

Modified Paschos-Wolfenstein relation and extraction of the weak mixing angle $\sin^2\theta_W$

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The NuTeV Collaboration reported an anomalously large weak mixing angle $\sin^2\theta_W$ in comparison with the standard model prediction. Neutrino and antineutrino charged- and neutral-current events are analyzed for extracting $\sin^2\theta_W$. Although the Paschos-Wolfenstein relation is not directly used in the analysis, it plays an important role in the determination. Noting that the target nucleus, iron, is not an isoscalar nucleus, we derive a leading-order expression for a modified Paschos-Wolfenstein relation for nuclei, which may have neutron excess. Then, using charge and baryon-number conservations for nuclei, we discuss a nuclear correction in the $\sin^2\theta_W$ determination. It is noteworthy that nuclear modifications are different between valence up- and down-quark distributions. We show this difference effect on the NuTeV $\sin^2\theta_W$ deviation.

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I. INTRODUCTION

The weak mixing angle $\sin^2\theta_W$ is one of the important quantities in the standard model. In the on-shell scheme, it is related to the W and Z masses by $\sin^2\theta_W = 1 - m_W^2/m_Z^2$. Collider experiments provide accurate values for these masses and the angle. According to a global analysis [1], it is $\sin^2\theta_W^{on-shell} = 0.2227 \pm 0.0004$ by excluding neutrino-nucleus scattering data.

The NuTeV Collaboration (Zeller *et al.*) reported recently [2] that the mixing angle should be significantly larger:

$$\sin^2\theta_W = 0.2277 \pm 0.0013(\text{stat}) \pm 0.0009(\text{syst}), \quad (1.1)$$

by using their neutrino and antineutrino scattering data. For extracting $\sin^2\theta_W$, it is known that the Paschos-Wolfenstein (PW) relation [3]

$$R^- = \frac{\sigma_{NC}^{\nu N} - \sigma_{NC}^{\bar{\nu} N}}{\sigma_{CC}^{\nu N} - \sigma_{CC}^{\bar{\nu} N}} = \frac{1}{2} - \sin^2\theta_W, \quad (1.2)$$

is useful because uncertainties from charm production and possible nuclear effects are much reduced. Here, $\sigma_{NC}^{\nu N}$ and $\sigma_{CC}^{\nu N}$ are the deep inelastic cross sections for neutral-current (NC) and charged-current (CC) neutrino interactions with the nucleon. The factor ρ , which is the relative strength between the neutral and charged currents, is taken as one. The NuTeV Collaboration measured charged and neutral current ratios, $R_\nu = \sigma_{NC}^{\nu N}/\sigma_{CC}^{\nu N}$ and $R_{\bar{\nu}} = \sigma_{NC}^{\bar{\nu} N}/\sigma_{CC}^{\bar{\nu} N}$ and then a Monte Carlo simulation is used for relating the data to $\sin^2\theta_W$. Fitting these ratios simultaneously, they end up using the PW-like relation although it is not directly employed in the analysis [4]. In this sense, it is mentioned in Ref. [5] that “the NuTeV result derives $\sin^2\theta_W$ from the Paschos-Wolfenstein.” The result suggests that the left-handed neutral current coupling should be smaller than expected. If it is confirmed, it should lead to a new physics finding [6]. The situation is summarized in the paper by Davidson *et al.* [6].

On the other hand, there are suggestions from a conservative point of view. Miller and Thomas commented [7] that the anomalous result could be explained by the shadowing difference between neutral and charged current reactions by using a vector meson dominance (VMD) model. The NuTeV Collaboration replied to their comments [5] that the explanation is not favored because the shadowing effects are subtracted out in the PW relation. Furthermore, the model cannot explain observed R_ν and $R_{\bar{\nu}}$ ratios, which are smaller than expected, and also the VMD model does not have proper Q^2 dependence. However, their Q^2 discussion is refuted by Melnitchouk and Thomas in Ref. [8]. Nuclear corrections are also discussed by Kovalenko, Schmidt, and Yang [9] by noting nuclear modifications of F_2 . However, one should note that such nuclear effects were taken into account in the NuTeV analysis in a slightly different way [10].

