$$\begin{split} A_6 &= s_1^{\alpha_1} \left(\frac{s_{12}}{s_1} \right)^{\alpha_2} \left(\frac{s_{31}}{s_1} \right)^{\alpha_3} \\ &\times \frac{1}{(s_1 s_2 / s_{12}) (s_3 s_1 / s_{31}) + 1 / s_1^{\alpha_v (0) - \alpha_1}} \end{split}$$

¹⁴These limits have been discussed and studied in the dual resonance model by C. E. DeTar and J. H. Weis, Phys. Rev. D <u>4</u>, 3141 (1971).

¹⁵Since the behavior of the amplitude is less specified, it is generally more difficult to rigorously exclude "nonsense" helicity singularities in limits like (2.4) with few variables asymptotic when they do not also occur in the Regge limits. These would be peculiar helicity singularities whose location differed from that of a given angular momentum singularity by an integer, but do not occur in the part of the amplitude with the angular momentum singularity.

 $^{16} The \ criterion \ \lambda \leq 0$ for an helicity limit being inside

the physical region has been emphasized by Abarbanel and Schwimmer (Ref. 1). For a nice presentation of kinematics in standard notation, see M. N. Misheloff, Phys. Rev. 184, 1732 (1969).

¹⁷Note that $\lambda(\overline{t_1, t_2}, t_3) \leq 0$, only if all t_i have the same sign. Then, given, say, t_2 and t_3 , $\lambda \leq 0$ for

 $|\sqrt{t_2} - \sqrt{t_3}| \le |t_1| \le |\sqrt{t_2} + \sqrt{t_3}|$.

¹⁸From Eq. (2.12) one sees that the triple-Regge limit $(s_2, s_3, s_{23} \rightarrow \infty)$ is obtained for fixed φ_2 and φ_3 , whereas, the Regge-helicity limit $(s_2, s_3, s_{23} \text{ fixed})$ is obtained for φ_2 and φ_3 varying with s_{12}/s_1 and s_{31}/s_1 . Thus, depending on the relative orientation of the states on . either side of $\sqrt{s_1}$, either one or the other limit is obtained inside the physical region.

¹⁹A. Patrascioiu, MIT report (unpublished).

 ^{20}C cannot be a constant since the coefficient of s_{12}/s_1 in (2.13) can be made arbitrarily large for $\lambda \le 0$ by scaling all the t_i to large values.

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Chiral Symmetry Limit in the SU(3) σ Model with Bilinear Breaking

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We add to the usual σ model a $(3^*, 3) + (3, 3^*)$ term which is bilinear in the scalar-meson fields and study the behavior of the tree-approximation solutions as the explicit symmetry breaking is turned off. We find that both $c \simeq -\sqrt{2}$ and $c \simeq 0$ occur for the same mass spectrum, 0-8 mixing, and F_K/F_{π} . Solutions with different c values approach the symmetry limit differently, but in this limit one can smoothly reach either a Goldstone pseudoscalar octet, or a symmetric vacuum, or neither for both values of c. No clear indication is found that solutions are near a critical value of the mass parameter.

I. INTRODUCTION

The assumption that expansions in symmetrybreaking parameters exist provides a basis for study of approximate symmetries of the stronginteraction Hamiltonian. For systems whose underlying symmetry is chiral $SU(3) \times SU(3)$,¹ the question of the nature of the symmetry limit and the existence of expansions in powers of symmetry-breaking parameters about such a limit is complicated by the Goldstone phenomenon² where the solutions of the symmetric theory do not exhibit the full symmetry of the Hamiltonian. Dashen³ and Dashen and Weinstein⁴ developed chiralsymmetry-breaking expansions and systematically exploited the technique in deriving a number of correlations among symmetry-breaking effects. Li and Pagels⁵ subsequently showed that Goldstone-boson intermediate states can give rise to singularities in the symmetry limit, thus invalidating many attempts to extrapolate "soft pion" theorems to the pion mass shell. In a study of Lagrangian models solved in the tree-graph approximation, Carruthers and Haymaker⁶ noticed a different phenomenon which has the same negative implication for expansions about a Goldstone symmetry solution. They found that vacuum expectation values in the σ model⁷ are multivalued func-

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tions of the Lagrangian symmetry-breaking parameters and that the power-series expansion of their solution had a radius of convergence much smaller than the value of the $SU(3) \times SU(3)$ -breaking parameter needed to produce correct pseudo-scalar masses.

Subsequent work on Lagrangian models has clarified the nature of the singularities.^{8,9,10} Carruthers and Haymaker¹⁰ have suggested pursuing the idea that the *n*-point functions are analytic functions of the breaking parameters in order to extract model-independent results. The effects of single-loop diagrams on the tree-approximation parameters have also been studied and found to be small.¹¹ In this connection, single-loop diagrams in the SU(2)×SU(2) σ model have been shown to be analytic in the symmetry limit,¹² indicating that only higher-order loop corrections contribute to the nonanalytic Li-Pagels⁵ behavior near the Goldstone limit. The singularities of the tree approximation itself may still be a problem, of course.

In this work, we extend the study of one- and two-point functions as multivalued functions of the chiral-symmetry-breaking parameters. We stay within the tree approximation and push the σ model further to include breaking terms in the Lagrangian which are bilinear in the basis fields [though still transforming as P = +, I = Y = 0 members of a $(3, 3^*) + (3^*, 3)$ representation^{13,14}]. Since the model with bilinear breaking terms is renormalizable,¹¹ loop corrections could be calculated in principle. However, the tree-approximation solutions themselves are varied and complex enough to discourage any idea of systematically evaluating loop effects. We look closely at the numerous tree-approximation solutions of this more complicated model to see what features, if any, are common to all solutions of both models.

In the following section, we review and extend the work on the σ model with linear $(3^*, 3) + (3, 3^*)$ breaking. In Sec. III we introduce the bilinear term and analyze the way the solutions to this model approach the SU $(3) \times$ SU(3)-symmetry limit. In Sec. IV, we summarize and discuss the implications of our results. The Appendix covers the procedure for obtaining solutions and includes necessary definitions and formulas.

II. LINEAR SYMMETRY BREAKING

In this section we shall review the tree-approximation solutions to the linear σ model with linear $(3, 3^*) + (3^*, 3)$ breaking.⁷ The Lagrangian density is

$$\mathcal{L} = -\frac{1}{2} (\partial_{\mu} u_i \partial_{\mu} u_i + \partial_{\mu} v_i \partial_{\mu} v_i) - \alpha I_2 - \alpha' (I_2)^2 - \beta I_3^* - \gamma I_4 + \delta a (u_0 + c u_8), \qquad (1)$$

where u_i and v_i are, respectively, scalar and pseudoscalar fields which belong to a $(3^*, 3)$ $+(3, 3^*)$ representation of SU(3)×SU(3). The invariants I_2 , I_3^+ , and I_4 are defined in the Appendix. Predictions of this model⁷ and of its generalizations¹⁵ have been extensively studied. In addition, the analytic structure of the solutions as functions of δ and c has been examined, and those symmetry limits have been determined which can be reached smoothly from solutions fitted to physical parameters, such as pseudoscalar masses and one of the I = Y = 0 members of the scalar nonet.¹⁶ The indications are^{6,10} that the linear model can connect smoothly only to the "normal" symmetry limit, $\delta = 0$, $\langle u_0 \rangle = 0$, $\langle u_8 \rangle = 0$, and all masses nonzero degenerate, or to the Goldstone pseudoscalar octet symmetry limit $\langle u_0 \rangle \neq 0$, $\delta = 0$, $\langle u_8 \rangle = 0$, pseudoscalar octet masses = 0, $m_{n'} \neq 0$, and the scalar nonet masses degenerate. We review this situation now and comment on the observation¹⁰ that one solution is near a "critical point" in the parameter α .

