# Plebanski action extended to a unification of gravity and Yang-Mills theory

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We study a unification of gravity with Yang-Mills fields based on a simple extension of the Plebanski action to a Lie group G which contains the local Lorentz group. The Coleman-Mandula theorem is avoided because the dynamics has no global spacetime symmetry. This may be applied to Lisi's proposal of an E8 unified theory, giving a fully E8 invariant action. The extended form of the Plebanski action suggests a new class of spin foam models.

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### I. INTRODUCTION

Ashtekar's formulation of general relativity [1] taught us to think of gravitational theories as theories of connections, on a bare manifold with no metric structure. This insight is deepened by the Plebanski form of the action [2] and related forms studied by Capovilla, Dell and Jacobson and others [3,4] which reveal that general relativity is a simple perturbation or constraint of a topological field theory. This taught us, for example, that the metric is not fundamental; it emerges from solutions to the classical field equations. These insights into the structure of the classical theory led to much progress on the quantization of gravity, in loop quantum gravity and spin foam models [5].

The idea that general relativity has its deepest formulation as a connection theory suggested immediately a new approach to the unification of general relativity with Yang-Mills theories. The group of the connection in the connection formulations of general relativity is the local Lorentz group SO(3, 1) or a subgroup of it. What if one just takes a larger group *G* containing SO(3, 1) and plugs it into the same action or Hamiltonian constraint that gives general relativity so that one now has a theory with local *G* invariance? Does one get a theory that contains general relativity coupled to Yang-Mills fields in a subgroup of G/SO(3, 1)?

This question was answered affirmatively, in the Hamiltonian formulation, by Peldan [6]. All that was required was a simple modification of the Hamiltonian constraint, to keep the field equations consistent. Peldan's approach was further studied by Peldan and Chakraborty [7], and Gambini, Olson and Pullin [8].<sup>1</sup>

More recently, it has been understood by Krasnov [10] and Bengtsson [11] that these extensions of general relativity have elegant formulations in terms of action principles which extend the Plebanski action in a natural way.

In this paper we study a very simple theory of this kind. We find that the addition of one simple term to the Plebanski action suffices to get a consistent dynamics for any G containing the local Lorentz group. There is a simple mechanism which breaks the symmetry down in such a way that the resulting dynamics is Yang-Mills coupled to general relativity plus corrections. These corrections involve gauge fields which mix local Lorentz transformations with local internal gauge transformations.

One reason for revived interest in these kinds of unifications is Lisi's recent proposal [12] of a unified theory, based on the same strategy, where G = E8. There are several open issues regarding this proposal, which would have to be resolved for the idea to rise to the level of a theory. At least two of them have to do with making the dynamics fully E8 invariant. Lisi proposes a form of the action based on an approach to writing general relativity as a gauge theory, invented by MacDowell and Mansouri [13]. In that scheme the metric and local Lorentz connection are unified in a de Sitter or anti-de Sitter connection. but the action is only invariant under a local Lorentz subgroup. Lisi's action is similar, with E8 playing the role of the de Sitter group. The action studied here, applied to E8, gives a fully gauge invariant action, which has solutions which spontaneously break the symmetry and give, when expanded around, the bosonic part of Lisi's action, plus corrections.<sup>2</sup>

In the next section we describe the general extension of the Plebanski action and show how it reduces to Einstein-Yang-Mills plus corrections. We review the standard SO(4)Plebanski formulation [4,15] and describe the simplest extension in this case along the lines of [10]. We then show how a simple ansatz leads to a low energy theory which is a unification of general relativity and Yang-Mills theory.

Another way in which Lisi's proposal breaks the gauge invariance is by a strategy of incorporating fermions by means of a Becchi-Rouet-Stora-Tyutin extension of the

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<sup>&</sup>lt;sup>1</sup>Other, possibly related approaches to unification of the gauge groups of general relativity and Yang-Mills theory are explored in [9].

<sup>&</sup>lt;sup>2</sup>It was also recently shown that the action proposed here also naturally incorporates Higgs fields. This and other aspects of the unification will be discussed in [14].

