

Solar and atmospheric neutrino oscillations and lepton flavor violation in supersymmetric models with right-handed neutrinos

J. Hisano

Theory Group, KEK, Oho 1-1, Tsukuba, Ibaraki 305-0801, Japan

Daisuke Nomura

Department of Physics, University of Tokyo, Tokyo 113-0033, Japan

(Received 27 October 1998; published 16 April 1999)

Taking the solar and the atmospheric neutrino experiments into account we discuss the lepton flavor violating processes, such as $\tau \rightarrow \mu \gamma$ or $\mu \rightarrow e \gamma$, in the minimal supersymmetric standard model with right-handed neutrinos (MSSMRN) and the supersymmetric SU(5) grand unified theory with right-handed neutrinos [SU(5)RN]. The predicted branching ratio of $\mu \rightarrow e \gamma$ in the MSSMRN with the Mikheyev-Smirnov-Wolfenstein (MSW) large angle solution is so large that it goes beyond the current experimental bound if the second-generation right-handed Majorana mass M_{ν_2} is greater than $\sim 10^{13} (\sim 10^{14})$ GeV for $\tan \beta = 30(3)$. When we take the MSW small angle solution, the $\mu \rightarrow e \gamma$ rate is at most about 1/100 of that of the MSW large angle solution. The “just so” solution implies 10^{-5} of that of the MSW large angle solution. Also, in the SU(5)RN the large $\mu \rightarrow e \gamma$ rate naturally follows from the MSW large angle solution, and the predicted rate is beyond the current experimental bound if the typical right-handed Majorana mass M_N is larger than $\sim 10^{13} (\sim 10^{14})$ GeV for $\tan \beta = 30(3)$, similarly to the MSSMRN. We show the multimass insertion formulas and their applications to $\tau \rightarrow \mu \gamma$ and $\mu \rightarrow e \gamma$. [S0556-2821(99)03909-0]

PACS number(s): 11.30.Hv, 12.60.Jv, 13.35.-r, 14.60.Pq

I. INTRODUCTION

Introduction of supersymmetry (SUSY) to the standard model (SM) is a solution for the naturalness problem on the radiative correction to the Higgs boson mass. The minimal supersymmetric standard model (MSSM) is considered as one of the most promising models beyond the standard model. Nowadays the signal of supersymmetry is being searched for by many experimental ways.

Lepton flavor conservation, lepton number conservation in each generation, is an exact symmetry in the SM; however, it may be violated in the MSSM [1]. The SUSY breaking mass terms for sleptons have to be introduced phenomenologically. Then, the mass eigenstates for sleptons may be different from those for leptons. This leads to the lepton flavor violating (LFV) rare processes, such as $\mu \rightarrow e \gamma$, $\tau \rightarrow \mu \gamma$, and so on. In fact, the experimental bounds on them have given a constraint on the slepton mass matrices.

The structure of the SUSY breaking mass matrices for sleptons depends on the mechanism to generate the SUSY breaking terms in the MSSM. One of the interesting mechanisms is the minimal supergravity (SUGRA) scenario. Similar to the slepton masses, arbitrary SUSY breaking masses for squarks are also strongly constrained from the flavor changing neutral current (FCNC) processes, such as $K^0 - \bar{K}^0$ mixing. In the minimal SUGRA scenario, the SUSY breaking masses for squarks, sleptons, and the Higgs bosons are expected to be given universally at the tree level, and we can escape from these phenomenological constraints.

However, the universality of the SUSY breaking masses for the scalar bosons is not stable for the radiative correction. Especially, if the physics below the gravitational scale $M_{\text{grav}} (\sim 10^{18} \text{ GeV})$ has the LFV interaction, the interaction in-

duces radiatively the LFV SUSY breaking masses for sleptons. Then, the LFV rare processes are sensitive to physics beyond the MSSM [2].

