# **Principles of wide bandwidth acoustic detectors and the single-mass dual detector**

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We apply the standard theory of the elastic body to obtain a set of equations describing the behavior of an acoustic gravitational wave detector, fully taking into account the three-dimensional properties of the mass, the readout, and the signal. We show that the advantages given by a dual detector made by two nested oscillators also can be obtained by monitoring two different acoustic modes of the same oscillator, thus easing the detector realization. We apply these concepts and by means of an optimization process we derive the main figures of such a single mass dual detector designed specifically for the frequency interval 2–5 kHz. Finally we calculate the standard quantum limit sensitivity of this detector.

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## **I. INTRODUCTION**

Direct detection of gravitational waves (GW), for physics and astrophysics study, is one of the great challenges of contemporary experimental physics. Resonant mass bar detectors  $[1-3]$  $[1-3]$  $[1-3]$ , the first historically to come to continuous operation, have been improved by 4 orders of magnitude in energy sensitivity during the last 40 years, so that they can detect energy changes of a 2300 kg bar as little as a few hundred of quanta of vibration at about 1 kHz with a bandwidth up to 100 Hz [[4](#page-12-1)[,5\]](#page-12-2). Methods to further enlarge the bandwidth by a multimode mechanical matching network were proposed in the past  $[6-8]$  $[6-8]$  $[6-8]$ , but, until now, only two-mechanical-mode systems have worked their way into operating detectors.

Even though the direct and unambiguous detection of GW remains an important goal in today's experimental physics, the scientific interest is moving toward the possibility of studying the physical and astrophysical features of the radiation. In fact, gravitational waves will yield unique information about the coherent bulk motions of matter generating the radiation, revealing features of their source which could not be investigated by electromagnetic, cosmic ray, or neutrino studies [[9\]](#page-12-5). In order for acoustic detectors to take part in this ''observatory phase,'' a substantial improvement should be achieved both in terms of absolute sensitivity and detection bandwidth. In fact only a detector sensitive in a wide frequency range, with a good sky coverage for most of the time, will permit to follow the star system evolution during a long time interval, improving the signal reconstruction and the detection probabilities. Wide band acoustic detectors were recently proposed and their expected sensitivity analytically evaluated in two

ideal configurations [[10](#page-12-6)[,11\]](#page-12-7); a precise analytical evaluation of the limit sensitivity of such detectors is still lacking, due to the complexity of the problem.

Present acoustic gravitational wave detectors consist of a test mass, which is strained under the effect of the gravitational wave. A position meter monitors a relevant dimension of the test mass, and thereby deduces the equivalent force signal. The sensitivity evaluations are traditionally made by modeling the test mass as a simple mechanical oscillator with a properly chosen effective mass. This approximation is very good when a narrow band detector sensitive around one of its internal resonances is considered: the detector behavior within its bandwidth is essentially determined by this resonant mode. On the contrary, this approximation is seriously limited when applied to a wide band detector. In this case the detector is also sensitive at frequencies where many internal resonant modes of the test mass will equally contribute to the output signal, each with different amplitudes and phases according to their resonant frequencies. The resulting test-mass dimension variation is then obtained as the sum of a number of contributions and can be enhanced or reduced by the phase relation of the internal resonant modes.

Acoustic detectors work at or between the resonant frequencies of the lowest quadrupolar normal modes of the test mass, which are set by the geometry and material properties. GW sensitive modes in a practical test mass (up to a few meters in diameter) range from about 1 to several kHz. As a general rule, large test masses give better sensitivities and lower frequencies, at the expenses of a greater realization difficulty. The current predictions of the expected rates of GW signals in the kHz band are in rapid evolution, and a full discussion of this topic goes beyond [\\*C](#page-0-1)orresponding author: bonaldi@science.unitn.it the sake of this paper. We point out that a variety of sources

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<span id="page-0-0"></span>

have been considered recently to give possibly significant GW emission in the acoustic frequency range. As an example we give the design guidelines for a detector sensitive in the 2–5 kHz range: it would observe a number of interesting phenomena  $[12-14]$  $[12-14]$  $[12-14]$  $[12-14]$  $[12-14]$  in a frequency range not easily covered by interferometric detectors due to the laser shot noise contribution [\[15\]](#page-12-10).

In Sec. II we develop the general mathematical framework of the complete test mass  $+$  readout system, leading to a set of formulas in which the response of the full detector is related to an external GW force field and to the readout backaction force. We also evaluate the standard quantum limit (SQL) sensitivity of the detector in the ideal case and propose an optimization strategy accounting for the presence of antiresonance frequencies in the transfer functions and the noise properties of the readout. Then in Sec. III we calculate the SQL sensitivity of a single-mass dual acoustic detector specifically designed to be sensitive in a broad frequency range. By an in-depth study of the phase relations between the acoustic modes, we obtain the geometrical conditions which optimize the detector sensitivity. As our calculation takes into account only plane strain resonant modes of the test mass, we validate the results by a finite element method (FEM) simulation of the full three-dimensional body. Conclusions and limits of the current work are summarized in Sec. IV.

# **II. WIDE BAND RESONANT GW DETECTOR GENERAL MODEL**

The detector consists of a test mass and a readout system, which monitors a properly chosen physical observable of the test mass, namely, an average of the surface displacement over a sensitive area. Provided that all external noise sources are reduced down to a negligible level, the test mass changes its dimensions due to the combined effects of the external GW force, the backaction force from the readout and the thermal noise force. In contrast with the lumped model case (which holds for a narrow band resonant detector), the effects of these forces need to be evaluated differently when the detector is considered as an elastic body. In fact, the GW equivalent force is applied on the volume of the body, while the backaction force is applied on the body surface. As a consequence, we need to define two distinct transfer functions for the system, one describing the effects of the GW force and the other the effects of the backaction (BA) force; they are called, respectively, GW transfer function and BA transfer function. A complete description of the detector requires the knowledge of the position and extension of the surfaces on the test mass which the readout samples. In the most general case the readout sensitivity will be position dependent, and a weight function must be defined in order to fully describe the interaction with the test mass.

In the following, vectors are indicated by bold characters.

### **A. Test mass mode expansion**

<span id="page-1-0"></span>The equations of the motion of an elastic body of density  $\rho$ , forced by a force density  $\mathbf{F}(\mathbf{r}, t)$ , can be summarized as [\[16\]](#page-12-11)

$$
\rho \frac{\partial^2 \mathbf{u}(\mathbf{r}, t)}{\partial t^2} - L[\mathbf{u}(\mathbf{r}, t)] = \mathbf{F}(\mathbf{r}, t), \quad (1)
$$

with the appropriate initial and boundary conditions. Here  $\mathbf{u}(\mathbf{r}, t)$  is the displacement field of the elastic body and  $L[\mathbf{u}(\mathbf{r}, t)]$  is defined as

$$
L[\mathbf{u}] = (\lambda + \mu)\nabla(\nabla \cdot \mathbf{u}) + \mu \nabla^2 \mathbf{u}.
$$
 (2)

The Lame coefficients  $\lambda$  and  $\mu$  depend on the Poisson ratio  $\sigma_p$  and on the Young modulus *Y* of the material:

$$
\lambda = \frac{Y\sigma_p}{(1 + \sigma_p)(1 - 2\sigma_p)}, \qquad \mu = \frac{Y}{2(1 + \sigma_p)}.
$$
 (3)

<span id="page-1-2"></span>We define the displacement normal modes  $w_n(r)$  as the solutions of the eigenvalue problem,

$$
- \rho \omega_n^2 \mathbf{w}_n = L[\mathbf{w}_n], \tag{4}
$$

with the boundary conditions on  $w_n$  given by the requirement that components of the stress normal to the test-mass surfaces vanish on the test-mass surfaces. The normal modes constitute an orthogonal complete system, and can be normalized to satisfy the condition

$$
\int_{V} \rho \mathbf{w}_{n}(\mathbf{r}) \cdot \mathbf{w}_{m}(\mathbf{r}) dV = M \delta_{nm},
$$
\n(5)

<span id="page-1-3"></span>where the volume integral is performed on the test-mass volume *V* and *M* is its total mass.

