# What happens if an unbroken flavor symmetry exists?

Yoshio Koide\*

Department of Physics, University of Shizuoka, 52-1 Yada, Shizuoka 422-8526, Japan (Received 3 August 2004; revised manuscript received 15 November 2004; published 21 January 2005)

Without assuming any specific flavor symmetry and/or any specific mass-matrix forms, it is demonstrated that if an unbroken flavor symmetry exists, we cannot obtain the Cabibbo-Kobayashi-Maskawa (CKM) quark mixing matrix V and the Maki-Nakagawa-Sakata (MNS) lepton mixing matrix U except for those between two families for the case with the completely undegenerated fermion masses, so that we can never give the observed CKM and MNS mixings. Only in the limit of  $m_{\nu 1} = m_{\nu 2}$  ( $m_d = m_s$ ), we can obtain three family mixing with an interesting constraint  $U_{e3} = 0$  ( $V_{ub} = 0$ ).

DOI: 10.1103/PhysRevD.71.016010

PACS numbers: 11.30.Hv, 12.15.Ff, 14.60.Pq

## I. INTRODUCTION

It is well known that the masses of the charged fermions rapidly increase as  $(u, d, e) \rightarrow (c, s, \mu) \rightarrow (t, b, \tau)$ . It has been considered that such rapid increasing of the mass spectra cannot be understood from an idea of "symmetry." However, nowadays, it is a popular idea to understand the observed quark and lepton mass spectra and mixing matrices from assuming a flavor symmetry which puts constraints on the Yukawa coupling constants.

In the present paper, without assuming any explicit flavor symmetry and/or any explicit mass-matrix forms, we will point out that if a flavor symmetry exists, we cannot obtain the observed Cabibbo-Kobayashi-Maskawa [1] (CKM) quark mixing matrix  $V_q$  and Maki-Nakagawa-Sakata [2] (MNS) lepton mixing matrix  $U_{\ell}$ , even if we can obtain reasonable mass spectra under the symmetry. You may think that this conclusion is not so remarkable and rather trivial, because anyone thinks that the flavor symmetry is badly broken. However, most investigations on the broken flavor symmetries are based on specific models, and we are not clearly aware that what problem happens if a flavor symmetry, in general, exists until a low energy scale  $\mu \sim 10^2$  GeV. In some phenomenological mass-matrix models, sometimes, the forms of the symmetry breaking are brought into the model by hand. As we note in the present paper, the flavor symmetry is defined for Yukawa interactions, and not for mass matrices, so that the form of the flavor symmetry breaking in the Yukawa interactions must not break the  $SU(2)_L$  symmetry prior to the spontaneous breakdown of  $SU(2)_L$ . (In the present paper, we discuss the world above  $10^2$  GeV, so that all the conventional fermion masses are exactly zero. Nevertheless, for convenience, we will sometimes use the terminology "mass matrix"  $M_f$  instead of the Yukawa coupling  $Y_f$ and "masses"  $m_{fi}$  ( $f = u, d, e, \nu$ ; i = 1, 2, 3) instead of the eigenvalues of  $Y_f$ .)

Even when we consider a broken flavor symmetry, it is important to consider the world in which the flavor symmetry is exactly unbroken. In the present paper, we will conclude that in such a world with an unbroken flavor symmetry, the CKM and MNS mixing matrices cannot describe flavor mixings except for those between two families when the fermion masses are completely different from each other, and that only when  $m_{\nu 1} = m_{\nu 2}$  ( $m_d = m_s$ ), the MNS matrix  $U_\ell$  (the CKM matrix  $V_q$ ) can describe a three family mixing with an interesting constraint  $(U_\ell)_{e3} = 0$  [ $(V_q)_{ub} = 0$ ]. This will suggest that our world with a broken flavor symmetry should be derived from what unbroken world.

First, let us consider that the up- and down-quark fields transform under a flavor symmetry as

$$u_{L} = U_{XL}^{u} u_{L}^{\prime}, \qquad u_{R} = U_{XR}^{u} u_{R}^{\prime}, \qquad d_{L} = U_{XL}^{d} d_{L}^{\prime},$$
  

$$d_{R} = U_{XR}^{d} d_{R}^{\prime}.$$
(1)

