Weakly Broken Subsymmetries of Hadrons

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The relation between explicit and spontaneous breaking of chiral $SU(3) \times SU(3)$ symmetry is analyzed. Because it is assumed that the transformations of the asymptotic fields, even if nonlinear, are the same up to the linear term as they are in the limit of no explicit symmetry breaking, it is possible to solve the equations of the Glashow-Weinberg model. Using only the pseudoscalar-meson masses and the ratio of the kaon and pion leptonic decay constants as input, the remaining parameters of the model are computable. Numerical results are obtained for M_K^2 , $Z_K^{1/2}/Z_\pi^{1/2}$, $\epsilon_8/\sqrt{2}\epsilon_0$, and $\langle 0 |u_8 | 0 \rangle / \sqrt{2}$ $\langle 0 |u_0 | 0 \rangle$. Figures display the dependence of M_K on the ratio F_K/F_π . Two noteworthy features of the analysis are: (1) The Glashow-Weinberg model admits two types of solutions simultaneously, one characterized by chiral $SU(2) \times SU(2)$ as a weakly broken subsymmetry [i.e., by. the hadron Hamiltonian's being approximately invariant under chiral $SU(2) \times SU(2)$, which is the view of Gell-Mann, Oakes, and Renner, and the other by SU(3) as a weakly broken subsymmetry, which is the view of Brandt and Preparata. (2) The limit of no explicit symmetry breaking is realized with the entire spinless chiral multiplet degenerate in mass, and massless if either the full symmetry or any subsymmetry is spontaneously broken. It is made plausible that recent studies of the limit of no explicit symmetry breaking which find an octet of pseudoscalar Goldstone bosons and an SU(3)-invariant vacuum do so because their assumptions or approximations guarantee that they must.

I. INTRODUCTORY REMARKS AND SUMMARY OF RESULTS

That the hadron Hamiltonian should consist of one part H_0 which is invariant under chiral SU(3) \times SU(3) and another part H' which transforms as a $(3, 3^*)$ + $(3^*, 3)$ representation of the group was suggested by Gell-Mann.¹ Explicitly, the hypothesis was

$$
H = H_0 + H'
$$
\n(1a)

with

$$
H' = \int d\vec{\mathbf{x}} \left[\epsilon_0 u_0(x) + \epsilon_8 u_8(x) \right], \tag{1b}
$$

where the $u_i(x)$ belong to a nonet of scalar densities, which together with a nonet of pseudoscalar densities $v_i(x)$ make up a $(3, 3^*) + (3^*, 3)$ representation of chiral $SU(3) \times SU(3)$.

This idea has been reconsidered by Glashow and Weinberg² (GW), who make an additional assumption about smoothness of certain vertex functions and find inequalities to be satisfied by the κ -meson mass, a condition on the K_{13} form factor $f_+(t)$ at $t = 0$, and a formula which relates the masses of the pseudoscalar nonet and of the κ meson and which reduces to the Gell-Mann-Qkubo mass formula when the vacuum is SU(3)-invariant. In these equations appear the leptonic decay constants of the pseudoscalar mesons and κ meson, needed ratios of which cannot be determined without introducing additional assumptions (GW use spectralfunction sum rules).

The $(3, 3^*)$ + $(3^*, 3)$ model has also been reconsidered by Gell-Mann, Oakes, and Renner³ (GOR), who neglect the effect of $\eta-\eta'$ mixing and of scalar mesons and who assume pole dominance of the axial-vector current divergences by π , K , and η mesons and approximate SU(3) symmetry for matrix elements of the scalar densities between onepseudoscalar-meson states. In this context, QQR find that the hadron Hamiltonian is approximately chiral-SU(2) \times SU(2)-invariant and the vacuum approximately SU(3)-invariant: Explicitly,

$$
a \simeq -0.89 \tag{2a}
$$

and

$$
b \simeq 0 \tag{2b}
$$

where we have used the definitions

 $a \equiv \epsilon_{\rm a}/\sqrt{2} \epsilon_{\rm o}$ (3a)

and

$$
b \equiv \lambda_{\rm B} / \sqrt{2} \; \lambda_0 \; , \tag{3b}
$$

with

$$
\lambda_i \equiv \langle 0 | u_i(x) | 0 \rangle = \delta_{i0} \lambda_0 + \delta_{i8} \lambda_8 . \tag{4}
$$

[Strictly speaking, GOR find

$$
\lambda_{\rm s} \simeq 0 \tag{5}
$$

rather than Eg. (2b), which is then a consequence as long as $\lambda_0 \neq 0$. For the leptonic decay constants defined by

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$$
\langle 0 | A^B_\mu(0) | P_i(\vec{k}) \rangle \equiv i k_\mu F_{Bi} / [(2\pi)^3 2\omega_{\vec{k}}]^{1/2} , \qquad (6a)
$$

where

$$
F_{Bi} = F_B \delta_{Bi} \quad \text{for } B = 1, 2, \dots, 7
$$
 (6b)

(large Latin indices run from 1 to 8, small Latin indices from 0 to 8), GOR obtain

$$
F_{\pi} \simeq F_K \simeq F_{\eta} ; \tag{7}
$$

and for the renormalization constants defined by

$$
\langle 0 | v_i(0) | P_j(\vec{k}) \rangle \equiv Z_{ij}^{1/2} / [(2\pi)^3 2\omega_{\vec{k}}]^{1/2}, \qquad (8a)
$$

where

$$
Z_{ij}^{1/2} = Z_i^{1/2} \delta_{ij} \text{ for } i = 1, 2, ..., 7
$$
 (8b)

they obtain

$$
Z_{08}^{1/2} \simeq 0 \tag{9a}
$$

ancl

$$
Z_{\pi}^{-1/2} \simeq Z_{K}^{-1/2} \simeq Z_{\eta}^{-1/2} . \tag{9b}
$$

Or for the ratios

$$
R \equiv F_K / F_\pi \tag{10a}
$$

and

$$
\zeta \equiv Z_K^{-1/2}/Z_{\pi}^{-1/2}, \qquad (10b)
$$

their results are

$$
R\simeq 1
$$

and

$$
\zeta \simeq 1\,. \tag{11b}
$$

The result given in Eq. (2a) is correlated by GOR with the relative smallness of the mass of the physical pion, which, if the parameter a were exactly equal to -1 [corresponding to no explicit breaking of chiral $SU(2)\times SU(2)$, would be a massless, Goldstone boson. The result given in Eq. (2b) is close to the limiting value $b = 0$, for which the vacuum would be invariant under SU(3).

