Pressure of massless hot scalar theory in the boundary effective theory framework

A. Bessa, F. T. Brandt, C. A. A. de Carvalho, and E. S. Fraga

¹Escola de Ciências e Tecnologia, Universidade Federal do Rio Grande do Norte, Caixa Postal 1524, 59072-970, Natal, RN, Brazil

²Instituto de Física, Universidade de São Paulo, Caixa Postal 66318, 05315-970, São Paulo, SP, Brazil

³Instituto de Física, Universidade Federal do Rio de Janeiro, Caixa Postal 68528, 21941-972, Rio de Janeiro, RJ, Brazil

(Received 9 November 2010; published 22 April 2011)

We use the boundary effective theory approach to thermal field theory in order to calculate the pressure of a system of massless scalar fields with quartic interaction. The method naturally separates the infrared physics, and is essentially nonperturbative. To lowest order, the main ingredient is the solution of the free Euler-Lagrange equation with nontrivial (time) boundary conditions. We derive a resummed pressure, which is in good agreement with recent calculations found in the literature, following a very direct and compact procedure.

DOI: 10.1103/PhysRevD.83.085024 PACS numbers: 11.10.Wx, 11.10.Gh

I. INTRODUCTION

In finite-temperature field theory it is known that a naive implementation of perturbation theory for the calculation of Feynman diagrams is ill-defined in the presence of massless bosons. This is due to the appearance of severe infrared divergences, brought about by the vanishing bosonic Matsubara mode in thermal propagators. These divergences plague the entire series, which becomes essentially meaningless [1]. As a result, one is forced to resort to resummation techniques that reorganize the perturbative series, and resum certain classes of diagrams, in order to extract sensible results. There are several ways of performing resummations and rewriting the degrees of freedom more efficiently in terms of quasiparticles; we refer the reader to the reviews [2–4] and to Ref. [5] for a discussion and a list of specific references. All such techniques are designed to partially tame the infrared divergences, creating a nonzero domain of validity for weak-coupling expansions. Nevertheless, the zero-mode problem remains, and the region of validity of resummed perturbative treatments can not be indefinitely enlarged.

In a recent paper [6], we have proposed an alternative approach to thermal field theories, denoted by boundary effective theory (BET). The central idea of the method is to respect the double integral structure of the partition function in the functional integral formalism,

$$Z = \int [D\phi_0(\mathbf{x})]\rho[\beta;\phi_0,\phi_0], \tag{1a}$$

where

$$\rho[\beta;\phi_0,\phi_0] = \int_{\phi(0,x) = \phi(\beta,x) = \phi_0(x)} [D\phi(\tau,x)] e^{-S[\phi]}$$
(1b)

is the diagonal element of the functional density matrix and $S[\phi]$ is the Euclidean action of the theory. This approach, based on the calculation of the density matrix, was already used in [7] to construct dimensionally-reduced effective

actions and in [5] to study the thermodynamics of scalar fields based on a semiclassical approximation. The functional density matrix formulation of quantum statistics was discussed in [8].

As it will be clear in the sequence of this article, the double integration approach is essentially different from the one where a single functional integration over periodic configurations $\phi(\tau, x)$ is performed. In the BET approach, the protagonist is the quantity $\phi_0(x)$ —the field eigenvalue in the functional Schrödinger field-representation: $\hat{\phi}|\phi_0(x)\rangle = \phi_0(x)|\phi_0(x)\rangle$. Indeed, any thermal observable can be constructed by integrating the appropriate functional of $\phi_0(x)$ over the fields $\phi_0(x)$ weighted by the corresponding diagonal element of the density matrix. The imaginary time evolution can be viewed as an intermediate step which calculates the weights for the effective theory of static ϕ_0 fields. That effective theory defines a quantum statistical problem which encodes all information about thermalization; one is led to compute correlations of the field $\phi_0(x)$, connecting the theory to physical quantities.

The field $\phi_0(x)$ has still another remarkable property: it is the zero (static) component of the dynamical field $\phi(\tau, x)$. Indeed, to each dynamical configuration $\phi(\tau, x)$ there corresponds a static configuration $\phi_0(x) = \phi(0, x) = \phi(\beta, x)$. We say that $\phi_0(x)$ is the (time) boundary value of $\phi(\tau, x)$. The difference $\phi(\tau, x) - \phi_0(x)$ vanishes at $\tau = 0$, β , and so it can be expanded in a sine-Fourier series with nonzero frequencies $\hat{\omega}_n = n\pi/\beta$, $n = 1, 2, \ldots$ From these considerations, we conclude that the effective theory encoded in $\rho[\beta; \phi_0, \phi_0]$ also contains all the infrared physics, and the double integral structure of Z naturally separates the potentially divergent modes.