It is not the purpose of this paper to examine the details of these previous studies. We rather try to address ourselves to the extraction of $\sin^2\theta_W$ from nuclear data in a model independent way as much as possible by resorting to charge and baryon-number conservations. Because the NuTeV target is mainly the iron nucleus, nuclear corrections should be carefully taken into account for a precise determination of $\sin^2\theta_W$. In this paper, we derive a modified PW relation for general nuclear targets. Then, we discuss a possible nuclear modification factor which could change the extracted $\sin^2\theta_W$ value.

This paper consists of the following. First, nuclear corrections of the PW relation are discussed in Sec. II. Then, possible effects on the extraction of $\sin^2\theta_W$ are explained in Sec. III. The results are summarized in Sec. IV.

II. MODIFIED PASCHOS-WOLFENSTEIN RELATION FOR NUCLEI

The PW relation was derived for the isoscalar nucleon; however, the used NuTeV target is mainly iron, which is not an isoscalar nucleus. The neutron excess may cause unexpected nuclear corrections, which should be carefully investigated. In this section, we derive a leading-order (LO) PW expression for general nuclei in a model independent way.

First, neutrino- and antineutrino-nucleus charged current

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cross sections are given in the LO by [11]

$$\begin{aligned} \frac{d\sigma_{CC}^{\nu A}}{dx dy} &= \sigma_0 x [d^A(x) + s^A(x) + \{\bar{u}^A(x) + \bar{c}^A(x)\}(1-y)^2], \\ \frac{d\sigma_{CC}^{\bar{\nu} A}}{dx dy} &= \sigma_0 x [\bar{d}^A(x) + \bar{s}^A(x) + \{u^A(x) + c^A(x)\}(1-y)^2]. \end{aligned} \quad (2.1)$$

Here, σ_0 is defined as $\sigma_0 = G_F^2 s / \pi$, by neglecting the factor Q^2 / M_W^2 from the propagator, with the Fermi coupling constant G_F and the center-of-mass squared energy s . The variables x and y are defined by the momentum transfer square Q^2 ($= -q^2$), the energy transfer q^0 , and the nucleon mass M as $x = Q^2 / (2Mq^0)$ and $y = q^0 / E_\nu$ or $q^0 / E_{\bar{\nu}}$. Nuclear quark and antiquark distributions are denoted $q^A(x)$ and $\bar{q}^A(x)$, respectively. They depend also on Q^2 ; however, the explicit Q^2 dependence is abbreviated in the parton distribution functions (PDFs) throughout this paper. Next, the neutral current cross section for neutrino scattering is given by [11]

$$\begin{aligned} \frac{d\sigma_{NC}^{\nu A}}{dx dy} &= \sigma_0 x [\{u_L^2 + u_R^2(1-y)^2\}\{u^A(x) + c^A(x)\} \\ &\quad + \{u_R^2 + u_L^2(1-y)^2\}\{\bar{u}^A(x) + \bar{c}^A(x)\} \\ &\quad + \{d_L^2 + d_R^2(1-y)^2\}\{d^A(x) + s^A(x)\} \\ &\quad + \{d_R^2 + d_L^2(1-y)^2\}\{\bar{d}^A(x) + \bar{s}^A(x)\}], \end{aligned} \quad (2.2)$$

where left- and right-handed couplings are expressed by the weak mixing angle as $u_L = 1/2 - (2/3)\sin^2\theta_W$, $u_R = -(2/3)\sin^2\theta_W$, $d_L = -1/2 + (1/3)\sin^2\theta_W$, and $d_R = (1/3)\sin^2\theta_W$. The cross section for antineutrino scattering is obtained by exchanging the left- and right-handed couplings.

