

Physics beyond the standard model: Focusing on the muon anomaly

Helder Chavez*

Instituto de Física, Universidade Federal do Rio de Janeiro, Caixa Postal 68528, 21945-910, Rio de Janeiro, Brazil

Cristine N. Ferreira†

*Núcleo de Física, Centro Federal de Educação Tecnológica de Campos,
Rua Dr. Siqueira, 273-Parque Dom Bosco, 28030-130, Campos dos Goytacazes, RJ, Brazil*

José A. Helayel-Neto‡

*Centro Brasileiro de Pesquisas Físicas,
Rua Dr. Xavier Sigaud 150, Urca 22290-180, Rio de Janeiro, RJ, Brazil
(Received 14 May 2006; published 7 August 2006)*

We present a model based on the implication of an exceptional E_6 -GUT symmetry for the anomalous magnetic moment of the muon. We follow a particular chain of breakings with Higgses in the **78** and **351** representations. We analyze the radiative correction contributions to the muon mass and the effects of the breaking of the so-called Weinberg symmetry. We also estimate the range of values of the parameters of our model.

DOI: [10.1103/PhysRevD.74.033006](https://doi.org/10.1103/PhysRevD.74.033006)

PACS numbers: 14.80.Bn, 12.10.Dm

I. INTRODUCTION

Among the known leptons, the muon is potentially interesting for several reasons. First, its relatively long lifetime of $2.2 \mu\text{s}$ ($c\tau = 658.65 \text{ m}$) makes it possible to perform precision measurements. Second, it is sensitive to new sectors of heavy particles and new interactions. In this sense, the muon anomaly has provided a stringent test for new theories of particle physics, since any new field or particle which couples to the muon must contribute to a_μ .

The most recent results reported by the Muon ($g - 2$) Collaboration [1] have triggered a renewal of interest on the theoretical prediction of the anomalous magnetic moment of the muon (commonly referred to as the muon anomaly), $a_\mu = \frac{g-2}{2}$, in the standard model (SM). This experimental value is claimed to show that there remains a discrepancy with the SM theoretical calculations at the confidence level of 2.3σ to 3.3σ [1,2], if the hadronic light-by-light contribution, $a_\mu^{\text{HLO}}(\text{LBL}) = 80(40) \times 10^{-11}$ [3], is used instead of $a_\mu^{\text{HLO}}(\text{LBL}) = 136(25) \times 10^{-11}$ [4], as a consequence that e^+e^- annihilation data are used to evaluate this contribution against hadronic τ decays data [5]. Among all contributions that yield corrections to the muon anomaly, the hadronic contributions are less accurate, due to the hadronic vacuum polarization effects in the diagrams which use data inputs coming from the e^+e^- annihilation cross section and the hadronic τ - decays. Also it is not clear, at present, whether the value from τ - decay data can be improved much further, due to the difficulty in evaluating more precisely the effect of isospin breaking [5].

In fact, these measurements have provided the highest accuracy of the validity of the different theories for strong, weak, and electromagnetic interactions because they have reached a fabulous relative precision of 0.5 parts per million (ppm) in the determination of a_μ . However, if this confidence level for the muon anomaly remains, it is possible that we are under a window open for a new physics at a high energy scale, Λ . The study of the muon anomaly becomes relevant because it is more sensitive to interactions that are not predicted in the SM but will be possibly reached at the CERN Large Hadron Collider (LHC), with $\sqrt{s} = 14 \text{ TeV}$.

On the theoretical side, if we take into account the effects of virtual massive particles in the diagrams contributing to the lepton anomaly, the ratios between the corrections to the anomalies are of the order $(\frac{m_\mu}{m_e})^2 \sim 4 \times 10^4$ for the muon and electron, and of the order $(\frac{m_\tau}{m_e})^2 \sim 1.2 \times 10^7$ for the tau and electron. The same huge enhancement factor would also affect the contributions coming from degrees of freedom beyond the SM, so that the measurement of the τ - anomaly would represent the best opportunity to detect new physics. Unfortunately, the very short lifetime of the τ - lepton which, precisely because of its high mass, can also decay into hadronic states, makes such a measurement impossible at present; this is the reason why there is an emphasis on the muon anomaly.

In this case, it becomes interesting to estimate the order of the correction of a_μ in the context of theories beyond the SM. This is done in terms of powers of $\frac{m_\mu}{\Lambda}$. This is related [6] to the validity or the breaking of the chiral symmetry for leptons together with the change of sign for m_μ . If this symmetry, which is referred to as Weinberg Symmetry (WS), is respected, then $\Delta a_\mu \sim (m_\mu/\Lambda)^2$; on the other hand, if it is broken, $\Delta a_\mu \sim m_\mu/\Lambda$. This is important

*Electronic address: helderch@if.ufrj.br†Electronic address: crisnfer@cetecampus.br‡Electronic address: helayel@cbpf.br

because in the latter case the explanation of the muon anomaly may be given by a new physics at a relatively high energy, whereas in the former it should appear at a scale close to the electroweak (EW) one.