The main result of Ref. [6] was the analytic calculation of the one-loop effective action for the boundary field in a scalar theory. Following the standard recipe, we used the saddle-point approximation but in two steps: first, we kept $\phi_0(x)$ fixed and expanded the action around the classical configuration $\phi_c[\phi_0]$, that has $\phi_0(x)$ as its boundary value, and integrated over quadratic fluctuations vanishing at

 $\tau=0,~\beta$; then, we expanded the resulting expression around the saddle point $\phi_0=0$, and took into account quadratic fluctuations of the boundary field. We showed that the one-loop effective action at finite temperature has the same expression as at zero temperature if written in terms of the classical field ϕ_c and if we trade free propagators at zero temperature by their finite temperature counterparts. Besides, we explained how to obtain a renormalized effective action in the case of a $\lambda \phi^4/4!$ theory.

Now we address the problem of computing the pressure of the single-well quartic theory in the (problematic) massless limit and testing the aforementioned advantages of our framework. The weak-coupling calculation of the pressure for the massless hot scalar theory is an enterprise of about 20 years so far [9–13], with recent results to order $g^8 \log g$ given in Ref. [14]. Moreover, attempts to reorganize the perturbative series have followed different paths, generally introducing one or more variational parameters in the pressure, to be maximized in the end. Among these nonperturbative methods we find screened perturbation theory (SPT), introduced in thermal field theory in Ref. [15], which can also be implemented in the framework called optimized perturbation theory (OPT) presented in Ref. [16], the linear δ expansion (LDE) [17], the so-called 2PI or Φ -derivable methods [18], and, more recently, a method based on nonperturbative renormalization group calculations [19]. All these methods, when applied to the thermodynamics of the scalar field, are remarkably more stable than the weakcoupling expansion (see Refs. [20,21] for results using SPT, Refs. [22] for LDE, and Refs. [23–26] for 2PI).

Because of the crucial role played by classical solutions $\phi_c[\phi_0]$ in the present approach (they are supposed to satisfy the Euler-Lagrange equation for arbitrary values of the coupling constant), the resulting pressure is essentially nonperturbative. Some technical difficulties arise, though. Decomposing the field as

$$\phi(\tau, \mathbf{x}) = \phi_c [\phi_0](\tau, \mathbf{x}) + \eta(\tau, \mathbf{x}),$$

where $\eta(0, x) = \eta(\beta, x) = 0$, introduces ultraviolet (UV) divergences in the calculation. In order to obtain a finite result, we use the renormalized effective action as derived in Ref. [6]. In addition, spurious UV divergences appear if one naively performs the saddle-point approximation and ignores higher order terms in the fluctuation $\eta(\tau, x)$ and in the fluctuations of the boundary field. A finite expression for the pressure is obtained by taking into account the first term in the expansion of the self-interaction. The procedure to avoid UV divergences can, in principle, be extended, allowing for a systematic calculation of higher order corrections.

One should distinguish the boundary field $\phi_0(x)$ from the conventional Matsubara zero mode. Indeed, the latter appears as the static component of a particular expansion, i.e., the Fourier expansion of the dynamical field in frequencies $\omega_n = 2n\pi/\beta$. The standard dimensionally-reduced effective theory for that field, a procedure known

in the literature as Dimensional Reduction (DR), is a high-temperature theory in character (see Refs. [12,27–31]). On the other hand, $\phi_0(x)$ is the physical field of the theory and, as such, its effective theory encompasses all the physical information on the system. Therefore, the procedure that we built (BET) yields an *alternative* dimensionally-reduced effective theory which is essentially different from DR.

We see that a *boundary field* rises naturally in the context of thermal theories. There, the word "boundary" precisely distinguishes and connects the field $\phi_0(x)$ to the associated dynamical field $\phi(\tau, x)$, which is the main quantity in conventional finite-temperature approaches. It should be clear that the use of that word has no relation to other uses, often in a topological sense, such as in holographic gauge-gravity duality, etc.

The structure of the paper is as follows: in Sec. II, we discuss the saddle-point approximation for the functional integration with fixed boundary configuration $\phi_0(x)$ and how to implement corrections to that approximation; besides, we write the partition function in terms of the effective action for the boundary field; in Sec. III, we perform the second functional integration using, again, the saddle-point method, and a renormalized expression for the pressure is obtained; finally, in Sec. IV, we present our conclusions.

