# Short-baseline electron neutrino disappearance, tritium beta decay, and neutrinoless double-beta decay

Carlo Giunti[\\*](#page-0-0)

INFN, Sezione di Torino, Via P. Giuria 1, I–10125 Torino, Italy

Marco Laveder<sup>[†](#page-0-1)</sup>

<span id="page-0-3"></span><span id="page-0-2"></span>Dipartimento di Fisica "G. Galilei," Università di Padova, and INFN, Sezione di Padova, Via F. Marzolo 8, I–35131 Padova, Italy (Received 31 May 2010; published 22 September 2010)

> We consider the interpretation of the MiniBooNE low-energy anomaly and the gallium radioactive source experiments anomaly in terms of short-baseline electron neutrino disappearance in the framework of  $3 + 1$  four-neutrino mixing schemes. The separate fits of MiniBooNE and gallium data are highly compatible, with close best-fit values of the effective oscillation parameters  $\Delta m^2$  and  $\sin^2 2\vartheta$ . The combined fit gives  $\Delta m^2 \ge 0.1$  eV<sup>2</sup> and  $0.11 \le \sin^2 2\theta \le 0.48$  at  $2\sigma$ . We consider also the data of the Bugey and Chooz reactor antineutrino oscillation experiments and the limits on the effective electron antineutrino mass in  $\beta$  decay obtained in the Mainz and Troitsk tritium experiments. The fit of the data of these experiments limits the value of  $\sin^2 2\theta$  below 0.10 at  $2\sigma$ . Considering the tension between the neutrino MiniBooNE and gallium data and the antineutrino reactor and tritium data as a statistical fluctuation, we perform a combined fit which gives  $\Delta m^2 \approx 2$  eV and  $0.01 \le \sin^2 2\theta \le 0.13$  at  $2\sigma$ . Assuming a hierarchy of masses  $m_1$ ,  $m_2$ ,  $m_3 \ll m_4$ , the predicted contributions of  $m_4$  to the effective neutrino masses in  $\beta$  decay and neutrinoless double- $\beta$  decay are, respectively, between about 0.06 and 0.49 and between about 0.003 and 0.07 eV at  $2\sigma$ . We also consider the possibility of reconciling the tension between the neutrino MiniBooNE and gallium data and the antineutrino reactor and tritium data with different mixings in the neutrino and antineutrino sectors. We find a  $2.6\sigma$  indication of a mixing angle asymmetry.

DOI: [10.1103/PhysRevD.82.053005](http://dx.doi.org/10.1103/PhysRevD.82.053005) PACS numbers: 14.60.Pq, 14.60.Lm, 14.60.St

## I. INTRODUCTION

Neutrino oscillations have been observed in solar, atmospheric, and long-baseline reactor and accelerator experiments. The data of these experiments are well fitted in the framework of three-neutrino mixing, in which the three flavor neutrinos  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$  are unitary linear combinations of three massive neutrinos  $\nu_1$ ,  $\nu_2$ , and  $\nu_3$  with the solar (SOL) and atmospheric (ATM) squared-mass differences

$$
\Delta m_{21}^2 = \Delta m_{\text{SOL}}^2 \simeq 8 \times 10^{-5} \text{ eV}^2, \tag{1}
$$

<span id="page-0-5"></span>
$$
|\Delta m_{31}^2| \simeq |\Delta m_{32}^2| = \Delta m_{\text{ATM}}^2 \simeq 2 \times 10^{-3} \text{ eV}^2, \quad (2)
$$

where  $\Delta m_{jk}^2 = m_j^2 - m_k^2$  and  $m_j$  is the mass of the neutrino<br> $\kappa$  (see Refs. [1, 8])  $\nu_i$  (see Refs. [[1](#page-11-0)[–8](#page-12-0)]).

Besides these well-established observations of neutrino oscillations, there are at least three anomalies which could be signals of short-baseline (SBL) neutrino oscillations generated by a larger squared-mass difference: the LSND  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$  signal [[9](#page-12-1)], the gallium radioactive source experi-<br>ments anomaly [10,111, and the MiniBooNE low-energy ments anomaly [[10,](#page-12-2)[11](#page-12-3)], and the MiniBooNE low-energy anomaly [\[12\]](#page-12-4). In this paper we consider the MiniBooNE and gallium anomalies, which can be explained by shortbaseline electron neutrino disappearance [[13](#page-12-5)–[15](#page-12-6)] in the effective framework of four-neutrino mixing, as explained in Secs. [II](#page-1-0) and [III](#page-4-0). On the other hand, the LSND anomaly is disfavored by the results of the MiniBooNE  $v_{\mu} \rightarrow v_{e}$ experiment [[12](#page-12-4),[16](#page-12-7)] and may require another explanation [\[17–](#page-12-8)[23\]](#page-12-9).

In Refs. [\[13](#page-12-5)[,15\]](#page-12-6) we proposed to explain the MiniBooNE low-energy anomaly [[12](#page-12-4),[16](#page-12-7)] through the disappearance of electron neutrinos due to very-short-baseline oscillations into sterile neutrinos generated by a squared-mass difference  $\Delta m^2$  larger than about 20 eV<sup>2</sup>. In that case, the analysis of the MiniBooNE data is simplified by the fact that the effective survival probability  $P_{\nu_e \to \nu_e}$  is practically constant in the MiniBooNE energy range from 200 to 3000 MeV. In this paper we extend the analysis of MiniBooNE data to lower values of  $\Delta m^2$ , considering the resulting energy dependence of the effective SBL electron neutrino and antineutrino survival probability

<span id="page-0-4"></span>
$$
P_{\substack{(-)\\ \nu_e \to \nu_e}}^{\text{SBL}}(L, E) = 1 - \sin^2 2\vartheta \sin^2 \left(\frac{\Delta m^2 L}{4E}\right),\tag{3}
$$

where  $L$  is the neutrino path length and  $E$  is the neutrino energy (CPT invariance implies that the survival probabilities of neutrinos and antineutrinos are equal; see Ref. [\[8](#page-12-0)]).

<span id="page-0-0"></span>[<sup>\\*</sup>A](#page-0-2)lso at Department of Theoretical Physics, University of Torino, Italy.

giunti@to.infn.it [†](#page-0-3)

<span id="page-0-1"></span>laveder@pd.infn.it

#### CARLO GIUNTI AND MARCO LAVEDER PHYSICAL REVIEW D 82, 053005 (2010)

The two-neutrino-like effective short-baseline survival probability in Eq. [\(3](#page-0-4)) is obtained in four-neutrino schemes (see Refs. [\[1](#page-11-0)[,4,](#page-11-1)[6](#page-12-10)[,7](#page-12-11)]), which are the simplest extension of three-neutrino mixing schemes which can accommodate the two small solar and atmospheric squared-mass differences in Eqs. [\(2\)](#page-0-5) and ([40](#page-9-0)), and one larger squared-mass difference for short-baseline neutrino oscillations,

$$
|\Delta m_{41}^2| = \Delta m^2 \gtrsim 0.1 \text{ eV}^2. \tag{4}
$$

<span id="page-1-1"></span>The existence of a fourth massive neutrino corresponds, in the flavor basis, to the existence of a sterile neutrino  $\nu_s$ .

In this paper we consider  $3 + 1$  four-neutrino schemes, since  $2 + 2$  four-neutrino schemes are disfavored by the combined constraints on active-sterile transitions in solar and atmospheric neutrino experiments [[4](#page-11-1)]. For simplicity, we consider only  $3 + 1$  four-neutrino schemes with

$$
m_1, m_2, m_3 \ll m_4,
$$
 (5)

<span id="page-1-4"></span>which give the  $\Delta m_{41}^2$  $\Delta m_{41}^2$  $\Delta m_{41}^2$  in Eq. (4) and appear to be more<br>natural, than the other possible  $3+1$  four-neutrino natural than the other possible  $3 + 1$  four-neutrino schemes in which either three neutrinos or all four neutrinos are almost degenerate at a mass scale larger than  $\sqrt{\Delta m^2}$  (see Refs. [[1](#page-11-0)[,4,](#page-11-1)[6](#page-12-10)[,7](#page-12-11)]).<br>In 3 + 1 four-neutrino

<span id="page-1-5"></span>In  $3 + 1$  four-neutrino schemes the effective mixing angle in the effective short-baseline electron neutrino survival probability in Eq.  $(3)$  $(3)$  is given by (see Refs. [\[1,](#page-11-0)[4](#page-11-1)[,6](#page-12-10),[7](#page-12-11)])

$$
\sin^2 2\vartheta = 4|U_{e4}|^2(1 - |U_{e4}|^2). \tag{6}
$$

In this paper we assume that the value of  $|U_{\mu 4}|^2$  is so small<br>that the offective short begaling muon neutring symival that the effective short-baseline muon neutrino survival probability is practically equal to unity and short-baseline  $\overrightarrow{v}_{\mu} \leftrightarrows \overrightarrow{v}_{e}$  are negligible.<sup>1</sup> This assumption is justified by the lack of any indication of  $\nu_\mu \rightarrow \nu_e$  transitions in the MiniBooNE experiment [[12](#page-12-4),[16](#page-12-7)] and the limits on shortbaseline muon neutrino disappearance found in the CDHSW [[24](#page-12-12)], CCFR [\[25\]](#page-12-13), and MiniBooNE [[26](#page-12-14)] experiments. We do not consider the MiniBooNE antineutrino data [\[27\]](#page-12-15), which have at present statistical uncertainties which are too large to constrain new physics [[15](#page-12-6)].

The plan of the paper is as follows. In Sec. [II](#page-1-0) we discuss the analysis of MiniBooNE data. In Sec. [III](#page-4-0) we present an update of the analysis of gallium data published in Ref. [[14](#page-12-16)] and the combined analysis of MiniBooNE and gallium data. In Sec. [IV](#page-5-0) we discuss the implications of the measurements of the effective electron neutrino mass in tritium  $\beta$ -decay experiments and their combination with reactor neutrino oscillation data. In Sec. [V](#page-6-0) we present the results of the combined analysis of MiniBooNE, gallium, reactor, and tritium data and in Sec. [VI](#page-8-0) we present the corresponding predictions for the effective masses measured in  $\beta$ -decay and neutrinoless double- $\beta$ -decay experiments. In Sec. [VII](#page-9-1) we calculate the mixing angle asymmetry between the neutrino and antineutrino sectors which could explain the tension between the neutrino and antineutrino data under our shortbaseline  $\nu_e$ -disappearance hypothesis. In Sec. [VIII](#page-10-0) we draw the conclusions.

### II. MINIBOONE

<span id="page-1-0"></span>The MiniBooNE experiment was made with the purpose of checking the indication of  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$  oscillations generated by a  $\Lambda m^2 \ge 0.1$  eV<sup>2</sup> found in the LSND experiment ated by a  $\Delta m^2 \ge 0.1$  eV<sup>2</sup> found in the LSND experiment [\[9\]](#page-12-1). The MiniBooNE Collaboration did not find any indication of such oscillations in the  $\nu_{\mu} \rightarrow \nu_{e}$  channel [\[12,](#page-12-4)[16\]](#page-12-7). On the other hand, the MiniBooNE Collaboration found an anomalous excess of low-energy  $\nu_e$ -like events in the data on the search for  $\nu_{\mu} \rightarrow \nu_{e}$  oscillations [[12](#page-12-4),[16](#page-12-7)], as shown in Fig. [1\(a\)](#page-2-0).

As in Refs. [\[13,](#page-12-5)[15\]](#page-12-6), we consider an explanation of the low-energy MiniBooNE anomaly based on the possible short-baseline disappearance of electron neutrinos, taking into account a possible overall normalization factor  $f_{\nu}$  of the calculated  $\nu_e$ -induced and misidentified  $\nu_\mu$ -induced events which contribute to the observed number of  $v_e$ -like events. The normalization factor  $f_v$  could be due mainly to the uncertainty of the calculated neutrino flux (see Ref. [\[28\]](#page-12-17)). Since the misidentified  $\nu_{\mu}$ -induced and  $\nu_e$ -induced events dominate, respectively, at low and high energies [see Fig. [1\(a\)\]](#page-2-0), the low-energy excess can be fitted with  $f_{\nu} > 1$  and the high-energy data can be fitted compensating  $f_{\nu} > 1$  with the disappearance of  $\nu_e$ 's.

