Limitations on the topological BF scheme in Riemann-Cartan spacetime with torsion

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Cartan's structure equations in the Riemann-Cartan framework and some topological invariants of gravity are reanalyzed from the perspective of *BF* theories. This is related to a variational approach to Chern-Simons terms and Bianchi identities employing Lagrange multipliers. Here, it is pointed out that the *BF* scheme has some *limitations* to the effect that a coupling to matter would *leave the minimal coupling prescription* of gauge theories. In the case of gravity, the field equations would, generically, become *higher order* with a coupling to the *relocalized* Belinfante-Rosenfeld energy-momentum current.

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I. INTRODUCTION

A rather new development in topological field theory is the *BF* formalism, which provides potentially interesting relations to higher-dimensional knots, cf. Refs. [1,2]. *BF* theory is a framework where the connection one-form $A = A_i dx^i$ and an auxiliary¹ two-form $B = B_{ij} dx^i \wedge dx^j/2$ are varied independently. In the Abelian case, the one-form *A* can be interpreted as electromagnetic potential. In its primordial form, it starts from the metric-independent Lagrangian four-form

$$L_{BF} = -B \wedge F = -B \wedge dA. \tag{1.1}$$

Independent variations with respect to A and B lead to dB = 0 and the *constraint* of vanishing field strength F := dA = 0. This topological model has no local degrees of freedom.

This pure BF system can be modified [3] via a boundary term such that the Lagrangian

$$\tilde{L}_{BF} = -B \wedge dA + \frac{1}{2}B \wedge B \cong -B \wedge dA + \frac{1}{2}dC \quad (1.2)$$

becomes, "on shell," quadratic² in B. Now, independent variations provide the definition of the field strength together with the corresponding *Bianchi identity*

$$B \cong dA := F, \qquad dB = dF \equiv 0, \tag{1.3}$$

respectively, in compliancy with the Poincaré lemma $dd \equiv 0$. It still defines a topological theory since, "on shell,"

Eq. (1.2) is equivalent to (1.1) amended by a boundary term dC derived from a *Chern-Simons three-form* C. For an Abelian connection, this is simply given by $C = A \wedge F$. In general, it is well known [7] that Bianchi-type identities can be recovered via the variation of the associate Pontrjagin term $dC = F \wedge F$, e.g. $\delta dC/\delta A = 2dF \equiv 0$ in the Abelian case.

Bianchi *identities* do not allow for couplings to source terms thus seriously limiting applications of such a topological scheme. In order to proceed to more *realistic physical models* admitting matter couplings as in Maxwell's theory, the corresponding *BF* Lagrangian

$$L_{\max} = -B \wedge {}^*dA + \frac{1}{2}B \wedge {}^*B + L_{\max}$$
(1.4)

necessarily involves the *Hodge dual* * depending on the metric, cf. Ref. [8]. Then, independent variations of (1.4) provide again the definition of the field strength B = dA := F but, as a bonus, the nontrivial physical field equation $d^*B = d^*F \cong j$.

However, it should not be overlooked that in a coupling to matter such a BF scheme would leave the minimal coupling prescription since it generates a current three-form

$$j := \frac{\delta L_{\text{matter}}}{\delta A} = \frac{\partial L_{\text{matter}}}{\partial A} + d \frac{\partial L_{\text{matter}}}{\partial dA}$$

$$= \Psi \wedge \frac{\partial L_{\text{matter}}}{\partial D\Psi} + D \frac{\partial L_{\text{matter}}}{\partial B},$$
 (1.5)

which "on shell," is conserved classically, i.e. $dj \approx 0$. In general, this includes *Pauli-type terms* generated by the variation of the Lagrangian with respect to dA, cf. (5.2.18) of [9]. Because of Eq. (1.3), this additional term is equivalent to the one generated by the variation with respect to *B*, as indicated in (1.5).

