3^\dagger \phi_1)(\phi_2^\dagger \phi_1) + \lambda_2(\phi_1^\dagger \phi_2)(\phi_3^\dagger \phi_2) + \lambda_3(\phi_2^\dagger \phi_3)(\phi_1^\dagger \phi_3) + \text{H.c.}, \quad (3)$$

where at least two of the coefficients $\lambda_1, \lambda_2, \lambda_3$ are nonzero (otherwise, the potential would have a continuous symmetry). This potential is symmetric under the phase rotations generated by

$$a = \frac{2\pi}{3}(0, 1, 2), \quad a^3 = 1, \quad (4)$$

In fact, the assignment of these charges to the three doublets is completely arbitrary, and the group generated by the generator a with permuted charges has the same action of the potential. In Eq. (4), we simply chose ϕ_1 to be the SM-like doublet.

Whether this symmetry is conserved or spontaneously broken depends on the pattern of the vacuum expectation values. If we insist on conservation of the \mathbb{Z}_3 symmetry, we require that $\langle \phi_1^0 \rangle = v_1 = v/\sqrt{2}$, $\langle \phi_2^0 \rangle = \langle \phi_3^0 \rangle = 0$. There is nothing surprising that the potential $V_0 + V_{\mathbb{Z}_3}$ can have a \mathbb{Z}_3 -symmetric global minimum upon an appropriate choice of coefficients. The question is whether it requires any fine-tuning or not. Below, we show that it does not, and the minimum of the type $(v_1, 0, 0)$ arises in a sizable part of the entire available parameter space of this model.

First, we note that if $(v_1, 0, 0)$ is an extremum of V_0 , then it is also an extremum of $V_0 + V_{\mathbb{Z}_3}$ because the extra terms contain ϕ_1 only linearly and quadratically. Therefore, when constructing a model example, one can first build V_0 with a global minimum at $(v_1, 0, 0)$ and then add a sufficiently weak $V_{\mathbb{Z}_3}$ so that this point remains a minimum.

Now, turning to minimization of V_0 , suppose that we search for the neutral minimum with a generic complex v.e.v. pattern (v_1, v_2, v_3) . By introducing $\rho_i = |v_i|^2 \geq 0$, we can rewrite V_0 as

$$V_0 = -M_i \rho_i + \frac{1}{2} \Lambda_{ij} \rho_i \rho_j = \frac{1}{2} \Lambda_{ij} (\rho_i - \mu_i)(\rho_j - \mu_j) + \text{const.} \quad (5)$$

Here, $M_i = (m_1^2, m_2^2, m_3^2)$, Λ_{ij} is constructed from λ_{ij} and λ'_{ij} in an obvious way, and $\mu_i = (\Lambda_{ij})^{-1} M_j$. Positivity condition on the potential guarantees that Λ_{ij} is a positive definite matrix; therefore, its inverse exists (we omit here the degenerate situations when the quartic potential has flat directions).

The allowed values of the ρ_i populate the first octant ($\rho_i \geq 0$) in the three-dimensional Euclidean space. Because of Eq. (5), the search for the global minimum can be reformulated as the search for the point in the first octant which lies closest to μ_i in the Euclidean metric defined by Λ_{ij} . Clearly, if μ_i itself lies inside the first octant ($\mu_1, \mu_2, \mu_3 \geq 0$), then the global minimum is at $\rho_i = \mu_i$. If μ_i lies outside the first octant, then the closest point lies either on a face or on an edge of the first octant, or at the origin. The global minimum is unique by the convexity arguments. It is now clear from this geometric construction that the entire space of all possible vectors μ_i can be broken into several regions of nonzero measure which correspond to all of these possibilities. In particular, the region corresponding to vacuum alignment of the type $(v_1, 0, 0)$ also has a nonzero measure and fills a sizeable part of the parameter space. In this sense, this vacuum pattern does not require any fine-tuning of the coefficients of the potential.