II. RENORMALIZED PARTITION FUNCTION IN TERMS OF THE BOUNDARY FIELD

Let us consider the Euclidean action,

$$S[\phi] = \int_0^\beta (d^4x)_E \left[\frac{1}{2} \partial_\mu \phi \partial^\mu \phi + \frac{m_0^2}{2} \phi^2 + U(\phi) \right], \quad (2)$$

where $(d^4x)_E$ is a shorthand for $d\tau d^3x$. In this paper, $U(\phi) = \lambda \phi^4/4!$ and $m_0 = 0$. For single-well potentials like $U(\phi)$, the unique saddle point of the action S is the trivial vacuum $\phi = 0$. However, in the density matrix approach, the functional domain of integration is partitioned in classes where all field configurations have the same time boundary. The restriction of S to one of those classes (for instance, the one with boundary value $\phi_0(x)$) has a nontrivial saddle point $\phi_c[\phi_0]$, solution of

$$\Box_E \phi_c(x) + m_0^2 \phi_c(x) + U'(\phi_c(x)) = 0,(3a)$$
$$\phi_c(0, \mathbf{x}) = \phi_c(\beta, \mathbf{x}) = \phi_0(\mathbf{x}),(3b)$$

where $\Box_E = -(\partial_\tau^2 + \nabla^2)$ is the Euclidean d'Alembertian operator and x denotes (τ, x) . The first functional integration, being performed over configurations inside a certain class, will be dominated by fluctuations in the vicinity of the saddle point of that class. In particular, the correspondence $\phi_0 \mapsto \phi_c[\phi_0]$ is 1:1 for single-well potentials. Therefore, one obtains an optimized spanning of the domain of integration in the calculation of Z by collecting the contributions from quadratic fluctuations around a line of

saddle point configurations, as suggested by Fig. 1. One is naturally led to a two-fold saddle-point approximation. It is worth remarking that $\phi_c(\tau, \mathbf{x})$ is, in general, a nonperiodic function of τ in the sense that $\partial_\tau \phi_c(0, \mathbf{x}) \neq \partial_\tau \phi_c(\beta, \mathbf{x})$. One can verify it even in the simple case of a free theory. This is another important difference between BET and other methods.

The explicit dependence of ϕ_c on ϕ_0 is not known, except in very special cases. In [6], we obtained the following recursive relation for the classical solution,

$$\phi_{c}[\phi_{0}](\tau, \mathbf{x}) = \int d^{3}\mathbf{x}' \phi_{0}(\mathbf{x}') [\partial_{\tau'} G_{0}(\tau, \mathbf{x}; \tau', \mathbf{x}')]_{\tau'=0}^{\tau'=\beta}$$

$$- \int_{0}^{\beta} d\tau' \int d^{3}\mathbf{x}' G_{0}(\tau, \mathbf{x}; \tau', \mathbf{x}') U'(\phi_{c}(\tau', \mathbf{x}')),$$
(4)

where G_0 is a Green function of the free operator,

$$(\Box_E + m_0^2)G_0(x, x') = \delta^{(4)}(x - x')$$
 (5a)

$$G_0(\tau, \mathbf{x}; 0, \mathbf{x}') = G_0(\tau, \mathbf{x}; \beta, \mathbf{x}') = 0.$$
 (5b)

For a fixed boundary configuration, the fluctuations η around ϕ_c vanish at $\tau=0$ and β ,

$$\phi(\tau, \mathbf{x}) = \phi_c(\tau, \mathbf{x}) + \eta(\tau, \mathbf{x}), \qquad \eta(0, \mathbf{x}) = \eta(\beta, \mathbf{x}) = 0.$$
(6)

In terms of η , the renormalized partition function reads

$$Z_{R} = \int [D\phi_{0}(\mathbf{x})] \int_{\eta(0,\mathbf{x})=\eta(\beta,\mathbf{x})=0} [D\eta(\tau,\mathbf{x})] e^{-S[\phi_{c}+\eta]+C.T},$$
(7)

where C.T. are counterterms to be chosen. One can think of each ϕ_c as a background field around which the dynamics of the fields η takes place.

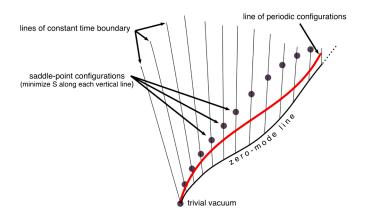


FIG. 1 (color online). Pictorial description of the partition of the functional domain of integration for Z in the BET approach. The whole domain is indexed by the zero-mode line. A given configuration $\phi(\tau, x)$ lives on the vertical fiber over the boundary value $\phi(0, x) = \phi(\beta, x)$ in the bottom. Inside each vertical line, the Euclidean action is minimized by the saddle-point configuration (bullet). The line of periodic configurations is also shown.

