W/Z bremsstrahlung as the dominant annihilation channel for dark matter

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Dark matter annihilation to leptons, $\chi\chi \rightarrow \ell \bar{\ell}$, is necessarily accompanied by electroweak radiative corrections, in which a *W* or *Z* boson is radiated from a final-state particle. Given that the *W* and *Z* gauge bosons decay dominantly via hadronic channels, it is thus impossible to produce final-state leptons without accompanying protons, antiprotons, and gamma rays. Significantly, while many dark matter models feature a helicity-suppressed annihilation rate to fermions, radiating a massive gauge boson from a final-state fermion removes this helicity suppression, such that the branching ratios Br($\ell\nu W$), Br($\ell^+\ell^-Z$), and Br($\nu\bar{\nu}Z$) dominate over Br($\ell\bar{\ell}$). *W*/*Z* bremsstrahlung thus allows indirect detection of many weakly interacting massive particle models that would otherwise be helicity suppressed, or ν^2 suppressed. Antiprotons and even antideuterons become consequential final-state particles. This is an important result for future dark matter searches. We discuss the implications of *W*/*Z* bremsstrahlung for "leptonic" dark matter models which aim to fit recent cosmic ray positron and antiproton data.

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I. INTRODUCTION

An abundance of cosmological and astrophysical evidence attests to the existence of dark matter (DM), whose presence is inferred via its gravitational influence [1–3]. However, the fundamental particle properties of DM remain essentially unknown. One important means of probing DM's particle nature is via indirect detection, whereby we search for products of DM annihilation (or decay) emanating from regions of DM concentration in the Universe today.

The dark matter annihilation cross section is often parametrized as $\langle v\sigma_A \rangle = a + bv^2 + \cdots$, where $\langle v\sigma_A \rangle$ is the thermally averaged annihilation cross section. The constant *a* comes from *s*-wave annihilation, while the velocity suppressed bv^2 term receives both *s*-wave and *p*-wave contributions; the *L*th partial wave contribution to the annihilation rate is suppressed as v^{2L} . Given that $v \sim 10^{-3}c$ in galactic halos, even the *p*-wave contribution is highly suppressed and thus only the *s*-wave annihilation rate of the sequence of the universe today. However, in many DM models the *s*-wave annihilation into a fermion pair $\chi\chi \to \bar{f}f$ is helicity suppressed by a factor $(m_f/M_\chi)^2$ (only $\to \bar{t}t$ modes remain of interest, and then only for a certain range of χ mass).

When computing DM annihilation signals, it is normally assumed that only the lowest order tree-level processes make a significant contribution. However, there are important exceptions to this statement. Dark matter annihilation into charged particles, $\chi \chi \rightarrow \bar{f} f$, is necessarily accompanied by the internal bremsstrahlung process $\chi \chi \rightarrow \bar{f} f \gamma$, where the photon may be radiated from one of the external particle legs (final-state radiation, FSR) or, possibly, from a virtual propagator (virtual internal bremsstrahlung, VIB). On the face of it, the radiative rate is down by the usual QED coupling factor of $\alpha/\pi \sim 500$. However, and significantly, photon bremsstrahlung can lift the helicity suppression of the *s*-wave process [4], which more than compensates for the extra coupling factor. Such a striking enhancement can arise when a symmetry of the initial state $\chi\chi$ is satisfied by the three-body final state $\bar{f}f\gamma$, but not by the two-body final state $\bar{f}f$. For bremsstrahlung of photons, only VIB is effective in lifting the helicity suppression, as FSR is dominated by soft or collinear photons (such that the two- and three-body final states have the same symmetry properties) as discussed in Ref. [5].

In this paper we examine electroweak bremsstrahlung [6–12], i.e., bremsstrahlung of Z or W^{\pm} electroweak gauge bosons to produce $\bar{f}fZ$ and $\bar{\ell}\nu W$ final states. The virtue for W/Z bremsstrahlung to lift initial-state velocity and final-state helicity suppressions, alluded to in [8,11], has not been previously explored. We show that W/Z bremsstrahlung can also lift suppression and become the dominant annihilation channel. Thus, W/Z bremsstrahlung allows indirect detection of many weakly interacting massive particle models that would otherwise be helicity suppressed, or ν^2 suppressed. This is an important result for future DM searches.

There are a number of important distinctions between electromagnetic (EM) and electroweak (EW) bremsstrahlung. An obvious one is that EM bremsstrahlung produces just photons, whereas EW bremsstrahlung and subsequent decay of the gauge bosons leads to leptons, hadrons, and gamma rays, offering correlated "multimessenger" signals for indirect dark matter searches. Another distinction is that W/Z bremsstrahlung from final-state particles (FSR) is sufficient to lift a suppression. This is due to the nonzero gauge boson masses, and the coupling of the gauge bosons

to the nonconserved axial current which leads to a different form for the polarization sum than in the case of the photon (or a gluon in the similar QCD process). In contrast, for the EM process, VIB is required for the photon to lift a suppression. Because an additional propagator appears for VIB, suppression-lifting EM bremsstrahlung is itself suppressed by an additional factor of M_{χ}^2/M_{η}^2 relative to electroweak's FSR, where M_{η} is the mass of the internal exchange particle. Only in the event of a near-degeneracy $M_{\chi} \sim M_{\eta}$ is this relative suppression of EM bremsstrahlung negligible.

DM annihilation to charged leptons has been the subject of much recent attention, due to recently measured cosmic ray anomalies which point to an excess of cosmic ray positrons above those that may be attributed to conventional astrophysical processes. PAMELA has observed a sharp excess in the $e^+/(e^- + e^+)$ fraction at energies beyond approximately 10 GeV [13], without a corresponding excess in the antiproton/proton data [14,15], while Fermi and HESS have reported more modest excesses in the $(e^- + e^+)$ flux at energies of order 1 TeV [16]. These signals have led to a reexamination of positron production in nearby pulsars [17], emission from supernova remnants [18], acceleration of e^+e^- in cosmic ray sources [19], and propagation in conventional cosmic ray models [20]. As an alternative to these astrophysical mechanisms, it has also been proposed that the excess e^+ and e^- are produced via dark matter annihilation in the galactic halo, with an abundance of DM models proposed to accomplish this end. A recent overview of e^{\pm} -excess data and possible interpretations is available in [21].

However, some of the most popular models suffer from helicity or v^2 suppression. A prototypical example of suppressed production of standard model (SM) fermion pairs is provided by supersymmetry: Majorana neutralinos annihilate into a pair of SM fermions via *t*- and *u*-channel exchange of SU(2)-doublet sfermions. To overcome the suppression, proponents of these models have invoked large "boost" factors. These boost factors may be astrophysical in origin, as with postulated local overdensities of dark matter, or they may arise from particle physics, as with the Sommerfeld enhancement that arises from light scalar exchange between dark matter particles. Although not ruled out, these factors do seem to be a contrivance designed to overcome the innate suppression.

A further problem with suppressed models is the overproduction of antiprotons from unsuppressed W/Z bremsstrahlung. Given that hadronic decay modes of the W and Z bosons will lead to significant numbers of both antiprotons and gamma rays, this will impact the viability of models that might otherwise have explained the observed positron excess. Even in models which do not feature a suppression, the W/Z bremsstrahlung has important phenomenological consequences, as the decay products of the gauge bosons make a pure leptonic e^+e^- signal impossible [11]. In Sec. II we discuss the circumstances under which dark matter annihilation may be suppressed, and in Sec. III explain how W/Z bremsstrahlung is able to circumvent such a suppression. In Sec. IV we consider a representative model, and explicitly calculate the cross sections for both the lowest order annihilation process, and for the W/Z bremsstrahlung process. We discuss implications of these results in Sec. V. Calculational details are collected in five Appendixes.

II. UNDERSTANDING SUPPRESSION USING FIERZ TRANSFORMATIONS

In this section we describe the origin of v^2 and helicity suppressions. We shall make use of Fierz transformation and partial wave decomposition to determine under what circumstances these suppressions will or will not arise.

Dark matter candidates may be scalar, fermionic, or vector in nature; if fermionic, they may be either Dirac or Majorana. Permissible annihilation models include s-, t-, and u-channel exchanges of a new particle, and the various possibilities are listed in Refs. [22–24]. In every case, it is useful to classify the partial waves available to the decay process, and to analyze the dependence on the mass of the SM particle pair in the final state. In this article, we focus on fermionic Majorana dark matter.

For fermionic dark matter, the natural projection of $2 \rightarrow 2$ processes into partial waves makes use of the Fierz transformation. In the next subsection we consider DM annihilation via the process $\chi \chi \rightarrow \bar{f}f$, and explain the use of Fierz transforms to convert the matrix elements for t/u-channel annihilation, which are of the form $(\bar{\chi}\Gamma_A l) \times$ $(\bar{l}\Gamma_B\chi)$, to a sum of s-channel amplitudes of the form $(\bar{\chi}\Gamma_1\chi)(l\Gamma_2 l)$. In the following subsection we then categorize the Fierzed s-channel amplitudes into partial waves and fermion-pair spin states, which determines whether the amplitudes are velocity suppressed, mass suppressed, or unsuppressed. In the third and final subsection, we put our findings together to determine which class of models will have a suppressed $2 \rightarrow 2$ annihilation. We show that in a certain popular class of suppressed models, the $2 \rightarrow 3$ W/Z-bremsstrahlung process is unsuppressed, and in fact dominant for $2M_{\chi} > M_W$. We will find in Sec. III that a generalization of the Fierz transformation offers useful insight into the nonsuppression of the $2 \rightarrow 3$ process.

A. Fierz transformations in the chiral basis

Helicity projection operators are essential in chiral gauge theories, so it is worth considering the reformulation of Fierz transformations in the chiral basis [25]. (A discussion of standard Fierz transformations may be found in, e.g. Ref. [26].) We place hats above the generalized Dirac matrices constituting the chiral basis. These matrices are

$$\{\hat{\Gamma}^B\} = \{P_R, P_L, P_R \gamma^{\mu}, P_L \gamma^{\mu}, \frac{1}{2} \sigma^{\mu\nu}\}, \text{ and} \\ \{\hat{\Gamma}_B\} = \{P_R, P_L, P_L \gamma_{\mu}, P_R \gamma_{\mu}, \frac{1}{2} \sigma_{\mu\nu}\},$$
(1)

where $P_R \equiv \frac{1}{2}(1 + \gamma_5)$ and $P_L \equiv \frac{1}{2}(1 - \gamma_5)$ are the usual helicity projectors. Notice that the dual of $P_R \gamma^{\mu}$ is $P_L \gamma_{\mu}$, and the dual of $P_L \gamma^{\mu}$ is $P_R \gamma_{\mu}$. The tensor matrices in this basis contain factors of $\frac{1}{2}$: $\hat{\Gamma}^T = \frac{1}{2}\sigma^{\mu\nu}$ and $\hat{\Gamma}_T = \frac{1}{2}\sigma_{\mu\nu}$. These facts result from the orthogonality and normalization properties of the chiral basis and its dual, as explained in detail in Appendix A.

$$\begin{pmatrix} (P_R)[P_R] \\ (P_L)[P_L] \\ (\hat{T})[\hat{T}] \\ (\gamma_5 \hat{T})[\hat{T}] \\ (P_R)[P_L] \\ (P_R)[P_L] \\ (P_L)[P_R] \\ (P_L \gamma^{\mu})[P_R \gamma_{\mu}] \\ (P_L \gamma^{\mu})[P_R \gamma_{\mu}] \\ (P_L \gamma^{\mu})[P_R \gamma_{\mu}] \\ (P_L \gamma^{\mu})[P_L \gamma_{\mu}] \end{pmatrix} = \frac{1}{4} \begin{pmatrix} 2 & 0 & 1 & 1 \\ 0 & 2 & 1 & -1 \\ 6 & 6 & -2 & 0 \\ 6 & -6 & 0 & 2 \\ & & & & \\ 6 & -6 & 0 & 2 \\ & & & & \\ & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & &$$

Nonexplicit matrix elements in (3) are zero, and we have introduced a shorthand \hat{T} for either $\hat{\Gamma}^T = \frac{1}{2}\sigma^{\mu\nu}$ or $\hat{\Gamma}_T = \frac{1}{2}\sigma_{\mu\nu}$.