(We do not consider the case in which  $U_{XL}^f$  and  $U_{XR}^f$  are proportional to a unit matrix **1**.) If the Lagrangian is invariant under the transformation (1), the Yukawa coupling constants  $Y_u$  and  $Y_d$  must satisfy the relations

$$(U_{XL}^{u})^{\dagger}Y_{u}U_{XR}^{u} = Y_{u}, \qquad (U_{XL}^{d})^{\dagger}Y_{d}U_{XR}^{d} = Y_{d}, \quad (2)$$

where  $U_{XL}^{u}(U_{XL}^{u})^{\dagger} = \mathbf{1}$ , and so on. Since these transformations must not break SU(2)<sub>L</sub> symmetry, we cannot consider a case with  $U_{XL}^{u} \neq U_{XL}^{d}$ . We must rigorously take

$$U_{XL}^u = U_{XL}^d \equiv U_X. \tag{3}$$

Therefore, the up- and down-quark mass matrices  $M_u = Y_u \langle H_u^0 \rangle$  and  $M_d = Y_d \langle H_d^0 \rangle$  must satisfy the relations

$$U_X^{\dagger} M_u M_u^{\dagger} U_X = M_u M_u^{\dagger}, \qquad U_X^{\dagger} M_d M_d^{\dagger} U_X = M_d M_d^{\dagger},$$
(4)

independently of  $U_{XR}^u$  and  $U_{XR}^d$ . (Exactly speaking, we should read  $M_f M_f^{\dagger}$  as  $Y_f Y_f^{\dagger}$ .)

A similar situation is required in the lepton sectors. Although, sometimes, in the basis where the charged lepton mass-matrix  $M_e$  is diagonal [i.e.,  $M_e = D_e \equiv \text{diag}(m_e, m_\mu, m_\tau)$ ], a "symmetry" for the neutrino mass-matrix  $M_\nu$  is investigated, such a prescription cannot be

<sup>\*</sup>Email address: koide@u-shizuoka-ken.ac.jp

### YOSHIO KOIDE

regarded as a field theoretical symmetry. For example, when we assume a permutation symmetry between neutrinos  $\nu_{L2}$  and  $\nu_{L3}$ , we can obtain a nearly bimaximal mixing [3]. However, the symmetry is applied only to neutrino sector  $M_{\nu}$ , and not to the charged lepton sector  $M_e = D_e$ . Therefore, we cannot regard this  $2 \leftrightarrow 3$  permutation rule as a symmetry in the field theoretical meaning, because it is badly broken the SU(2)<sub>L</sub> symmetry. (Of course, as we comment later, by assuming a specific mechanism, it is possible to introduce different flavor-symmetry-breaking forms between  $M_e$  and  $M_{\nu}$ . However, in most phenomenological models, the mechanism has not been given explicitly.)

In the lepton sectors, we must consider that under the transformations

$$\nu_{L} = U_{X}\nu'_{L}, \qquad \nu_{R} = U^{\nu}_{XR}\nu'_{R}, \qquad e_{L} = U_{X}e'_{L}, 
e_{R} = U^{e}_{XR}e'_{R},$$
(5)

the Yukawa coupling constants which are defined by  $\overline{e}_L Y_e e_R$ ,  $\overline{\nu}_L Y_D^{\nu} e_R$ , and  $\overline{\nu}_R^c Y_M^{\nu} \nu_R$  ( $\nu_R^c \equiv C \overline{\nu}_R^T$ ) are invariant as follows:

$$U_{X}^{\dagger}Y_{e}U_{XR}^{e} = Y_{e}, \qquad U_{X}^{\dagger}Y_{D}^{\nu}U_{XR}^{\nu} = Y_{D}^{\nu}, \qquad (6)$$

$$U_{XR}^T Y_M^{\nu} U_{XR} = Y_M^{\nu}.$$

In other words, the mass matrices  $M_e M_e^{\dagger}$  and  $M_{\nu}$  are invariant under the transformation  $U_X$  as

$$U_X^{\dagger} M_e M_e^{\dagger} U_X = M_e M_e^{\dagger}, \tag{7}$$

$$U_X^{\dagger} M_{\nu} U_X^* = M_{\nu}, \tag{8}$$

independently of the forms  $U_{XR}^{\nu}$  and  $U_{XR}^{e}$ , where we assumed the seesaw mechanism [4]  $M_{\nu} \propto Y_{D}^{\nu}(Y_{M}^{\nu})^{-1}(Y_{D}^{\nu})^{T}$ . [Even when we do not assume the seesaw mechanism, as long as the effective neutrino mass matrix is given by  $\overline{\nu}_{L}M_{\nu}\nu_{L}^{e}$ , the mass matrix must obey the constraint (8).]