Recently, the Super-Kamiokande experiment has given us a convincing result [3] that the atmospheric neutrino anomaly [4] comes from the neutrino oscillation. From the zenith-angle dependence of ν_e and ν_μ fluxes following neutrino mass square difference and mixing angle are expected:

$$\Delta m_{\nu_\mu \nu_X}^2 \simeq (10^{-3} - 10^{-2}) \text{ eV}^2, \quad \sin^2 2\theta_{\nu_\mu \nu_X} \gtrsim 0.8. \quad (1)$$

From the negative result for $\nu_e - \nu_\mu$ oscillation in the CHOOZ experiment [5] it is natural to consider $\nu_X = \nu_\tau$ from above results, and the tau neutrino mass is given as

$$m_{\nu_\tau} \simeq (3 \times 10^{-2} - 1 \times 10^{-1}) \text{ eV}, \quad (2)$$

provided mass hierarchy $m_{\nu_\tau} \gg m_{\nu_\mu}$.

The simplest model to generate the neutrino masses is the seesaw mechanism [6]. The neutrino mass Eq. (2) leads to the right-handed neutrino masses below $\sim (10^{14} - 10^{15})$ GeV, even if the Yukawa coupling constant for the tau neutrino mass is of the order of one. This means that a LFV interaction exists below the gravitational scale. Then it is expected that the LFV large mixing for sleptons between the second- and the third-generations is generated radiatively in the minimal SUGRA scenario, and that the LFV rare processes may occur with rates accessible by future experiments [7,8]. In fact, the branching ratio of $\tau \rightarrow \mu \gamma$ in the MSSM with the right-handed neutrinos can reach the present experimental bound [9].

The solar neutrino deficit [10] may also come from neutrino oscillation between $\nu_\mu - \nu_e$. The Mikheyev-Smirnov-Wolfenstein (MSW) solution [11] due to the matter effect in the sun is natural for its explanation, and the observation favors

$$\begin{aligned} \Delta m_{\nu_e \nu_Y}^2 &\simeq (8 \times 10^{-6} - 3 \times 10^{-4}) \text{ eV}^2, \\ \sin^2 2\theta_{\nu_e \nu_Y} &\gtrsim 0.5, \end{aligned} \quad (3)$$

or

$$\begin{aligned} \Delta m_{\nu_e \nu_Y}^2 &\simeq (4 \times 10^{-6} - 1 \times 10^{-5}) \text{ eV}^2, \\ \sin^2 2\theta_{\nu_e \nu_Y} &\simeq (10^{-3} - 10^{-2}). \end{aligned} \quad (4)$$

If the solar neutrino anomaly comes from so-called ‘‘just so’’ solution [12], neutrino oscillation in vacuum, the following mass square difference and mixing angle are expected [13],

$$\begin{aligned} \Delta m_{\nu_e \nu_Y}^2 &\simeq (6 \times 10^{-11} - 1 \times 10^{-10}) \text{ eV}^2, \\ \sin^2 2\theta_{\nu_e \nu_Y} &\gtrsim 0.5. \end{aligned} \quad (5)$$

It is natural to consider $\nu_Y = \nu_\mu$, combined with the atmospheric neutrino observation. If one of the large angle solutions for the solar neutrino anomaly is true, the large mixing $\theta_{\nu_\mu \nu_e}$ may imply the LFV large mixing for sleptons between the first and the second generations.

In this article we investigate the LFV processes in the supersymmetric models with the right-handed neutrinos, assuming the minimal SUGRA scenario. We take the above results for the atmospheric and solar neutrinos. In Sec. II after discussing the origin of the observed mixing angles we calculate the $\tau \rightarrow \mu \gamma$ and $\mu \rightarrow e \gamma$ branching ratios under the assumption of the MSSM with the right-handed neutrinos. It is argued that the $\mu \rightarrow e \gamma$ rate depends on the solar neutrino solutions, and that especially the MSW large angle solution naturally leads to a large $\mu \rightarrow e \gamma$ rate. In Sec. III we consider them in the SU(5) SUSY grand unified theory (GUT) with the right-handed neutrinos. Here also it is shown that the large $\mu \rightarrow e \gamma$ rate naturally follows from the MSW large angle solution. Section IV is for our conclusion. In Appendix A we give our convention used in this article. In Appendices B and C we show the multimass insertion formulas and their applications to $\tau \rightarrow \mu \gamma$ and $\mu \rightarrow e \gamma$, which are useful for estimating the LFV amplitudes and understanding the qualitative behavior of the LFV rates. In Appendix D the renormalization group equations (RGE’s) relevant for our discussion are given.