The importance of this transformation for us is that it converts *t*-channel and *u*-channel exchange graphs into *s*-channel form, for which it is straightforward to evaluate the partial waves. The block-diagonal structures, contained in the middle matrix, show that "mixing" occurs only within the subsets $\{P_R \otimes P_R, P_L \otimes P_L, \hat{T} \otimes \hat{T}, \gamma_5 \hat{T} \otimes \hat{T}\}$, and $\{P_R \otimes P_L, P_R \gamma^{\mu} \otimes P_L \gamma_{\mu}\}$. The Fierz transform matrix is idempotent, meaning its square is equal to the identity matrix. This follows from the fact that two Fierz rearrangements return the process to its initial ordering. A consequence of the block-diagonal form is that each sub-block is itself idempotent.

In Eq. (3) we have included one nonmember of the basis set, namely $\gamma_5 \hat{T}$; it is connected to \hat{T} via the relation

$$\gamma_5 \sigma^{\mu\nu} = \frac{i}{2} \epsilon^{\mu\nu\alpha\beta} \sigma_{\alpha\beta}. \tag{4}$$

Explicit use of $\gamma_5 \hat{T}$ in Eq. (3) is an efficient way to express the chiral Fierz transformation.

So far we have not used the qualifier in the assumption that the dark matter is Majorana. Majorana particles are invariants under charge conjugation C, which implies that vector and tensor bilinears are disallowed. Another way of understanding this is to note that interchanging the two identical Majorana particles in a *t*-channel diagram generates an accompanying *u*-channel diagram with a relative minus sign (from fermion anticommutation). When Fierzed, these two amplitudes cancel for *V* and *T* couplings Using completeness of the basis (see Appendix A), one arrives at a master formula which expands the outer product of two chiral matrices in terms of their Fierzed forms:

$$(\hat{\Gamma}^D)[\hat{\Gamma}_E] = \frac{1}{4} \operatorname{Tr}[\hat{\Gamma}^D \hat{\Gamma}^C \hat{\Gamma}_E \hat{\Gamma}_B](\hat{\Gamma}^B][\hat{\Gamma}_C), \qquad (2)$$

where the parentheses symbols are a convenient shorthand for matrix indices [27] (see Appendix A for details). Evaluating the trace in Eq. (2) leads to the Fierz transformation matrix in the chiral basis:



(exactly so in the four-Fermi limit where the differing momenta in the *t*- and *u*-channel propagators can be ignored—refer to Appendix B for details). We must thus drop *V* and *T* couplings appearing in the Fierzed bilinears of the χ current.

B. Origin of v^2 and helicity suppressions

One can use partial wave decomposition (see, e.g., the textbooks [28–30], or the convenient summary in the Appendix of [8]) to expand the scattering amplitudes as a sum of angular momentum components. Partial waves do not interfere, and the *L*th partial wave contribution to the total cross section σv is proportional to v^{2L} . The annihilating χ particles are very nonrelativistic today, so an unsuppressed *s* wave (L = 0), if present, will dominate the annihilation cross section. The DM virial velocity within our Galaxy is about 10^{-3} (in units of *c*), leading to a suppression of $v^2 \sim 10^{-6}$ for *p*-wave processes.

On the other hand, the SM fermions produced in the $2 \rightarrow 2$ annihilation are highly relativistic (except possibly for $t\bar{t}$ production). For many annihilation channels the spin state of the fermion pair gives rise to a helicity suppression by a factor of $(m_l/M_\chi)^2$, where m_l is the fermion mass.

Unfortunately, many popular models for annihilation of Majorana dark matter to charged leptons are subject to one or more of these two suppressions, the v^2 and/or $(m_\ell/M_\chi)^2$ suppressions. This includes some of the models proposed to accommodate the positron and e^+e^- excesses observed in PAMELA, Fermi-LAT, and HESS data. In Sec. III, we show that in the class of models which have suppressed rates for $\chi\chi \rightarrow \ell^+\ell^-$, the $2 \rightarrow 3$ graph obtained by adding

PHYSICAL REVIEW D 83, 013001 (2011)

TABLE I. Extreme nonrelativistic and extreme-relativistic limits for *s*-channel bilinears. In order for a term with an initial-state DM bilinear and a final-state lepton bilinear to remain unsuppressed, the DM bilinear must have a nonzero entry in the appropriate cell of the "w = 0 limit" columns, and the lepton bilinear must have a nonzero term in the appropriate cell of the "M = 0 limit" columns. Otherwise, the term is suppressed. (The tensor and pseudotensor are not independent, but rather are related by $\gamma_5 \sigma^{\mu\nu} = \frac{i}{2} \epsilon^{\mu\nu\alpha\beta} \sigma_{\alpha\beta}$.) We recall that antiparallel spinors correspond to parallel particle spins (and antiparallel particle helicities for the M = 0 current), and vice versa. Amplitudes are shown for $\bar{u}\Gamma_D v = [\bar{v}\Gamma_D u]^*$. The twofold \pm ambiguities reflect the twofold spin assignments for parallel spins, and separately for antiparallel spins.

s-channel bilinear $\bar{\Psi}\Gamma_D\Psi$		v	= 0 limit	M = 0 limit				
		Parallel spinors	Antiparallel spinors	Parallel spinors	Antiparallel spinors			
Scalar	$\bar{\Psi}\Psi$	0	0	\sqrt{s}	0			
Pseudoscalar	$\bar{\Psi}i\gamma_5\Psi$	-2iM	0	$-i\sqrt{s}$	0			
Axial vector	$\bar{\Psi}\gamma_5\gamma^0\Psi$	2M	0	0	0			
	$\bar{\Psi}\gamma_5\gamma^j\Psi$	0	0	0	$\sqrt{s}(\pm\delta_{i1}-i\delta_{i2})$			
Vector	$\bar{\Psi}\gamma^0\Psi$	0	0	0	0			
	$\bar{\Psi} \gamma^j \Psi$	$\mp 2M\delta_{i3}$	$-2M(\delta_{i1} \mp i\delta_{i2})$	0	$-\sqrt{s}(\delta_{i1} \mp i\delta_{i2})$			
Tensor	$ar{\Psi}\sigma^{0j}\Psi$	$\mp 2iM\delta_{i3}$	$-2iM(\delta_{i1} \pm \delta_{i2})$	$-i\sqrt{s}\delta_{i3}$	0			
	$ar{\Psi}\sigma^{jk}\Psi$	0	0	$\pm \sqrt{s} \delta_{i1} \delta_{k2}$	0			
Pseudotensor	$\bar{\Psi}\gamma_5\sigma^{0j}\Psi$	0	0	$\pm i\sqrt{s\delta_{i3}}$	0			
	$ar{\Psi} \gamma_5 \sigma^{jk} \Psi$	$\mp 2M\delta_{j1}\delta_{k2}$	$-2M(\delta_{j2}\delta_{k3}\mp i\delta_{j3}\delta_{k1})$	$-\sqrt{s}\delta_{j1}\delta_{k2}$	0			

a radiative W^{\pm} or Z to the final-state particles of the $2 \rightarrow 2$ graph becomes dominant. The radiated W's and Z's will decay to, among other particles, antiprotons. Since an excess generation of antiprotons is not observed by PAMELA, this class of models is ruled out by the present work.

Consider products of *s*-channel bilinears of the form $(\bar{\chi}\Gamma_1\chi)(\bar{l}\Gamma_2 l)$. To further address the question of which products of currents are suppressed and which are not, we may set v^2 to zero in the χ current, and m_{ℓ}^2 to zero in the lepton current, and ask whether the product of currents is suppressed. If the product of currents is nonzero in this limit, the corresponding amplitude is unsuppressed. In Table I we give the results for the product of all standard Dirac bilinears. (The derivation of these results is outlined in Appendix C.) Suppressed bilinears enter this table as zeros.¹

One can read across rows of this table to discover that the only unsuppressed s-channel products of bilinears for the $2 \rightarrow 2$ process are those of the pseudoscalar, vector, and tensor. [For completeness, we also show results for the pseudotensor bilinears, although the pseudotensor is not independent of the tensor, as a result of Eq. (4).] For Majorana dark matter, the vector and tensor bilinears are disallowed by charge-conjugation arguments and one is left with just the unsuppressed pseudoscalar.

C. Class of models for which $\chi \chi \to \ell \bar{\ell}$ annihilation is suppressed

We now put the results of the previous two subsections together to explain which class of models has a v^2 and/or $(m_\ell/M_\chi)^2$ suppressed $2 \rightarrow 2$ annihilation. We have seen that, for Majorana DM, *s*-channel annihilation with a *P* coupling is unsuppressed, while *S* and *A* contributions are suppressed (and *V* and *T* forbidden). Let us now consider *t*-channel or *u*-channel processes.

Any *t*-channel or *u*-channel diagram that Fierz's to an *s*-channel form containing a pseudoscalar coupling will have an unsuppressed L = 0 *s*-wave amplitude. From the matrix in Eq. (3), one deduces that such will be the case for any *t*- or *u*-channel current product on the left side which finds a contribution in the 1st, 2nd, 5th, or 7th columns of the right side. This constitutes the *t*- or *u*-channel tensor, same-chirality scalar, and opposite-chirality vector products (rows 1–4, and 6 and 8 on the left). On the other hand, the *t*- or *u*-channel opposite-chirality scalars or same-chirality vectors (rows 5, 7, 9, and 10 on the left) do *not* contain a pseudoscalar coupling after Fierzing to *s*-channel form. Rather, it is the suppressed axial vector and vector (Dirac fermions only) that appear.

Interestingly, a class of the most popular models for fermionic dark matter annihilation to charged leptons falls into this latter, suppressed, category. It is precisely the opposite-chirality *t*- or *u*-channel scalar exchange that appears in these models, an explicit example of which will be discussed below. Thus it is rows 5 and 7 in Eq. (3) that categorize the model we will analyze. After Fierzing to *s*-channel form, it is seen that the Dirac bilinears are opposite-chirality vectors (i.e., *V* or *A*). Dropping the vector term from the χ current we see that

¹It is seen that the only bilinears in the table without velocity suppression are those of the pseudoscalar, the three-vector part of the vector, the zeroth component of the axial vector, and the time-space part of the tensor (or equivalently, the space-space part of the pseudotensor). It is also seen that the only bilinears without fermion mass suppression are the scalar, pseudoscalar, three-vector parts of the vector and axial vector, and the tensor.

the $2 \rightarrow 2$ process couples an axial vector χ current to a relativistic SM fermion current which is an equal mixture of *A* and *V*. Accordingly, this model has an *s*-wave amplitude occurring only in the L = 0, J = 1, S = 1 channel, with the spin flip from S = 0 to S = 1 (or equivalently, the mismatch between zero net chirality and one unit of helicity) costing a fermion mass insertion and a $(m_f/M_{\chi})^2$ suppression in the rate.

Let us pause to explain why this *t*- or *u*-channel scalar exchange with opposite fermion chiralities at the vertices is so common. It follows from a single popular assumption, namely, that the dark matter is a gauge-singlet Majorana fermion. As a consequence of this assumption, annihilation to SM fermions, which are SU(2) doublets or singlets, requires either an *s*-channel singlet boson or a *t*- or *u*-channel singlet or doublet scalar that couples to χ -*f*. In the first instance, there is no symmetry to forbid a new force between SM fermions, a disfavored possibility. In the second instance, unitarity fixes the second vertex as the Hermitian adjoint of the first. Since the fermions of the SM are left-chiral doublets and right-chiral singlets, one gets chiral opposites for the two vertices of the *t* or *u* channel.

Supersymmetry provides an analog of such a model. In this case the dark matter consists of Majorana neutralinos, which annihilate to SM fermions via the exchange of ("right"- and "left"-handed) SU(2)-doublet slepton fields. In fact, the implementation in 1983 of supersymmetric photinos as dark matter provided the first explicit calculation of *s*-wave suppressed Majorana dark matter [31]. However, the class of models described above is more general than the class of supersymmetric models.

To illustrate our arguments, we choose a simple example of the class of model under discussion. This is provided by the leptophilic model proposed in Ref. [32] by Cao, Ma, and Shaughnessy. Here the DM consists of a gauge-singlet Majorana fermion χ which annihilates to leptons via the SU(2)-invariant interaction term

$$f(\nu\ell^{-})_{L}\varepsilon \left(\begin{array}{c} \eta^{+} \\ \eta^{0} \end{array} \right) \chi + \text{H.c.} = f(\nu_{L}\eta^{0} - \ell_{L}^{-}\eta^{+}) \chi + \text{H.c.},$$
(5)

where f is a coupling constant, ε is the 2 × 2 antisymmetric matrix, and (η^+, η^0) form the new SU(2) doublet scalar which mediates the annihilation. (This model was originally discussed in Ref. [33], and an expanded discussion of its cosmology may be found in Ref. [34].)