Having established that the required vacuum pattern is generically possible, we now switch to a simple version of the model. This is done only to simplify the presentation of the argument; if needed, the calculations can be repeated

for a generic potential. Namely, we now choose $m_1^2 > 0$, $m_2^2, m_3^2 < 0$ and take $\Lambda_{ij} = 2\lambda_0\delta_{ij}$, so that the potential becomes

$$V = -m_1^2(\phi_1^\dagger\phi_1) + |m_2^2|(\phi_2^\dagger\phi_2) + |m_3^2|(\phi_3^\dagger\phi_3) \\ + \lambda_0[(\phi_1^\dagger\phi_1)^2 + (\phi_2^\dagger\phi_2)^2 + (\phi_3^\dagger\phi_3)^2] \\ + \lambda_1(\phi_3^\dagger\phi_1)(\phi_2^\dagger\phi_1) + \lambda_2(\phi_1^\dagger\phi_2)(\phi_3^\dagger\phi_2) \\ + \lambda_3(\phi_2^\dagger\phi_3)(\phi_1^\dagger\phi_3) + \text{H.c.} \quad (6)$$

By construction, its global minimum is at $\langle\phi_i^0\rangle = (v/\sqrt{2}, 0, 0)$, where $v^2 = m_1^2/\lambda_0$. In order to find the mass matrices, we write the doublets as

$$\phi_1 = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v + h + iG^0) \end{pmatrix}, \quad \phi_2 = \begin{pmatrix} w_2^+ \\ z_2 \end{pmatrix}, \\ \phi_3 = \begin{pmatrix} w_3^+ \\ z_3 \end{pmatrix}. \quad (7)$$

Here, h is the SM-like Higgs boson, G^0 , and G^+ are the would-be Goldstone bosons, while w_2^+ , w_3^+ and z_2 , z_3 are charged and neutral Higgs bosons, respectively. Fields $w_{2,3}^+$ and $z_{2,3}$ have well-defined \mathbb{Z}_3 charges: $q = 1$ for w_2^+ , z_2 as well as for w_3^+ , z_3^* and $q = 2$ ($= -1 \bmod 3$) for w_3^+ , z_3 and w_2^- , z_2^* . Note that in contrast to the usual practice, we describe the neutral Higgs bosons in the second and third doublet by complex fields rather than pairs of neutral fields. The reason is that, by construction, the fields corresponding to the real and imaginary pairs of z 's have identical masses and coupling constants. In any process which can arise in this model, these two fields are emitted and absorbed simultaneously, so they can be described by a single complex field z_i .

The SM-like Higgs boson has mass $m_h^2 = 2m_1^2$, while the masses of the charged Higgs bosons are $m_{w_2^\pm}^2 = |m_2^2|$, $m_{w_3^\pm}^2 = |m_3^2|$. Neutrals with equal q can mix, which indeed happens at $\lambda_1 \neq 0$. We describe the resulting mass eigenstates by complex fields d and D ($m_d < m_D$), both having $q = 1$:

$$d = \cos\alpha z_2 + \sin\alpha e^{-i\beta} z_3^*, \\ D = -\sin\alpha e^{i\beta} z_2 + \cos\alpha z_3^*, \\ \tan 2\alpha = \frac{|\lambda_1|}{\lambda_0} \frac{m_1^2}{|m_2^2| - |m_3^2|}, \quad \beta = \arg\lambda_1, \\ m_{D,d}^2 = \frac{|m_2^2| + |m_3^2|}{2} \pm \frac{1}{2} \sqrt{(|m_2^2| - |m_3^2|)^2 + \frac{|\lambda_1|^2}{\lambda_0^2} m_1^4}. \quad (8)$$

Note that within this model, we have

$$m_d < m_{w_2^\pm}, \quad m_{w_3^\pm} < m_D, \\ m_{w_2^\pm}^2 + m_{w_3^\pm}^2 = m_d^2 + m_D^2. \quad (9)$$

The triple and quartic interaction terms arising from V_0 and $V_{\mathbb{Z}_3}$ specify the dynamics of the dark-matter candidates. The lightest particle from the second and third Higgs generations is d , and it is stabilized against decaying into the SM particles by the \mathbb{Z}_3 symmetry. As for heavier particles, triple interactions lead to their decays such as $D \rightarrow dh$, $D \rightarrow dZ$, and $D \rightarrow w_2^+ W^-$, $D \rightarrow w_3^- W^+$ if allowed kinematically. If the mass splitting between the d and D is small, then these processes involve virtual h , Z , etc., which then decay into the SM particles. In this aspect, D decays are similar to weak decays. Charged Higgs bosons $w_{2,3}^\pm$ will also decay to d or d^* plus SM particles.

