## Neutrino mass and the $SU(2)_R$ breaking scale

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The neutrino sector in a left-right extension of the standard model depends on how  $SU(2)_R$  is broken. I list all possible scenarios, including the ones where the Majorana  $\nu_R$  mass is naturally much smaller than the  $SU(2)_R$  breaking scale, which is desirable for generating the proper baryon asymmetry of the Universe. The best such choice is identified and discussed.

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In the standard model of particle interactions, the neutrino is part of a left-handed doublet  $(\nu, l)_L$  under  $SU(2)_L$  $\times U(1)_Y$ . Whereas the charged lepton must have a righthanded singlet counterpart  $l_R$ , the singlet  $\nu_R$  is not mandatory [because it is trivial under  $SU(2)_L \times U(1)_Y$ ] and is absent in the minimal version of the model. On the other hand, its existence is usually assumed so that  $\nu_L$  may acquire a naturally small Majorana mass as  $\nu_R$  gets a large Majorana mass [again because it is trivial under  $SU(2)_L \times U(1)_Y$ ] in the famous canonical seesaw mechanism [1,2]. Where does  $\nu_R$  come from and what is the magnitude of its Majorana mass? The simplest answer [2] is that  $U(1)_Y$  is actually a remnant of  $SU(2)_R \times U(1)_{B-L}$  under which  $(\nu, l)_R$  is a doublet, and the large Majorana  $\nu_R$  mass comes from the vacuum expectation value (VEV) of a scalar  $SU(2)_R$  triplet, which also breaks  $SU(2)_R \times U(1)_{B-L}$  to  $U(1)_Y$ . This scenario has dominated the thinking on neutrino mass for over 20 years, but it is not the only possibility, even if the existence of  $\nu_R$  is conceded. (Mechanisms without  $\nu_R$  are also possible and just as natural [3].) It may not even be the best possibility as far as leptogenesis [4] is concerned, because the  $SU(2)_R$  gauge interactions will tend to diminish the  $\nu_R$ number density in the early Universe.

Under  $SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ , the quarks and leptons transform as

$$q_L = (u,d)_L \sim (3,2,1,1/3), \tag{1}$$

$$q_R = (u,d)_R \sim (3,1,2,1/3), \tag{2}$$

$$l_L = (\nu, e)_L \sim (1, 2, 1, -1), \tag{3}$$

$$l_R = (\nu, e)_R \sim (1, 1, 2, -1), \tag{4}$$

where the electric charge is given by

$$Q = I_{3L} + I_{3R} + \frac{1}{2}(B - L).$$
(5)

To break  $SU(2)_R \times U(1)_{B-L}$  to  $U(1)_Y$ , there are two possibilities. One is to use the scalar doublet

$$\Phi_R = (\phi_R^+, \phi_R^0) \sim (1, 1, 2, 1), \tag{6}$$

the other is to use the scalar triplet

$$\xi_R = (\xi_R^{++}, \xi_R^{+}, \xi_R^{0}) \sim (1, 1, 3, 2).$$
(7)

The subsequent breaking of  $SU(2)_L \times U(1)_Y$  to  $U(1)_{em}$  may be achieved with either a scalar doublet

$$\Phi_L = (\phi_L^+, \phi_L^0) \sim (1, 2, 1, 1), \tag{8}$$

or a scalar bidoublet

$$\eta = \begin{pmatrix} \eta_1^0 & \eta_2^+ \\ \eta_1^- & \eta_2^0 \end{pmatrix} \sim (1, 2, 2, 0), \tag{9}$$

where  $I_{3L} = 1/2, -1/2$  for the rows, and  $I_{3R} = -1/2, 1/2$  for the columns. The existence of a scalar triplet

$$\xi_L = (\xi_L^{++}, \xi_L^{+}, \xi_L^{0}) \sim (1, 3, 1, 2) \tag{10}$$

may also be contemplated but its VEV must be much smaller than that of  $\Phi_L$  or  $\eta$  to be consistent with the precisely determined values of  $\sin^2 \theta_W$  and the masses of the W and Z bosons. Neutrino masses are sensitive to which of these 5 scalars are chosen, resulting in 5 basic scenarios, as described below.