We consider the 78 and 351 Higgs representations of the E_6 grand-unified theory (GUT). The representations between square brackets refer to the E_6 -group, those between brackets refer to $SO(10) \otimes \bar{U}(1)$, and the ones between parentheses correspond to the $SU(5) \otimes \tilde{U}(1)$ group. The symmetry breaking pattern [7–10] is depicted below.

$$\begin{array}{c}
 E_6 \\
 [78]\{1, 0\} \\
 \downarrow \\
 SO(10) \otimes \bar{U}(1) \\
 [351]\{1, -8\} \\
 \downarrow \\
 SO(10) \\
 [78]\{45, 0\}(1, 0) \\
 \downarrow \\
 SU(5) \otimes \tilde{U}(1) \\
 [351]\{16, -5\}(1, -5) \\
 \downarrow \\
 SU(5) \\
 [351]\{54, 4\}(24, 0), \\
 [351]\{144, 1\}(24, 5) \\
 \downarrow \\
 SU(3)_C \otimes SU(2)_L \otimes U(1) \\
 [351]\{10, -2\} \\
 \downarrow \\
 SU(3)_C \otimes U(1)_{e.m.}
 \end{array} \quad (1)$$

The order of magnitude of the contribution is $\Delta a_\mu \sim m_\mu/m_M$, where m_M is the mass of the exotic fermion. This fermion is analogous to the ordinary muon contained in the [27] representation of fermions in $\{10, -2\}$ under $SO(10) \times \bar{U}(1)$. This connection makes sense if the radiative correction to the muon mass is small and if there occurs breaking of WS. On the other hand, if the muon mass is only due to radiative corrections, the right mixing angle between leptons is zero and WS is not broken.

Our paper is organized as follows. In the Sec. II, we discuss the WS in the SM in connection with the order of magnitude of the muon anomaly. In Sec. III, we present our model, considering the sequences of breakings of symmetries (1). In Sec. IV, we analyze the question of the radiative mass of the muon due to the mixings with the massive fermion that occur in the breaking chain $SU(5) \rightarrow SU(3)_C \otimes SU(2)_L \otimes U(1)$ with $\{144, 1\}$ Higgs; in Sec. V, we analyze WS in the context of our model and, finally, in Sec. VI, we present our general conclusions.

II. WS AND THE ANOMALOUS MAGNETIC MOMENT IN THE SM

The WS is a well-known property [6] of the SM of particle physics. In this section, we briefly review its main points, since this result is connected with the order of magnitude of the Δa_μ contribution in the E_6 model. The mass term $m_\mu \bar{\mu} \mu$ breaks chiral symmetry; the field redefinition below changes the sign of the mass term:

$$\mu \rightarrow \gamma_5 \mu, \quad m_\mu \rightarrow -m_\mu, \quad (2)$$

where μ is the field variable associated to the muon.

If the WS Eq. (2) is valid, the corrections to a_μ must be of even powers of the ratio of m_μ to a larger scale Λ :

$$a_\mu = c_0 \left(\frac{m_\mu}{\Lambda} \right)^0 + c_2 \left(\frac{m_\mu}{\Lambda} \right)^2 + \dots \quad (3)$$

The effective interaction that gives a nonzero contribution to the muon anomalous magnetic moment is $a_\mu \frac{e}{4m_\mu} \bar{\mu} \sigma_{\alpha\beta} \mu F^{\alpha\beta}$; for the SM version, it may be written as

$$\mathcal{L}_{\text{eff}} = a_\mu \frac{e}{4m_\mu} \left(\bar{\Psi} \sigma^{\alpha\beta} \mu_R \frac{f_0 \varphi_V}{m_\mu} + \text{H.c.} \right) F_{\alpha\beta}, \quad (4)$$

with a Higgs field doublet

$$\varphi = \begin{pmatrix} 0 \\ \varphi_1 \end{pmatrix} = \varphi_V + \begin{pmatrix} 0 \\ h_1/\sqrt{2} \end{pmatrix}$$

such that

$$\bar{\Psi}_L = \begin{pmatrix} \nu \\ \mu \end{pmatrix}_L, \quad \varphi_V = \begin{pmatrix} 0 \\ \frac{v_1}{\sqrt{2}} \end{pmatrix}, \quad f_0 \frac{v_1}{\sqrt{2}} = m_\mu. \quad (5)$$

Now, to have the WS invariance (2) in the SM, one must perform the transformations

$$\Psi_L \rightarrow \gamma_5 \Psi_L = -\Psi_L, \quad \mu_R \rightarrow \gamma_5 \mu_R = \mu_R, \quad \varphi \rightarrow -\varphi. \quad (6)$$

We can prove that the neutral current Lagrangian density reads as

$$\mathcal{L}_{\text{NC}} = -e \bar{\mu} \gamma^\alpha \mu A_\alpha - \frac{g}{2 \cos \theta_W} \bar{\mu} \gamma^\alpha (v_z - a_z \gamma^5) \mu Z_\alpha; \quad (7)$$

the charged current Lagrangian density is written as

$$\mathcal{L}_{\text{CC}} = \frac{g}{2\sqrt{2}} [\bar{\nu}_\mu \gamma^\alpha (1 - \gamma^5) \mu W_\alpha^{(+)} + \bar{\mu} \gamma^\alpha (1 - \gamma^5) \nu_\mu W_\alpha^{(-)}], \quad (8)$$

and the Yukawa sector

$$\begin{aligned} \mathcal{L}_{\text{Yuk}} &= -f_0 (\bar{\mu}_R \varphi^\dagger \mu_L + \bar{\mu}_L \varphi \mu_R) \\ &= -\frac{1}{\sqrt{2}} f_0 (v_1 + h_1) \bar{\mu} \mu, \end{aligned} \quad (9)$$