<span id="page-1-1"></span>The solution of Eq.  $(1)$  may then be written as a superposition of the normal modes:

$$
\mathbf{u}(\mathbf{r},t) = \sum_{n} \mathbf{w}_n(\mathbf{r}) q_n(t).
$$
 (6)

The sum runs over an infinite number of elements, as normal modes constitute an infinite enumerable set within the continuum mechanics mathematical framework [[17\]](#page-12-12). The functions  $q_n(t)$  in Eq. [\(6\)](#page-1-1) represent the time development of the (*n*)th mode, with initial value given by

$$
q_n(0) = \frac{1}{M} \int_V dV \rho \mathbf{u}(\mathbf{r}, 0) \cdot \mathbf{w}_n(\mathbf{r}). \tag{7}
$$

The evaluation of the explicit form of  $q_n(t)$  is straightfor-ward [\[18\]](#page-12-13) when the driving force can be factorized as

$$
F(\mathbf{r}, t) = G_t(t)\mathbf{G}_r(\mathbf{r}).\tag{8}
$$

In fact Eq. [\(1\)](#page-1-0) becomes

$$
\rho \sum_n \mathbf{w}_n(\mathbf{r}) \frac{\partial^2 q_n(t)}{\partial t^2} - \sum_n q_n(t) L[\mathbf{w}_n(\mathbf{r})] = G_t(t) \mathbf{G}_r(\mathbf{r}),
$$

and we can apply Eq. ([4\)](#page-1-2), multiply by  $w_m$  and integrate using Eq.  $(5)$ , obtaining

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$$
M\frac{\partial^2 q_m(t)}{\partial t^2} + M\omega_m^2 q_m(t) = G_t(t) \int_V dV \mathbf{G}_r(\mathbf{r}) \cdot \mathbf{w}_m(\mathbf{r}).
$$

The time development of the (*m*)th mode is the same of a forced harmonic oscillator. In the frequency domain (here and in the following we indicate the Fourier transform with a tilda) we have then

$$
\tilde{q}_m(\omega) = \frac{1}{M} \frac{\tilde{G}_t(\omega)}{(\omega_m^2 - \omega^2)} \int_V dV \mathbf{G}_r(\mathbf{r}) \cdot \mathbf{w}_m(\mathbf{r}). \quad (9)
$$

In the hypothesis of dissipation due to the material structure, the system losses can be modeled in the frequency domain by including explicitly a damping term [[19](#page-12-14)]. We have then

<span id="page-2-6"></span>
$$
\tilde{q}_m(\omega) = \frac{1}{M} \frac{\tilde{G}_l(\omega)}{(\omega_m^2 - \omega^2) + i \omega_m^2 \phi_m(\omega)} \int_V dV \mathbf{G}_r(\mathbf{r}) \cdot \mathbf{w}_m(\mathbf{r}).
$$
\n(10)

The function  $\phi_n(\omega)$  represents the phase lag between the mass displacement at a given frequency and a monochromatic driving force. Experimentally,  $\phi_n(\omega)$  is found to be roughly constant (this is usually referred to as ''structural damping") and, for low loss materials, it is  $\phi_n(\omega) \ll 1$ . It can be shown that the normal mode expansion is possible only if the damping term is homogeneous over the testmass volume, as any inhomogeneity of the structure damping causes a coupling between different modes [[20](#page-12-15)]. When the losses are frequency independent, the material quality factor  $Q = 1/\phi$  can be used equivalently in place of the phase lag.

## **B. Readout**

<span id="page-2-0"></span>To detect the external force we measure the resulting strain on some test-mass surface *S*. In the small displacement approximation the observable physical quantity, *X*, of the system may be defined as

$$
X(t) = \int_{S} ds \mathbf{P}(\mathbf{r}) \cdot \mathbf{u}(\mathbf{r}, t).
$$
 (11)

Here  $P(r)$  is a weighting function and the integral is performed on the chosen surface *S* of the test mass. We can decide freely the portions of the test-mass surface which are sampled and their relative weight in the construction of our output variable  $X(t)$ : the spatial form of the weight function  $P(r)$  reflects our measurement strategy and detection scheme. For example, in an optical readout,  $P(r)$ is proportional to the beam spot power profile [\[21](#page-12-16)], while in a capacitive readout [[22](#page-12-17)] we have  $P(r) = 1/S_0$  on the surface  $S_0$  of the electrodes and  $P(r) = 0$  outside (as border effects are usually negligible).

By Eqs.  $(6)$  $(6)$  and  $(11)$  the observable physical quantity of the system is

$$
X(t) = \sum_{n} q_n(t) \int_{S} ds \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r}),
$$
 (12)

<span id="page-2-1"></span>or, in the frequency domain

$$
\tilde{X}(\omega) = \frac{\tilde{G}_t(\omega)}{M} \sum_n \frac{\left[\int_V dV \mathbf{G}_r(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r})\right] \left[\int_S dS \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r})\right]}{(\omega_n^2 - \omega^2) + i\omega_n^2 \phi_n(\omega)}.
$$
\n(13)

This equation allows us to evaluate the motion of the system when the specific form of the driving force spatial dependence and of the readout weight function are given. Then we evaluate the detector response when driven by the relevant forces: a GW, the readout backaction force, and the thermal noise.

In the case of a linearly polarized gravitational wave propagating along the *z* axis, the force density applied on the system may be written in cylindrical coordinates  $(r, \theta)$ as [\[23\]](#page-12-18)

<span id="page-2-5"></span>
$$
G_t(t) = \frac{1}{2}\rho \ddot{h}(t),
$$
  
\n
$$
\mathbf{G}_r(\mathbf{r}) \equiv \mathbf{W}(\mathbf{r}) = r \cos(2\theta + \psi)\mathbf{i}_r - r \sin(2\theta + \psi)\mathbf{i}_\theta,
$$
\n(14)

where  $h(t)$  is the metric perturbation associated to the GW,  $\psi = 0$  for the GW + polarization and  $\psi = \pi/4$  for the  $\times$ polarization. When we substitute these in Eq.  $(13)$  $(13)$  $(13)$ , the system response to a metric perturbation *h* becomes

$$
\tilde{X}(\omega) = \tilde{h}(\omega) H_{\rm GW}(\omega),\tag{15}
$$

<span id="page-2-4"></span>where the detector output is fully described by the GW transfer function  $H_{\text{GW}}(\omega)$ :

$$
H_{\rm GW}(\omega)
$$
  
=  $\frac{1}{2V} \sum_{n} \frac{-\omega^2 [ \int_V dV \mathbf{W}(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r}) ] [ \int_S dS \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r}) ]}{(\omega_n^2 - \omega^2) + i \omega_n^2 \phi_n(\omega)}$ . (16)

<span id="page-2-2"></span>The second relevant force applied to the system is the readout backaction. Being exerted by the readout, this force is applied proportionally to the weight function  $P(r)$ :

$$
\mathbf{F}_{\mathrm{BA}}(t,r) = F_{\mathrm{BA}}(t)\mathbf{P}(\mathbf{r}).\tag{17}
$$

The test-mass response to this input force can again be evaluated by Eq.  $(13)$ , if the backaction surface force Eq. [\(17\)](#page-2-2) is considered in place of the volume force field  $G_t(t)G_r(r)$ :

$$
\tilde{X}(\omega) = T_{BA}(\omega)\tilde{F}_{BA}(\omega)
$$
\n
$$
= \frac{\tilde{F}_{BA}(\omega)}{M} \sum_{n} \frac{[\int_{s} ds \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_{n}(\mathbf{r})]^{2}}{(\omega_{n}^{2} - \omega^{2}) + i\omega_{n}^{2} \phi_{n}(\omega)}.
$$
\n(18)