Using these equations, we obtain a nuclear PW relation as

$$\begin{aligned} R_A^- &= \frac{\sigma_{NC}^{\nu A} - \sigma_{NC}^{\bar{\nu} A}}{\sigma_{CC}^{\nu A} - \sigma_{CC}^{\bar{\nu} A}} = \{1 - (1-y)^2\} [(u_L^2 - u_R^2)\{u_v^A(x) + c_v^A(x)\} \\ &\quad + (d_L^2 - d_R^2)\{d_v^A(x) + s_v^A(x)\}] / [d_v^A(x) + s_v^A(x) \\ &\quad - (1-y)^2\{u_v^A(x) + c_v^A(x)\}]. \end{aligned} \quad (2.3)$$

Here, the valence quark distributions are defined by $q_v^A \equiv q^A - \bar{q}^A$. The name, valence quark, is conventionally used for valence up- and down-quark distributions, but we use the same nomenclature for strange and charm distributions by applying the same definitions, $s_v^A \equiv s^A - \bar{s}^A$ and $c_v^A \equiv c^A - \bar{c}^A$. Of course, there is no net strangeness and charm in ordinary nuclei, so that ‘‘valence’’ strange and charm distributions should satisfy $\int dx s_v^A(x) = 0$ and $\int dx c_v^A(x) = 0$. However, these equations do not mean that the distributions themselves vanish: $s_v^A(x) \neq 0$ and $c_v^A(x) \neq 0$.

The valence-quark distributions u_v^A and d_v^A are expressed by the weight functions w_{u_v} and w_{d_v} at any Q^2 :

$$\begin{aligned} u_v^A(x) &= w_{u_v}(x, A, Z) \frac{Z u_v(x) + N d_v(x)}{A}, \\ d_v^A(x) &= w_{d_v}(x, A, Z) \frac{Z d_v(x) + N u_v(x)}{A}, \end{aligned} \quad (2.4)$$

where u_v and d_v are the distributions in the proton, Z is the atomic number, N is the neutron number, and A is the mass number of a nucleus. The weight functions are defined at fixed Q^2 ($= 1 \text{ GeV}^2$) in Ref. [12]; however, they are used at any Q^2 throughout this paper. The explicit Q^2 dependence is abbreviated as for the PDFs. Although these equations are originally motivated by the isospin symmetry in nuclei for the virtual $w_{u_v} = w_{d_v} = 1$ case, we do not rely on such an assumption. This is because nuclear modifications, including isospin violation [13,14] and nuclear charge-symmetry breaking [15], could be taken into account by the weight functions in any case. Therefore, the expressions are given without losing any generality.

Next, we define neutron excess constant $\hat{\varepsilon}_n$ and a related function $\varepsilon_n(x)$ by

$$\hat{\varepsilon}_n = \frac{N-Z}{A}, \quad \varepsilon_n(x) = \hat{\varepsilon}_n \frac{u_v(x) - d_v(x)}{u_v(x) + d_v(x)}. \quad (2.5)$$

Then, a difference between the weight functions is defined by

$$\varepsilon_v(x) = \frac{w_{d_v}(x, A, Z) - w_{u_v}(x, A, Z)}{w_{d_v}(x, A, Z) + w_{u_v}(x, A, Z)}. \quad (2.6)$$

The function ε_v depends also on A , Z , and Q^2 , but these factors are abbreviated in writing ε_v . Substituting Eqs. (2.4), (2.5), and (2.6) together with the coupling constants into Eq. (2.3), we obtain

$$\begin{aligned} R_A^- &= \left[\left(\frac{1}{2} - \sin^2\theta_W \right) \{1 + \varepsilon_v(x)\varepsilon_n(x)\} + \frac{1}{3}\sin^2\theta_W \{ \varepsilon_v(x) \right. \\ &\quad \left. + \varepsilon_n(x) \} + \left(\frac{1}{2} - \frac{2}{3}\sin^2\theta_W \right) \varepsilon_s(x) \right. \\ &\quad \left. + \left(\frac{1}{2} - \frac{4}{3}\sin^2\theta_W \right) \varepsilon_c(x) \right] / \left[1 + \varepsilon_v(x)\varepsilon_n(x) \right. \\ &\quad \left. + \frac{1 + (1-y)^2}{1 - (1-y)^2} \{ \varepsilon_v(x) + \varepsilon_n(x) \} \right. \\ &\quad \left. + \frac{2\{ \varepsilon_s(x) - (1-y)^2 \varepsilon_c(x) \}}{1 - (1-y)^2} \right]. \end{aligned} \quad (2.7)$$

Here, ε_s and ε_c are defined by $\varepsilon_s = s_v^A / [w_v(u_v + d_v)]$ and $\varepsilon_c = c_v^A / [w_v(u_v + d_v)]$ with $w_v = (w_{d_v} + w_{u_v})/2$. We would like to stress that the LO expression in Eq. (2.7) has been derived without using any model dependent factor and any serious approximation.