The most reliable experimental meson information is, of course, the pseudoscalar masses. The pseudoscalar mass-squared matrix $m_{ij}^2 = \partial^2 \mathcal{L} / \partial v_i \partial v_j$, after diagonalization in the 0-8 sector, provides four equations (see Appendix) for the four quantities

$$b = \frac{\langle 0 | u_{8} | 0 \rangle}{\sqrt{2} \langle 0 | u_{0} | 0 \rangle},$$

$$x_{1} = 2\alpha + 4\alpha' \lambda_{p}^{2} (1 + 2b^{2}),$$

$$x_{2} = 4\sqrt{6} \beta \lambda_{p},$$

$$x_{3} = 8\gamma \lambda_{p}^{2}, \quad \lambda_{p} \equiv \langle u_{0} \rangle_{\text{physical}}.$$
(2)

In particular, m_{π}^{2} , m_{K}^{2} , and $m_{\eta}^{2} + m_{\eta}^{2}$ enable one to express the latter three parameters, x_{1} , x_{2} , and x_{3} , in terms of b_{p} . The quantity $m_{\eta}^{2} - m_{\eta}^{2}$ then gives essentially¹⁷ a quadratic equation for b. We label the solutions to this equation $b(\text{physical}) \equiv b_{p}$. A set of values of x_{1} , x_{2} , and x_{3} are then found for each b_{p} , but the value of $\lambda_{p} = \langle u_{0} \rangle_{\text{physical}}$ for the physical masses cannot be determined separately from the coefficients of the invariants without additional input, such as F_{π} , to set the scale of the vacuum-breaking effects.¹⁸ The necessary conditions for stability, $\partial \mathcal{L} / \partial u_{0} = 0$ and $\partial \mathcal{L} / \partial u_{8} = 0$,¹⁹ are

$$\frac{\delta a}{\lambda_{p}} = x_{1} + x_{2}(1 - b^{2}) + \frac{2}{3}x_{3}(1 + 6b^{2} - 2b^{3}), \qquad (3)$$

$$\frac{\delta ac}{\lambda_p} = \sqrt{2} b \left[x_1 - x_2 (1+b) + 2x_3 (1-b+b^2) \right], \qquad (4)$$

respectively. Equations (3) and (4) yield values for $\delta a/\lambda_p$ and c for *each* b_p value. For the scalar mass matrix $\partial^2 \mathcal{L}/\partial u_i \partial u_j = M^2_{ij}$, the values of $M_{\pi_N}^2$

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bp	c _p	M_{π_N}	M_{κ}	$\tan \theta_P$	$M_{\epsilon}^{2} + M_{\epsilon}'^{2}$	M _e ,	M_{ϵ}	$\tan \theta_{s}$
-0.218	-1.30	0.906	0.888	0.052	1.70 1.40	$1.140 \\ 1.050$	0.630	0.955 0.682
-0.352	-1.33	0.902	0.725	0.228	1.70 1.40	1.225 1.095	$\begin{array}{c} 0.456 \\ 0.440 \end{array}$	$\begin{array}{c} 1.290\\ 1.146\end{array}$

TABLE I. Solutions to linear-breaking model. Input masses are $m_{\pi}^2 = 0.019$, $m_K^2 = 0.246$, $m_{\pi}^2 = 3.01$, and $m_{\eta'}^2 = 0.917$. All masses are in GeV.

^a These are input values. Both c_{p} values go with each $M_{\epsilon}^{2} + M_{\epsilon}^{2}$ value.

 $=\partial^{2} \mathcal{L}/\partial u_{1,2,3}^{2}$ and $M_{\kappa}^{2} = \partial^{2} \mathcal{L}/\partial u_{4,5,6,7}^{2}$ are predicted for each b_p solution. The isoscalar- and scalarnonet masses squared, $M_{\epsilon}{}^2$ and $M_{\epsilon'}{}^2$, are determined, and the value of one or of a combination of the two assumed in order to separate α and α' . The latter information is necessary to study the behavior of the solutions as functions of δ and/or c. For each of several reasonable values of M_{ϵ}^{2} $+M_{\epsilon'}^2$, the two solutions of the model which correspond to the pseudoscalar mass values (in units of GeV² always unless specified otherwise) m_{π}^{2} =0.019, m_{κ}^2 = 0.246, m_n^2 = 0.301, and m_n^2 = 0.917 are shown in Table I. The solutions differ in some details, but the features that $c \approx -\sqrt{2}$ (near the $SU(2) \times SU(2)$ -invariant Hamiltonian) and that b_{ϕ} is small and negative (small vacuum breaking relative to chiral breaking) are common to both. We are primarily interested in following the solutions as functions of δ in the region around δ (physical) $\equiv \delta_{p}$ down to $\delta = 0$, and we shall keep $c = c_{physical}$ for this study.⁸ We ask whether or not the physical solution connects smoothly to a symmetric limit as $\delta \rightarrow 0$, and, if so, what the character of the symmetric solution is. A reasonable criterion of "smoothness" was presented by Carruthers and Haymaker,⁶ who require that all masses squared remain positive (a no-ghost or stability criterion)



FIG. 1. Linear-breaking case $b_{\rho} = -0.218$, $M_{\epsilon}^2 + M_{\epsilon} r^2 = 1.7 \text{ GeV}^2$ for $\hat{\lambda}$ vs δ region including physical point, symmetry point, and singularity nearest symmetry point. $\hat{\lambda}$ is dimensionless and δ is actually $\delta a / \lambda_{\rho}$ in units of GeV², so $\delta_{\rho} \approx 10m_{\pi}^2$.

and that no singularities of $\lambda(\delta)$, $b(\delta)$, or masses are encountered in the path $\delta \rightarrow 0$. We note that the mass positivity requirement usually implies that no singularities occur, since singularities are generally accompanied by a zero of M_{ϵ} or $M_{\epsilon'}$.⁹

The actual procedure we use is dictated by calculational convenience. We study λ , δ , and masses squared as functions of $b = b(\delta) \equiv \langle u_{\beta} \rangle / \sqrt{2} \langle u_{\beta} \rangle$ (with x_1 , α , $\alpha' \lambda_p^2$, x_2 , x_3 , and c fixed) in the neighborhood of $b_p = b(\delta_p)$. The results, however, are plotted as $\lambda = \langle u_0 \rangle$ vs δ , since the behavior is more clearly shown in this way. In Figs. 1 and 2 we show the dependence of the vacuum expectation values $\lambda \equiv \langle u_0 \rangle$ on the Hamiltonian symmetrybreaking parameter, δ , for the solutions b_{b} = -0.218 and $b_p = -0.352$, respectively. The scalar-mass values used are such that $M_e^2 + M_e^2$ = 1.7 GeV^2 , a convenient input combination for computational purposes. Both solutions have the same character, that is, they connect smoothly to $\langle u_0 \rangle \neq 0, \langle u_8 \rangle = 0$ Goldstone pseudoscalar-octetsymmetry limits as $\delta \rightarrow 0$. The one $b_{p} = -0.218$ yields $M_e = 630$ MeV and $M_{e'} = 1140$ MeV; the b_p = -0.352 yields M_{ϵ} = 456 MeV and $M_{\epsilon'}$ = 1204 MeV. However, both solutions also exhibit singularities at negative δ values for which $|\delta|_{\text{singularity}} < \delta(\text{physi-}$ cal), thus prohibiting a power-series expansion about $\delta = 0$ from reaching δ (physical). The singularity occurs when $M_{e'} = 0.^9$ These solutions have identical features to the solution studied by Carruthers and Haymaker⁶ for the model with $\alpha = 0$.