where $m_\mu = f_0 \frac{v_1}{\sqrt{2}}$ is the muon mass and the interactions are invariant under the transformations of Eq. (6). Therefore, the corrections to a_μ are of the type of Eq. (3) with the EW scale, Λ . The first term is the electromagnetic contribution $c_0 = \frac{\alpha}{2\pi} + \dots$, computed recently up to $(\alpha/\pi)^5$ [11]; the second term, $c_2(\frac{m_\mu}{\Lambda})^2 \sim a_\mu^{\text{QED}} \times 1$, $7 \times 10^{-6} \simeq 2 \times 10^{-9}$, corresponds to the weak contribution.

III. AN ALTERNATIVE E_6 -MODEL FOR THE MUON ANOMALY

The exceptional group E_6 [12] was proposed as an alternative to SU(5) and SO(10) models, and it is actually, in many aspects, the preferred gauge group for grand unification. In this section, let us discuss the pattern of breakings (1) based on the [78] and [351] representations. The ordinary fermions of the SM are contained in the $\{16, 1\} \subset 27$ -dimensional representation:

$$[27] = \{16, 1\} \oplus \{10, -2\} \oplus \{1, 4\}. \quad (10)$$

There are 11 additional fermions with respect to the SM fermions. For the first generation, these particles are

$$\underbrace{\Psi_L}_{\{1,4\}} \oplus \underbrace{(\mathbf{D}^C \mathbf{N} \mathbf{E})_L}_{\{5,-2\}} \oplus \underbrace{(\mathbf{D} \mathbf{N}^C \mathbf{E}^C)_L}_{\{5,2\}}. \quad (11)$$

$\underbrace{\hspace{10em}}_{\{10,-2\}}$

The gauge bosons are contained in the adjoint 78-dimensional representation, that, with respect to $\text{SO}(10) \otimes \bar{\text{U}}(1)$, is decomposed as below:

$$[78] = \{45, 0\} \oplus \{16, -3\} \oplus \{1, 0\} \oplus \{\bar{16}, 3\}. \quad (12)$$

For the first generation, the exotic fermions of the 10 representation of SO(10) can acquire mass from the Higgs $\{54, 4\}$ of the [351] representation of E_6 , because $\{10\} \otimes \{10\} = \{54\} \oplus \{45\} \oplus \{1\}$. The mass terms are of the type [13]

$$\varphi_2(\mathbf{54}, \mathbf{24}) \left(D^c D - \frac{3}{2} E^c E - \frac{3}{2} N^c N \right). \quad (13)$$

In this same representation, $\{144, 1\}$, let us mix these fermions with the ordinary ones, because both components contain a 24 of SU(5), which has one invariant component under $\text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_Y$: $\{16\} \otimes \{10\} = \{144\} \oplus \{\bar{16}\}$. This mixing term is given by

$$\varphi_3(\mathbf{144}, \mathbf{24}) \left(d^c D - \frac{3}{2} E^c e - \frac{3}{2} N^c \nu \right). \quad (14)$$

Observe that in both Higgs, $\varphi_2(\mathbf{54}, \mathbf{24})$ and $\varphi_3(\mathbf{144}, \mathbf{24})$, being singlets $(1, 1, 0)$ under $\text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_Y$, we shall assume, take different values of expectation around his quantum fields h_2 and h_3 :

$$\varphi_2(\mathbf{54}, \mathbf{24}) = \frac{1}{\sqrt{2}}(v_2 + h_2) \quad (15)$$

$$\varphi_3(\mathbf{144}, \mathbf{24}) = \frac{1}{\sqrt{2}}(v_3 + h_3), \quad (16)$$

where the vacuum expectation values v_3 and v_2 we will assume satisfy the relation $v_3 \leq v_2$.

On the other hand, the ordinary fermions of the SM get masses from the Higgs $\{10, -2\}$, because the Yukawa term that conserves the $\bar{\text{U}}(1)$ charge is

$$\{16\} \otimes \{16\} = \{10\} \oplus \{126\} \oplus \{120\}, \quad (17)$$

and this Higgs is in the [351]. This mass term is

$$H(\mathbf{10}, \bar{\mathbf{5}})(d^c d + e^c e + N^c L). \quad (18)$$

In order to explain the notation, here $\varphi'(\mathbf{a}, \mathbf{24})$ stands for the component of the Higgs representation, φ' , where the label \mathbf{a} indicates the transformation under SO(10) and the label $\mathbf{24}$ -component refers to SU(5); similarly, for $H(\mathbf{10}, \bar{\mathbf{5}})$ In fact, this Higgs $H(\mathbf{10}, \bar{\mathbf{5}})$ is indeed that one of the SM $\varphi_1(\mathbf{1}, \mathbf{2}, 1/2)$ under $\text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_Y$ which is, as we already said before, written as

$$\varphi_1(\mathbf{1}, \mathbf{2}, 1/2) = \frac{1}{\sqrt{2}}(v_1 + h_1). \quad (19)$$

Now, let us extend this for the second generation of fermions, and call M the supermassive fermion in analogy to the ordinary muon of the SM.

