

Chiral symmetry breaking and the Lorentz nature of confinement

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We address the question of the Lorentz nature of the effective interquark interaction in QCD which leads to the formation of the QCD string between color charges. In particular, we start from a manifestly vectorial fundamental interaction mediated by gluons and demonstrate that, as soon as chiral symmetry is broken spontaneously, the effective interquark interaction acquires a self-consistently generated scalar part which is eventually responsible for the formation of the QCD string. We demonstrate this explicitly for a heavy-light quarkonium, using the approach of the Schwinger-Dyson-type equation and the quantum-mechanical Hamiltonian method of the QCD string with quarks at the ends.

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I. INTRODUCTION

In this paper we discuss one of the long-standing problems in QCD—namely, the problem of the Lorentz nature of the long-range confining interquark interaction. QCD is believed to be a stringlike theory at large distances; that is, the long-range interquark interaction is expected to be generated by an extended object—the QCD string. Such a phenomenological picture appears rather successful in various studies of hadronic properties. An important step was made in the framework of the vacuum correlators method (VCM) [1] in which the Lagrangian of the QCD string with quarks at the ends can be derived naturally starting from the fundamental QCD Lagrangian [2]. In addition, the nonperturbative spin-dependent forces in heavy and light quarkonia were found, following the formalism established in Ref. [3]. Thus, for the spin-orbit interaction, a celebrated representation,

$$V_{\text{so}}(r) = \left(\frac{\vec{\sigma}_q \vec{l}_q}{4m_q^2} - \frac{\vec{\sigma}_{\bar{q}} \vec{l}_{\bar{q}}}{4m_{\bar{q}}^2} \right) \left(\frac{1}{r} \frac{\partial \varepsilon}{\partial r} + \frac{2}{r} \frac{\partial V_1}{\partial r} \right) + \frac{1}{2m_q m_{\bar{q}}} (\vec{\sigma}_q \vec{l}_q - \vec{\sigma}_{\bar{q}} \vec{l}_{\bar{q}}) \frac{1}{r} \frac{\partial V_2}{\partial r}, \quad (1)$$

was introduced in Ref. [3], where $\varepsilon(r)$ is the static confining potential, and a general relation (Gromes relation [4]) is valid,

$$\varepsilon' + V_1' - V_2' = 0. \quad (2)$$

For a purely scalar interaction, one obtains at large r 's

$$V_1' = -\varepsilon', \quad V_2' = 0, \quad (3)$$

and this was demonstrated explicitly for the Gaussian approximation for the field correlators in Ref. [5]. For the case of a vector confinement, for example, for the Coulomb potential, one would find

$$V_1' = 0, \quad V_2' > 0, \quad (4)$$

and the coefficient at the spin-orbit term would have the opposite sign. Phenomenology of the heavy-quarkonia spectrum favors the first possibility, Eq. (3), so that one

has evidence that nature prefers scalar interquark interaction, at least for heavy quarks. On the lattice, numerous data also support the first possibility (see Ref. [6] for recent results and the vast bibliography). In the meantime, any quantum-mechanical approach meets severe problems with the description of another celebrated phenomenon—spontaneous breaking of chiral symmetry, which is known to take place in QCD. A full quantum field theory based treatment has to be exploited for this purpose. An example of such a treatment, also based on the VCM, is given by the Schwinger-Dyson-type approach to heavy-light quarkonia suggested in Ref. [7], and the Lorentz nature of confinement for heavy quarks was studied in this formalism in Refs. [7–9]. On the other hand, for light quarks, vectorlike confining interaction would have resulted in the well-known Klein paradox and, then, in problems with building the spectrum of hadrons. No evidence for such problems exists so far. In this paper we prove that, indeed, even for light quarks, if chiral symmetry is broken spontaneously, the effective interquark interaction acquires a self-consistently generated scalar part. This result is quite general and holds regardless of the explicit form of the interquark kernel, provided it is confining and thus leads to spontaneous breaking of chiral symmetry. A link is established between the VCM and potential quark models [10,11] which we refer to as generalized Nambu-Jona-Lasinio (GNJL) models, the latter being widely used for studies of low-energy phenomena in QCD. This is an important outcome of our work since the vast results obtained in the literature in the framework of such quark models are valid for our situation as well—what we do is approach the same problem from another side. Thus we demonstrate that, starting from the fundamental QCD Lagrangian and using the Gaussian approximation for the background field correlators, one can derive a Schwinger-Dyson-type equation for the heavy-light quarkonium which, at large interquark distances, reduces to a Diraclike equation with an effective interquark interaction which contains a dynamically generated scalar part, as a consequence of chiral symmetry breaking. At the same

time, a Schrödingerlike equation with the Hamiltonian of the QCD string with quarks at the ends (in the form of the Salpeter equation) arises naturally from the same Schwinger-Dyson-type equation, if the scalar interaction dominates [12]. We conclude, therefore, that this is the dynamical scalar interaction responsible for the QCD string formation.

II. SCHWINGER-DYSON-TYPE EQUATION FOR A HEAVY-LIGHT QUARKONIUM

In this section we consider a heavy-light quarkonium consisting of a static antiquark and a quark, whose mass is unconstrained and can take any value (we shall be mostly interested in the case of the massless quark). Our starting point is the heavy-light Green's function $S_{q\bar{Q}}$ written in Euclidean space as [7]

$$S_{q\bar{Q}}(x, y) = \frac{1}{N_C} \int D\psi D\psi^\dagger DA_\mu \exp\left\{-\frac{1}{4} \int d^4x F_{\mu\nu}^2 - \int d^4x \psi^\dagger (-i\hat{\partial} - im - \hat{A})\psi\right\} \psi^\dagger(x) S_{\bar{Q}}(x, y|A) \psi(y), \quad (5)$$

where $S_{\bar{Q}}(x, y|A)$ is the propagator of the static antiquark placed at the origin. For further analysis it is convenient to fix the modified Fock-Schwinger gauge [13],

$$\vec{x} \vec{A}(x_4, \vec{x}) = 0, \quad A_4(x_4, \vec{0}) = 0, \quad (6)$$

which ensures that the gluonic field vanishes at the trajectory of the static particle. Then the static antiquark Green's function is simply

$$S_{\bar{Q}}(x, y|A) = S_{\bar{Q}}(x, y) = i \frac{1 - \gamma_4}{2} \theta(x_4 - y_4) e^{-M(x_4 - y_4)} + i \frac{1 + \gamma_4}{2} \theta(y_4 - x_4) e^{-M(y_4 - x_4)}. \quad (7)$$