Note that the constraints (4) [and also (7) and (8)] do not always mean that the matrix forms  $M_u M_u^{\dagger}$  and  $M_d M_d^{\dagger}$  are identical each other. Indeed, in the present paper, we consider a general case in which the eigenvalues and mixing matrices between  $M_u M_u^{\dagger}$  and  $M_d M_d^{\dagger}$  are different from each other. Nevertheless, the conditions (4) [and also (7) and (8)] will put very strong constraints on the CKM mixing matrix  $V_q = (U_L^u)^{\dagger} U_L^d$  [and also the MNS mixing matrix  $U_{\ell} = (U_L^e)^{\dagger} U_L^{\nu}$ ], where  $U_L^f$  ( $f = u, d, e, \nu$ ) are defined by

$$(U_L^f)^{\dagger} M_f M_f^{\dagger} U_L^f = D_f^2 \equiv \text{diag}(m_{f1}^2, m_{f2}^2, m_{f3}^2)$$
  
(f = u, d, e), (9)

$$(U_L^{\nu})^{\dagger} M_{\nu} (U_L^{\nu})^* = D_{\nu} \equiv \operatorname{diag}(m_{\nu 1}, m_{\nu 2}, m_{\nu 3}).$$
(10)

The purpose of the present paper is to see whether or not it is possible to consider such the flavor symmetry without an  $SU(2)_L$  symmetry breaking. Of course, further conditions

$$(U_{XR}^{f})^{\dagger} M_{f}^{\dagger} M_{f} U_{XR}^{f} = M_{f}^{\dagger} M_{f} \qquad (f = u, d, e),$$
(11)

will give more strict constraints on the mass matrices  $M_f$ . However, even apart from such an additional constraint, by using only the constraints (4), (7), and (8) we will obtain a severe conclusion that such a symmetry cannot lead to the observed CKM mixing matrix  $V_q$  and MNS mixing matrix  $U_{\ell}$ .

## II. OBSTACLES IN THE CKM AND MNS MIXING MATRICES

First, we investigate relations in the quark sectors under the conditions (4). Since we can rewrite the left-hand of Eq. (9) by using Eq. (4) as

$$(U_L^f)^{\dagger} M_f M_f^{\dagger} U_L^f = (U_L^f)^{\dagger} U_X^{\dagger} M_f M_f^{\dagger} U_X U_L^f$$
$$= (U_L^f)^{\dagger} U_X^{\dagger} U_L^f D_f^2 (U_L^f)^{\dagger} U_X U_L^f, \quad (12)$$

for f = u, d, we obtain the relation

$$(U_X^f)^{\dagger} D_f^2 U_X^f = D_f^2, \tag{13}$$

where

$$U_X^f = (U_L^f)^\dagger U_X U_L^f.$$
(14)

Therefore, the matrix  $U_X^f$  which satisfies Eq. (13) must be a diagonal matrix with a form

$$U_X^f = P_X^f \equiv \operatorname{diag}(e^{i\delta_1^f}, e^{i\delta_2^f}, e^{i\delta_3^f}), \qquad (15)$$

unless the masses are not degenerated. Therefore, from (14), we obtain

$$U_X = U_L^u P_X^u (U_L^u)^{\dagger} = U_L^d P_X^d (U_L^d)^{\dagger},$$
(16)

which leads to a constraint on the CKM matrix  $V_q = (U_I^u)^{\dagger} U_I^d$ :

$$P_X^u = V_q P_X^d (V_q)^{\dagger}. \tag{17}$$

The constraint (17) (i.e.,  $P_X^u V_q = V_q P_X^d$ ) requires

$$(e^{i\delta_i^u} - e^{i\delta_j^d})(V_q)_{ij} = 0 \qquad (i, j = 1, 2, 3).$$
(18)

Only when  $\delta_i^u = \delta_j^d$ , we can obtain  $(V_q)_{ij} \neq 0$ . For the case  $\delta_1^u = \delta_2^u = \delta_3^u \equiv \delta_u$  (also  $\delta_1^d = \delta_2^d = \delta_3^d \equiv \delta_d$ ), the matrix  $P_X^u = \mathbf{1}e^{i\delta_u}$  (and also  $P_X^d = \mathbf{1}e^{i\delta_d}$ ) leads to a trivial result  $U_X = \mathbf{1}e^{i\delta_u} = \mathbf{1}e^{i\delta_d}$ ), so that we do not consider such a case. Therefore, from the requirement (17), we cannot consider such a case as all elements of  $V_q$  are not zero. For example, if we can take  $(V_q)_{ii} \neq 0$  for i = 1, 2, 3 by taking  $\delta_i^u = \delta_i^d \equiv \delta_i$ , since we must choose, at least, one of  $\delta_i$  differently from others, we obtain a mixing

matrix between only two families, e.g.,  $(V_q)_{13} = (V_q)_{31} =$  $(V_q)_{23} = (V_q)_{32} = 0$  for a case of  $\delta_1 = \delta_2 \neq \delta_3$ :

$$V = \begin{pmatrix} * & * & 0\\ * & * & 0\\ 0 & 0 & 1 \end{pmatrix}.$$
 (19)