In the vicinity of ϕ_c , the action is approximately quadratic,

$$S[\phi_c + \eta] = S[\phi_c] + \frac{1}{2} \int (d^4x)_E \eta(x) [\Box_E + m_0^2 + U''(\phi_c(x))] \eta(x) + \mathcal{O}(\eta^3).$$
 (8)

It is convenient to introduce the Green function:

$$[\Box_E + m_0^2 + U''(\phi_c(x'))]G[\phi_c](x; x') = \delta^{(4)}(x - x') \quad (9a)$$

$$G[\phi_c](\tau, x; 0, x') = G[\phi_c](\tau, x; \beta, x') = 0. \quad (9b)$$

In particular, $G[0] = G_0$, the free propagator defined in Eq. (5). The Gaussian integration formally yields

$$Z_R[\beta] \approx \int [\mathcal{D}\phi_0(\mathbf{x})] e^{-S[\phi_c] + \text{C.T}} (\det G[\phi_c])^{1/2}. \quad (10)$$

We should mention that the present saddle-point approximation can be good even in a strong-coupling regime. In fact, the classical solution is supposed to be exact for arbitrary values of the coupling constant. Besides, one can systematically improve the saddle-point approximation by expanding the cubic and quartic η interactions and calculating diagrams with lines of $G[\phi_c]$:

$$Z_R[\beta] = \int [\mathcal{D}\phi_0(\mathbf{x})] e^{-S[\phi_c] + \text{C.T.}} (\det G[\phi_c])^{1/2} e^{-A[\phi_c]},$$
(11)

where

$$e^{-A[\phi_c]} = 1 - \frac{\lambda}{8} \int (d^4x)_E G^2[\phi_c](x, x) + \frac{C_1}{2} \int (d^4x)_E G[\phi_c](x, x) + \mathcal{O}(3 \text{ loops}), \quad (12)$$

with C_1 being a mass counterterm. The corresponding loop expansion is often called semiclassical series.

In [6], we have shown that the renormalized 1-loop effective action for the boundary field ϕ_0 is

$$\beta \Gamma_R[\phi_0] = S[\phi_c] + \frac{1}{2} \operatorname{Tr} \log(G^{-1}[\phi_c]) + \frac{1}{2} \operatorname{Tr} \log(\mathcal{C}[\phi_c]) - \text{C.T.}, \tag{13}$$

where

$$C[\phi_c] = \frac{\delta^2 S[\phi_c]}{\delta \phi_0^2} = [\partial_\tau \partial_{\tau'} G[\phi_c]]_0^\beta \tag{14}$$

is the 1-loop contribution from quadratic fluctuations of the field ϕ_0 .

Comparing (11) and

$$e^{-\beta\Gamma_{R}[\phi_{0}]} = e^{-S[\phi_{c}] + \text{C.T.}} (\det G[\phi_{c}])^{1/2} (\det \mathcal{C}[\phi_{c}])^{-1/2},$$
(15)

one can write

BESSA et al.

$$Z_R[\beta] = (\det \mathcal{C}_0)^{1/2} \int [\mathcal{D}\phi_0(\mathbf{x})] e^{-\beta \Gamma_R[\phi_0]} e^{-S_I[\phi_c]}, \quad (16)$$

where $C_0 = C[0]$ and

$$S_I[\phi_c] = A[\phi_c] - \frac{1}{2}\operatorname{Tr}\log(\mathcal{C}[\phi_c]) + \frac{1}{2}\operatorname{Tr}\log(\mathcal{C}_0). \tag{17}$$

In this paper, the terms involving S_I and the interacting part of $S[\phi_c]$ will be dealt with order by order in a loop expansion that will be discussed in the next section.

