# **Heavy neutrino mixing effects in helicity amplitudes for the process**  $\mu^+\mu^-\rightarrow W^+W^-$

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The helicity amplitudes for the process  $\mu^+\mu^-\rightarrow W^+W^-$  are reevaluated with the inclusion of heavy neutrino mixing with the standard light neutrino. The effects of mixing are parametrized as a fractional change  $\Delta \tilde{M}^{* \nu}$  in the helicity amplitude for the *t*-channel  $\nu_{\mu}$  exchange. The behavior of  $\Delta \tilde{M}^{* \nu}$  is examined. It is found that for a heavy neutrino mass  $m_N=10$  GeV and scattering angle  $\theta=90^\circ$ ,  $\Delta \tilde{M}^{* \nu}=0.007|U_{\nu N}|^2$  at  $\sqrt{s}$ = 200 GeV and  $\Delta \tilde{M}^{* \nu}$  = 0.00002|U<sub>vN</sub>|<sup>2</sup> at  $\sqrt{s}$  = 800 GeV. The heavy neutrino mixing effects tend to vanish in the limit  $m_N \ll \sqrt{s}$  even if the mixing element  $U_{\nu N}$  is large.

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### **I. INTRODUCTION**

Neutral heavy leptons or heavy neutrinos (HNs) are predicted by various extensions of the standard model (SM)  $[1-3]$ . These HNs can mix with standard light neutrinos [3]. Although recent searches for HNs reveal no evidence in the mass range 0.25–2.0 GeV, the existence of more heavy neutrinos is not ruled out  $[4]$ .

The physics potential of high energy muon colliders  $[5,6]$ has opened the prospect for a detailed study of *W*-boson pair production in the process  $\mu^+\mu^-\rightarrow W^+W^-$ . In the SM this occurs at the tree level through the dominant *s*-channel  $\gamma$ , *Z* exchange [Fig. 1(a)] and *t*-channel neutrino ( $v<sub>\mu</sub>$ ) exchange [Fig. 1(b)]. The Higgs-boson exchange [Fig. 1(c)] cancels the bad high energy behavior for the production of longitudinally polarized *W* bosons in case of massive muons (growth of *s*-wave scattering amplitude as  $\sqrt{s}$ ). If mixing between light and heavy neutrinos exists, it results in two effects: (i) the  $W_{\nu\mu}$  couplings are modified (reduced) and (ii) an additional *t*-channel heavy neutrino  $(N<sub>u</sub>)$  exchange [Fig.  $1(d)$ ] contributes. With a view to discern the heavy neutrino mixing effects, we evaluate in this Brief Report the helicity amplitudes for the process  $\mu^+\mu^-\rightarrow W^+W^-$  with the inclusion of heavy neutrino mixing. We parametrize the mixing effects as a fractional change in the helicity amplitude contribution from the neutrino exchange diagram. We retain the muon mass terms to examine the effects of mixings on the cancellation of bad high energy behavior.

# **II. HEAVY NEUTRINO MIXING AND MODIFIED HELICITY AMPLITUDES**

### **A. Mixing model**

A general discussion of mixing between known neutrino fields and new heavy neutrino fields is available in Ref.  $[7]$ . When mixings are allowed, the neutrino mass eigenstates  $(v_{\mu}, N_{\mu})$  are related to the weak interaction eigenstates  $(\nu^0_\mu, N^0_\mu)$  by a unitary transformation [8]

$$
\begin{pmatrix} \nu_{\mu} \\ N_{\mu} \end{pmatrix} = \begin{pmatrix} U_{\nu\nu} & U_{\nu N} \\ U_{N\nu} & U_{NN} \end{pmatrix} \begin{pmatrix} \nu_{\mu}^{0} \\ N_{\mu}^{0} \end{pmatrix}, \tag{1}
$$

where the matrix *U* diagonalizes the neutrino mass matrix. The elements  $U_{ij}$  are constrained by the relations

$$
|U_{\nu\nu}|^2 = 1 - |U_{\nu N}|^2, \quad |U_{NN}|^2 = 1 - |U_{N\nu}|^2,
$$
  

$$
U_{\nu\nu}U_{N\nu}^* = -U_{\nu N}U_{N\nu}^*.
$$
 (2)