As discussed above, the *u*- and *t*-channel amplitudes for DM annihilation to leptons, of the form $(\bar{\chi}P_L l)(\bar{l}P_R\chi)$, become pure $(\bar{\chi}P_L\gamma^{\mu}\chi)(\bar{l}P_R\gamma_{\mu}l)$ under the chiral Fierz transformation. The product of the Majorana and fermion bilinears then leads to an AA term and an AV term. However, reference to Table I shows that neither of these terms leads to an unsuppressed amplitude: in all cases, either the lepton bilinear is suppressed by m_ℓ , the DM bilinear by v, or both are suppressed. Thus, Majorana DM annihilation to a lepton pair is suppressed in this model, in accordance with the explicit calculation in Ref. [32].

III. LIFTING THE SUPPRESSION

Allowing the lepton bilinear to radiate a W or Z boson (as shown in Fig. 1) does yield an unsuppressed amplitude. In the rate, there will be the usual radiative suppression factor of $\frac{\alpha_2}{4\pi} \sim 10^{-3}$. But, this will be partially compensated by a 3-body phase space factor $\sim (M_{\chi}/M_W)^2/8\pi^2$ relative to 2-body massless phase space, which exceeds unity for dark matter masses exceeding \sim TeV.² More importantly, the v^2 suppression for Majorana annihilation to 2-body final states will be lifted by the 3-body W-bremsstrahlung process. In Sec. IV we show, by explicit calculation, that the $2 \rightarrow 3$ radiative process that leads to antiprotons dominates for any M_{χ} that allows the W to be produced on shell, i.e., for $2M_{\chi} > M_W$.

The next inevitable question is "Why is the radiative $2 \rightarrow 3$ process unsuppressed?" To answer this question, we invoke a more general Fierz rearrangement applicable to $2 \rightarrow N$ processes, $N \ge 3$. The relevant equation, derived in Appendix A, states that any 4×4 matrices X and Y may be expressed as

$$(\mathbf{X})[\mathbf{Y}] = (\mathbf{X}\mathbb{1})[\mathbb{1}\mathbf{Y}] = \frac{1}{4}(\mathbf{X}\Gamma^{B}\mathbf{Y}][\Gamma_{B})$$
$$= \frac{1}{4^{2}}\operatorname{Tr}[\mathbf{X}\Gamma^{B}\mathbf{Y}\Gamma_{C}](\Gamma^{C}][\Gamma_{B}),$$
(6)

where the Dirac matrices here are taken in the standard basis defined in Eq. (A1).

From Table I we see that setting Γ^C to $\gamma_5 \gamma^0$, the only structure available to a nonrelativistic Majorana current other than the pseudoscalar, and Γ_B to either γ^j or $\gamma_5 \gamma^j$, provides an unsuppressed product of the Majorana dark matter bilinear and the charged-lepton bilinear. Moreover, for the W/Z-bremsstrahlung process, X and Y in the general Fierz equation are the un-Fierzed couplings P_L and $q^{-2}P_R q P_L \epsilon$, respectively. So we will have shown that the radiative process is unsuppressed if we can show that $q^{-2} \text{Tr}[P_L (\gamma^j \text{ or } \gamma_5 \gamma^j) P_R q P_L \epsilon \gamma_5 \gamma_0]$ is unsuppressed. This trace reduces to $q^{-2} \text{Tr}[P_R \gamma_0 \gamma^j \epsilon q]$. The expansion of this trace as scalar products contains terms such as $q_0 \cdot \epsilon^j$ and $(\epsilon \times q)^j$, which are nonzero and unsuppressed by fermion masses. Thus, the $2 \rightarrow 3$ process contains an unsuppressed *s*-wave amplitude.

Physically, the unsuppression works because the gauge boson carries away a unit of angular momentum, allowing a fermion spin flip such that there is no longer a mismatch

²When $M_{\chi}^2 \gg M_W^2$, the rate for single W production is dominated by infrared and collinear divergences, leading to a suppressed factor $\ln^2(\frac{s}{4M_W^2}) - 3\ln(\frac{s}{4M_W^2})$ [10,35] instead of our $(\frac{s}{4M_W^2})$. Moreover, the rate for multiple production of W's becomes so large that resummation techniques are necessary.



FIG. 1. *t*-channel (a) and (c) and *u*-channel (b) and (d) contributions to $\chi \chi \rightarrow e^+ \nu W^-$. Emission from the scalar propagator is not included, as it is suppressed by $1/M_{\eta}^2$. Note that all fermion momenta flow with the arrow except p_2 , so $q_1 = p_1 + Q$, $q_2 = -p_2 - Q$.

between the chirality of the leptons and their allowed twoparticle spin state.

One may ask why emission of a gamma ray rather than a W/Z boson is less effectual. It has been known for some time [4,5] that gamma-ray emission in the final state does produce an unsuppressed s-wave contribution, but at second order rather than lowest order in the inverse mass-squared M_{η}^{-2} of the *t*- and *u*-channel exchange particle(s). The reason is that gamma-ray emission from the final-state fermions (FSR) is dominated by infrared and collinear singularities, each of which puts the intermediate lepton on shell (virtuality $q^2 \rightarrow 0$). Including the q^{-4} from the squared propagator in the phase space integral [see Eq. (20)], one gets the factor $\int_{M_V^2}^{s} \frac{dq^2}{q^6} (s - q^2)(q^2 - M_V^2)$, where M_V is the mass of the radiated boson (photon or Wor Z). For a gamma ray, with $M_V^2 = 0$, one readily sees the infrared and collinear singularities in $\int_0 \frac{dq^2}{q^4}$. An on-shell particle is observable, so the spin states of the $q^2 \rightarrow 0$ intermediate fermion do not interfere. Thus, as $q^2 \rightarrow 0$, the trace for gamma emission, $\operatorname{Tr}[\gamma_0 \gamma^j \epsilon q] =$ $\operatorname{Tr}[\gamma_0 \gamma^j (P_R^2 + P_L^2) \notin q]$ goes over to $\operatorname{Tr}[\gamma_0 \gamma^j P_R] \times$ $\operatorname{Tr}[P_R \notin q] + \operatorname{Tr}[\gamma_0 \gamma^j P_L] \operatorname{Tr}[P_L \notin q]$. The first trace in each term of this sum vanishes. Consequently, the gammaemission amplitude remains suppressed at order M_{η}^{-2} . However, at order M_{η}^{-4} , the gamma ray may be emitted from the internal particle η (VIB). For VIB, phase space does not favor $q^2 = 0$, and an unsuppressed amplitude results.

The emission of a massive W (or Z) boson contrasts significantly from the emission of a massless photon. With the W emission, the relevant phase space integral over virtuality q^2 is $\int_{M_W^2}^s \frac{dq^2}{q^6}(s-q^2)(q^2-M_W^2)$. The minimum virtuality of the intermediate fermion is $q^2 = M_W^2$, and the mean virtuality for $s \gg M_W^2$ is greater again by the factor $2 \ln(s/M_W^2)$. With no infrared or collinear singularities for W/Z emission, an unsuppressed amplitude results already at order M_{η}^{-2} .

Before looking at an explicit example in which electroweak bremsstrahlung is seen to lift a suppression, we pause to summarize some important facts for the $2 \rightarrow 2$ annihilation process:

- (i) Fierz transformation is used to reexpress *t* and *u*-channel amplitudes of the form $(\bar{\chi}\hat{\Gamma}^A l)(\bar{l}\hat{\Gamma}_B\chi)$ as a sum of *s*-channel (not to be confused with *s*-wave) amplitudes of the form $(\bar{\chi}\hat{\Gamma}^C\chi)(\bar{l}\hat{\Gamma}_D l)$.
- (ii) For Majorana dark matter, only S, P, and A s-channel bilinears are allowed, with the V and T bilinears forbidden by the self-conjugate properties of Majorana particles.
- (iii) Considering the product of an *s*-channel χ current with an *s*-channel fermion current, we find that the pseudoscalar is the only member of the set (S, P, A) which is unsuppressed. The other combinations are either helicity (m_{ℓ}/M_{χ}) or velocity (v) suppressed.
- (iv) The annihilation process $\chi \chi \rightarrow \ell \bar{\ell}$ via *t* and *u*-channel exchange of a scalar is suppressed.

Importantly, electroweak bremsstrahlung lifts this suppression at lowest order in the propagator mass squared $(M_{\eta}^{-2}$ in amplitude), whereas photon bremsstrahlung lifts the suppression at the next order $(M_{\eta}^{-4}$ in amplitude).

Amplification of the latter remark is the purpose of this paper.

IV. EXPLICIT CALCULATION OF SUPPRESSION LIFTING WITH ELECTROWEAK BREMSSTRAHLUNG

To explicitly demonstrate that emission of a W^{\pm} or Z boson does lift helicity suppression, we calculate the cross section for $\chi \chi \rightarrow e^{\pm \nu} W^{\pm}$ below in the leptophilic model of Ref [32]. The interaction term for this model is that given above in Eq. (5).

A. Example of helicity-suppressed rate

In the model of Ref. [32], the cross section for the $2 \rightarrow 2$ process $\chi \chi \rightarrow e^+ e^-$ or $\nu \bar{\nu}$ with Majorana DM is given as

$$\upsilon\sigma = \frac{f^4 \upsilon^2 r^2}{24\pi M_{\chi}^2} (1 - 2r + 2r^2),\tag{7}$$

where $m_l \simeq 0$ and $M_{\eta^{\pm}} = M_{\eta^0}$ have been assumed, and $r = M_{\chi}^2/(M_{\eta}^2 + M_{\chi}^2)$. The suppressions discussed in Sec. II are apparent in Eq. (7). The helicity-suppressed *s*-wave term is absent in the $m_l = 0$ limit, and thus only the v^2 -suppressed term remains.

This $2 \rightarrow 2$ cross section can be calculated by inclusion of two Feynman diagrams, a *t*-channel exchange of η and the associated *u*-channel exchange obtained by crossing the Majorana particles. The relative sign between the graphs is negative, due to the fermion exchange. Summing and squaring, one has three terms including the interference term. Alternatively, one may Fierz transform the fermion bilinears in the two contributing amplitudes. The relative minus sign is compensated by the special Majorana minus sign described in Eq. (B2). Reference to Eq. (3) then shows that one gets $(P_L)[P_R] \rightarrow \frac{1}{2}(P_L \gamma^{\mu}] \times$ $[P_R \gamma_\mu) \times 2$, where the final factor of 2 counts the two contributing amplitudes, which are identical in the four-Fermi limit $M_{\eta}^2 \gg t$ and u. We are left with just one amplitude, $\frac{f^2}{M_{\eta}^2} [\bar{v}(k_2)(\frac{1}{2}\gamma_5)v(p_2)][\bar{u}(p_1)P_L\gamma_{\mu}v(p_2)]]$. The surviving Dirac structure for the Majorana current is pure axial vector, since the vector (and tensor) part of a Majorana current vanishes. With just a single product of bilinears, the remaining part of the $2 \rightarrow 2$ calculation is straightforward. One arrives at

$$\upsilon \sigma = \frac{f^4 M_{\chi}^2}{16\pi M_{\eta}^4} \left[\frac{m_l^2}{s} + \frac{2}{3} \upsilon^2 + \mathcal{O}(\upsilon^4) \right], \tag{8}$$

in agreement with the four-Fermi, $m_{\ell} = 0$ limit of Eq. (7). Here, the helicity suppression of the *s*-wave amplitude, proportional to a helicity flip, in turn proportional to a mass insertion, is manifest.

B. *W* emission and unsuppressed *S* wave

We now turn to the calculation of the cross section for the process $\chi\chi \rightarrow e^+ \nu W^-$ (equal to that for $\chi\chi \rightarrow e^- \bar{\nu}W^+$). The four contributing Feynman diagrams are shown in Fig. 1. Note that we consider bremsstrahlung only from the final-state particles (FSR), and neglect emission from the virtual scalar (VIB). Strictly speaking, the distinction between FSR and VIB is somewhat artificial in the sense that the partition depends upon the choice of gauge. However, we shall work in unitary gauge, in which emission from the internal line is suppressed by a further power of M_{η}^2 due to the additional scalar propagator; consequently, we expect our results to be valid to order M_{η}^{-2} in amplitude, i.e. order M_{η}^{-4} in rate.