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connection. In Sec. III A I propose an alternative way to incorporate the fermions, which would not break the gauge symmetry. It is based on proposals that matter degrees of freedom arise in loop quantum gravity as a result of the phenomena of disordered locality discovered by Markopoulou [16,17]. This is in fact a version of a proposal of Misner and Wheeler from 1957 that matter might be nothing but the mouths of Planck scale wormholes [18], and in the context of loop quantum gravity, this was argued previously to give fermions [19]. In this kind of mechanism, the full local symmetry remains present, but the cost is that for every generator of *G* there is both a gauge boson and a fermion degree of freedom.

Regarding the quantization, the class of actions proposed here suggest a particularly simple class of spin foam models. This is discussed in Sec. III B.

There are many aspects of the theories proposed here to be developed. In general, and in regard to *E*8, there are many open issues and many things to work out. For example in the latter there are open issues concerning how the structure of the generations is realized. Nonetheless, these approaches are worth exploring because it is very possible that they give rise to consistent, finite quantum theories. Many of the kinematical results of loop quantum gravity will go over to them, and as for dynamics, they involve simple modifications of by now well studied spin foam models.

We close the introduction by noting that the well-known Coleman-Mandula no-go theorem [20] is avoided because that only applies to an S-Matrix whose symmetries include global Poincare invariance. This theory, like general relativity, has no global symmetries, the Poincare symmetry acts only on the ground state not the action, and only in the limit in which the cosmological constant is zero. In fact, there is a nonzero cosmological constant, as it is related to parameters of the theory. By the time the S matrix in Minkowski spacetime could be defined in this theory one will be studying only small perturbations of a ground state in a certain limit and the symmetry will only apply in that limit and approximation. As we shall see below, the symmetry will already be broken by the time that approximation and limit are defined, in such a way that Coleman-Mandula theorem could be satisfied in its domain of applicability.

#### **II. PLEBANSKI ACTION FOR A GENERAL GROUP**

We will study here a general extension of the Plebanski action based on a connection valued in a semisimple G which contains the Euclidean Lorentz group SO(4). (For simplicity we consider only the case of Euclidean signature.)

We begin by reviewing the Plebanski action for general relativity, based on the group SO(4) [15], parametrized by antisymmetric pairs of four dimensional vector indices, [ab], where a = 1, ...4,

$$S^{\text{Pleb}} = \frac{1}{G} \int_{\mathcal{M}} B^{ab} \wedge F_{ab} - \frac{1}{2} \phi_{abcd} B^{ab} \wedge B^{cd}.$$
 (1)

The integral is over a four manifold,  $\mathcal{M}$ .  $F^{ab}$  is the curvature of the spacetime connection  $A^{ab}$  that gauges the local Lorentz symmetry,  $B^{ab}$  is a two-form valued in the Lie algebra of SO(4) and  $\phi_{abcd}$  is a multiplet of scalar fields that satisfy

$$\phi_{abcd} = \phi_{cdab} = -\phi_{bacd}.$$
 (2)

We do the variation subject to the constraint that

$$\boldsymbol{\epsilon}^{abcd}\boldsymbol{\phi}_{abcd} = \boldsymbol{\Lambda},\tag{3}$$

where  $\Lambda$  is the cosmological constant. For the derivation of the Einstein field equations from (1), see [4,15]; how this works will also be clear shortly.

We now extend this by embedding SO(4) in a larger Lie algebra, G, and adding one term to the action. The new, extended, Plebanski action is (generators of G are labeled by I, J = 1, ..., n.)

$$S^{G} = \frac{2}{G} \int_{\mathcal{M}} \left[ B^{J} \wedge F_{J} - \frac{1}{2} \phi_{JK} B^{J} \wedge B^{K} + \frac{g}{2} \phi^{KL} \phi_{KL} B^{J} \wedge B_{J} \right].$$
(4)

Now,  $A^I$  is a *G* connection,  $B^I$  is a two-form valued in the Lie algebra of *G* and  $\phi_{IJ} = \phi_{JI}$  is a multiplet of scalar fields. These are no longer constrained by (3). *A* has dimensions of inverse length, and *B* is dimensionless, so  $\phi$  has dimensions of (length)<sup>-2</sup>. *g* is a new coupling constant with dimensions of length<sup>2</sup>.