In the case of symmetric dark matter, the main process leading to depletion of dark matter after electroweak symmetry breaking is the direct annihilation $dd^* \rightarrow X_{\text{SM}}$, and the semiannihilation reaction discussed below is only a correction to this process. Still, it might be possible that this correction leads to a sizable departure of the kinetics of the dark matter in the early Universe and affects the relic abundance at the freeze-out, [13].

The situation becomes more interesting in models of asymmetric dark matter [11], in which an asymmetry between d and d^* is generated at a higher energy scale. It is possible, for example, that upon electroweak phase transition, almost all d 's annihilate with d^* into the SM sector, leaving behind a certain concentration of dark-matter candidates d . In the present epoch, d can scatter elastically, $dd \rightarrow dd$ with $\sigma_{\text{el.}} \propto \lambda_0^2$, but they can also initiate semiannihilation processes such as $dd \rightarrow d^* X_{\text{SM}}$ with a subsequent annihilation of d^* with a d . This possibility originates from the following quartic terms in the scalar potential:

$$\frac{1}{\sqrt{2}} hddd \cos\alpha \sin\alpha e^{i\beta} (\lambda_2 \cos\alpha + \lambda_3^* \sin\alpha e^{i\beta}) + \text{H.c.} \quad (10)$$

Depending on λ 's, this process can be as efficient as the direct annihilation in the usual annihilating dark-matter models, or it can be suppressed by the small coupling constant.

Finally, the same interaction terms also generate the triple annihilation processes $ddd \rightarrow h \rightarrow X_{\text{SM}}$, whose rate is, however, suppressed at small densities with respect to the semiannihilation.

IV. AVOIDING SEMIANNIHILATION

The presence of the $hddd$ terms in the interaction Lagrangian in the previous example, which were responsible for the two-particle semiannihilation process, was due to the \mathbb{Z}_3 symmetry group. One can wonder whether two-particle semiannihilation can be avoided by employing a \mathbb{Z}_p group with larger p . In this present section, we show that it is indeed possible in a model with four Higgs doublets.

According to Ref. [20], one can encode in the 4HDM scalar sector any group Z_p with $p \leq 8$. Note that in order to avoid continuous symmetry, one must accompany the phase-symmetric part of the potential V_0 with at least three distinct terms transforming nontrivially under phase rotations. In Table I, we give the list of these symmetry groups together with examples of the three interaction terms and the phase rotations which generate the corresponding group.

Two remarks concerning this table are in order. First, we stress that all these groups are realizable, so that if the three terms in each line are written down with nonzero coefficients, there is no phase transformation other than the multiple of the generator and the common overall phase shift which leaves them invariant. Second, since the potential is invariant under the common phase shift of all doublets, one can freely add additional equal phases to the generators shown in the third column, and, in addition, one can permute the doublets. For example, the last line of this table can be replaced by $(3^\dagger 2)(1^\dagger 2)$, $(3^\dagger 1)(4^\dagger 1)$, $(4^\dagger 3)^2$, which is symmetric under the Z_8 group generated by phase rotations $\frac{2\pi}{8}(0, 1, 2, 6)$.

The patterns of phase shifts given in this table allow for construction of various dark sectors with different possibilities for dark-matter dynamics. Here, we do not aim at a complete classification of these possibilities but would like only to show that there are examples in 4HDM in which semiannihilation of dark-matter candidates is also forbidden.

To this end, let us consider the Z_7 -symmetric 4HDM with the potential $V = V_0 + V_{Z_7}$ with

$$V_{Z_7} = \lambda_1(\phi_4^\dagger \phi_1)(\phi_3^\dagger \phi_1) + \lambda_2(\phi_4^\dagger \phi_2)(\phi_1^\dagger \phi_2) + \lambda_3(\phi_4^\dagger \phi_3)(\phi_2^\dagger \phi_3) + \text{H.c.} \quad (11)$$

As before, we assume that only the first doublet couples to fermions. Using the technique of the previous section, one can easily construct the potential V_0 with the global minimum at $(v/\sqrt{2}, 0, 0, 0)$. Then, expanding the doublets similarly to Eq. (7), one observes that z_3 and z_4^* have Z_7 charges $q = 3$ and mix via the λ_1 term leading to mass eigenstates d and D . In addition, we have the field z_2 with Z_7 charge

$q = 2$ and electrically charged Higgs bosons $w_{2,3,4}^\pm$ with appropriate Z_7 charges.