It is easy now to perform integration of the gluonic field in Eq. (5) to arrive at

$$S_{q\bar{Q}}(x, y) = \frac{1}{N_C} \int D\psi D\psi^\dagger \exp\left\{-\int d^4x L_{\text{eff}}(\psi, \psi^\dagger)\right\} \psi^\dagger(x) S_{\bar{Q}}(x, y) \psi(y), \quad (8)$$

with $L_{\text{eff}}(\psi, \psi^\dagger)$ being the effective Lagrangian of the light quark moving in the field of the static antiquark source:

$$\begin{aligned} \int d^4x L_{\text{eff}}(\psi, \psi^\dagger) &= \int d^4x \psi^\dagger_\alpha(x) (-i\hat{\partial} - im) \psi^\alpha(x) + \int d^4x \psi^\dagger_\alpha(x) \gamma_\mu \psi^\beta(x) \langle A_{\mu\beta}^{\alpha} \rangle \\ &+ \frac{1}{2} \int d^4x_1 d^4x_2 \psi^\dagger_{\alpha_1}(x_1) \gamma_{\mu_1} \psi^{\beta_1}(x_1) \psi^\dagger_{\alpha_2}(x_2) \gamma_{\mu_2} \psi^{\beta_2}(x_2) \langle A_{\mu_1\beta_1}^{\alpha_1}(x_1) A_{\mu_2\beta_2}^{\alpha_2}(x_2) \rangle + \dots, \end{aligned} \quad (9)$$

where all α 's and β 's are fundamental color indices, and the irreducible correlators $\langle A_{\mu_1\beta_1}^{\alpha_1}(x_1) \dots A_{\mu_n\beta_n}^{\alpha_n}(x_n) \rangle$ of all orders enter. The first correlator, $\langle A_{\mu\beta}^{\alpha} \rangle$, obviously vanishes due to the gauge and Lorentz invariances of the vacuum. In what follows we assume the Gaussian dominance to take place in the QCD vacuum and thus we keep only the bilocal correlator

$$\langle A_{\mu\beta}^{\alpha}(x) A_{\nu\delta}^{\gamma}(y) \rangle \equiv 2(\lambda_a)_{\beta}^{\alpha} (\lambda_a)_{\delta}^{\gamma} K_{\mu\nu}(x, y), \quad (10)$$

and neglect contributions of all higher correlators.¹ Then, using the relation $(\lambda_a)_{\beta}^{\alpha} (\lambda_a)_{\delta}^{\gamma} = \frac{1}{2} \delta_{\delta}^{\alpha} \delta_{\beta}^{\gamma} - \frac{1}{2N_C} \delta_{\beta}^{\alpha} \delta_{\delta}^{\gamma}$ and considering the large- N_C limit, we rewrite the effective

¹This approximation leads to an exact Casimir scaling; that is, the ratio of any two potentials between static sources in different representations of the color group is given by the ratio of the Casimir operators evaluated for the given two representations [14]. The Casimir scaling was tested on the lattice [15] and it was found to manifest itself with a very high accuracy, which evidences a suppression of higher gluonic correlators as compared to the Gaussian one and thus justifies the approximation made above.

light-quark Lagrangian in the form:

$$\begin{aligned} L_{\text{eff}}(\psi, \psi^\dagger) &= \psi^\dagger_\alpha(x) (-i\hat{\partial} - im) \psi^\alpha(x) \\ &+ \frac{1}{2} \int d^4y \psi^\dagger_\alpha(x) \gamma_\mu \psi^\beta(x) \psi^\dagger_\beta(y) \gamma_\nu \psi^\alpha(y) \\ &\times K_{\mu\nu}(x, y), \end{aligned} \quad (11)$$

which leads to the Schwinger-Dyson-type equation [7]

$$\begin{aligned} (-i\hat{\partial}_x - im)S(x, y) - i \int d^4z M(x, z)S(z, y) &= \delta^{(4)}(x - y), \\ -iM(x, z) &= K_{\mu\nu}(x, z) \gamma_\mu S(x, z) \gamma_\nu, \end{aligned} \quad (12)$$

for the color trace of the light-quark Green's function $S(x, y) = \frac{1}{N_C} \langle \psi^\beta(x) \psi^\dagger_\beta(y) \rangle$.

In order to evaluate the quark kernel $K_{\mu\nu}(x, y)$ we notice that a celebrated property of the radial gauges [gauge (6) obviously belonging to this class] is a possibility to express the gluonic field A in terms of the field strength tensor F . For the gauge (6) such a relation reads

$$A_4^a(x_4, \vec{x}) = \int_0^1 d\alpha x_i F_{i4}^a(x_4, \alpha \vec{x}), \quad (13)$$

$$A_i^a(x_4, \vec{x}) = \int_0^1 \alpha x_k F_{ki}^a(x_4, \alpha \vec{x}) d\alpha, \quad i = 1, 2, 3,$$

so that the kernel $K_{\mu\nu}$ can be expressed in terms of the field strength correlator $\langle F_{\mu\nu}^a(x) F_{\lambda\rho}^b(y) \rangle$, for which we use the parametrization [1]

$$\langle F_{\mu\nu}^a(x) F_{\lambda\rho}^b(y) \rangle = \frac{\delta^{ab}}{N_C^2 - 1} D(x-y) (\delta_{\mu\lambda} \delta_{\nu\rho} - \delta_{\mu\rho} \delta_{\nu\lambda}) + \Delta^{(1)}, \quad (14)$$

where the second term $\Delta^{(1)}$ is a full derivative and it does not contribute to confinement and therefore will not be

considered below. The profile function $D(x-y)$ decreases in all directions of the Euclidean space, and this decrease is governed by the gluonic correlation length T_g . Lattice simulations give rather small values of $T_g \approx 0.2 \div 0.3$ fm [6,16], so that the profile $D(x-y)$ has the support at $y \approx x$. The term proportional to $D(x-y)$ in (14) contributes to the area law with the string tension [1]

$$\sigma = 2 \int_0^\infty d\tau \int_0^\infty d\lambda D(\tau, \lambda). \quad (15)$$

Then, in view of Eqs. (13) and (14), the quark kernel $K_{\mu\nu}(x, y) = K_{\mu\nu}(x_4 - y_4, \vec{x}, \vec{y})$ is found to be ($\tau = x_4 - y_4$)

$$\begin{cases} K_{44}(\tau, \vec{x}, \vec{y}) = (\vec{x} \vec{y}) \int_0^1 d\alpha \int_0^1 d\beta D(\tau, |\alpha \vec{x} - \beta \vec{y}|), \\ K_{i4}(\tau, \vec{x}, \vec{y}) = K_{4i}(\tau, \vec{x}, \vec{y}) = 0, \\ K_{ik}(\tau, \vec{x}, \vec{y}) = ((\vec{x} \vec{y}) \delta_{ik} - y_i x_k) \int_0^1 d\alpha \int_0^1 d\beta D(\tau, |\alpha \vec{x} - \beta \vec{y}|). \end{cases} \quad (16)$$