Thus, for any choice of  $\delta_i^u$  and  $\delta_j^d$ , condition (18) cannot lead to the observed CKM mixing matrix.

For the lepton sectors, the situation is the same. From Eqs. (8) and (10), we obtain the constraint

$$(U_X^{\nu})^{\dagger} D_{\nu} (U_X^{\nu})^* = D_{\nu}, \qquad (20)$$

where

$$U_X^{\nu} = (U_L^{\nu})^{\dagger} U_X U_L^{\nu}.$$
 (21)

Again, if we assume that the neutrino masses are not degenerated, we obtain that the matrix  $U_X^{\nu}$  must be diagonal, and it is given by

$$U_X^{\nu} = P_X^{\nu} \equiv \operatorname{diag}(e^{i\delta_1^{\nu}}, e^{i\delta_2^{\nu}}, e^{i\delta_3^{\nu}}), \qquad (22)$$

because the constraint (20) leads to

$$(m_{\nu i}e^{-i\phi_{ij}} - m_{\nu j}e^{i\phi_{ij}})|(U_X^{\nu})_{ij}| = 0, \qquad (23)$$

where we have put  $(U_X^{\nu})_{ij} = |(U_X^{\nu})_{ij}|e^{i\phi_{ij}}$ . Here, differently from the matrix (15), the phases  $\delta_i^{\nu}$  are constrained as  $\delta_i^{\nu} = 0$  or  $\delta_i^{\nu} = \pi$  (*i* = 1, 2, 3) from the condition (20). From the relations (14) and (21), we obtain

$$U_X = U_L^e P_X^e (U_L^e)^{\dagger} = U_L^{\nu} P_X^{\nu} (U_L^{\nu})^{\dagger}, \qquad (24)$$

so that the MNS matrix  $U_{\ell} = (U_L^e)^{\dagger} U_L^{\nu}$  must satisfy the constraint

$$P_X^e = U_\ell P_X^\nu (U_\ell)^\dagger, \tag{25}$$

i.e.,

$$(e^{i\delta_i^e} - e^{i\delta_j^\nu})(U_\ell)_{ij} = 0 \qquad (i, j = 1, 2, 3).$$
(26)

Again, only when  $\delta_i^e = \delta_j^\nu$ , we can obtain  $(U_\ell)_{ij} \neq 0$ , and

V =

$$\begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - s_{23}c_{12}s_{13}e^{i\delta} & c_{23}c_{12} - s_{23}s_{12}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{23}s_{12} - c_{23}c_{12}s_{13}e^{i\delta} & -s_{23}c_{12} - c_{23}s_{12}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix},$$

and  $P_M$  is a Majorana phase matrix

$$P_M = \text{diag}(e^{i\alpha}, e^{i\beta}, e^{i\gamma}). \tag{32}$$

For the case with the form (27) of  $U_X^{\nu}$ , we obtain

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{12} = -c_{13}\{c_{23}(c_{12}^{2}e^{-i\phi} + s_{12}^{2}e^{i\phi})s + s_{13}s_{23}[(e^{i\phi} - e^{-i\phi})c_{12}s_{12}s + c - 1]e^{-i\delta}\},$$
(33)

we cannot consider a case in which all elements of  $U_{\ell}$  are not zero. We only obtain a mixing matrix between two families.

Thus, the requirements (4) [and also (7) and (8)] lead to a serious trouble in the CKM matrix  $V_q$  (the MNS matrix  $U_{\ell}$ ), even if we can suitably give the observed mass spectra. The similar conclusion has already been derived by Low and Volkas [5] although they have demonstrated it by using explicit mass-matrix forms.

## III. CASE OF $m_{\nu 1} = m_{\nu 2}$

In order to seek for a clue to a possible symmetry breaking, let us go on a phenomenological study.