Recently, in a different scheme which also uses the $(3, 3^*) + (3^*, 3)$ model for broken chiral SU(3) \times SU(3) symmetry, Brandt and Preparata⁴ (BP) have instead found

$$
a \simeq -0.17 \tag{12a}
$$

 $R \simeq 1.2$, (12b)

and

$$
\xi \simeq 0.92 m_{K}^{2}/m_{\pi}^{2} \simeq 12. \qquad (12c)
$$

The result in Eq. (12a) is close to the limiting value $a=0$, which corresponds to no explicit breaking of SU(3) symmetry. The possibility of $a \approx 0$ was, of course, the original suggestion of Gell-Mann, ' but is controverted in the GOR reformulation.

The disparity between the GOR and BP versions

of Gell-Mann's model,

$$
a_{\text{GOR}} \simeq -0.89 \; , \quad \zeta_{\text{GOR}} \simeq 1 \; , \quad b_{\text{GOR}} \simeq 0 \; , \tag{13}
$$

vs

$$
a_{\text{BP}} \simeq -0.17
$$
, $\xi_{\text{BP}} \simeq 0.92 m_{K}^{2}/m_{\pi}^{2}$, $b_{\text{BP}} \simeq ?$, (14)

is due to their proximity to the limiting situations of $b=0$ vs $a=0$, SU(3)-invariant vacuum vs SU(3)invariant Hamiltonian. Both GQR-like and BP-like results may be obtained as solutions of the GW model. We find this to be so in our analysis.

Already Auvil and Deshpande' have shown that the GW model admits, under the conditions of no η - η' mixing $(Z_{08}^{1/2}=0)$ and of Gell-Mann-Okubo mass splitting for the 0^- octet, as possible limiting solutions

$$
b=0, \quad a=-0.89 \ , \quad \zeta=1 \ , \quad R=1 \ , \tag{15}
$$

and

(1la)

$$
a=0
$$
, $b=-0.89$, $\xi=m_{K}^{2}/m_{\pi}^{2}$, $R=1$. (16)

(We note that these values for ζ and R are not the only possibilities, and that we do not agree completely with their conclusions concerning the mass and leptonic decay constant of the κ meson.) The GOR results almost coincide with (15), and the BP results at least tend toward (16).

We consider herein what may be called a GW-like theory; By a GW theory we mean a theory of broken chiral symmetry which has three essential components: (1) Gell-Mann's hypothesis of $(3, 3^*)$ $+(3*,3)$ transformation properties for the symmetry-breaking part of the hadron Hamiltonian, (2) the possibility of spontaneous breaking of both SU(3) \times SU(3) and SU(3) symmetries, and (3) both 0⁻ and 0' mesons. The GOR and BP theories do not satisfy all these criteria: The scalar mesons and the SU(3) asymmetry of the vacuum have been either eliminated (GOR) or ignored (BP).

In GW theories, in order to get results for a, b , masses, etc. , hitherto there have been two ways to get needed relations among the leptonic decay constants: Either spectral-function sum rules are used to determine R and the other relevant ratios constants: Either spectral-function sum rules as
used to determine R and the other relevant ratios
of leptonic decay constants,^{2,6} or the renormaliza tion constants are constrained to be all equal due to various smoothness requirements 5,7 or all equal to unity in a tree-diagram approximation.^{8,9} When eque
all
8,9 ever the renormalization constants are so constrained, certain of the GW equations reduce to useful relations among the leptonic decay constants. However, as may be expected from Eqs. (15) and as we shall see later, the constraint $\zeta = 1$, or even $\xi \approx 1$, automatically prejudices a GW theory to favor a GOR world in which b is small and $a \approx -1$.

Therefore, the fact that Dutt, Eliezer, and Nanda,⁷ Intertucted the ract that Dutt, Effect, and Nair
Carruthers and Haymaker,⁸ and Olshansky⁹ find that the limit of no explicit symmetry breaking, ϵ_0 + 0 and ϵ_8 + 0, is realized in the GOR manner is perhaps not really a prediction of their analyses, but rather a self-consistent result. [Actually Olshansky also finds another possibility, $\lambda_0 = \lambda_8 = 0$ and all 0^{\pm} masses degenerate (the normal solution), with the choice between the GQR solution and the normal solution to be settled ultimately by the value of the π -N σ commutator.] It would be very interesting to.see, if possible, the results of an analysis as in Refs. 8 and 9 with the renormalization constants unconstrained.

Our theory is GW-like in that it contains the three essential components, except that the first component is modified, as will be discussed below. Because we are willing to make a commitment as to how the asymptotic (in- or out-) fields transform, we can obtain not only the GW equations, but also relations among the leptonic decay constants, so that we are able to solve the GW equations for arbitrary values of R , with no use of spectral-function sum rules and no prior constraint on ζ .

Although we do not believe that the hadron Hamiltonian transforms strictly as a $(3, 3^*) + (3^*, 3)$ reptonian transforms strictly as a $(3, 3^*) + (3^*, 3)$ rep-
resentation,^{10,11} we nevertheless determine the consequences of assuming that, at least for vacuumto-one-spinless-meson and vacuum-to-vacuum matrix elements of the commutators $[G_A, H]$ and $[G_A, [G_A, H]]$, respectively [where G_A is a generator of chiral $SU(3) \times SU(3)$, we can take H' to effectively belong to a $(3, 3^*) + (3^*, 3)$ representation $[i.e., we use Eqs. (1) only under these circum$ stances].

We limit our attention to the nonets of 0^{\pm} mesons, and we assume that, even when the hadron Hamiltonian is asymmetric, we can take the transformations of the asymptotic fields of the 0^{\pm} mesons to be the same, up to the terms linear in the fields, as they are¹² when H is invariant. If H were invariant, then, of course, the transformations of asymptotic fields are necessarily purely linear since the Hamiltonian when expressed in terms of asymptotic fields (which are free) is bilinear, no local nonlinear transformations of the fields can leave it invariant. When H is not required to be invariant, we can no longer expect the transformations of asymptotic fields to be at most linear (or even purely local), but, for simplicity, we assume that the coefficients of the leading local terms are
not appreciably modified.¹³ not appreciably modified.