The Schwinger-Dyson-type equation (12) is an essentially nonlinear equation. Its linearized form, with the Green's function $S(x, z)$ substituted by the free-quark Green's function $S_0(x, z)$ in the mass operator $M(x, z)$, can be used if the quark is heavy. This was done in Refs. [8,9], where the effective potential for the heavy-quark interaction with the static antiquark source was built, including the nonperturbative spin-orbit interaction. The leading correction to this potential due to the proper string dynamics was identified in Ref. [17]. It was noticed in Ref. [9], however, that the given linearization procedure is self-consistent only if the product of the quark mass and the gluonic correlation length is large, $mT_g \gg 1$. In the case $mT_g \ll 1$, the series of corrections to the leading regime blows up and no conclusion concerning the dynamics of the system can be made [9]. This procedure is useless, therefore, for the purposes of the present paper which is aimed at consideration of the light (massless) quark with its effective mass generation due to the phenomenon of spontaneous breaking of chiral symmetry. Thus we study this case using a different approach.

Below we use two simplifications: (i) we neglect the spatial part of the kernel K_{ik} and (ii) we neglect corrections due to the finiteness of the correlation length T_g . The first approximation utilizes the fact that, although the spatial part of the kernel is important for the correct account of the QCD string rotation, it is not decisive for the Lorentz nature of confinement, yielding only unnecessary complications. The second approximation is justified in view of the results of the lattice simulations which give, as was mentioned before, quite small values of T_g . The latter simplification allows us to approximate the kernel (16) by an instantaneous kernel and thus, for the Fourier trans-

form of K in time, to neglect its dependence on the energy,

$$K_{44}(\omega, \vec{x}, \vec{y}) \equiv K(\omega, \vec{x}, \vec{y}) = K(\vec{x}, \vec{y}) = (\vec{x} \vec{y}) \int_0^1 d\alpha \int_0^1 d\beta \int_{-\infty}^\infty d\tau D(\tau, |\alpha \vec{x} - \beta \vec{y}|). \quad (17)$$

Then the mass operator can be written as

$$M(x, y) = \delta(x_4 - y_4) M(\vec{x}, \vec{y}), \quad (18)$$

$$M(\vec{x}, \vec{y}) = \frac{1}{2} K(\vec{x}, \vec{y}) \gamma_4 \Lambda(\vec{x}, \vec{y}),$$

where, following Ref. [7], we introduced the quantity

$$\begin{aligned} \Lambda(\vec{x}, \vec{y}) &\equiv \sum_{n=-\infty}^{\infty} \psi_n(\vec{x}) \text{sign}(n) \psi_n^\dagger(\vec{y}) \\ &= 2i \int \frac{d\omega}{2\pi} S(\omega, \vec{x}, \vec{y}) \gamma_4 \\ &= 2i S(x_4 - y_4, \vec{x}, \vec{y}) \gamma_4|_{x_4=y_4}, \end{aligned} \quad (19)$$

which is convenient for studies of the Lorentz nature of the interquark interaction. In addition, in the limit $T_g \rightarrow 0$, the profile function $D(\tau, \lambda)$ takes a singular form $D(\tau, \lambda) = 2\sigma \delta(\tau) \delta(\lambda)$ (see also Ref. [18] for the discussion of the singular limit of some stochastic model). For such a profile one finds readily for the kernel (17)

$$K(\vec{x}, \vec{y}) = 2\sigma (\vec{x} \vec{y}) \int_0^1 d\alpha \int_0^1 d\beta \delta(|\alpha \vec{x} - \beta \vec{y}|). \quad (20)$$

Evaluation of the integrals in Eq. (20) is trivial and gives

$$K(\vec{x}, \vec{y}) = 2\sigma \min(|\vec{x}|, |\vec{y}|) = \sigma (|\vec{x}| + |\vec{y}| - |\vec{x} - \vec{y}|), \quad (21)$$

if the vectors \vec{x} and \vec{y} are collinear [the kernel vanishes otherwise, as required by the delta function in Eq. (20)]. The requirement of collinearity of the vectors \vec{x} and \vec{y} ensures that the interaction between the light quark and the static antiquark is due to an infinitely thin string—a two-dimensional object embedded into the four-dimensional space. In order to proceed we simplify the form of the kernel (21) and relax the constraint of collinearity. Thus we approximate the kernel as

$$K(\vec{x}, \vec{y}) = \sigma(|\vec{x}| + |\vec{y}| - |\vec{x} - \vec{y}|). \quad (22)$$

The quark kernel (22) possesses a number of important properties:

- (i) It allows us to pass over back to Minkowski space which will be used from now onward in this paper.
- (ii) It admits a clear interpretation. Indeed, the kernel can be split into two parts: the local part $-\sigma|\vec{x} - \vec{y}|$ which is responsible for the self-interaction of the light quark, and the nonlocal part $\sigma(|\vec{x}| + |\vec{y}|)$ which describes the interaction of the light quark with the static source. Such a form of the kernel is a consequence of the gauge condition (6) which decouples the static particle from the system and brings all the information about the antiquark to the kernel $K(\vec{x}, \vec{y})$.
- (iii) It admits a natural generalization from the linearly rising potential to the potential of a generic form $V(r)$. In order to emphasize this important property we keep the potential as $V(r)$ in all formulas below. Nevertheless, every time we need to specify the form of the potential, the linear confinement is understood as the most phenomenologically justified candidate.
- (iv) With the kernel (22) we establish a link between the VCM and the GNJL models for QCD with instantaneous quark kernels which have a long history in the literature [10,11] and which can be viewed as nonlocal divergence-free generalizations of the Nambu-Jona-Lasinio model [19] (recent exhaustive studies of the mesonic spectrum in this model can be found in Ref. [20]). Below we employ the chiral angle approach which is widely used in such potential quark models.

Combining Eqs. (12), (18), (19), and (22), we arrive at the Schwinger-Dyson-type equation for the heavy-light quarkonium in the form

$$\begin{aligned} \left(-i\gamma_0 \frac{\partial}{\partial t} + i\vec{\gamma} \frac{\partial}{\partial \vec{x}} - m\right) S(t, \vec{x}, \vec{y}) - \int d^3z M(\vec{x}, \vec{z}) S(t, \vec{z}, \vec{y}) \\ = \delta(t) \delta^{(3)}(\vec{x} - \vec{y}), \end{aligned} \quad (23)$$

where

$$\begin{aligned} M(\vec{x}, \vec{z}) &= -\frac{i}{2} K(\vec{x}, \vec{z}) \gamma_0 \Lambda(\vec{x}, \vec{z}), \\ \Lambda(\vec{x}, \vec{z}) &= 2i \int \frac{d\omega}{2\pi} S(\omega, \vec{x}, \vec{y}) \gamma_0, \end{aligned} \quad (24)$$

and the quark kernel is

$$K(\vec{x}, \vec{y}) = V(|\vec{x}|) + V(|\vec{y}|) - V(|\vec{x} - \vec{y}|). \quad (25)$$

In the next section we study the properties of this equation.