<span id="page-2-3"></span>The BA transfer function of the system, relative to the backaction force, is then

$$
T_{\text{BA}}(\omega) = \frac{1}{M} \sum_{n} \frac{\left[\int_{s} ds \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_{n}(\mathbf{r})\right]^{2}}{(\omega_{n}^{2} - \omega^{2}) + i\omega_{n}^{2} \phi_{n}(\omega)},
$$
(19)

and the resulting displacement can be considered as the

### **C. Readout optimization and SQL**

The readout noise properties are defined in terms of the single sided power spectral densities  $S_{xx}(\omega)$  and  $S_{ff}(\omega)$ . Here  $S_{xx}(\omega)$  is the equivalent input displacement noise power spectral density contributed by the readout;  $S_{ff}(\omega)$ is the equivalent force noise power spectrum density due to the readout that drives the system through the BA transfer function  $T_{BA}$  [Eq. ([19](#page-2-3))]. Under the hypothesis of uncorrelated noise sources, the readout contribution to the noise on the observable *X* is

$$
S_{XX}(\omega) = S_{xx}(\omega) + |T_{BA}(\omega)|^2 S_{ff}(\omega).
$$
 (20)

<span id="page-3-0"></span>A useful figure for expressing the readout performance is its energy resolution expressed as number of energy quanta:

$$
\epsilon_r(\omega) = \frac{\sqrt{S_{xx}(\omega)S_{ff}(\omega)}}{\hbar}.
$$
 (21)

The uncertainty relation for a continuous linear measurement requires that [\[24\]](#page-12-19)

$$
\epsilon_r(\omega) \ge 1,\tag{22}
$$

and the readout is called ''quantum limited'' when this limit is achieved. It is also helpful to define the readout noise stiffness  $\kappa_r$  as

$$
\kappa_r(\omega) = \sqrt{S_{ff}(\omega)/S_{xx}(\omega)}.
$$
 (23)

<span id="page-3-3"></span>If we now write the power spectral densities  $S_{xx}(\omega)$  and  $S_{ff}(\omega)$  in terms of  $\epsilon_r$  and  $\kappa_r$ ,

$$
S_{xx}(\omega) = \hbar \epsilon_r(\omega) / \kappa_r(\omega), \qquad (24)
$$

$$
S_{ff}(\omega) = \hbar \epsilon_r(\omega) \kappa_r(\omega), \qquad (25)
$$

the Eq. ([20](#page-3-0)) becomes

$$
S_{XX}(\omega) = \hbar \epsilon_r(\omega) \bigg[ \frac{1}{\kappa_r(\omega)} + |T_{BA}(\omega)|^2 \kappa_r(\omega) \bigg].
$$
 (26)

It is straightforward to demonstrate that for each  $\omega$  the readout noise contribution is minimized, giving the SQL [\[24\]](#page-12-19) for the detector, if the following equations are both satisfied:

$$
\epsilon_r(\omega) = 1,\tag{27}
$$

$$
\kappa_r(\omega) = \frac{1}{|T_{\text{BA}}(\omega)|}.
$$
 (28)

<span id="page-3-1"></span>Equivalently the following must hold:

$$
S_{xx}(\omega) = \hbar |T_{BA}(\omega)|, \qquad S_{ff}(\omega) = \frac{\hbar}{|T_{BA}(\omega_b)|}.
$$
 (29)

These relations show that, as obvious, the best detector performances require the use of a quantum limited readout and, less trivial, the noise stiffness must be properly matched to the test-mass mechanical impedance  $T_{BA}(\omega)$ . The total displacement noise becomes

$$
S_{XX}(\omega) = 2\hbar |T_{BA}(\omega)|, \tag{30}
$$

<span id="page-3-2"></span>and can be transformed into an equivalent input GW spectral density through the GW transfer function  $H_{\text{GW}}(\omega)$  [Eq. [\(16](#page-2-4))]. The detector sensitivity is then

$$
S_{hh}(\omega) = 2h \frac{|T_{\text{BA}}(\omega)|}{|H_{\text{GW}}(\omega)|^2}.
$$
 (31)

We point out that the SQL here derived is calculated taking into account the volume properties of test mass, signal, and readout: this is in contrast to the usual derivation of the SQL for a GW detector, in which a onedimensional model is considered.

In principle a wide bandwidth detector, designed to be sensitive in a angular frequency range  $\omega_{\min} \div \omega_{\max}$  could reach the SQL at every frequency, provided that a readout satisfying Eq. ([29](#page-3-1)) on the full interval  $\omega_{\min} \div \omega_{\max}$  can be designed. The elastic body BA transfer function  $T_{BA}(\omega)$  is a rapidly varying function (as we show below in Fig.  $4(b)$ ): sharp magnitude peaks, corresponding to the resonant normal modes of the system, alternate with magnitude dips called antiresonances. Unfortunately, the noise properties of a practical wide bandwidth readout have a smooth frequency dependence over the detector sensitivity range and cannot match these structures of the BA transfer function, preventing the possibility of reaching the SQL on a wide frequency range.

The origin of antiresonances is well explained within the modal expansion model. We show in the appendix that in the backaction response  $T_{BA}(\omega)$  there must be an antiresonance following a resonance, without exception and regardless of the complexity of structures. An optimized readout should have a high backaction force noise spectral density at the antiresonances frequencies, where the detector is quite insensitive. On the contrary it must have low backaction noise at frequencies where resonances of  $T_{BA}(\omega)$  occur, in order to avoid the resonance amplification of the readout force noise. The optimization process will approach the best performances by averaging among these conflicting requirements.

The frequency dependence of the readout noise properties depends on its actual design. In order to show the optimization strategy, we work out here the simple case of a readout with constant noise properties:

$$
\epsilon_r(\omega) = \epsilon_0, \tag{32}
$$

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$$
\kappa_r(\omega) = \kappa_0. \tag{33}
$$

The readout contribution to the equivalent input displacement noise on the variable *X* becomes now

$$
S_{XX}(\omega) = \hbar \epsilon_0 \left[ \frac{1}{\kappa_0} + |T_{BA}(\omega)|^2 \kappa_0 \right],\tag{34}
$$

<span id="page-4-0"></span>and the detector sensitivity is simply

$$
S_{hh}(\omega) = \frac{\hbar \epsilon_0}{|H_{\text{GW}}(\omega)|^2} \left[ \frac{1}{\kappa_0} + |T_{\text{BA}}(\omega)|^2 \kappa_0 \right].
$$
 (35)

<span id="page-4-4"></span>We find again that for best performances the use of a quantum limited readout  $(\epsilon_0 = 1)$  is required, and the noise stiffness  $\kappa_0$  must be properly matched to the readout. The optimization strategy necessarily depends on the characteristic of the expected GW sources. In fact, for a given GW signal  $\tilde{h}(\omega)$ , the measurement's signal to noise ratio *S/N* is defined as

$$
\frac{S}{N} = \sqrt{\frac{2}{\pi} \int_0^{+\infty} \frac{|\tilde{h}(\omega)|^2}{S_{hh}(\omega)} d\omega}.
$$
 (36)

#### **D. Thermal noise**

In order to reach the sensitivity shown in Eq.  $(35)$ , technical noises must be reduced down to a negligible level. Among all possible technical noise sources, the one due to the fluctuations of the body volume (the socalled brownian or thermal noise) often ends up as a barrier preventing the achievement of the SQL.