We have chosen to examine the details of both solutions because the only argument against the b_p = -0.352 solution at this stage is its somewhat low- κ -mass prediction of 725 MeV and, with the input combination $M_{\epsilon^2} + M_{\epsilon'}{}^2 = 1.7 \text{ GeV}^2$, the low value of M_{ϵ} . It is, however, not clear whether or not one should try to compare to better than 100– 200 MeV with the illusive scalar-mass spectrum. An additional reason for treating this solution is the sensitivity of the b_p values to the choice of average kaon mass, as pointed out by Chan and Haymaker.¹¹

Let us now choose several different $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ values to study the sensitivity of the conclusions about the nature of the symmetry limit to the values of the scalar masses. The value $M_{e^2} + M_{e^{\prime}}^2$ = 2.0 GeV² yields the *same* essential features as shown in Figs. 1 and 2, and no conclusions are altered. However, the choice $M_e^2 + M_e^2 = 1.4$ GeV² leads to new behavior for the $b_p = -0.218$ solution for which $M_{\epsilon} = 552$ MeV and $M_{\epsilon'} = 1046$ MeV, but the same behavior as before for the solution $b_p = -0.352$, for which $M_e = 440$ MeV and $M_{e'} = 1095$ MeV. The curves are shown in Figs. 3 and 4 for $b_p = -0.218$ and $b_p = -0.352$, respectively. The former solution ($b_p = -0.218$, Fig. 3) has changed character, now connecting to a normal $\langle \langle u_0 \rangle = 0$, $\langle u_{\beta} \rangle = 0$ as $\delta \rightarrow 0$) solution with degenerate pseudoscalar and scalar nonets. A continuation from $\boldsymbol{\delta}$ =0 to $\delta = \delta_p$ is blocked by a singularity, at which $M_{\epsilon'} = 0$, ⁹ at $\delta \approx 0.007$. The details of this region are



FIG. 3. Case $b_p = -0.218$, $M_e^2 + M_{e'}^2 = 1.4 \text{ GeV}^2$.

shown in the inset in Fig. 3. The effect exhibited is essentially that pointed out by Olshansky.⁸ There is a slight difference here, since we reevaluate all of the coefficients for the new values M_{ϵ} and $M_{\epsilon'}$ rather than view the effect as being dependent on α alone. This phenomenon has been likened to a "critical point" in the solution to the theory.¹⁰ The suggestion was made that modelindependent results might be obtained by hypothesizing that one is near a critical point in the meson mass parameter α and then applying Wilson's renormalization-group techniques.²⁰

However, we wish to point out that the other solution (see Fig. 4) for which $b_p = -0.352$ and c= -1.32, does *not* exhibit this critical-point effect, and retains the Goldstone-limit character shown in Figs. 2 and 4, for $M_e^2 + M_{e'}^2$ values in the range $1.2 \le M_e^2 + M_{e'}^2 \le 2.0$ GeV², with no accompanying transition from Goldstone to normal symmetry limit. The transition will occur for this solution, just as for the solution $b_p = -0.218$, c= -1.30, but it occurs for a *very* low value of the scalar mass combination $M_e^2 + M_{e'}^2$. Only the b_p = -0.218 solution, therefore, provides support for the notion that the meson system is *close* to a transition from a Goldstone solution to a normal



FIG. 4. Case $b_{p} = -0.352$, $M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 1.4 \text{ GeV}^{2}$.

solution, the latter separated by singularities from the physical values of masses and couplings. We shall return to this point in the next section, where we discuss the effects which arise when bilinear chiral-symmetry-breaking terms are added to the Lagrangian, Eq. (1).

III. BILINEAR (3, 3*)+(3*, 3)-BREAKING EFFECTS

A Lagrangian with exactly the same $SU(3) \times SU(3)$ structure as the original σ model, Eq. (1), is obtained by adding a term $\delta d(U_0 + cU_8)$, bilinear in the fields u_i and v_i , which is parallel in SU(3)space to the bilinear breaking term of Eq. (1). The scalar quantities U_0 and U_8 are members of the (3, 3*) + (3*, 3) component in the reduction of the product of the basic multiplet with itself. Explicitly we have

$$U_{0} = 4(\frac{2}{3})^{1/2}(u_{0}^{2} - v_{0}^{2}) - 2(\frac{2}{3})^{1/2} \sum_{i=1}^{3} (u_{i}^{2} - v_{i}^{2}),$$

$$U_{8} = -4(\frac{2}{3})^{1/2}(u_{0}u_{8} - v_{0}v_{8}) + 4\sum_{i=1}^{8} d_{8ii}(u_{i}^{2} - v_{i}^{2}).$$
(5)

The mass formulas, extremum equations, and divergence equations are all modified in the presence of the new terms. We relegate to the Appendix the details of the equations and their solutions. The essential point is that we have one new parameter, d, and must therefore provide one more piece of information. We choose to work with the mass of the scalar, I=1 particle π_N , which may be tentatively identified with the I=1, $J^{P}=(0^{+}?)$ effect which has been observed at about 980 MeV. The standard σ model predicts this mass to be about 900 MeV, somewhat lower than one would like. While our main purpose is not to "fix up" the scalar mass spectrum, it is of interest to determine the behavior of the tree-approximation solutions when the π_N mass is lifted to a possibly more realistic value and to compare the behavior with the results of the linear-breaking case. Qualitatively, the most striking difference between the solutions of the modified model and those of the original one is the added complexity of the former. We find that there are eight solutions for each set of mass m_{π^2} , m_{K^2} , \underline{m}_{η^2} , $m_{\eta'}$, $m_{\eta'}$, and M_{π_N} . The quantity b_p $(b = \langle u_8 \rangle / \sqrt{2} \langle u_0 \rangle)$, the ratio of SU(3) octet to singlet vacuum breaking, satisfies a quartic equation. For each of the four b_p values which solves the quartic equation that determines the b_p 's, there are now two corresponding values of c_p , the ratio of SU(3) octet to singlet breaking in the Lagrangian.

Before entering into the detailed results of our solutions to the model with the bilinear breaking

terms added, we note several new possibilities that this expanded model allows which are evident from the general tree-approximation equations themselves. First of all, as is clear from the bilinear terms in the axial-vector divergence equations

$$\partial_{\mu}A_{\mu}^{i} = \delta(\frac{2}{3})^{1/2} \left[a + 4(\frac{2}{3})^{1/2} d \langle u_{0} \rangle \right] v_{i} + 4(\frac{2}{3})^{1/2} \delta d(\tilde{u}_{0}v_{i} + v_{0}\tilde{u}_{i} + \sum_{j, k=1}^{8} d_{ijk} \tilde{u}_{j} v_{k}) + \text{octet breaking}, \qquad (6)$$

where $u_0 = \langle u_0 \rangle + \tilde{u}_0$, $u_8 = \langle u_8 \rangle + \tilde{u}_8$, one can discuss broken-chiral-symmetry relations from the point of view of Dashen's power-counting arguments³ even within the tree approximation. Even in σ models, as mentioned by Carruthers and Haymaker,⁶ this is the essential question in regard to the validity of low-energy theorems. More technical points are (i) that the $\langle u_8 \rangle = 0$, $c \neq 0$ solution, impossible in the linear-breaking case, is allowed in the bilinear-breaking version; and (ii) that $\delta d U_0$ contributes to splitting the π^0 and η' masses and, in contrast with the linear-breaking case, the I_3^+ is no longer essential to keeping π and η' nondegenerate in the SU(3) limit, $c = 0.^{21}$

To begin the discussion of the solutions of the model with the added term, we choose the input value $M_{\pi N}^2 = 0.8229 \text{ GeV}^2$ (equal to the "preferred" one⁶ predicted by the model when only linear-breaking terms are present; see Sec. II). In Table II we list all of the eight solutions found for the mass inputs $m_{\pi}^2 = 0.019$, $m_K^2 = 0.246$, $m_{\eta}^2 = 0.301$, $m_{\eta'}^2 = 0.917$, and $M_{\pi N}^2 = 0.8229$. Of the eight solutions, one is identical to the solution of the linear-breaking model with the same mass spectrum, having $b_p = -0.218$, c = -1.30, and d = 0. Another solution is close to the other linear-model solution which had $M_{\pi N}^2 = 0.816$, $b_p = -0.354$, and c = -1.32. As we let $M_{\pi N}^2$ decrease to this latter value, 0.816 GeV², we recapture this other linear-breaking solution, with d = 0, as one of the eight solutions to the modified model.