Since the observed neutrino data [6-9] have shown  $\Delta m_{\rm solar}^2 \ll \Delta m_{\rm atm}^2$ , it is interesting to consider a limit of  $m_{\nu 1} = m_{\nu 2}$ . In this case, the conclusion (22) [and (26)] is not correct any more, because the constraint (20) allows a case with  $(U_X^{\nu})_{12} \neq 0$  and  $(U_X^{\nu})_{21} \neq 0$ :

$$U_X^{\nu} = \begin{pmatrix} c & -s & 0\\ s & c & 0\\ 0 & 0 & 1 \end{pmatrix}, \tag{27}$$

or

$$U_X^{\nu} = \begin{pmatrix} -c & s & 0\\ s & c & 0\\ 0 & 0 & 1 \end{pmatrix},$$
(28)

where  $c = \cos\theta$  and  $s = \sin\theta$ . (Again, each element must be real.) Therefore, we must check the relation

$$P_X^e = U_\ell U_X^\nu (U_\ell)^\dagger, \tag{29}$$

with the forms of (27) and (28) of  $U_X^{\nu}$ , instead of (22).

Now, we explicitly calculate  $U_{\ell}U_{\chi}^{\nu}(U_{\ell})^{\dagger}$  by using a general form of  $U_{\ell}$ 

$$U_{\ell} = V P_M, \tag{30}$$

where

$$\begin{array}{ccc} c_{13}s_{12} & s_{13}e^{-i\delta} \\ {}_{3}e^{i\delta} & c_{23}c_{12} - s_{23}s_{12}s_{13}e^{i\delta} & s_{23}c_{13} \\ e^{i\delta} & -s_{23}c_{12} - c_{23}s_{12}s_{13}e^{i\delta} & c_{23}c_{13} \end{array} \right),$$
(31)

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{13} = c_{13}\{s_{23}(c_{12}^{2}e^{-i\phi} + s_{12}^{2}e^{i\phi})s - s_{13}c_{23}[(e^{i\phi} - e^{-i\phi})c_{12}s_{12}s + c - 1]e^{-i\delta}\},$$
(34)

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{23} = c_{23}s_{23}[(e^{i\phi} - e^{-i\phi})c_{12}s_{12}(1 + s_{13}^{2})s + c_{13}^{2}(1 - c)] - s_{13}s[s_{23}^{2}(c_{12}^{2}e^{-i\phi} + s_{12}^{2}e^{i\phi})e^{i\delta} + c_{23}^{2}(c_{12}^{2}e^{i\phi} + s_{12}^{2}e^{-i\phi})e^{-i\delta}],$$
(35)

#### YOSHIO KOIDE

where  $\phi = \beta - \alpha$ . If  $c_{13} \neq 0$ , there is no solution which gives zeros for all the elements (33)–(35), except for a trivial solution with c = 1 (i.e.,  $U_X = \mathbf{1}$ ). If  $c_{13} = 0$ , there is a solution for suitable choice of  $\phi$  and  $\delta$ , and then, the matrix V takes the form

$$V = \begin{pmatrix} 0 & 0 & e^{-i\delta} \\ * & * & 0 \\ * & * & 0 \end{pmatrix}.$$
 (36)

Of course, the form (36) is ruled out. Thus, the case (27) cannot lead to any interesting form of  $U_{\ell}$ .

On the other hand, for the case (28), we obtain

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{12} = c_{13}\{c_{23}[(c_{12}^{2}e^{-i\phi} - s_{12}^{2}e^{i\phi})s + 2c_{12}s_{12}c] + s_{13}s_{23}[1 + (c_{12}^{2} - s_{12}^{2})c - (e^{i\phi} + e^{-i\phi})c_{12}s_{12}s]e^{-i\delta}\},$$
(37)

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{13} = -c_{13}\{s_{23}[(c_{12}^{2}e^{-i\phi} - s_{12}^{2}e^{i\phi})s + 2c_{12}s_{12}c] \\ - s_{13}c_{23}[1 + (c_{12}^{2} - s_{12}^{2})c \\ - (e^{i\phi} + e^{-i\phi})c_{12}s_{12}s]e^{-i\delta}\},$$
(38)

$$(U_{\ell}U_{X}^{\nu}U_{\ell}^{\dagger})_{23} = c_{23}s_{23}[(e^{i\phi} + e^{-i\phi})c_{12}s_{12}(1 + s_{13}^{2})s + c_{13}^{2} - (c_{12}^{2} - s_{12}^{2})(1 + s_{13}^{2})c] + s_{13}s_{23}^{2}[(c_{12}^{2}e^{-i\phi} - s_{12}^{2}e^{i\phi})s + 2c_{12}s_{12}c]e^{i\delta} - s_{13}c_{23}^{2}[(c_{12}^{2}e^{i\phi} - s_{12}^{2}e^{-i\phi})s + 2c_{12}s_{12}c]e^{-i\delta}.$$
(39)