The heavy neutrino mixing when included in the SM theory does not change the (i)  $WW\gamma$ , (ii)  $WWZ$ , (iii)  $\gamma\mu\mu$ , and (iv)  $Z\mu\mu$  couplings. However, the  $W_{\nu\mu}$  coupling is modified. The SM interaction Lagrangian, relevant for ascertaining changes in  $W_{\nu\mu}$  coupling in weak eigenstate basis, is

$$
L_{\text{int}} = \frac{e}{\sqrt{2} \sin \theta_W} \left[ \left( \bar{v}_{\mu L}^0 \gamma^\mu \mu_L^0 \right) W_\mu^+ + \text{H.c.} \right],\tag{3}
$$

where *e* is the positron charge and  $\theta_W$  is the weak mixing angle. Using Eq.  $(1)$ , the  $L<sub>int</sub>$  in the mass eigenstate basis becomes  $[9]$ 



FIG. 1. (a) Dominant *s*-channel  $\gamma$ - and *Z*-exchange diagram, (b) *t*-channel  $v_u$ -exchange diagram, (c) Higgs-boson exchange diagram, and  $\overline{d}$  additional *t*-channel  $N_\mu$ -exchange diagram.

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$(\lambda,\overline{\lambda})$	$A_{\lambda\overline{\lambda}}^{\gamma,Z}$	$B_{\lambda\overline{\lambda}}$	$C_{\lambda\overline{\lambda}}$	$A'^{\gamma,Z}_{\lambda\overline{\lambda}}$	$B'_{\lambda\overline{\lambda}}$		$H_{\lambda\overline{\lambda}}$
$(+,0)$ $(0,-)$	$2\gamma$	$2\gamma$	$2(1+\beta)/\gamma$			$1-(\beta+\sigma)\beta^2$ $1-(\beta-\sigma)\beta^2$	
$(0,+)$ $(-,0)$	$2\gamma$	$2\gamma$	$2(1-\beta)/\gamma$	1		$1+(\beta+\sigma)\beta^2$ $1+(\beta-\sigma)\beta^2$	
$(+,+)$ $(-,-)$	1		$1/\gamma^2$	$1/2\gamma$	$\cos \theta/2\gamma$	$\cos \theta/2\gamma^3$	$-1/4\nu$
(0,0)	$2\gamma^2+1$	$2v^2$	$2/\gamma^2$	$(2\gamma^2+1)/2\gamma$	$\gamma(\beta + \cos \theta)$	$\cos \theta / \gamma^3$	$\gamma(1+\beta^2)/4$

TABLE I. The explicit form of the  $A_{\lambda\overline{\lambda}}^{\gamma,Z}$ ,  $A_{\lambda\overline{\lambda}}^{\gamma,Z}$ ,  $B_{\lambda\overline{\lambda}}$ ,  $B_{\lambda\overline{\lambda}}^{\gamma}$ ,  $C_{\lambda\overline{\lambda}}^{\gamma}$ ,  $C_{\lambda\overline{\lambda}}^{\gamma}$ , and  $H_{\lambda\overline{\lambda}}^{\gamma}$  coefficients.

$$
L_{\text{int}} = \frac{e U_{\nu\nu}^*}{\sqrt{2} \sin \theta_W} \left[ \left( \overline{\nu}_{\mu L} \gamma^{\mu} \mu_L \right) W_{\mu}^+ + \text{H.c.} \right]
$$

$$
+\frac{e U_{\nu N}^*}{\sqrt{2} \sin \theta_W} [(\bar{N}_{\mu L} \gamma^\mu \mu_L) W_\mu^+ + \text{H.c.}]. \tag{4}
$$

The heavy neutrino mixing (i) reduces the  $W_{\nu\mu}$  coupling by a factor  $U_{\nu\nu}^*$  and (ii) induces a  $W N \mu$  coupling which allows an additional *t*-channel heavy neutrino  $(N_\mu)$  exchange [Fig.  $1(d)$ ] contribution.