In Refs. [[13](#page-12-5)[,15](#page-12-6)] we considered only very-short-baseline  $\nu_e$  disappearance due to a  $\Delta m^2 \ge 20 \text{ eV}^2$ , which generates a survival probability  $P_{\nu_e \to \nu_e}$  which is constant in the MiniBooNE energy range, from 200 to 3000 MeV. In this paper we extend the analysis to lower values of  $\Delta m^2$ , considering the resulting energy dependence of the survival probability. In this case, the theoretical number of  $\nu_e$ -like events in the *j*th energy bin is given by

$$
N_{\nu_e,j}^{\text{the}} = N_{\nu_e,j}^{\nu_e,\text{the}} + N_{\nu_e,j}^{\nu_\mu,\text{the}},\tag{7}
$$

<span id="page-1-2"></span>where

$$
N_{\nu_e,j}^{\nu_e, \text{the}} = f_{\nu} P_{\nu_e \to \nu_e}^{(j)} N_{\nu_e,j}^{\nu_e, \text{cal}} \tag{8}
$$

<span id="page-1-3"></span>is the number of  $\nu_e$ -induced events and

$$
N_{\nu_e,j}^{\nu_\mu, \text{the}} = f_\nu N_{\nu_e,j}^{\nu_\mu, \text{cal}} \tag{9}
$$

is the number of misidentified  $\nu_{\mu}$ -induced events. Here  $N_{\nu_e,j}^{\nu_e, \text{cal}}$  and  $N_{\nu_e,j}^{\nu_\mu, \text{cal}}$  are, respectively, the number of  $\nu_e$ -induced and misidentified  $\nu_\mu$ -induced events calculated

<sup>&</sup>lt;sup>1</sup>In 3 + 1 four-neutrino schemes the effective short-baseline upon neutrino survival probability has the form in Eq. (3) with muon neutrino survival probability has the form in Eq. [\(3](#page-0-4)) with  $\sin^2 2\theta$  replaced by  $\sin^2 2\theta_{\mu\mu} = 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2)$ . The effective short-baseline  $\overleftrightarrow{\nu}_{\mu} \leftrightarrows \overleftrightarrow{\nu}_{e}$  transition probability is given by  $P_{\substack{(-)\\ \nu_{\mu} \leq \nu_{e}}}^{\text{SBL}}(\mathcal{L}, E) = \sin^2 2\vartheta_{e\mu} \sin^2 \left(\frac{\Delta m^2 \mathcal{L}}{4E}\right), \text{ with } \sin^2 2\vartheta_{e\mu} =$  $4|U_{e4}|^2|U_{\mu 4}|^2$  (see Refs. [[1](#page-11-0),[4,](#page-11-1)[6,](#page-12-10)[7\]](#page-12-11)).

<span id="page-2-3"></span>

<span id="page-2-0"></span>FIG. 1. Expected number of  $\nu_e$  events compared with MiniBooNE data, represented by the black points. The energy bins are numbered with the index j. The uncertainty is represented by the vertical error bars, which represent the sum of statistical and uncorrelated systematic uncertainties. (a) Expected number of  $v_e$ -like events  $N_{v_e,j}^{\text{cal}}$  calculated by the MiniBooNE Collaboration.  $N_{v_e,j}^{\text{cal}}$ is given by the sum of the  $\nu_e$ -induced events  $(N_{\nu_e,j}^{\nu_e,\text{cal}})$  and the misidentified  $\nu_\mu$ -induced events  $(N_{\nu_e,j}^{\nu_\mu,\text{cal}})$ . (b) Best-fit value of the number of  $\nu_e$ -like events  $N_{\nu_e,j}^{\text{the}}$  obtained with the hypothesis of  $\nu_e$  disappearance.  $N_{\nu_e,j}^{\text{the}}$  is given by the sum of  $N_{\nu_e,j}^{\nu_e,\text{the}} = f_{\nu} P_{\nu_e \to \nu_e}^{(j)} N_{\nu_e,j}^{\nu_e,\text{cal}}$  and  $N_{\nu_e,j}^{\nu_\mu, \text{the}} = f_\nu N_{\nu_\mu}^{\text{cal}}$ . The best-fit values of  $f_\nu$ , sin<sup>2</sup>2 $\vartheta$ , and  $\Delta m^2$  are those in the first column of Table [I](#page-3-0) (MB<sub> $\nu$ </sub>).

by the MiniBooNE Collaboration for the jth energy bin [\[29](#page-12-18)[,30\]](#page-12-19).  $P_{\nu_e \to \nu_e}^{(j)}$  is the survival probability of electron neutrinos in Eq. [\(3](#page-0-4)) averaged in the jth energy bin. The average in each bin is calculated using the ntuple file of 17 037 predicted muon-to-electron neutrino full transmutation events given in Ref. [[29](#page-12-18)], which contains information on reconstructed neutrino energy, true neutrino energy, neutrino baseline, and event weight for each event.

The MiniBooNE measurement of a ratio  $1.21 \pm 0.24$  of detected and predicted charged-current quasielastic  $\nu_{\mu}$ events [[31](#page-12-20)] allows a value of  $f<sub>\nu</sub>$  as large as about 15%. In Ref. [[15](#page-12-6)] we used this estimate of the uncertainty of  $f_{\nu}$ in order to constrain its value in the least-squares analysis. Here we use directly the  $\nu_{\mu}$  data given in Ref. [[29](#page-12-18)] for the construction of the MiniBooNE least-squares function

<span id="page-2-2"></span>
$$
\chi^2_{\text{MB}\nu} = \sum_{j=1}^{11} \left( \frac{N_{\nu_e,j}^{\text{the}} - N_{\nu_e,j}^{\text{exp}}}{\sigma_{\nu_e,j}} \right)^2 + \sum_{j,k=1}^{8} (N_{\nu_\mu,j}^{\text{the}} - N_{\nu_\mu,j}^{\text{exp}})
$$

$$
\times (V_{\nu_\mu}^{-1})_{jk} (N_{\nu_\mu,k}^{\text{the}} - N_{\nu_\mu,k}^{\text{exp}}). \tag{10}
$$

Here  $N_{\nu_e,j}^{\text{exp}}$  are the numbers of measured  $\nu_e$ -like events in 11 reconstructed neutrino energy bins and  $N_{\nu_{\mu},j}^{\text{exp}}$  are the numbers of measured  $\nu_{\mu}$  charged-current quasielastic events in 8 reconstructed neutrino energy bins. The theo<span id="page-2-1"></span>retical number of  $\nu_{\mu}$  events in the *j*th energy bin is given by

$$
N_{\nu_{\mu},j}^{\text{the}} = f_{\nu} N_{\nu_{\mu},j}^{\text{cal}},\tag{11}
$$

where  $N_{\nu_{\mu},j}^{\text{cal}}$  is the number of  $\nu_{\mu}$  events calculated by the MiniBooNE Collaboration [\[29\]](#page-12-18). In order to take into account the correct statistical uncertainty corresponding to the rescaling of the number of  $\nu_{\mu}$  events due to  $f_{\nu}$  in Eq. ([11](#page-2-1)), we used the covariance matrix  $V_{\nu_n}$  given by

$$
(V_{\nu_{\mu}})_{jk} = (V_{\nu_{\mu}}^{\text{cal}})_{jk} + (f_{\nu} - 1)N_{\nu_{\mu},j}^{\text{cal}} \delta_{jk}
$$
(12)

where  $V_{\nu_{\mu}}^{\text{cal}}$  is the 8  $\times$  8 covariance matrix of  $\nu_{\mu}$  events presented by the MiniBooNE Collaboration in Ref. [[29\]](#page-12-18). We did not use the complete  $19 \times 19$  covariance matrix of  $\nu$  and  $\nu$  events given in Ref. [29] because the correla $v_e$  and  $v_u$  events given in Ref. [\[29\]](#page-12-18) because the correlations involving  $\nu_e$  events have been obtained without taking into account the energy-dependent disappearance of electron neutrinos that we want to test. Assuming the correlations given in that  $19 \times 19$  covariance matrix would<br>suppress the energy dependence of  $P^{(j)}$ . Therefore, for suppress the energy dependence of  $P_{\nu_e \to \nu_e}^{(j)}$ . Therefore, for the uncertainties  $\sigma_{v_e, j}$  in Eq. ([10](#page-2-2)) we used only the diagonal elements of the  $\nu_e$  covariance matrix  $V_{\nu_e}^{\text{cal}}$  given in Ref. [[29](#page-12-18)], corrected by the change of statistical uncertainty corresponding to the variation of expected events due to  $f_{\nu}$ 

### CARLO GIUNTI AND MARCO LAVEDER PHYSICAL REVIEW D 82, 053005 (2010)

<span id="page-3-0"></span>TABLE I. Values of  $\chi^2$ , number of degrees of freedom (NDF) and goodness of fit (GoF) for the fit of different combinations of MiniBooNE (MB<sub>v</sub>), gallium (Ga), and reactor (Re) data. The first three lines correspond to the case of  $f_v = 1$  and no oscillations (Null hyp.). The following six lines correspond to the case  $f_{\nu} > 1$  and  $\nu_e$  disappearance (Our hyp.). The last three lines give the parameter goodness-of-fit (PG) [\[32\]](#page-12-21). In the MB<sub>v</sub> column, the value of  $\chi^2$  in the Null. hyp., the number of degrees of freedom in the Null. hyp. (which is equal to the number of energy bins), and the value of  $\chi^2_{min}$  in Our hyp. are shown as the sum of the contributions of the first three low-energy u, bins and the other u, and u, energy bins. In the MB + Ga and three low-energy  $v_e$  bins and the other  $v_e$  and  $v_\mu$  energy bins. In the MB<sub>v</sub> + Ga and  $(MB_\nu + Ga) + (Re + {^3H})$  columns the value of  $\chi^2_{\text{min}}$  in Our hyp. is shown as the sum of the contribution of the first three MiniBooNE low-energy  $\nu_e$  bins and the other contributions.

		$MB_{\nu}$	Ga	$MB_{\nu}$ + Ga	$Re + {}^{3}H$	$(MB_{\nu} + Ga) + (Re + {}^{3}H)$
Null hyp.	$\chi^2$	$14.3 + 5.4$	9.4		51.5	
	<b>NDF</b>	$3 + 16$	4		58	
	GoF	0.41	0.051		0.71	
Our hyp.	$\chi^2_{\rm min}$	$2.0 + 7.6$	1.8	$2.2 + 9.2$	49.1	$4.1 + 63.4$
	<b>NDF</b>	16	2	20	56	78
	GoF	0.89	0.40	0.93	0.73	0.80
	$\sin^2 2\vartheta_{\rm bf}$	0.32	0.27	0.28	0.042	0.062
		1.84	2.09	1.92	1.85	1.85
	$\frac{\Delta m^2_{\rm bf}}{f_\nu^{\rm bf}}$	1.26		1.25		1.17
PG	$\Delta \chi^2_{\rm min}$			0.098	0.01	6.97
	<b>NDF</b>			2	2	2
	GoF			0.95	0.99	0.03

and  $P_{\nu_e \to \nu_e}^{(j)}$  in Eqs. ([7](#page-1-2))–[\(9\)](#page-1-3):

$$
\sigma_{\nu_{e},j}^{2} = (V_{\nu_{e}}^{\text{cal}})_{jj} + N_{\nu_{e},j}^{\text{the}} - N_{\nu_{e},j}^{\text{cal}}, \tag{13}
$$

with  $N_{\nu_e,j}^{\text{cal}} = N_{\nu_e,j}^{\nu_e,\text{cal}} + N_{\nu_e,j}^{\nu_\mu,\text{cal}}$ .