The strange quark (s_v^A) effects are investigated in Ref. [14], and the neutron-excess effects, namely the ε_n terms, are taken into account in the NuTeV analysis [10]. In addition to these corrections, we notice in Eq. (2.7) that another correction factor ε_v contributes to the deviation from the PW relation $1/2 - \sin^2\theta_W$ due to *the difference between the valence up- and down-quark modifications in a nucleus*. This fact needs to be clarified. There are two restrictions on the valence-quark distributions in a nucleus. One is the baryon number conservation [16], and the other is the charge conservation [12]. Nuclear baryon number and charge have to be A and Z , and they are expressed in the parton model as

$$A = \int dx A \sum_q \frac{1}{3} (q^A - \bar{q}^A) = \int dx \frac{A}{3} (u_v^A + d_v^A),$$

$$Z = \int dx A \sum_q e_q (q^A - \bar{q}^A) = \int dx \frac{A}{3} (2u_v^A - d_v^A), \quad (2.8)$$

where A is multiplied in the integrands because the nuclear parton distributions are defined by those per nucleon, and the relation $\int dx s_v(x) = \int dx c_v(x) = 0$ is used in obtaining the right-hand sides. Substituting Eq. (2.4) into Eq. (2.8), we obtain

$$\int dx (u_v + d_v) [\Delta w_v + w_v \varepsilon_v(x) \varepsilon_n(x)] = 0, \quad (2.9)$$

$$\int dx (u_v + d_v) [\Delta w_v \{1 - 3\varepsilon_n(x)\} - w_v \varepsilon_v(x) \{3 - \varepsilon_n(x)\}] = 0, \quad (2.10)$$

where Δw_v is defined by $\Delta w_v = w_v - 1$. Although it is not straightforward to determine $\varepsilon_v(x)$ from Eqs. (2.9) and (2.10), it indicates that $\varepsilon_v(x)$ is finite.

In this way, we find that nuclear modifications are, in general, different between the valence up- and down-quark distributions because of the baryon number and charge conservations. It gives rise to the factor ε_v . This kind of detailed nuclear effects cannot be accounted simply by investigating electron and muon scattering data and by fitting charged current cross-section data for the same target [10]. Because the physics associated with the ε_v factor is missing in the NuTeV analysis, it may cause a significant effect on the $\sin^2\theta_W$ determination.

III. EFFECTS ON $\sin^2\theta_W$ DETERMINATION

The angle $\sin^2\theta_W$ can be extracted by using Eq. (2.7) together with the experimental data of R_A^- . In order to find whether or not the ε_n corrections could explain the deviation from the standard model value for $\sin^2\theta_W$, we approximate the relation by considering that the corrections are small ($\varepsilon_v \ll 1$). Retaining only the leading correction of ε_v in Eq. (2.7), we obtain

$$R_A^- = \frac{1}{2} - \sin^2\theta_W - \varepsilon_v(x) \left\{ \left(\frac{1}{2} - \sin^2\theta_W \right) \frac{1 + (1-y)^2}{1 - (1-y)^2} - \frac{1}{3} \sin^2\theta_W \right\} + O(\varepsilon_v^2) + O(\varepsilon_n) + O(\varepsilon_s) + O(\varepsilon_c). \quad (3.1)$$

The higher-order and other corrections $O(\varepsilon_v^2)$, $O(\varepsilon_n)$, $O(\varepsilon_s)$, and $O(\varepsilon_c)$ are not explicitly written. As far as present neutrino data suggest, the ‘‘valence’’ strange distribution should be small, and the measurements indicate that such a correction increases the NuTeV $\sin^2\theta_W$ deviation according to Ref. [14]. Therefore, at least at this stage, the finite s_v^A and c_v^A distributions effects, $O(\varepsilon_s)$ and $O(\varepsilon_c)$, are not the favorable explanation. Although accurate measurements may clarify the details of the distributions s_v^A and c_v^A in future, they are not discussed in the following. The neutron-excess effects $O(\varepsilon_n)$ are included in the NuTeV analysis [10], so that they are not the source of the $\sin^2\theta_W$ deviation.