The case can lead to a nontrivial solution for  $s_{13} = 0$ ,  $\phi = \beta - \alpha = 0$ , and

$$\cos(2\theta_{12} + \theta) = 1, \tag{40}$$

i.e.,

$$U_{\ell} = \begin{pmatrix} c_{12} & s_{12} & 0\\ -c_{23}s_{12} & c_{23}c_{12} & s_{23}\\ s_{23}s_{12} & -s_{23}c_{12} & c_{23} \end{pmatrix} P_{M}.$$
 (41)

It should be noted that in the limit of  $m_{\nu 1} = m_{\nu 2}$ , the Majorana phases in  $P_M$  must be  $\alpha = \beta$ .

The similar result  $(U_{\ell})_{13} = 0$  has also been derived by Low and Volkas [5] although their interest was in the "trimaximal mixing" and they have assumed a specific flavor symmetry. In the present general study, we can obtain  $s_{13} = 0$ , but  $s_{12}$  and  $s_{23}$  are still free. The result (41) is a conclusion which is derived model-independently.

Note that the case (28) satisfies  $(U_X^{\nu})^2 = 1$ , so that the flavor transformation  $U_X$  also satisfies

$$(U_X)^2 = \mathbf{1}.$$
 (42)

This suggests that an approximate flavor symmetry in the lepton sectors is a discrete symmetry  $Z_2$ .

Inversely, for the neutrino mass spectra with  $m_{\nu 1} \neq m_{\nu 2}$ , if we take the operator  $U_X = U_L^{\nu} U_X^{\nu} (U_L^{\nu})^{\dagger}$  with the form (28) of  $U_X^{\nu}$ , we obtain

$$(U_X^{\nu})^{\dagger} D_{\nu} (U_X^{\nu})^* = D_{\nu} + (m_{\nu 2} - m_{\nu 1}) s \begin{pmatrix} s & c & 0 \\ c & -s & 0 \\ 0 & 0 & 0 \end{pmatrix},$$
(43)

which leads to

$$U_X^{\dagger} M_{\nu} U_X^* = M_{\nu} + (m_{\nu 2} - m_{\nu 1}) sB, \qquad (44)$$

where the symmetry breaking term B is given by

$$B = U_L^{\nu} \begin{pmatrix} s & c & 0 \\ c & -s & 0 \\ 0 & 0 & 0 \end{pmatrix} (U_L^{\nu})^T.$$
(45)

The matrix B is rewritten as

$$B = U_L^e \begin{pmatrix} 0 & c_{23} & -s_{23} \\ c_{23} & 0 & 0 \\ -s_{23} & 0 & 0 \end{pmatrix} (U_L^e)^{\dagger}, \qquad (46)$$

by using the relation  $U_X = U_L^e P_X^e (U_X^e)^{\dagger}$  and the constraint (40). Of course, the result (45) shows that in the limit of  $m_{\nu 1} = m_{\nu 2}$  and/or s = 0, the operation  $U_X$  becomes that of the exact symmetry. The forms (45) and (46) of the symmetry breaking term will give a clue to a possible form of the flavor symmetry breaking. However, in order to fix the values of  $s_{23}$  and  $s_{12}$  (or *s*), we must put a further assumption. In the present paper, we do not give such a speculation any more.

If we apply the similar discussion to the quark sector in the limit of  $m_d = m_s$ , we can obtain  $|V_{ub}| = 0$ . This may be taken as the reason of  $|V_{ub}|^2 \ll |V_{cb}|^2$ ,  $|V_{us}|^2$ .