By adjusting free parameters, one can easily make d the lightest among the particles which transform nontrivially under Z_7 . Then, the other particles will either eventually decay to d or d^* plus SM particles or will be stable representing an additional contribution to dark matter. Again, if the asymmetry between d or d^* exists and if the rate of their annihilation is high, then after the freeze-out, we are left with the gas predominantly made of d^* 's.

The subsequent microscopic dynamics depends on the interactions between d^* 's and z_2 's which follow from Eq. (11). The relevant terms are $d z_2 z_2$ times a SM field from the λ_2 term and $ddd z_2^*$ from the λ_3 term. One- or two-particle processes such as $d \rightarrow z_2^* z_2^*$, $dd \rightarrow d^* z_2$ and $dd \rightarrow z_2 z_2 z_2$ are all kinematically forbidden. Multiple collision kinetics depends on whether $m_{z_2} < 3m_d$ or not. If z_2 is not too heavy, then the ‘‘triple semiannihilation,’’ $ddd \rightarrow z_2 X_{\text{SM}}$, is kinematically allowed and will create a population of z_2 even if it were absent before. However, z_2 will get depleted by the semiannihilation process $z_2 z_2 \rightarrow d^* X_{\text{SM}}$. So, if one starts with a certain concentration of z_2 , d and their antiparticles, then z_2 will die off at a higher rate than d^* 's. In stationary conditions, the terminal concentrations will be those equilibrating the rates of the following 6d tree-level scattering with intermediate z_2 's:

$$6d \rightarrow z_2 z_2 X_{\text{SM}} \rightarrow d^* X_{\text{SM}}, \quad (12)$$

and the subsequent annihilation of d^* . The net result of this chain will be the ‘‘7d-burning process,’’ $7d \rightarrow X_{\text{SM}}$, the bottleneck in this chain being the triple- d process $ddd \rightarrow z_2 X_{\text{SM}}$.

On the other hand, if $m_{z_2} > 3m_d$, then $ddd \rightarrow z_2$ is kinematically forbidden, while the inverse process leads to a quick z_2 decay. In this case, one can still burn d^* 's via the tree-level process with intermediate virtual z_2 's:

$$dddd \rightarrow d^* d^* z_2 z_2 \rightarrow d^* d^* X_{\text{SM}}, \quad (13)$$

The net result will be the same 7d burning, but the bottleneck process is now the 4d collision, whose rate is even stronger suppressed.

V. DISCUSSION AND CONCLUSIONS

The main purpose of this paper is to demonstrate that there already exists a phenomenological template for scalar dark-matter models stabilized by cyclic groups larger than Z_2 . This template uses several electroweak Higgs doublets decoupled from fermions, and it represents one of the simplest extensions of the standard model. Remarkably, models with few doublets can easily accommodate dark sectors which are stabilized by a large list of discrete groups and which display various kinds of microscopic dynamics. In particular, we gave explicit examples of dark sectors where the bottleneck process leading to depletion of asymmetric dark matter can be a two-particle,

TABLE I. Cyclic groups realizable as symmetry groups in the scalar sector of 4HDM; $(a^\dagger b)$ is a short notation for $(\phi_a^\dagger \phi_b)$.