As mentioned in Sec. I and Ref. [5], the NuTeV $\sin^2\theta_W$ is ‘‘derived’’ from the PW-like relation indirectly. It is obvious from Eq. (3.1) that there is an ε_n -type correction to $\sin^2\theta_W$, and it may explain, at least partially, the deviation from the standard model. However, the correction is essentially unknown at this stage in the sense that there is no significant data to find the difference between valence up- and down-quark modifications (ε_v). In order to investigate whether or not the ε_v correction is large enough, we should inevitably use some prescription for describing the ε_v factor. In the following, we introduce two different descriptions as examples.

A. A prescription for the conservations—description 1

It is not straightforward to find a solution $\varepsilon_v(x)$ to satisfy Eqs. (2.9) and (2.10). For an approximate estimate, the higher-order corrections $\varepsilon_v \varepsilon_n$ are neglected in these equations. Then, substituting Eq. (2.9) into Eq. (2.10), we obtain

$$\varepsilon_v(x) = -\hat{\varepsilon}_n \frac{u_v(x) - d_v(x)}{u_v(x) + d_v(x)} \frac{\Delta w_v(x)}{w_v(x)}, \quad (3.2)$$

by considering a special case that the integrand vanishes. Of course, this is not a unique solution, but this estimate should be able to provide information about the magnitude of the correction.

B. A χ^2 analysis of nuclear PDFs—description 2

Global χ^2 analysis results could be used for calculating $\varepsilon_v(x)$. A χ^2 analysis for determining nuclear parton distribution functions (NPDFs) is reported in Ref. [12], and obtained distributions can be calculated by using subroutines at the web site in Ref. [17]. Before using the NPDF code, we would like to remind the reader what has been done for the valence distributions. At $Q^2 = 1 \text{ GeV}^2$, the weight functions are assumed to be

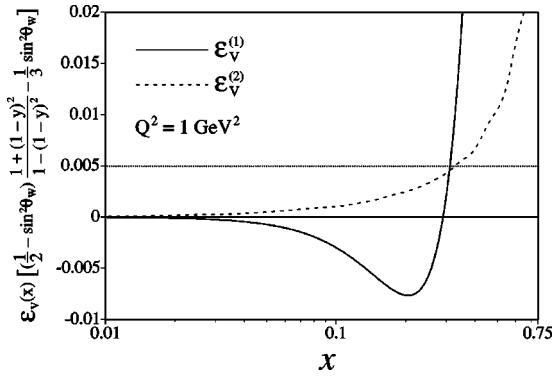


FIG. 1. The ε_ν correction term is evaluated at $Q^2=1 \text{ GeV}^2$. The solid curve is calculated by Eq. (3.2), and the dashed curve is obtained by the χ^2 fit code [12,17]. The NuTeV deviation 0.005 is shown by the dotted line just for comparison.

$$w_i(x) = 1 + \left(1 - \frac{1}{A^{1/3}}\right) \frac{a_i + b_v x + c_v x^2 + d_v x^3}{(1-x)^{\beta_v}}, \quad (3.3)$$

where i denotes u_v or d_v , and a_{u_v} , b_v , c_v , and d_v are the parameters to be determined by the χ^2 fit. Because it is the first χ^2 analysis attempt for nuclei, the parameter number has to be reduced as many as possible. It is the reason why the common parameters are chosen for β_v , b_v , c_v , and d_v . However, in order to satisfy both the charge and baryon-number conservations, at least one parameter should be different. Because the parameters a_{u_v} and a_{d_v} are determined so as to satisfy these conservations, this description of $\varepsilon_\nu(x)$ surely satisfies Eqs. (2.9) and (2.10). However, one should note that this $\varepsilon_\nu(x)$ may not be valid if it has more complicated x dependence than the one calculated from Eq. (3.3).