#### **B. Helicity amplitudes**

The helicity amplitudes for the process

$$
\mu^-(k,\sigma) + \mu^+(\overline{k},\overline{\sigma}) \to W^-(q,\lambda) + W^+(\overline{q},\overline{\lambda})
$$
 (5)

(where the arguments indicate the four-momenta and helicities of the respective particles) are calculated following the technique described by Hagiwara *et al.* for the process  $e^+e^- \rightarrow W^+W^-$  [10]. We include heavy neutrino mixing and the Higgs-boson contribution and retain the muon mass terms to see the effects of heavy neutrino mixings on the cancellation of bad high energy behavior  $[11]$ . Following the Ref.  $[10]$ , we separate the contributions to the helicity amplitudes from the various graphs in Fig. 1 as  $[12]$ 

$$
M_{\sigma,\bar{\sigma}\lambda\bar{\lambda}}(\theta) = \sqrt{2}e^2 \tilde{M}_{\sigma,\bar{\sigma}\lambda\bar{\lambda}}(\theta) d_{\Delta\sigma,\Delta\lambda}^{J_0}(\theta), \qquad (6)
$$

where  $\Delta\lambda = \lambda - \overline{\lambda}$ ,  $\Delta\sigma = (\sigma - \overline{\sigma})/2$ ,  $J_0 = \max(|\Delta\sigma|, |\Delta\lambda|)$ , and  $\theta$  is the scattering angle of  $W^-$  with respect to the  $\mu^-$  direction in the  $\mu^+ \mu^-$ c.m. frame. The  $d_{\Delta\sigma,\Delta\lambda}^{J_0}$  are the *d* functions [13].  $J_0$  is the minimum angular momentum of the system. The amplitude  $\tilde{M}_{\sigma\bar{\sigma}\lambda\bar{\lambda}}$  includes contributions from the  $\gamma$ , *Z*,  $\nu_{\mu}$ ,  $N_{\mu}$ , and *H* exchange diagrams of Fig. 1, i.e.,

$$
\widetilde{M} = \widetilde{M}^{\gamma} + \widetilde{M}^Z + \widetilde{M}^{* \nu} + \widetilde{M}^{* N} + \widetilde{M}^H, \tag{7}
$$

where the asterisk on the  $\nu$  and  $N$  contributions is to remind us of mixing-modified terms. The  $\tilde{M}^{\gamma}$ ,  $\tilde{M}^{\gamma}$ , and  $\tilde{M}^H$  contributions are not affected by the heavy neutrino mixings. The explicit forms of these contributions [with  $\beta$ 

 $=\sqrt{1-(4m_W^2/s)}$ ,  $\gamma = \sqrt{s}/2m_W$ ,  $\sqrt{s}$  = total c.m. energy, and  $m<sub>\mu</sub>$ ,  $m<sub>W</sub>$ ,  $m<sub>Z</sub>$ , and  $m<sub>H</sub>$  the masses of the muon, *W*, *Z*, and *H* bosons, respectively] are

$$
\widetilde{M}^{\gamma} = -\beta \delta_{|\Delta \sigma|, I} A^{\gamma}_{\lambda \widetilde{\lambda}} \delta_{J_{0,1}} - \beta \delta_{\Delta \sigma, 0} \frac{\sqrt{2} m_{\mu}}{m_{W}}
$$

$$
\times A^{\prime \gamma}_{\lambda \widetilde{\lambda}} (\delta_{J_{0}, 1} + \cos \theta_{J_{0}, 0}), \qquad (8)
$$