The result of the minimization of  $\chi^2_{\text{MB}\nu}$  is shown in<br>a 1(b) in which the solid histogram corresponds to the Fig. [1\(b\)](#page-2-0), in which the solid histogram corresponds to the best-fit values of  $f_{\nu}$ , sin<sup>2</sup>2 $\vartheta$ , and  $\Delta m^2$  in the first column of Table [I.](#page-3-0) From Fig. [1\(b\)](#page-2-0), one can see that the fit is acceptable for all the  $\nu_e$  energy bins, including the first three bins which are out of fit in Fig.  $1(a)$ . [I](#page-3-0)n Table I we give separately the contribution to  $\chi^2_{\text{MB}}$  of the first three low-<br>energy u bins and the sum of the contributions of the other energy  $v_e$  bins and the sum of the contributions of the other  $\nu_e$  energy bins and all the  $\nu_\mu$  energy bins. In this way one can see that with  $f_{\nu} = 1$  and  $P_{\nu_e \to \nu_e} = 1$  (null hypothesis), although the global value  $\chi^2 = 19.7$  is compatible with the number of degrees of freedom,  $NDF = 19$ , almost all the  $\chi^2$  is due to the anomalous contribution 14.3 of the first three low-energy  $v_e$  bins, whereas the other 16  $v_e$  and  $v_\mu$ energy bins are overfitted, with the excessively small  $\chi^2$ contribution of 5.4. This overfitting, which is probably due to an overestimate of the uncertainties, remains in the fit of the data with our hypothesis of  $f_{\nu} > 1$  and  $\nu_e$  disappearance. On the other hand, our hypothesis clearly explains the low-energy anomaly reducing the  $\chi^2$  contribution of the first three low-energy  $v_e$  bins to the acceptable best-fit value of 2.0.

Figure [2](#page-3-1) shows the allowed regions in the  $\sin^2 2\theta -$ <br> $m^2$  plane and the marginal  $\Delta v^2 = v^2 - v^2$ , 's for  $\Delta m^2$  plane and the marginal  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$ 's for<br>sin<sup>2</sup>2.2 and  $\Delta m^2$  from which one can infer the correspond- $\sin^2 2\vartheta$  and  $\Delta m^2$ , from which one can infer the corresponding uncorrelated allowed intervals. One can see that the indication in favor of neutrino oscillations is not strong,

being at the level of about 79% C.L.  $(1.2\sigma)$ . It is interesting to notice that the best-fit value of  $\Delta m^2$  is about 2 eV<sup>2</sup>, which is approximately the same best-fit value obtained in Ref. [[14\]](#page-12-16) from the fit of the neutrino data of gallium radioactive source experiments and the antineutrino data of the Bugey and Chooz reactor experiments under the hypothesis of  $\nu_e$  and  $\bar{\nu}_e$  disappearance. The results of the

<span id="page-3-1"></span>

FIG. 2. Allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane and marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  and  $\Delta m^2$  obtained from the fit of MiniBooNE neutrino data. The best-fit point is indicated by a +.

combined analysis of MiniBooNE neutrino data and the data of these other experiments is discussed in the following sections.

## <span id="page-4-2"></span><span id="page-4-0"></span>III. GALLIUM RADIOACTIVE SOURCE EXPERIMENTS

The GALLEX [[11](#page-12-3)[,33](#page-12-22)[,34\]](#page-12-23) and SAGE [\[10](#page-12-2)[,35–](#page-12-24)[37\]](#page-12-25) Collaborations tested the respective gallium solar neutrino detectors in so-called ''gallium radioactive source experiments'' which consist of the detection of electron neutrinos produced by intense artificial  ${}^{51}Cr$  and  ${}^{37}Ar$  radioactive sources placed inside the detectors. Taking into account the uncertainty of the cross section of the detection process  $v_e$  + <sup>71</sup>Ga  $\rightarrow$  <sup>71</sup>Ge + e<sup>-</sup> estimated in Ref. [\[38\]](#page-12-26), the ratios  $\overline{R}$  of measured and predicted  $^{71}$ Ge event rates are

<span id="page-4-4"></span><span id="page-4-3"></span>
$$
R_{\rm Cr1}^{\rm GALLEX} = 0.95^{+0.11}_{-0.12},\tag{14}
$$

$$
R_{\text{Cr2}}^{\text{GALLEX}} = 0.81_{-0.11}^{+0.10},\tag{15}
$$

$$
R_{Cr}^{\text{SAGE}} = 0.95_{-0.12}^{+0.12},\tag{16}
$$

$$
R_{\text{Ar}}^{\text{SAGE}} = 0.79_{-0.10}^{+0.09},\tag{17}
$$

<span id="page-4-5"></span><span id="page-4-1"></span>and the average ratio is

$$
R^{Ga} = 0.86^{+0.05}_{-0.05}.\tag{18}
$$

Thus, the number of measured events is about  $2.7\sigma$  smaller than the prediction.

The theoretical prediction of the rate is based on the calculation of the detection cross section presented in Ref. [[38](#page-12-26)]. It is possible that a part of the observed deficit is due to an overestimation of this cross section [\[10](#page-12-2)[,37,](#page-12-25)[39\]](#page-12-27), because only the cross section of the transition from the ground state of  $^{71}$ Ga to the ground state of  $^{71}$ Ga is known with precision from the measured rate of electron capture decay of  $71$ Ge to  $71$ Ga. Electron neutrinos produced by  $51$ Cr and 37Ar radioactive sources can be absorbed also through transitions from the ground state of  $71$ Ga to two excited states of  $^{71}$ Ge, with cross sections which are inferred using a nuclear model from  $p + {}^{71}Ga \rightarrow {}^{71}Ge + n$  measurements [\[40\]](#page-12-28). This calculation has large uncertainties [\[41](#page-12-29)[,42\]](#page-12-30). However, since the contribution of the transitions to the two excited states is only 5% [[38](#page-12-26)], even the complete absence of such transitions would reduce  $R<sup>Ga</sup>$  to about  $0.90^{+0.05}_{-0.05}$ , leaving an anomaly of about  $1.8\sigma$ .

Here we consider the electron neutrino disappearance explanation of the gallium radioactive source experiments anomaly [\[13–](#page-12-5)[15](#page-12-6)[,43](#page-12-31)–[45](#page-12-32)] (another interesting explanation through quantum decoherence in neutrino oscillations has been proposed in Ref. [[21](#page-12-33)]).

In Ref. [[14](#page-12-16)] we have analyzed the data of the gallium radioactive source experiments in terms of the effective



FIG. 3. Allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane and marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  and  $\Delta m^2$  obtained from the combined fit of the results of the two GALLEX <sup>51</sup>Cr radioactive source of the results of the two GALLEX <sup>51</sup>Cr radioactive source<br>experiments and the SAGE <sup>51</sup>Cr and <sup>37</sup>Ar radioactive source experiments and the SAGE <sup>51</sup>Cr and <sup>37</sup>Ar radioactive source<br>experiments. The best-fit point corresponding to  $v^2$  is indiexperiments. The best-fit point corresponding to  $\chi^2_{\text{min}}$  is indicated by a + cated by a +.

survival probability in Eq. ([3\)](#page-0-4). Here we update that analysis taking into account the revised value of  $R_{\text{Crl}}^{\text{GALLEX}}$  in<br>Eq. (14) published recently in Ref. [11] and taking into Eq. ([14](#page-4-1)) published recently in Ref. [[11](#page-12-3)] and taking into account the asymmetric uncertainties of  $R_{\text{Crl}}^{\text{GALLEX}}$ ,<br> $R_{\text{GALLEX}}$  and  $R_{\text{SAGE}}^{\text{SAGE}}$  (which have been symmetrized for  $R_{\rm cr2}^{\rm GALLEX}$ , and  $R_{\rm Ar}^{\rm SAGE}$  (which have been symmetrized for simplicity in the analysis presented in Ref. [141] simplicity in the analysis presented in Ref. [\[14](#page-12-16)]). Following the method described in Ref. [\[14\]](#page-12-16), we obtained the best-fit values of  $\sin^2 2\vartheta$  and  $\Delta m^2$  in the second column of Table [I](#page-3-0) and the allowed regions in the  $\sin^2 2\theta - \Delta m^2$ plane shown in Fig. [3.](#page-4-2) The indication in favor of neutrino oscillations is at the level of about 98% C.L.  $(2.3\sigma)$ .

From Table [2](#page-3-1) and the comparison of Figs. [2](#page-3-1) and [3](#page-4-2) one can see that the fits of MiniBooNE and gallium data lead to remarkably similar results: the best-fit values of the oscillation parameters are very close and the allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane are highly compatible. This is certainly an impressive success of our hypothesis of electron neutrino disappearance.

The results of the combined fit of MiniBooNE and gallium data are shown in the third column of Table [I](#page-3-0) and in Fig. [4.](#page-5-1) The separate data sets are well fitted by the electron neutrino disappearance hypothesis: the  $\chi^2$  contribution of the first three MiniBooNE low-energy  $v_e$  bins is 2.2, that of the other 16 MiniBooNE  $\nu_e$  and  $\nu_\mu$  energy bins is 7.4, and that of the 4 gallium data is 1.9. The consistency of the combined fit is also supported by the excellent value

<span id="page-5-1"></span>

FIG. 4. Allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane and marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  and  $\Delta m^2$  obtained from the combined fit of the results of MiniBooNE neutrino data and the data of the gallium radioactive source experiments. The best-fit point corresponding to  $\chi^2_{\text{min}}$  is indicated by a +.

of the parameter goodness of fit. Combining the two data sets improves the indication in favor of neutrino oscillations to the level of about 99.5% C.L.  $(2.8\sigma)$ .

## <span id="page-5-0"></span>IV. REACTOR AND TRITIUM EXPERIMENTS

The indication of electron neutrino disappearance that we have found from the analysis of MiniBooNE and gallium data must be confronted with the results of reactor electron antineutrino experiments. Assuming CPT invariance, the survival probabilities of neutrinos and antineutrinos are equal (see Ref. [\[8](#page-12-0)]). Thus, we can combine directly the results presented in the previous section with the results of the analysis of the data of the Bugey and Chooz reactor experiments obtained in Ref. [[14](#page-12-16)]. We are encouraged in this task by the coincidence of the best-fit value of  $\Delta m^2$  at about 2 eV<sup>2</sup>.