The remaining six solutions, however, are completely different from either of the solutions discussed in Sec. II. The cases with b = -20.65 are clearly outside of currently popular ideas about chiral-symmetry breaking, as it is impossible to make them even roughly compatible with estimates of b and c from $F_K/F_{\pi}f_{+}(0)$ or from a quark-model description of baryon mass splittings. In addition, if we let the value of $M_{\pi_N}^2$ increase, we find that the real branches b = -0.3945 and -20.65 meet and split off again into complex conjugate branches as $M_{\pi_N}^2 \approx 0.90$ GeV² is reached. For these reasons, we choose not to consider further the last four solutions of Table II. The behavior just described

b _p	C _p	M _{πN}	$\tan \theta_{P}$	$M_{\epsilon}^{2} + M_{\epsilon}$, ^{2 a}	M _e ,	Μ _ε	$\tan \theta_{S}$
-0.218	-0.308	0.888	0.052	1.70	1.140	0.630	0.955
	-1.30			1.20	0.993	0.456	0.508
+0.120	-0.010	0.766	0.184	1.70	1.050	0.736	-6.07
	-1.17			1.20	0.810	0.715	-6.67
-0.390	-0.60	0.715	0.244	1.70	1.23	0.413	1.204
	-1.32			1.20	1.02	0.391	0.982
-20.55	-1.48	0,550	0.947	1.70	1.30	0,120	0.054
	4.28			1.20	1.09	0.123	0,063

TABLE II. Solutions to bilinear-breaking model. The mass value $M_{\pi_N} = 0.906$ is assumed, identical to one *predicted* for the $b_p = -0.218$ solution of linear case $(M_{\pi_N}^2 = 0.823)$. All masses are in GeV.

^a These are input values. Both c_p values go with each $M_e^2 + M_e^2$ value.

gives some added justification for discarding the $b_p = -0.352$ case of the linear model, since it disappears as a viable candidate if one tries to "lift" the scalar spectrum by adding a bilinear breaking term. At the same time, of course, new candidates for "physical" solutions are introduced, and it is these that we wish to discuss in the remainder of this section.

The most striking feature of Table II is that for each of the b_p values, $b_p = -0.218$ and $b_p = 0.120$, one has *both* of the currently competing pictures of chiral-symmetry breaking as possible solutions. Namely, for *each* b_p *either* $c \approx -\sqrt{2}$ (Ref. 14) [near SU(2)×SU(2)] or c small²² [near SU(3) Lagrangian symmetry]. This effect becomes even more pronounced as $M_{\pi_N}^2$ is allowed to rise. Thus the restriction to $c \approx -\sqrt{2}$ of the usual σ model, a result characteristic of strict pole dominance of all nonvanishing divergences,²³ disappears when deviations from operator partial conservation of axialvector current (PCAC) are allowed by the introduction of bilinear-breaking terms.

Our primary interest is in the behavior of solutions as symmetry breaking is turned off. We see in Fig. 5, where $\hat{\lambda} (\hat{\lambda} = \langle u_0 \rangle / \langle u_0 \rangle_{\text{physical}})$ is plotted against $\hat{\delta} = \delta / \delta_{\text{physical}}$, that the b = -0.219 and c = -0.3 solution chooses the Goldstone pseudoscalar octet in the symmetry limit, $\hat{\delta} \rightarrow 0$. Note that one can subsequently reach the normal limit by circling two branch points. This behavior is found for all of the values $1.2 \text{ GeV}^2 \leq M_e^2 + M_e^2 \leq 1.7 \text{ GeV}^2$. There is no singularity of λ between $\hat{\delta} = 1$ and $\hat{\delta} = 0$. Moreover, there are no real δ singularities in the whole range $-1 \leq |\hat{\delta}| \leq 1$, and $\hat{\lambda}$ may be expandable about $\hat{\delta} = 0$ in this entire region, which includes the physical solution. Clearly, from Fig.



FIG. 5. Bilinear-breaking cases $M_{\pi_N}^2 = 0.823 \text{ GeV}^2$, $b_p = -0.218$, c = -0.308, $M_e^2 + M_{e'}^2 = 1.7 \text{ GeV}^2$, (solid line) and $M_e^2 + M_{e'}^2 = 1.2 \text{ GeV}^2$ (dashed line). $\hat{\lambda} = \lambda/\lambda_p$ and $\hat{\delta} = \delta/\delta_p$ are dimensionless.

5, a linear fit to $\hat{\lambda}$ in this region is a good approximation.

When we examine the pseudoscalar masses, however, we find that they cannot be described by a linear expansion in $\hat{\delta}$ from $\hat{\delta} = 0$ to $\hat{\delta} = 1$ (physical $\hat{\delta} = 1$ with our normalization). This is shown in Fig. 6, where the pion mass is plotted against $\hat{\delta}$. The pion mass squared rises from zero at $\hat{\delta} = 0$. reaches a local maximum before $\hat{\delta} = 1$, and then falls sharply through the physical value $m_{\pi}^2 = 0.019$ GeV². In addition, as the value of $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ is lowered from 1.7 GeV^2 , the position of the local minimum, which occurs at small negative $\hat{\delta}$ values when $M_{\epsilon'}^2 + M_{\epsilon'}^2$ is above 1.7 GeV², moves toward more positive $\hat{\delta}$ values and occurs at about $\hat{\delta} = 0.5$ when $M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 1.2 \text{ GeV}^{2}$. This behavior is shown in Fig. 7, and the negative mass-squared values that occur between the physical point at δ = 1 and the symmetry limit at δ = 0 clearly violate the criterion that the symmetry limit be reached via physically admissable values of the observable quantities. Thus the b = -0.218 and c = 0.3 solution is unable to connect naturally to an $SU(3) \times SU(3)$ symmetry for small scalar mass values. This effect persists as $M_{\pi_N}^2$ is raised to 0.95.

The behavior of the solution having $b_p = 0.12$ (see Ref. 24) and c = -1.17 is shown in Fig. 8. With $M_e^2 + M_e^2 = 1.7 \text{ GeV}^2$, the Goldstone pseudoscalar octet is reached smoothly as $\delta \rightarrow 0$. A branch point in $\hat{\lambda}(\delta)$ occurs at $\hat{\delta} = -0.42$, so a power-series expansion about $\hat{\delta} = 0$ cannot represent the physical solution. As $M_e^2 + M_e^2$ is lowered to 1.2 GeV², the symmetry limit changes from Goldstone to normal, and the latter case is shown by the dashed line in Fig. 8. This critical behavior is nearly the same as that found in the case b = -0.218, c = -1.30, but with the interesting difference that the branch points that occur at small positive $\hat{\delta}$ in the latter case are absent. In summary, the description $b_p = +0.12$, $c_p = -1.17$ connects smoothly to the



FIG. 6. $m_{\pi}^2 \operatorname{vs} \hat{\delta}$ for solid-line solution of Fig. 5 $(M_{\epsilon}^2 + M_{\epsilon})^2 = 1.7 \text{ GeV}^2$.

limit $\delta = 0$, b = 0, $\lambda \neq 0$ (Goldstone pseudoscalar octet) when $M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 1.7 \text{ GeV}^{2}$, while it connects smoothly to the limit $\delta = 0$, b = 0, $\lambda = 0$ (normal limit) when $M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 1.2 \text{ GeV}^{2}$.

The $b_{\rho} = +0.12$ and c = -0.09 solution (not shown in a figure) is completely isolated from the symmetry limit for all values of $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ which we have tried (1.2 to 2.0 GeV²). Negative pion masses appear at about $\hat{\delta} = 0.9$ as $\hat{\delta}$ is reduced from $\hat{\delta} = 1$, the physical point. We have not investigated the details of the approach to the symmetry limit for this case.

A summary of the interesting solutions is presented in Table III for the value $M_{\pi_N}^2 = 0.822 \text{ GeV}^2$. At this value of $M_{\pi_N}^2$ the b = -0.218, c = -1.30, d = 0 linear-model solution is reproduced, as expected, while the remaining solutions are peculiar to the bilinear-breaking model.

Details of Solutions for Choice $M_{\pi_N}^2 = 0.950 \text{ GeV}^2$

As mentioned earlier, the solutions $b_p = -0.395$ and $b_p = -20.6$ become physically uninteresting (meet and become complex) as $M_{\pi_N}^2$ is allowed to rise. Taking the round figure $M_{\pi_N}^2 = 0.950 \text{ GEV}^2$ $(M_{\pi_N} = 0.974 \text{ GeV})$, we examine the real solutions which evolve from the $b_p = -0.218$ and $b_p = +0.12$ cases discussed above. These roots move closer together, remaining real, to values -0.142 and +0.013, respectively, as $M_{\pi_N}^2$ is lifted to 0.950 GeV². For each *b*(physical), the two *c* roots again come in a "small" [near SU(3) Hamiltonian] and a "large"¹⁴ [$c \approx -\sqrt{2}$, near SU(2)×SU(2) Hamiltonian] size. A summary of the solutions is presented in Table IV for the scalar mass inputs M_{e}^{2} $+M_{\epsilon'}^2 = 1.7 \text{ GeV}^2$ and $M_{\epsilon'}^2 + M_{\epsilon'}^2 = 2.0 \text{ GeV}^2$. We note that both the $b_p = -0.142$ and $b_p = 0.013$ solutions produce acceptable values for the scalar nonet masses. The mixing between ϵ and ϵ' is, however, considerably different in these two cases.