The field equations are

$$F_J = \phi_{JK} B^K - g \phi^{KL} \phi_{KL} B^J, \tag{5}$$

$$\mathcal{D} \wedge B^K = 0, \tag{6}$$

$$B^{(J} \wedge B^{K)} = 2g\phi^{JK}B^L \wedge B_L. \tag{7}$$

We can note that a theory is most natural when all the constraints that need to be imposed on its degrees of freedom follow from variation of the action. It is then satisfying that a constraint such as (3) no longer needs to be imposed.

Let us now consider the consequences of the last field equation (7) which resulted from varying the  $\phi^{IK}$ .

Tracing (7) we find, so long as  $B^L \wedge B_L \neq 0$ ,

$$\phi_K^K = \frac{1}{2g}.\tag{8}$$

We also see that generally<sup>3</sup>

<sup>&</sup>lt;sup>3</sup>The ratio of two 4-forms is a scalar which can be computed in any coordinate system.

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$$\phi^{JK} = \frac{1}{2g} \frac{B^{(J} \wedge B^{K)}}{B^L \wedge B_L}.$$
(9)

Putting (5) and (6) together we find

$$(\mathcal{D}\phi_{JK}) \wedge B^{K} = 2g\phi^{KL}\mathcal{D}\phi_{KL} \wedge B_{J}.$$
(10)

So far all equations are fully *G* invariant. I will now exhibit a natural symmetry breaking mechanism that distinguishes an SO(4) subalgebra, leading to a coupling of general relativity to Yang-Mills fields in the quotient G/SO(4).

To exhibit the symmetry breaking mechanism we first decompose *G* into its *SO*(4) subalgebra and remaining generators in *G*/*SO*(4). As before, we label the generators of *SO*(4) by antisymmetric pairs of four dimensional vector indices, [ab], where a = 1, ...4. The remaining generators in *G*/*SO*(4) we label by i = 7, 8...n. Then we have  $I = \{[ab], , i\}$ .

We look for a solution to the equations of motion to leading order in g which breaks the symmetry from Gdown to SO(4). To find it, we consider the component of (7) in the SO(4) directions

$$B^{ab} \wedge B^{cd} = 2g\phi^{abcd}B^L \wedge B_L. \tag{11}$$

We require a solution to this to leading order in g for which  $B^L \wedge B_L$  is nonvanishing. This can be solved by taking as an ansatz a form discussed in [21],

$$\phi^{abcd} = \frac{1}{2g} \epsilon^{abcd} W + \frac{\rho}{2g} \delta^{a[c} \delta^{d]b}$$
(12)

with the rest of the components of  $\phi^{IJ}$  of order zero in g and higher. Here W is a scalar dimensionless function of the fields. This breaks the symmetry, as it gives, to zeroth order in g,

$$\frac{B^{ab} \wedge B_{cd}}{B^L \wedge B_I} = \epsilon^{ab}_{\ \ cd} W + \rho \delta^{[a}_c \delta^{b]}_d. \tag{13}$$

Tracing (11) we find that

$$\phi^{ab}_{\ ab} = \frac{3\rho}{g} = \frac{1}{2g} \frac{B^{ab} \wedge B_{ab}}{B^L \wedge B_L}.$$
 (14)

It is consistent to assume that  $B^i$  are of higher order in g so that to leading order

$$B^L \wedge B_L = \frac{1}{2} (B^{ab} \wedge B_{ab}) + O(g), \tag{15}$$

which implies that

$$\rho = \frac{1}{3} + O(g) \tag{16}$$

We next solve the remaining components of (7) to order zero in g. To solve (11) by ansatz (12) we write, to leading order in g,

$$B^{ab} = \gamma \Sigma^{ab} + \Sigma^{*ab}, \tag{17}$$

where  $\Sigma^{ab} = e^a \wedge e^b$ . It follows that

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$$B^{ab} \wedge B_{ab} = 24\gamma e, \tag{18}$$

where *e* is the determinant of the frame field.  $\gamma$  is a parameter which labels the solutions (17), and must not vanish. We also exclude the case  $\gamma = \pm 1$ , which is the self-dual case.