The spontaneous breakdown of the symmetry is introduced¹² via the appearance, in the commutators of the generators with the asymptotic fields, of c numbers C_0 and C_8 added to the $I = Y=0$ scalar asymptotic fields $S_0(x)$ and $S_8(x)$: $C_0 = C_8 = 0$ is required for a chiral-SU(3) \times SU(3)-invariant vacuum, $C_a = 0$ for an SU(3)-invariant vacuum, $C_a = 2\sqrt{2}C₀$ for a chimeral-SU(3)-invariant¹⁴ vacuum, and C_8 $=-\sqrt{2}C_0$ for a chiral-SU(2)×SU(2)-invariant vacuum. These c numbers are expressible¹² in terms of the leptonic decay constants of the pseudoscalar nonet and of the κ meson, so that, for example, C_8 = 0 implies F_K = 0 and F_{π} = F_K , etc. Since there are only two c numbers, there can only be two independent leptohic decay constants among the five that appear; and it is just the consequent relations among the five leptonic decay constants which enable us to go beyond the results of the usual GW theories, without constraining the renormalization constants.

The effects of the existence of the c numbers in the limit of no explicit breaking of chiral $SU(3)$ \times SU(3) symmetry, in which case the nonlinear and/or nonlocal terms in the commutators of the generators with the asymptotic fields [the unwritten terms indicated by the dots in Eqs. (22)] vanish, can be determined by using the explicit expression for H in terms of asymptotic fields to determine the implications of $0 = \dot{G}_A = -i[G_A, H]$ (and also
by examining $\langle 0 | [G_A, \phi, ||0 \rangle]^{15}$ where ϕ_i is an asymp by examining $\langle 0 | [G_A, \phi_i] | 0 \rangle$,¹⁵ where ϕ_i is an asymp totic field). In the case of no mixing, for the SU(3) generator T_A , we have

$$
\dot{T}_A = f_{Aij} \int d\vec{x} (M_i{}^2 C_j S_i + m_i{}^2 P_i P_j + M_i{}^2 S_i S_j) ,
$$

while for the chiral generator X_A , we have

$$
\dot{X}_A = d_{Aij} \int d\vec{x} [m_i{}^2 C_j P_i + (m_i{}^2 - M_j{}^2) P_i S_j].
$$

Requiring the coefficients of the bilinear terms in \dot{T}_A to vanish implies mass degeneracy within the 0^{\degree} and 0^{\degree} octets: $\dot{T}_k = 0$ implies $m_{\pi}^2 = m_K^2$, m_K^2 $=m_{\eta}^2$, $M_{\delta}^2 = M_{\kappa}^2$, and $M_{\kappa}^2 = M_{\sigma}^2$; requiring the bilinear coefficients in X_A to vanish implies mass degeneracy within the entire 0^{\pm} chiral multiplet: $m_{\eta}^2 - m_{\delta}^2$, and $m_{\eta'}^2 - m_{\delta}^2$, $X_R = 0$ implies $m_{\pi}^2 - 1$
 $m_{K^2}^2 - M_{\delta}^2$, $m_{K^2}^2 = M_{\delta}^2$, $m_{K^2}^2 = M_{\delta}^2$, and admittedly, unsurprising results, since, when H is invariant, we assign the 0^{\pm} asymptotic fields to a linear representation. Since the c numbers are expressible in terms of the leptonic decay constants, requiring the coefficients of the linear terms in \dot{G}_A to vanish implies that either the 0⁺ mesons are massive and their leptonic decay constants are zero (the normal case) or the 0^{\pm} mesons are massless and their leptonic decay con-

FIG. 1. Plot of M_{κ}^2 vs R for case I, mixing angle $\theta = 0$.

 $\theta = -10.6^\circ$.

FIG. 2. Plot of M_{κ}^2 vs R for case II, mixing angle $\theta = 10.6^\circ$.

FIG. 4. Plot of M_{κ}^2 vs R for case IV, $\xi_1 = 1$ and $\zeta_2 = m_K^2/n$

stants need not be zero or equal (the spontaneous breakdown case). Note that if $an\theta$ subsymmetry is spontaneously broken (i.e., the relevant leptoni decay constant is different from zero), all eighteen $0[±]$ mesons are massless due to the degeneracy. Qur view of the limit of no explicit symmetry breaking differs from the popular view due to $GOR₁³$ who have only π , K, and η mesons massless, with F_{π} $=F_K\neq 0$ and $F_K=0$. But, as we have mentioned before, they are working in a theory where η' and 0^+ mesons and SU(3) vacuum asymmetry are neglected.

Upon deriving the GW equations and our relations among the leptonic decay constants, we are able to obtain equations for M_k^2 , for ζ , for a, for b, and for $C_8/\sqrt{2}C_0$, which require as basic input data only the π , K, η , and η' masses and the F_{κ}/F_{π} ratio R. For the κ -mass equation, we consider four cases: (I) no mixing; (II) positive mixing angle; (III) negative mixing angle; and (IV) $\xi = 1$, or $\xi = m_{K}^{2}/m_{\pi}^{2}$, constraints which yield a well-known equation which expresses M_{κ}^2 simply in terms of m_{κ}^2 , m_{π}^2 , and R. We exhibit plots (see Figs. 1-4) of M_{κ}^2 as a function of R , showing the existence of certain upper and lower bounds for M_{κ}^2 and showing that certain values of R correspond to negative M_r^2 and are consequently forbidden, and also numerical results (see Tables I—IV), using a few interesting values of R, for M_{κ}^2 , ξ , a, b, and $C_8/\sqrt{2}C_0$.