(i)  $\xi_R + \eta$ . This is the canonical scenario where  $\nu_L$  pairs up with  $\nu_R$  through the VEV's of the bidoublet  $\eta$  to form a Dirac mass  $m_D$  while  $\nu_R$  picks up a large Majorana mass  $m_R$ through the VEV of the  $SU(2)_R$  triplet  $\xi_R$ . The famous seesaw mass matrix

$$\mathcal{M}_{\nu} = \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix} \tag{11}$$

is obtained, with  $m_R$  of order the  $SU(2)_R$  breaking scale. The zero of this matrix comes from the fact that there is no  $\xi_L$ .

(ii)  $\xi_R + \eta + \xi_L$ . This is the canonical left-right symmetric scenario, where  $\xi_L \leftrightarrow \xi_R$  is often imposed as a symmetry of the theory. Since the VEV of  $\xi_L$  contributes to the Majorana  $\nu_L$  mass, the neutrino mass matrix of Eq. (11) becomes

$$\mathcal{M}_{\nu} = \begin{pmatrix} m_L & m_D \\ m_D & m_R \end{pmatrix}. \tag{12}$$

This means that the canonical seesaw formula is corrected to read

$$m_{\nu} = m_L - \frac{m_D^2}{m_P}.$$
 (13)

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However  $m_L$  is routinely argued to be small because  $\langle \xi_L^0 \rangle$  is of order  $\langle \eta_1^0 \rangle \langle \eta_2^0 \rangle / \langle \xi_R^0 \rangle$  provided  $m_{\xi_L}^2$  is positive and of order  $v_R^2$ . In this case,  $m_L$  may be larger or smaller than  $m_D^2/m_R$ , or the two terms may be of comparable magnitude. Leptogenesis in this scenario has been studied in a number of recent papers [5].

For the many practitioners of the canonical seesaw mechanism,  $m_L$  is implicitly assumed to be negligible. On the other hand, if  $m_L$  is the dominant term, then  $\nu_R$  may be dispensed with. In other words, we have just the standard  $SU(2)_L \times U(1)_Y$  model with the simple addition of a Higgs triplet [6]. Again assuming  $m_{\xi_L}^2$  to be positive and large, we have [7]

$$\left< \xi_L^0 \right> = -\,\mu \left< \phi_L^0 \right>^2 / m_{\xi_I}^2 \,, \tag{14}$$

where  $\mu$  is the  $\xi_L^{\dagger} \Phi_L \Phi_L$  coupling. This mechanism without any  $\nu_R$  is also a completely satisfactory explanation of the smallness of  $m_{\nu}$ .

(iii)  $\Phi_R + \eta$ . Here the VEV of  $\Phi_R$  breaks  $SU(2)_R \times U(1)_{B-L}$  to  $U(1)_Y$ , and all fermions obtain Dirac masses from  $\eta$ . Since there is no  $\xi_R$  or  $\xi_L$ , the neutrino is apparently a Dirac particle in this scenario. Thus  $m_D$  has to be orders of magnitude smaller than any other Dirac mass. This is theoretically disfavored, and it is seldom discussed in the literature.

(iv)  $\Phi_R + \eta + \Phi_L$ . This is the left-right symmetric version of (iii). Again the neutrino mass appears to be purely Dirac. However, the coexistence of  $\Phi_L$  and  $\eta$  allows for an interesting extension of the usual left-right model, especially in the context of  $E_6$ . One of the complications of using a scalar bidoublet in a left-right extension of the standard model is that two different VEV's, i.e.  $\langle \eta_{1,2}^0 \rangle$ , contribute to any given fermion mass, thus implying the existence of flavor changing neutral currents (FCNC) in the scalar sector [8]. This is not a problem if the  $SU(2)_R$  breaking scale is very high as in models with a large Majorana  $\nu_R$  mass. In models where the neutrino mass is purely Dirac, the  $SU(2)_R$  breaking scale is not necessarily very high, so FCNC becomes the limiting constraint on the scale of  $SU(2)_R$  breaking. This constraint may be relaxed if there exists [9] an exotic quark h of charge -1/3 such that  $(u,h)_R$  is an  $SU(2)_R$  doublet instead of the usual  $(u,d)_R$ . Then  $m_u$  comes from  $\eta_1^0$ ,  $m_d$  comes from  $\phi_L^0$ , and  $m_h$  comes from  $\phi_R^0$ , with no FCNC in the scalar sector. This turns out to be a natural possibility [10] in the superstring-inspired  $E_6$  model. As for the lepton sector, the Dirac mass partner of  $\nu_L$  is then a new field which is a singlet, whereas the  $SU(2)_R$  partner of  $e_R$  (usually called  $\nu_R$ ) is now a different particle. Because there are more neutral fermions in this extension, Majorana masses for  $\nu_L$  may again be generated [11].