The transition from normal to Goldstone limit again occurs for the case b = -0.142, c = -1.29,



FIG. 7. $m_{\pi}^2 \operatorname{vs} \hat{\delta}$ for dashed-line solution of Fig. 5. $(M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.2 \text{ GeV}^2)$.



FIG. 8. $\hat{\lambda}$ vs $\hat{\delta}$ for the bilinear-breaking cases $M_{\pi_N^2} = 0.823 \text{ GeV}^2$, $b_p = 0.12$, c = -1.17. $M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.7 \text{ GeV}^2$ (solid line) and $M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.2 \text{ GeV}^2$ (dashed line). There is no singularity in dashed path.

but at higher scalar mass values than before. This is shown in Fig. 9, where the solid line $(M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 2.0 \text{ GeV}^{2})$ is the Goldstone limit and the dashed line $(M_{\epsilon}^{2} + M_{\epsilon'}^{2} = 1.7 \text{ GeV}^{2})$ is the normal limit. The basic character of this solution, which is the one that tends in the limit d = 0 (no bilinear breaking) to the linear-breaking-model case studied in detail by several authors,^{6,8} remains unchanged as $M_{\pi_{N}}^{2}$ is increased.

For the alternative c value (-0.178), the behavior of the solution near $\hat{\delta} = 0$ is shown in Fig. 10. The Goldstone nature of the limit is unchanged by varying $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ over a wide range, but the limit is screened from the physical point by a region of instability (pion mass squared negative) for all $M_{\epsilon}^2 + M_{\epsilon'}^2$ values less than about 2.6 GeV². The pion mass squared is plotted vs $\hat{\delta}$ for the cases $M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.7$ GeV² and $M_{\epsilon}^2 + M_{\epsilon'}^2 = 2.0$ GeV² in Fig. 11.

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Next, we see in Fig. 12 that the $b_p = 0.013$, c = -1.25 case connects smoothly to the Goldstone limit for both values of $M_e^2 + M_{e'}^2$ tried. The transition to a normal limit occurs when $M_e^2 + M_{e'}^2$ is lowered to 0.8 GeV². The small-c case, c = -0.001, has highly erratic behavior, and the pion mass squared becomes negative for $\hat{\delta} \approx 0.99$,

TABLE III. Summary of symmetry limits of the interesting solutions to the bilinearbreaking model with $M_{\pi_N}^2 = 0.823 \text{ GeV}^2$. Smooth limit means that no instabilities and/or singularities are encountered as $\delta \rightarrow 0$. Goldstone limit is $\langle u_0 \rangle \neq 0$, b = 0 as $\delta \rightarrow 0$; normal limit means $\langle u_0 \rangle = 0$, $\langle u_8 \rangle = 0$ as $\delta \rightarrow 0$. $\langle u_0 \rangle_{\text{physical}} \equiv \lambda(\delta_p)$ expansion exists if there are no $\lambda(\delta)$ singularities in $|\delta| \leq |\delta_p|$.

$M_{\pi N}^{2} = 0.8$	323 GeV^2 c_p	$M_{\epsilon}^{2} + M_{\epsilon}'^{2}$ a (GeV ²)	Smooth limit?	Character of limit	$\langle u_0 \rangle$ expansion?
-0.218	-1.30	1.70 1.20	yes no	Goldstone normal	no no
<i>.</i>	-0.30	1.70 1.20	yes no	Goldstone Goldstone	maybe maybe
+0.120	-1.17	1.70 1.20	yes yes	Goldstone normal	no maybe
	-0.09	1.70 1.20	no no	•••	no no

$M_{\pi}^{2} = 0.$	950 GeV^2			$M_e^2 + M_e^2$ a	Mei	Me		
b_p^N	c _p	M_{κ}	$\tan \theta_P$	(GeV^2)	(GeV)	(GeV)	$ an heta_S$	
-0.143	-0.178 -1.29	1.15	-0.015	$1.70\\2.00$	$\begin{array}{c} 1.130\\ 1.192 \end{array}$	$0.650 \\ 0.760$	$0.525 \\ 0.776$	
0.013	-0.001 -1.25	0.845	0.186	$1.70\\2.00$	$\begin{array}{c} 1.040\\ 1.173\end{array}$	$\begin{array}{c} 0.783 \\ 0.792 \end{array}$	5.44 9.64	

TABLE IV. Real solutions to bilinear-breaking model with $M_{\pi_N}^2 = 0.950 \text{ GeV}^2$.

^a Input

very near the physical point. We have not followed this highly unusual case in detail, since there is clearly no possibility that one can talk about a smooth symmetry limit for any reasonable choice of input parameters.

A recapitulation of the solutions is presented in Table V. Comparing with Table III, we see that the nature of each solution is essentially unchanged by the increase in $M_{\pi_N}^2$ from 0.822 GeV² to 0.95 GeV². The only effect is that scalar-mass-dependent changes in the character of the $\delta \rightarrow 0$ limit occur at different values of $M_{\epsilon}^2 + M_{\epsilon'}^2$ as $M_{\pi_N}^2$ is increased. For example, the first solution in Table III, where $M_{\pi_N}^2 = 0.822$ GeV², changes from a "normal" limit, but one that can only be reached by circling two branch points, to a smooth "Goldstone" limit at about $M_{\epsilon}^2 + M_{\epsilon'}^2 \approx 1.5$ GeV², while, referring to Table V, this solution makes the change at $M_{\epsilon}^2 + M_{\epsilon'}^2 \approx 1.9$ GeV² when $M_{\pi_N}^2 = 0.95$ GeV².

IV. DISCUSSION

In this section we shall summarize the essential features of the solutions to the linear- and bilinear-breaking σ models, elaborating on several



FIG. 9. Bilinear-breaking cases $M_{\pi_N}^2 = 0.95 \text{ GeV}^2$, $b_p = -0.142$, c = -1.29, $M_e^2 + M_{e'}^2 = 1.7 \text{ GeV}^2$ (dashed line) and $M_e^2 + M_{e'}^2 = 2.0 \text{ GeV}^2$ (solid line). The s-shaped curve for the former case involves two branch points.

features which are possible only in bilinear breaking, and comment on the model dependence of the approach to symmetry limits.

(i) Linear-breaking model. There are two solutions to this model, one of which has been discussed by Carruthers and Haymaker⁶ for the special case in which the mass term $u_i^2 + v_i^2$ is dropped ($\alpha = 0$). (Following an observation by Olshansky,⁸ Carruthers and Haymaker subsequently discussed the influence of the mass term, but in the context of a model with no octet breaking.¹⁰) We looked at both solutions in detail in Sec. II. In agreement with the analysis of Refs. 8 and 10, both solutions failed to have a smooth Goldstone limit when the originally assumed value of $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ $=1.7 \text{ GeV}^2$ was decreased to a value small enough that the "critical" value of α was exceeded. However, the solution with $b_p = -0.218$ (see Ref. 25) became unstable when M_{ϵ} and $M_{\epsilon'}$ were in the neighborhood of 0.500 GeV and 1.0 GeV, respectively, while the $b_{p} = -0.352$ case lost the Goldstone limit at much lower values of the scalar masses. Whether or not the model is near a critical¹⁰ value of the mass parameter depends upon the solution one chooses. The fact that the $b_p = -0.218$ solution, which alternates between the Goldstone pseudoscalar octet limit and the "pseudonormal"²⁶ limit as M_{ϵ} and $M_{\epsilon'}$ are varied around 0.500 GeV and 1.0 GeV, respectively, gives a more plausible picture of ϵ and ϵ' masses, and mixing might be suggestive. However, the solution into which it develops



FIG. 10. Bilinear-breaking cases $M_{\pi_N}^2 = 0.95 \text{ GeV}^2$, $b_p = -0.142$, c = -0.178, $M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.7 \text{ GeV}^2$ (dashed line) and $M_{\epsilon}^2 + M_{\epsilon'}^2 = 2.0 \text{ GeV}^2$ (solid line).