III. THE LORENTZ NATURE OF CONFINEMENT

Let us investigate the properties of the quantity $\Lambda(\vec{x}, \vec{y})$ and demonstrate the way it acquires the contribution with the matrix structure $\propto \gamma_0$, since this phenomenon exactly constitutes spontaneous breaking of chiral symmetry.

It was argued in Ref. [7] that Eq. (12) admits linearization via the substitution

$$\Lambda(\vec{x}, \vec{y}) \approx \gamma_0 \delta^{(3)}(\vec{x} - \vec{y}) + \dots, \quad (26)$$

where the ellipsis denotes the terms which are subleading at large distances. Such a substitution was justified then by an explicit computation of this quantity using the spectrum of the resulting linearized equation (23). In Appendix A we give some details of the derivation of Eq. (26), taken from Ref. [7]. The scalar Lorentz nature of the effective interaction follows immediately from the form of Eq. (23) and the matrix structure of $\Lambda(\vec{x}, \vec{y})$ given in Eq. (26). We conclude therefore that Eq. (23) does admit a solution given by the scalar interaction generated in a self-consistent manner. Let us consider in detail this self-consistent generation of the scalar effective interquark interaction.

The separation of the kernel into the local and nonlocal parts mentioned above is the key feature which allows us to proceed. Indeed, let us consider the local part of the kernel first and omit the nonlocal part. Then Eq. (12) reduces to the Dyson equation for the light-quark propagator,

$$(\gamma_0 p_0 - \vec{\gamma} \vec{p} - m - \Sigma(\vec{p})) S(p_0, \vec{p}) = 1, \quad (27)$$

where the mass operator $\Sigma(\vec{p})$ does not depend on the energy due to the instantaneous nature of the interaction and can be evaluated as

$$\Sigma(\vec{p}) = -i \int \frac{d^4k}{(2\pi)^4} V(\vec{p} - \vec{k}) \gamma_0 S(k_0, \vec{k}) \gamma_0. \quad (28)$$

Equations (27) and (28) together lead to the self-consistent nonlinear equation for the quark mass operator [10],

$$\Sigma(\vec{p}) = -i \int \frac{d^4k}{(2\pi)^4} V(\vec{p} - \vec{k}) \gamma_0 \frac{1}{\gamma_0 k_0 - \vec{\gamma} \vec{k} - m - \Sigma(\vec{k})} \gamma_0. \quad (29)$$

If we now parametrize the mass operator in the form

$$\Sigma(\vec{p}) = [A_p - m] + (\vec{\gamma} \hat{p}) [B_p - p], \quad (30)$$

with A_p and B_p being two auxiliary functions, then Eq. (29) gives the self-consistency conditions for such a parametrization,

$$\begin{aligned} A_p &= m + \frac{1}{2} \int \frac{d^3k}{(2\pi)^3} V(\vec{p} - \vec{k}) \sin\varphi_k, \\ B_p &= p + \frac{1}{2} \int \frac{d^3k}{(2\pi)^3} (\hat{p} \cdot \hat{k}) V(\vec{p} - \vec{k}) \cos\varphi_k, \end{aligned} \quad (31)$$

where the angle φ_p —known as the chiral angle—is introduced to obey the condition

$$A_p \cos\varphi_p = B_p \sin\varphi_p, \quad (32)$$

which, given the relations (31), plays the role of the mass-gap equation for the chiral angle. Historically, the chiral angle φ_p is defined such that $\varphi_p(p=0) = \frac{\pi}{2}$ and $\varphi_p(p \rightarrow \infty) = 0$. In Fig. 1 we plot the chiral angle—the solution to the mass-gap equation (32) for the linear confinement. The interested reader can find the details of the chiral angle formalism in Refs. [10,11,21]. A comprehensive analysis of the properties of the mass-gap equation and its solutions for various powerlike potentials is given in Ref. [22].

It is an easy task now to evaluate the quantity $\Lambda(\vec{p}, \vec{q})$, which is the double Fourier transform of $\Lambda(\vec{x}, \vec{y})$:

$$\Lambda(\vec{p}, \vec{q}) = 2i \int \frac{d\omega}{2\pi} S(\omega, \vec{p}, \vec{q}) \gamma_0 = (2\pi)^3 \delta^{(3)}(\vec{p} - \vec{q}) U_p, \quad (33)$$

where

$$U_p = \beta \sin\varphi_p - (\vec{\alpha} \cdot \hat{p}) \cos\varphi_p, \quad \beta = \gamma_0, \quad \vec{\alpha} = \gamma_0 \vec{\gamma}. \quad (34)$$

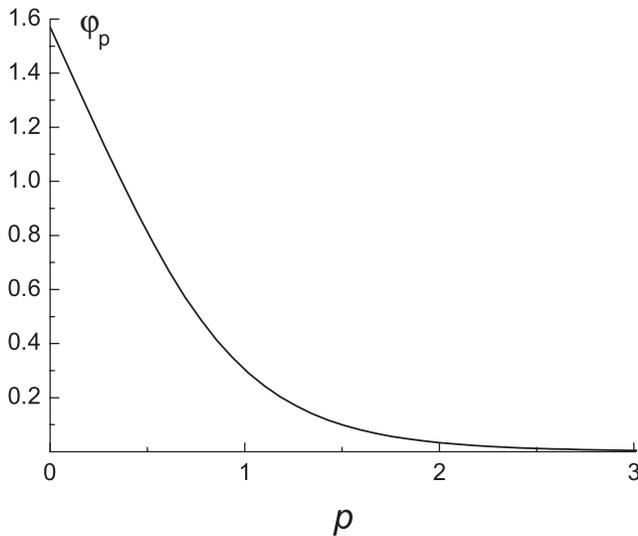


FIG. 1. The profile of the solution to the mass-gap equation (32) with $m = 0$ and $V(r) = \sigma r$. The momentum p is given in units of $\sqrt{\sigma}$.