If the breakings of symmetry are due to a [351], when the GUT symmetry is broken, the mass eigenstates μ_o and \hat{M} are determined by the expectation values of the (SO(10), SU(5)) multiplets $\varphi_2(\mathbf{54}, \mathbf{24})$ and $\varphi_3(\mathbf{144}, \mathbf{24})$, through the mixture of left and right components [13,14]:

$$\begin{pmatrix} \mu_{L,R} \\ M_{L,R} \end{pmatrix} = \begin{pmatrix} \cos\theta_{L,R} & \sin\theta_{L,R} \\ -\sin\theta_{L,R} & \cos\theta_{L,R} \end{pmatrix} \begin{pmatrix} \mu_{L,R}^0 \\ \hat{M}_{L,R} \end{pmatrix}, \quad (20)$$

where $\theta_{L,R}$ are the left and right mixing angles, respectively. It is possible that the mixing angle θ_R is small, of the order $\sim m_\mu/m_M$, where m_M is the mass of the heavy muon, M , however, due to the weak universality, the angle θ_L between μ_L and M_L is expected to be the same mixing angle for ν^μ and the neutral exotic lepton N ; but θ_L can still be large [15].

The fermion-Higgs interaction Lagrangian is given by

$$L = \frac{f_0}{\sqrt{2}} \bar{\mu}_L \mu_R (h_1 + v_1) + \frac{f_1}{\sqrt{2}} \bar{M}_L M_R (h_2 + v_2) + \frac{f_2}{\sqrt{2}} \bar{\mu}_L M_R (h_3 + v_3) + \frac{f_3}{\sqrt{2}} \bar{M}_L \mu_R (h_1 + v_1) + \text{H.c.}, \quad (21)$$

where some of the f_i 's could be vanishing. The previous expression can be written as below:

$$L = (\bar{\mu}_L \quad \bar{M}_L) \frac{1}{\sqrt{2}} \begin{pmatrix} f_0 v_1 & f_2 v_3 \\ f_3 v_1 & f_1 v_2 \end{pmatrix} \begin{pmatrix} \mu_R \\ M_R \end{pmatrix}. \quad (22)$$

The mass matrix reads as

$$\mathbf{M} = \frac{1}{\sqrt{2}} \begin{pmatrix} f_0 v_1 & f_2 v_3 \\ f_3 v_1 & f_1 v_2 \end{pmatrix}. \quad (23)$$

As usual, the previous matrix mass is diagonalized by a biunitary transformation [14,16] $U_L^\dagger \mathbf{M} U_R = \mathbf{M}_{\text{diag}}$, where $U_{L,R}$ is given in (20). From $U_L^\dagger \mathbf{M} \mathbf{M}^\dagger U_L = \mathbf{M}_{\text{diag}}^2$, it is possible to find

$$\tan(2\theta_L) = \frac{2(f_0 f_3 v_1^2 + f_1 f_2 v_2 v_3)}{(f_3^2 - f_0^2) v_1^2 + f_1^2 v_2^2 - f_2^2 v_3^2}; \quad (24)$$

on the other hand, from $U_R^\dagger \mathbf{M}^\dagger \mathbf{M} U_R = \mathbf{M}_{\text{diag}}^2$, we obtain

$$\tan(2\theta_R) = \frac{2(f_0 f_2 v_1 v_3 + f_1 f_3 v_1 v_2)}{f_1^2 v_2^2 + f_2^2 v_3^2 - (f_3^2 + f_0^2) v_1^2}. \quad (25)$$

In the limit for which all the couplings f_i are equal and $v_3 \simeq v_2 \gg v_1$, we have to $\tan(2\theta_L) \rightarrow \infty$,

$$\tan(2\theta_R) \simeq \frac{2v_1 v_2}{v_2^2 - v_1^2} \quad (26)$$

or to the angles θ_L and θ_R the values $\theta_L = \frac{\pi}{4}$, $\theta_R \sim \frac{v_1}{v_2}$. As it can be seen, in this case θ_R is small.

The part of the interaction Lagrangian for the quantum fluctuations can be written as

$$\begin{aligned} L = & \frac{f_0}{\sqrt{2}} \bar{\mu}_L \mu_R h_1 + \frac{f_1}{\sqrt{2}} \bar{M}_L M_R h_2 + \frac{f_2}{\sqrt{2}} \bar{\mu}_L M_R h_3 \\ & + \frac{f_3}{\sqrt{2}} \bar{M}_L \mu_R h_1 + \text{H.c.}; \end{aligned} \quad (27)$$

after the mixing equations (20), we obtain the changing-flavor Lagrangian:

$$\begin{aligned} L_{\text{FCh}} = & \frac{f_0}{\sqrt{2}} (c_{LSR} \bar{\mu}_L^0 \hat{M}_R + c_{RSL} \bar{\tilde{M}}_L \mu_R^0) h_1 \\ & + \frac{f_1}{\sqrt{2}} (-s_{LCR} \bar{\mu}_L^0 \hat{M}_R - c_{LSR} \bar{\tilde{M}}_L \mu_R^0) h_2 \\ & + \frac{f_2}{\sqrt{2}} (c_{RCL} \bar{\mu}_L^0 \hat{M}_R - s_{LSR} \bar{\tilde{M}}_L \mu_R^0) h_3 \\ & + \frac{f_3}{\sqrt{2}} (-s_{LSR} \bar{\mu}_L^0 \hat{M}_R + c_{LCR} \bar{\tilde{M}}_L \mu_R^0) h_1 + \text{H.c.}, \end{aligned} \quad (28)$$