$$
\tilde{M}^Z = \beta \left( \delta_{|\Delta \sigma, 1|} - \frac{\delta_{\Delta \sigma, -1}}{2 \sin^2 \theta_W} \right) A_{\lambda \tilde{\lambda} S}^Z - \frac{s}{m_Z^2}
$$
\n
$$
\times \delta_{J_0, 1} + \beta \delta_{\Delta \sigma, 0} \frac{\sqrt{2} m_\mu}{m_W} \left( 1 - \frac{1}{4 \sin^2 \theta_W} \right)
$$
\n
$$
\times A_{\lambda \tilde{\lambda}}^{\Delta' Z} \frac{s}{s - m_Z^2} (\delta_{J_0, 1} + \cos \theta \delta_{J_0, 0}), \tag{9}
$$

$$
\widetilde{M}^H = -\delta_{\Delta\sigma,0} \frac{\sqrt{2}m_{\mu}}{m_W} \frac{1}{2\sin^2\theta_W} H_{\lambda\widetilde{\lambda}} \frac{s}{s - m_H^2} \delta_{J_0,0}.
$$
\n(10)

Here the first term in  $\tilde{M}^{\gamma}, \tilde{M}^{\gamma}$  is the same as that in Ref. [10]. The second term in  $\tilde{M}^{\gamma}$ ,  $\tilde{M}^Z$  and the contribution from  $\tilde{M}^H$ arise due to the retention of muon mass terms  $[11]$ . The coefficients  $A_{\lambda\overline{\lambda}}$ ,  $A'_{\lambda\overline{\lambda}}$ , and  $H_{\lambda\overline{\lambda}}$  are given in Table I. The neutrino-mixing-modified helicity amplitudes are

$$
\widetilde{M}^{* \nu} = |U_{\nu \nu}|^2 \widetilde{M}^{\nu} \tag{11}
$$

and

$$
\widetilde{M}^{*N} = |U_{\nu N}|^2 \frac{1 + \beta^2 - 2\beta \cos \theta}{1 + \beta^2 - 2\beta \cos \theta + (4m_{N}^2/s)} \widetilde{M}^{\nu}, \quad (12)
$$

with

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$$
\tilde{M}^{\nu} = \frac{-\sqrt{2}}{\sin^2 \theta_W} \frac{\delta_{J_0,2}}{1 + \beta^2 - 2\beta \cos \theta} \left( \delta_{\Delta \sigma, -1} + \delta_{\Delta \sigma,0} \sqrt{\frac{2}{3}} \frac{m_{\mu}}{\sqrt{s}} \right)
$$
  
+ 
$$
\frac{1}{2 \sin^2 \theta_W \beta} \left[ \delta_{\Delta \sigma, -1} \left( B_{\lambda \tilde{\lambda}} - \frac{C_{\lambda \tilde{\lambda}}}{1 + \beta^2 - 2\beta \cos \theta} \right) \right]
$$

$$
\times \delta_{J_0,1} + \delta_{\Delta \sigma,0} \frac{\sqrt{2}m_{\mu}}{m_W} \frac{1}{2} \left( B_{\lambda \tilde{\lambda}}' - \frac{C_{\lambda \tilde{\lambda}}'}{1 + \beta^2 - 2\beta \cos \theta} \right)
$$

$$
\times (\delta_{J_0,1} + \delta_{J_0,0}) \Bigg|.
$$
(13)

Here the  $\tilde{M}^{\nu}$  is the contribution of the *t*-channel neutrino exchange in the absence of neutrino mixing. The terms without  $m<sub>u</sub>$  are the same as those in Ref. [10]. The extra terms with  $\delta_{\Delta\sigma,0}$  arise due to retention of muon mass. The coefficients  $B_{\lambda\overline{\lambda}}$ ,  $B'_{\lambda\overline{\lambda}}$ ,  $C_{\lambda\overline{\lambda}}$ , and  $C'_{\lambda\overline{\lambda}}$  are given in Table I.