Since a general GW theory with no constraints on Since a general GW theory with no constraints of yields a quadratic equation for ζ ,¹⁶ we obtain two solutions: (1) a solution similar to that favored by GOR: $\zeta_1 = O(1)$; and (2) a solution similar to that favored by BP: $\zeta_2 = O(m_{K}^{2}/m_{\pi}^{2})$. Corresponding to these solutions, we get, e.g., with no mixing and with $R = 1.28$, the results $a_1 \approx -0.91$, $b_1 \approx -0.14$ and $a_2 \approx -0.14$, $b_2 \approx -0.91$; we also get $C_8/\sqrt{2}C_0 \approx -0.16$, this result being independent of the value of ζ . The fact that $a_1 = b_2$ and $a_2 = b_1$ is no accident – it can be

TABLE II. Results for case II, mixing angle $\theta = 10.6$ °.

	R M_{κ}^{2} (GeV ²) ξ_1 ξ_2 $a_1 = b_2$ $b_1 = a_2$ $C_8/\sqrt{2}C_0$				R M_{κ}^{2} (GeV ²) ξ_{1} ξ_{2} $a_{1} = b_{2} b_{1} = a_{2} C_{8}/\sqrt{2}C_{0}$			
	$1.18 \t 0.540$		0.896 14.6 -0.915 -0.0368 -0.107	1.18	1.51		$1\quad 13.05\quad -0.906\quad -0.107\quad -0.107$	
1.23	0.547		0.886 14.7 -0.919 -0.0564 -0.133	1.23	1.23		$1 \quad 13.05 \quad -0.909 \quad -0.133 \quad -0.133$	
1.28	0.551		0.881 14.8 -0.923 -0.0785 -0.157	1.28	1.06		$1 \quad 13.05 \quad -0.913 \quad -0.157 \quad -0.157$	
	1.34 0.555		0.881 14.8 -0.926 -0.107 -0.185	1.34	0.914		$1 \quad 13.05 \quad -0.916 \quad -0.185 \quad -0.185$	

TABLE III. Results for case III, mixing angle $\theta = -10.6^{\circ}$.

\boldsymbol{R}	M_{κ}^{2} (GeV ²) ξ_1 ξ_2 $a_1 = b_2$ $b_1 = a_2$ $C_8/\sqrt{2}C_0$			
1.18	2.70			1.17 11.1 -0.890 -0.204 -0.107
1.23	2.33	1.27 10.3		$-0.886 - 0.272 - 0.133$
1.28	2.09			$1,41$ 9.28 - 0.879 - 0.348 - 0.157
1.34	1.90			1.65 7.90 -0.865 -0.448 -0.185

shown that there is a single equation relating a and b with only m_{π} , m_{K} , m_{η} , and $m_{\eta'}$ as parameters, which is symmetric under the interchange $a \leftrightarrow b$. This same symmetry in the $a-b$ plane is manifested in the Mathur-Okubo $plot^{14}$ of the allowed domains for broken chiral symmetry in the GW model.

We can determine under what circumstances the simple case IV ($\zeta = 1$ or $\zeta = m_{\kappa}^{2}/m_{\pi}^{2}$) could serve as an approximation to cases I, II, or III (mixing angle $\theta=0$, $\theta>0$, or $\theta<0$). We find that, for reasonable values of R , case IV can approximate only case I (and then only over a very short range); specifically, we obtain

$$
for R \approx 1.32, \qquad (17a)
$$

$$
(M_{\kappa}^{2})_{IV} \simeq (M_{\kappa}^{2})_{I} \simeq 0.96 \text{ GeV}^{2}, \qquad (17b)
$$

$$
\xi_1 \simeq 1.00
$$
 and $\xi_2 \simeq 13.05 = m_{K}^{2}/m_{\pi}^{2}$, (17c)

$$
a_1 = b_2 \approx -0.91
$$
 and $b_1 = a_2 \approx -0.18$, (17d)

$$
C_8/\sqrt{2}C_0 \simeq -0.18 \ . \tag{17e}
$$

However, the mass required, $M_{\nu} \approx 980$ MeV, lies between the values, ~890 MeV and ~1200 MeV,
which have been suggested.¹⁷ which have been suggested.¹⁷

Although our preliminary results, Eqs. (29), (33), and (35), were also obtained by Auvil and Deshpande,⁵ apart from their limiting GOR-like and BPlike solutions, our Eqs. (15) and (16), their further work is restricted by the condition $\zeta = 1$.

II. THEORY

Since the hadron Hamiltonian, when expressed in terms of the asymptotic fields, must describe free hadrons, for the 0^{\pm} mesons we must have

$$
H = (\text{kinetic terms}) + H', \qquad (18)
$$

TABLE IV. Results for case IV, $\xi_1 = 1$ and $\xi_2 = m_K^2/m_{\pi}^2$.

	R M_{κ}^{2} (GeV ²) ξ_1 ξ_2			$a_1 = b_2$, $b_1 = a_2$, $C_8/\sqrt{2}C_0$
1.18 1.23 1.28 1.34	1.51 1.23 1.06 0.914		$1\quad 13.05\quad -0.906\quad -0.107$ $1 \quad 13.05 \quad -0.909 \quad -0.133$ $1 \quad 13.05 \quad -0.913 \quad -0.157$ $1 \quad 13.05 \quad -0.916 \quad -0.185 \quad -0.185$	-0.107 -0.133 -0.157

with

 $\boldsymbol{6}$

$$
H' = \int d\vec{x} \frac{1}{2} (m_{\pi}^{2} \pi^{2} + m_{\kappa}^{2} K^{2} + m_{\eta}^{2} \eta^{2} + m_{\eta'}^{2} \eta'^{2} + M_{\delta}^{2} \delta^{2} + M_{\kappa}^{2} \kappa^{2} + M_{\sigma}^{2} \sigma^{2} + M_{\sigma'}^{2} \sigma'^{2}).
$$
\n(19)

The hadron Hamiltonian as given by Eqs. (18) and (19) is exactly equal to the expression for the hadron Hamiltonian in terms of interpolating fields.