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¹In four dimensions, *B* resembles the two-form potential for the gauge-invariant field strength or excitation H = dB, the Kalb-Ramond *axion* three-form.

²In three dimensions, non-Abelian *BF* systems with a cubic term $-B \wedge F + B \wedge B \wedge B/3$ are directly related to Chern-Simons theories departing from the three-form $C := A \wedge F - A \wedge A \wedge A/3$, cf. Refs. [4,5]. The resulting field equations $F = -B \wedge B$ together with the Bianchi identity $DB = DF \equiv 0$ correspond, in 3D gravity [6], to those with a cosmological term.

II. GRAVITY IN RIEMANN-CARTAN SPACETIME

In the case of topological gravity, the structure equations³ for the curvature $R_{\alpha\beta}$ in Riemann-Cartan (RC) spacetime and its Bianchi identity can be recovered similarly as in the *BF* scheme.

The Lagrangian corresponding to the *second Bianchi identity*

$$DR_{\alpha}{}^{\beta} \equiv 0 \tag{2.1}$$

is the boundary term

$$dC_{\rm RR} = -\frac{1}{2} R_{\alpha}{}^{\beta} \wedge R_{\beta}{}^{\alpha}. \tag{2.2}$$

As is well known, its integration yields a number proportional to the topological invariant of Pontrjagin. Because of the Poincaré lemma $dd \equiv 0$, there do not arise higherorder terms⁴ in this chain, cf. our scrupulous related method [7] using *Lagrange multipliers* as well as Ref. [12]. Such a topological theory can, similarly to (1.2), be given the generalized *BF* structure

$$B_{\alpha}{}^{\beta} \wedge R_{\beta}{}^{\alpha} - \frac{1}{2} B_{\alpha}{}^{\beta} \wedge B_{\beta}{}^{\alpha} = B_{\alpha}{}^{\beta} \wedge R_{\beta}{}^{\alpha} + dC_{\text{RR}}.$$
(2.3)

However, for torsion the situation is more subtle: The linear or Lorentz connection $\Gamma_{\alpha\beta} = -\Gamma_{\beta\alpha} = \Gamma^{\{\}}_{\alpha\beta} - K_{\alpha\beta} = \Gamma^{\{\}}_{\alpha\beta} + e_{\alpha}]T_{\beta} + (e_{\alpha}]e_{\beta}]T_{\gamma}) \wedge \vartheta^{\gamma}$ can be regarded as a "deformation" [4] of the unique Levi-Civita connection $\Gamma^{\{\}\alpha\beta}$ of Riemannian geometry via the *contortion* $K_{\alpha\beta} = -K_{\beta\alpha}$ implicitly related to torsion via $T^{\alpha} = K^{\alpha}{}_{\beta} \wedge \vartheta^{\beta}$. In order to account for the torsion content, one can consider a *change of variables* in the gravitational Lagrangian with optional matter couplings, i.e.

$$L(\vartheta^{\alpha}, \Gamma_{\beta}{}^{\gamma}, \Psi, D\Psi) \longrightarrow \hat{L}(\vartheta^{\alpha}, T^{\beta}, \Psi, D\Psi) + \mu_{\alpha} \wedge (B^{\alpha} - D\vartheta^{\alpha}) + \frac{\theta_{T}}{2\ell^{2}} B_{\alpha} \wedge B^{\alpha},$$
(2.4)

thereby leaving, as in the Maxwell case, the minimal coupling scheme. As is explained in more detail⁵ in Section 5.6 of Ref. [9], when torsion is regarded as an *independent B*-type two-form in the variational procedure, the first Cartan structure equation

$$B^{\alpha} = D\vartheta^{\alpha} := d\vartheta^{\alpha} + \Gamma_{\beta}{}^{\alpha} \wedge \vartheta^{\beta} = T^{\alpha}$$
(2.5)