# **III. HEAVY NEUTRINO MIXING EFFECTS AND HIGH ENERGY BEHAVIOR**

We combine the heavy neutrino mixing modified helicity amplitudes  $\tilde{M}^{*\nu}$  and  $\tilde{M}^{*\nu}$  in the following form [using Eq.  $(2)!$ 

$$
\widetilde{M}^{* \nu} + \widetilde{M}^{*N} = \widetilde{M}^{\nu} (1 - \Delta \widetilde{M}^{* \nu}), \tag{14}
$$

where

$$
\Delta \widetilde{M}^{* \nu} = |U_{\nu N}|^2 \left[ 1 - \frac{1 + \beta^2 - 2\beta \cos \theta}{1 + \beta^2 - 2\beta \cos \theta + (4m_N^2/2)} \right].
$$
\n(15)

The parameter  $\Delta \tilde{M}^{*\nu}$  aggregates the heavy neutrino mixing effects although heavy neutrino mixing adds one more *t*-channel heavy neutrino exchange diagram, but the overall effect is to reduce the neutrino exchange contribution in the helicity amplitude  $[Eq. (14)]$ . We note the following.

(i) In the presence of heavy neutrino mixing, the helicity amplitudes  $\tilde{M}^{\gamma}$ ,  $\tilde{M}^{\gamma}$ , and  $\tilde{M}^H$  remain unaffected, while the *t*-channel neutrino exchange helicity amplitude  $\tilde{M}^{\nu}$  is modified to  $\tilde{M}^{\nu} (1 - \Delta \tilde{M}^{*\nu}).$ 

(ii) The fractional change  $\Delta \tilde{M}^{*\nu}$  in the helicity amplitude for *t*-channel neutrino exchange, occurring due to heavy neutrino mixing, is always less than  $|U_{\nu N}|^2$ , where  $U_{\nu N}$  is the heavy neutrino mixing parameter [see Eq.  $(15)$ ]. In Fig. 2, the dependence of  $\Delta \tilde{M}^{*v}/|U_{vN}|^2$  on (a)  $\sqrt{s}$  for different  $m_N$ values, (b)  $m_N$  for different  $\sqrt{s}$  values, and (c)  $\theta$  for a fixed  $\sqrt{s}$  and different  $m_N$  values is shown.

(iii) In the high energy limit, the heavy neutrino mixing effects tend to vanish (even if mixing is large); that is, for  $\sqrt{s} \ge m_N$ ,  $4m_N^2/s \to 0$ , and as a result  $\Delta \tilde{M}^{* \nu} \to 0$ .

(iv) Since the muons are massive and may therefore be found in wrong helicity states, the *s*-wave scattering amplitude which exists in this case (case  $J_0=0$ ) grows as  $\sqrt{s}$  for the production of longitudinally polarized *W* bosons  $[(\lambda, \overline{\lambda})]$ 



FIG. 2. The variation of  $\Delta \widetilde{M}^{* \nu} / |U_{\nu N}|^2$  with (a)  $\sqrt{s}$ , the c.m. energy, (b)  $m_N$ , heavy neutrino mass, and (c) scattering angle  $(\theta)$ .

 $= (0,0)$ . In the SM, this is canceled by the *H*-exchange contribution. To see the effects of heavy neutrino mixing on this cancellation, we consider  $\mu_{R(L)}^+ \mu_{R(L)}^- \to W_L^+ W_L^-$ , that is the  $J_0=0$ ,  $\Delta \sigma=0$  case. For  $\theta=90^\circ$ , from Eqs. (8)–(13) we have  $\tilde{M}^{\gamma} = 0$ ,  $\tilde{M}^Z = 0$ ,

$$
\widetilde{M}^H = -\frac{m_\mu}{m_W^2} \frac{1}{4\sqrt{2}\sin^2\theta_W} \left(1 - \frac{2m_W^2}{s}\right) \left(\frac{s}{s - m_H^2}\right) \sqrt{s}
$$

and

$$
\widetilde{M}^{* \nu} + \widetilde{M}^{*N} = \frac{m_{\mu}}{m_W^2} \frac{1}{4\sqrt{2} \sin^2 \theta_W} \left(1 - \Delta \widetilde{M}^{* \nu}\right) \sqrt{s}.
$$