In terms of the SU(3) bases, we have

$$
H' = \int d\vec{x} \frac{1}{2} (m_i^2 P_i^2 + \Delta P_0 P_8 + M_j^2 S_j^2 + \Delta_S S_0 S_8),
$$
\n(20)

where there is a sum on i and j from 0 to 8 and where we have introduced singlet-octet mixing:

$$
P_8 = \eta \cos \theta + \eta' \sin \theta ,
$$

\n
$$
P_0 = -\eta \sin \theta + \eta' \cos \theta ,
$$

\n
$$
m_8^2 = m_{\eta}^2 \cos^2 \theta + m_{\eta'}^2 \sin^2 \theta ,
$$

\n
$$
m_0^2 = m_{\eta}^2 \sin^2 \theta + m_{\eta'}^2 \cos^2 \theta
$$

and

$$
\Delta = (m_{\eta'}^2 - m_{\eta}^2) \sin 2\theta
$$

with analogous equations for the scalar nonet. The question of whether or not we should use the Gell-Mann-Okubo mass formula in this theory will be discussed shortly. If we do not impose it, then not all the parameters introduced by mixing can be specified by knowing only the masses of π , K , η , and η' mesons; hence, we neglect mixing and put $\theta = 0$ (our case I). If we do impose it, then all the parameters introduced by mixing can be specified as usual,

$$
m_{8}^{2} = \frac{1}{3} (4m_{K}^{2} - m_{\pi}^{2}),
$$

$$
m_{0}^{2} = m_{\eta}^{2} + m_{\eta'}^{2} - m_{8}^{2},
$$

and

$$
\Delta = \pm 2[(m_{n'}^2 - m_{\rm s}^2)(m_{\rm s}^2 - m_{n}^2)]^{1/2},
$$

where the plus or minus sign corresponds to positive or negative mixing angle θ (cases II and III).

If Gell-Mann's hypothesis' is correct, then we should be able to equate the right-hand sides of Eqs. (Ib) and (20). Actually the right-hand side of Eq. (20) contains a small amount of $SU(3)\times SU(3)$ singlet, but that is of no consequence since H' is only used in commutators with the generators. We put.

$$
\epsilon_0 u_0 + \epsilon_8 u_8 = \frac{1}{2} (m_1^2 P_i^2 + \Delta P_0 P_8 + M_j^2 S_j^2 + \Delta_S S_0 S_8),
$$
\n(21)

with the stipulation that this "equality" only be used as described in Sec. I.

The scalar and pseudoscalar densities in the the-

ory are assigned to the $(3, 3^*)+(3^*,3)$ representation. For the 0^{\pm} asymptotic fields, we make the assumption that

$$
[T_A, S_i] = i f_{Ai}(S_i + C_i) + \cdots,
$$
 (22a)

$$
[T_A, P_i] = i f_{Ai} P_j + \cdots, \qquad (22b)
$$

$$
[X_A, S_i] = -i d_{Ai} P_j + \cdots, \qquad (22c)
$$

and

$$
[X_A, P_i] = id_{Aij}(S_j + C_j) + \cdots,
$$
 (22d)

where

$$
C_i = \delta_{i0} C_0 + \delta_{i8} C_8 \tag{23}
$$

is a c number. Since the Hamiltonian is not invariant, we admit the possibility of nonlinear terms (indicated by the dots). However, for the sake of simplicity, we tentatively retain the same coefficients for the leading terms as would be necessary were the Hamiltonian invariant.

If we believed Eq. (21) to be a strict equality, then we could use Eqs. (22) to compute $\langle K^{\dagger}[[T_{\kappa^+},\dot{T}_{\kappa^+}]|K^+\rangle$. This would yield the Gell-Mann– Okubo mass formula. However, we do not know a priori that the hadron Hamiltonian, Eqs. (18) -(20), contains no SU(3) representations higher than the octet. Hence we allow for the possibility that the Gell-Mann-Okubo mass formula may (cases II and III} or may not (case I) hold. For case IV, it makes no difference.

The c numbers are necessarily present in the above commutators, Eqs. (22), because, for example,

$$
\langle 0|[X_{\pi}(x_0), \pi(x)]|0\rangle = [\langle 0|A_0^{\pi}(0)|\pi(\vec{0})\rangle
$$

– $\langle \pi(\vec{0})|A_0^{\pi}(0)|0\rangle][(2\pi)^3/2m_{\pi}]^{1/2}$
= iF_{π} . (24)

Thus, from Eqs. (22a) and (22d) we obtain

 $F_{\kappa}=\frac{1}{2}\sqrt{3}C_{8}$ (25a)

and

$$
F_{Ai} = d_{Aij} C_j , \qquad (25b)
$$

so that, with Eq. (23),

$$
F_{\pi} = (\sqrt{2} C_0 + C_8) / \sqrt{3} \,, \tag{26a}
$$

$$
F_{K} = (\sqrt{2} C_0 - \frac{1}{2} C_8) / \sqrt{3} , \qquad (26b)
$$

$$
C_8/\sqrt{2}C_0 = -2(R-1)/(2R+1) , \qquad (27)
$$

$$
F_{\pi} = F_K + F_{\kappa} \quad , \tag{28a}
$$

$$
F_{88} = \frac{1}{3}(4F_K - F_\pi), \qquad (28b)
$$

and

$$
F_{80} = \frac{2}{3}\sqrt{2}(F_{\pi} - F_{K}).
$$
 (28c)

From Eq. (27) it is clear that the limits of invariance of the vacuum under SU(3), $R \rightarrow 1$, under chimeral SU(3), $R \rightarrow 0$, and under chiral SU(2) \times SU(2), $R \rightarrow \infty$, are realized when $C_s/\sqrt{2}C_0$ is equal to 0, 2, and -1, respectively. Equations (28) are just what enable us to go beyond the usual GW theories.