(v)  $\Phi_R + \Phi_L$ . This is the simplest way of breaking  $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  to  $U(1)_{em}$ . However, since the bidoublet  $\eta$  is absent, there are apparently no fermion masses. On the other hand, this creates a unique opportunity, i.e. the possibility that all fermion masses, be they Dirac or Majorana, come from dimension-5 operators instead [12,13], i.e. operators suppressed by presumably the Planck mass. In

the neutrino mass matrix of Eq. (12),  $m_L$  comes from  $(l_L \Phi_L)^2$ ,  $m_R$  comes from  $(l_R \Phi_R)^2$ , and  $m_D$  comes from  $(\bar{l}_L \Phi_L^{\dagger})(l_R \Phi_R)$ . The smallness of the Majorana neutrino mass compared to all Dirac fermion masses may then be attributed to the smallness of  $v_L \equiv \langle \phi_L^0 \rangle$  compared to  $v_R \equiv \langle \phi_R^0 \rangle$ .

There is another important consequence of this scenario. Because the Majorana  $\nu_R$  mass is now given by  $v_R^2/\Lambda$ , where  $\Lambda$  may be of order the Planck mass, say 10<sup>19</sup> GeV, it will be very much smaller than the  $SU(2)_R$  breaking scale, i.e.

$$m_R \sim \frac{v_R^2}{\Lambda} \ll v_R \,. \tag{15}$$

This means that in the early Universe, at temperatures comparable to  $m_R$ , the  $SU(2)_R$  gauge interactions of  $\nu_R$  are strongly suppressed and can safely be ignored. This is a crucial requirement [14] for leptogenesis through the decay of  $\nu_R$  [4]. Recent detailed analyses [15] of this mechanism for obtaining a realistic baryon asymmetry of the Universe and its relationship to the neutrino mass matrix all assume this implicitly.

Going back to scenarios (iii) and (iv), and allowing for the existence of  $(l_L \Phi_L)^2$  in (iv) and that of  $(l_R \Phi_R)^2$  in both (iii) and (iv), the seesaw neutrino mass matrices of Eqs. (11) and (12) are again reproduced for (iii) and (iv), respectively. This means that for a natural understanding of successful leptogenesis, the  $SU(2)_R$  model to be adopted should be one with an  $SU(2)_R$  doublet rather than a triplet. Scenario (v) requires  $v_R$  to be very high [12], of order the grand-unification scale, because of  $m_t$ . Scenario (iv) is a modification of (v) but without the  $v_R$  constraint, because fermion masses may now come from  $\eta$ . Scenario (iii) is a special case of (iv) and has the desirable original form of the seesaw neutrino mass matrix, i.e. Eq. (11) and not Eq. (12) as in (iv) and (v). This previously neglected model should then be put forward as the model of choice for understanding both neutrino mass and leptogenesis. [If Eq. (12) is used instead of Eq. (11), the extra parameters needed to specify the  $3 \times 3 m_L$  matrix would disable the connection [15] between neutrino mass and leptogenesis.

The scalar sector of scenario (iii) consists of only  $\Phi_R$  and  $\eta$ . Whereas  $SU(2)_R \times U(1)_{B-L}$  is broken down to  $U(1)_Y$  by the VEV of the doublet  $\Phi_R$ ,  $SU(2)_L \times U(1)_Y$  is broken down to  $U(1)_{em}$  by the VEV's of the bidoublet  $\eta$  which also provide Dirac masses for all the fermions. For example, consider the quark Yukawa couplings,

$$\mathcal{L}_{Y} = h_{1} \bar{q}_{L} \eta q_{R} + h_{2} \bar{q}_{L} \tilde{\eta} q_{R} + \text{H.c.}, \qquad (16)$$

where

$$\tilde{\eta} = \sigma_2 \eta^* \sigma_2 = \begin{pmatrix} \bar{\eta}_2^0 & -\eta_1^+ \\ -\eta_2^- & \bar{\eta}_1^0 \end{pmatrix}.$$
(17)