group	interaction terms	phase rotations
Z_2	$(1^\dagger 2), (1^\dagger 3), (1^\dagger 4)^2$	$\frac{2\pi}{2}(0, 0, 0, 1)$
Z_3	$(3^\dagger 2), (1^\dagger 3)(4^\dagger 3), (1^\dagger 4)(1^\dagger 2)$	$\frac{2\pi}{3}(0, 1, 1, 2)$
Z_4	$(3^\dagger 2), (1^\dagger 3)(4^\dagger 3), (1^\dagger 4)^2$	$\frac{2\pi}{4}(0, 1, 1, 2)$
Z_5	$(4^\dagger 3)(2^\dagger 3), (3^\dagger 2)(1^\dagger 2), (4^\dagger 1)(3^\dagger 1)$	$\frac{2\pi}{5}(0, 1, 2, 3)$
Z_6	$(4^\dagger 3)(2^\dagger 3), (3^\dagger 2)(1^\dagger 2), (1^\dagger 4)^2$	$\frac{2\pi}{6}(0, 1, 2, 3)$
Z_7	$(4^\dagger 1)(3^\dagger 1), (4^\dagger 3)(2^\dagger 3), (4^\dagger 2)(1^\dagger 2)$	$\frac{2\pi}{7}(0, 2, 3, 4)$
Z_8	$(4^\dagger 3)(2^\dagger 3), (4^\dagger 2)(1^\dagger 2), (1^\dagger 4)^2$	$\frac{2\pi}{8}(0, 2, 3, 4)$

three-particle or four-particle semiannihilation. We stress that these models do not require any serious fine-tuning. We only ask for the presence of terms invariant under the chosen symmetry group but do not constrain coefficients in front of these terms.

In certain aspects, these models resemble the inert doublet model [3,4], but in the other, they rely on symmetry patterns which arise only with several doublets. In this respect, such models can be viewed as “multi-inert” doublet models, although this name of course does not completely specify the microscopic dynamics.

Exploring the observational consequences of each sort of microscopic dynamics is a separate task. It should include study of the dark-matter kinetics in two situations. First, one obviously needs to track down the dark-matter evolution in the expanding Universe after the electroweak phase transition and determine the freeze-out abundances. Analysis of Ref. [13] already proves that semiannihilation processes can be important, but it remains to be understood how sensitive the evolution is to the exact microscopic dynamics.

We would like to stress that studying this problem in the context of multi-Higgs-doublet models can be much subtler than it looks at first glance due to multiple phase transitions near and below the electroweak scale. Indeed, even in the two-Higgs-doublet model, a single electroweak phase transition can split into a sequence of several phase transitions of different nature, both in the general case [21] and in the Inert doublet model [22]. One can expect that even longer chains of phase transitions can be possible in multidoublet models. Note that the last among these phase transitions can, in principle, happen at temperatures much lower than the nominal electroweak temperature scale. Consequently, the Universe might have evolved through a sequence of vacua with different, and perhaps exotic, properties. Phase transitions between these phases could have led to complete restructuring of the particle mass spectrum, both within the SM and in the dark sector; particles which are stable in one phase can be unstable in another. All these delicate details, as well as the thermodynamics of the phase transitions themselves, can modify

the evolution of the dark sector. None of the existing evolution codes can adequately address these intricacies.

The second situation where the microscopic dark-matter dynamics can make an impact is the present epoch evolution at astrophysical sites of elevated dark matter concentrations (Galactic centers, interiors of compact stars, etc. [23]). Since multiparticle processes are involved, the sensitivity to the dark-matter density will be different from that of the usual two-particle annihilating or with truly nonannihilating dark matter.

For example, it is known that dark matter with a sufficient elastic cross section can get captured inside neutron stars [23]. In models with asymmetric truly nonannihilating scalar dark matter, its accumulation can lead to formation of the Bose-Einstein condensate (which means that the occupation number in the phase space can become large) or even to collapsing in a tiny black hole [24]. In a certain region of parameter space, this black hole will destroy the host neutron star sufficiently quickly compared to the typical neutron star lifetime; therefore, this region is excluded by observations. In the case of multi-inert dark matter accumulated inside a neutron star, multiple annihilation processes will effectively enter the game as the density reaches a certain threshold, precluding black-hole formation and avoiding the above constraints.

In conclusion, we showed that multi-Higgs-doublet models can naturally accommodate scalar dark-matter candidates protected by the group \mathbb{Z}_p . For a model with N doublets, the values of p can be as large as 2^{N-1} . These models do not require any significant fine-tuning and can lead to a variety of forms of microscopic dynamics among the dark matter candidates (allowing or forbidding semiannihilation, offering different routes to multiparticle annihilation, etc.).

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