In the case of a system at equilibrium with a bath at temperature  $T$ , with a single coordinate  $p$ , and a force  $F$ that represents the interaction between the system and the externals, the single sided thermal noise power spectrum of the coordinate is predicted by the fluctuation-dissipation theorem [[25](#page-12-20)] to be

$$
S_{pp}(\omega) = -\frac{4k_B T}{\omega} \Im m[T_F(\omega)], \tag{37}
$$

where  $T_F(\omega)$  is the transfer function

$$
T_F(\omega) = \frac{\tilde{p}(\omega)}{\tilde{F}(\omega)}.
$$
\n(38)

When the system coordinate *X* is a linear combination weighted by a function  $P(r)$ , as in Eq. ([11](#page-2-0)), the thermal noise on *X* can be evaluated by the system response to a force  $F(t)$ **P**(**r**) applied to the test-mass surface [\[26\]](#page-12-21). The thermal noise power spectral density on the output variable is then given by the BA transfer function  $T_{BA}(\omega)$ [Eq. [\(19\)](#page-2-3)]:

<span id="page-4-1"></span>
$$
S_{XX}(\omega, T) = -\frac{4k_B T}{\omega} \Im m[T_{BA}(\omega)]
$$
  
= 
$$
-\frac{4k_B T}{\omega} \sum_n \frac{-\omega_n^2 \phi_n(\omega)}{M}
$$
  

$$
\times \frac{|\int_S ds \mathbf{P}(\mathbf{r}) \cdot \mathbf{w}_n(\mathbf{r})|^2}{(\omega_n^2 - \omega^2)^2 + \omega_n^4 \phi_n^2(\omega)}.
$$
 (39)

The mechanical thermal noise of the system can be reduced in principle below the readout contributed noise by reducing the detector temperature *T*. If the thermal noise evaluated in Eq. [\(39](#page-4-1)) must be well below the readout SQL noise given by Eq.  $(30)$  $(30)$ , then

$$
-\frac{4k_B T}{\omega} \Im m[T_{BA}(\omega)] \ll 2\hbar |T_{BA}(\omega)|. \tag{40}
$$

<span id="page-4-2"></span>Note that the BA transfer function  $T<sub>BA</sub>$  enters in both sides of Eq. ([40](#page-4-2)) but with different meanings: the left-hand side is essentially the incoherent superposition of the thermal noises generated by each normal mode of the system, while the right-hand side is the coherent superposition of the normal mode responses.

The effect of the detector thermal noise on the  $S/N$  can be easily evaluated if the sensitivity  $S_{hh}$  [Eq. [\(35\)](#page-4-0)] is generalized as

<span id="page-4-3"></span>
$$
S_{hh}(\omega, T) = \frac{\hbar \epsilon_0 [\frac{1}{\kappa_0} + |T_{BA}(\omega)|^2 \kappa_0] + S_{XX}(\omega, T)}{|H_{\text{GW}}(\omega)|^2}, \quad (41)
$$

where the detector thermal noise  $S_{XX}(\omega, T)$  is given by Eq. [\(39](#page-4-1)). In the case of structural damping, the thermal noise contribution to  $S_{hh}(\omega, T)$  is proportional to the product  $T\phi$ .

## **III. A CASE STUDY: THE SINGLE-MASS DUAL DETECTOR**

On the basis of the optimization process described in Sec. II C, we can estimate the SQL sensitivity of specific detector configurations, provided that the eigenvalue problem Eq. [\(4](#page-1-2)) for the test mass is solved. As an example, we consider a hollow cylinder test mass [Fig.  $1(a)$ ]. If we limit our evaluation to GW propagating along the *z* axis, the symmetry axis of the system, the corresponding force does not depend on *z*: the system response can be well described by plane strain solutions, where the displacements are functions of *x* and *y* only and the displacement along *z* vanishes. The analytical plain strain solutions are well known [[27](#page-12-22)] and we use them to evaluate the detector sensitivity; we make use of FEM simulations to confirm that the main features of the detector are maintained when a real three-dimensional test mass is considered.

We show in the following that, thanks to the effect of its lowest order quadrupolar resonant modes, this *single* testmass detector offers the same advantages of the two-nested mass configuration (namely, the dual detector), recently proposed  $[10,11,28]$  $[10,11,28]$  $[10,11,28]$  $[10,11,28]$ : the large bandwidth, set by the fre-

<span id="page-5-2"></span>

<span id="page-5-0"></span>FIG. 1. (a) Test mass and readout: the displacements induced by a GW propagating along the *z* axis are measured in eight regions (in black) spanning the whole cylinder height *L*. (b) Section of the detector: the angle  $\alpha$  determines the extent of the measuring surfaces and represents a compromise between the requirements of wide sampling area and the possibility of implementing eight independent sampling areas on the test mass. The value  $\alpha = 0.2$  rad is used throughout the paper.

quency difference between the lowest order GW sensitive modes of each test mass, and the backaction reduction effect, determined by the antiresonance frequency placed between these modes. To fully exploit the dual detector properties, we consider to implement a pair of selective readouts. These readouts [\[11\]](#page-12-7) are sensitive to the two independent components of the quadrupolar modes and reject other classes of modes, which are not excited by GWs.

The selective readouts  $R_{\Theta}$  and  $R_{\infty}$  (Fig. [2](#page-5-1)) measures the inner diameters of the hollow cylinder and combines them to obtain  $X_{\oplus} = d_2 - d_1$  and  $X_{\otimes} = d_4 - d_3$ . In each diameter evaluation the test-mass inner surface displacement is averaged over two opposite areas spanning the full cylinder length. These detection schemes are implemented by the weight function

<span id="page-5-1"></span>

FIG. 2. Sections of the detector showing the detail of the readout measurement strategy. The measuring surfaces are arranged in two independent readout configurations: the readout  $R_{\Theta}$  measures the difference  $X_{\Theta} = d_2 - d_1$ , while  $R_{\Theta}$  measures the difference  $X_{\infty} = d_4 - d_3$ . It will be shown that  $R_{\infty}$  and  $R_{\infty}$ are sensitive, respectively, to the  $+$  and  $\times$  components of a GW propagating along the *z* axis. Both readouts are mainly sensitive to the quadrupolar modes of the test mass.

$$
\mathbf{P}_{\mathbf{\Theta}}(\mathbf{r}) = \delta(r - r_{\text{int}})P_{\theta}(\theta)\mathbf{i}_r \tag{42}
$$

<span id="page-5-4"></span>
$$
\mathbf{P}_{\mathbf{\otimes}}(\mathbf{r}) = \delta(r - r_{\text{int}})P_{\theta}\left(\theta + \frac{\pi}{4}\right)\mathbf{i}_{r},\tag{43}
$$

with

$$
P_{\theta} = \frac{1}{S_0} \sum_{m=0}^{1} \sum_{n=0}^{4} (-1)^{n+m} \Theta \left[ \theta + (-1)^m \alpha - n \frac{\pi}{2} \right], \quad (44)
$$

where  $\Theta(x)$  represents the unit step function and  $2\alpha$  is the angle subtended by each measured surface [Fig. [1\(b\)](#page-5-0)]. The value  $\alpha = 0.2$  rad is used in the following: it represents a good compromise between the requirements of wide sampling area and the possibility of implementing two readouts rotated by  $\pi/4$  on the same test mass. The normalization  $S_0$  is the area of each measured surface, evaluated as  $S_0 \approx 2\alpha r_{\text{int}}L$ ; as shown in Fig. [1](#page-5-2),  $r_{\text{int}}$  is the inner radius of the hollow cylinder. It is straightforward to demonstrate that, according to Eq. ([41](#page-4-3)), the detector sensitivity does not depend on the chosen normalization  $S_0$ .