Hence

$$m_u = h_1 v_1 + h_2 v_2^*, \tag{18}$$

$$m_d = h_1 v_2 + h_2 v_1^*, \tag{19}$$

where  $v_{1,2} \equiv \langle \eta_{1,2}^0 \rangle$ . Because

$$\operatorname{Tr} \tilde{\eta}^{\dagger} \tilde{\eta} = \operatorname{Tr} \eta^{\dagger} \eta = \bar{\eta}_{1}^{0} \eta_{1}^{0} + \eta_{1}^{+} \eta_{1}^{-} + \bar{\eta}_{2}^{0} \eta_{2}^{0} + \eta_{2}^{+} \eta_{2}^{-},$$
(20)

whereas the independent scalar quartic terms  $f_1 \Phi_R^{\dagger} \tilde{\eta}^{\dagger} \tilde{\eta} \Phi_R$ and  $f_2 \Phi_R^{\dagger} \eta^{\dagger} \eta \Phi_R$  contain  $f_1 v_R^2 \bar{\eta}_1^0 \eta_1^0$  and  $f_2 v_R^2 \bar{\eta}_2^0 \eta_2^0$ , respectively, the effective scalar potential after  $\Phi_R$  has been integrated out has different mass terms for  $\eta_1^0$  and  $\eta_2^0$ , i.e.

$$(m^2 + f_1 v_R^2) \,\overline{\eta}_1^0 \,\eta_1^0 + (m^2 + f_2 v_R^2) \,\overline{\eta}_2^0 \,\eta_2^0. \tag{21}$$

This means that unless  $f_1=f_2$ , it is impossible to make both coefficients negative and of order the electroweak breaking scale. In other words, either  $\eta_1^0$  or  $\eta_2^0$  must remain heavy, i.e. of order  $v_R$ . Let  $m^2+f_1v_R^2=-\mu^2$ , then  $m_2^2=(f_2-f_1)v_R^2-\mu^2$ , and  $v_2/v_1$  is expected to be suppressed by a factor of order  $v_1^2/m_2^2$ , as shown for example in Ref. [16]. As a result, the contributions of  $v_2$  to Eqs. (18) and (19) are negligible and the suppression of FCNC is achieved.

The model so far has only Dirac fermion masses. Majorana neutrino masses would normally come from the wellknown dimension-5 operator [17]  $(l_L \Phi_L)^2$ , but since  $\Phi_L$  is absent, only  $(l_R \Phi_R)^2$  is available. Thus the original seesaw matrix of Eq. (11) is obtained, and  $m_R$  is guaranteed to be suppressed relative to  $v_R$  as shown in Eq. (15). Given the particle content of scenario (iii) and the acceptance of higher-dimensional operators,  $m_L$  actually gets a contribution from the dimension-7 operator  $(\bar{l}_L \eta \tilde{\Phi}_R)^2 / \Lambda^3$ . However its magnitude is  $v_1^2 v_R^2 / \Lambda^3$  which is smaller than the double seesaw [12] mass of  $v_1^2 \Lambda / v_R^2$  by the factor  $(v_R / \Lambda)^4$ .

Consider next the supersymmetric version of scenario (iii). Using the convention that all superfields are lefthanded,  $q_R$  is replaced by  $q^c \sim (3^*, 1, 2, -1/3)$  and  $l_R$  by  $l^c \sim (1,1,2,1)$ . The Higgs sector now consists of the superfields  $\eta$ ,  $\Phi_R$ , and  $\Phi_R^c \sim (1,1,2,-1)$ . An extra unbroken discrete  $Z_2$  symmetry is imposed, under which quark and lepton superfields are odd, but Higgs superfields are even. This serves to distinguish  $l^c$  from  $\Phi_R$ , and leads to the usual R parity of most supersymmetric extensions of the standard model.

To break  $SU(2)_R$  at a high scale without breaking the supersymmetry, consider the following superpotential:

$$W = M \epsilon_{ij} \Phi_{Ri} \Phi_{Rj}^c + \frac{1}{2\Lambda} (\epsilon_{ij} \Phi_{Ri} \Phi_{Rj}^c)^2.$$
(22)