FIG. 11. Plot of m_{π}^2 in GeV² vs $\hat{\delta}$ for solutions of Fig. 10. Dashed line is $M_{\epsilon}^2 + M_{\epsilon'}^2 = 1.7 \text{ GeV}^2$ case and solid line is $M_{\epsilon'}^2 + M_{\epsilon'}^2 = 2.0 \text{ GeV}^2$ case.

as the bilinear term is introduced is only one of several reasonable candidates in this more complicated model.

(ii) Bilinear-breaking model. There are eight solutions possible in this model, though four are complex and of no interest in the case $M_{\pi_N} \approx 0.980$ GeV, as discussed in the latter part of Sec. III. In fact, the solution $b_p = -0.352$ to the linear model, which had been discounted in previous work,⁶ evolves into one of the complex solutions as M_{π_N} is raised from the (predicted) value $M_{\pi_N} = 0.904$ GeV of the linear model to (assumed) values above 0.940 GeV in the bilinear model. This is perhaps another justification for preferring the $b_p = -0.218$ solution⁶ as a guide to general results in the linear-breaking studies. However, the number of viable solutions certainly increases when the extra term is added to £. The most interesting new twist is the appearance of two values of c for each value of b_{p} ; a "small" $c \ (c \approx -0.2$, supporting the Brandt and Preparata¹⁹ point of view) and a "large"



FIG. 12. Plot of $\hat{\lambda}$ vs $\hat{\delta}$ for $M_{\pi N}^2 = 0.95$ GeV², $b_p = +0.013$, c = -1.25, $M_e^{2} + M_e^{2} = 1.7$ GeV² (dashed) and $M_e^{2} + M_e^{2} = 2.0$ GeV² (solid). m_{π}^2 dependence on $\hat{\delta}$ for latter case is shown at bottom.

 $c \ (c \approx -\sqrt{2})$, supporting the Gell-Mann-Oakes-Renner and Glashow-Weinberg point of view). All masses, the mixing angles, and F_K/F_{π} are the same for these two widely differing values of c. Depending upon the assumed values for M_{π_N} and $M_{\epsilon}^2 + M_{\epsilon'}^2$, both solutions, one or the other, or neither may connect smoothly to a Goldstone SU(3) ×SU(3) limit. The value of c found from the tree approximation does not depend upon the way in which the symmetry limit is achieved, and one can have both, either, or neither of the Brandt-Preparata and Gell-Mann-Oakes-Renner pictures of SU(3)×SU(3) Hamiltonian breaking compatible with the pseudoscalar octet as Goldstone particles as $\delta \rightarrow 0$.

In principle, bilinear breaking affords two interesting possibilities which are not present in the linear-breaking model. Solutions are actually found among those studied which closely corre-

TABLE V. Summary of symmetry limits of solutions to bilinear model with $M_{\pi_N}^2 \approx 0.950$ GeV².

$\frac{M_{\pi_N}^2}{b_p^2} = 0.9$	0.00 GeV^2	$M_{\epsilon}^{2} + M_{\epsilon} r^{2}$ a (GeV ²)	Smooth limit?	Character of limit	$\langle u_0 \rangle$ expansion?
-1.42	-1.29	$1.70\\2.00$	no yes	normal Goldstone	no no
	-0.178	$1.70\\2.00$	no no	Goldstone Goldstone	maybe maybe
0.013	-1.25	$1.70\\2.00$	yes yes	Goldstone Goldstone	no no
	-0.001	$\begin{array}{c} 1.70 \\ 2.00 \end{array}$	no no	•••	no no

^a Inputs.

spond to these two cases. We can have (1) b = 0, $c \approx -\sqrt{2}$ and we find a case $b_p = 0.01$, $c_p = -1.25$ when $M_{\pi_N} = 0.975$ GeV. Note that $m_{\pi, K, \eta}^2 \sim \delta$ as δ \rightarrow 0 in this solution. This is almost a literal reproduction of the picture in the original Gell-Mann-Oakes-Renner paper.¹⁴ We can also have (2) $m_{\pi, K, \eta}^2 \sim \delta^2$ as $\delta \rightarrow 0$, and this is the case for the solution $b \approx -0.2$, $c \approx -0.2$ [approximate SU(3) symmetry of the Lagrangian]. Let us briefly discuss the origin of these effects. (a) In the σ model which has breaking linear in the fields, the extremum equation $\partial \mathcal{L} / \partial u_{8} = 0$ requires the condition λ' $=0 \Rightarrow c=0$ (or $b=\lambda'/\sqrt{2} \lambda=0 \Rightarrow c=0$). However, if $a/d = 4\lambda_b (\frac{2}{3})^{1/2}$ (see Ref. 27) this requirement is lifted in the model with breaking which is bilinear in the basis fields. (b) If we write

$$\mathcal{K} = \mathcal{K}_{\text{symmetric}} - (u_0' + u_8' c), \qquad (7)$$

then we can express the pseudoscalar masses

$$m_{i}^{2} = \frac{1}{F_{i}^{2}} \left[\left(\frac{2}{3} \right)^{1/2} + c d_{8ii} \right] \\ \times \left[\left(\frac{2}{3} \right)^{1/2} \langle 0 | u_{0}' | 0 \rangle + d_{8ii} \langle 0 | u_{8}' | 0 \rangle \right].$$
(8)

In the usual σ model, $u'_0 = \delta u_0$ and $u'_8 = \delta u_8$, where δ is the explicit SU(3)×SU(3)-breaking parameter. In the Goldstone limit where $\langle u_0 \rangle$ and $F_i \rightarrow \text{constant}$ and $\langle u_8 \rangle \rightarrow 0$, then $m_i^2 \sim \text{constant} \times \delta$. This is the only way in which the Goldstone limit can be achieved in the model with linear breaking. With breaking which is bilinear in the basis fields u_i and v_i , we can write

$$u_0' = \delta(au_0 + dU_0),$$

$$u_0' = \delta(au_0 + dU_0).$$
(9)

Since

$$U_0 = 4(\frac{2}{3})^{1/2} u_0^2 + \cdots , \qquad (10)$$

it is possible to have

$$\langle 0 | u_0' | 0 \rangle = \delta \left[a \langle u_0 \rangle + 4 \left(\frac{2}{3} \right)^{1/2} d \langle u_0 \rangle^2 + \cdots \right]$$
(11)

and

$$\langle 0 | u_0' | 0 \rangle \sim \delta^2$$

if

$$\langle 0 | u_0 | 0 \rangle |_{\delta=0} = -\frac{a}{4(\frac{2}{3})^{1/2}d}$$
.

Referring to Eq. (4.2), this means $m_i^2 \sim \delta^2$ as $\delta \rightarrow 0$. This quadratic behavior is observed to occur in the Brandt-Preparata solution $b \approx -0.2$, $c \approx -0.2$, as we have mentioned. It is interesting that, of all the solutions studied, this is one of several which are free of singularities on the real line $|\hat{\delta}| \leq 1$ (see Table III). However, though a straight line $\lambda = \lambda_0 + (d\lambda/d\delta)\delta$ describes the vacuum dependence on δ very well, the pseudoscalar masses require at least cubic terms in δ for a description of their behavior in the interval $|\hat{\delta}| \leq 1$. Implications of this curious behavior may be worth pursuing.

In summary, we refer to Table VI, where we see that it is as likely as not that the tree-approximation solutions simply cannot be connected smoothly to a meaningful $SU(3) \times (3)$ limit. In those cases where the limit *can* be obtained smoothly, the radius of convergence of a power-series expansion of "one-point" functions, and consequently masses, couplings, etc., is generally too small to provide a description of the physical parameters. In one of the possible exceptions to this observation, the masses, though probably analytic in δ , require at least cubic terms in δ to reproduce their behavior. Thus, our study does not lend support to the hope that a model with a smooth $SU(3) \times SU(3)$ limit can be unambiguously identified, and then provide a useful expansion point for successfully describing low-energy hadron phenomena.