Numerical results

In the first description, ε_ν is evaluated numerically by using Eq. (3.2) with the NPDF code in Ref. [17] for calculating u_v , d_v , w_{u_v} , and w_{d_v} at given Q^2 . In the second one, ε_ν is calculated by the definition in Eq. (2.6) with the weight functions w_{u_v} and w_{d_v} , which are numerically calculated by the NPDF code. In the NuTeV measurements [4,5], averages of the kinematical variables are given by $\langle E \rangle = 120$ and 112 GeV , $\langle Q^2 \rangle = 25.6$ and 15.4 GeV^2 , $\langle x \rangle = 0.22$ and 0.18 for neutrinos and antineutrinos, respectively. The ranges of x and Q^2 are $0.01 < x < 0.75$ and $1 < Q^2 < 140 \text{ GeV}^2$. Although the ν and $\bar{\nu}$ incident energies are different, we use the averaged value $\langle E \rangle = 116 \text{ GeV}$ for connecting x , y , and Q^2 : $y = Q^2 / (2Mx\langle E \rangle)$.

In Fig. 1, the Q^2 value is fixed at $Q^2=1 \text{ GeV}^2$, and the correction term in Eq. (3.1) is evaluated as a function of x . The neutron excess $\hat{\varepsilon}_n = 4/56$ and $\sin^2\theta_W = 0.2227$ are used in the calculations. The solid ($\varepsilon_\nu^{(1)}$) and dashed ($\varepsilon_\nu^{(2)}$) curves are obtained by the first and second descriptions, respectively. The dotted line indicates the NuTeV $\sin^2\theta_W$ deviation ($0.2277 - 0.2227 = 0.0050$) just for a comparison. Both curves increase rapidly as x becomes larger, and this is mostly a kinematical effect due to the factor $1/[1 - (1 - y)^2]$.

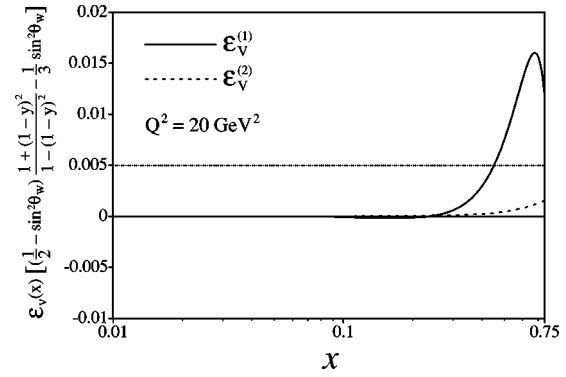


FIG. 2. The correction term is evaluated at $Q^2=20 \text{ GeV}^2$.

Another effect at large x comes from the Fermi-motion-like factor $1/(1-x)^{\beta_v}$ in Eq. (3.3). The $\varepsilon_\nu^{(2)}$ correction term is very small with the following reason. The x dependence in Eq. (3.3) indicates that $\varepsilon_\nu^{(2)}$ is almost independent of x except for large x . The first terms in Eqs. (2.9) and (2.10) are valence-quark distributions multiplied by Δw_v , and these integrals almost vanish. The obtained $\varepsilon_\nu^{(2)}$ is roughly proportional to these small integral values. One should note that this result could be artificially small due to the weak x dependence assumption. On the other hand, the magnitude of $\varepsilon_\nu^{(1)}$ is large in general and it has completely different x dependence. Because $\varepsilon_\nu^{(1)}$ is directly proportional to the valence-quark modification (Δw_v) in Eq. (3.2), the function changes sign at $x \approx 0.3$, which is the transition point from antishadowing to the EMC-effect region. In the revised analysis [18,19], the valence modifications are slightly different; however, the essential features are the same. According to the first description, the function becomes comparable magnitude to the NuTeV deviation.