II. THE LEPTON FLAVOR VIOLATION IN THE MSSM WITH RIGHT-HANDED NEUTRINOS

Before starting to investigate the LFV rare processes in the MSSM with the right-handed neutrinos (MSSMRN), we discuss the origin of the large mixing of neutrino between

$\nu_\tau - \nu_\mu$ or $\nu_\mu - \nu_e$. The MSSMRN is the simplest supersymmetric model to explain the neutrino masses, and following discussion is valid to the extension. The superpotential of the Higgs and lepton sector is given as

$$\begin{aligned} W_{\text{MSSMRN}} &= f_{\nu_{ij}} H_2 \bar{N}_i L_j + f_{e_{ij}} H_1 \bar{E}_i L_j + \frac{1}{2} M_{\nu_i \nu_j} \bar{N}_i \bar{N}_j \\ &+ \mu H_1 H_2, \end{aligned} \quad (6)$$

where L is a chiral superfield for the left-handed lepton, and \bar{N} and \bar{E} are for the right-handed neutrino and the charged lepton. H_1 and H_2 are for the Higgs doublets in the MSSM. Here, i and j are generation indices. After redefinition of the fields, the Yukawa coupling constants and the Majorana masses can be taken as

$$\begin{aligned} f_{\nu_{ij}} &= f_{\nu_i} V_{Dij}, \\ f_{e_{ij}} &= f_{e_i} \delta_{ij}, \\ M_{\nu_i \nu_j} &= U_{ik}^* M_{\nu_k} U_{kj}^\dagger, \end{aligned} \quad (7)$$

where V_D and U are unitary matrices. In this model the mass matrix for the left-handed neutrinos (m_ν) becomes

$$(m_\nu)_{ij} = V_{Dik}^\dagger (\bar{m}_\nu)_{kl} V_{Dlj}, \quad (8)$$

where

$$\begin{aligned} (\bar{m}_\nu)_{ij} &= m_{\nu_i D} [M^{-1}]_{ij} m_{\nu_j D} \\ &\equiv V_{Mik}^\dagger m_{\nu_k} V_{Mkj}. \end{aligned} \quad (9)$$

Here, $m_{\nu_i D} = f_{\nu_i} v \sin \beta / \sqrt{2}$ and V_M is a unitary matrix.¹ We assume $f_{\nu_3} \gtrsim f_{\nu_2} \gtrsim f_{\nu_1}$, similar to the quark sector, and $m_{\nu_\tau} \gg m_{\nu_\mu} \gg m_{\nu_e}$. Also, we take the Yukawa coupling and the Majorana masses for the right-handed neutrinos real for simplicity.

When we consider only the tau and the mu neutrino masses, we parametrize two unitary matrices as

$$V_D = \begin{pmatrix} \cos \theta_D & \sin \theta_D \\ -\sin \theta_D & \cos \theta_D \end{pmatrix}, \quad V_M = \begin{pmatrix} \cos \theta_M & \sin \theta_M \\ -\sin \theta_M & \cos \theta_M \end{pmatrix}. \quad (10)$$

The observed large angle $\theta_{\nu_\mu \nu_\tau}$ is a sum of θ_D and θ_M . However, in order to derive large θ_M we need to fine-tune the independent Yukawa coupling constants and the mass parameters. The neutrino mass matrix (\bar{m}_ν) in the second and the third generations is written explicitly as