If we use Eq. (21) and the transformations of both the 0^{\pm} densities and the 0^{\pm} asymptotic fields to evaluate $\langle 0[[G_{\Delta},H]]1 \rangle$, we get

$$
\epsilon_{\kappa} Z_{\kappa}^{1/2} = M_{\kappa}^{2} F_{\kappa} , \qquad (29a)
$$

$$
\epsilon_{\pi} Z_{\pi}^{-1/2} = m_{\pi}^2 F_{\pi} , \qquad (29b)
$$

$$
\epsilon_K Z_K^{-1/2} = m_K^2 F_K , \qquad (29c)
$$

$$
\epsilon_{88} Z_{88}^{1/2} + \epsilon_{80} Z_{08}^{1/2} = m_8^2 F_{88} + \frac{1}{2} \Delta F_{80} , \qquad (29d)
$$

and

$$
\epsilon_{88} Z_{80}^{1/2} + \epsilon_{80} Z_{00}^{1/2} = m_0^2 F_{80} + \frac{1}{2} \Delta F_{88} , \qquad (29e)
$$

where we have used the definitions

 $\epsilon_{ij} \equiv d_{ijk} \epsilon_k$, $(30a)$

$$
\epsilon_{\pi} \equiv \epsilon_{11} = \epsilon_{22} = \epsilon_{33} , \qquad (30b)
$$

$$
\epsilon_K \equiv \epsilon_{44} = \epsilon_{55} = \epsilon_{66} = \epsilon_{77} , \qquad (30c)
$$

$$
\epsilon_{\kappa} \equiv \frac{1}{2} \sqrt{3} \, \epsilon_{8} \,, \tag{30d}
$$

and

$$
Z_{\kappa}^{\ 1/2} \equiv \langle 0 | u_{\kappa}(0) | \kappa(\vec{k}) \rangle \big[(2\pi)^3 2\omega_{\vec{k}} \big]^{1/2} . \tag{31}
$$

If we expand¹⁸ the scalar and pseudoscalar densities $u_i(x)$ and $v_i(x)$ in terms of asymptotic fields,

$$
u_i(x) = \lambda_i + Z_{ij}^{(s) 1/2} S_j(x)
$$
\n
$$
+ \int dy dz : P_j(y) P_k(z) : F_{ijk}(xyz)
$$
\n
$$
+ \int dy dz : S_j(y) S_k(z) : G_{ijk}(xyz) + \cdots
$$
\n
$$
= \epsilon_{\kappa} \lambda_{\kappa} = m_{\kappa}^2 F_{\kappa}^2,
$$
\n
$$
+ \int dy dz : S_j(y) S_k(z) : G_{ijk}(xyz) + \cdots
$$
\n(32a) (32a)

and

(32a)

$$
v_i(x) = Z_{ij}^{1/2} P_j(x) + \int dy dz : S_j(y) P_k(z) : H_{ijk}(xyz) + \cdots,
$$

(32b)

and evaluate, again using the transformation properties of both the 0^{\pm} densities and the 0^{\pm} asymptotic fields, the matrix elements $\langle 0| [T, u] |0 \rangle$ and

$$
\langle 0 \| X, v \| 0 \rangle, \text{ we find}
$$

$$
\lambda_{\kappa} = F_{\kappa} Z_{\kappa}^{1/2}, \qquad (33a)
$$

$$
\lambda_{\pi} = F_{\pi} Z_{\pi}^{1/2},\tag{33b}
$$

$$
\lambda_K = F_K Z_K^{-1/2},\tag{33c}
$$

$$
\lambda_{88} = F_{88} Z_{88}^{1/2} + F_{80} Z_{80}^{1/2} , \qquad (33d)
$$

and

$$
\lambda_{80} = F_{80} Z_{00}^{1/2} + F_{88} Z_{08}^{1/2} , \qquad (33e)
$$

where we have used the definitions

$$
\lambda_{ij} \equiv d_{ijk} \lambda_k \tag{34a}
$$

$$
\lambda_{\pi} \equiv \lambda_{11} = \lambda_{22} = \lambda_{33} , \qquad (34b)
$$

$$
\lambda_K \equiv \lambda_{44} = \lambda_{55} = \lambda_{66} = \lambda_{77} , \qquad (34c)
$$

and

$$
(29a) \qquad \qquad \lambda_{\kappa} \equiv \frac{1}{2} \sqrt{3} \lambda_{8} \,.
$$

Although Eqs. (29) are, within the context of our assumptions, exact, in Eqs. (33) we have neglected effects coming from commutators of the generators with the bilinear and higher terms in the expansions for u and v . The terms we have neglected can be shown to arise from multiparticle intermediate states. How Eqs. (33) can be derived by considering only single-particle intermediate states has been demonstrated by Auvil and Deshpande.⁵ That our approximation is sufficient for our purposes will become clear shortly. In a theory with the renormalization constants constrained to be all diagonal and to be all equal or all equal to unity, Eqs. (33) simplify and imply the very useful Eqs. (28). In the limit of no explicit symmetry breaking, Eqs. (33) are just a straightforward application of a general formula derived in Ref. 15, and are exact.

Eliminating the renormalization constants by simultaneous use of Eqs. (29) and (33) yields

$$
\epsilon_{\kappa} \lambda_{\kappa} = M_{\kappa}^2 F_{\kappa}^2 , \qquad (35a)
$$

$$
\epsilon_{\pi} \lambda_{\pi} = m_{\pi}^2 F_{\pi}^2 , \qquad (35b)
$$

$$
\epsilon_K \lambda_K = m_K^2 F_K^2 \,, \tag{35c}
$$

and

 2Γ $2 \pm m$ 2Γ 2Γ $88\lambda_{88} + \epsilon_{80}\lambda_{80} = m_8^2F_{88}^2 + m_0^2F_{80}^2 + \Delta F_{88}F_{80}$ (35d)

Even in a theory with the renormalization constants constrained, these relations still hold, of course. It is interesting to note that these same equations (35) could have been derived directly by evaluating, using Eq. (21) and the transformation properties of both the 0^{\pm} densities and the 0^{\pm} asymptotic fields, the matrix elements $\langle 0|[T_{\kappa}, \dot{T}_{\kappa}]|0\rangle$, $\langle 0[[X_{\pi}, \dot{X}_{\pi}]]0 \rangle$, $\langle 0[[X_K, \dot{X}_K]]0 \rangle$, and $\langle 0[[X_8, \dot{X}_8]]0 \rangle$. Thus, the assumptions $[Eqs. (21)$ and $(22)]$ used in such a derivation and in the derivation of Eqs. (29) are clearly compatible with the assumption [Eqs. (22)] and one-particle-pole approximation used in the derivation of Eqs. (33).