needs to be enforced by a term involving the Lagrangian multiplier two-form μ_{α} . Then torsion couples via $\delta B^{\alpha} \wedge [\mu_{\alpha} + (\theta_T/\ell^2)B_{\alpha}]$ to the former Lagrange multiplier and, finally, the variation of the translational gauge field, the soldered coframe ϑ^{α} together with (2.5), induces a *relocalization* of the canonical energy-momentum current $\Sigma_{\alpha} := \delta L/\delta \vartheta^{\alpha}$ into

$$\sigma_{\alpha} = \Sigma_{\alpha} - D\mu_{\alpha} + e_{\beta} [(T^{\beta} \wedge \mu_{\alpha})]$$
$$\cong \Sigma_{\alpha} - D\mu_{\alpha} - \frac{\ell^{2}}{\theta_{T}} e_{\beta}](\mu^{\beta} \wedge \mu_{\alpha}), \qquad (2.6)$$

generalizing the familiar Belinfante-Rosenfeld symmetrization of general relativity (GR). Since $\tau_{\alpha\beta} = \vartheta_{[\alpha} \wedge \mu_{\beta]}$ is the spin current, Eq. (2.6) reveals the physical interpretation of μ_{α} as *spin energy potential* of matter. This relocalization [13,14] can be traced back to the *translational* nature of energy momentum and torsion ("translational" curvature) in the Lagrange-Noether machinery for the affine group and, for $\theta_T \neq 0$, a quadratic *contact-type* interaction to the spin energy potential emerges "on shell."

Since torsion involves the exterior derivative of the coframe ϑ^{α} , the above change of variables *cannot be regarded as a point-transformation* in the canonical formalism, and therefore its complete elimination would, in general, generate a change of the gravitational gauge energy-momentum current E_{α} according to

$$2D\left(e^{\beta}DH_{\alpha\beta} - \frac{1}{4}\vartheta_{\alpha}e^{\gamma}e^{\delta}DH_{\gamma\delta}\right) - E_{\alpha} = \Sigma_{\alpha} - D\mu_{\alpha},$$
(2.7)

where $H_{\alpha\beta} := -\partial L/\partial R^{\alpha\beta}$ are curvature excitations subjected to the constraint $T^{\alpha} = 0$. If the two higher derivative Cotton-type terms in the left-hand side of (2.7), cf. Eq. (5.8.25) of Ref. [9], are nonvanishing, the gravitational field equation becomes third order in the Levi-Civita connection $\Gamma^{\{\alpha\beta\}}$, i.e. fourth order in the holonomic metric with an induced coupling to the symmetric Belinfante-Rosenfeld energy-momentum current $\Sigma_{\alpha} - D^{\{\beta\}}\mu_{\alpha}$, cf. Ref. [14].

³Incidentally, the 3 + 1 decomposition of Cartan's structure equations and Bianchi identities in Ref. [10] are well known and can also be performed in a rather general slicing of spacetime, cf. Appendices B, C, and D of Ref. [11]. The tangential parts of both Bianchi identities are "constraints," which are preserved during time evolution, as proven in Eqs. (D.5) and (D.6) of Ref. [11]. The so-called "reducibility" equations (10) of Ref. [10] are contained in Eqs. (2.15) and (2.16) of Ref. [7] as tangential pieces. There, also the reducibility of the de Rham chain for curvature and torsion is explicitly derived, cf. the "Identities of the identities" in Sec. II of Ref. [7].

⁴This can also be seen by applying the Ricci formula to the first Bianchi identity (2.1), i.e. $DDR_{\alpha}{}^{\beta} = -R_{\mu}{}^{\beta} \wedge R_{\alpha}{}^{\mu} + R_{\alpha}{}^{\mu} \wedge R_{\mu}{}^{\beta} \equiv 0.$

⁵The nonmetricity $Q_{\alpha\beta} := -Dg_{\alpha\beta}$ is always put to zero here following the constraint formalism of Sec. 5.8.1 of Ref. [9], although a metric-affine framework would easily allow to liberate this constraint.