On the positive side, it is certainly true that if any smooth limit exists in these models it is either "normal" or "pseudoscalar octet of Goldstones." A κ -Goldstone role is not possible,²⁸ for example. It is also worth pointing out that of the three leading candidates for reasonable, physical solutions to the bilinear-breaking model, one with reasonable scalar mass spectra shows the "critical" switch between Goldstone and "pseudonormal" or normal limits as the relative size of the mass parameter, α , is varied over a "physically reasonable" range such that $M_{\epsilon} \approx 0.5$ GeV, $M_{\epsilon'} \approx 1$ GeV.

Regarding the question of whether or not the addition to the usual σ model of terms which are bi-

TABLE VI. Condensation of Tables III and V for "most reasonable" solutions with $M_{\epsilon}^{2}+M_{\epsilon}$, $^{2}=1.2$, 1.7, and 2.0 GeV².

	M	$T_{\pi v}^{2} = 0.822 \text{ GeV}$	$\sqrt{2}$	$M_{\pi v}^{2} = 0.950 \text{ GeV}^{2}$		
$M_{\epsilon}^{2} + M_{\epsilon} r^{2}$	b = -0.218	b = -0.218	b = -0.120	b = -0.143	b = -0.143	b = -0.01
(GeV ²)	c = -0.300	c = -1.30	c = -1.17	c = -0.178	c = -1.29	c = -1.25
1.20	no	no	normal	no	no	normal
1.70	Goldstone	Goldstone	Goldstone	no	no	Goldstone
2.00	Goldstone	Goldstone	Goldstone	no	Goldstone	Goldstone

linear in the fields and transform as $(3^*, 3) + (3, 3^*)$ members is "harmless" (by which we mean that flexibility is provided without affecting general results) we find that substantial changes occur in the nature of the new solutions compared to the nature of the linear-model solutions. Several of these solutions may provide as reasonable descriptions of (poorly known) masses and mixings as the "preferred" $b_p = -0.218$ linear-breaking-model solution.⁶ In view of the poorly determined K_{l3} parameters²⁹ and of π -N σ terms,³⁰ it is not clear where one may turn to find additional evidence which may discriminate among the possibilities. Consequently, we feel that inferences about possible chiralsymmetry limits and about possible "critical point" effects which are drawn from the usual $\boldsymbol{\sigma}$ model (linear breaking) may not be supported by variants of the model and should be viewed with caution.

APPENDIX

The Lagrangian we work with in Sec. III reads

$$\mathcal{L} = -\frac{1}{2} (\partial_{\mu} u_{i} \partial_{\mu} u_{i} + \partial_{\mu} v_{i} \partial_{\mu} v_{i}) - \alpha I_{2} - \alpha' (I_{2})^{2} - \beta I_{3} - \gamma I_{4} + \delta [a(u_{0} + cu_{8}) + d(U_{0} + cU_{8})].$$
(A1)

The results of the linear model in all subsequent formulas are recovered by setting d = 0. Written in terms of the basis fields, the invariants and bilinear $(3^*, 3) + (3, 3^*)$ members are

$$I_{2} = u_{i}^{2} + v_{i}^{2} \quad (\text{``mass'' term}),$$

$$I_{3}^{+} = 4d_{ijk}(u_{i} u_{j} u_{k} - 3v_{i} v_{j} u_{k}) - 6\sqrt{6} u_{0}u_{i} u_{i}$$

$$+ 6\sqrt{6} u_{0}v_{i} v_{i} + 12\sqrt{6} v_{0}v_{i} u_{i}$$

$$+ 6\sqrt{6} u_{0}^{3} - 18\sqrt{6} v_{0}u_{0},$$
(A2)

$$I_{4} = 2(d_{ijn} d_{nkl} + i f_{ijn} d_{nkl} + i d_{ijn} f_{nkl} - f_{ijn} f_{nkl})$$

$$\times (u_{i} u_{j} u_{k} u_{l} + v_{i} v_{j} v_{k} v_{l} + 4u_{i} u_{j} v_{k} v_{l} - 2u_{i} v_{j} u_{k} v_{l}),$$

and

$$U_{0} = 4(\frac{2}{3})^{1/2}(u_{0}^{2} - v_{0}^{2}) - 2(\frac{2}{3})^{1/2} \sum_{j=1}^{8} (u_{j}^{2} - v_{j}^{2}),$$
(A3)
$$U_{8} = -4(\frac{2}{3})^{1/2}(u_{0}u_{8} - v_{0}v_{8}) + 4 \sum_{j=1}^{8} d_{8jj}(u_{j}^{2} - v_{j}^{2}).$$

The divergences of the axial-vector currents are given by

$$\partial_{\mu}A^{i}_{\mu} = \delta(\frac{2}{3})^{1/2} \left[a + 4(\frac{2}{3})^{1/2} d \langle u_{0} \rangle \right] v_{i}$$

$$+ 4(\frac{2}{3})^{1/2} \delta d \left(\tilde{u}_{0}v_{i} + v_{0}\tilde{u}_{i} + \sum_{j,k=1}^{8} d_{ijk} \tilde{u}_{j} v_{k} \right)$$

$$+ \text{ octet breaking ,}$$
(A4)

$$u_0 = \langle u_0 \rangle + \tilde{u}_0.$$

The locations of the extrema of the classical potential determine the vacuum expectation values of the fields u_0 and u_g in the lowest order, and are given by the equations

$$\delta a + 8\delta d(\frac{2}{3})^{1/2} \lambda + \frac{8\delta dc}{\sqrt{3}} \lambda b$$

= $\lambda \Big[2\alpha + 4\alpha' \lambda^2 (1 + 2b^2) + 4\sqrt{6} \beta \lambda (1 - b^2) + \frac{16}{3} \gamma \lambda^2 (1 + 6b^2 - 2b^3) \Big]$ (A5)

from

$$\frac{\partial \mathcal{L}}{\partial u_0} = 0,$$

and

$$\delta ac - 4\delta d(\frac{2}{3})^{1/2} \lambda(c + \sqrt{2}b) - 8(\frac{2}{3})^{1/2} \delta dc \lambda b$$
$$= \sqrt{2}b\lambda [2\alpha + 4\alpha'\lambda^2(1+2b^2) - 4\sqrt{6}\beta\lambda(1+b)]$$
$$+ 16\gamma\lambda^2(1-b+b^2)]$$
(A6)

from

$$\frac{\partial \mathcal{L}}{\partial u_8} = 0,$$

where

$$\lambda = \langle 0 | u_0 | 0 \rangle$$

and

$$b = \frac{\langle 0 | u_8 | 0 \rangle}{\sqrt{2} \langle 0 | u_0 | 0 \rangle}.$$

In order to fix parameters, the expressions for masses are necessary, and we have

$$m_{\pi}^{2} = x_{1} + (x_{2} - x_{4}) - 2(bx_{2} - x_{5}) + \frac{2}{3}x_{3}(1 + b)^{2},$$

$$m_{K}^{2} = x_{1} + (x_{2} - x_{4}) + (bx_{2} - x_{5}) + \frac{2}{3}x_{3}(1 - b + 7b^{2}),$$

$$m_{\pi}^{2} + m_{\pi}r^{2} = 2x_{1} - (x_{2} - x_{4}) + 2(bx_{2} - x_{5}) + \frac{2}{3}x_{3}(2 - 2b + 5b^{2}),$$

$$m_{\pi}r^{2} - m_{\pi}r^{2} = \begin{cases} 3(x_{2} - x_{3}) + 2(x_{3} - x_{3}) - \frac{2}{3}x_{3}(2 - b) \end{cases}$$