By means of Eqs. (30) and (34) it is possible to show that Eqs. (35) can be combined into¹⁹

$$
4(M_{\kappa}{}^{2}F_{\kappa}{}^{2} + m_{\kappa}{}^{2}F_{\kappa}{}^{2})
$$

= 3(m_{\delta}{}^{2}F_{8\delta}{}^{2} + m_{0}{}^{2}F_{8\delta}{}^{2} + \Delta F_{8\delta}F_{8\delta}) + m_{\pi}{}^{2}F_{\pi}{}^{2}. (36)

Suppose we do not impose the Gell-Mann-Okubo mass formula and we take $\theta = 0$ (our case I); if we put in the experimental 0^- masses and then, reput in the experimental σ masses and then, $r =$
calling Eqs. (28), study the limit $C_8 \rightarrow 0$ and $|C_0| < \infty$ $(i.e., R-1),$ which is the limit of an SU(3)-invari ant vacuum, we find $M_{\kappa}^2 \rightarrow -\infty$. On the other hand, suppose we do impose the Gell-Mann-Okubo mass formula and we consider $\theta \neq 0$ (our cases II and III); if we repeat the procedure, we find that Eq. (36) just degenerates into the Gell-Mann-Okubo mass formula while M_{κ}^2 drops out and is thus unspecified.

The explicit functional forms we find from Eq. (36) using Eqs. (28) and the experimental data²⁰ for the 0^- masses are: for $\theta = 0$ (case I),

$$
M_{\kappa}^{2}=0.767+\frac{0.110}{R-1}-\frac{0.0158}{(R-1)^{2}};
$$
 (37a)

for $\theta > 0$ (case II),

$$
M_{\kappa}^{2} = 0.571 - \frac{0.00583}{R - 1};
$$
 (37b)

for θ <0 (case III),

$$
M_{\kappa}^{2} = 0.990 + \frac{0.308}{R - 1} \tag{37c}
$$

In Figs. 1, 2, and 3 we have plotted these expressions for M_k^2 as a function of R. At $R = 1$, the curves exhibit the behavior described above. We can, if we wish, establish certain bounds on R since those values corresponding to M_{κ}^2 <0 must be unphysical. Since the experimental value of M_{κ}^2

is still rather uncertain,¹⁷ we attempt no predictio for R . However, we can rule out case II if we regard $R < 1$ as unlikely.

Since Eqs. (30) imply $\epsilon_{\pi} = \epsilon_K + \epsilon_{\kappa}$, we conclude from Eqs. (29) that²

$$
m_{\pi}^{2}F_{\pi}Z_{\pi}^{-1/2} = m_{K}^{2}F_{K}Z_{K}^{-1/2} + M_{K}^{2}F_{K}Z_{K}^{-1/2};
$$
 (38)

since Eqs. (34) imply $\lambda_{\pi} = \lambda_K + \lambda_K$, we conclude from Eqs. (33) that²

$$
F_{\pi}Z_{\pi}^{-1/2} = F_K Z_K^{-1/2} + F_{\kappa} Z_{\kappa}^{-1/2} . \tag{39}
$$

These equations are analogous in form to our earlier result

$$
F_{\pi} = F_K + F_{\kappa} \tag{28a}
$$

If we put $Z_{\pi}^{-1/2} = Z_K^{-1/2} = Z_{\kappa}^{-1/2}$ (i.e., $\xi = 1$), then Eq. (39)reduces to Eq. $(28a)$, and Eq. (38) can be used to compute M_{κ}^2 ; if we put the various $Z^{1/2}$'s proportional to the corresponding masses squared (i.e., $\zeta = m_{\kappa}^2/m_{\pi}^2$ and $Z_{\kappa}^{1/2}/Z_{\pi}^{1/2} = M_{\kappa}^2/m_{\pi}^2$, then Eq. (38) reduces to Eq. (28a), and Eq. (39) can be used to compute M_{κ}^2 . Thus, for either $\zeta = 1$ or $\zeta = m_{\kappa}^2/2$ m_{π}^{2} , Eqs. (38) and (39) reduce to Eq. (28a) and to

$$
M_{\kappa}^{2} = \frac{Rm_{K}^{2} - m_{\pi}^{2}}{R - 1}
$$

= 0.246 + $\frac{0.227}{R - 1}$ (40)

In Fig. 4, we have M_{κ}^2 as a function of R for this case IV. At $R = 1$, M_{κ}^2 is, as in cases II and III, unspecified.

It is more interesting, however, to eliminate $Z_{\kappa}^{1/2}$ and F_{κ} by combining Eqs. (38), (39), and (28a) to get a single quadratic equation for ζ in terms of M_{κ}^{2} , m_{κ}^{2} , m_{π}^{2} , and R. The solutions are

$$
\zeta = (m_{\pi}^{2} + m_{K}^{2}R^{2} + M_{\kappa}^{2}(1 - R)^{2} \pm \{ [m_{\pi}^{2} + m_{K}^{2}R^{2} + M_{\kappa}^{2}(1 - R)^{2}]^{2} - 4m_{\pi}^{2}m_{K}^{2}R^{2}]^{1/2} \} / 2m_{\pi}^{2}R .
$$
 (41)

We will denote the solution with the negative square root as ζ_1 and that with the positive as ζ_2 . If we use the κ -mass equations, Eqs. (36) and (40), to determine $\lim_{R\to 1} M_{\kappa}^2 (1 - R)^2$, we find: For case I (no Gell-Mann-Okubo mass formula) and $R = 1$ that ζ_1 \approx 1 and $\zeta_2 \approx m_{\kappa}^2/m_{\pi}^2$; for cases II and III (with Gell-Mann-Okubo mass formula) and $R = 1$ that ζ = 1 and $\zeta_2 = m_{\kappa}^2/m_{\pi}^2$; and for case IV (without need for Gell-Mann-Okubo mass formula) and $R = 1$ that $\zeta_1 = 1$ and $\zeta_2 = m_{K}^2/m_{\pi}^2$. The results for case IV are, of course, just what we assumed in order to get Eq. (40).

Using Eqs. $(29b)$, $(29c)$, and $(30a)$ – $(30c)$, we get

$$
a = -2(m_{K}^{2}R - m_{\pi}^{2}\zeta)/(2m_{K}^{2}R + m_{\pi}^{2}\zeta); \qquad (42)
$$

using Eqs. (33b), (33c), and $(34a)$ - $(34c)$, we get

$$
b = -2(R\zeta - 1)/(2R\zeta + 1) \,.
$$
 (43)

For case I and $R=1$, we have $a_1 = b_2 \approx -0.9$ and a_2 $=b_1 \approx -0.05$; for cases II, III, and IV and $R=1$, we have $a_1 = b_2 = -0.89$ and $a_2 = b_1 = 0$. The GOR results (without mixing, but with the Gell-Mann-Okubo mass formula) seem to coincide with the solution (ζ_1, a_1, b_1) for cases II, III, and IV at $R = 1$.