III. PRESSURE IN THE BET APPROACH

Our strategy is to perform the second functional integration using again the saddle-point approximation. The saddle point of $\Gamma_R[\phi_0]$ is the trivial configuration $\phi_0=0$. Therefore, it is natural to expand Γ_R in terms of the following n-point functions:

$$\Gamma_{R}[\phi_{0}] = \sum_{n=1}^{\infty} \frac{1}{n!} \int \Gamma_{R}^{(n)}(\mathbf{x}_{1},...,\mathbf{x}_{n}) \phi_{0}(\mathbf{x}_{1}) ... \phi_{0}(\mathbf{x}_{n}) d^{3}\mathbf{x}_{1} ... d^{3}\mathbf{x}_{n},$$
(18)

where

$$\Gamma_R^{(n)}(\mathbf{x}_1, \dots, \mathbf{x}_n) = \frac{\delta^{(n)} \Gamma_R[\phi_0]}{\delta \phi_0(\mathbf{x}_1) \dots \delta \phi_0(\mathbf{x}_n)} \bigg|_{\phi_0 = 0}.$$
(19)

Up to quadratic order in ϕ_0 , we have (see Ref. [6])

$$\beta\Gamma_{R}[\phi_{0}] \approx \beta\Gamma_{R}^{(0)} + \frac{1}{2} \int \beta\Gamma_{R}^{(2)}(\mathbf{x}_{1}, \mathbf{x}_{2})\phi_{0}(\mathbf{x}_{1})\phi_{0}(\mathbf{x}_{2})d^{3}x_{1}d^{3}x_{2},$$
(20)

where $\Gamma_R^{(0)} = -V\pi^2/(90\beta^4)$ is essentially *minus* the pressure of an ideal gas of free massless bosons, V is the volume, and

$$\frac{\beta}{(2\pi)^3 \delta(\mathbf{p}_1 + \mathbf{p}_2)} \Gamma_R^{(2)}(\mathbf{p}_1, \mathbf{p}_2; \mu)
= 2|\mathbf{p}_1| \tanh \frac{\beta|\mathbf{p}_1|}{2} + \beta m^2(|\mathbf{p}_1|; \beta),$$
(21)

with

$$\beta m^2(k;\beta) = \frac{\lambda}{24\beta} \frac{\tanh \beta k/2}{\beta k} \left(1 + \frac{\beta k}{\sinh \beta k} \right). \tag{22}$$

The mass counterterm chosen was

$$C_1 = \frac{\lambda}{2} \int^{\Lambda} \frac{d^4 q}{(2\pi)^4} \Delta_F^0(q), \tag{23}$$

where $\Delta_F^0(q) = 1/q^2$ is the zero-temperature (massless) free propagator in four-dimensional Euclidean Fourier space.

Substituting Eq. (20) in (16) with $S_I = 0$, and performing the quadratic integration over ϕ_0 , one obtains the saddle-point approximation for Z,

$$Z_{\rm sp}[\beta] = e^{-\beta \Gamma_R^{(0)}} (\det \mathcal{C}_0)^{1/2} (\det \beta \Gamma_R^{(2)})^{-1/2} = e^{\beta V P_{\rm sp}}.$$
 (24)

In Appendix A, we show that

$$P_{\rm sp}(\beta) = \frac{\pi^2}{90\beta^4} - \frac{1}{2\beta} \lim_{\Lambda \to \infty} \int_{-\infty}^{\Lambda} \frac{d^3k}{(2\pi)^3} \log\left(1 + \beta m^2(k;\beta)\right) \times \frac{\coth\beta k/2}{2k}.$$
 (25)

The second term on the right hand side (rhs) of Eq. (25) can be identified with a series of daisy diagrams, where the petal is given by (22). As one can easily check, the $\mathcal{O}(\lambda)$ term in $P_{\rm sp}$ is UV divergent. The 2-loop diagram carries the divergence. We will show that such a spurious divergence is removed when we consistently add the remaining 2-loop corrections to the saddle-point approximation.

We have to be careful to identify the good propagator to represent the contraction of two ϕ_0 fields. We know that such a propagator should be calculated at the saddle point $\phi_0 = 0$. However, it is not obvious if the interacting mass should enter or not its definition. We claim that, in order to be consistent with the one-loop calculation of the effective action, we must use \mathcal{C}_0 defined in (14) at $\phi_c = 0$ as the propagator.

Notice that all quantities in this calculation depend on ϕ_0 through ϕ_c . Therefore, it is convenient to define the contraction of two fields ϕ_c . In [7], it was shown that the propagator $\hat{\mathcal{C}}_0$ which is associated with that contraction satisfies

$$\Delta_F = G_0 + \hat{\mathcal{C}}_0, \tag{26}$$

where Δ_F is the usual free thermal propagator, and G_0 is defined in Eq. (5). A pragmatic argument in favor of using lines of \mathcal{C}_0 (or $\hat{\mathcal{C}}_0$) to build 2-loop corrections is that it solves the problem of UV divergences of the saddle-point approximation, reproducing the correct result for the $\mathcal{O}(\lambda)$ perturbative contribution. We show that in Appendix B.