In addition to reactor neutrino experiments, also tritium  $\beta$ -decay experiments give information on the masses and mixing of neutrinos through the measurement of the electron energy spectrum in the process

$$
{}^{3}\text{H} \rightarrow {}^{3}\text{He} + e^{-} + \bar{\nu}_{e}. \tag{19}
$$

<span id="page-5-5"></span>The most accurate measurements of the effective electron neutrino mass (see Refs. [\[2](#page-11-2),[8](#page-12-0)])

$$
m_{\beta} = \left(\sum_{k} |U_{ek}|^2 m_k^2\right)^{1/2} \tag{20}
$$

<span id="page-5-4"></span>have been performed in the Mainz [[46\]](#page-12-34) and Troitsk [\[47\]](#page-12-35):

$$
m_{\beta}^2 = -0.6 \pm 2.2 \pm 2.1 \text{ eV}^2 \quad \text{(Mainz)}, \qquad (21)
$$

$$
m_{\beta}^{2} = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^{2} \quad \text{(Troitsk)}.\tag{22}
$$

<span id="page-5-3"></span>These measurements can be interpreted and combined in order to derive upper bounds for the effective mass  $m<sub>\beta</sub>$ through a  $\chi^2$  analysis in the physical region  $m_\beta^2 \ge 0$ . In<br>Fig. 5 we plotted the corresponding  $\Lambda x^2$  as a function of Fig. [5](#page-5-2) we plotted the corresponding  $\Delta \chi^2$ 's as a function of  $m_{\beta}$ . One can see that

$$
m_{\beta} \le 2.3 \text{ eV}
$$
 (Mainz, 95% C.L.), (23)

$$
m_{\beta} \le 2.0 \text{ eV}
$$
 (Troitsk, 95% C.L.), (24)

<span id="page-5-6"></span>in approximate agreement with the corresponding values in Refs. [\[46](#page-12-34)[,47\]](#page-12-35). The combined upper bound is

$$
m_{\beta} \le 1.8 \text{ eV} \quad \text{(Mainz + Troitsk, 95\% C.L.).} \tag{25}
$$

In  $3 + 1$  four-neutrino schemes, taking into account that the mass splittings among  $m_1$ ,  $m_2$ , and  $m_3$  in Eqs. [\(2](#page-0-5)) and [\(40\)](#page-9-0) are negligible for a measurement of  $m<sub>\beta</sub>$  at the scale of 0.1–1 eV, the effective mass is given by

<span id="page-5-2"></span>

FIG. 5.  $\Delta \chi^2$  as a function of  $m_\beta$ . The horizontal lines corre-<br>spond to the indicated value of confidence level. The dashed and spond to the indicated value of confidence level. The dashed and dotted lines have been obtained, respectively, from the results in Eqs. ([21](#page-5-3)) and [\(22\)](#page-5-4) of the Mainz and Troitsk tritium  $\beta$ -decay experiments. The solid line is the result of the combined fit.

$$
m_{\beta}^2 \simeq (1 - |U_{e4}|^2) m_1^2 + |U_{e4}|^2 m_4^2 = m_1^2 + |U_{e4}|^2 \Delta m_{41}^2.
$$
\n(26)

<span id="page-6-2"></span>In  $3 + 1$  schemes of the type in Eq. ([5](#page-1-4)), we have

$$
m_{\beta} \ge |U_{e4}| \sqrt{\Delta m^2}.
$$
 (27)

The effective mixing angle in short-baseline electron neutrino disappearance experiments is related to  $|U_{e4}|$  by Eq. ([6\)](#page-1-5). Inverting this relation and taking into account that the value of  $|U_{e4}|$  must be small in order to fit the data of solar neutrino experiments with neutrino oscillations, we have

$$
|U_{e4}|^2 = \frac{1}{2}(1 - \sqrt{1 - \sin^2 2\vartheta}).
$$
 (28)

<span id="page-6-3"></span>In Fig. [6](#page-6-1) we show the limits in the  $\sin^2 2\theta - \Delta m^2$  plane obtained from Eqs. [\(27\)](#page-6-2) and [\(28\)](#page-6-3) and the results in Eqs. ([21](#page-5-3)) and ([22\)](#page-5-4) of the Mainz and Troitsk tritium  $\beta$ -decay experiments. Notice that for small values of  $\sin^2 2\theta$  the bounds are practically linear in the log-log plot in Fig. [6,](#page-6-1) because in this case  $|U_{e4}|^2 \approx \sin^2 2\vartheta/4$  and the inequality in Eq. (27) leads to the inequality in Eq. ([27](#page-6-2)) leads to

$$
\log \Delta m^2 \lesssim 2 \log 2 + 2 \log m_{\beta}^{\text{ub}} - \log \sin^2 2 \vartheta, \qquad (29)
$$

where  $m_\beta^{\text{ub}}$  is the upper bound for  $m_\beta$ .

<span id="page-6-1"></span>

FIG. 6. Allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane and marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  and  $\Delta m^2$  obtained from the combined fit of the results of the Bugey and Chooz reactor experiments and the results of the Mainz and Troitsk tritium  $\beta$ -decay experiments. The three lines in the upper-right corner are the exclusion curves obtained from the results of the Mainz and Troitsk tritium  $\beta$ -decay experiments alone. The best-fit point corresponding to  $\chi^2_{\text{min}}$  is indicated by a +.

In the fourth column of Table [I](#page-3-0) and in Fig. [6](#page-6-1) we report the results of the analysis of the data of the Bugey and Chooz reactor experiments presented in Ref.  $[14]$  $[14]$  $[14]$ <sup>2</sup> with the addition in the analysis of the results of the Mainz and Troitsk tritium  $\beta$ -decay experiments, which affect the high- $\Delta m^2$  region. As already commented on in Ref. [[14\]](#page-12-16), the reactor data are compatible with both the null hypothesis of the absence of electron antineutrino disappearance and our hypothesis of electron antineutrino disappearance, with a hint in favor of electron antineutrino oscillations due to a  $\Delta m^2$  of about 2 eV.

#### V. COMBINED ANALYSIS

<span id="page-6-0"></span>The results of the combined analysis of MiniBooNE, gallium, reactor, and tritium data are presented in the last column of Table [I](#page-3-0) and in Fig. [7.](#page-7-0) One can see that the goodness of fit is high. The separate data sets are fitted fairly well by the electron neutrino disappearance hypothesis: the  $\chi^2$  contribution of the first three MiniBooNE lowenergy  $v_e$  bins is 4.1, that of the other 16 MiniBooNE  $v_e$ and  $v_{\mu}$  energy bins is 7.5, that of the 4 gallium data is 6.3, that of the 56 reactor degrees of freedom is 49.0, and that of the 2 tritium degrees of freedom is 0.57. On the other hand, the 3% parameter goodness of fit of the combined analysis of neutrino MiniBooNE and gallium data and antineutrino reactor and tritium data is rather low.

This low compatibility of the neutrino and antineutrino data sets is illustrated in Fig. [8](#page-7-1), where we have plotted the marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  obtained with the analysis of different data sets. $3$  One can see that

- (1) The neutrino MiniBooNE and gallium data agree to indicate a value of  $\sin^2 2\theta$  between about 0.11 and 0.48 at  $2\sigma$ .
- (2) The antineutrino reactor data indicate a value of  $\sin^2 2\vartheta$  smaller than about 0.10 at  $2\sigma$ . The tritium data are practically irrelevant for the determination of  $\sin^2 2\vartheta$ .
- (3) The combined analysis is dominated by the reactor data and indicates a value of  $\sin^2 2\theta$  between about 0.01 and 0.13 at  $2\sigma$ .

The discrepancy between the neutrino and antineutrino determinations of  $\sin^2 2\theta$  is about  $2\sigma$ , in rough agreement with the above-mentioned 3% parameter goodness of fit of the combined analysis. In fact, the  $2\sigma$  disagreement be-

<sup>&</sup>lt;sup>2</sup>As an erratum, let us notice that in the fourth column of Table III in Ref. [\[14\]](#page-12-16) there is a small mistake in the evaluation of the parameter goodness of fit. The correct values are  $\Delta \chi_{min}^2 = 0.52$  and GoF = 0.47. We also notice that in the version of Ref. [14] published in Phys. Rev. D the value of  $\sin^2 2\theta_{1.5}$  for the Ref. [[14](#page-12-16)] published in Phys. Rev. D the value of  $\sin^2 2\theta_{\rm bf}$  for the  $Ga + Bu + Ch$  analysis (last column of Table III in Ref. [\[14\]](#page-12-16)) is different from the correct one, which is 0.054 (see the arXiv version of Ref.  $[14]$  $[14]$  $[14]$ ).

We thank the anonymous referee of Phys. Rev. D for suggesting this interesting figure.

<span id="page-7-0"></span>

FIG. 7. Allowed regions in the  $\sin^2 2\theta - \Delta m^2$  plane and marginal  $\Delta \chi^2$ 's for sin<sup>2</sup>2 $\vartheta$  and  $\Delta m^2$  obtained from the combined fit of the results of MiniBooNE, gallium, reactor, and tritium experiments. The best-fit point corresponding to  $\chi^2_{\text{min}}$  is indicated by  $2 +$  The three lines in the upper-right corner give the cated by a +. The three lines in the upper-right corner give the  $1\sigma$ ,  $2\sigma$ , and  $3\sigma$  limits in the sin<sup>2</sup> $2\vartheta - \Delta m^2$  plane obtained from Eqs.  $(20)$  and  $(28)$  and the results in Eqs.  $(21)$  $(21)$  $(21)$  and  $(22)$  of the Mainz and Troitsk tritium  $\beta$ -decay experiments.

<span id="page-7-1"></span>

FIG. 8. Marginal  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  for  $\sin^2 2\vartheta$  obtained from<br>the analysis of different combinations of MiniBooNE gallium the analysis of different combinations of MiniBooNE, gallium, reactor, and tritium data. The marginal  $\Delta \chi^2$  obtained from the analysis of reactor data alone and that obtained from the combined analysis of reactor and tritium data are shown by the same line, since they practically coincide.

<span id="page-7-2"></span>

FIG. 9. Expected number of MiniBooNE  $\nu_e$  events in the bestfit result of the combined analysis of MiniBooNE, gallium, reactor, and tritium data (last column in Table [I\)](#page-3-0). The notation is the same as in Fig. [1.](#page-2-3)

tween the neutrino and antineutrino data sets is entirely due to the different requirements on the value of  $\sin^2 2\theta$ , whereas they nicely agree on a best-fit value of  $\Delta m^2$  at about 2  $eV^2$ .

Since the 3% parameter goodness of fit of the combined analysis shows a tension between the neutrino and antineutrino data (under our  $\nu_e$ -disappearance hypothesis) but is not sufficiently small to reject with confidence the compatibility of the neutrino and antineutrino data sets,4 in the following part of this section and in Sec. [VI](#page-8-0) we consider the results and implications of the combined analysis. In Sec. [VII](#page-9-1) we consider a possible difference between the effective mixing angles in the neutrino and antineutrino sectors.

Although the combined analysis of neutrino and antineutrino data favors smaller values of  $\sin^2 2\vartheta$  than those obtained from the analysis of MiniBooNE and gallium data alone, the fit of the MiniBooNE and gallium data remains better than in the case of no oscillations and  $f_{\nu} = 1$ .

Figure [9](#page-7-2) shows the fit of MiniBooNE  $\nu_e$  data corresponding to the best-fit result of the combined analysis. One can see that the fit of the first three low-energy bins is not as good as that in Fig. [1\(b\),](#page-2-0) but it is nevertheless acceptable and much better than that in Fig. [1\(a\).](#page-2-0)

<sup>&</sup>lt;sup>4</sup>For example, the review on Statistics in the 2000 edition of the Review of Particle Physics [\[48\]](#page-12-36) says that if the goodness of fit "is larger than an agreed-upon value (0.001, 0.01, or 0.05 are common choices), the data are consistent with the assumptions.''

For the gallium source experiments, the best-fit values of the oscillation parameters give  $R_{\text{C1}}^{\text{GALLEX}} = R_{\text{C2}}^{\text{GALLEX}} = 0.97$   $R_{\text{SAGE}} = 0.96$  and  $R_{\text{SAGE}} = 0.96$  Therefore the 0.97,  $R_{\text{CF}}^{\text{SAGE}} = 0.96$ , and  $R_{\text{AT}}^{\text{SAGE}} = 0.96$ . Therefore, the experimental values of  $R_{\text{CT}}^{\text{GALEX}}$  and  $R_{\text{CT}}^{\text{SAGE}}$  in Eqs. [\(14\)](#page-4-1) and ([16](#page-4-3)) are fitted very well and the gallium  $\chi^2$  contribution of 6.3 is almost equally due to the loose fits of  $R_{Cr2}^{GALLEX}$ and  $R_{\text{Ar}}^{\text{SAGE}}$  in Eqs. ([15](#page-4-4)) and [\(17\)](#page-4-5).<br>Considering the combined

Considering the combined fit of the results of MiniBooNE, gallium, reactor, and tritium data as a fair indication in favor of a possible short-baseline electron neutrino disappearance generated by the effective mixing parameters  $\Delta m^2 \simeq 2$  eV and  $0.01 \lesssim \sin^2 2\vartheta \lesssim 0.13$ , in the next section we present the corresponding predictions for the effective neutrino masses in  $\beta$ -decay and neutrinoless double- $\beta$ -decay experiments which could be measured in future experiments.

## <span id="page-8-0"></span>VI. PREDICTIONS FOR BETA-DECAY AND NEUTRINOLESS DOUBLE-BETA-DECAY **EXPERIMENTS**

In this section we present predictions for the effective neutrino masses in  $\beta$ -decay and neutrinoless  $double$ - $\beta$ -decay experiments obtained as a consequence of the combined fit of MiniBooNE, gallium, reactor, and tritium data discussed in the previous section.