III. NIEH-YAN TOPOLOGICAL TERM

As an example of a partial BF structure in RC spacetime with torsion, let us consider the so-called Nieh-Yan (NY) term [15,16]

$$dC_{\rm TT} = \frac{1}{2\ell^2} (T^{\alpha} \wedge T_{\alpha} + R_{\alpha\beta} \wedge \vartheta^{\alpha} \wedge \vartheta^{\beta}), \qquad (3.1)$$

which violates parity [11,17]. It simply can be obtained by multiplying the *first Bianchi identity*

$$DT^{\alpha} \equiv R_{\beta}^{\ \alpha} \wedge \vartheta^{\beta} \tag{3.2}$$

of RC spacetime with ϑ_{α} from the left. *Vice versa*, this identity can be recovered via the variation of (3.1) with respect to the coframe, as carefully demonstrated in Ref. [7]. Since (3.2) constitute a purely geometric identity, it does *not* admit couplings to extended matter. Even for models with "distributional" matter, like cosmic strings [12], nonvanishing distributional torsion and curvature compensate each other for (3.2) to hold exactly [18]. The four-form on the right-hand side of the NY identity (3.1) even vanishes, as explicitly demonstrated in Eqs. (2.21) and (2.22) of Ref. [19].

Again, in order to obtain more realistic models, one of the field strengths in (3.1) needs to be converted via a duality rotation into its Hodge dual. Then the NY term suggest two options [20] for a viable gravitational Lagrangian: Hilbert's original choice

$$L_{\rm HE} = -\frac{1}{2\ell^2} R^{\{\}}_{\alpha\beta} \wedge^* (\vartheta^{\alpha} \wedge \vartheta^{\beta}) = -\frac{1}{2\ell^2} R^{\{\}} \eta, \quad (3.3)$$

where $R_{\alpha\beta}^{\{\}}$ denotes the Riemannian curvature for vanishing torsion and $R^{\{\}} := {}^*(R^{\{\}\alpha\beta} \wedge \eta_{\beta\alpha})$ the Riemannian curvature scalar, as in GR. Formally, it can be put into the *BF* scheme [12] when choosing $B_{\alpha\beta} = \eta_{\alpha\beta} = {}^*(\vartheta^{\alpha} \wedge \vartheta^{\beta})$ albeit the no-avoidance of the Hodge dual⁶. The purely torsion-square Lagrangian

$$L_{\parallel} := \frac{1}{2\ell^2} T^{\alpha} \wedge \left(-{}^{(1)}T_{\alpha} + 2{}^{(2)}T_{\alpha} + \frac{1}{2}{}^{(3)}T_{\alpha} \right) \quad (3.4)$$

where the torsion excitation $H_{\alpha}^{\parallel} := -\partial L_{\parallel}/\partial T^{\alpha} = (1/\ell^2)\eta_{\alpha\beta\gamma}K^{\beta\gamma}$ is dual to the contortion one-form $K_{\alpha\beta}$. It leads to proper *teleparallelism* (GR_{||}) when constrained by vanishing RC curvature, i.e. $R_{\alpha\beta} = 0$ via a Lagrangian multiplier term $\lambda_{\alpha\beta} \wedge R^{\alpha\beta}$, as was suggested already by Einstein. Because of the geometric identity

$$L_{\parallel} \equiv L_{\rm HE} + \frac{1}{2\ell^2} R_{\alpha\beta} \wedge^* (\vartheta^{\alpha} \wedge \vartheta^{\beta}) + \frac{1}{2\ell^2} d(\vartheta^{\alpha} \wedge^* T_{\alpha}),$$
(3.5)

 GR_{\parallel} with the teleparallel constraint $R_{\alpha\beta} = 0$ is classically *equivalent* to GR up to a boundary term $d(\vartheta^{\alpha} \wedge {}^*T_{\alpha})$ constructed from the Hodge dual ${}^*T_{\alpha}$ of the torsion, cf. Ref. [11].