Finally, the renormalized pressure is given by

$$\begin{split} P_{\mathrm{BET}}(\beta) &= \frac{\pi^2}{90\beta^4} - \frac{\lambda}{1152\beta^4} - \frac{1}{2\beta} \lim_{\Lambda \to \infty} \int^{\Lambda} \frac{d^3k}{(2\pi)^3} \Big\{ \log \Big(1 \\ &+ \frac{\beta m^2(k;\beta)}{2} \frac{\coth \beta k/2}{k} \Big) \\ &- \frac{\beta m^2(k;\beta)}{2} \frac{\coth \beta k/2}{k} \Big\}. \end{split} \tag{27}$$

Figures 2–4 plot the pressure normalized by the ideal pressure as a function of $g = \sqrt{\lambda}$ obtained using different methods. Figure 2 compares the pressure from BET with that from weak-coupling calculations up to $g^8 \log g$, according to Ref. [14]. The weak-coupling expansion already includes resummation from order g^3 on. We see that the BET approach is in good agreement with the most recent results.

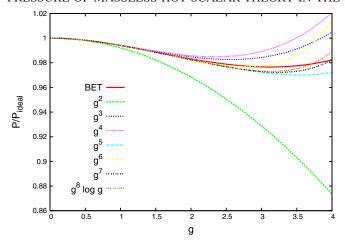


FIG. 2 (color online). Pressure normalized by P_{ideal} as a function of the coupling constant $g = \sqrt{\lambda}$ in BET and weak-coupling (see [14]) formalisms. The renormalization scale is $\mu = 2\pi T$.

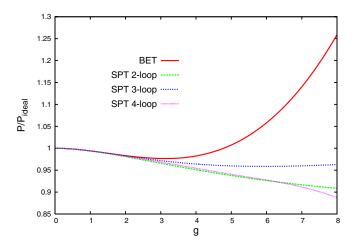


FIG. 3 (color online). Comparison of the normalized pressure obtained using BET and SPT calculations at two, three, and four loops reported in [21]. The renormalization scale is $\mu = 2\pi T$.

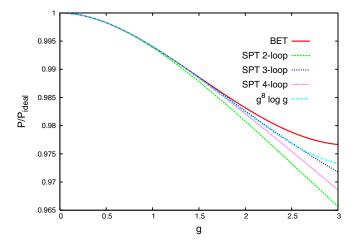


FIG. 4 (color online). Comparison of BET, $g^8 \log g$ weak-coupling (Ref. [14]), and SPT (Ref. [21]) normalized pressures in the range [0, 3]. The renormalization scale is $\mu = 2\pi T$.

Figures 3 and 4 compare BET with screening perturbation theory (SPT) calculations at two, three, and four loops from Ref. [21] over different ranges. In Fig. 3, the coupling g goes from 0 to 8. It shows a complete mismatch between BET and the SPT curves for, say, $g \gtrsim 5$. However, in that range the SPT curves are not reliable either.

In Fig. 4, we add the $g^8 \log g$ weak-coupling curve to the comparison. Around g = 3, SPT curves present quite a large oscillation as one goes from two to three and then to four loops. We conclude that, in the range where SPT shows convergence, the curve from BET seems to behave remarkably well.

IV. CONCLUSIONS

The boundary effective theory was introduced as an alternative approach to quantum statistical mechanics. One of its main features is providing a natural separation of the zero-mode (static) sector, leading to the construction of its effective theory which results from integrating over all the (imaginary time) dynamical modes. In previous work, the one-loop effective action for the zero mode had been calculated. In the present article, we have shown that BET is also a powerful method to attack a crucial problem of the thermodynamics of bosonic fields: infrared divergences.

The different strategies currently known to deal with such IR problems rely on some sort of resummation of diagrams of naive perturbation theory. The present calculation of the pressure using BET has the advantage of performing a highly nontrivial resummation in the scope of a natural and systematic procedure. In fact, the effective theory for the zero mode generated the whole series of daisy diagrams very naturally. Besides, in contrast with those built from thermal field theory, the daisy diagrams built in the context of BET fit quite well recent results in the literature obtained using rather involved techniques of screening perturbation theory at 4 loops.

A distinctive feature of BET is the role played by field configurations which are not strictly periodic in the imaginary time. Indeed, following the double integral scheme for the partition function, we have shown that there is a line of saddle-point configurations dominating the functional integral and all of them, but the trivial one, are nonperiodic.