It is easy to verify, by use of Eq. (41) that $a_1 = b_2$ and $a_2 = b_1$ for any R. Alternatively, if one exploits Eqs. (35b)-(35d) in such a way as to eliminate R , one gets an equation (with only the $0⁻$ masses as parameters) which is manifestly symmetric under the interchange $a \rightarrow b$ and which for each b has a possibility of as many as two real solutions for a and vice versa. Therefore, $a_1 = b_2$ and $a_2 = b_1$ is not unexpected.

In Tables I-IV, we give the results for ξ , a , b , and $C_{\rm g}/\sqrt{2}C_{\rm o}$ for a few interesting values of R in the cases I-IV. It is obvious that, at least for $1.34 \ge R \ge 1.18$, solution 1 is more nearly GOR-like [weakly broken chiral $SU(2) \times SU(2)$ symmetry] while solution 2 is more nearly BP-like [weakly broken SU(3) symmetry].

By simultaneously solving Eqs. (40) and (37) , we can find for which values of R case IV coincides with cases I-III. Cases IV and I coincide at R \approx 0.91 (M_{κ}^{2} < 0) and at $R \approx$ 1.32 (further details were given in Sec. I); however, since near $R = 1.32$ Fig. 1 exhibits much more curvature than Fig. 4, it is clear that case IV can serve as an approximation to case I only over a very small range of R . Case IV and II coincide at $R \approx 1.72$ (not interesting) and at $R = 1$. Cases IV and III coincide at $R \approx 0.89$ $(M_{r}^{2}<0)$ and at $R=1$.

III. FINAL REMARKS

The results which follow when our Eqs. (28) hold, are obtainable because we have made a conjecture,

¹M. Gell-Mann, Phys. Rev. 125, 1067 (1962).

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- 3M. Gell-Mann, R. J. Oakes, and B. Renner, Phys. Rev. 175, 2195 (1968).
- 4 R. A. Brandt and G. Preparata, Ann. Phys. (N.Y.) 61, 119 (1970); Phys. Rev. Letters 26, 1605 (1971); R. A.
- Brandt, CERN Report No. TH-1402, 1971 (unpublished). ⁵P. R. Auvil and N. G. Deshpande, Phys. Rev. 183, 1463 (1969).
	- 6 J. Cleymans, Nuovo Cimento 65A, 72 (1970).
- ⁷R. Dutt, S. Eliezer, and P. Nanda, Phys. Rev. D $\frac{4}{3}$, 3759 (1971).

 ${}^{8}P$. Carruthers and R. W. Haymaker, Phys. Rev. D 4, 1808 (1971).

 $_{\circ}^{\circ}$ R. Olshansky, Phys. Rev. D $\frac{4}{5}$, 2440 (1971).

¹⁰For some of the conjectured alternatives to or modifications of Gell-Mann's hypothesis, see the discussion in S. Okubo, Nuovo Cimento 7A, 765 (1972), and the relevant references cited therein.

 11 For a consideration of this point in the context of a scheme in which the asymptotic fields transform linearly, see N. Papastamatiou, H. Umezawa, and D. J. Welling, Phys. Rev. D 3, 2267 (1971).

 12 L. Bessler, T. Muta, H. Umezawa, and D. Welling, Phys. Rev. D 2, 349 (1970).

Eqs. (22), as to how the asymptotic 0^{\pm} fields might transform to linear order. We have imposed no constraint on the renormalization constants.

It is now manifest how a GW-like theory of broken chiral $SU(3) \times SU(3)$ symmetry may admit either³ chiral $SU(2) \times SU(2)$ or⁴ SU(3) as a weakly broken subsymmetry. It should be clear that any argument^{$7-9$} which appears to support the GOR realization of the limit of no explicit symmetry breaking, but has constrained the renormalization constants to be all equal or all equal to unity, has only obtained a self-consistent result.

It would be interesting to do, if possible, an analysis as in Refs. ⁸ or ⁹ without requiring the renormalization constants to be all equal, or at least to see what happens if all the renormalization constants are proportional to their corresponding masses squared (as in the BP scheme). In any event, it would be useful to see if contact with experimental results can be made without prior commitment to a specific viewpoint (i.e., neither necessarily a GQR theory nor necessarily a BP theory), in order to determine whether or not experiment is able to choose between a GOR world or a BP world in the context of such an impartial theory.

¹³The possibility of "weighted" linear transformations has been studied by J. Rest and D. J. Welling, Phys. Rev. D 3, 2783 (1971); 5, 1983 (1972).

 $14V$. S. Mathur and S. Okubo, Phys. Rev. D 1, 3468 (1970).

¹⁵L. Bessler, Lett. Nuovo Cimento 1, 247 (1971).

- 16 This quadratic equation, but with the ratios of leptonic decay constants determined from spectral-function sum rules, was also used in Ref. 6.
- 17 See, e.g., V. E. Barnes, Purdue University Report No. COO-1428-293, 1971 (unpublished).
- 18 Such expansions, which have been called "dynamical maps," have been used extensively elsewhere. See, e.g., J. Rest, V. Srinivasan, and H. Umezawa, Phys. Rev. D 3, 1890 (1971). For their use in the context of chiral symmetry, see Refs. 12 and 15. For a discussion of the general principle, see R. Hagedorn, Introduction to Field Theory and Dispersion Relations (Pergamon, Oxford, England, 1964), p. 27.

 19 A similar mass formula appears in footnote 8 of Ref. 2.

From Particle Data Group, Rev. Mod. Phys. 43, S1 (1971), we get the masses m_π^2 =0.01885 (average), $m_{\rm K}^2$ = 0.246 (average), m_{η}^2 = 0.301, and $m_{\eta'}^2$ = 0.917, all in units of GeV^2 . For the cases where we use mixing, these masses require $\theta \approx \pm 10.6^{\circ}$.