Note that an extra nonrenormalizable term has been added. In this case, the scalar potential becomes

$$V = \sum_{j} \left| M \epsilon_{ij} \Phi_{Ri} + \frac{1}{\Lambda} (\epsilon_{i'j'} \Phi_{Ri'} \Phi_{Rj'}^{c}) \epsilon_{ij} \Phi_{Ri} \right|^{2} + \sum_{i} \left| M \epsilon_{ij} \Phi_{Rj}^{c} + \frac{1}{\Lambda} (\epsilon_{i'j'} \Phi_{Ri'} \Phi_{Rj'}^{c}) \epsilon_{ij} \Phi_{Rj}^{c} \right|^{2} + V_{g},$$

$$(23)$$

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where  $V_g$  comes from the gauge interactions of  $\Phi_R$  and  $\Phi_R^c$ . Since the neutral components of  $\Phi_R$  and  $\Phi_R^c$  have opposite values of  $I_{3R}$  and B-L, the condition  $V_g=0$  at its minimum is satisfied if  $\langle \Phi_R \rangle = \langle \Phi_R^c \rangle$  in the above. A supersymmetric vacuum (V=0) is thus obtained with

$$v_R = \langle \Phi_R \rangle = \langle \Phi_R^c \rangle = \sqrt{M\Lambda}. \tag{24}$$

This shows that *M* of Eq. (22) may be identified with  $m_R$  of Eq. (15), i.e. the large Majorana mass of  $\nu_R$ .

In this scenario,  $SU(2)_R$  is broken at the scale  $v_R$  of Eq. (24). Below it, a consistent supersymmetric extension of the standard model survives, but with a singlet neutrino of Majorana mass  $\sim M \ll v_R$ . This singlet neutrino couples to  $(\nu \eta_2^0 - e \eta_2^+)$  but its interaction with the  $SU(2)_R$  gauge bosons is very much suppressed at the time of the early Universe when its temperature is comparable to M. Its decay will thus generate a lepton asymmetry [4,15] which gets converted [18] into the present observed baryon asymmetry of the Universe through sphalerons during the electroweak phase transition.

So far the scale  $v_R$  has not been determined. There are two possible approaches. One is to assume that it has to do with gauge-coupling unification of the minimal supersymmetric standard model (MSSM) [19], in which case it should be  $10^{16}$  GeV. This implies a singlet neutrino mass of order  $(10^{16})^2/10^{19} = 10^{13}$  GeV. The other is to use present neutrino data [20–22] together with the requirement that the canonical seesaw matrix of Eq. (11) yields a satisfactory baryon asymmetry of the Universe through  $v_R$  decay, from which a lower bound on  $v_R$  may be obtained. Recent indications [15] are that the smallest  $m_R$  is of order  $10^8$  GeV, which implies that  $v_R$  is at least of order  $10^{14}$  GeV.

The bidoublet  $\eta$  contains two electroweak doublets and they are just right for the unification of gauge couplings in the MSSM. However, two such bidoublets are usually assumed in a supersymmetric model in order to have realistic quark and lepton masses. [Because of supersymmetry, we cannot use  $\tilde{\eta}$  as the second bidoublet as in Eqs. (18) and (19).] In that case, there are four electroweak doublets and two would have to be heavy (i.e. at the  $v_R$  scale) not to spoil the unification of gauge couplings. An alternative possibility is to keep only one bidoublet and invoke a flavornondiagonal soft supersymmetry breaking scalar sector to account for the observed quark and lepton mass matrices [23].

Scenario (iii) is distinguished by the absence of a  $\Phi_L$ doublet. This is a problem if  $SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$  is embedded in a larger symmetry such as SO(10) or  $[SU(3)]^3$ , because any scalar multiplet of this larger symmetry would also contain  $\Phi_L$  if it contains  $\Phi_R$ . In that case, Eq. (12) is obtained, where  $m_L$  comes from  $(l_L \Phi_L)^2$ . As long as  $m_L \ll m_D^2/m_R$ , which holds if  $\Lambda$  is of order 10<sup>19</sup> GeV, this is an acceptable scenario as well.

In conclusion, to understand both neutrino masses in terms of the original canonical seesaw mechanism, i.e. Eq. (11), and the success of leptogenesis through  $\nu_R$  decay, the

simplest and most natural model is  $SU(3)_C \times SU(2)_L$   $\times SU(2)_R \times U(1)_{B-L}$  with a scalar sector consisting of only an  $SU(2)_R$  doublet  $\Phi_R$  and an  $SU(2)_L \times SU(2)_R$  bidoublet  $\eta$ . The dimension-5 operator  $(l_R \Phi_R)^2/2\Lambda$  leads to a large Majorana mass for  $\nu_R$  such that  $m_R \approx v_R^2/\Lambda \ll v_R$ , which is desirable for generating the proper baryon asymmetry of the Universe. A successful supersymmetric version of this model has also been discussed.

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