We multiply (13) and (11) by  $\epsilon_{abcd}$ , which tells us that

$$W = \frac{1+\gamma^2}{12\gamma}.$$
 (19)

We now consider the equations satisfied by  $B^i$ . To zeroth order in g we have

$$B^{ab} \wedge B^i = 0, \tag{20}$$

which becomes for solutions

$$(\gamma \Sigma^{ab} + \Sigma^{*ab}) \wedge B^i = 0. \tag{21}$$

For each  $i \in G/SO(4)$  these are six linear equations in the six components of  $B^i$ , and they imply, so long as the metric is nondegenerate,

$$B^i = 0 \tag{22}$$

to zeroth order in g.

Thus we see that the ansatz (12) has broken the symmetry. The  $n B^{I}$  all start out on the same footing representing the full *G* gauge symmetry. But if n > 6 the equations (7) to zeroth order in *g* 

$$B^{(J} \wedge B^{K)} = 0 \tag{23}$$

cannot be solved to make more than 6 of the  $B^{I}$  nonvanishing. This is bound to happen at some *n* because there are 6ncomponents of  $B^{I}$  but n(n + 1)/2 equations in (23). But we see that as soon as there are six nonvanishing  $B^{I}$  valued in the Lorentz algebra, each additional  $B^{i}$  must satisfy the 6 linear equations (21). These are six linear equations for the six components of each  $B^{i}$ , so the result is that they all vanish to zeroth order in g. They will be nonvanishing at O(g). Hence this is a mechanism to break G down to the Lorentz algebra at a scale given by g.

To higher order in g we will have

$$B^{ab} = \gamma \Sigma^{ab} + \Sigma^{*ab} + gb^{ab} + \dots, \qquad (24)$$

$$B^i = gb^i. (25)$$

We will not here develop the solutions to higher order but it is straightforward to do so. What we have so far is enough to see that we get the fields for general relativity coupled in some nontrivial way to additional matter fields represented by the fields with indices wholly or partly in G/SO(4). There will also be corrections to the Einstein and Yang-Mills equations given by a power series in g. We note that the terms proportional to g, quadratic in  $\phi_{IK}$ , are necessary because otherwise the only solutions (7) would lead to most of the fields vanishing.

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To see how the coupled Einstein-Yang-Mills dynamics arises, we can go use the equation of motion for  $\phi^{IJ}$ , Eq. (7), and use it to eliminate the  $\phi^{IJ}$ . This is fair to do since we have so far just used the consequences of the equations of motion for the  $\phi_{IJ}$ . Since in our ansatz  $B^K \wedge$  $B_K$  is nonvanishing, we use (9) to eliminate  $\phi_{IJ}$ . The result is an action equivalent to (4) on solutions to (7) where  $B^K \wedge B_K$  is nonvanishing:

$$S^{\mathrm{no}\phi} = \frac{4}{G} \int B^{I} \wedge F_{I} - \frac{1}{8g} \frac{(B^{I} \wedge B^{J})(B_{I} \wedge B_{J})}{B^{K} \wedge B_{K}}.$$
 (26)

We can divide this into a gravity and matter part

$$S^{\rm no\phi} = S^{\rm no\phi}_{\rm grav} + S^{\rm no\phi}_{\rm matter}.$$
 (27)

The gravity part is

$$S_{\text{grav}}^{\text{no}\phi} = \frac{2}{G} \int B^{ab} \wedge F_{ab} - \frac{1}{8g} \frac{(B^{ab} \wedge B^{cd})(B_{ab} \wedge B_{cd})}{B^{ef} \wedge B_{ef}}$$
$$= \frac{2}{G} \int \left( \Sigma^{ab*} \wedge F_{ab} + \gamma \Sigma^{ab} \wedge F_{ab} - \frac{\Lambda}{2}e \right), \quad (28)$$

where

$$\Lambda = \frac{\gamma}{g} [1 + 36W^2]. \tag{29}$$

We see that we recover the Palatini action with the Holst term, where  $(2\gamma)^{-1}$  is the Immirzi parameter and  $G = 8\pi G_{\text{Newton}}$ , where  $G_{\text{Newton}}$  is Newton's constant.<sup>4</sup> So our ansatz leads to standard gravity with a parity breaking topological term.