Figure [10](#page-8-1) shows the residual  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  as a set function of the contribution  $|U_{e4}| \sqrt{\Delta m^2}$  to the effective<br>mass  $m_e$  in *B*-decay experiments [see Eq. (27)]. Since mass  $m_\beta$  in  $\beta$ -decay experiments [see Eq. ([27](#page-6-2))]. Since from the last column of Table [I](#page-3-0) we have  $\sin^2 2\theta_{\rm bf} \ll 1$ , we obtain

$$
|U_{e4}|_{\text{bf}}^2 \simeq \frac{\sin^2 2\vartheta_{\text{bf}}}{4} = 0.016,\tag{30}
$$

and the best-fit value of  $|U_{e4}| \sqrt{\Delta m^2}$  is

$$
(|U_{e4}|\sqrt{\Delta m^2})_{\rm bf} = 0.17 \text{ eV},\tag{31}
$$

and

$$
0.06 \le |U_{e4}| \sqrt{\Delta m^2} \le 0.49 \text{ eV} \quad \text{at } 2\sigma. \tag{32}
$$

This prediction is relevant for the KATRIN experiment [\[49\]](#page-12-37), which is under construction and scheduled to start in 2012. The expected sensitivity of about 0.2 eV at 90% C.L. may be sufficient to observe a positive effect if  $|U_{e4}| \sqrt{\Delta m^2}$  is sufficiently large, as allowed by  $\Delta \chi^2$  in<br>Fig. 10 Fig. [10.](#page-8-1)

If massive neutrinos are Majorana particles, neutrinoless double- $\beta$  decay is possible, with a decay rate proportional to the effective Majorana mass (see Refs. [[2,](#page-11-2)[8](#page-12-0)[,50–](#page-12-38)[52](#page-12-39)])

$$
m_{2\beta} = |\sum_{k} U_{ek}^2 m_k|.\tag{33}
$$

The results of the combined fit of MiniBooNE, gallium, reactor, and tritium data discussed in Sec. [IV](#page-5-0) allow us to

<span id="page-8-1"></span>

FIG. 10.  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  as a function of the contribution  $|U| \sqrt{\Delta m^2}$  to the effective *B* decay electron neutrino mass *m*  $|U_{e4}| \sqrt{\Delta m^2}$  to the effective  $\beta$ -decay electron neutrino mass  $m_\beta$ <br>in four neutrino, schemes, obtained, from the analysis of in four-neutrino schemes obtained from the analysis of MiniBooNE and gallium data (dashed line), from the analysis of reactor and tritium data (dotted line), and from the combined analysis of the two sets of data (solid line).

estimate the contribution of the heaviest massive neutrino  $\nu_4$  to  $m_{2\beta}$ , which is approximately given by  $|U_{e4}|^2 \sqrt{\Delta m^2}$ ,<br>taking into account the mass hierarchy in Eq. (5) taking into account the mass hierarchy in Eq. ([5\)](#page-1-4).

Figure [11](#page-9-2) shows  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  as a function of the contribution  $|U_{e4}|^2 \sqrt{\Delta m^2}$  in four-neutrino schemes to  $m_{2\beta}$ .<br>The best-fit value is The best-fit value is

$$
(|U_{e4}|^2 \sqrt{\Delta m^2})_{\text{bf}} = 0.02 \text{ eV},\tag{34}
$$

<span id="page-8-2"></span>and

$$
0.003 \le |U_{e4}|^2 \sqrt{\Delta m^2} \le 0.07 \text{ eV} \quad \text{at } 2\sigma. \tag{35}
$$

<span id="page-8-3"></span>This range must be confronted with the expected contributions to  $m_{2\beta}$  coming from the three light massive neutrinos<br> $\cdots$  and  $\cdots$  Assuming a biggaraby of masses  $\nu_1$ ,  $\nu_2$ , and  $\nu_3$ . Assuming a hierarchy of masses,

$$
m_1 \ll m_2 \ll m_3 \ll m_4,\tag{36}
$$

which is the most natural case compatible with the hierarchy in Eq. [\(5](#page-1-4)), we have

$$
m_{2\beta} \simeq |U_{e2}^2 \sqrt{\Delta m_{\text{SOL}}^2} + U_{e3}^2 \sqrt{\Delta m_{\text{ATM}}^2} + U_{e4}^2 \sqrt{\Delta m^2}|, \tag{37}
$$

where we have neglected the contribution of the lightest massive neutrino  $\nu_1$ . From the 3 $\sigma$  upper limits of the three-<br>neutrino mixing parameters given in Ref. [4], we obtain

<span id="page-9-2"></span>

FIG. 11.  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  as a function of the contribution  $|U| = 2\sqrt{\Delta m^2}$  to the effective peutripoless double  $\beta$  decay  $|U_{e4}|^2 \sqrt{\Delta m^2}$  to the effective neutrinoless double- $\beta$  decay<br>Majorana mass means in four-neutrino schemes obtained from Majorana mass  $m_{2\beta}$  in four-neutrino schemes obtained from<br>the analysis of MiniBooME and sallium data (dashed line) the analysis of MiniBooNE and gallium data (dashed line), from the analysis of reactor and tritium data (dotted line), and from the combined analysis of the two sets of data (solid line).

$$
2 \times 10^{-3} \le |U_{e2}|^2 \sqrt{\Delta m_{\text{SOL}}^2} \le 4 \times 10^{-3} \text{ eV}, \qquad (38)
$$

$$
|U_{e3}|^2 \sqrt{\Delta m_{\text{ATM}}^2} \lesssim 3 \times 10^{-3} \text{ eV.}
$$
 (39)

Therefore, strong cancellations between the contributions of  $\nu_2$  and  $\nu_3$  are possible (albeit not likely [[53\]](#page-12-40)), whereas the range in Eq. [\(35\)](#page-8-2) disfavors strong cancellations between the contributions of  $\nu_2$  and  $\nu_3$  and the contribution of  $\nu_4$ . In this case,  $m_{2\beta} \simeq |U_{e4}|^2 \sqrt{\Delta m^2}$  leading to a pos-<br>sible observation of neutrinoless double-*B* decay in future sible observation of neutrinoless double- $\beta$  decay in future experiments which will be sensitive to values of  $m_{2,0}$ smaller than  $10^{-1}$  eV (e.g. CUORE [\[54\]](#page-12-41), EXO [\[55\]](#page-12-42), SuperNEMO [[56\]](#page-12-43); see the review in Ref. [\[52\]](#page-12-39)).

On the other hand, if neutrinoless double- $\beta$  decay experiments which are sensitive to values of  $m_{2\beta}$  of the order<br>of  $10^{-1}$  eV, (e.g. CUORICINO, 1571, CERDA, 1581) of  $10^{-1}$  eV (e.g. CUORICINO [[57](#page-12-44)], GERDA [\[58\]](#page-12-45), Majorana [[59\]](#page-12-46); see the review in Ref. [[52\]](#page-12-39)) will see a positive signal, maybe compatible with the signal asserted in Ref. [[60](#page-12-47)], the mass hierarchy in Eq. ([36](#page-8-3)) will become unlikely and the favorite  $3 + 1$  four-neutrino schemes will be those in which the three light neutrinos  $\nu_1$ ,  $\nu_2$ , and  $\nu_3$ are almost degenerate at the mass scale of  $m_{2\beta}$ .

## VII. MIXING ANGLE ASYMMETRY?

<span id="page-9-1"></span>The tension between neutrino and antineutrino data discussed in Sec. [V](#page-6-0) could be due to a difference of the effective mixing angles in the neutrino and antineutrino sectors. Such a difference could be due to a violation of the fundamental CPT symmetry or to another unknown mechanism. Phenomenological analyses of different masses and mixings for neutrinos and antineutrinos have been presented in several publications [\[61](#page-12-48)[–74\]](#page-12-49).

In this section we consider the possibility that neutrinos and antineutrinos have different effective masses and mixings in short-baseline  $\nu_e$  and  $\bar{\nu}_e$  disappearance experi-<br>ments. We fit the neutrino and antineutrino data with the ments. We fit the neutrino and antineutrino data with the survival probabilities

$$
P_{\nu_e \to \nu_e}^{\text{SBL}}(L, E) = 1 - \sin^2 2\vartheta_\nu \sin^2 \left(\frac{\Delta m_\nu^2 L}{4E}\right),\tag{40}
$$

$$
P_{\bar{\nu}_e \to \bar{\nu}_e}^{\text{SBL}}(L, E) = 1 - \sin^2 2\vartheta_{\bar{\nu}} \sin^2 \left(\frac{\Delta m_{\bar{\nu}}^2 L}{4E}\right).
$$
 (41)

<span id="page-9-0"></span>The results for the two fits are those presented in Fig. [4](#page-5-1) and the third column in Table [I](#page-3-0) for neutrinos ( $MB_\nu + Ga$ ) and Fig. [6](#page-6-1) and the fourth column in Table [I](#page-3-0) for antineutrinos  $(Re + 3H)$ .

Since the fit of the data does not require a difference of  $\Delta m_{\nu}^2$  and  $\Delta m_{\nu}^2$ , we consider only the mixing angle asymmetry metry

$$
A_{\sin^2 2\vartheta} = \sin^2 2\vartheta_\nu - \sin^2 2\vartheta_{\bar{\nu}}.\tag{42}
$$

Figure [12](#page-9-3) shows the marginal  $\Delta \chi^2$  as a function of  $A_{\sin^2 2\vartheta}$ .

<span id="page-9-3"></span>

FIG. 12. Marginal  $\Delta \chi^2 = \chi^2 - \chi^2_{\text{min}}$  as a function of the mix-<br>ing angle asymmetry  $A_{\perp}$  as a ing angle asymmetry  $A_{\sin^2 2\vartheta}$ .

The best-fit value of  $A_{\sin^2 2\vartheta}$  is

$$
A_{\sin^2 2\vartheta}^{\text{bf}} = 0.23, \tag{43}
$$

and the  $2\sigma$  allowed range of  $A_{\sin^2 2\vartheta}$  is

$$
0.06 \le A_{\sin^2 2\vartheta} \le 0.45, \tag{44}
$$

but there is no limit on the asymmetry at  $3\sigma$ . The statistical significance of  $A_{\sin^2 2\vartheta} > 0$  is 99.14% C.L. (2.6 $\sigma$ ).

It is interesting to note that a difference between neutrino and antineutrino mixings can be tested in  $\beta$ -decay experiments by searching for different effective neutrino masses in  $\beta^-$  and  $\beta^+$  decays. The prediction for the contribution  $(|U_{e4}| \sqrt{\Delta m^2})_{\bar{\nu}_e}$  to the effective electron anti-<br>neutrino mass in  $B^-$  decays from the analysis of antineuneutrino mass in  $\beta^-$  decays from the analysis of antineutrino reactor and tritium data can be obtained from the dotted line in Fig. [10](#page-8-1): the best fit is

$$
(|U_{e4}|\sqrt{\Delta m^2})_{\bar{\nu}_e}^{\text{bf}} = 0.14 \text{ eV}, \tag{45}
$$

and

$$
(|U_{e4}|\sqrt{\Delta m^2})_{\bar{\nu}_e} \le 1.07 \text{ eV} \quad \text{at } 2\sigma. \tag{46}
$$

For the contribution  $(|U_{e4}|\sqrt{\Delta m^2})_{v_e}$  to the effective electron peutrino mass in  $R^+$  decays we must consider the tron neutrino mass in  $\beta^+$  decays we must consider the dashed line in Fig. [10](#page-8-1), which has been obtained from the analysis of MiniBooNE and gallium neutrino data: the best fit is

$$
(|U_{e4}|\sqrt{\Delta m^2})_{\nu_e}^{\text{bf}} = 0.38 \text{ eV}, \qquad (47)
$$

<span id="page-10-1"></span>and

$$
(|U_{e4}|\sqrt{\Delta m^2})_{\nu_e} \ge 0.21 \text{ eV} \quad \text{at } 2\sigma. \tag{48}
$$

Unfortunately the existing and foreseen experiments are  $\beta$ <sup>-</sup>-decay experiments for which the contribution  $(|U_{e4}|\sqrt{\Delta m^2})_{\bar{v}_e}$  to the effective electron antineutrino mass<br>is expected to be small. The future  $B^-$ -decay experiment is expected to be small. The future  $\beta^-$ -decay experiment will use either tritium (KATRIN [\[49\]](#page-12-37)) or  $^{187}$ Re (MARE [\[75\]](#page-12-50)). If the mixing difference between the neutrino and antineutrino sectors will be confirmed with highest confidence by future neutrino oscillation data it will be interesting to study the possibility of making  $\beta^+$ -decay experiments for the search of the effective electron antineutrino mass, for which the dashed line in Fig. [10](#page-8-1) and Eq. ([48](#page-10-1)) give a reachable lower limit.