### **A. Detector response**

The plain strain normal modes of the motion of the cylinder in the  $x - y$  plane are a solution to the eigenstate problem [Eq. ([4](#page-1-2))], when the following additional boundary conditions are applied:

$$
u_z \equiv 0, \qquad \frac{\partial (u_r, u_z, u_\theta)}{\partial z} \equiv 0, \tag{45}
$$

where  $(u_r, u_\theta, u_z)$  are components of the displacement  $\mathbf{u}(\mathbf{r})$ in a cylindrical coordinate system. The analytical expressions of the normal modes of a hollow cylinder in the plane strain approximation are functions of the kind [\[27\]](#page-12-22):

<span id="page-5-3"></span>
$$
\mathbf{w}_{a,n}^{\mathbf{\Theta}}(\mathbf{r}) = f_{a,n}(r) \cos(a\theta) \mathbf{i}_r + g_{a,n}(r) \sin(a\theta) \mathbf{i}_\theta,
$$
  

$$
\mathbf{w}_{a,n}^{\mathbf{\Theta}}(\mathbf{r}) = -f_{a,n}(r) \sin(a\theta) \mathbf{i}_r + g_{a,n}(r) \cos(a\theta) \mathbf{i}_\theta,
$$
 (46)

where the functions  $f_{a,n}$ ,  $g_{a,n}$  are linear combinations of Bessel functions of the coordinate  $r$ , with coefficients determined by the boundary conditions. The integer *a* represents the angular symmetry of the mode, while *n* identifies the mode order within the angular family; modes within a given family are ordered by increasing frequency with increasing *n*. The orthogonal displacement fields  $\mathbf{w}_{a,n}^{\oplus}$ ,  $\mathbf{w}_{a,n}^{\otimes}$  represent the same radial distribution of the deformation, mutually rotated by  $\frac{\pi}{2a}$ : for this reason they share the same eigenvalue  $\omega_{a,n}$ . We call  $\nu_{a,n} = \omega_{a,n}/(2\pi)$  the resonant frequency of the mode. The displacement  $\mathbf{w}_{a,n}^s$  is an element of an orthogonal complete system and is normalized to satisfy the condition (with  $s, t = \theta$  or  $\otimes$ ):

$$
\int_{V} dV \rho \mathbf{w}_{a,n}^{s}(\mathbf{r}) \cdot \mathbf{w}_{b,m}^{t}(\mathbf{r}) = M \delta_{st} \delta_{ab} \delta_{nm}.
$$
 (47)

Any plane strain displacement  $\mathbf{u}_p$  may be written as linear superposition of these basis functions, with timedependent coefficients determined by the force acting on the body:

$$
\mathbf{u}_{p}(\mathbf{r}, t) = \sum_{s, a, n} \mathbf{w}_{a,n}^{s}(\mathbf{r}) q_{a,n}^{s}(t). \tag{48}
$$

The plain strain approximation holds in presence of an external force when this vanishes, with its first derivatives, along the *z* axis:

$$
\mathbf{G}_r(\mathbf{r}) \cdot \mathbf{i}_z = 0, \qquad \frac{\partial \mathbf{G}_r(\mathbf{r})}{\partial z} = 0. \tag{49}
$$

Moreover, the same conditions must hold for the weight function of the readout used to measure the displacement:

$$
\mathbf{P}(\mathbf{r}) \cdot \mathbf{i}_z = 0, \qquad \frac{\partial \mathbf{P}(\mathbf{r})}{\partial z} = 0. \tag{50}
$$

The plane strain solutions are exact for an infinite length cylinder, as the fulfillment of plane strain condition requires the application, over the terminal sections, of tension or pressure adjusted so as to keep constant the length of all longitudinal filaments (that is the body height). In the case of a finite length cylinder the solutions can be considered exact only if a surface force opposite to the internal stress is applied to the end surfaces of the cylinder. According to the Saint-Venant principle [\[16\]](#page-12-11), we infer that the plane strain approximation is good when the length of the cylinder is large compared to its diameter. In this case, stress and strain in the interior are practically independent of the presence of a force distribution applied on the ends, in all the portions of the cylinder except comparatively small slices near its ends.

On the basis of the solutions Eq.  $(46)$ , for instance, we can evaluate the effect of  $a +$  polarized GW propagating along the *z* axis, say  $W^+(r)$ , given by Eq. [\(14\)](#page-2-5) with  $\psi = 0$ . This wave can excite only the family of the  $w^{\oplus}$  plane strain quadrupolar  $(a = 2)$  modes, as we have

$$
\int_{V} dV \mathbf{W}^{+}(\mathbf{r}) \cdot \mathbf{w}_{2,n}^{\otimes}(\mathbf{r}) = 0,
$$
\n(51)

due to the quadrupolar symmetry of the force field. The resulting displacement would be the superposition of the displacement contributed by an infinite number of modes with  $a = 2$ , but we find that the main contribution is given by the two modes of lowest frequency,  $\mathbf{w}_{2,1}^{\oplus}(\mathbf{r})$  and  $\mathbf{w}_{2,2}^{\oplus}(\mathbf{r})$ . These modes induce an elliptical strain of the inner surface of the cylinder, as shown in Fig. [3.](#page-6-0) We remind the reader that these modal shapes remain unchanged over the whole frequency range (namely, also far from their resonance), whatever the frequency of the driving force, while their relative strength varies according to Eq. [\(10\)](#page-2-6).

As discussed in the appendix, the relative phase between the two contributions depends on the sign of the modal constants, which in this case, [according to Eq.  $(16)$ ] are determined by the convolution of the GW spatial force  $W(r)$  over the modal shape *and* by the readout weight function  $P_{\oplus}(\mathbf{r})$  [Eq. [\(42\)](#page-5-4)]:

<span id="page-6-1"></span>
$$
C_{2,n}^{\mathbf{\Theta}} \propto \left[ \int_{V} dV \mathbf{W}(\mathbf{r}) \cdot \mathbf{w}_{2,n}^{\mathbf{\Theta}}(\mathbf{r}) \right] \left[ \int_{S} ds \mathbf{P}_{\mathbf{\Theta}}(\mathbf{r}) \cdot \mathbf{w}_{2,n}^{\mathbf{\Theta}}(\mathbf{r}) \right].
$$
 (52)

<span id="page-6-0"></span>

FIG. 3. Modal shape of the first two plane strain quadrupolar modes. The continuous lines show the deformation of internal and external surfaces and of twelve radial lines traced ideally inside the cylinder section. The shape of the cross section at null deformation is shown as a dashed line. Both modes induce deformations of the inner surface of the cylinder and then contribute to the detector output. (a) First quadrupolar mode  $\mathbf{w}_{2,1}^{\mathbf{\Theta}}(\mathbf{r})$ : both the inner and outer surfaces change their shape. The cylinder thickness  $r_{ext} - r_{int}$  remains approximately constant (flexural mode). (b) Second quadrupolar mode  $\mathbf{w}_{2,2}^{\mathbf{\Theta}}(\mathbf{r})$ : the outer surface remains essentially unchanged and the deformation involves only inner portions of the cross section. This deformation is then obtained by a change of the cylinder thickness (bulk mode).

The two modes induce a similar deformation of the inner surface, but given by different deformations of the cylinder interior: the  $\mathbf{w}_{2,1}^{\oplus}$  is a flexural mode and it does not change the cylinder thickness  $r_{ext} - r_{int}$ , while  $\mathbf{w}_{2,2}^{\oplus}$  is a bulk mode which changes the thickness. It is then not surprising that the analytical evaluation of Eq. [\(52\)](#page-6-1) results in two modal constants with *opposite sign*. For this reason, when the detector is driven by a GW at frequency  $\nu_{\text{GW}}$ , we recover the characteristics of a dual detector  $[10]$ ; namely, the contribution of the two modes

- (i) subtracts at low frequencies ( $\nu_{\text{GW}} < \nu_{2,1}$ ), when they follow in phase the driving force.
- (ii) is enhanced when the modes get out of phase in the interval ( $\nu_{2,1} \div \nu_{2,2}$ ), thanks to the  $\pi$  phase lag of the mode  $\nu_{2,1}$  which is driven above its resonant frequency. This frequency range determines approximately the detector bandwidth.
- (iii) subtracts at higher frequencies ( $\nu$ <sub>GW</sub> >  $\nu$ <sub>2,2</sub>) when both modes get out of phase because they are driven above their resonance.