### **IV. CONCLUDING REMARKS**

In conclusion, we have noticed that when we assume a flavor symmetry, we must use the same operation  $U_X$  simultaneously for the up-quarks  $u_{Li}$  and down-quarks  $d_{Li}$  (and also for the charged leptons  $e_{Li}$  and neutrinos  $\nu_{Li}$ ), and we have demonstrated that the existence of such an operation  $U_X$  without an SU(2)<sub>L</sub> breaking leads to some serious obstacles:

- (i) If the masses  $(m_u, m_c, m_t)$  and  $(m_d, m_s, m_b)$ [ $(m_e, m_\mu, m_\tau)$  and  $(m_{\nu 1}, m_{\nu 2}, m_{\nu 3})$ ] are exactly different from each others (here, the masses means the eigenvalues of the Yukawa coupling  $Y_f$ ), we cannot obtain the CKM mixing matrix  $V_q = (U_L^u)^{\dagger} U_L^d$  [the MNS mixing matrix  $U_\ell (U_L^e)^{\dagger} U_L^\nu$ ] except for two family mixing.
- (ii) Only when  $m_{\nu 1} = m_{\nu 2} (m_d = m_s)$ , the MNS matrix  $U_\ell$  (the CKM matrix  $V_q$ ) can describe a three family mixing with an interesting constraint  $(U_\ell)_{e3} = 0$   $[(V_q)_{ub} = 0]$ .

### WHAT HAPPENS IF AN UNBROKEN FLAVOR SYMMETRY ...

Thus, if we consider that any flavor symmetry survives keeping unbroken until  $10^2$  GeV, we meet the serious obstacles mentioned above. It is one way to adopt a model with no flavor symmetry in order to evade the present severe conclusions. However, we know the fact (the degree of freedom of "rebasing") that we cannot physically distinguish two mass-matrix sets  $(M_u, M_d)$  and  $(M'_u, M'_d)$ , where  $(M'_u, M'_d)$  is obtained from  $(M_u, M_d)$  by a common flavor-basis rotation for the  $SU(2)_L$  doublet fields. (The situation is the same in the lepton sector.) Only when there is a flavor symmetry, the mass-matrix forms  $(M_u, M_d)$  in a specific flavor basis have a meaning, because the operator of the flavor rotation does not commute with the flavorsymmetry operator  $U_X$ . Therefore, the idea of a flavor symmetry is still attractive to most mass-matrix-modelbuilders.

If we want to investigate a model with a flavor symmetry, our results (18) and (26) demand that the flavor symmetry should be completely broken at a high energy scale  $M_X$ , so that we cannot have any flavor symmetry below  $\mu = M_X$ . We have to seek for a flavor symmetry breaking mechanism under the condition that the original Lagrangian (including the symmetry breaking mechanism) is exactly invariant under the SU(2)<sub>L</sub>.

For example, let us consider a two Higgs doublet model, or a  $\overline{5}_L \leftrightarrow \overline{5}'_L$  model [10]. In such a model, the effective Yukawa coupling constants  $Y^f$  below  $\mu = M_X$  are given by a linear combination of two Yukawa coupling constants with different textures  $Y^f_A$  and  $Y^f_B$ ,

$$Y^f = c^f_A Y^f_A + c^f_B Y^f_B, (47)$$

so that  $Y^f(Y^f)^{\dagger}$  do not satisfy the flavor-symmetry condition

$$U_X^{\dagger} Y^f (Y^f)^{\dagger} U_X = Y^f (Y^f)^{\dagger}, \qquad (48)$$

although  $Y_A^f(Y_A^f)^{\dagger}$  and  $Y_B^f(Y_B^f)^{\dagger}$  must satisfy the conditions

$$U_{XA}^{\dagger}Y_{A}^{f}(Y_{A}^{f})^{\dagger}U_{XA} = Y_{A}^{f}(Y_{A}^{f})^{\dagger}, U_{XB}^{\dagger}Y_{B}^{f}(Y_{B}^{f})^{\dagger}U_{XB}$$
$$= Y_{B}^{f}(Y_{B}^{f})^{\dagger}, \qquad (49)$$

respectively, even if at  $\mu < M_X$ . In other words, there is no operator  $U_X$  which satisfies the condition (48). Thus, we can break the flavor symmetry without breaking the SU(2)<sub>L</sub> symmetry. Of course, we should note that the matrices  $Y_A^f$  and  $Y_B^f$  have to satisfy the conditions (49) simultaneously for f = u and f = d [f = e and  $f = \nu$ ]. For currently proposed phenomenological mass-matrix models with a flavor symmetry breaking, it is important to check whether the mass-matrix forms with a broken flavor symmetry are still invariant or not under the SU(2)<sub>L</sub> symmetry.