$$\Delta a_\mu^{\text{FCh}} = \frac{1}{128\pi^2} m_\mu^2 \sum_{i=1}^3 \xi_i^2 \int_0^1 dx \frac{(x^2 - x^3) + \frac{m_M}{m_\mu} x^2}{m_\mu^2 x^2 + (m_M^2 - m_\mu^2)x + M_i^2(1-x)} = \frac{1}{128\pi^2} \frac{m_\mu}{m_M} \sum_{i=1}^3 \xi_i^2 G(z_i), \quad (32)$$

denoting $z_i = M_i^2/m_M^2$, where $\xi_i^2 = \alpha_i^2 - \beta_i^2$ with $\alpha_i = \tilde{\beta}_i + \tilde{\alpha}_i$, $\beta_i = \tilde{\beta}_i - \tilde{\alpha}_i$, the function $G(z_i) = [(1-z_i)^2 - 2z_i(1-z_i) - 2z_i^2 \ln z_i]/[2(1-z_i)^3]$ is plotted in Fig. 2. Let us see two cases of interest:

- (a) $m_M \simeq M_2 \simeq M_3 \gg M_1$. If we consider the rough case in that $f_1 = f_2$, we have $\xi_3^2 = \xi_2^2$ with the reasonable value $G(z_{2,3}) \simeq 0.3$ and $G(z_1) \simeq 0.5$. In this case the total contribution is

where $c_{L,R} = \cos\theta_{L,R}$ and $s_{L,R} = \sin\theta_{L,R}$. We label the neutral mass eigenstates of the Higgses by H_1, H_2, H_3 whose masses are M_1, M_2, M_3 , respectively. Then, suitable rotations between fields h_1, h_2, h_3 must diagonalize the mass matrix in the potential $V(h_1, h_2, h_3)$. We suppose (from now on) that $M_1 \ll M_3 \simeq M_2$, assuming conservation of CP , the matrix of rotations will be real. In the limit $v_1 \ll v_3 \leq v_2$, the state $h_1 \rightarrow H_1$ is weak and any appreciable mixing between scalars will only appear between h_2 and h_3 :

$$\begin{pmatrix} h_2 \\ h_3 \end{pmatrix} = \begin{pmatrix} \cos\vartheta & -\sin\vartheta \\ \sin\vartheta & \cos\vartheta \end{pmatrix} \begin{pmatrix} H_2 \\ H_3 \end{pmatrix}, \quad (29)$$

with ϑ being the angle of rotation that allows the diagonalization of the matrix. With these mixings of neutral scalars fields, the flavor-changing Lagrangian (28) now takes the form:

$$L_{\text{FCh}}^{\text{eff}} = \frac{1}{2\sqrt{2}} \sum_{i=1}^3 \bar{\mu}^0 [\tilde{\beta}_i + \tilde{\alpha}_i - \gamma_5(\tilde{\beta}_i - \tilde{\alpha}_i)] \hat{M} H_i + \text{H.c.}, \quad (30)$$

where

$$\begin{aligned} \tilde{\alpha}_1 &= f_0 c_{LSR} - f_3 s_{LSR}, & \tilde{\beta}_1 &= f_0 c_{RSL} + f_3 c_{LCR}, \\ \tilde{\alpha}_2 &= -f_1 c_{RSL} \cos\vartheta + f_2 c_{RCL} \sin\vartheta \\ \tilde{\beta}_2 &= -f_1 c_{LSR} \cos\vartheta - f_2 s_{LSR} \sin\vartheta, \\ \tilde{\alpha}_3 &= f_1 c_{RSL} \sin\vartheta + f_2 c_{LCR} \cos\vartheta, \\ \tilde{\beta}_3 &= f_1 c_{LSR} \sin\vartheta - f_2 s_{LSR} \cos\vartheta. \end{aligned} \quad (31)$$

The generic diagram with Higgs interchange contributing to the anomaly of the muon is shown in Fig. 1. In fact, the explicit calculation [17] of the one-loop contribution yielded by Eq. (29) gives the results (in the limit $m_M/m_\mu \gg 1$):

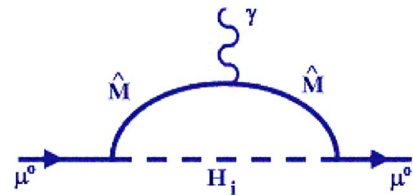


FIG. 1 (color online). Contributions with Higgs interchange to the muon anomalous magnetic moment.

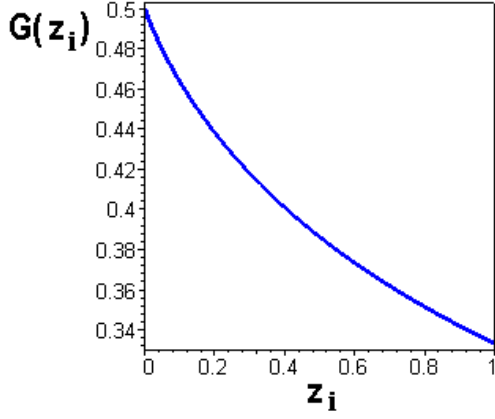


FIG. 2 (color online). Plot of $G(z_i)$ as a function of z_i , where $z_i = M_i^2/m_M^2$.

$$\Delta a_\mu^{\text{FCh}} = \frac{1}{128\pi^2} \frac{m_\mu}{m_M} (0.5 \times \xi_1^2 + 0.6 \times \xi_2^2), \quad (33)$$

then, to complete the anomaly value [2] $\Delta a_\mu = 25.2 \times 10^{-10}$, we have

$$7.4 \times 10^{-3} \leq \xi_1^2 + 1.2 \times \xi_2^2 \leq 0.64, \quad (34)$$

where we consider $115 \text{ GeV} \leq m_M \leq 10 \text{ TeV}$.