¹ $\langle h_1 \rangle = (v \cos \beta / \sqrt{2}, 0)^\dagger$ and $\langle h_2 \rangle = (0, v \sin \beta / \sqrt{2})^\dagger$ with $v \simeq 246 \text{ GeV}$.

$$(\bar{m}_\nu) = \frac{1}{1 - \frac{M_{\nu_2\nu_2}^2 M_{\nu_3\nu_3}^2}{M_{\nu_2\nu_2} M_{\nu_3\nu_3}}} \begin{pmatrix} \frac{m_{\nu_2 D}^2}{M_{\nu_2\nu_2}} & -\frac{m_{\nu_2 D} m_{\nu_3 D}}{M_{\nu_2\nu_3}} \frac{M_{\nu_2\nu_3}^2}{M_{\nu_2\nu_2} M_{\nu_3\nu_3}} \\ -\frac{m_{\nu_2 D} m_{\nu_3 D}}{M_{\nu_2\nu_3}} \frac{M_{\nu_2\nu_3}^2}{M_{\nu_2\nu_2} M_{\nu_3\nu_3}} & \frac{m_{\nu_3 D}^2}{M_{\nu_3\nu_3}} \end{pmatrix}. \quad (11)$$

If the following relations are imposed,

$$\frac{m_{\nu_3 D}^2}{M_{\nu_3\nu_3}} \simeq \frac{m_{\nu_2 D}^2}{M_{\nu_2\nu_2}} \simeq \frac{m_{\nu_2 D} m_{\nu_3 D}}{M_{\nu_2\nu_3}}, \quad (12)$$

the neutrino mass hierarchy $m_{\nu_\tau} \gg m_{\nu_\mu}$ and $\theta_M \simeq \pi/4$ can be derived. However, it is difficult to explain the relation among the independent coupling constants and masses without some mechanism or symmetry. Also, the hierarchy $m_{\nu_3 D} \gg m_{\nu_2 D}$ suppresses the mixing angle θ_M as

$$\tan 2\theta_M \simeq 2 \left(\frac{m_{\nu_2 D}}{m_{\nu_3 D}} \right) \left(\frac{M_{\nu_2\nu_3}}{M_{\nu_2\nu_2}} \right), \quad (13)$$

as far as the Majorana masses for the right-handed neutrinos do not have stringent hierarchical structure as Eq. (12). Therefore, in the following discussion we assume that the large mixing angle between ν_τ and ν_μ comes from θ_D and that U is a unit matrix. Similarly, it is natural to consider that the large mixing angle between ν_μ and ν_e in the MSW solution or the ‘‘just so’’ solution for the solar neutrino anomaly comes from V_D .

The existence of the large mixing angles in V_D may lead to radiative generation of sizable LFV masses for the sleptons in the minimal SUGRA scenario. Though the SUSY breaking masses for the left-handed slepton are flavor-independent at tree level, the Yukawa interaction for the neutrino masses induces radiatively the LFV off-diagonal components in the left-handed slepton mass matrix.

The SUSY breaking terms for the Higgs and lepton sector in the MSSMRN are in general given as

$$\begin{aligned} -\mathcal{L}_{\text{SUSY breaking}} = & (m_{\tilde{L}}^2)_{ij} \tilde{L}_i^\dagger \tilde{L}_j + (m_{\tilde{e}}^2)_{ij} \tilde{e}_{Ri}^* \tilde{e}_{Rj} + (m_{\tilde{\nu}}^2)_{ij} \tilde{\nu}_{Ri}^* \tilde{\nu}_{Rj} \\ & + \tilde{m}_{h_1}^2 h_1^\dagger h_1 + \tilde{m}_{h_2}^2 h_2^\dagger h_2 \\ & + \left(A_{\nu}^{ij} h_2 \tilde{\nu}_{Ri}^* \tilde{L}_j + A_e^{ij} h_1 \tilde{e}_{Ri}^* \tilde{L}_j \right. \\ & \left. + \frac{1}{2} B_{\nu}^{ij} \tilde{\nu}_{Ri}^* \tilde{\nu}_{Rj} + B_h h_1 h_2 + \text{H.c.} \right), \quad (14) \end{aligned}$$