We do not consider here neutrinoless double- $\beta$  decay in the case of a neutrino-antineutrino mixing difference, since the Majorana nature of neutrinos requires a treatment which goes well beyond the purposes of this paper (see Ref. [\[76\]](#page-13-0)).

### VIII. CONCLUSIONS

<span id="page-10-0"></span>In this paper we have discussed a neutrino oscillation interpretation of the MiniBooNE low-energy anomaly and the gallium radioactive source experiments anomaly in the framework of  $3 + 1$  four-neutrino mixing schemes. We have shown that the combined fit of MiniBooNE and gallium data indicates a possible short-baseline electron neutrino disappearance generated by effective oscillation parameters  $\Delta m^2 \ge 0.1$  eV<sup>2</sup> and  $0.11 \le \sin^2 2\vartheta \le 0.48$  at  $2\sigma$ , with the best fit at  $\Delta m^2 \approx 2$  eV<sup>2</sup> and sin<sup>2</sup> $2\vartheta \approx 0.3$  (see Fig. [4\)](#page-5-1).

We have also considered the data of the Bugey and Chooz reactor neutrino oscillation experiments and the results of the Mainz and Troitsk tritium  $\beta$ -decay experiments, which imply an upper bound on the effective electron neutrino mass of about 2 eV [see Fig. [5](#page-5-2) and the combined upper bound in Eq. [\(25](#page-5-6))]. As already discussed in Ref. [\[14\]](#page-12-16), the Bugey data give a faint indication of a possible short-baseline electron neutrino disappearance generated by effective oscillation parameters  $\Delta m^2 \simeq$  $2 eV^2$  and  $\sin^2 2\theta \approx 0.04$ , which is compatible with Chooz and Tritium data (see Fig. [6\)](#page-6-1).

In Sec. [V](#page-6-0) we have discussed the tension between the neutrino MiniBooNE and gallium data and the antineutrino reactor and tritium data. Considering such tension as a statistical fluctuation, we have presented the results of the combined analysis of MiniBooNE, gallium, reactor,<br>and tritium data:  $\Delta m^2 \approx 2 \text{ eV}^2$  and  $0.01 \le \sin^2 2\theta \le$ and tritium data:  $\Delta m^2 \approx 2 \text{ eV}^2$  and  $0.01 \le \sin^2 2\vartheta \le$ <br>0.13 at  $2\sigma$  with the best fit at  $\Delta m^2 \approx 2 \text{ eV}^2$  and  $\sin^2 2\vartheta \approx$ 0.13 at  $2\sigma$ , with the best fit at  $\Delta m^2 \approx 2 \text{ eV}^2$  and  $\sin^2 2\vartheta \approx 0.06$  (see Fig. 7) 0:06 (see Fig. [7](#page-7-0)).

In Sec. [VI,](#page-8-0) we have presented predictions for the effective neutrino masses in  $\beta$ -decay and neutrinoless  $double$ - $\beta$ -decay experiments obtained as a consequence of the combined analysis of MiniBooNE, gallium, reactor, and tritium data, assuming the hierarchy of masses in Eq. [\(5](#page-1-4)). The predicted interval for the contribution of  $m_4$ to the effective neutrino mass in  $\beta$  decay is between about 0.06 and 0.49 eV at  $2\sigma$ . The upper part of this interval may be reached by the KATRIN experiment [\[49\]](#page-12-37). For neutrinoless double- $\beta$  decay we obtained a prediction for the contribution of  $m<sub>4</sub>$  to the effective neutrino mass between about 0.003 and 0.07 eV at  $2\sigma$ , which may be reached in future experiments (see Ref. [\[52\]](#page-12-39)).

We also considered, in Sec. [VII,](#page-9-1) the possibility of reconciling the tension between the neutrino MiniBooNE and gallium data and the antineutrino reactor and tritium data discussed in Sec. [V](#page-6-0) with different mixings in the neutrino and antineutrino sectors. We found a  $2.6\sigma$  indication of a mixing angle asymmetry (99.14% C.L.). We pointed out the possibility of checking the mixing difference between the neutrino and antineutrino sectors by measuring different effective electron antineutrino and neutrino masses in  $\beta^-$ - and  $\beta^+$ -decay experiments.

The indication in favor of short-baseline disappearance of electron neutrinos implies the possible existence of a light sterile neutrino which could have important consequences in physics [[18](#page-12-51)[–20](#page-12-52)[,22,](#page-12-53)[77](#page-13-1)[–83\]](#page-13-2), astrophysics [\[84–](#page-13-3) [92](#page-13-4)], and cosmology [\[93–](#page-13-5)[98](#page-13-6)].

As far as the effective number of neutrino species in cosmology,  $N_{\text{eff}}$ , is concerned, the analysis of 7-years WMAP data has provided the following result:  $N_{\text{eff}}$  = 4.34 $^{+0.86}_{-0.88}$  (68% C.L.) [\[99](#page-13-7)]. In 2011 the Planck experiment will measure  $N_{\text{eff}}$  with a factor of 4 improvement in accuracy with respect to present data [\[100,](#page-13-8)[101](#page-13-9)]. In other words, the possibility of the existence of a fourth light sterile neutrino could be pursued with  $5\sigma$  significance.

Finally, we would like to encourage all experiments which can investigate the hypothesis of short-baseline electron neutrino disappearance.

Starting from 2010, at the same  $L/E$  of MiniBoone, the magnetic off-axis near detector at 280 m of the T2K experiment [[102](#page-13-10)] will count  $\nu_e$  events with expected higher statistics and similar  $\nu_{\mu}$  background contamination. A test of short-baseline oscillations may be done, although the accuracy suffers from the scarce knowledge of the neutrino flux and of the neutrino cross section at 1 GeV energies [\[15](#page-12-6)[,103](#page-13-11)].

A better measurement will be possible with the new CERN-PS neutrino beam [[104,](#page-13-12)[105\]](#page-13-13), thanks to the presence of 2 detectors at 140 m (NEAR) and 885 m (FAR). At  $\Delta m^2 \approx 2$  eV<sup>2</sup>, the oscillation length is about 1 km for 1 GeV neutrino energies. Therefore, one can reduce the systematic error of the Monte Carlo predictions by normalizing the high-energy part of the  $v_e$  spectrum at the NEAR location. In addition, a better  $\nu_{\mu}$  background rejection will be possible using the liquid argon technology. The interesting possibility of a  $\nu_e$  tagging in the CERN-PS beam was also studied in this context [\[106\]](#page-13-14).

New measurements with a radioactive source could be made in the SAGE experiment [\[10\]](#page-12-2), with the Borexino detector and with the future LENS detector [[79](#page-13-15)]. At  $\Delta m^2 \approx 2$  eV<sup>2</sup>, the oscillation length is about 1 m for 1 MeV neutrino energies. Therefore Borexino could measure the oscillation pattern over a distance of 4 m (the Borexino radius) using the well-known  $\nu_e$ -e scattering process and with a vertex resolution that at the moment is about 15 cm [\[107\]](#page-13-16).

At the Gran Sasso laboratories a very interesting measurement could be realized by using the ICARUS 600 ton detector and new low-cost and high-power proton cyclotrons under development for commercial use [\[108\]](#page-13-17). These provide electron neutrino beams with energy up to 52 MeV from muon decay at rest. A low-energy  $\nu_e$  disappearance experiment (as well as  $\nu_{\mu}^{\left(1\right)}$  disappearance and  $\nu_{\mu}^{\left(2\right)}\nu_{e}^{\left(1\right)}$  measurements) can be performed with such devices due to the full efficiency of the ICARUS detector at 20 MeVenergies. The expected event rate is about 400 charged-current electron neutrino events per year per ton with the ICARUS detector located at 50 m from the source [[109\]](#page-13-18). The number of events is calculated assuming  $10^{15}\nu_e$ 's per year and a fully efficient detector. Since at  $\Delta m^2 \approx 2 \text{ eV}^2$  the oscillation length is about 20 m for a 20 MeV neutrino energy, it is possible to measure the full oscillation pattern along the beam direction inside the ICARUS volume.

The disappearance of electron neutrinos can be investigated with high accuracy in future near-detector betabeam [[110](#page-13-19)] and neutrino factory [\[73](#page-12-54)[,111](#page-13-20)] experiments in which the neutrino fluxes will be known with high precision.

Furthermore, the MiniBooNE low-energy anomaly may be clarified by the ArgoNeuT, MicroBooNE [\[112](#page-13-21)], and BooNE experiments [[113](#page-13-22)], and the magnetic off-axis near detector at 280 m of the T2K experiment has the unique opportunity to measure the charge of the events of the low-energy anomaly [[114](#page-13-23)].

## ACKNOWLEDGMENTS

We would like to thank E. Bellotti, C. Giganti, A. Longhin, F. Pietropaolo, and A. Rubbia for interesting discussions and suggestions.

Note added.—After the completion of this work, two important experimental results were presented at the Neutrino 2010 Conference:

- (1) The MiniBooNE Collaboration presented updated results on the search for short-baseline  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$ <br>oscillations which are compatible with the I SND oscillations which are compatible with the LSND signal [\[115](#page-13-24),[116](#page-13-25)]. The inclusion of these data in our framework will require a separate analysis in which the assumption of negligible  $|U_{\mu4}|^2$  is relaxed [\[117\]](#page-13-26).
- (2) The MINOS Collaboration presented an indication of a possible difference between the effective mixings of neutrinos and antineutrinos in long-baseline  $\nu_{\mu}$  and  $\bar{\nu}_{\mu}$  disappearance [[118](#page-13-27)]. This indication is<br>analogous to that discussed in Sec. VII analogous to that discussed in Sec. [VII.](#page-9-1)

- <span id="page-11-2"></span><span id="page-11-0"></span>[1] S. M. Bilenky, C. Giunti, and W. Grimus, [Prog. Part. Nucl.](http://dx.doi.org/10.1016/S0146-6410(99)00092-7) Phys. 43[, 1 \(1999\)](http://dx.doi.org/10.1016/S0146-6410(99)00092-7).
- [2] S. M. Bilenky, C. Giunti, J. A. Grifols, and E. Masso, [Phys.](http://dx.doi.org/10.1016/S0370-1573(03)00102-9) Rep. 379[, 69 \(2003\).](http://dx.doi.org/10.1016/S0370-1573(03)00102-9)
- [3] C. Giunti and M. Laveder, in Developments in Quantum Physics—2004, edited by F. Columbus and V.

Krasnoholovets (Nova Science, Hauppauge, NY, 2003), pp. 197–254.