FIG. 3. The variation of  $|\widetilde{M}^{*\nu} + \widetilde{M}^{*\nu} + \widetilde{M}^{H}|$  with  $\sqrt{s}$  for various values of the heavy neutrino mass  $m_N$  and mixing parameter  $|U_{\nu N}|^2$ for the process  $\mu_{R(L)}^+ \mu_{R(L)}^- \to W_L^+ W_L^-$  at  $\theta = 90^\circ$ .

We plot in Fig. 3 the helicity amplitude  $|\tilde{M}^{*} + \tilde{M}^{*N}|$  $+\tilde{M}^H$  as a function of  $\sqrt{s}$  for heavy neutrino mass  $m_N$  $=$  50 GeV, 100 GeV and  $|U_{\nu N}|^2$  = 0.1 (small mixing),  $|U_{\nu N}|^2$  = 0.9 (large mixing) [2] along with SM expectations. We take  $m_{\mu}$ =105.658 MeV,  $m_W$ =80.41 GeV [2]. For the Higgs-boson mass, we take a representative value  $m_H$  $=125 \text{ GeV}$  [14]. Our conclusions are summarized below.

### **IV. CONCLUSIONS**

We considered heavy neutrino mixing effects in the helicity amplitudes for the process  $\mu^+\mu^-\rightarrow W^+W^-$ . We parametrized the mixing effects of heavy neutrino mixing as a fractional change  $\Delta \tilde{M}^{*\nu}$  in the helicity amplitude contribution from the *t*-channel  $v_{\mu}$  exchange diagram. We find the following.

(i)  $\Delta \tilde{M}^* V < |U_{\nu N}|^2$ , where  $U_{\nu N}$  is the heavy neutrino mixing parameter.

(ii) The dependence of  $\Delta \tilde{M}^{*\nu}$  on (a)  $\sqrt{s}$ , (b)  $m_N$ , and (c)  $\theta$  is as shown in Fig. 2.

(iii) We note from the data for Fig. 2 that for  $m_N$  $=10 \text{ GeV}$  at  $\sqrt{s} = 200 \text{ GeV}$ ,  $\Delta \tilde{M}^{*v} = 0.007 |U_{vN}|^2$  and for  $m_N = 10 \text{ GeV}$  at  $\sqrt{s} = 800 \text{ GeV}$ ,  $\Delta \tilde{M}^{* \nu} = 0.00002 |U_{\nu}y|^2$ . Thus, even if the neutrino mixing parameter  $U_{\nu N}$  is large, say,  $|U_{\nu N}|^2 \sim 1$ , then in the high energy limit  $(\sqrt{s})$ =800 GeV) the fractional change  $\Delta \tilde{M}^{*\nu}$  ~ 0.002% for  $m_N$ = 10 GeV and  $\Delta \tilde{M}^*$ <sup> $\nu$ </sup> ~ 0.8% for  $m_N$ = 50 GeV. At extremely high energies (i.e.,  $\sqrt{s} \gg m_N$ ), the heavy neutrino mixing effects tend to vanish  $(\Delta \tilde{M}^{*\nu} \rightarrow 0)$ .

 $(iv)$  From Fig. 3 we note that at low energies the neutrino mixing effectively reduces the neutrino exchange contribution, enhancing the value of  $|\tilde{M}^*|$  neutrino)  $+ \tilde{M}$ (Higgs boson) contribution in comparison to the SM value. At high energies, the neutrino mixing effects decrease. In the limit  $\sqrt{s} \rightarrow \infty$ , as  $\Delta \tilde{M}^{* \nu} \rightarrow 0$ , the neutrino mixing effects tend to vanish even if  $|U_{\nu N}|^2$  is large.

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