We retain the assumptions $m_l \simeq 0$ and $M_{\eta^{\pm}} = M_{\eta^0}$. The matrix element for the top-left diagram is

$$\mathcal{M}_{a} = \frac{igf^{2}}{\sqrt{2}q_{1}^{2}} \frac{1}{t_{1} - M_{\eta}^{2}} (\bar{\upsilon}(k_{2})P_{L}\upsilon(p_{2})) \\ \times (\bar{\upsilon}(p_{1})\gamma^{\mu}P_{L}q_{1}u(k_{1}))\epsilon_{\mu}^{Q}, \qquad (9)$$

where $i \frac{g}{\sqrt{2}} \gamma^{\mu} P_L$ is the coupling at the $\ell \nu W$ vertex, and t_1, t_2, u_1 , and u_2 are the standard Mandelstam variables,

$$t_{1} = (k_{1} - q_{1})^{2} = (p_{2} - k_{2})^{2},$$

$$t_{2} = (k_{1} - p_{1})^{2} = (-q_{2} - k_{2})^{2},$$

$$u_{1} = (k_{2} - q_{1})^{2} = (p_{2} - k_{1})^{2},$$

$$u_{2} = (k_{2} - p_{1})^{2} = (-q_{2} - k_{1})^{2}.$$

(10)

Upon applying Eq. (6) to Fierz transform the matrix element, we obtain

$$\mathcal{M}_{a} = \frac{igf^{2}}{\sqrt{2}q_{1}^{2}} \frac{1}{t_{1} - M_{\eta}^{2}} \epsilon_{\mu}^{Q} \frac{1}{4} [(\bar{v}(k_{2})u(k_{1})) \\ \times (\bar{u}(p_{1})P_{L}\gamma^{\mu}P_{L}\dot{q}_{1}v(p_{2})) \\ + (\bar{v}(k_{2})\gamma_{5}u(k_{1}))(\bar{u}(p_{1})P_{L}\gamma_{5}\gamma^{\mu}P_{L}\dot{q}_{1}v(p_{2})) \\ + (\bar{v}(k_{2})\gamma_{5}\gamma_{\alpha}u(k_{1}))(\bar{u}(p_{1})\gamma^{\alpha}\gamma^{\mu}P_{L}\dot{q}_{1}v(p_{2}))] \\ = \frac{igf^{2}}{\sqrt{2}q_{1}^{2}} \frac{1}{t_{1} - M_{\eta}^{2}} \epsilon_{\mu}^{Q} \frac{1}{4} (\bar{v}(k_{2})\gamma_{5}\gamma_{\alpha}u(k_{1})) \\ \times (\bar{u}(p_{1})P_{L}\gamma^{\alpha}\gamma^{\mu}\dot{q}_{1}v(p_{2})).$$
(11)

The first two terms after the first equality are zero due to the helicity projection operators, leaving only an axialvector term. (Vector and tensor χ bilinears have been omitted, as they will cancel between *u*- and *t*-channel diagrams in the heavy M_n limit, as discussed above.) Note that although this matrix element resembles that of an *s*-channel annihilation process, the γ matrices in the lepton bilinear would be in a different order for a true *s*-channel annihilation process involving W/Z bremsstrahlung from one of the final-state leptons.

Similarly, the matrix element for the top-right diagram can be written as

$$\mathcal{M}_{b} = \frac{-igf^{2}}{\sqrt{2}q_{1}^{2}} \frac{1}{u_{1} - M_{\eta}^{2}} \frac{1}{4} (\bar{v}(k_{2})\gamma_{5}\gamma_{\alpha}u(k_{1})) \\ \times (\bar{u}(p_{1})P_{L}\gamma^{\alpha}\gamma^{\mu}\dot{q}_{1}v(p_{2}))\epsilon_{\mu}^{Q}, \qquad (12)$$

and those for the bottom diagrams,

$$\mathcal{M}_{c} = \frac{-igf^{2}}{\sqrt{2}q_{2}^{2}} \frac{1}{t_{2} - M_{\eta}^{2}} \frac{1}{4} (\bar{\upsilon}(k_{2})\gamma_{5}\gamma_{\alpha}u(k_{1})) \\ \times (\bar{\upsilon}(p_{1})P_{L}\dot{q}_{2}\gamma^{\mu}\gamma^{\alpha}\upsilon(p_{2}))\epsilon_{\mu}^{Q},$$
(13)

$$\mathcal{M}_{d} = \frac{igf^{2}}{\sqrt{2}q_{2}^{2}} \frac{1}{u_{2} - M_{\eta}^{2}} \frac{1}{4} (\bar{v}(k_{2})\gamma_{5}\gamma_{\alpha}u(k_{1})) \\ \times (\bar{u}(p_{1})P_{L}\dot{q}_{2}\gamma^{\mu}\gamma^{\alpha}v(p_{2}))\epsilon_{\mu}^{Q}.$$
(14)

Performing the sum over spins and polarizations, we find

$$\sum_{\text{spin,pol.}} |\mathcal{M}|^{2} = \sum_{\text{spin,pol.}} |(\mathcal{M}_{a} + \mathcal{M}_{c}) - (\mathcal{M}_{b} + \mathcal{M}_{d})|^{2} \\ = \left(\frac{gf^{2}}{\sqrt{2}}\right)^{2} \frac{1}{16} \operatorname{Tr}[(k_{2} + M_{\chi})\gamma_{\alpha}(k_{1} + M_{\chi})\gamma_{\beta}] \left(g_{\mu\nu} - \frac{Q_{\mu}Q_{\nu}}{M_{W}^{2}}\right) \left(\frac{1}{q_{1}^{4}} \left(\frac{1}{t_{1} - M_{\eta}^{2}} + \frac{1}{u_{1} - M_{\eta}^{2}}\right)^{2} \operatorname{Tr}[\not{p}_{1}\gamma^{\alpha}\gamma^{\mu}\not{q}_{1}\not{p}_{2}\dot{q}_{1}\gamma^{\nu}\gamma^{\beta}P_{R}] \\ - \frac{1}{q_{1}^{2}q_{2}^{2}} \left(\frac{1}{t_{1} - M_{\eta}^{2}} + \frac{1}{u_{1} - M_{\eta}^{2}}\right) \left(\frac{1}{t_{2} - M_{\eta}^{2}} + \frac{1}{u_{2} - M_{\eta}^{2}}\right) (\operatorname{Tr}[\not{p}_{1}\gamma^{\alpha}\gamma^{\mu}\not{q}_{1}\not{p}_{2}\gamma^{\beta}\gamma^{\nu}\not{q}_{2}P_{R}] \\ + \operatorname{Tr}[\not{p}_{1}\not{q}_{2}\gamma^{\mu}\gamma^{\alpha}\not{p}_{2}\dot{q}_{1}\gamma^{\nu}\gamma^{\beta}P_{R}]) + \frac{1}{q_{2}^{4}} \left(\frac{1}{t_{2} - M_{\eta}^{2}} + \frac{1}{u_{2} - M_{\eta}^{2}}\right)^{2} \operatorname{Tr}[\not{p}_{1}\not{q}_{2}\gamma^{\mu}\gamma^{\alpha}\not{p}_{2}\gamma^{\beta}\gamma^{\nu}\not{q}_{2}P_{R}]\right).$$
(15)

We evaluate this in terms of scalar products using the standard Dirac algebra, leading to a result too lengthy to record here.

The thermally averaged rate is given by

$$vd\sigma = \frac{1}{2s} \int \frac{1}{4} \sum_{\text{spin,pol.}} |\mathcal{M}|^2 dLips^3, \qquad (16)$$

where the $\frac{1}{4}$ arises from averaging over the spins of the initial χ pair, and $v = \sqrt{1 - \frac{4M_{\chi}^2}{s}}$ is the mean dark matter relative velocity, as well as the dark matter single-particle velocity in the center of mass (c.m.) frame.³

The three-body Lorentz invariant phase space is

$$dLips^{3} = (2\pi)^{4} \frac{d^{3}\vec{p}_{1}}{2E_{1}} \frac{d^{3}\vec{p}_{2}}{2E_{2}} \frac{d^{3}\vec{Q}}{2E_{W}} \frac{\delta^{4}(P - p_{1} - p_{2} - Q)}{(2\pi)^{9}}$$
(17)

and $P = k_1 + k_2$. This factorizes into the product of two two-body phase space integrals, convolved with an integral over the fermion propagator momentum,

$$dLips^{3} = \int_{M_{W}^{2}}^{s} \frac{dq_{1}^{2}}{2\pi} \left(\frac{d^{3}\vec{q}_{1}}{2E_{q_{1}}} \frac{d^{3}\vec{p}_{2}}{2E_{2}} \frac{\delta^{4}(P-q_{1}-p_{2})}{(2\pi)^{2}} \right) \\ \times \left(\frac{d^{3}\vec{p}_{1}}{2E_{1}} \frac{d^{3}\vec{Q}}{2E_{W}} \frac{\delta^{4}(q_{1}-Q-p_{1})}{(2\pi)^{2}} \right) \\ = \int_{M_{W}^{2}}^{s} \frac{dq_{1}^{2}}{2\pi} dLips^{2}(P^{2}, q_{1}^{2}, p_{2}^{2}) dLips^{2}(q_{1}^{2}, Q^{2}, p_{1}^{2}).$$
(18)

Evaluating the two-body phase space factors in their respective center of momentum frames, and using $p_1^2 = p_2^2 = 0$, we have

$$dLips^{2}(x^{2}, y^{2}, 0) = \frac{x^{2} - y^{2}}{8\pi x^{2}} \frac{d\bar{\Omega}}{4\pi}.$$
 (19)

This allows us to write the three-body phase space as

$$dLips^{3} = \frac{1}{2^{6}(2\pi)^{4}} \times \int_{M_{W}^{2}}^{s} dq_{1}^{2} \frac{(s-q_{1}^{2})(q_{1}^{2}-Q^{2})}{sq_{1}^{2}} d\phi d\cos\theta_{P} d\cos\theta_{q},$$
(20)

where ϕ is the angle of intersection of the plane defined by $\chi \chi \rightarrow ee^*$ with that defined by $e\nu W$, and θ_P and θ_q are defined in P (c.m.) and q rest frames, respectively.

³Informative discussions of the meaning of v are given in [36], and, including thermal averaging, in [37].

We evaluate the scalar products that arise from Eq. (15) in terms of the invariants q_1^2 , $Q^2 = M_W^2$, s, t_1 , and u_1 , and the angles θ_P , θ_q , and ϕ . We then use Eq. (16) to evaluate the cross section. As we have neglected diagrams suppressed by M_{η}^{-2} relative to those in Fig. 1, we present our results to leading order in M_{η}^{-4} (i.e., we take $M_{\eta}^2 \gg t_1$, t_2 , u_1 , u_2). To leading order in powers of M_{χ} and M_W in the numerator, we find

$$\upsilon \sigma = \frac{g^2 f^4}{512 M_W^2 M_\eta^4 \pi^3} \Big\{ M_\chi^4 \Big(\frac{1}{3} \ln \Big[\frac{4M_\chi^2}{M_W^2} \Big] - \frac{7}{18} \Big) \\ + M_\chi^2 M_W^2 \Big(\ln \Big[\frac{4M_\chi^2}{M_W^2} \Big] \Big\{ 1 + \ln \Big[\frac{2M_W M_\chi}{M_W^2 + 4M_\chi^2} \Big] \Big\} - 1 \\ + \mathrm{Li}_2 \Big[\frac{4M_\chi^2}{M_W^2 + 4M_\chi^2} \Big] - \mathrm{Li}_2 \Big[\frac{M_W^2}{M_W^2 + 4M_\chi^2} \Big] \Big\} + \mathcal{O}(M_W^4) \Big\}.$$
(21)

The Spence function (or "dilogarithm") is defined as

$$\operatorname{Li}_{2}(z) \equiv -\int_{0}^{z} \frac{d\zeta}{\zeta} \ln|1-\zeta| = \sum_{k=1}^{\infty} \frac{z^{k}}{k^{2}}.$$

The full expression (retaining subleading terms in M_{χ} in the numerator) is specified in Appendix D. Clearly, the leading terms are neither helicity nor velocity suppressed.