- <span id="page-11-1"></span>[4] M. Maltoni, T. Schwetz, M. Tortola, and J. Valle, [New J.](http://dx.doi.org/10.1088/1367-2630/6/1/122) Phys. 6[, 122 \(2004\)](http://dx.doi.org/10.1088/1367-2630/6/1/122).
- [5] G. L. Fogli, E. Lisi, A. Marrone, and A. Palazzo, [Prog.](http://dx.doi.org/10.1016/j.ppnp.2005.08.002) [Part. Nucl. Phys.](http://dx.doi.org/10.1016/j.ppnp.2005.08.002) 57, 742 (2006).
- <span id="page-12-10"></span>[6] A. Strumia and F. Vissani, [arXiv:hep-ph/0606054.](http://arXiv.org/abs/hep-ph/0606054)
- <span id="page-12-11"></span>[7] M. C. Gonzalez-Garcia and M. Maltoni, [Phys. Rep.](http://dx.doi.org/10.1016/j.physrep.2007.12.004) 460, 1 [\(2008\)](http://dx.doi.org/10.1016/j.physrep.2007.12.004).
- <span id="page-12-0"></span>[8] C. Giunti and C. W. Kim, Fundamentals of Neutrino Physics and Astrophysics (Oxford University Press, Oxford, UK, 2007), pp. 1–728.
- <span id="page-12-1"></span>[9] A. Aguilar et al. (LSND Collaboration), [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.64.112007) 64, [112007 \(2001\)](http://dx.doi.org/10.1103/PhysRevD.64.112007).
- <span id="page-12-2"></span>[10] J.N. Abdurashitov et al. (SAGE Collaboration), [Phys.](http://dx.doi.org/10.1103/PhysRevC.80.015807) Rev. C 80[, 015807 \(2009\).](http://dx.doi.org/10.1103/PhysRevC.80.015807)
- <span id="page-12-3"></span>[11] F. Kaether, W. Hampel, G. Heusser, J. Kiko, and T. Kirsten, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2010.01.030) 685, 47 (2010).
- <span id="page-12-4"></span>[12] A. A. Aguilar-Arevalo (MiniBooNE Collaboration), [Phys.](http://dx.doi.org/10.1103/PhysRevLett.102.101802) Rev. Lett. 102[, 101802 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.102.101802)
- <span id="page-12-5"></span>[13] C. Giunti and M. Laveder, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.77.093002) 77, 093002 [\(2008\)](http://dx.doi.org/10.1103/PhysRevD.77.093002).
- <span id="page-12-16"></span>[14] M. A. Acero, C. Giunti, and M. Laveder, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.78.073009) 78, [073009 \(2008\)](http://dx.doi.org/10.1103/PhysRevD.78.073009) [[arXiv:hep-ph/0405172\]](http://arXiv.org/abs/hep-ph/0405172).
- <span id="page-12-6"></span>[15] C. Giunti and M. Laveder, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.80.013005) 80, 013005 [\(2009\)](http://dx.doi.org/10.1103/PhysRevD.80.013005).
- <span id="page-12-7"></span>[16] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), Phys. Rev. Lett. 98[, 231801 \(2007\).](http://dx.doi.org/10.1103/PhysRevLett.98.231801)
- <span id="page-12-8"></span>[17] A. Bueno, M. Campanelli, M. Laveder, J. Rico, and A. Rubbia, [J. High Energy Phys. 06 \(2001\) 032.](http://dx.doi.org/10.1088/1126-6708/2001/06/032)
- <span id="page-12-51"></span>[18] M. Maltoni and T. Schwetz, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.76.093005) 76, 093005 [\(2007\)](http://dx.doi.org/10.1103/PhysRevD.76.093005).
- [19] S. Goswami and W. Rodejohann, [J. High Energy Phys. 10](http://dx.doi.org/10.1088/1126-6708/2007/10/073) [\(2007\) 073.](http://dx.doi.org/10.1088/1126-6708/2007/10/073)
- <span id="page-12-52"></span>[20] T. Schwetz, [J. High Energy Phys. 02 \(2008\) 011.](http://dx.doi.org/10.1088/1126-6708/2008/02/011)
- <span id="page-12-33"></span>[21] Y. Farzan, T. Schwetz, and A. Y. Smirnov, [J. High Energy](http://dx.doi.org/10.1088/1126-6708/2008/07/067) [Phys. 07 \(2008\) 067.](http://dx.doi.org/10.1088/1126-6708/2008/07/067)
- <span id="page-12-53"></span>[22] G. Karagiorgi, Z. Djurcic, J. Conrad, M. H. Shaevitz, and M. Sorel, Phys. Rev. D 80[, 073001 \(2009\)](http://dx.doi.org/10.1103/PhysRevD.80.073001).
- <span id="page-12-9"></span>[23] S. Hollenberg, O. Micu, and H. Pas, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.80.053010) 80, [053010 \(2009\)](http://dx.doi.org/10.1103/PhysRevD.80.053010).
- <span id="page-12-12"></span>[24] F. Dydak et al. (CDHSW Collaboration), [Phys. Lett.](http://dx.doi.org/10.1016/0370-2693(84)90688-9) 134B[, 281 \(1984\)](http://dx.doi.org/10.1016/0370-2693(84)90688-9).
- <span id="page-12-13"></span>[25] I. E. Stockdale et al. (CCFR Collaboration), [Z. Phys. C](http://dx.doi.org/10.1007/BF01642480) 27, [53 \(1985\).](http://dx.doi.org/10.1007/BF01642480)
- <span id="page-12-14"></span>[26] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), Phys. Rev. Lett. 103[, 061802 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.103.061802)
- <span id="page-12-15"></span>[27] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), Phys. Rev. Lett. 103[, 111801 \(2009\).](http://dx.doi.org/10.1103/PhysRevLett.103.111801)
- <span id="page-12-17"></span>[28] S. E. Kopp, Phys. Rep. 439[, 101 \(2007\).](http://dx.doi.org/10.1016/j.physrep.2006.11.004)
- <span id="page-12-18"></span>[29] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), [http://www-boone.fnal.gov/for\\_physicists/data\\_release/](http://www-boone.fnal.gov/for_physicists/data_release/lowe/) [lowe/.](http://www-boone.fnal.gov/for_physicists/data_release/lowe/)
- <span id="page-12-19"></span>[30] W.C. Louis (private communication).
- <span id="page-12-20"></span>[31] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), Phys. Rev. Lett. 100[, 032301 \(2008\).](http://dx.doi.org/10.1103/PhysRevLett.100.032301)
- <span id="page-12-21"></span>[32] M. Maltoni and T. Schwetz, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.68.033020) 68, 033020 [\(2003\)](http://dx.doi.org/10.1103/PhysRevD.68.033020).
- <span id="page-12-22"></span>[33] P. Anselmann et al. (GALLEX Collaboration), [Phys. Lett.](http://dx.doi.org/10.1016/0370-2693(94)01586-2) B 342[, 440 \(1995\)](http://dx.doi.org/10.1016/0370-2693(94)01586-2).
- <span id="page-12-23"></span>[34] W. Hampel et al. (GALLEX Collaboration), [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(97)01562-1) 420[, 114 \(1998\).](http://dx.doi.org/10.1016/S0370-2693(97)01562-1)
- <span id="page-12-24"></span>[35] J.N. Abdurashitov et al. (SAGE Collaboration), [Phys.](http://dx.doi.org/10.1103/PhysRevLett.77.4708) Rev. Lett. 77[, 4708 \(1996\).](http://dx.doi.org/10.1103/PhysRevLett.77.4708)
- [36] J.N. Abdurashitov et al. (SAGE Collaboration), [Phys.](http://dx.doi.org/10.1103/PhysRevC.59.2246) Rev. C 59[, 2246 \(1999\)](http://dx.doi.org/10.1103/PhysRevC.59.2246).
- <span id="page-12-25"></span>[37] J. N. Abdurashitov et al., Phys. Rev. C 73[, 045805 \(2006\).](http://dx.doi.org/10.1103/PhysRevC.73.045805)
- <span id="page-12-26"></span>[38] J. N. Bahcall, Phys. Rev. C **56**[, 3391 \(1997\)](http://dx.doi.org/10.1103/PhysRevC.56.3391).
- <span id="page-12-27"></span>[39] G. Fogli, E. Lisi, A. Marrone, and A. Palazzo, 3rd NO-VE International Workshop on: Neutrino Oscillations in Venice: 50 Years after the Neutrino Experimental Discovery, Venice, Italy, 2006.
- <span id="page-12-28"></span>[40] D. Krofcheck et al., [Phys. Rev. Lett.](http://dx.doi.org/10.1103/PhysRevLett.55.1051) **55**, 1051 (1985).
- <span id="page-12-29"></span>[41] N. Hata and W. Haxton, [Phys. Lett. B](http://dx.doi.org/10.1016/0370-2693(95)00598-F) 353, 422 (1995).
- <span id="page-12-30"></span>[42] W. C. Haxton, *[Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(98)00581-4)* 431, 110 (1998).
- <span id="page-12-31"></span>[43] J. N. Bahcall, P. I. Krastev, and E. Lisi, [Phys. Lett. B](http://dx.doi.org/10.1016/0370-2693(95)00111-W) 348, [121 \(1995\)](http://dx.doi.org/10.1016/0370-2693(95)00111-W).
- [44] M. Laveder, [Nucl. Phys. B, Proc. Suppl.](http://dx.doi.org/10.1016/j.nuclphysbps.2007.02.037) **168**, 344 (2007).
- <span id="page-12-32"></span>[45] C. Giunti and M. Laveder, [Mod. Phys. Lett. A](http://dx.doi.org/10.1142/S0217732307025455) 22, 2499 [\(2007\)](http://dx.doi.org/10.1142/S0217732307025455).
- <span id="page-12-34"></span>[46] C. Kraus et al., [Eur. Phys. J. C](http://dx.doi.org/10.1140/epjc/s2005-02139-7) 40, 447 (2005).
- <span id="page-12-35"></span>[47] V. M. Lobashev, Nucl. Phys. **A719**[, C153 \(2003\)](http://dx.doi.org/10.1016/S0375-9474(03)00985-0).
- <span id="page-12-36"></span>[48] D. Groom *et al.*, Eur. Phys. J. C **15**, 1 (2000).
- <span id="page-12-37"></span>[49] M. Beck, [arXiv:0910.4862.](http://arXiv.org/abs/0910.4862)
- <span id="page-12-38"></span>[50] S. R. Elliott and P. Vogel, [Annu. Rev. Nucl. Part. Sci.](http://dx.doi.org/10.1146/annurev.nucl.52.050102.090641) 52, [115 \(2002\)](http://dx.doi.org/10.1146/annurev.nucl.52.050102.090641).
- [51] S. R. Elliott and J. Engel, J. Phys. G 30[, R183 \(2004\)](http://dx.doi.org/10.1088/0954-3899/30/9/R01).
- <span id="page-12-39"></span>[52] I. Avignone, T. Frank, S.R. Elliott, and J. Engel, [Rev.](http://dx.doi.org/10.1103/RevModPhys.80.481) Mod. Phys. 80[, 481 \(2008\)](http://dx.doi.org/10.1103/RevModPhys.80.481).
- <span id="page-12-40"></span>[53] C. Giunti, Phys. Rev. D **61**[, 036002 \(2000\)](http://dx.doi.org/10.1103/PhysRevD.61.036002).
- <span id="page-12-41"></span>[54] M. Pedretti et al., [Int. J. Mod. Phys. A](http://dx.doi.org/10.1142/S0217751X08042183) 23, 3395 (2008).
- <span id="page-12-43"></span><span id="page-12-42"></span>[55] N. Ackerman (EXO Collaboration), [arXiv:0909.1826.](http://arXiv.org/abs/0909.1826)
- [56] I. Nasteva et al. (SuperNEMO Collaboration), Proc. Sci., EPS-HEP2009 (2009) 463 [[arXiv:0909.3167](http://arXiv.org/abs/0909.3167)].
- <span id="page-12-44"></span>[57] C. Arnaboldi et al. (CUORICINO Collaboration), [Phys.](http://dx.doi.org/10.1103/PhysRevC.78.035502) Rev. C 78[, 035502 \(2008\).](http://dx.doi.org/10.1103/PhysRevC.78.035502)
- <span id="page-12-45"></span>[58] I. Abt et al. (GERDA Collaboration), [arXiv:hep-ex/](http://arXiv.org/abs/hep-ex/0404039) [0404039.](http://arXiv.org/abs/hep-ex/0404039)
- <span id="page-12-46"></span>[59] C.E. Aalseth et al. (MAJORANA Collaboration), [J. Phys.](http://dx.doi.org/10.1088/1742-6596/203/1/012057) Conf. Ser. 203[, 012057 \(2010\).](http://dx.doi.org/10.1088/1742-6596/203/1/012057)
- <span id="page-12-47"></span>[60] H. V. Klapdor-Kleingrothaus and I. V. Krivosheina, [Mod.](http://dx.doi.org/10.1142/S0217732306020937) Phys. Lett. A 21[, 1547 \(2006\)](http://dx.doi.org/10.1142/S0217732306020937).
- <span id="page-12-48"></span>[61] H. Murayama and T. Yanagida, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(01)01136-4) 520, 263 [\(2001\)](http://dx.doi.org/10.1016/S0370-2693(01)01136-4).
- [62] G. Barenboim, L. Borissov, J. Lykken, and A. Y. Smirnov, [J. High Energy Phys. 10 \(2002\) 001.](http://dx.doi.org/10.1088/1126-6708/2002/10/001)
- [63] S. M. Bilenky, M. Freund, M. Lindner, T. Ohlsson, and W. Winter, Phys. Rev. D 65[, 073024 \(2002\)](http://dx.doi.org/10.1103/PhysRevD.65.073024).
- [64] G. Barenboim, L. Borissov, and J. Lykken, Phys. Lett. B 534, 106 (2002).
- [65] A. Strumia, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(02)02042-7) **539**, 91 (2002).
- [66] J. N. Bahcall, V. Barger, and D. Marfatia, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(02)01714-8) 534[, 120 \(2002\).](http://dx.doi.org/10.1016/S0370-2693(02)01714-8)
- [67] H. Murayama, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2004.06.106) **597**, 73 (2004).
- [68] V. Barger, D. Marfatia, and K. Whisnant, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2003.10.004) 576[, 303 \(2003\).](http://dx.doi.org/10.1016/j.physletb.2003.10.004)
- [69] H. Minakata and S. Uchinami, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.72.105007) 72, 105007 [\(2005\)](http://dx.doi.org/10.1103/PhysRevD.72.105007).
- [70] M.C. Gonzalez-Garcia, M. Maltoni, and T. Schwetz, Phys. Rev. D 68[, 053007 \(2003\).](http://dx.doi.org/10.1103/PhysRevD.68.053007)
- [71] S. Antusch and E. Fernandez-Martinez, [Phys. Lett. B](http://dx.doi.org/10.1016/j.physletb.2008.06.005) 665, [190 \(2008\)](http://dx.doi.org/10.1016/j.physletb.2008.06.005).
- <span id="page-12-54"></span>[72] A. D. Dolgov, [Phys. At. Nucl.](http://dx.doi.org/10.1134/S1063778810040022) **73**, 588 (2010).
- [73] C. Giunti, M. Laveder, and W. Winter, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.80.073005) 80, [073005 \(2009\).](http://dx.doi.org/10.1103/PhysRevD.80.073005)
- <span id="page-12-49"></span>[74] G. Barenboim and J. D. Lykken, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.80.113008) 80, 113008 [\(2009\)](http://dx.doi.org/10.1103/PhysRevD.80.113008).
- <span id="page-12-50"></span>[75] A. Monfardini et al., [Prog. Part. Nucl. Phys.](http://dx.doi.org/10.1016/j.ppnp.2005.11.004) 57, 68 (2006).