- (b) $M_2 \simeq M_3 \gg M_1 \simeq m_M$. The principal contribution comes from H_1

$$\Delta a_\mu^{\text{FCh}}(H_1) \simeq \frac{1}{128\pi^2} \frac{m_\mu}{m_M} \xi_1^2 G(z_1), \quad (35)$$

and in this case $G(z_{2,3}) \rightarrow 0$. We can find the limits of ξ_1^2 over the range masses indicated $7 \times 10^{-3} \leq \xi_1^2 \leq 0.61$, as illustrated in Fig. 3.

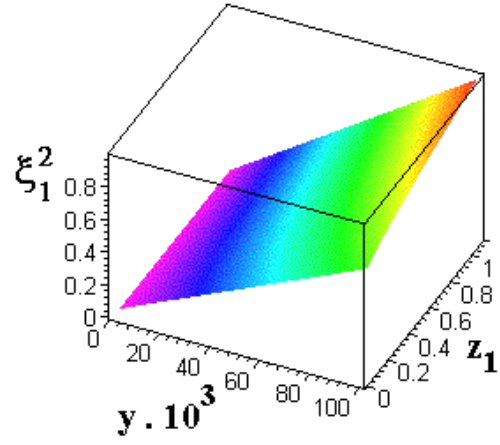


FIG. 3 (color online). Space of values of ξ_1^2 in the range of masses $115 \text{ GeV} \leq m_M \leq 10 \text{ TeV}$, $115 \text{ GeV} \leq M_1 \leq 700 \text{ GeV}$ to explain the anomaly value, where $y = \frac{m_\mu}{m_M}$.

IV. RADIATIVE CORRECTIONS TO THE MUON MASS

Another interesting possibility is to suppose a situation in which the muon mass comes only from radiative corrections. There are models of this type [18,19] in the literature. In Ref. [19], the authors, working out an $SU(3) \otimes SU(3) \otimes U(1)$ model, introduce some symmetries to avoid the light fermions from acquiring their masses at tree level through their couplings to the SM Higgs boson with non-zero vacuum expectation value; as a consequence, the muon gets its mass from the radiative corrections induced by other particles.

The one-loop correction to the muon mass is obtained by removing the photon line from the diagram Fig. 1. The amplitude for this diagram is

$$\begin{aligned} \Sigma(p) = & -i\kappa^2 \left[\int \frac{d^4 q}{(2\pi)^4} \frac{m_M (|\alpha_l|^2 - |\beta_l|^2) + (|\alpha_l|^2 + |\beta_l|^2) q_\mu \gamma^\mu}{(q^2 - m_M^2)((p-q)^2 - M_l^2)} + \int \frac{d^4 q}{(2\pi)^4} \right. \\ & \left. \times \frac{(\alpha_l \beta_l^\dagger + \beta_l \alpha_l^\dagger) q_\mu \gamma^\mu \gamma_5 + m_M (\alpha_l \beta_l^\dagger - \beta_l \alpha_l^\dagger) \gamma_5}{(q^2 - m_M^2)((p-q)^2 - M_l^2)} \right], \quad (36) \end{aligned}$$

where $\kappa = \frac{1}{2\sqrt{2}}$ and $l = 1, 2, 3$. Let us suppose that M_2 is the maximal energy scale for our model, then, as $m_\mu \ll m_M, M_2$, we obtain the following expression for the radiatively induced muon mass:

$$m_\mu^{1\text{-loop}} = \frac{(\alpha_2^2 - \beta_2^2)}{8(4\pi)^2} m_M F(z_2), \quad (37)$$

$$F(z_2) = 1 - \frac{1}{z_2 - 1} \ln z_2, \quad (38)$$

where $z_2 = M_2^2/m_M^2$. Notice that, for $M_2 \simeq M_3 \gg m_M, M_1$ (or $z_{2,3} \gg 1$), the function $F(z_2)$ takes its asymptotic value equal to 1, then

$$m_\mu^{\text{loop}}(H_2, H_3) \simeq \frac{\xi_2^2}{128\pi^2} m_M, \quad (39)$$

and for the case $M_2 \simeq M_3 \simeq m_M \gg M_1$ the function $F(z_2) \simeq 1 - m_M^2/M_2^2$. To assure small radiative mass for the muon, for example, of 0.1–10 MeV with $115 \text{ GeV} \leq m_M \leq 10 \text{ TeV}$, it is necessary that $1.0 \times 10^{-3} \leq \xi_2^2 \leq 1.3 \times 10^{-3}$.