The effectiveness of the *W*-strahlung processes in lifting suppression of the annihilation rate can be seen Fig. 2, where we plot the ratio of the *W*-strahlung cross section to that of the lowest order process, $R_W = v\sigma(\chi\chi \rightarrow e^+\nu W^-)/v\sigma(\chi\chi \rightarrow e^+e^-)$. We see that the *W* strahlung rate rises with DM mass, to quickly dominate over the lowest order annihilation process. The *W*-bremsstrahlung rate rises approximately as M_{χ}^4 . As M_{χ}^2 increases, eventu-



FIG. 2 (color online). The ratio $R = v\sigma(\chi\chi \rightarrow e^+\nu W^-)/v\sigma(\chi\chi \rightarrow e^+e^-)$ for the example model [32], with $M_\eta^2 \gg M_\chi^2$. We have used $v = 10^{-3}c$, appropriate for the galactic halo.

ally phase space allows multi-W/Z radiative production, with such a large rate that resummation techniques become necessary. The onset of multi-W/Z dominance has been discussed in [6–8].

To obtain the energy spectrum of the W, we compute the differential cross section in terms of E_W by making the transformation

$$d\cos(\theta_q) \rightarrow \frac{-4\sqrt{sq^2}}{(s-q^2)(q^2-M_W^2)} dE_W.$$
 (22)

We find [38], again to leading order M_{η}^{-4} ,

$$\frac{\upsilon d\sigma}{dE_W} = \frac{g^2 f^4}{512E_W M_W^2 M_\eta^4 \pi^3} \left\{ 2E_W \sqrt{E_W^2 - M_W^2} (M_W^2 - 6E_W^2) + 8E_W M_\chi - 2M_\chi^2 + (4E_W^4 - 8E_W^3 M_\chi) + (2E_W^2 - M_W^2) (2M_\chi^2 + M_W^2)) \right\} \times \ln \left[\frac{E_W + \sqrt{E_W^2 - M_W^2}}{E_W - \sqrt{E_W^2 - M_W^2}} \right] \right\}.$$
(23)

The W spectrum per $\chi \chi \rightarrow e \nu W$ event is given in Fig. 3. We use the scaling variable $x_W \equiv E_W/M_{\chi}$, and plot

$$dN/dx_W \equiv \left(\frac{1}{\sigma_{e^+\nu W^-}}\right) \frac{d\sigma_{e^+\nu W^-}}{dx_W}$$

The kinematic range of x_w is $\left[\frac{M_w}{M_\chi}, 1 + \frac{M_w^2}{4M_\chi^2}\right]$, with the lower limit corresponding to a *W* produced at rest, and the upper limit corresponding to parallel lepton momenta balancing the opposite *W* momentum. As evident in Fig. 3, the *W*



FIG. 3 (color online). The W spectrum per $\chi \chi \rightarrow e\nu W$ event for the example model, with $M_{\chi} = 300$ GeV and $M_{\eta}^2 \gg M_{\chi}^2$.

boson spectrum has a broad energy distribution, including a significant component at high energy $E_W \sim M_{\chi}$.

The energy spectrum of the primary leptons is calculated in similar fashion. We present the analytic result in Appendix D (along with more detailed expressions for $v\sigma$ and $vd\sigma/dE_W$). Here the range of the scaling variable $x_\ell \equiv E_\ell/M_\chi$ is $[0, 1 - \frac{M_W^2}{4M_\chi^2}]$. Both limits arise when one lepton has zero energy and the other is produced back-toback with the W. The lepton spectrum is shown in Fig. 4. Note that this lepton spectrum is valid for either e^+ or ν from the annihilation $\chi\chi \rightarrow e^+\nu W^-$, and for either e^- or $\bar{\nu}$ from the annihilation $\chi\chi \rightarrow e^-\bar{\nu}W^+$. The primary lepton spectrum in Fig. 4 features a sharp cutoff near $E_\ell = M_\chi$, and a dip in the spectrum that is due to an absorptive interference effect.

To obtain the full lepton spectrum, the contributions from the subsequent decays of the gauge bosons to leptons must be included. (The contribution from the lowest order $2 \rightarrow 2$ process $\chi \chi \rightarrow e^+ e^-$ or $\nu \bar{\nu}$ is negligible. We also neglect final-state leptons resulting from μ decay and from the τ decay chain. These leptons are softer than those we consider.) For leptons from W decay, the range of the scaling variable x_{ℓ} is $\left[\frac{M_W^2}{4M_{\chi}^2}, 1\right]$. These limits arise when all four final-state lepton momenta are collinear. Particle spectra from the W decay may be calculated in a simple but approximate way, as we describe in Appendix E leading to Eq. (E8). The resulting secondary lepton spectrum is shown in Fig. 5. Not surprisingly, the spectrum of secondary leptons is softer than the spectrum of primary leptons.

When combining the primary lepton and secondary lepton spectra, the relative weights are model dependent. For example, the primary ℓ spectrum is weighted by BR $(\chi\chi \rightarrow W\nu\ell) + 2BR(\chi\chi \rightarrow Z\ell\ell)$, while the secondary ℓ spectrum is weighted by BR $(\chi\chi \rightarrow W + X) \times$ BR $(W \rightarrow \nu\ell) + BR(\chi\chi \rightarrow Z + X) \times BR(Z \rightarrow \ell\ell)$.



FIG. 4 (color online). The primary lepton spectrum per $\chi \chi \rightarrow e \nu W$ for the example model, with $M_{\chi} = 300 \text{ GeV}$ and $M_{\eta}^2 \gg M_{\chi}^2$.



FIG. 5 (color online). The secondary lepton spectrum (i.e., from $W \to \nu_{\ell} \ell$) per W for the example model, with $M_{\chi} = 300 \text{ GeV}$ and $M_{\eta}^2 \gg M_{\chi}^2$. (The branching ratio for $W \to \nu l$, 11% per flavor, is not included here.)

We note that the final charged-lepton spectra will be modified by cosmic propagation effects. The injected e^{\pm} will suffer rapid energy losses from synchrotron and inverse Compton processes on the Universe's background magnetic and radiation fields (see, e.g., Ref. [39] for a recent analysis). On the other hand, the injected neutrinos do not interact with the environment, and so their spectra remain unmodified.

C. Unsuppressed Z emission

Consider the process producing the $\bar{\nu}\nu Z$ final state. The cross sections for the Z-strahlung processes are related to those for W strahlung in a simple way: The amplitudes producing $\bar{\nu}\nu Z$ arise from the same four graphs of Fig. 1, where e, W, and η^+ are replaced everywhere by ν and Z and η_0 , respectively. The calculation of the amplitudes, and their interferences, thus proceeds in an identical fashion. After making the replacement $M_W \rightarrow M_Z$, the cross section for the annihilation process $\chi \chi \rightarrow \nu \bar{\nu} Z$ differs from that for $\chi \chi \rightarrow e^+ \nu W^-$ by only an overall normalization factor,

$$v\sigma_{\nu\bar{\nu}Z} = \frac{1}{(2\cos^2\theta_W)} \times v\sigma_{e^+\nu W^-}|_{M_W \to M_Z}$$
$$\simeq 0.65 \times v\sigma_{e^+\nu W^-}|_{M_W \to M_Z}.$$
(24)

Consider now the e^+e^-Z final state. Again, the amplitudes arise from the same four basic graphs of Fig. 1. Since only the left-handed leptons couple to the dark matter via the SU(2) doublet η , only the left-handed component of e^- participates in the interaction with the Z. Therefore, the couplings of the charged leptons to Z and W take the same form, up to a normalization constant. We thus find

$$\nu \sigma_{e^+e^-Z} = \frac{2(\sin^2 \theta_W - \frac{1}{2})^2}{\cos^2 \theta_W} \times \nu \sigma_{e^+\nu W^-}|_{M_W \to M_Z}$$
$$\simeq 0.19 \times \nu \sigma_{e^+\nu W^-}|_{M_W \to M_Z}.$$
(25)

V. CONCLUSIONS

In an attempt to explain recent anomalies in cosmic ray data in a dark matter framework, various nonstandard properties have been invoked such as dominant annihilation to leptons in so-called leptophilic models. When the dark matter is Majorana in nature, such annihilations invariably are confronted by suppressions of such processes via either *p*-wave velocity suppression or helicity suppression. With the aid of Fierz transformation technology, which we have presented in some detail, we have elucidated the general circumstances where suppressions may be encountered.

It has been known for some time that photon bremsstrahlung may have a dramatic effect on such suppressions. We have shown that once one considers the inclusion of three-body final states due to electroweak bremsstrahlung, one may also lift these suppressions and obtain rates which may be several orders of magnitude beyond those without such radiative corrections. In fact, barring an unexpected mass degeneracy, the EW bremsstrahlung lifts the suppression at 1 order lower in a certain small ratio of squared masses than does EM bremsstrahlung, as explained in the text.

Such radiative processes may be lethal for models attempting to produce positrons without overproducing antiprotons due to the subsequent hadronization of the radiated gauge bosons. Given that electroweak bremsstrahlung is the dominant annihilation channel for the DM models under consideration, and both W and Z decay to hadrons with a branching ratio of approximately 70%, a large hadronic component is unavoidable. Importantly in the context of recent cosmic ray data, there will be sizable antiproton production. We also note that dark matter searches triggering on antideuterons will find a sample in the W- and Z-bremsstrahlung processes. The Aleph experiment has measured an antideuteron production rate of $5.9 \pm 1.9 \times 10^{-6}$ antideuterons per hadronic decay of the Z [40]. We expect the rate for antideuteron production in W decay to be similar.

Even for models which do not suffer a suppression of the lowest order process, we see that it is impossible to have purely leptonic annihilation products, including "leptophilic" models in which the dark matter has direct couplings only to leptons. In a broader context the results presented here show the importance that may be played by electroweak bremsstrahlung in future searches of indirect dark matter detection. For any DM model for which electroweak bremsstrahlung makes an important contribution to observable fluxes, there will be large, correlated fluxes of e^{\pm} , neutrinos, hadrons, and gamma rays. We

will explore the detection of these signals in a future article [41].

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APPENDIX A: FUNDAMENTALS OF FIERZING

In this paper we have made use of standard Fierz transformations, helicity-basis Fierz transformations, and generalizations of the two. In this Appendix, we derive these transformations. The procedure for standard Fierz transformation can be found in, e.g., [26], while more general Fierz transformations are laid down in [25]. The starting point is to define a basis { Γ^B } and a dual basis { Γ_B }, each spanning 4 × 4 matrices over the complex number field C, such that an orthogonality relation holds. The standard Fierz transformation uses the "Hermitian" bases

$$\{\Gamma^B\} = \{\mathbb{1}, i\gamma_5, \gamma^{\mu}, \gamma_5\gamma^{\mu}, \sigma^{\mu\nu}\}, \text{ and} \\ \{\Gamma_B\} = \{\mathbb{1}, (-i\gamma_5), \gamma_{\mu}, (-\gamma_5\gamma_{\mu}), \frac{1}{2}\sigma_{\mu\nu}\},$$
(A1)

respectively. Because of their Lorentz and parity transformation properties, these basis matrices and their duals are often labeled as *S* and \tilde{S} (scalars), *P* and \tilde{P} (pseudoscalars), *V* and \tilde{V} (vectors, four for *V*, four for \tilde{V}), *A* and \tilde{A} (axial vectors, four for *A*, four for \tilde{A}), and *T* and \tilde{T} (antisymmetric tensors, six for *T*, six for \tilde{T}). As usual, spacetime indices are lowered with the Minkowski metric, $\gamma_5 = \gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3$, $\sigma^{\mu\nu} \equiv \frac{i}{2}[\gamma^{\mu}, \gamma^{\nu}]$, (and $\gamma^5\sigma^{\mu\nu} = \frac{i}{2}\epsilon^{\mu\nu\alpha\beta}\sigma_{\alpha\beta}$). Note the change of sign between the basis and dual for the *P* and *A* matrices. The bases are Hermitian in that $\gamma^0\Gamma_B^{\dagger}\gamma^0 = \Gamma_B$, so that the associated Dirac bilinears satisfy $[\bar{\Psi}_1\Gamma^B\Psi_2]^{\dagger} = \bar{\Psi}_2\Gamma^B\Psi_1$ and $[\bar{\Psi}_1\Gamma_B\Psi_2]^{\dagger} = \bar{\Psi}_2\Gamma_B\Psi_1$. Importantly, we have $\Gamma_B = (\Gamma^B)^{-1}$ in the sense of the accompanying orthogonality relation

$$\operatorname{Tr}[\Gamma_C \Gamma^B] = 4\delta_C^B, \quad B, C = 1, \dots, 16.$$
(A2)