The resulting GW transfer function is shown in Fig.  $4(a)$ : the sensitivity peaks due to the quadrupolar modes can be seen clearly. Figure  $4(b)$  shows the BA transfer function  $T<sub>BA</sub>$ , evaluated on the basis of the readout weight function  $P_{\varphi}(\mathbf{r})$  with  $\alpha = 0.2$  rad. An antiresonance shows up within the range  $\nu_{2,1} \div \nu_{2,2}$ , as expected on the basis of the discussion in the appendix. The frequency of the antiresonance depends not only on the modal constants of the first two quadrupolar modes but also by the contribution of all modes sensed by the readout. We observed that a conver-

<span id="page-7-1"></span>

<span id="page-7-0"></span>FIG. 4. Transfer functions for the  $R_{\oplus}$  readout with  $\alpha = 0.2$  rad describing the dynamical response of a hollow cylinder singlemass dual detector (material: molybdenum, outer radius  $r_{ext}$  = 0.5 m, height  $L = 3$  m, form ratio  $r_{\text{int}}/r_{\text{ext}} = 0.3$ ). (a) Analytical (continuous line) and FEM evaluated (gray line) GW transfer function for a signal propagating along *z* axis. Dashed line: 3D FEM evaluated response under  $a + polarized$ GW traveling along the *y* axis. (b) BA transfer function. Continuous line: analytical plain strain evaluation. Gray line: 3D FEM simulation.

gence better than  $1\%$  is obtained if all modes with frequency below 30 kHz are considered in the evaluation (about 15 modes).

In Fig. [4](#page-7-1) we also show the transfer functions (both BA and GW type) obtained by a three-dimensional FEM harmonic analysis. These harmonic analysis are performed in the range  $0 \div 10$  kHz using the ANSYS [\[29\]](#page-12-24) *full method* of solving dynamic equations, proved to be an accurate method of evaluating the elastic body dynamical impedance [[30](#page-12-25)]. The FEM code calculates the transfer functions from the displacement averaged over the readout surface according to the weight function Eq.  $(42)$  $(42)$  $(42)$ . To evaluate  $T_{BA}$ , a unit external force is applied onto the readout surfaces according to Eq. ([17\)](#page-2-2). The GW volume force field Eq. [\(14\)](#page-2-5) is applied to evaluate  $H<sub>GW</sub>$ . As expected, the resonant modes found in the plane strain approximation are essentially confirmed by the 3D FEM evaluation, but other modes appear that cannot be obtained in the plane strain framework. Thanks to the readout capability of rejecting non-GW sensitive modes, the displacement induced by these additional modes is efficiently averaged out and the detector dynamical performances do not differ from that evaluated in the plain strain approximation. This fact confirms that the body dynamics is modified only weakly by the absence of the balancing force applied at the bottom and top surfaces of the cylinder.

The FEM code is not limited to plane strain evaluation and also can calculate the sensitivity pattern of the detector for GWs arriving from an arbitrary direction  $(\theta, \varphi)$  with an arbitrary polarization  $\psi$ . We point out that the use of FEM to calculate the directional sensitivity is applicable to acoustic detectors of arbitrary shape, in particular, to all of the resonant spherical and bar detectors. Following the approach introduced in Ref. [\[31\]](#page-12-26), we separate the radiation into its two linear polarization components, described by  $\psi = 0$  and  $\psi = \pi/4$  in the wave front plane of the GW. When  $(\theta, \varphi) = (0, 0)$  these force fields reduce, respectively, to the  $+$  and  $\times$  GW described by Eq. [\(14\)](#page-2-5).

The frequency response to a GW is scaled down for a wave from any direction with respect to the *z*-propagating case, as shown in Fig.  $4(a)$  in the case of a wave traveling along the *y* axis. We evaluated the scaling factor as a function of the signal direction  $(\theta, \phi)$  and polarization  $\psi$ at a fixed frequency (3000 Hz). For each readout  $R_{\oplus}$  and  $R_{\infty}$ , the normalized output for a circularly polarized GW is obtained by quadratically averaging the response to  $\psi = 0$ and  $\psi = \pi/4$  waves. The antenna pattern shown in Fig.  $5(a)$  refers to circularly polarized signals and shows the maximum response between the two readouts.

The dual detector is then omnidirectional, in the sense that no blind directions show up in its antenna pattern. On the other hand it is not isotropic, and the lowest response for a circularly polarized GW is still 30% of the response to an optimally oriented GW, i.e. propagating along *z*. In general, the two linear polarization components of the GW excite two independent linear combinations of  $\mathbf{w}_{2,n}^{\mathcal{D}}(\mathbf{r})$  and  $\mathbf{w}_{2,n}^{\infty}(\mathbf{r})$ , which are separately read by  $R_{\mathcal{D}}$  and  $R_{\infty}$ , respectively. Therefore, it is relevant to point out that the single-mass dual detector can resolve both amplitude polarizations of the impinging GW if its propagation direction  $(\theta, \phi)$  is known [[32](#page-12-27)]. This is not the case of the operating interferometric or resonant bar detectors, which measures only one linear polarization component of the GW.

The presence of two independent readout channels which monitor different signal polarizations also enhance the possibilities of setting up vetoing strategies in joint observations with other GW detectors, by exploiting Riemann tensor properties to discriminate noise events. In addition, the two independent readouts are affected by independent intrinsic noise sources, such as thermal and amplifier noises, so that they can be considered as coming from different detectors as long as environmental disturbances are negligible.

### **B. Detector design and optimization**

In this section we determine the optimal detector figures on the basis of the analysis developed in Sec. II. As usual, a low dissipation material is required to reduce the effect of



<span id="page-8-0"></span>FIG. 5. Antenna pattern of a hollow cylinder single-mass dual detector (material molybdenum, outer radius  $r_{ext} = 0.5$  m, height  $L = 3$  m, form ratio  $r_{int}/r_{ext} = 0.3$ ,  $\alpha = 0.2$  rad). (a) Antenna pattern of the detector equipped with the readouts  $R_{\oplus}$  and  $R_{\otimes}$ . The antenna pattern is represented by the modulus of the vector to the plotted surface. Per each direction, we show the maximum of the antenna patterns of the readout channels for circularly polarized signals. (b) Section of the antenna pattern across the  $x - y$  plane. This is the worst case for the detector, which is maximally sensitive for GW traveling along the *z* axis; nonetheless, no blind directions show up. The response ranges from 30% to 55% of the optimal case.

the thermal noise. Molybdenum represents an interesting choice [\[33\]](#page-12-28), as it shows low mechanical losses ( $T\phi$  < 5  $\times$  $10^{-9}$  K) for acoustic modes down to 50 mK. Moreover, thanks to its higher density and sound velocity values, molybdenum would allow better sensitivities than an aluminum detector dimensioned to operate in the same frequency range.

In this section we choose to optimize the detector geometry by requiring a specific frequency band, namely, the interval 2–5 kHz. This requirement fixes the frequency of the first two quadrupolar modes as  $v_{2,1} \approx 2$  kHz and  $v_{2,2} \approx$ 5 kHz. As shown in Fig. [6,](#page-8-1) the frequency  $\nu_{2,2}$  depends essentially on the outer radius  $r_{ext}$  and on the material sound velocity  $v_s$ . A frequency  $v_{2,2} \sim 5$  kHz could be obtained with a molybdenum test mass with  $r_{ext}$ 0*:*5 m. The height of the cylindrical test mass is fixed to  $L = 3$  m. The inner radius, through the form factor  $r_{\text{int}}/r_{\text{ext}}$ , determines the frequency of the first quadrupolar mode and the placement of the antiresonance within the bandwidth. In order to better understand the effect of the detector geometry on the sensitivity, we evaluate the sensitivity of the detector for different values of the form factor in the range  $0.1 < r_{\text{in}}/r_{\text{ext}} < 0.8$ .