(A7)

$$m_{\eta'}{}^{2} - m_{\eta}{}^{2} = \left\{ \left[3(x_{2} - x_{4}) + 2(x_{2}b - x_{5}) - \frac{2}{3}x_{3}b(2 - b) \right]^{2} + 8\left[(x_{2}b - x_{5}) + \frac{2}{3}x_{3}b(2 - b) \right]^{2} \right\}^{1/2}$$

for the pseudoscalars, and

$$M_{\pi_N}^2 = x_1 - (x_2 - x_4) + 2(bx_2 - x_5) + 2x_3(1 + b)^2,$$

$$M_{\kappa}^2 = x_1 - (x_2 - x_4) - (x_2b - x_5) + 2x_3(1 - b + b^2),$$

$$M_{\epsilon}^2 + M_{\epsilon'}^2 = 2x_1 + 4\alpha'\lambda^2(1 + 2b^2) + (x_2 - x_4)$$

$$- 2(x_2b - x_5) + 2x_3(2 - 2b + 5b^2),$$
 (A8)

$$M_{\epsilon'}{}^2 - M_{\epsilon}{}^2 = \left\{ \left[8\alpha'\lambda^2(1-2b^2) + x_2(3+2b) + 2x_3(2-b)b \right]^2 + 8\left[(-x_2b + x_5) + 2x_3(2-b) + 8\alpha'\lambda^2 b \right]^2 \right\}^{1/2} \right\}^{1/2}$$

for scalars. Here we have defined the quantities

$$\begin{aligned} x_1 &= 2\alpha + 4\alpha'^2 \lambda^2 (1 + 2b^2) ,\\ x_2 &= 4\sqrt{6} \beta \lambda ,\\ x_3 &= 8\gamma \lambda^2 , \qquad (A9)\\ x_4 &= 4(\frac{2}{3})^{1/2} \delta d ,\\ x_5 &= \frac{\delta dc}{\sqrt{3}} . \end{aligned}$$

The inputs m_{π}^{2} , m_{K}^{2} , m_{η}^{2} , $m_{\eta'}^{2}$, and $M_{\pi N}^{2}$ determine b, x_{3} , $bx_{2} - x_{5}$, $x_{2} - x_{4}$, and x_{1} . M_{κ}^{2} is then predicted. Specifying $M_{\epsilon}^{2} + M_{\epsilon'}^{2}$ then fixes α and $\lambda^{2}\alpha'$. Our procedure is to express x_{3} , $bx_{2} - x_{5}$, $x_{2} - x_{4}$, and x_{1} in terms of m_{π}^{2} , m_{K}^{2} , $m_{\eta}^{2} + m_{\eta'}^{2}$, $M_{\pi N}^{2}$, and b and then to substitute into the $m_{\eta'}^{2} - m_{\eta}^{2}$ equation. This yields a quartic equation for b. For each b solution, values of x_{3} , $bx_{2} - x_{5}$, $x_{2} - x_{4}$, and x_{1} are computed and the value of M_{κ}^{2} found. The values of M_{ϵ}^{2} and $M_{\epsilon'}^{2}$ can be computed separately at this point. The extremum equations (A5) and (A6) are used to evaluate c, which satisfies a quadratic equation. Thus two c values are found for each of the four b values. The individual values of x_{2} , x_{4} , and x_{5} are determined by the expression

$$x_4 = \frac{\sqrt{2} (2-b)}{4(1+b)(2\sqrt{2}-c)} - \frac{\sqrt{2} b M_{\kappa}^2}{4(1+b)c}$$
(A10)

and the now known values of c, b, $bx_2 - x_5$, and $x_2 - x_4$. The value of $\delta a/\lambda_p$ can be determined from

$$x_{6} \equiv \frac{\delta a}{\lambda_{p}} = \frac{1}{2} \left[\frac{(2-b)m_{\kappa}^{2}}{2\sqrt{2}-c} + \frac{bM_{\kappa}^{2}}{c} \right]$$
(A11)

for each b, c combination. Equations (A10) and (A11) follow by combining the extremum equations and mass formulas. The dimensionless quantity x_4/x_6 measures the ratio of bilinear to linear breaking.

To study $\lambda(\delta)$, $b(\delta)$, and masses (δ) we fix x_1 ,

 x_2 , x_3 , x_4 , x_5 , and x_6 (where $x_5/x_4 = c/\sqrt{2}$) and return to the extremum equations and mass formulas with the substitutions

$$\alpha - \alpha , \quad x_3 - x_3 \hat{\lambda}^2 ,$$

$$\alpha' \lambda_p^2 - \alpha' \hat{\lambda}^2 , \quad x_4 - x_4 \hat{\delta} ,$$

$$x_2 - x_2 \hat{\lambda}^2 , \quad x_5 - x_5 \hat{\delta} ,$$

(A12)

where $\hat{\lambda} \equiv \lambda/\lambda_p$ and $\hat{\delta} \equiv \delta/\delta_p$. That is, since the scales of λ and δ are not fixed, we can only study the dependence of the dimensionless quantity $\hat{\lambda}$ on the dimensionless parameter $\hat{\delta}$. We find a cubic equation for $\hat{\lambda}$ as a function of b, and $\hat{\delta}$ is determined from knowledge of $\hat{\lambda}$. Expressing $\hat{\lambda}$ directly as a function of $\hat{\delta}$ is too complicated, so we choose to find values of $\hat{\lambda}$ and $\hat{\delta}$ by varying b over a range which includes b_p and b = 0. Working in this way we are able to map out $\hat{\lambda}(\hat{\delta})$ and $b(\hat{\delta})$ in the region including $\hat{\delta} = 0$ and $\hat{\delta} = 1$. The physical root is defined as the one for which $\hat{\lambda} = 1$ and $\hat{\delta} = 1$ at $b = b_p$. Our primary interest is the behavior of this root in the region around $\hat{\delta} = 0$ which includes $\hat{\delta} = 1$.

Mixing angles. We define the $\eta - \eta'$ and $\epsilon - \epsilon'$ mixing angles, θ_P and θ_S , as follows:

$$\eta' = \cos\theta_P v_0 + \sin\theta_P v_8,$$

$$\eta = -\sin\theta_P v_0 + \cos\theta_P v_8$$
(A13)

and

$$\epsilon' = \cos\theta_{s} u_{0} + \sin\theta_{s} u_{8},$$

$$\epsilon = -\sin\theta_{s} u_{0} + \cos\theta_{s} u_{8}.$$
(A14)

The mixing angle is given in terms of the masses as

$$\tan\theta = \frac{-2m_{08}^2}{m_{88}^2 - m_{00}^2 - \left[(m_{88}^2 - m_{00}^2)^2 + 4m_{08}^4\right]^{1/2}}.$$
(A15)

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(Ref. 13) sum rule and predicts an approximately correct value for $m_{\eta'}^2 - m_{\eta}^2$. He suggests that the value for $m_{\eta'}^2 - m_{\eta}^2$ can be made to agree with experiment by addition of a bilinear term such as the one we study in Sec. III. Chan and Haymaker (Ref. 11) note that *b* can be considered a variable parameter over a range -0.2 to -0.35 because of its sensitive dependence on the uncertain value of the kaon average mass.

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PHYSICAL REVIEW D

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Constraints on Higgs Fields in Unified Weak-Electromagnetic Gauge Theories*

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In renormalizable unified gauge theories of the weak and electromagnetic interactions, for a given gauge group we ask the following question: Which representations of the scalar Higgs particles can reproduce a vector-meson mass spectrum in which one vector meson, the photon, remains massless and all other vector mesons acquire mass? Given the Higgs representation, we then examine what restrictions are placed on the charge operator of the theory. The examples of SU(3) and $SU(3) \times U(1)$ are worked out in detail. The possible Higgs representations for all the unexceptional classical Lie groups are given. For each such group, there are some low-dimensional representations for which no solution exists.

I. INTRODUCTION AND SUMMARY

The development of renormalizable gauge theories which attempt to unify the weak and electromagnetic interactions¹ has opened a Pandora's box of possible models. Once a gauge group is chosen, the model builder must decide on a representation of scalar Higgs particles, whose nonzero vacuum expectation values (VEV) spontaneously break the gauge symmetry, and assign known and probably unknown fermions to group multiplets in such a way that he satisfies the constraints imposed by the experimental cross sections, masses, moments, decay rates, etc.

In this paper, we present simple criteria for determining, given the underlying group, which representations of the Higgs particles can reproduce a vector-meson mass spectrum in which one vector meson, the photon, remains massless and all other vector mesons acquire mass. Kibble has discussed the choice of Higgs particles in the context of the strong interactions, thus requiring

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