Because the average Q^2 is much larger than 1 GeV^2 in the NuTeV experiment, the corrections are also calculated at $Q^2=20 \text{ GeV}^2$ by noting the kinematical limit $y < 1$. The results are shown in Fig. 2. In comparison with the $Q^2=1 \text{ GeV}^2$ results, the effects are much suppressed. This is again due to the kinematical factor $1/[1 - (1 - y)^2]$. For example, this factor is 55 for $x=0.5$ and $Q^2=1 \text{ GeV}^2$; however, it becomes 3.0 for $x=0.5$ and $Q^2=20 \text{ GeV}^2$. This is the reason why both distributions become smaller. Although the $\varepsilon_\nu^{(2)}$ effect is too small to explain the deviation at $Q^2=20 \text{ GeV}^2$, the $\varepsilon_\nu^{(1)}$ is still comparable magnitude.

In order to investigate these effects on the NuTeV $\sin^2\theta_W$, the analysis should be done with the Monte Carlo code with the experimental data. However, in order to show the order of magnitude, we simply average the obtained curves over the x range (Δx) from 0.01 or $x_{min} = Q^2 / (2M\langle E \rangle)$ to 0.75. Because the data are centered at about $x=0.2$, this kind of simple average could overestimate the effects coming from the large x region. The calculated results are shown by the solid curves in Fig. 3. As already found in Figs. 1 and 2, the effects are very large at small Q^2 and become smaller as Q^2 becomes larger.

If the simple x average is taken, the effects look large. On the other hand, there is a method to take the NuTeV kine-

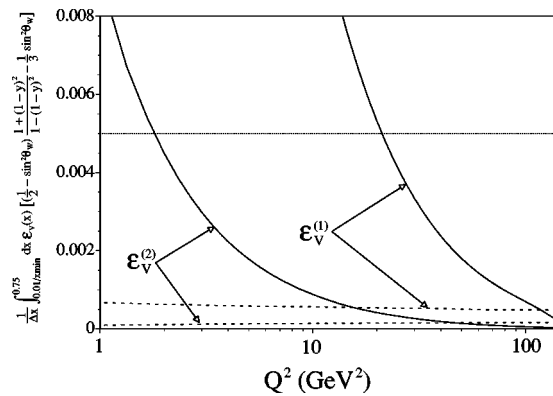


FIG. 3. The solid curves indicate the corrections averaged over the x range. The dashed ones are obtained by taking the NuTeV kinematics into account.

matics into account [22] by using the functionals in Fig. 1 of Ref. [14]. Although the physics motivation is completely different, the present ϵ_v distribution could be simulated by the NuTeV distributions $u_v^p - d_v^n$ and $d_v^p - u_v^n$. It is interesting that their “isospin-violating distributions” could effectively contain the present nuclear effect. If such a correspondence is made, the effect on $\sin^2\theta_W$ is calculated by using the NuTeV functionals [4,22]. The results are shown by the dashed curves in Fig. 3. We find that the effects become significantly smaller due to the lack of large x events in the

NuTeV experiment. Therefore, as far as the considered two descriptions are concerned, the nuclear modifications are not large enough to explain the whole NuTeV deviation.

It is, however, still too early to conclude that the present mechanism is completely excluded, because the nuclear modification difference between u_v^A and d_v^A is not measured at all. It may not be possible to find this nuclear effect until a neutrino factory [19,20] or the NuMI project [21] is realized. On the other hand, if these facilities are built, the “nucleon” cross sections (hence $\sin^2\theta_W$) could be measured with proton and deuteron targets with minor nuclear corrections. The NuTeV also reported that R_ν and $R_{\bar{\nu}}$ are unexpectedly smaller [5]. We are also investigating this issue, and hopefully it will be clarified in the near future.

IV. SUMMARY

We have derived a modified Paschos-Wolfenstein relation for nuclei. Using this relation, we investigated the possibility that the NuTeV $\sin^2\theta_W$ deviation could be explained by the nuclear parton distributions in iron. In particular, we pointed out that nuclear modifications are different between valence up- and down-quark distributions. The difference partially explains the NuTeV deviation although it may not be large enough to explain the whole deviation. Because such modifications are not measured, it is important to investigate them in future.

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