As discussed in Sec. II C, the optimization process requires a balancing of the noise figures. For this sake we must evaluate as a function of  $\kappa_0$  the minimum detectable value of a specific GW signal by the use of Eq.  $(36)$ . As

<span id="page-8-1"></span>

FIG. 6. Frequency of the normal modes in the plane strain approximation. The left axis shows the reduced value of the frequency, while the right axis shows the actual value of the frequency in the case of a molybdenum test mass with  $r_{ext}$ 0.5 m and  $v_s = 5650$  m/s. The frequency  $v_{2,1}$  of the first quadrupolar mode (flexural mode) is strongly dependent on the form factor, while the frequency  $\nu_{2,2}$  of the second quadrupolar mode (bulk mode) is essentially constant. The antiresonance frequency (a.r.) is evaluated accounting for the contribution of the number of modes needed to obtain a 5% convergence. We point out that while the mode frequencies are solutions of the eigenvalue problem Eq. [\(4](#page-1-2)) and do not depend on the chosen readout, the antiresonance frequency is evaluated specifically for the selective readout shown in Fig.  $1(b)$ . A number of non-GW sensitive modes are also shown as dotted lines.

<span id="page-8-2"></span>burst GW signals are expected in the required frequency range, a sine-Gaussian is considered as test wave form

$$
h(t) = h_0 \sin[\omega_0(t - t_0)]e^{-[(t - t_0)^2/\tau^2]}, \tag{53}
$$

with central frequency  $f_0 = \omega_0/2\pi$  and  $\tau = 2/f_0$ . For GW detectors it is customary to quote performance in terms of the minimum detectable value (at  $S/N = 1$ ) of the root sum square amplitude  $h_{res}$ :

$$
h_{rss} = \sqrt{\int_{-\infty}^{+\infty} dt |h(t)|^2} = \sqrt{\frac{1}{\pi} \int_{0}^{+\infty} d\omega |h(\omega)|^2}.
$$
 (54)

In order to estimate the detector performances on its full bandwidth, four different central frequencies were used for the burst of Eq.  $(53)$ . The calculation also was performed as a function of the amount of dissipation ( $T\phi$  value) considered in the test mass. As shown in Fig. [7](#page-9-0) for the case  $r_{\text{int}}/r_{\text{ext}} = 0.3$  and  $\epsilon_0 = 1$ , the optimal value of the noise stiffness is  $\kappa_0 \approx 1.7 \times 10^{11} \text{ N/m}$ . This number is independent of the  $T\phi$  value and, according to Eq. [\(24\)](#page-3-3), corresponds to a readout displacement power noise of  $S_{xx} \approx 6 \times 10^{-46} \text{ m}^2/\text{Hz}$  and backaction force power noise of  $S_{ff} \approx 1.8 \times 10^{-23} \text{ N}^2/\text{Hz}.$ 

We see in Fig. [7](#page-9-0) that, at the optimal  $\kappa_0$ , the thermal noise induced by a finite dissipation value  $T\phi = 10^{-8}$  K reduces the sensitivity of less than 20%. This dissipation is a good compromise between sensitivity and feasibility and can be assumed as the reference  $T\phi$  value for the case  $\epsilon_0 = 1$ . In the general case, the optimal  $T\phi$  value is determined by the

<span id="page-9-0"></span>

FIG. 7. Minimum detectable (at  $SNR = 1$ ) root sum square amplitude *hrss* of a sine-Gaussian test wave form, impinging the detector at four different values of the central frequency  $f_0$ . The calculation was performed for a detector form factor  $r_{\text{int}}/r_{\text{ext}}=$ 0.3, equipped with a quantum limited readout ( $\epsilon_0 = 1$ ). The value  $\kappa_0 \approx 1.7 \times 10^{11}$  N/m, indicated by the dotted vertical line, gives the best *hrss* value averaged over the four test frequencies. This optimal value of the noise stiffness is independent of the  $T\phi$  value and corresponds to a readout displacement power noise of  $S_{xx} \approx 6 \times 10^{-46} \text{ m}^2/\text{Hz}$  and backaction force power noise of  $S_{ff} \approx 1.8 \times 10^{-23} \text{ N}^2/\text{Hz}.$ 

readout energy sensitivity  $\epsilon_0$ . In fact, according to Eq. ([41\)](#page-4-3), the readout contributed noise grows linearly with  $\epsilon_0$  and we find that a value  $T\phi \leq 10^{-8} \epsilon_0$  K is needed to limit the thermal noise contribution within the 20% of the readout noise. On the contrary, the optimal value of the noise stiffness  $\kappa_0$  is mainly determined by the BA transfer function  $T_{BA}(\omega)$  and remains unchanged.

Figure [8](#page-9-1) shows the detector performances at the optimal noise stiffness averaged over the four test frequencies and then plotted against the form factor. Different values of energy sensitivity were considered with their corresponding optimal value of  $T\phi$ . The noise stiffness was optimized for each point on the graph and found essentially constant

<span id="page-9-1"></span>

FIG. 8. The detector performances for different values of the readout energy sensitivity is averaged over the four test frequencies and shown as a function of the form factor  $r_{\text{in}}/r_{\text{ext}}$ . The error bars indicate the scattering of the sensitivity around the average value, a measure of the uniformity of the performances over the full bandwidth. A dissipation  $T\phi = 10^{-8} \epsilon_0$  K was assumed.

<span id="page-9-2"></span>

FIG. 9. Predicted spectral strain sensitivities of the molybdenum detector: weight 22 ton, outer diameter 1 m, inner diameter 0.3 m, and height 3 m. The curves at different values of the readout energy sensitivity  $\epsilon_0$  are evaluated with optimal values of noise stiffness  $\kappa_0 = 1.7 \times 10^{11}$  N/m and with dissipation  $T\phi = 10^{-8} \epsilon_0$  K. The corresponding displacement noise power of the optimal readout is  $S_{xx} = 6 \times 10^{-46} \epsilon_0 \text{ m}^2/\text{Hz}$  and backaction force power noise of  $S_{ff} \approx 1.8 \times 10^{-23} \epsilon_0 \text{ N}^2/\text{Hz}$ . The curve with  $\epsilon_0 = 1$  represents the detector SQL.

 $\kappa_0 \approx 1.7 \times 10^{11}$  N/m. The dependence on the inner radius is not strong, but the best sensitivity is obtained for 0*:*2 *<*  $r_{\text{in}}/r_{\text{ext}}$  < 0.5, that is when the antiresonance is located approximately in the middle of the bandwidth. The GW sensitivity  $S_{hh}$  of this optimal detector (molybdenum,  $r_{\text{int}} = 0.15 \text{ m}, r_{\text{ext}} = 0.5 \text{ m}, \text{height } 3 \text{ m}$  is shown in Fig. [9.](#page-9-2)

## **IV. CONCLUSIONS**

In this paper we discuss wide band (i.e., nonresonant) acoustic GW detectors in the framework of the elastic theory of continuous bodies, and we take fully into account the three-dimensional properties of signal, test mass, and readout. Once the readout (i.e. its spatial weight function) is chosen, transfer functions are calculated for the GW tidal

force, the backaction noise pressure exerted by the readout and the thermal noise of the test mass. The SQL sensitivity limit of acoustic GW detectors is then derived.

On the basis of this analysis we design a new kind of acoustic GW wide band detector: the single-mass dual detector, formed by a hollow cylinder equipped with two sets of selective readout displaced  $\pi/4$  apart. This detector would offer the same advantages of a dual detector, namely, the kHz-wide frequency range of sensitivity and the backaction reduction effect: moreover, we show that such a detector has no blind directions for impinging GWs, being in this sense omnidirectional even if not isotropic. These findings are analytically derived in the plain strain approximation and then extended to the full threedimensional body by means of FEM numerical analysis. We also show that the single-mass dual detector can measure simultaneously both the amplitude and the polarization of the impinging GW. From the practical viewpoint, the single-mass dual detector removes the technological challenge of nesting two independent resonators which is one of the key issues of the recently proposed dual detectors.

The general treatment of a GW acoustic detector is then applied to the single-mass dual detector in an optimization process that gives the mass dimensions and readout noise properties resulting in the best sensitivity in the 2–5 kHz frequency range. We show that a molybdenum hollow cylinder single-mass dual (outer radius  $r_{ext} = 0.5$  m, height  $L = 3$  m, form ratio  $r_{int}/r_{ext} = 0.3$ , and readout angular aperture  $\alpha = 0.2$  rad) can show a sensitivity at angular aperture  $\alpha = 0.2$  rad) can show a sensitivity at<br>the level of  $1 \div 2 \cdot 10^{-23} / \sqrt{Hz}$  in the frequency interval  $2 \div 5$  kHz and truly free from noise resonances, provided that  $T\phi < 10^{-8}$  K, as already achieved for this material.