We saw that the separation of the field in its static and dynamical parts could lead to problems in the ultraviolet limit. This technical point has already been addressed in the calculation of the one-loop effective action. The solution is to perform a parallel calculation of the two functional integrations in the definition of the partition function. That led to a prescription for the present calculation which, we hope, can serve as a guide to extend the method to higher orders.

An interesting extension of the present work would be the application of BET to a scalar theory with N field components. The pressure per degree of freedom in the large-N limit is exactly known and coincides with the twoloop SPT result (see Ref. [23]). In order to calculate the pressure in that limit using BET, one needs the corresponding renormalized effective action. A preliminary analysis indicates that the full calculation involves subtle issues regarding the nontrivial features of the semiclassical propagator and will be postponed to a future publication.

ACKNOWLEDGMENTS

E. S. F. thanks J. O. Andersen and J. Pawlowski for fruitful discussions. The authors also thank J. O. Andersen and L. Kyllingstad for providing tables for their results for the pressure for comparison. This work was partially supported by CAPES, CNPq, FAPERJ, FAPESP and FUJB/UFRJ.

APPENDIX A

In [6], we show that $C_0(k) = 2k \tanh(k\beta/2)$. Expressing log det C_0 as Tr log C_0 , we obtain

$$\log \det \mathcal{C}_0 = \int_0^{\Lambda} \frac{d^3k}{(2\pi)^3} \log[2k \tanh(k\beta/2)]. \tag{A1}$$

Analogously, from (21) it follows that

$$\log \det \beta \Gamma_R^{(2)} = \int_{-\infty}^{\infty} \frac{d^3k}{(2\pi)^3} \log[2k \tanh \beta k/2 + \beta m^2(k;\beta)]. \tag{A2}$$

Therefore,

 $\log[(\det \mathcal{C}_0)^{1/2}(\det \boldsymbol{\beta} \Gamma_R^{(2)})^{-1/2}]$

$$= -\frac{1}{2} \int_{-\infty}^{\infty} \frac{d^3k}{(2\pi)^3} \log\left(1 + \beta m^2(k;\beta) \frac{\coth\beta k/2}{2k}\right).$$
 (A3)

Using that expression in (24), one obtains (25).

APPENDIX B

We start by expanding the interacting terms directly in (11):

$$e^{-\int (d^4x)_E U(\phi) + \text{C.T.}} \approx 1 - \frac{\lambda}{24} \int (d^4x)_E \phi_c^4(x) + \frac{C_1}{2} \int (d^4x)_E \phi_c^2(x)$$
 (B1)

and

$$e^{-A[\phi_c]} \approx 1 - \frac{\lambda}{8} \int (d^4x)_E G_0^2(x, x) + \frac{C_1}{2} \int (d^4x)_E G_0(x, x).$$
(B2)

Contracting the fields ϕ_c using the proper symmetry factors, we obtain

$$Z_{\text{BET}}[\beta] \approx Z_{\text{sp}}[\beta] \left[1 - \frac{\lambda}{8} \int (d^4 x)_E (G_0^2(x, x) + \hat{C}_0^2(x, x)) + \frac{C_1}{2} \int (d^4 x)_E (G_0(x, x) + \hat{C}_0(x, x)) \right].$$
(B3)

The 2-loop contribution from $Z_{\rm sp}[\beta]$ is obtained expanding (see Ref. [6])

$$(\det G[\phi_c])^{1/2} \approx (\det G_0)^{-1/2} \left[1 - \frac{\lambda}{4} \int (d^4 x)_E G_0(x, x) \phi_c^2(x) \right],$$
 (B4)

and contracting the fields ϕ_c . Collecting the 2-loop contributions for $P_{\text{BET}} = (\log Z_{\text{BET}})/\beta V$, we obtain

$$P_{2\text{-loop}} = -\frac{\lambda}{8\beta V} \int (d^4x)_E (G_0(x, x) + \hat{C}_0(x, x))^2 + \frac{C_1}{2\beta V} \int (d^4x)_E (G_0(x, x) + \hat{C}_0(x, x)).$$
(B5)

Using (26), we have

$$P_{2\text{-loop}} = -\frac{\lambda}{8\beta V} \int (d^4x)_E \left(\Delta_F(x, x) - \frac{2C_1}{\lambda}\right)^2 + D,$$
(B6)

where D is a zero-temperature infinite constant which can be set to zero. Finally, using (23) and performing the remaining integration, we obtain that the $\mathcal{O}(\lambda)$ contribution to the pressure is finite and reproduces the perturbative result

$$P_{2\text{-loop}} = -\frac{\lambda}{1152\beta^4}.$$
 (B7)

- [1] M. Le Bellac, *Thermal Field Theory* (Cambridge University Press, Cambridge, England, 2000); J.I. Kapusta and C. Gale, *Finite-Temperature Field Theory: Principles and Applications* (Cambridge University Press, Cambridge, England, 2006); A. Das, *Finite Temperature*
- Field Theory (World Scientific, Singapore, 1997); J. Zinn-Justin, *Quantum Field Theory and Critical Phenomena* (Oxford University Press, Oxford, England, 1993).
- [2] J. P. Blaizot, E. Iancu, and A. Rebhan, arXiv:hep-ph/ 0303185.