Note that the factor of $\frac{1}{2}$ in the definition of $\tilde{T} = \frac{1}{2} \sigma_{\mu\nu}$ (but not in $T = \sigma^{\mu\nu}$) provides the normalization required by Eq. (A2):

$$\operatorname{Tr}[\Gamma_B \Gamma^B]_{(\text{nosum})} = \sum_C \operatorname{Tr}[\Gamma_C \Gamma^B] = 4.$$
(A3)

The orthogonality relation allows us to expand any 4×4 complex matrix X in terms of the basis as

$$X = X_B \Gamma^B = X^B \Gamma_B, \text{ with}$$
$$X_B = \frac{1}{4} \operatorname{Tr}[X\Gamma_B], \text{ and } X^B = \frac{1}{4} \operatorname{Tr}[X\Gamma^B],$$
i.e.,
$$X = \frac{1}{4} \operatorname{Tr}[X\Gamma^B]\Gamma_B = \frac{1}{4} \operatorname{Tr}[X\Gamma_B]\Gamma^B.$$
(A4)

One readily finds that the particular matrix element $(X)_{ab}$ satisfies

$$(\mathbf{X})_{cd}\delta_{db}\delta_{ac} = \frac{1}{4} [(\mathbf{X})_{cd}(\Gamma_B)_{dc}](\Gamma^B)_{ab}.$$
 (A5)

Since each element $(X)_{cd}$ is arbitrary, Eq. (A5) is equivalent to a completeness relation

$$(\mathbb{1})[\mathbb{1}] = \frac{1}{4}(\Gamma_B][\Gamma^B] = \frac{1}{4}(\Gamma^B][\Gamma_B), \qquad (A6)$$

where we have adopted Takahashi's notation [27] where matrix indices are replaced by parentheses $(\cdot \cdot \cdot)$ and brackets $[\cdot \cdot \cdot]$, in an obvious way. Thus, any 4×4 matrices X and Y may be expressed as

$$(\mathbf{X})[\mathbf{Y}] = (\mathbf{X}\mathbb{1})[\mathbb{1}\mathbf{Y}] = \frac{1}{4}(\mathbf{X}\Gamma^{B}\mathbf{Y}][\Gamma_{B})$$
$$= \frac{1}{4^{2}}\operatorname{Tr}[\mathbf{X}\Gamma^{B}\mathbf{Y}\Gamma_{C}](\Gamma^{C}][\Gamma_{B}).$$
(A7)

This equation is presented as Eq. (6) in the main text. Alternatively, we may express any 4×4 matrices X and Y as

$$(\mathbf{X})[\mathbf{Y}] = (\mathbf{X}\mathbb{1})[\mathbf{Y}\mathbb{1}] = \frac{1}{4}(\mathbf{X}\Gamma^{B}][\mathbf{Y}\Gamma_{B})$$
$$= \frac{1}{4^{3}}\operatorname{Tr}[\mathbf{X}\Gamma^{B}\Gamma_{C}]\operatorname{Tr}[\mathbf{Y}\Gamma_{B}\Gamma^{D}](\Gamma^{C}][\Gamma_{D}). \quad (A8)$$

The rhs's of Eqs. (A7) and (A8) offer two useful options for Fierzing matrices. The first option sandwiches both lhs matrices into one of the two spinor bilinears, and ultimately into a single long trace. The second option sandwiches each lhs matrix into a separate spinor bilinear, and ultimately into separate trace factors. Equation (A7) seems to be more useful than (A8). One use we will make of Eq. (A7) will be to express chiral vertices in terms of Fierzed standard vertices. But first we reproduce the standard Fierz transformation rules by setting $X = \Gamma^D$ and $Y = \Gamma_E$ in Eq. (A7), to wit:

$$(\Gamma^{D})[\Gamma_{E}] = \frac{1}{4^{2}} \operatorname{Tr}[\Gamma^{D}\Gamma^{B}\Gamma_{E}\Gamma_{C}](\Gamma^{C}][\Gamma_{B}).$$
(A9)

(An additional minus sign arises if the matrices are sandwiched between anticommuting field operators, rather than between Dirac spinors.) Evaluation of the trace in Eq. (A9) for the various choices of (B, C) leads to the oft-quoted result [26] PHYSICAL REVIEW D 83, 013001 (2011)

$$\begin{pmatrix} (S)[S]\\ (V)[\tilde{V}]\\ (T)[\tilde{T}]\\ (A)[\tilde{A}]\\ (P)[\tilde{P}] \end{pmatrix} = \frac{1}{4} \begin{pmatrix} 1 & 1 & 1 & 1 & 1 \\ 4 & -2 & 0 & 2 & -4 \\ 6 & 0 & -2 & 0 & 6 \\ 4 & 2 & 0 & -2 & -4 \\ 1 & -1 & 1 & -1 & 1 \end{pmatrix} \begin{pmatrix} (S)[S)\\ (V][\tilde{V})\\ (T][\tilde{T})\\ (A][\tilde{A}]\\ (P][\tilde{P}) \end{pmatrix}.$$
(A10)

More relevant for us, as will be seen, is the ordering *P*, *S*, *A*, *V*, and *T*, which leads to a Fierz matrix obtained from the one above with the swapping of matrix indices $1 \rightarrow 2 \rightarrow 4 \rightarrow 3 \rightarrow 5 \rightarrow 1$. The result is

$$\begin{pmatrix} (P)[\tilde{P}]\\ (S)[\tilde{S}]\\ (A)[\tilde{A}]\\ (V)[\tilde{V}]\\ (T)[\tilde{T}] \end{pmatrix} = \frac{1}{4} \begin{pmatrix} 1 & 1 & -1 & -1 & 1\\ 1 & 1 & 1 & 1 & 1\\ -4 & 4 & -2 & 2 & 0\\ -4 & 4 & 2 & -2 & 0\\ 6 & 6 & 0 & 0 & -2 \end{pmatrix} \begin{pmatrix} (P)[\tilde{P})\\ (S][\tilde{S})\\ (A][\tilde{A})\\ (V][\tilde{V})\\ (T][\tilde{T}) \end{pmatrix}.$$
(A11)

(The zeros make it clear that Fierzing induces no coupling between tensor interactions and vector and axial-vector interactions.) As an example of how to read this matrix,

$$(A)[\tilde{A}] = -(P][\tilde{P}] + (S][\tilde{S}] - \frac{1}{2}(A][\tilde{A}] + \frac{1}{2}(V][\tilde{V}], \quad (A12)$$

or, multiplying by spinors and giving the explicit forms of the gamma matrices,

$$(\bar{u}\gamma_{5}\gamma^{\mu}u)(\bar{v}(-\gamma_{5}\gamma_{\mu})v)$$

$$= -(\bar{u}i\gamma_{5}v)(\bar{v}(-i\gamma_{5})u)$$

$$+ (\bar{u}v)(\bar{v}u) - \frac{1}{2}(\bar{u}\gamma_{5}\gamma^{\mu}v)(\bar{v}(-\gamma_{5}\gamma_{\mu})u)$$

$$+ \frac{1}{2}(\bar{u}\gamma^{\mu}v)(\bar{v}\gamma_{\mu}u). \qquad (A13)$$

The Fierz matrix M for the standard basis is nonsingular, and hence has five nonzero eigenvalues λ_j . Since two swaps of Dirac indices return the indices to their original order, the matrix is idempotent, with $M^2 = 1$, or equivalently, $M^{-1} = M$. Accordingly, the five eigenvalues satisfy $\lambda_j^2 = 1$, so individual eigenvalues must be $\lambda_j = \pm 1$. Also, the corresponding eigenvectors are invariant under the interchange of two Dirac indices. In Table II we list the eigenvalues and "Fierz-invariant" eigenvectors.

Helicity projection operators are often present in theories where the DM couple to the SU(2) lepton doublet, so it

TABLE II. Fierz-invariant combinations in the standard basis.

Fierz-invariant combination	Eigenvalue
$\overline{3(S \otimes \tilde{S} + P \otimes \tilde{P}) + T \otimes \tilde{T}}$ $2(S \otimes \tilde{S} - P \otimes \tilde{P}) + (V \otimes \tilde{V} + A \otimes \tilde{A})$	+1 +1
$ \begin{split} V &\otimes \tilde{V} - A \otimes \tilde{A} \\ S &\otimes \tilde{S} + P \otimes \tilde{P} - T \otimes \tilde{T} \\ 2(S \otimes \tilde{S} - P \otimes \tilde{P}) - (V \otimes \tilde{V} + A \otimes \tilde{A}) \end{split} $	

is worth considering Fierz transformations in the more convenient chiral basis.

One derivation of chiral Fierz transformations utilizes the following chiral bases (hatted) [25]:

$$\{\hat{\Gamma}^B\} = \{P_R, P_L, P_R \gamma^{\mu}, P_L \gamma^{\mu}, \frac{1}{2} \sigma^{\mu\nu}\}, \text{ and}$$
$$\{\hat{\Gamma}_B\} = \{P_R, P_L, P_L \gamma_{\mu}, P_R \gamma_{\mu}, \frac{1}{2} \sigma_{\mu\nu}\},$$
(A14)

where $P_R \equiv \frac{1}{2}(1 + \gamma_5)$ and $P_L \equiv \frac{1}{2}(1 - \gamma_5)$ are the usual helicity projectors. The orthogonality property between the chiral basis and its dual is

$$\operatorname{Tr}[\hat{\Gamma}_{C}\hat{\Gamma}^{B}] = 2\delta_{C}^{B}, \quad B, C = 1, \dots, 16,$$
(A15)

which implies the normalization

$$\operatorname{Tr}\left[\hat{\Gamma}_{B}\hat{\Gamma}^{B}\right]_{(\text{nosum})} = \sum_{C} \operatorname{Tr}\left[\hat{\Gamma}_{C}\hat{\Gamma}^{B}\right] = 2.$$
(A16)

Notice that because $\{\gamma_5, \gamma^{\mu}\} = 0$, the dual of $P_R \gamma^{\mu}$ is $P_L \gamma_{\mu}$, and the dual of $P_L \gamma^{\mu}$ is $P_R \gamma_{\mu}$. Notice also that the normalization for the chiral bases necessitates factors of $\frac{1}{2}$ in both $\hat{T} = \frac{1}{2} \sigma^{\mu\nu}$ and $\tilde{T} = \frac{1}{2} \sigma_{\mu\nu}$, in contrast to the tensor elements of the standard bases, given in Eq. (A1).

In the chiral basis, one is led to a general expansion

$$X = \frac{1}{2} \operatorname{Tr}[X\hat{\Gamma}^{B}]\hat{\Gamma}_{B} = \frac{1}{2} \operatorname{Tr}[X\hat{\Gamma}_{B}]\hat{\Gamma}^{B}, \qquad (A17)$$

and to a completeness relation

$$(1)[1] = \frac{1}{2}(\hat{\Gamma}_B][\hat{\Gamma}^B] = \frac{1}{2}(\hat{\Gamma}^B][\hat{\Gamma}_B].$$
 (A18)

Thus, any 4×4 matrices X and Y may be expressed as

$$\begin{aligned} (\mathbf{X})[\mathbf{Y}] &= (\mathbf{X}\mathbb{1})[\mathbb{1}\mathbf{Y}] = \frac{1}{2}(\mathbf{X}\hat{\Gamma}_{B}\mathbf{Y}][\hat{\Gamma}^{B}) \\ &= \frac{1}{4}\mathrm{Tr}[\mathbf{X}\hat{\Gamma}^{C}\mathbf{Y}\hat{\Gamma}_{B}](\hat{\Gamma}^{B}][\hat{\Gamma}_{C}). \end{aligned} \tag{A19}$$

TABLE III. Fierz-invariant combinations in the chiral basis.