Equation (34) gives the answer to the question of the Lorentz nature of confinement in the heavy-light quarkonium. Indeed, for low-lying states with the relative momentum p being small, the chiral angle φ_p is close to $\pi/2$, so that the matrix $U_p = \beta$, and this immediately leads one to Eq. (26). Notice that the contribution to $\Lambda(\vec{x}, \vec{y})$ proportional to the matrix γ_0 appeared entirely due to chiral symmetry breaking described in terms of the nontrivial chiral angle φ_p —see Fig. 1 ($\varphi_p \equiv 0$ for the massless quark and without chiral symmetry breaking²). This is the regime found in Ref. [7] and mentioned in the beginning of this section. In the next section we demonstrate that this regime exactly has to take place in order to allow one to describe the quarkonium using the Salpeter equation. The opposite situation of the vanishing chiral angle, which realizes for highly excited bound states, is discussed in detail in Ref. [23].

Obviously, inclusion of the spatial part of the kernel (16) as well as relaxing other simplifying assumptions made in the course of this section do not change the main conclusion of this section—namely that, as soon as chiral symmetry is broken spontaneously, an effective scalar interaction appears in a self-consistent manner.

IV. SCALAR CONFINEMENT AND THE SALPETER BOUND-STATE EQUATION

In the previous section we considered the approach to the heavy-light quark-antiquark system based on the Schwinger-Dyson-type equation (12). An alternative approach to (heavy-light) quarkonia, based on the spinless Salpeter equation,

$$[\sqrt{p^2 + m^2} + \sigma r] \psi = E \psi, \quad (35)$$

is also celebrated in the literature [see, for example, Ref. [24] and Appendix B for the derivation of Eq. (35) in the formalism of the QCD string with quarks at the ends [2], which is also based on the VCM [1]]. The purpose of the present section is to demonstrate that Eq. (35) and its more sophisticated versions, like the Hamiltonian of the QCD string with quarks at the ends (see Appendix B), are consistent with the bound-state equation (40) under the assumption of the scalar confinement dominance in the effective interquark interaction.

We return now to the full Eq. (23) and rewrite it in the form of the bound-state equation for the bispinor wave function Ψ ,

$$(\vec{\alpha} \cdot \hat{p} + \beta m) \Psi(\vec{x}) + \beta \int d^3z M(\vec{x}, \vec{z}) \Psi(\vec{z}) = E \Psi(\vec{x}), \quad (36)$$

²The situation is trivial for the heavy quark when chiral symmetry is broken explicitly. Indeed, in this case the chiral angle acquires the contribution $\arctan(m/p)$ and thus $\varphi_p \approx \pi/2$ for $p \ll m$.

with the nonlocal part of the kernel included together with the local one. Notice that the heavy-light bound-state equation in the form Eq. (36) can be derived independently in the formalism of the GNJL quark models [23].

Passing over to the momentum space and using the mass-gap equation in the form

$$E_p U_p = \tilde{\alpha} \vec{p} + \beta m + \frac{1}{2} \int \frac{d^3 k}{(2\pi)^3} V(\vec{p} - \vec{k}) U_k, \quad (37)$$

one can rewrite Eq. (36) as

$$\begin{aligned} E_p U_p \Psi(\vec{p}) + \frac{1}{2} \int \frac{d^3 k}{(2\pi)^3} V(\vec{p} - \vec{k}) [U_p + U_k] \Psi(\vec{k}) \\ = E \Psi(\vec{p}), \end{aligned} \quad (38)$$

where the quantity $E_p = A_p \sin \varphi_p + B_p \cos \varphi_p$ is the dressed quark dispersive law which substitutes the free dispersion $\sqrt{\vec{p}^2 + m^2}$ and which appears as a result of the quark self-interaction.

Equation (38) is subject to a Foldy-Wouthuysen transformation, which was built in a closed form in Ref. [23]. The corresponding operator is

$$T_p = \exp \left[-\frac{1}{2} (\hat{\gamma} \hat{p}) \left(\frac{\pi}{2} - \varphi_p \right) \right], \quad \Psi(\vec{p}) = T_p \begin{pmatrix} \psi(\vec{p}) \\ 0 \end{pmatrix}, \quad (39)$$

and the resulting Schrödingerlike equation, which stems from Eq. (38) after the Foldy-Wouthuysen transformation with the operator (39), reads

$$\begin{aligned} E_p \psi(\vec{p}) + \int \frac{d^3 k}{(2\pi)^3} V(\vec{p} - \vec{k}) [C_p C_k + (\vec{\sigma} \hat{p})(\vec{\sigma} \hat{k}) S_p S_k] \psi(\vec{k}) \\ = E \psi(\vec{p}), \end{aligned} \quad (40)$$

where we used the shorthand notations $C_p = \cos \frac{1}{2} (\frac{\pi}{2} - \varphi_p)$ and $S_p = \sin \frac{1}{2} (\frac{\pi}{2} - \varphi_p)$; $\vec{\sigma}$ are Pauli matrices, and \hat{p} and \hat{k} are unity vectors for \vec{p} and \vec{k} , respectively.

Let us consider Eq. (40) in the regime $\varphi_p \approx \frac{\pi}{2}$. Then $C_p = 1$, $S_p = 0$ and the interaction term in Eq. (40) reduces to the plain potential σr , in coordinate space. Then the resulting equation reads

$$[E_p + \sigma r] \psi = E \psi, \quad (41)$$

where E_p plays the role of the kinetic energy operator for the quark. This equation has the form of the Salpeter equation (35). Notice, however, that it is not always sufficient to keep the kinetic term for light quarks in the form of the free-particle energy $\sqrt{\vec{p}^2 + m^2}$, as in Eq. (35). Strictly speaking, the dispersive law of the light quark E_p is generated dynamically and has to be treated with care. This is especially important for small interquark momenta, where E_p can even become negative. As a result, the lowest states in the spectrum—the pions and the kaons—cannot

be described using the simple Eq. (35) and the like. These states are to be considered using the full Bethe-Salpeter equation with the two-component mesonic wave function [10,21] or in the framework of the full Schwinger-Dyson-type equation, similar to the heavy-light equation (12) [25,26]. Progress in adapting the Salpeter equation based approach to the description of lightest mesons was achieved in Ref. [27] in the framework of a matrix Hamiltonian technique. Apart from the aforementioned problem with the pions and kaons, the quark dispersive law E_p in Eq. (41) can be substituted, with a good accuracy, by the free-quark energy, so that the Salpeter equation (35) is readily reproduced.

In addition, as seen from Eq. (40), the simple Salpeter equation (35) fails for highly excited states as well. Indeed, the relative interquark momentum is large in excited mesons, so that the chiral angle vanishes asymptotically (see Fig. 1). As a result, the effective interaction in Eq. (40) becomes vectorial [see Eq. (40) with $C_p = S_p = 1/\sqrt{2}$] and it does not reduce to a plain potential anymore [23].