- [3] U. Kraemmer and A. Rebhan, Rep. Prog. Phys. 67, 351 (2004).
- [4] J. O. Andersen and M. Strickland, Ann. Phys. (N.Y.) 317, 281 (2005).
- [5] A. Bessa, C. A. A. de Carvalho, E. S. Fraga, and F. Gelis, J. High Energy Phys. 08 (2007) 007.
- [6] A. Bessa, F.T. Brandt, C.A.A. de Carvalho, and E.S. Fraga, Phys. Rev. D 82, 065010 (2010).
- [7] C. A. A. de Carvalho, J. M. Cornwall, and A. J. da Silva, Phys. Rev. D 64, 025021 (2001).
- [8] A. Bessa, C. A. A. de Carvalho, and E. S. Fraga, Phys. Rev. E 81, 011103 (2010).
- [9] R. R. Parwani, Phys. Rev. D **45**, 4695(E) (1992); **48**, 5965 (1993).
- [10] J. Frenkel, A. V. Saa, and J. C. Taylor, Phys. Rev. D 46, 3670 (1992).
- [11] R.R. Parwani and H. Singh, Phys. Rev. D 51, 4518 (1995).
- [12] E. Braaten and A. Nieto, Phys. Rev. D 51, 6990 (1995).
- [13] A. Gynther, M. Laine, Y. Schroder, C. Torrero, and A. Vuorinen, J. High Energy Phys. 04 (2007) 094.
- [14] J. O. Andersen, L. Kyllingstad, and L. E. Leganger, J. High Energy Phys. 08 (2009) 066.
- [15] F. Karsch, A. Patkos, and P. Petreczky, Phys. Lett. B 401, 69 (1997).
- [16] S. Chiku and T. Hatsuda, Phys. Rev. D 58, 076001 (1998).
- [17] A. Okopinska, Phys. Rev. D 35, 1835 (1987); A. Duncan and M. Moshe, Phys. Lett. B 215, 352 (1988).

- [18] J. M. Luttinger and J. C. Ward, Phys. Rev. 118, 1417 (1960); G. Baym, Phys. Rev. 127, 1391 (1962); J. M. Cornwall, R. Jackiw, and E. Tomboulis, Phys. Rev. D 10, 2428 (1974).
- [19] J. P. Blaizot, A. Ipp, and N. Wschebor, Nucl. Phys. A 849, 165 (2011).
- [20] J. O. Andersen, E. Braaten, and M. Strickland, Phys. Rev. D 63, 105008 (2001).
- [21] J.O. Andersen and L. Kyllingstad, Phys. Rev. D 78, 076008 (2008).
- [22] M. B. Pinto and R. O. Ramos, Phys. Rev. D 60, 105005 (1999); 61, 125016 (2000); R. L. S. Farias, G. Krein, and R. O. Ramos, Phys. Rev. D 78, 065046 (2008).
- [23] J. P. Blaizot, E. Iancu, and A. Rebhan, Phys. Rev. Lett. 83, 2906 (1999); Phys. Lett. B 470, 181 (1999); Phys. Rev. D 63, 065003 (2001).
- [24] A. Peshier, Phys. Rev. D 63, 105004 (2001).
- [25] E. Braaten and E. Petitgirard, Phys. Rev. D 65, 041701 (2002); 65, 085039 (2002).
- [26] J. Berges, S. Borsanyi, U. Reinosa, and J. Serreau, Phys. Rev. D 71, 105004 (2005).
- [27] P. H. Ginsparg, Nucl. Phys. B 170, 388 (1980).
- [28] T. Appelquist and R. D. Pisarski, Phys. Rev. D 23, 2305 (1981).
- [29] S. Nadkarni, Phys. Rev. D 27, 917 (1983).
- [30] N. P. Landsman, Nucl. Phys. B 322, 498 (1989).
- [31] K. Kajantie, M. Laine, K. Rummukainen, and M.E. Shaposhnikov, Nucl. Phys. B 458, 90 (1996).