Finally, we would like to comment on the results (41) and (42) in the case of  $m_{\nu 1} = m_{\nu 2}$ . This suggests a possibility that we can reasonably understand the observed smallness of  $|(U_\ell)_{13}|$  [11] and  $|(V_q)_{ub}|$  [12] if we consider a model with a flavor symmetry of  $Z_2$  type,  $(U_X)^2 = 1$ , and with  $m_{\nu 1} = m_{\nu 2}$  and  $m_{d1} = m_{d2}$  at  $\mu > M_X$ . This will give a promising clue to possible features of the unbroken flavor symmetry at  $\mu > M_X$  and its symmetry breaking form.

# ACKNOWLEDGMENTS

The author would like to thank J. Sato for helpful conversations. This work was supported by the Grant-in-Aid for Scientific Research, the Ministry of Education, Science and Culture, Japan (Grant No. 15540283).

- N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963); M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
- [2] Z. Maki, M. Nakagawa, and S. Sakata, Prog. Theor. Phys. 28, 870 (1962); B. Pontecorvo, Pis'ma Zh. Eksp. Theor. Fiz. 33, 549 (1957); Sov. Phys. JETP 26, 984 (1968).
- [3] T. Fukuyama and H. Nishiura, in *Proceedings of the International Workshop on Masses and Mixings of Quarks and Leptons, Shizuoka, Japan, 1997*, edited by Y. Koide (World Scientific, Singapore, 1998), p. 252; E. Ma and M. Raidal, Phys. Rev. Lett. 87, 011802 (2001); C. S. Lam, Phys. Lett. B 507, 214 (2001); K. R. S. Balaji, W. Grimus, and T. Schwetz, Phys. Lett. B 508, 301 (2001); W. Grimus and L. Lavoura, Acta Phys. Pol. B 32, 3719 (2001).
- [4] T. Yanagida, in *Proceedings of the Workshop on the* Unified Theory and Baryon Number in the Universe, edited by O. Sawada and A. Sugamoto (KEK, Tsukuba,

1979), p. 95; M. Gell-Mann, P. Ramond, and R. Slansky, in *Supergravity*, edited by P. van Nieuwenhuizen and D. Freedman (North-Holland, Amsterdam, 1979), p. 315; G. Senjanović and R. N. Mohapatra, Phys. Rev. Lett. **44**, 912 (1980); J. Schchter and J. W. F. Valle, Phys. Rev. D **22**, 2227 (1980).

- [5] C.I. Low and R.R. Volkas, Phys. Rev. D 68, 033007 (2003).
- [6] J. N. Bahcall, M. C. Gonzalez-Garcia, and C. Penã-Garay, J. High Energy Phys. 0108 (2001) 014; G. L. Fogli, E. Lisi, D. Montanino, and A. Palazzo, Phys. Rev. D 64, 093007 (2001); V. Barger, D. Marfatia, and K. Whisnant, Phys. Rev. Lett. 88, 011302 (2002); P. I. Krastev and A. Yu. Smirnov, Phys. Rev. D 65, 073022 (2002); SNO Collaboration, Q. R. Ahmad *et al.*, Phys. Rev. Lett. 89, 011302 (2002); P. C. de Holanda and A. Yu. Smirnov, Astropart. Phys. 21, 287 (2004); A. Bandyopadhyay *et al.*, Phys. Lett. B 583, 134 (2004); G. L. Fogli *et al.*,

Phys. Lett. B 583, 149 (2004).

- [7] KamLAND Collaboration, K. Eguchi *et al.*, Phys. Rev. Lett. **90**, 021802 (2003); **92**, 071301 (2004).
- [8] Super-Kamiokande Collaboration, Y. Fukuda *et al.*, Phys. Rev. Lett. **81**, 1562 (1998); M. Shiozawa, in Proceedings of Neutrino 2002, Munich, Germany, 2002 (http:// neutrino.t30.physik.tu-muenchen.de/); C. Saji, in Proceedings of the 5th Workshop on "Neutrino Oscillations and their Origin" (NOON2004), Tokyo,

Japan, 2004 (http://www-sk.icrr.u-tokyo.ac.jp/noon2004/).

- [9] K2K Collaboration, M. H. Ahn *et al.*, Phys. Rev. Lett. **90**, 041801 (2003).
- [10] M. Bando and T. Kugo, Prog. Theor. Phys. 101, 1313 (1999). Also see J. Sato and T. Yanagida, Nucl. Phys. Proc. Suppl. 77, 293 (1999).
- [11] M. Apollonio et al., Phys. Lett. B 466, 415 (1999).
- [12] Particle Data Group, S. Eidelman *et al.*, Phys. Lett. B **592**, 1 (2004) (http://pdg.lbl.gov).