We conclude this section by stating that, contrary to naive expectations (potential is added to the energy), the form of the Salpeter equation (35) does not suggest that the effective interquark interaction in the meson is vectorial. Furthermore, we demonstrate that this equation arises naturally from the full Schwinger-Dyson-type equation for the heavy-light quarkonium under the assumption that chiral symmetry is broken (explicitly or spontaneously) and, as a result, the effective scalar interquark interaction is generated in a self-consistent manner.

V. CONCLUSIONS

In this paper we address the problem of the Lorentz nature of confinement in QCD. We consider a heavy-light quark-antiquark system as a testing ground and exploit the Schwinger-Dyson-type equation derived for the Green's function of such a system using the VCM. We demonstrate explicitly that the stringlike picture of the interquark interaction at large distances [in the form of the Salpeter equation (35)] appears due to chiral symmetry breaking. In particular, we prove that the Salpeter equation (35) appears self-consistently in the Schwinger-Dyson approach to the heavy-light quarkonium if chiral symmetry is broken, explicitly or spontaneously, and the effectively generated scalar potential dominates in the effective interquark interaction. This implies that the genuine Lorentz nature of the confining interaction in this Salpeter equation (as well as in the Hamiltonian of the QCD string with quarks at the ends) is scalar, which is the main result of this work. This solves the problem of the Klein paradox which is known to operate for systems with vectorial interaction growing with the distance. We conclude that there is no room for such a problem in QCD.

The reported result is robust since it is only based on a quite general consideration and is stable across the whole

variety of quark kernels. Furthermore, our conclusions acquire additional support from the fact that exactly the same bound-state equation for the heavy-light quarkonium can be derived independently in the framework of the GNJL quark models which have a long history in the literature and are known to give deep insight into physics of chiral symmetry breaking.

Finally, let us mention that for light quarks and without spontaneous breaking of chiral symmetry one would have a vanishing chiral angle and, consequently, no effective scalar interaction. This situation is believed to take place in QCD above the temperature of the chiral symmetry restoration transition or for highly excited hadrons (see Ref. [28] for a review). Properties of the interquark interaction in these situations deserve a special investigation and will be the subject of future publications (see, for example, Ref. [29]).

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APPENDIX A: DERIVATION OF EQ. (26)

Following the approach suggested in Ref. [7], we assume (and justify this assumption *a posteriori*) that the Schwinger-Dyson equation (23) possesses a solution which gives for the quark mass operator [see Eq. (24)] a form described by a local scalar $[U(\vec{x})]$ and a local vector $[V(\vec{x})]$ potential,

$$M(\vec{x}, \vec{y}) = [U(\vec{x}) + \gamma_0 V(\vec{x})] \delta^{(3)}(\vec{x} - \vec{y}). \quad (\text{A1})$$

Then one can rewrite Eq. (23) in the form of a Dirac equation for the wave function $\psi(\vec{x})$,

$$(\vec{\alpha} \vec{p} + \beta[m + U(\vec{x})] + V(\vec{x}))\psi(\vec{x}) = E\psi(\vec{x}) \quad (\text{A2})$$

or, in components,

$$\left(\psi = \frac{1}{r} \begin{pmatrix} G_n \Omega_{jlm} \\ iF_n \Omega_{j'l'm} \end{pmatrix} \right), \quad (\text{A3})$$

$$\begin{cases} \frac{dG_n}{dr} + \frac{\kappa}{r} G_n - (E_n + m + U - V)F_n = 0, \\ \frac{dF_n}{dr} - \frac{\kappa}{r} F_n + (E_n - m - U - V)G_n = 0. \end{cases}$$

One can use a simple trick to guess the matrix structure of the function $\Lambda(\vec{x}, \vec{y})$,

$$\Lambda(\vec{x}, \vec{y}) \equiv \sum_{n=-\infty}^{\infty} \psi_n(\vec{x}) \text{sign}(n) \psi_n^\dagger(\vec{y}), \quad (\text{A4})$$

built with the help of the solutions to Eq. (A2). Indeed,

according to its definition, $\Lambda(\vec{x}, \vec{y})$ can be naturally split into two parts,

$$\Lambda^{(V)}(\vec{x}, \vec{y}) = \Lambda_+^{(V)}(\vec{x}, \vec{y}) - \Lambda_-^{(V)}(\vec{x}, \vec{y}), \quad (\text{A5})$$

where \pm stand for the summation over positive and negative eigenvalues, respectively. Also, for future convenience, we used the superscript (V) . A similar decomposition is valid for the reversed sign of the vector interaction,

$$\Lambda^{(-V)}(\vec{x}, \vec{y}) = \Lambda_+^{(-V)}(\vec{x}, \vec{y}) - \Lambda_-^{(-V)}(\vec{x}, \vec{y}). \quad (\text{A6})$$

Now we notice the following symmetry inherent to the system (A3), $(V, E_n, \kappa, G, F) \leftrightarrow (-V, -E_n, -\kappa, F, G)$, and find

$$\Lambda_+^{(V)} \propto (GiF) \begin{pmatrix} G^* \\ -iF^* \end{pmatrix} = \begin{pmatrix} GG^* & \\ & FF^* \end{pmatrix}, \quad (\text{A7})$$

$$\Lambda_-^{(-V)} \propto \begin{pmatrix} FF^* & \\ & GG^* \end{pmatrix}.$$

Hence one can rewrite Eq. (A5) as

$$\Lambda^{(V)} = [\Lambda_+^{(V)} - \Lambda_-^{(-V)}] + [\Lambda_-^{(-V)} - \Lambda_+^{(V)}] \\ = \gamma_0 \sum_{E_n > 0} (GG^* - FF^*) + \delta\Lambda, \quad (\text{A8})$$

where the correction $\delta\Lambda$ vanishes for $V = 0$. Therefore, for a purely scalar confinement, the matrix structure of $\Lambda(\vec{x}, \vec{y})$ is, indeed, given by the matrix γ_0 . In order to establish its spatial structure one can use the WKB calculation performed in Ref. [7]. We omit the lengthy calculation which can be found in Ref. [7] and quote here the final result:

$$\Lambda(\vec{x}, \vec{y}) \approx \gamma_0 \frac{\sigma}{\pi^2 \sqrt{xy}} K_0(\sigma \sqrt{xy} |x - y|) \delta(1 - \cos\theta_{xy}), \quad (\text{A9})$$

where K_0 is the MacDonald function. It is easy to check that

$$\int d^3y \Lambda(\vec{x}, \vec{y}) = \gamma_0, \quad (\text{A10})$$

and, therefore, for $|\vec{x}|, |\vec{y}| \gg \frac{1}{\sigma|\vec{x}-\vec{y}|}$, $\Lambda(\vec{x}, \vec{y})$ can be approximated by the three-dimensional delta function peaked at $\vec{y} = \vec{x}$. Thus we arrive at Eq. (26). Moreover, for $V \neq 0$, the same WKB method reproduces Eq. (A9) and gives the decrease of the term $\delta\Lambda$ in Eq. (A8) at large distances, so that Eq. (26) holds [7].