The nonresonant readouts for a dual detector would be evolutions both conceptually and in technology of the resonant readouts used in the GW bar detectors: the optomechanical one [\[21\]](#page-12-16), based on Fabry-Perot cavities, and the capacitive one [\[5](#page-12-2)], based on SQUID amplifiers. The main improvements needed to make these readouts suitable in a dual detector configuration can be summarized as

- (a) *The wide band sensitivity.—*The deformation of the test mass which forms the detector needs to be measured by devices that do not show mechanical resonances in the kHz-wide frequency interval of high sensitivity.
- (b) *The selectivity.—*The dual detector properties can be fully exploited by a readout scheme which is geometrically selective to the fundamental quadrupolar modes. This selectivity also is effective in reducing down to a negligible level the acoustic modes not sensitive to GW that would appear in the detector bandwidth. Moreover, the readout system must sense the deformation of the resonant masses on a wide surface, in order to be less sensitive to the acoustic modes resonating at higher frequencies, which do not carry any gravitational signal. In this

way, the thermal noise of the detector is minimized, while preserving the sensitivity to the signal.

Without the resonant amplification stage, the displacement sensitivity of readout employed in running GW bar ment sensitivity of readout employed in running GW bardetectors ranges in the  $10^{-19}$  m/ $\sqrt{Hz}$  decade over a bandwidth of hundreds of Hz centered at about 1 kHz. In order to meet the requirement of a dual detector, an improvement is needed by about a factor of  $10<sup>4</sup>$  and these performances must be extended to cover the wider dual bandwidth. For passive readouts, as the capacitor-SQUID system used in the AURIGA detector  $[2]$  $[2]$ , a factor of  $10<sup>3</sup>$  improvement in the displacement sensitivity is foreseeable after specific R&D devoted to increase the bias field (presently limited to about  $1\%$  of the material breakdown electric field  $[34]$  $[34]$  $[34]$ ) and to increase the capacitor area (the transducer efficiency scales as the square root of the capacitance) and after a reduction of the operating temperature down to the 0.1 K range (the SQUID amplifier additive noise is linearly proportional to the thermodynamic temperature). For active readouts as the optomechanical one [\[21\]](#page-12-16), the readouts as the optomechanical one [21], the<br> $10^{-23}$  m/ $\sqrt{Hz}$  range is reachable in principle by increasing both the finesse (up to  $10<sup>6</sup>$ ) and the light power (up to a few Watts); however, such figures could end up not to be compatible with a cryogenic operation. Therefore, for both technologies we can consider as a reasonable goal a bout technologies we can consider as a reasonable goal a sensitivity at the level of  $10^{-22}$  m/ $\sqrt{Hz}$  over the bandwidth of a dual detector. A factor of 10 further could be gained via broadband mechanical lever amplifiers, as currently under investigation.

The wide-area and selectivity requirements could be satisfied easily by capacitive readouts. In fact, they are currently implemented on surfaces of  $0.05 \text{ m}^2$ , and many capacitors could be properly connected to obtain the required total surface and selectivity, as shown in Ref. [[11\]](#page-12-7). For the optomechanical readout, the implementation of selectivity based on crossed cavities is under study; in order to meet the requirement of the wide area, some of us have proposed a new configuration of the Fabry-Perot optical resonator, the folded Fabry-Perot [[35](#page-12-30)], and its experimental investigation has started already [[36](#page-12-31)].

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## **APPENDIX: MODAL ANALYSIS**

The complex behavior observed in a typical transfer function [Fig.  $10(a)$ ] is not surprising, as it is well known that structures with low modal density [\[37\]](#page-12-32) exhibit a pronounced resonance-antiresonance behavior in their transfer

<span id="page-11-3"></span>

<span id="page-11-2"></span>FIG. 10. Typical transfer function for a four-modes system where the driving force is applied with the same spatial weight of the readout. The resonance-antiresonance behavior is due to the superposition of the single mode responses and, in the hypotheses of low modal density and small damping, an antiresonance is invariably present between the two resonances. (a) Continuous line: transfer function obtained when the *mode 1* and *mode 2* contributions are only considered. Dashed line: transfer function obtained from the contribution of all the four modes; the position of the first antiresonance is changed by the presence of the higher frequency modes contributions. (b) Transfer functions of each mode.

function [[38](#page-12-33)]. Antiresonances provide important information on the dynamic behavior of the structure under consideration and have been applied to several engineering problems, as reported in [[39](#page-12-34)] and references therein.

Considerable insight may be gained by considering the origin of antiresonances. Within the modal expansion model, the system transfer function can be written in the form:

$$
T(\omega) = \sum_{n} \frac{C_n}{(\omega_n^2 - \omega^2) + i D_n(\omega)},
$$
 (A1)

where the explicit form of the modal constants  $C_n$  and the dissipative factor  $D_n$  can be obtained by a straightforward comparison with Eq.  $(13)$  $(13)$  $(13)$ . As usual, the resonant frequencies represent the zeros of the denominator polynomial, that in the case of small damping are approximately given by the eigenvalues  $\omega_n$ . The eigenvalues are determined by the mechanical properties of the system and by the boundary conditions applied when solving Eq. [\(4\)](#page-1-2). The antiresonance frequencies are the zeros of the numerator polynomial resulting from the sum, and their position on the frequency axis depends on the values of the modal constants  $C_n$ . In Fig. [10](#page-11-3), we show  $|T(\omega)|$  as a sum of the contributions of each resonant mode for a four-mode system with  $C_i > 0$ .

Some of the important features of the transfer function can be understood by using a simpler example with just two modes and no dissipation:

$$
T(\omega) = \frac{C_1(\omega_2^2 - \omega^2) + C_2(\omega_1^2 - \omega^2)}{(\omega_2^2 - \omega^2)(\omega_1^2 - \omega^2)}.
$$
 (A2)

If both modal constants are positive, a zero of the numerator polynomial  $\omega_{ar}$  is invariably present in the range  $[\omega_1, \omega_2]$  between the two resonances, at a frequency depending on the relative weight between the modal constants  $C_1$  and  $C_2$ :

$$
\omega_{\rm ar} = \sqrt{\frac{\omega_1^2 + \frac{C_1}{C_2} \omega_2^2}{1 + \frac{C_1}{C_2}}}. \tag{A3}
$$

Positive modal constants appear in all transfer functions where the driving force is applied with the same spatial weight of the readout, as for the BA transfer function  $T<sub>BA</sub>$ [Eq. ([19](#page-2-3))]. In this case there must be an antiresonance following a resonance, without exception and regardless of the complexity of structures [\[38](#page-12-33)]. On the other hand, modal constants with opposite signs result in a smooth minimum between the two resonances, located at

$$
\omega_{\min} = \sqrt{\frac{\omega_1^2}{1 + \sqrt{-\frac{C_1}{C_2}}} + \frac{\omega_2^2}{1 + \sqrt{-\frac{C_2}{C_1}}}}
$$
(A4)

Modal constants of the detector GW transfer function [Eq.  $(16)$ ] can be negative or positive, depending on the integrals performed over the modes. In this case, the sign is essentially determined by the phase relation between a specific mode and the GW excitation force and by the phase relation between this mode and the weight function. Then we expect a mixture of antiresonances and smooth minima.

<span id="page-11-1"></span>We point out that this simple model is only indicative: the actual position of antiresonances and smooth minima between resonances depends on the contribution of the other modes, as shown in Fig. [10.](#page-11-3)

<span id="page-11-0"></span><sup>[1]</sup> ALLEGRO http://gravity.phys.lsu.edu. [2] AURIGA http://www.auriga.lnl.infn.it.

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