Fierz-invariant combination	Eigenvalue
$\overline{3(P_R \otimes P_R + P_L \otimes P_L) + \hat{T} \otimes \tilde{\hat{T}}}$	+1
$2P_R \otimes P_L + P_R \gamma^\mu \otimes P_L \gamma_\mu$	+1
$2P_L \otimes P_R + P_L \gamma^\mu \otimes P_R \gamma^\mu_\mu$	+1
$P_R \otimes P_R + P_L \otimes P_L - \hat{T} \otimes \tilde{\hat{T}}$	-1
$2P_R \otimes P_L - P_R \gamma^\mu \otimes P_L \gamma_\mu$	-1
$2P_L \otimes P_R - P_L \gamma^\mu \otimes P_R \gamma_\mu$	-1
$P_R \gamma^\mu \otimes P_R \gamma_\mu$	-1
$P_L \gamma^\mu \otimes P_L \gamma^{\dot{\mu}}_{\mu}$	-1

Substituting $X = \hat{\Gamma}^D$ and $Y = \hat{\Gamma}_E$ into Eq. (A19), one gets

$$(\hat{\Gamma}^D)[\hat{\Gamma}_E] = \frac{1}{4} \operatorname{Tr}[\hat{\Gamma}^D \hat{\Gamma}^C \hat{\Gamma}_E \hat{\Gamma}_B](\hat{\Gamma}^B][\hat{\Gamma}_C).$$
(A20)

Evaluating the trace in Eq. (2) leads to the chiral-basis analog of (A10) or (A11), presented in Eq. (3) of the main text.

As a check, we note that the matrix M in Eq. (3) is idempotent, $M^2 = 1$, as it must be. The eigenvalues are therefore ± 1 . Eigenvalues and Fierz-invariant eigenvectors for the chiral basis are given in Table III. The final two eigenvectors in the table simply express again the invariance of $V \pm A$ interactions under Fierz transposition of Dirac indices. This invariance is also evident in the diagonal nature of the bottom two rows of the matrix Eq. (3).

One may instead want the Fierz transformation that takes chiral bilinears to standard bilinears. Since models are typically formulated in terms of chiral fermions, a projection onto standard *s*-channel bilinears would be well suited for a partial wave analysis. Because different partial waves do not interfere with one another, the calculation simplifies in terms of *s*-channel partial waves.

Setting $X = \hat{\Gamma}^D$ and $Y = \hat{\Gamma}_E$ in Eq. (A7), we readily get

$$(\hat{\Gamma}^D)[\hat{\Gamma}_E] = \frac{1}{4^2} \operatorname{Tr}[\hat{\Gamma}^D \Gamma^B \hat{\Gamma}_E \Gamma_C](\Gamma^C][\Gamma_B).$$
(A21)

We (should) get the same result by resolving the rhs vector in Eq. (3) into standard basis matrices. The result is

$ \begin{pmatrix} (P_{R})[P_{R}] \\ (P_{L})[P_{L}] \\ (P_{R}\gamma^{\mu})[P_{L}\gamma_{\mu}] \\ (P_{L}\gamma^{\mu})[P_{R}\gamma_{\mu}] \\ (\hat{T})[\hat{T}] \\ (\gamma_{5}\hat{T})[\hat{T}] \\ (P_{R})[P_{L}] \\ (P_{L})[P_{R}] \\ (P_{L}\gamma^{\mu})[P_{R}\gamma_{\mu}] \\ (P_{L}\gamma^{\mu})[P_{L}\gamma_{\mu}] \end{pmatrix} =$	$=\frac{1}{8}$	(1) (1) (4) (4) (6) (0) (0)	$ \begin{array}{c} 1 \\ -4 \\ -4 \\ 6 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 1 \\ -1 \\ 4 \\ -4 \\ 0 \\ 6 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 1 \\ -1 \\ -4 \\ 4 \\ 0 \\ 6 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 1 \\ 1 \\ 0 \\ -2 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 1 \\ -1 \\ 0 \\ 0 \\ 2 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \end{array} $	$ \begin{array}{c} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 1 \\ -2 \\ -2 \\ -2 \end{array} $	$ \begin{array}{c} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ -1 \\ -2 \\ -2 \end{array} $	$ \begin{array}{c} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 1 \\ -1 \\ -2 \\ 2 \end{array} $	$ \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ -1 \\ 1 \\ -2 \\ 2 \end{pmatrix} $		$(1][1)(\gamma_{5}][\gamma_{5})(\gamma_{5}][1)(1][\gamma_{5})(T][\tilde{T})(\gamma_{5}T][\tilde{T})(\gamma^{\mu}][\gamma_{\mu})\gamma_{5}\gamma^{\mu}][\gamma_{5}\gamma_{\mu})(\gamma^{\mu}][\gamma_{5}\gamma_{\mu})(\gamma^{\mu}][\gamma_{5}\gamma_{\mu})$			(A22)
---	----------------	---	---	--	--	--	--	--	--	--	---	--	---	--	--	------	---

All relations are invariant under the simultaneous interchanges $P_R \leftrightarrow P_L$ and $\gamma_5 \rightarrow -\gamma_5$. The matrix in (A22), relating two different bases, is not idempotent. In fact, it is singular.

APPENDIX B: CANCELLATION OF VECTOR AND TENSOR AMPLITUDES FOR MAJORANA FERMIONS

Majorana particles are invariants under charge conjugation C. Accordingly, the Majorana field creates and annihilates the same particle. This implies that for each t-channel diagram, there is an accompanying u-channel diagram, obtained by interchanging the momentum and spin of the two Majorana fermions. The relative sign between the t- and u-channel amplitudes is -1 in accord with Fermi statistics. For example, consider the Fierzed (i.e., s channel) bilinear for χ annihilation: $\bar{v}(k_1, s_1) \times \Gamma_B u(k_2, s_2)$. The associated Fierzed bilinear from the $(k_1 \leftrightarrow k_2)$ -exchange graph, with its relative minus sign, is $-\bar{v}(k_2, s_2)\Gamma_B u(k_1, s_1)$. Constraints relating the fourcomponent Dirac spinors to their underlying twocomponent Majorana spinors must be imposed. These constraints, any one of which implies the other three, are

$$u(p, s) = C\bar{v}^{T}(p, s), \qquad \bar{u}(p, s) = -v^{T}(p, s)C^{-1}, v(p, s) = C\bar{u}^{T}(p, s), \qquad \bar{v}(p, s) = -u^{T}(p, s)C^{-1}.$$
(B1)

Here, C is the charge-conjugation matrix. These Majorana conditions on the spinors allow us to rewrite the exchange bilinear as (suppressing spin labels for brevity of notation)

$$-\bar{v}(k_{2})\Gamma_{B}u(k_{1})) = u^{T}(k_{2})C^{-1}\Gamma_{B}C\bar{v}^{T}(k_{1})$$
$$= [\bar{v}(k_{1})(C^{-1}\Gamma_{B}C)^{T}u(k_{2})]^{T}$$
$$= \bar{v}(k_{1})(\eta_{B}\Gamma_{B})u(k_{2}).$$
(B2)

For the final equality, we have used (i) the fact that the transpose symbol can be dropped from a number, and (ii) the identity $(C^{-1}\Gamma_B C)^T = (\eta_B(\Gamma_B)^T)^T = \eta_B\Gamma_B$, where $\eta_B = +1$ for $\Gamma =$ scalar, pseudoscalar, axial vector, and $\eta_B = -1$ for $\Gamma =$ vector or tensor.

In the four-Fermi or heavy propagator limit, where the differing momenta in the *t*- and *u*-channel propagators can be ignored, one obtains an elegant simplification. Subtracting the *u*-channel amplitude from the *t*-channel amplitude, one arrives at the weighting factor $(1 + \eta_B)$, which is 2 for *S*, *P* and *A* couplings, and zero for *V* and *T* couplings. Thus, we must drop *V* and *T* couplings appearing in the Fierzed bilinears of the χ current. What this means for the model under discussion is that after Fierzing, only the axial-vector coupling of the χ current remains, and the factor of $1 + \eta_A = 2$ is multiplied by the (7–8) element $= \frac{1}{2}$ in the Fierz matrix of Eq. (3) to give a net weight of 1.

APPENDIX C: NONRELATIVISTIC AND EXTREME-RELATIVISTIC LIMITS OF FERMION BILINEARS

We work in the chiral representation of the Dirac algebra, and we follow the notation of [28]. Accordingly,

$$\gamma_0 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \qquad \vec{\gamma} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix}, \qquad \gamma_5 = \begin{pmatrix} -1 & 0 \\ 0 & 1 \end{pmatrix}.$$
(C1)

The rest-frame four-spinor is

$$u(\vec{p}=0) = \sqrt{M} \begin{pmatrix} \xi \\ \xi \end{pmatrix}, \tag{C2}$$

where ξ is a two-dimensional spinor. The spinor with arbitrary momentum is obtained by boosting. One gets

$$u(p) = \begin{pmatrix} \sqrt{p \cdot \sigma} \xi \\ \sqrt{p \cdot \overline{\sigma}} \xi \end{pmatrix}, \tag{C3}$$

where $\sigma \equiv (1, \vec{\sigma})$ and $\bar{\sigma} \equiv (1, -\vec{\sigma})$.

In a standard fashion, we choose the up and down spin eigenstates of σ_3 as the basis for the two-spinors. These basis two-spinors are

$$\xi_{+} \equiv \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \qquad \xi_{-} \equiv \begin{pmatrix} 0 \\ 1 \end{pmatrix}. \tag{C4}$$

In terms of the chosen basis, we have for the nonrelativistic (NR) u spinors,

$$u_{\pm} \xrightarrow{\mathrm{NR}} \sqrt{M} \begin{pmatrix} \dot{\xi}_{\pm} \\ \dot{\xi}_{\pm} \end{pmatrix}. \tag{C5}$$

We get the extreme-relativistic (ER) limit of the u spinors from Eq. (C3). After a bit of algebra, one finds

$$u_{+} \xrightarrow{\mathrm{ER}} \sqrt{2E} \begin{pmatrix} 0\\0\\\xi_{+} \end{pmatrix}, \qquad u_{-} \xrightarrow{\mathrm{ER}} \sqrt{2E} \begin{pmatrix} \xi_{-}\\0\\0 \end{pmatrix}.$$
 (C6)

The arbitrary v spinor is given by

$$v(p) = \left(\begin{array}{c} \sqrt{p \cdot \sigma} \eta \\ -\sqrt{p \cdot \overline{\sigma}} \eta \end{array} \right)$$
(C7)

In the Dirac bilinear the two-spinor η is independent of the two-spinor xi, and so it is given an independent name, η . However, the basis η^{\pm} remains ξ^{\pm} as defined above. It is the minus sign in the lower components of v relative to the upper components that distinguishes v in Eq. (C7) from u in Eq. (C3) in a fundamental way. After a small amount of algebra, one finds the limits

$$v_{\pm} \xrightarrow{\mathrm{NR}} v_{\pm}(\vec{p}=0) = \sqrt{M} \begin{pmatrix} \eta_{\pm} \\ -\eta_{\pm} \end{pmatrix}, \qquad (C8)$$

and

$$v_{+} \xrightarrow{\mathrm{ER}} \sqrt{2E} \begin{pmatrix} 0\\ 0\\ -\eta_{+} \end{pmatrix}, \qquad v_{-} \xrightarrow{\mathrm{ER}} \sqrt{2E} \begin{pmatrix} \eta_{-}\\ 0\\ 0 \end{pmatrix}.$$
 (C9)

Finally, we apply the above to determine the values of Dirac bilinears in the NR and ER limits. The $\bar{u} \equiv u^{\dagger} \gamma_0$ and $\bar{v} \equiv v^{\dagger} \gamma_0$ conjugate spinors are easily found from the *u* and *v* spinors. We let Γ denote any of the Hermitian basis Dirac matrices {1, $i\gamma_5$, γ^{μ} , $\gamma_5\gamma^{\mu}$, $\sigma^{\mu\nu}$ }. Then, the NR limit of $\bar{u}(p_1)\Gamma v(p_2)$ is just

$$\bar{u}(p_1)\Gamma v(p_2) \xrightarrow{\text{NR}} M \left[(\xi_1, \xi_1)\Gamma \begin{pmatrix} \eta_2 \\ -\eta_2 \end{pmatrix} \right].$$
(C10)

Nonrelativistic results for the various choices of basis Γ 's and spin combinations are listed in Table I of the text.

To give a succinct formula for the ER limit of $\bar{u}(p_1)\Gamma v(p_2)$, we take $\hat{p}_1 = -\hat{p}_2 = \hat{3}$, i.e. we work in a frame where \hat{p}_1 and \hat{p}_2 are collinear, and we quantize the spin along this collinear axis. The result is

$$\bar{u}(p_1)\Gamma v(p_2) \xrightarrow{\text{ER}} \sqrt{4E_1E_2} \bigg[\xi_1(\Lambda_+, \Lambda_-)\Gamma \bigg(\begin{array}{c} \Lambda_+ \\ -\Lambda_- \end{array} \bigg) \eta_2 \bigg],$$
(C11)

where the matrices Λ_{\pm} are just up and down spin projectors along the quantization axis $\hat{3}$:

$$\lambda_{+} = \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix}, \qquad \Lambda_{-} = \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix}. \tag{C12}$$

Extreme-relativistic results for the various choices of basis Γ 's and spin combinations are listed in Table I of the text.