APPENDIX B: ROTATING QCD STRING AND THE SPINLESS SALPETER EQUATION

In this appendix we give a brief derivation of Eq. (35) in the formalism of the QCD string with quarks at the ends which is also derived in the framework of VCM. Following

the method of Ref. [2], we start from the in- and out-states of the quark-antiquark meson,

$$\begin{aligned} \Psi_{q\bar{q}}^{(\text{in,out})}(x, y|A) &= \bar{\Psi}_{\bar{q}}(x)\Phi(x, y)\Psi_q(y), \\ \Phi(x, y) &= P \exp\left(ig \int_y^x dz_\mu A_\mu^a t^a\right), \end{aligned} \quad (\text{B1})$$

and build its Green's function,

$$\begin{aligned} G_{q\bar{q}} &= \langle \Psi_{q\bar{q}}^{(\text{out})}(\bar{x}, \bar{y}|A)\Psi_{q\bar{q}}^{(\text{in})}(x, y|A)^\dagger \rangle_{q\bar{q}A} \\ &= \langle \text{Tr} S_q(\bar{x}, x|A)\Phi(x, y)S_{\bar{q}}(y, \bar{y}|A)\Phi(\bar{y}, \bar{x}) \rangle_A, \end{aligned} \quad (\text{B2})$$

where S_q and $S_{\bar{q}}$ are the propagators of the quark and the antiquark, respectively, in the background gluonic field. Averaging over the background field is performed using the minimal area law assumption for the isolated Wilson loop,

$$\left\langle \text{Tr} P \exp\left(ig \oint_C dz_\mu A_\mu\right) \right\rangle_A \sim \exp(-\sigma S_{\text{min}}), \quad (\text{B3})$$

which is usually assumed for the stochastic QCD vacuum (see, for example, Ref. [1]) and is found on the lattice. Here S_{min} is the area of the minimal surface swept by the quark and antiquark trajectories,

$$S_{\text{min}} = \int_0^T dt \int_0^1 d\beta \sqrt{(\dot{w}w')^2 - \dot{w}^2 w'^2}, \quad (\text{B4})$$

where, for the profile function of the string $w_\mu(t, \beta)$, we adopt the straight-line ansatz:

$$w_\mu(t, \beta) = \beta x_{1\mu}(t) + (1 - \beta)x_{2\mu}, \quad (\text{B5})$$

$x_{1,2}(t)$ being the four-coordinates of the quarks at the ends of the string. We choose to consider the system in the laboratory frame and also to synchronize the quark times,

$$x_{10} = x_{20} = t. \quad (\text{B6})$$

The resulting Lagrangian of the string reads

$$L_{\text{str}} = -\sigma r \int_0^1 d\beta \sqrt{1 - [\vec{n} \times (\beta \dot{\vec{x}}_1 + (1 - \beta)\dot{\vec{x}}_2)]^2}, \quad (\text{B7})$$

where $\vec{r} = \vec{x}_1 - \vec{x}_2$, $\vec{n} = \vec{r}/r$. This interaction Lagrangian is to be supplied by the quark kinetic terms $-m_1\sqrt{1 - \dot{\vec{x}}_1^2} - m_2\sqrt{1 - \dot{\vec{x}}_2^2}$. Then, with the help of the auxiliary (einbein) field technique, used to get rid of the square roots in the kinetic terms (the einbeins $\mu_{1,2}$ [30]) and in the string term

[the continuous einbein $\nu(\beta)$ [2]] in the Lagrangian (B7), one can proceed to the Hamiltonian of the system (see Ref. [2] for the details of the derivation),

$$\begin{aligned} H &= \sum_{i=1}^2 \left[\frac{p_r^2 + m_i^2}{2\mu_i} + \frac{\mu_i}{2} \right] + \int_0^1 d\beta \left[\frac{\sigma^2 r^2}{2\nu} + \frac{\nu}{2} \right] \\ &\quad + \frac{\vec{L}^2}{2r^2[\mu_1(1 - \zeta)^2 + \mu_2\zeta^2 + \int_0^1 d\beta \nu(\beta - \zeta)^2]}, \\ \zeta &= \frac{\mu_1 + \int_0^1 d\beta \nu \beta}{\mu_1 + \mu_2 + \int_0^1 d\beta \nu}. \end{aligned} \quad (\text{B8})$$

Extrema in the einbein fields are understood either in the Hamiltonian (B8) or, alternatively, in its spectrum. In the latter case the einbein field method is a variety of the celebrated variational method in quantum mechanics.

Now, if the contribution of the string to the total inertia of the rotating system [denominator of the last, angular-momentum-dependent term in the Hamiltonian (B8)] is neglected, then the extrema in all einbeins can be taken analytically yielding for the Hamiltonian (this procedure is exact for $L = 0$)

$$H = \sqrt{\vec{p}^2 + m_1^2} + \sqrt{\vec{p}^2 + m_2^2} + \sigma r, \quad (\text{B9})$$

or, in the one-particle limit ($m_1 \equiv M \rightarrow \infty$, $m_2 \equiv m$),

$$H = \sqrt{\vec{p}^2 + m^2} + \sigma r, \quad (\text{B10})$$

where we omitted the infinite contribution of the static particle mass M . After a canonical quantization of the Hamiltonian (B10) we reproduce the spinless Salpeter equation (35). The proper string dynamics in Hamiltonian (B8) as well as spin-dependent terms are discussed in the literature—see, for example, Ref. [31] and Refs. [5,32], respectively. Discussion of the proper string dynamics in the formalism of the Schwinger-Dyson-type equation (12) can be found in Ref. [17]. Calculations of various hadronic spectra in the framework of the Hamiltonian (B8) supplied by the perturbative exchange and by the spin-dependent terms demonstrate a good accuracy of the predictions (see, for example, recent results for the spectrum of heavy-light D , D_s , B , and B_s mesons [33]). Notice that, in the case of light quarks, the major contribution to the spectrum of the Hamiltonian (B8) comes from the confining QCD string. Therefore, this case can be referred to as the case of the ‘‘heavy’’ string (as opposed to the case of heavy quarks when the proper string dynamics gives only small corrections).