Both pieces on the right-hand side of the dimensionless boundary type four-form (3.2) have found tentative applications before: The pseudoscalar curvature $R_{\alpha\beta} \wedge \vartheta^{\alpha} \wedge$ $\vartheta^{\beta} = \eta^{\alpha\beta\gamma\delta}R_{\alpha\beta\gamma\delta}\eta$ has been used for generating a *selfdual* or chiral reformulation of GR with its Hilbert-Einstein Lagrangian⁷ proportional to curvature scalar $R_{\alpha\beta} \wedge$ $^*(\vartheta^{\alpha} \wedge \vartheta^{\beta})$. This was anticipated already by Plebanski [22], Hojman *et al.* [23] and Dolan [24], but only later "rediscovered" by Holst [25] without references to earlier work. On the other hand, the torsion-squared term $T^{\alpha} \wedge T_{\alpha}$ in Eq. (3.1) has been employed to induce a *chiral* reformulation of the *teleparallelism equivalent* of GR, cf. Ref. [11] where the limiting case of vanishing RC curvature is consistently enforced⁸ again via Lagrange multipliers.

In the supersymmetric extension of EC theory, the NY boundary term (3.1) induces a chiral formulation [27] of simple ($\mathcal{N} = 1$) supergravity, whereas the *translational* Chern-Simons three-form $C_{\text{TT}} := \vartheta^{\alpha} \wedge T_{\alpha}/2\ell^2$ is instrumental for first order models of topological gravity [6] in 3D.

Interesting enough, the generalized BF scheme (1.2) can also be employed to induce a "breaking" [28] of the de Sitter gauge symmetry down to Einstein's GR with cosmological constant. This and its relation to BRST quantization of gravity [29,30] needs to be seen.

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⁶In fact, there have been attempts, to restrict oneself to the Lie dual ^(*) only, as is the case of the Lie dual $R_{\alpha\beta}^{(*)} := 1/2\eta_{\alpha\beta\gamma\delta}R^{\gamma\delta}$ of the curvature featuring, e.g. in the Euler invariant. However, it is not always realized, cf. for example Eq. (33) of Ref. [10], that the Lie dual $\eta_{\alpha\beta} := \frac{1}{2}\eta_{\alpha\beta\gamma\delta}\vartheta^{\gamma} \wedge \vartheta^{\delta} = *(\vartheta^{\alpha} \wedge \vartheta^{\beta})$ of the "unit" two-form $\vartheta^{\alpha} \wedge \vartheta^{\beta}$ is *equivalent* to its Hodge dual * as a consequence of the soldering of the coframe ϑ^{α} , cf. Eq. (3.7.8) of Ref. [9].

⁷The currently widespread usage of the label "Palatini action" or "method" has been questioned in Ref. [21] from the historical point of view and, therefore, appears to be a misnomer.

⁸For instance, it is claimed that the pseudoscalar action S_5 and the torsion-squared action S_6 of Ref. [10] "...define the same dynamical system." However, according to (3.1) they *differ* by the boundary term $2\ell^2 dC_{\text{TT}}$, which invariantly characterizes nontrivial topologies [26]. In fact, integration over the boundary three-sphere at infinity yields the invariant $n_{\text{NY}} := \int_{R^4 \cup \infty} dC_{\text{TT}} = \int_{S_5^3} C_{\text{TT}} = 6\pi^2 k$, where k is the winding or instanton number of Pontrjagin. The Lagrangian S_7 is of the *BF* type (1.2), but merely leads to the *truncated* Bianchi identity $DT^{\alpha} = {}^* dd \vartheta^{\alpha} \equiv 0$ of a teleparallel spacetime, where $\Gamma^{\alpha\beta} = {}^* 0$ locally.

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