There is another diagram that can contribute to the radiative mass of the muon, as shown in Fig. 4. The result was estimated [19,20] as

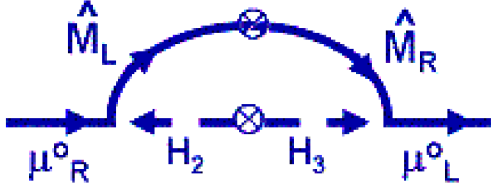


FIG. 4 (color online). Diagram of radiative correction to muon mass with mixing between heavy scalars.

$$m_\mu^{1\text{-loop}} \simeq \frac{\varepsilon}{16\pi^2} m_M \left[\frac{M_2^2}{m_M^2 - M_2^2} \ln\left(\frac{m_M^2}{M_2^2}\right) - \frac{M_3^2}{m_M^2 - M_3^2} \ln\left(\frac{m_M^2}{M_3^2}\right) \right], \quad (40)$$

where ε is a parameter function of Yukawa couplings [that can read from (29) and (30)] and of the mixing angle ϑ . However, the contribution of \hat{M} , H_2 , and H_3 given by Fig. 4 for the limit natural $m_M \ll M_2 \simeq M_3$, $m_\mu^{1\text{-loop}}$ is essentially zero.

In our model, the ordinary fermions are massless at the tree level in the GUT scale (i.e. no bare m_μ^0 is possible due to symmetry), but it couples to the heavy fermion \hat{M} through the mixing with scalars, according to the breaking $SU(5) \otimes \tilde{U}(1) \rightarrow SU(5)$. If we suppose this, then the only diagrams that contribute to the anomaly are those with the interchange of H_2 and H_3 in Fig. 1. To simplify, let us suppose the case $M_2 \simeq M_3$ and $f_1 = f_2$ from which $\xi_2^2 = \xi_3^2$; then, the contribution with the H_2 -interchange is

$$\Delta a_\mu^{\text{Fch}} = \frac{\xi_2^2}{128\pi^2} \frac{m_\mu}{m_M} G(z_2); \quad (41)$$

but, from (36) and (37), we can write for $M_2 \gg m_M$

$$m_\mu^{1\text{-loop}} \simeq \frac{\xi_2^2}{128\pi^2} m_M F(z_2); \quad (42)$$

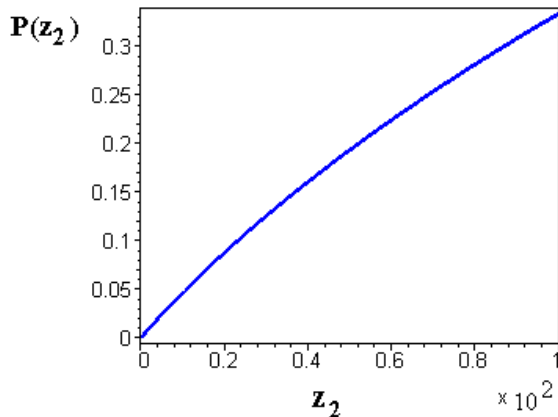


FIG. 5 (color online). Plot of $P(z_2)$. Note that $P(z_2)$ is roughly $\mathcal{O}(1)$ on the values range considered.

combining these equations, we obtain

$$\Delta a_\mu^{\text{Fch}} \simeq \frac{m_\mu^2}{M_2^2} \frac{z_2[(1-z_2)^2 - 2z_2(1-z_2) - 2z_2^2 \ln z_2]}{2(1-z_2)^3}, \quad (43)$$

where the function $P(z_2) = z_2[(1-z_2)^2 - 2z_2(1-z_2) - 2z_2^2 \ln z_2]/2(1-z_2)^3$ is plotted in Fig. 5. In this way, if the mass of the muon is of radiative origin we obtain $\Delta a_\mu^{\text{Fch}} \sim m_\mu^2/M_2^2$. An analogous result was obtained by Marciano using a toy model [21].

V. WEINBERG SYMMETRY INVARIANCE

In terms of the mixing angles $\theta_{L,R}$, from the biunitary diagonalization $U_L^\dagger \mathbf{M} U_R = \mathbf{M}_{\text{diag}}$, we find for the masses

$$m_\mu = \frac{1}{\sqrt{2}} [(c_L f_0 - s_L f_3) v_1 c_R - (c_L f_2 v_3 - s_L f_1 v_2) s_R], \quad (44)$$

$$m_M = \frac{1}{\sqrt{2}} [(s_L f_0 + c_L f_3) v_1 s_R + (s_L f_2 v_3 + c_L f_1 v_2) c_R], \quad (45)$$

where $\theta_{L,R}$ are given in (24) and (25), respectively. Under the WS in (6): $\varphi \rightarrow -\varphi$ (equivalently $v_1 \rightarrow -v_1$), θ_L is invariant, but $\theta_R \rightarrow -\theta_R$, then $m_\mu \rightarrow -m_\mu$ and m_M is invariant $m_M \rightarrow m_M$. Now, let us remember that μ and M are in the same fundamental representation [27] of E_6 . This entails that under WS invariance, we will have $M_L \rightarrow -M_L$, $M_R \rightarrow M_R$. Then, the mass eigenstates transform as

$$\begin{aligned} \mu_L^0 &= c_L \mu_L - s_L M_L \rightarrow -\mu_L^0, \\ \mu_R^0 &= c_R \mu_R - s_R M_R \rightarrow \mu_R^0, \\ \hat{M}_L &= s_L \mu_L + c_L M_L \rightarrow -\hat{M}_L, \\ \hat{M}_R &= s_R \mu_R + c_R M_R \rightarrow \hat{M}_R. \end{aligned} \quad (46)$$

Thus, the WS invariance is ensured only when $\theta_R \rightarrow 0$ or when $v_2 \gg v_1$. Consequently, the last transformations imply $m_\mu \rightarrow -m_\mu$, but not $m_M \rightarrow -m_M$ and then one may expect a linear correction to the muon magnetic moment as (31). This analysis do not apply if the muon gets its mass by radiative corrections from other particles.