- [1] H. G. Dosch, Phys. Lett. B **190**, 177 (1987); H. G. Dosch and Yu. A. Simonov, Phys. Lett. B **205**, 339 (1988); Yu. A. Simonov, Nucl. Phys. **B307**, 512 (1988).
- [2] A. Yu. Dubin, A. B. Kaidalov, and Yu. A. Simonov, Phys. Lett. B **323**, 41 (1994); Phys. Lett. B **343**, 310 (1995); E. L. Gubankova and A. Yu. Dubin, Phys. Lett. B **334**, 180 (1994).
- [3] E. Eichten and F. L. Feinberg, Phys. Rev. D **23**, 2724 (1981).
- [4] D. Gromes, Z. Phys. C **26**, 401 (1984); N. Brambilla, D. Gromes, and A. Vairo, Phys. Rev. D **64**, 076010 (2001).
- [5] Yu. A. Simonov, Nucl. Phys. **B324**, 67 (1989); Yad. Fiz. **66**, 363 (2003) [Phys. At. Nucl. **66**, 338 (2003)]; A. M. Badalian and Yu. A. Simonov, Yad. Fiz. **59**, 2247 (1996) [Phys. At. Nucl. **59**, 2164 (1996)].
- [6] M. Koma, Y. Koma, and H. Wittig, Nucl. Phys. **B769**, 79 (2007).
- [7] Yu. A. Simonov, Yad. Fiz. **60**, 2252 (1997) [Phys. At. Nucl. **60**, 2069 (1997)]; Phys. Rev. D **65**, 094018 (2002); Yu. A. Simonov and J. A. Tjon, Phys. Rev. D **62**, 014501 (2000).
- [8] N. Brambilla and A. Vairo, Phys. Lett. B **407**, 167 (1997).
- [9] Yu. S. Kalashnikova and A. V. Nefediev, Phys. Lett. B **414**, 149 (1997).
- [10] A. Amer, A. Le Yaouanc, L. Oliver, O. Pene, and J.-C. Raynal, Phys. Rev. Lett. **50**, 87 (1983); A. Le Yaouanc, L. Oliver, O. Pene, and J.-C. Raynal, Phys. Lett. **134B**, 249 (1984); Phys. Rev. D **29**, 1233 (1984); A. Le Yaouanc, L. Oliver, S. Ono, O. Pene, and J.-C. Raynal, Phys. Rev. D **31**, 137 (1985).
- [11] P. Bicudo and J. E. Ribeiro, Phys. Rev. D **42**, 1611 (1990); **42**, 1625 (1990); **42**, 1635 (1990); P. Bicudo, Phys. Rev. Lett. **72**, 1600 (1994); Phys. Rev. C **60**, 035209 (1999); P. J. A. Bicudo, A. V. Nefediev, and J. E. F. T. Ribeiro, Phys. Rev. D **65**, 085026 (2002).
- [12] A. V. Nefediev and Yu. A. Simonov, Pis'ma Zh. Eksp. Teor. Fiz. **82**, 633 (2005) [JETP Lett. **82**, 557 (2005)].
- [13] I. I. Balitsky, Nucl. Phys. **B254**, 166 (1985).
- [14] J. Ambjørn, P. Olesen, and C. Peterson, Nucl. Phys. **B240**, 189 (1984); **B240**, 533 (1984); N. A. Campbell, I. H. Jorjusz, and C. Michael, Phys. Lett. **167B**, 91 (1986).
- [15] G. S. Bali, Nucl. Phys. B, Proc. Suppl. **83**, 422 (2000); Phys. Rev. D **62**, 114503 (2000); V. I. Shevchenko and Yu. A. Simonov, Phys. Rev. Lett. **85**, 1811 (2000).
- [16] M. Campostrini, A. Di Giacomo, and G. Mussardo, Z. Phys. C **25**, 173 (1984); M. Campostrini, A. Di Giacomo, and S. Olejnik, Z. Phys. C **31**, 577 (1986); G. Bali, N. Brambilla, and A. Vairo, Phys. Lett. B **421**, 265 (1998).
- [17] A. V. Nefediev, Pis'ma Zh. Eksp. Teor. Fiz. **78**, 801 (2003) [JETP Lett. **78**, 349 (2003)].
- [18] V. I. Zakharov, AIP Conf. Proc. **756**, 182 (2005).
- [19] Y. Nambu and G. Jona-Lasinio, Phys. Rev. **122**, 345 (1961).
- [20] R. F. Wagenbrunn and L. Ya. Glozman, Phys. Rev. D **75**, 036007 (2007).
- [21] A. V. Nefediev and J. E. F. T. Ribeiro, Phys. Rev. D **70**, 094020 (2004).
- [22] P. J. A. Bicudo and A. V. Nefediev, Phys. Rev. D **68**, 065021 (2003).
- [23] Yu. S. Kalashnikova, A. V. Nefediev, and J. E. F. T. Ribeiro, Phys. Rev. D **72**, 034020 (2005).
- [24] T. J. Allen and M. G. Olsson, Phys. Rev. D **68**, 054022 (2003); T. J. Allen, M. G. Olsson, J. R. Schmidt, S. Veseli, and Yu Yuan, Phys. Rev. D **70**, 054012 (2004).
- [25] Yu. A. Simonov, Yad. Fiz. **67**, 868 (2004) [Phys. At. Nucl. **67**, 846 (2004)].
- [26] Yu. A. Simonov, Yad. Fiz. **67**, 868 (2004) [Phys. At. Nucl. **67**, 846 (2004)].
- [27] Yu. A. Simonov, Phys. At. Nucl. **68**, 709 (2005).
- [28] L. Ya. Glozman, Phys. Rep. **444**, 1 (2007).
- [29] A. V. Nefediev and Yu. A. Simonov, arXiv:hep-ph/0703306 [Phys. At. Nucl. (to be published)].
- [30] L. Brink, P. Di Vecchia, and P. Howe, Nucl. Phys. **B118**, 76 (1977); Yu. S. Kalashnikova and A. V. Nefediev, Yad. Fiz. **60**, 1529 (1997) [Phys. At. Nucl. **60**, 1389 (1997)].
- [31] V. L. Morgunov, A. V. Nefediev, and Yu. A. Simonov, Phys. Lett. B **459**, 653 (1999).
- [32] Yu. A. Simonov, in *Proceedings of the XVII International School of Physics "QCD: Perturbative or Nonperturbative," Lisbon, 1999*, edited by L. S. Ferreira, P. Nogueira, and J. I. Silva-Marcos (World Scientific, Singapore, 2000), p. 60.
- [33] Yu. S. Kalashnikova and A. V. Nefediev, Phys. Lett. B **492**, 91 (2000); **530**, 117 (2002); Yad. Fiz. (to be published); Yu. S. Kalashnikova, A. V. Nefediev, and Yu. A. Simonov, Phys. Rev. D **64**, 014037 (2001); A. M. Badalian, B. L. G. Bakker, and Yu. A. Simonov, Phys. Rev. D **75**, 116001 (2007); arXiv:hep-ph/0610193 [Phys. Rev. D (to be published)].