To work out the matter part it is convenient to define

$$F^{i} = \frac{1}{2} F^{i}_{ab} \Sigma^{ab}, \qquad B^{i} = \frac{1}{2} B^{i}_{ab} \Sigma^{ab}, \qquad (30)$$

where  $F_{ab}^{i}$  and  $B_{ab}^{i}$  are functions. Then we can expand the matter part of (26) as

$$S_{\text{matter}}^{\text{no}\phi} = \frac{4}{G} \int e \left( 2B_{ab}^{i} F_{ab}^{i*} - \frac{1}{16g} \left( WB_{iab} B^{iab} + \frac{2}{3} B_{ab}^{i} B_{ab}^{i*} \right) \right).$$
(31)

We can solve the equations of motion for  $B_{ab}^i$  and then plug the solution into the action, and we find that

$$S_{\text{matter}}^{\text{no}\phi} = -\frac{1}{4} \int \left( \frac{e}{g_{\text{YM}}^2} F^{iab} F_{iab} + \theta F^i \wedge F_i \right), \quad (32)$$

where

$$\frac{1}{g_{\rm YM}^2} = \frac{256gW}{G(W^2 - \frac{4}{9})^2} \bigg[ W^2 + \frac{23}{54} \bigg],\tag{33}$$

and

$$\theta = \frac{4g}{9G(W^2 - \frac{4}{9})^2} \left[ W^2 - \frac{133}{3}W + \frac{4}{9} \right].$$
 (34)

By comparing with (29), we see that roughly  $g_{YM}^2 \approx G\Lambda$ so the Yang-Mills theory is weakly coupled when the cosmological constant is small. It is also interesting to note that our ansatz naturally leads to a parity breaking topological term in the Yang-Mills sector, as it does for the gravitational part of the action. However, the  $\theta$  angle is large.

It is also interesting to take into account the fact that, as shown in [22], the ansatz (12) is only consistent for special values of the Immirzi parameter, which are  $\gamma_0 = \pm \sqrt{\frac{2}{\sqrt{3}} - 1}$ . For these values,  $W^2 = \frac{1}{108}$ .

It is also interesting to note that we get the Yang-Mills form (32) for the dynamics of all the components of the connection,  $A^i$ ,with components in G/SO(4). These include Yang-Mills fields in the largest compact subgroup of the quotient G/SO(4), which we can call H. However, note that  $F^i$  are the components of the full G curvatures, which for  $i \in H$  are the Yang-Mills field strength plus corrections. Note also that there are generally connection and curvature components in  $G/(SO(4) \oplus H)$  which will not be Yang-Mills fields as they will transform nontrivially under both H and SO(4). The details of this are studied in [14] where it is shown that the degrees of freedom in these off diagonal connections contain the Higgs fields for H.

The scheme we have described works for any gauge group so it will work for G = E8. This gives a fully E8 covariant dynamics for Lisi's unification proposal [12].

#### **III. DISCUSSION**

#### A. Fermions

In Lisi's proposal the fermions of the standard model are described by a Becchi-Rouet-Stora-Tyutin extension of the connection. There is within loop quantum gravity another alternative, which is that fermions and scalars arise due to the possibility of disordered locality. As described in [16] this phenomenon arises because there are two notions of locality in loop quantum gravity. Microscopic locality is determined by the connections of the spin network graphs; two nodes are neighbors if they are connected. In the semiclassical limit a classical metric will emerge, generally from a superposition of spin network states. This gives rise to a notion of macrolocality. As described in [16,17] these two notions of locality disagree when two nodes xand y, which are far from each other in terms of the classical emergent metric, are connected by a single edge  $e_{xy}$ . Furthermore, in those papers it is argued that disordered locality is generically to be expected in the semiclassical limit. Consider a graph as in Fig. 1 which is regular and therefore may occur in the superposition of states making up a semiclassical state associated with a flat metric. There is in loop quantum gravity no apparent

<sup>&</sup>lt;sup>4</sup>In a related paper [22] it is shown that the ansatz made here is consistent only if  $\gamma$  is fixed to specific values.