VI. GENERAL CONCLUSIONS

To conclude, it is possible to explain the muon anomaly in our model based on E_6 through the breaking chain (1), using only Higgses in [78] and [351] representations with a minimal set of Higgses to be singlets and doublet under the

SM symmetry. We find a linear relation between masses for the muon anomaly, if the radiative correction to muon mass, due to mixing with heavy fermion, is small and WS is broken. On the other hand, we find a quadratic relation between masses whenever we suppose that the muon has its mass generated only by radiative corrections in the GUT scale, since, in this case, WS is conserved.

ACKNOWLEDGMENTS

H. Ch. acknowledges the IF of the UFRJ and LAFEX-CBPF/MCT for the kind hospitality and FAPERJ for financial help. C. N. F. and J. A. H.-N. express their gratitude to CNPq-Brazil for the invaluable financial help.

-
- [1] B. Lee Roberts (Muon ($g - 2$) Collaborations), hep-ex/0501012; G. W. Bennet *et al.* (BNL Muon $g - 2$ Collaboration), Phys. Rev. Lett. **92**, 161802 (2004); E. Sichtermann *et al.* (BNL Muon ($g - 2$) Collaboration), hep-ex/0309008; G. W. Bennet *et al.*, Phys. Rev. Lett. **89**, 129903 (2002).
 - [2] J. F. de Trocóniz and F. J. Ynduráin, Phys. Rev. D **71**, 073008 (2005); M. Knecht, Lect. Notes Phys. **629**, 37 (2004); V. Barger, Ch. Kao, P. Langacker, and Hye-Sung Lee, Phys. Lett. B **614**, 67 (2005); A. Czarnecki, Nucl. Phys. B, Proc. Suppl. **144**, 201 (2005); M. Passera, J. Phys. G **31**, R75 (2005).
 - [3] A. Nyffeler, Acta Phys. Pol. B **34**, 5197 (2003).
 - [4] K. Melnikov and A. Vainshtein, Phys. Rev. D **70**, 113006 (2004).
 - [5] M. Davier, S. Eidelman, A. Höcker, and Z. Zhang, Eur. Phys. J. C **31**, 503 (2003); T. Teubner, Eur. Phys. J. C **33**, 653 (2004); A. Höcker, hep-ph/0410081.
 - [6] S. Weinberg, *The Quantum Theory of Fields* (Cambridge University Press, Cambridge, England, 1995), Vol. I, p. 520.
 - [7] R. W. Robinett, Phys. Rev. D **26**, 2396 (1982).
 - [8] R. W. Robinett and J. L. Rosner, Phys. Rev. D **26**, 2388 (1982).
 - [9] H. Chavez and L. Masperi, New J. Phys. **4**, 65 (2002).
 - [10] H. Chavez, L. Masperi, and M. Orsaria, Spacetime & Substance **2**, 111 (2001).
 - [11] Toichiro Kinoshita and Makiko Nio, Phys. Rev. D **73**, 053007 (2006).
 - [12] F. Gürsey, P. Ramond, and P. Sikivie, Phys. Lett. **60B**, 177 (1976); Y. Achiman and B. Stech, Phys. Lett. **77B**, 389 (1978); Q. Shafi, Phys. Lett. **79B**, 301 (1978); R. Barbieri, D. V. Nanopoulos, and A. Masiero, Phys. Lett. **104B**, 194 (1981); J. L. Hewett and Th. G. Rizzo, Phys. Rep. **183**, 193 (1989).
 - [13] R. Barbieri and D. V. Nanopoulos, Phys. Lett. **91B**, 369 (1980).
 - [14] Th. Rizzo, Phys. Rev. D **34**, 1438 (1986).
 - [15] G. Couture and J. N. Ng, Phys. Rev. D **35**, 70 (1987); **33**, 1912 (1986); R. W. Robinett, Phys. Rev. D **33**, 1908 (1986); Th. Rizzo, Phys. Rev. D **33**, 3329 (1986).
 - [16] R. N. Mohapatra and P. B. Pal, *Massive Neutrinos in Physics and Astrophysics*, Lectures Notes in Physics Vol. 41 (World Scientific, Singapore, 1991).
 - [17] J. Leveille, Nucl. Phys. **B137**, 63 (1978).
 - [18] E. Ma, D. Ng, J. Pantaleone, and G. G. Wong, Phys. Rev. D **40**, 1586 (1989).
 - [19] N. V. Cortez, Jr. and M. D. Tonasse, Phys. Rev. D **72**, 073005 (2005).
 - [20] T. W. Kephart and H. Päs, Phys. Rev. D **65**, 093014 (2002).
 - [21] A. Czarnecki and W. J. Marciano, Phys. Rev. D **64**, 013014 (2001).