APPENDIX D: FULL CROSS-SECTION RESULTS

We present here the full results of the cross-section calculations for the process $\chi\chi \rightarrow e^{\pm \nu}W^{\pm}$, including terms of all orders in M_{χ} . In Sec. IV we presented only the leading order terms, which dominate in the large M_{χ} limit. For M_{χ} not too much heavier than M_W , it is important to retain subleading terms.

The total cross section for $\chi \chi \rightarrow e^{\pm (-)} W^{\pm}$ is given by

$$\begin{split} \upsilon\sigma_{e^{+}\nu W^{-}} &= \frac{g^{2}f^{4}}{2^{13}M_{W}^{4}M_{\eta}^{4}\pi^{3}} \Big\{ \Big(\frac{7}{32} \frac{M_{W}^{8}}{M_{\chi}^{4}} - \frac{7}{9} \frac{M_{W}^{6}}{M_{\chi}^{2}} + 4M_{W}^{4} - 16M_{W}^{2}M_{\chi}^{2} - \frac{56}{9}M_{\chi}^{4} \Big) + \ln \Big[\frac{4M_{\chi}^{2}}{M_{W}^{2}} \Big] \Big(\frac{1}{16} \frac{M_{W}^{8}}{M_{\chi}^{4}} + \frac{4}{3} \frac{M_{W}^{6}}{M_{\chi}^{2}} - 2M_{W}^{4} \\ &+ 16M_{W}^{2}M_{\chi}^{2} + \frac{16}{3}M_{\chi}^{4} + 8M_{W}^{2}(M_{W}^{2} + 2M_{\chi}^{2}) \ln \Big[\frac{2M_{W}M_{\chi}}{M_{W}^{2} + 4M_{\chi}^{2}} \Big] \Big) + 8M_{W}^{2}(M_{W}^{2} + 2M_{\chi}^{2}) \Big(\text{Li}_{2} \Big[\frac{4M_{\chi}^{2}}{M_{W}^{2} + 4M_{\chi}^{2}} \Big] \\ &- \text{Li}_{2} \Big[\frac{M_{W}^{2}}{M_{W}^{2} + 4M_{\chi}^{2}} \Big] \Big) + \mathcal{O}(\upsilon^{2}, M_{\eta}^{-2}, m_{\ell}^{2}) \Big\}. \end{split}$$
(D1)

The W energy spectrum is

$$\frac{\upsilon d\sigma_{e^+\nu W^-}}{dE_W} = \frac{g^2 f^4}{512 E_W M_W^2 M_\eta^4 \pi^3} \left\{ 2E_W \sqrt{E_W^2 - M_W^2} (M_W^2 - 6E_W^2 + 8E_W M_\chi - 2M_\chi^2) + (4E_W^4 - 8E_W^3 M_\chi^2 + (2E_W^2 - M_W^2)(2M_\chi^2 + M_W^2)) \ln \left[\frac{E_W + \sqrt{E_W^2 - M_W^2}}{E_W - \sqrt{E_W^2 - M_W^2}} \right] + \mathcal{O}(\upsilon^2, M_\eta^{-2}, m_\ell^2) \right\}, \quad (D2)$$

while the lepton spectrum (for either the charged lepton or the neutrino) is

$$\begin{aligned} \frac{\upsilon d\sigma_{e^+ \nu W^-}}{dE_{\ell}} &= \frac{g^2 f^4}{2^{18} M_{\eta}^4 (M_{\chi} - E_e) \pi^3} \Big\{ \frac{E_e (4(M_{\chi} - E_e)M_{\chi} - M_W^2)}{M_W^2 M_{\chi}^4 (M_W^2 + 4E_e M_{\chi})(M_{\chi} - E_e)^4} \times \Big(7 \times 2^8 E_e^7 M_W^4 + 2^8 E_e^6 M_{\chi}^3 \Big(M_W^2 - \frac{139}{3} M_{\chi}^2\Big) \\ &- 2^5 E_e^5 M_{\chi}^2 \Big(M_W^4 + \frac{146}{3} M_{\chi}^2 M_W^2 - 984 M_{\chi}^4\Big) - 2^4 E_e^4 M_{\chi} (M_W^6 - 13 M_W^4 M_{\chi}^2 - 232 M_{\chi}^4 M_W^2 + 2704 M_{\chi}^6) \\ &+ E_e^3 \Big(-M_W^8 + \frac{164}{3} M_W^6 M_{\chi}^2 - 560 M_W^4 M_{\chi}^4 - 4672 M_{\chi}^6 M_W^2 + \frac{191 \times 2^9}{3} M_{\chi}^8\Big) + 8 E_e^2 M_{\chi} \Big(\frac{1}{3} M_W^8 - 6 M_W^6 M_{\chi}^2 \\ &+ 116 M_W^4 M_{\chi}^4 + \frac{1376}{3} M_{\chi}^6 M_W^2 - 1600 M_{\chi}^8\Big) - 2 E_e M_{\chi}^2 (M_W^8 M_{\chi}^2 - 4 M_W^6 M_{\chi}^2 + 400 M_W^4 M_{\chi}^4 + 960 M_{\chi}^6 M_W^2 - 2^{10} M_{\chi}^8) \\ &+ 2^8 M_{\chi}^3 (M_W^2 + 2 M_{\chi}^2) \Big) + 2^8 \Big(2 M_{\chi}^2 + M_W^2 - \frac{4 E_e^2 (M_{\chi} - E_e)^2}{M_W^2} \Big) \ln \Big[\frac{M_W^2 M_{\chi}}{(M_{\chi} - E_e)(M_W^2 + 4E_e M_{\chi})}\Big] \Big\} \\ &+ \mathcal{O}(\nu^2, M_{\eta}^{-2}, m_{\ell}^2). \end{aligned}$$

APPENDIX E: APPROXIMATE SPECTRUM FOR BOOSTED W DECAY PRODUCTS

If any possible polarization of the produced *W* is neglected, then a simple calculation results for the spectra of the final-state particles from *W* decay. The lab frame spectra of the decay product (of type or "flavor" *F*) depends on a one-dimensional convolution of the isotropic spectrum in the *W* rest frame (RF energy E'), $\frac{dN_F}{dE_{F'}}$, with the *W* spectrum in the lab frame, $\frac{dN}{dE_W}$. We now develop this convolution.

Given the energy distribution $dN_W/d\gamma$ of produced W's (with $\gamma = E_W/M_W$), and the energy distribution dN_F/dE'_F of decay particle F in the W rest frame, normalized to the multiplicity of F per W decay (i.e., there is a branching ratio $W \rightarrow F$ multiplier implicit in dN_F/dE'_F) and assumed to be isotropic,⁴ one gets the spectrum dN_F/dE_F of particle F in the lab via

$$\frac{dN_F(E)}{dE} = \int_{-1}^{1} \frac{d\cos\theta'}{2} \int d\gamma \frac{dN_W}{d\gamma} \\ \times \int dE' \frac{dN_F}{dE'} \delta(E - [\gamma E' + \beta \gamma p' \cos\theta']), \quad (E1)$$

with $p' = \sqrt{E'^2 - m_F^2}$, $\beta \gamma = \sqrt{\gamma^2 - 1}$. The $\cos \theta'$ integral is easily done, and one gets

$$\frac{dN_F(E)}{dE} = \frac{1}{2} \int_1^\infty \frac{d\gamma}{\sqrt{\gamma^2 - 1}} \frac{dN_W}{d\gamma} \int_{E'_-}^{E'_+} \frac{dE'}{p'} \frac{dN_F}{dE'}, \quad (E2)$$

with $E'_{\pm} = \gamma E \pm \beta \gamma p$. Equivalently, we get

$$\frac{dN_F(E)}{dE} = \frac{1}{2} \int_{m_F}^{\infty} \frac{dE'}{p'} \frac{dN_F}{dE'}, \int_{\gamma_-}^{\gamma_+} \frac{d\gamma}{\sqrt{\gamma^2 - 1}} \frac{dN_W}{d\gamma}, \quad (E3)$$

with $\gamma_{\pm} = (EE' \pm pp')/m_F^2$ and $p = \sqrt{E^2 - m_F^2}$. This formulation neglects interferences between identical particles produced in both the primary and secondary channels, if any.

As given, Eq. (E3) applies to any particle type in the W's final state. For example, it could be used to calculate the antiproton or antineutron spectrum from W production and decay, if the fragmentation functions for $W \rightarrow \bar{p}$ or \bar{n} , i.e. $f(x_{\bar{B}} \equiv 2E_{\bar{B}}/M_W)$ were input.

Here we perform the convolution for the especially simple case of *W* decay to two massless particles, say ν_e and *e*. For massless leptons, we have

$$\frac{dN_{\nu}}{dE'} = \frac{dN_e}{dE'} = \mathrm{BR}(W \to \nu e)\delta\left(E' - \frac{1}{2}M_W\right), \quad (E4)$$

with $\gamma_+ = (E_W/M_W)_{\text{max}} = (s + M_W^2)/2\sqrt{s}M_W \approx (4M_\chi^2 + M_W^2)/4M_\chi M_W$, and $\gamma_- = (4E^2 + M_W^2)/4EM_W$. The spectrum in the lab is given by Eq. (E3) and becomes just

$$\frac{dN_{\nu}}{dE} = \frac{dN_e}{dE} = \frac{(\text{BR})}{M_W} \int_{\gamma_-}^{\gamma_+} \frac{d\gamma}{\sqrt{\gamma^2 - 1}} \frac{dN_W}{d\gamma}.$$
 (E5)

The *W* spectrum shown in Fig. 3 is approximately one-half of an ellipse, suggesting the fit

$$\left(\frac{\ln(\frac{dN}{dx_W}) - \ln 0.07}{\ln 2.0 - \ln 0.07}\right)^2 + \left(\frac{x_W - 0.65}{0.50(1.01 - 0.29)}\right)^2 = 1, \quad (E6)$$

valid for $0.29 \leq x_W \leq 1.01$. Solving for dN/dx_W then gives

$$\frac{dN}{dx_W} = 0.07 \left(\frac{2.0}{0.07}\right)^{\sqrt{1-7.7(x_W - 0.65)^2}}.$$
 (E7)

Substituting into Eq. (E5) $\gamma = \frac{M_x}{M_W} x_W$, $\frac{dN}{d\gamma} = \frac{M_W}{M_\chi} \frac{dN}{dx_W}$, and $\frac{dN}{dx_W}$ given in Eq. (E7), we obtain the desired onedimensional integral for the secondary lepton spectrum, per *W*:

$$\frac{dN_{\nu}(x_{\ell})}{dx_{\ell}} = 0.07(\text{BR}) \int_{x_{-}}^{x_{+}} \frac{dx_{W}}{\sqrt{x_{W}^{2} - (\frac{M_{W}}{M_{\chi}})^{2}}} \times \left(\frac{2.0}{0.07}\right)^{\sqrt{1 - 7.7(x_{W} - 0.65)^{2}}}.$$
 (E8)

The integration limits are

$$x_{+} = 1 + \frac{M_{W}^{2}}{4M_{\chi}^{2}},$$

and

$$x_{-} = \left(x_{\ell} + \frac{M_W^2}{4x_{\ell}M_{\chi}^2}\right)$$

The range of x_{ℓ} for the leptons from W decay is $\left[\frac{M_W^2}{4M_{\chi}^2}, 1\right]$. The relevant branching ratios [42] are BR($W \rightarrow \nu e$) = 11%, BR($Z \rightarrow \nu \bar{\nu}$) = 6.7%, and BR($Z \rightarrow \ell^+ \ell^-$) = 3.4%, each per single flavor mode, e, μ , or τ . We show the resulting lepton spectrum, without the BR factor, in Fig. 5.

⁴If the *W* polarization is not neglected, then the *W* decay amplitude includes Wigner functions $d_{\mu_i\mu_f}^1(\theta)$, which introduce a linear $\cos\theta$ or $\sin\theta$ term into Eq. (E1).

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