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FIG. 1 (color online). Disordered locality: A regular graph contaminated by nonlocal links.

energy cost to contaminating that latticelike graph with nonlocal links as shown in the figure. Nor is there an incompatibility with the semiclassicality of the state. As there are many more ways to add a link to a lattice that connects two far away nodes than two nearby nodes, there is an instability for the formation of such nonlocal links as the universe expands from Planck scales. Moreover, once inserted in a graph, nonlocal links are trapped, as they can only be eliminated if two of them annihilate by the coincidence of their ends arriving by local moves at neighboring nodes. The proposal is then that these act as Planck scale wormholes, carrying quantum numbers associated with gauge fields carried by the nonlocal link.

Let us consider observations made by local observers in the neighborhood of x. From their point of view the edge  $e_{xy}$  simply comes to an end; that is, it appears to connect to a one valent node. But ends, or one valent nodes, in loop quantum gravity represent matter degrees of freedom. Thus, the dislocations due to disordered locality appear in the semiclassical limit as matter degrees of freedom.

Let us suppose that the gauge group is  $SU(2) \otimes H$ , where *H* is an internal gauge symmetry. Then the edge  $e_{xy}$  carries representations of these groups, (j, r). Local observers will describe  $e_{xy}$  as a particle of spin *j* and charge *r*.

This leads to a picture in which for every generator of G, the gauge symmetry, the semiclassical limit has a gauge field plus a set of particle excitations given by the representations of G.

This emergence of particle states is so far kinematical; more work needs to be done to ensure that in the low energy limit these particles have the correct dynamics and statistics. Earlier work, described in [19], indicates that in the Hamiltonian constraint formulation of loop quantum gravity, Planck scale wormholes do behave as spinors.

We then turn to some comments on the quantum dynamics.

### B. A class of spin foam models

I would like to comment that the form of the extended Plebanski action may have a simple quantization in terms of a spin foam model. Let us consider the action in the form (26). The Euclidean path integral naively will have the form,

$$Z = \int dAdB e^{(1/G)} \int B^{I} \wedge F_{I} - (\Lambda/8G)((B^{I} \wedge B^{J})(B_{I} \wedge B_{J})/B^{K} \wedge B_{K})}.$$
(35)

The  $B \wedge F$  term alone would give rise to a topological field theory with a spin foam formulation of the form of

$$Z = \sum_{r,i} \prod_{\text{4 simplices}} \{15 - j\}_G,$$
(36)

where the four dimensional manifold has been triangulated whose faces are labeled by representations r and the tetrahedra by the intertwiners i of the group G.  $\{15 - j\}_G$  are the 15j symbols which are functions of the labels on each four simplex. The usual Barrett-Crane strategy is to modify this by constraining the sums over representations and intertwiners to the balanced representations. Instead, the action (26) suggests using the simplicity constraints as Gaussian weights so we have

$$Z = \sum_{r,i} \prod_{\substack{\text{simplices}}} \{15\text{-}j\}_G e^{-(\Lambda/8G)((B^I \wedge B^J)(B_I \wedge B_J)/B^K \wedge B_K)},$$
(37)

where the  $B^{I}$  are functions of the representation labels on each face.<sup>5</sup>

## C. Open issues

I would like to close by listing a few out of many open issues facing this kind of unification.

- (i) The kinematical quantum theory can now be developed along loop quantum gravity lines for a general *G*, as well as for the particular case of *E*8.
- (ii) The spin foam quantization may also be explored based on the proposal discussed here. It will be interesting to see if the ultraviolet convergence results from the Barrett-Crane model also apply here.
- (iii) The proposal of matter as the ends of long-distance links needs more development. One needs to check whether the spin foam dynamics gives the right dynamics for the fermions in the case of graviweak unification or a larger unification. There are also open issues regarding spin and statistics; these may be addressed by generalized or topological spinstatistics theorems.
- (iv) Another application of the present formulation is to an SO(8) based unification of gravity and the electroweak gauge theory, along the lines discussed in [24].
- (v) We have shown that general relativity follows from a particular ansatz for solutions (12), but this does not show that there are not other gravitational degrees of freedom which could be excited. An analysis in [25]

 $<sup>{}^{5}</sup>$ A related form has been considered by Speziale for 2 + 1 gravity coupled to a Yang-Mills field [23].

indicates that there are; their physics is studied in [26]. Another interesting line of development would be to extend just the chiral part of the Plebanski action to a unification of gravity and Yang-Mills gauge symmetry.

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