### CARLO GIUNTI AND MARCO LAVEDER PHYSICAL REVIEW D 82, 053005 (2010)

- <span id="page-13-0"></span>[76] G. Barenboim, J. F. Beacom, L. Borissov, and B. Kayser, [Phys. Lett. B](http://dx.doi.org/10.1016/S0370-2693(02)01947-0) 537, 227 (2002).
- <span id="page-13-1"></span>[77] A. de Gouvea, J. Jenkins, and N. Vasudevan, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.75.013003) 75[, 013003 \(2007\)](http://dx.doi.org/10.1103/PhysRevD.75.013003).
- [78] G. Karagiorgi et al., Phys. Rev. D 75[, 013011 \(2007\);](http://dx.doi.org/10.1103/PhysRevD.75.013011) [80](http://dx.doi.org/10.1103/PhysRevD.80.099902), [099902\(E\) \(2009\).](http://dx.doi.org/10.1103/PhysRevD.80.099902)
- <span id="page-13-15"></span>[79] C. Grieb, J. Link, and R. S. Raghavan, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.75.093006) 75, [093006 \(2007\)](http://dx.doi.org/10.1103/PhysRevD.75.093006).
- [80] D. C. Latimer, J. Escamilla, and D. J. Ernst, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.75.042501) 75[, 042501 \(2007\)](http://dx.doi.org/10.1103/PhysRevC.75.042501).
- [81] A. Donini et al., [J. High Energy Phys. 12 \(2007\) 013.](http://dx.doi.org/10.1088/1126-6708/2007/12/013)
- [82] A. Bandyopadhyay and S. Choubey, [arXiv:0707.2481.](http://arXiv.org/abs/0707.2481)
- <span id="page-13-2"></span>[83] A. de Gouvea and T. Wytock, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.79.073005) 79, 073005 [\(2009\)](http://dx.doi.org/10.1103/PhysRevD.79.073005).
- <span id="page-13-3"></span>[84] R.L. Awasthi and S. Choubey, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.76.113002) 76, 113002 [\(2007\)](http://dx.doi.org/10.1103/PhysRevD.76.113002).
- [85] S. Choubey, [J. High Energy Phys. 12 \(2007\) 014.](http://dx.doi.org/10.1088/1126-6708/2007/12/014)
- [86] D. Boyanovsky, H. J. de Vega, and N. Sanchez, [Phys. Rev.](http://dx.doi.org/10.1103/PhysRevD.77.043518) D 77[, 043518 \(2008\)](http://dx.doi.org/10.1103/PhysRevD.77.043518).
- [87] G. Gentile, H. S. Zhao, and B. Famaey, [arXiv:0712.1816](http://arXiv.org/abs/0712.1816) [Mon. Not. R. Astron. Soc. (to be published)].
- [88] G. W. Angus, [Mon. Not. R. Astron. Soc.](http://dx.doi.org/10.1111/j.1365-2966.2008.14341.x) 394, 527 (2009).
- [89] A. Donini and O. Yasuda, [arXiv:0806.3029.](http://arXiv.org/abs/0806.3029)
- [90] L. S. Kisslinger, E. M. Henley, and M. B. Johnson, [Mod.](http://dx.doi.org/10.1142/S0217732309031806) Phys. Lett. A 24[, 2507 \(2009\)](http://dx.doi.org/10.1142/S0217732309031806).
- [91] L. S. Kisslinger and S. Pakvasa, [arXiv:0906.4117.](http://arXiv.org/abs/0906.4117)
- <span id="page-13-4"></span>[92] C. Giunti and Y. Li, Phys. Rev. D 80[, 113007 \(2009\)](http://dx.doi.org/10.1103/PhysRevD.80.113007).
- <span id="page-13-5"></span>[93] O. Civitarese and M.E. Mosquera, [Phys. Rev. C](http://dx.doi.org/10.1103/PhysRevC.77.045806) 77, [045806 \(2008\)](http://dx.doi.org/10.1103/PhysRevC.77.045806).
- [94] A. Melchiorri et al., [J. Cosmol. Astropart. Phys. 01 \(2009\)](http://dx.doi.org/10.1088/1475-7516/2009/01/036) [036.](http://dx.doi.org/10.1088/1475-7516/2009/01/036)
- [95] M. A. Acero and J. Lesgourgues, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.79.045026) 79, 045026 [\(2009\)](http://dx.doi.org/10.1103/PhysRevD.79.045026).
- [96] T. M. Nieuwenhuizen, [Europhys. Lett.](http://dx.doi.org/10.1209/0295-5075/86/59001) 86, 59001 (2009).
- [97] G. W. Angus, B. Famaey, and A. Diaferio, [arXiv:0906.3322.](http://arXiv.org/abs/0906.3322)
- <span id="page-13-6"></span>[98] Y. I. Izotov and T. X. Thuan, [Astrophys. J.](http://dx.doi.org/10.1088/2041-8205/710/1/L67) 710, L67 [\(2010\)](http://dx.doi.org/10.1088/2041-8205/710/1/L67).
- <span id="page-13-7"></span>[99] E. Komatsu et al., [arXiv:1001.4538.](http://arXiv.org/abs/1001.4538)
- <span id="page-13-8"></span>[100] K. Ichikawa, T. Sekiguchi, and T. Takahashi, [Phys. Rev. D](http://dx.doi.org/10.1103/PhysRevD.78.083526) 78[, 083526 \(2008\)](http://dx.doi.org/10.1103/PhysRevD.78.083526).
- <span id="page-13-10"></span><span id="page-13-9"></span>[101] S. Galli et al., [arXiv:1005.3808.](http://arXiv.org/abs/1005.3808)
- [102] T. Le (T2K Collaboration), [arXiv:0910.4211.](http://arXiv.org/abs/0910.4211)
- <span id="page-13-11"></span>[103] C. Giganti and A. Longhin (private communication).
- <span id="page-13-12"></span>[104] B. Baibussinov et al., [arXiv:0909.0355.](http://arXiv.org/abs/0909.0355)
- <span id="page-13-13"></span>[105] A. Rubbia CERN PS, 2010, [http://indico.cern.ch/](http://indico.cern.ch/getFile.py/access?contribId=4&resId=0&materialId=slides&confId=87234) [getFile.py/access?contribId=4&resId=0&materialId=](http://indico.cern.ch/getFile.py/access?contribId=4&resId=0&materialId=slides&confId=87234) [slides&confId=87234.](http://indico.cern.ch/getFile.py/access?contribId=4&resId=0&materialId=slides&confId=87234)
- <span id="page-13-14"></span>[106] L. Ludovici and F. Terranova, [arXiv:1004.2904.](http://arXiv.org/abs/1004.2904)
- <span id="page-13-16"></span>[107] E. Bellotti (private communication).
- <span id="page-13-17"></span>[108] J. Alonso et al., [arXiv:1006.0260.](http://arXiv.org/abs/1006.0260)
- <span id="page-13-18"></span>[109] R. Lazauskas and C. Volpe, [arXiv:1004.0310.](http://arXiv.org/abs/1004.0310)
- <span id="page-13-19"></span>[110] S. K. Agarwalla, P. Huber, and J. M. Link, [J. High Energy](http://dx.doi.org/10.1007/JHEP01(2010)071) [Phys. 01 \(2010\) 071.](http://dx.doi.org/10.1007/JHEP01(2010)071)
- <span id="page-13-20"></span>[111] J. Bernabeu et al., [arXiv:1005.3146.](http://arXiv.org/abs/1005.3146)
- <span id="page-13-21"></span>[112] M. Soderberg (ArgoNeuT Collaboration), Acta Phys. Pol. B 40, 2665 (2009).
- <span id="page-13-22"></span>[113] I. Stancu et al., [arXiv:0910.2698.](http://arXiv.org/abs/0910.2698)
- <span id="page-13-23"></span>[114] M. Laveder, International Workshop Search for Baryon and Lepton Number Violations (LBNL Berkeley, CA, 2007), [http://inpa.lbl.gov/blnv2/files/Friday/Session10/](http://inpa.lbl.gov/blnv2/files/Friday/Session10/Laveder.pdf) [Laveder.pdf](http://inpa.lbl.gov/blnv2/files/Friday/Session10/Laveder.pdf).
- <span id="page-13-24"></span>[115] R. Van de Water, Neutrino 2010, Athens, Greece.
- <span id="page-13-25"></span>[116] A. A. Aguilar-Arevalo et al. (MiniBooNE Collaboration), [arXiv:1007.1150.](http://arXiv.org/abs/1007.1150)
- <span id="page-13-26"></span>[117] C. Giunti and M. Laveder (unpublished).
- <span id="page-13-27"></span>[118] P. Vahle, Neutrino 2010, Athens, Greece.