where \tilde{L} , \tilde{e}_R , and $\tilde{\nu}_R$ represent the left-handed slepton, and the right-handed charged slepton, and the right-handed neutrino. Also, h_1 and h_2 are the doublet Higgs bosons. In the minimal SUGRA scenario at the gravitational scale the SUSY breaking masses for sleptons, squarks, and the Higgs bosons are universal, and the SUSY breaking parameters associated with the supersymmetric Yukawa couplings or

masses (A or B parameters) are proportional to the Yukawa coupling constants or masses. Then, the SUSY breaking parameters in Eq. (14) are given as

$$(m_{\tilde{L}}^2)_{ij} = (m_{\tilde{e}}^2)_{ij} = (m_{\tilde{\nu}}^2)_{ij} = \delta_{ij} m_0^2,$$

$$\tilde{m}_{h_1}^2 = \tilde{m}_{h_2}^2 = m_0^2,$$

$$A_{\nu}^{ij} = f_{\nu_{ij}} a_0, \quad A_e^{ij} = f_{e_{ij}} a_0,$$

$$B_{\nu}^{ij} = M_{\nu_{ij}} b_0, \quad B_h = \mu b_0. \quad (15)$$

In order to know the values of the SUSY breaking parameters at the low energy, we have to include the radiative corrections to them. We can evaluate them by the RGE's. We present them in Appendix D, and here we discuss only the qualitative behavior of the solution using the logarithmic approximation. The SUSY breaking masses of squarks, sleptons, and the Higgs bosons at the low energy are enhanced by gauge interactions, and the corrections are flavor-independent and proportional to square of the gaugino masses. On the other hand, Yukawa interactions reduce the SUSY breaking masses. If the Yukawa coupling is LFV, the radiative correction to the SUSY breaking parameters is LFV. The LFV off-diagonal components for $(m_{\tilde{L}}^2)$, $(m_{\tilde{e}}^2)$, and A_e^{ij} are given at the low energy as

$$(m_{\tilde{L}}^2)_{ij} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{Dki}^* V_{Dkj} f_{\nu_k}^2 \log \frac{M_{\text{grav}}}{M_{\nu_k}},$$

$$(m_{\tilde{e}}^2)_{ij} \simeq 0,$$

$$A_e^{ij} \simeq -\frac{3}{8\pi^2} a_0 f_{e_{ij}} V_{Dki}^* V_{Dkj} f_{\nu_k}^2 \log \frac{M_{\text{grav}}}{M_{\nu_k}},$$

where $i \neq j$. In these equations, the off-diagonal components of $(m_{\tilde{L}}^2)$ and A_e are generated radiatively while those of $(m_{\tilde{e}}^2)$ are not. This is because the right-handed leptons have only one kind of the Yukawa interaction f_e and we can always

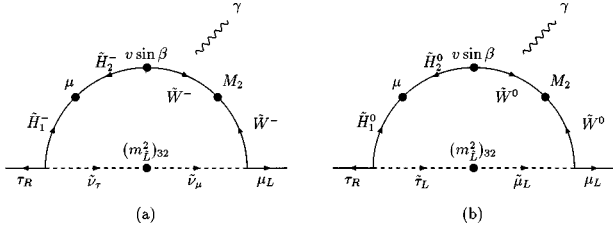


FIG. 1. The Feynman diagrams which give dominant contributions to $\tau^+ \rightarrow \mu^+ \gamma$ when $\tan \beta \geq 1$ and the off-diagonal elements of the right-handed slepton mass matrix are negligible, as in the MSSM with the right-handed neutrinos. In the diagrams, $(m_{\tilde{L}}^2)_{32}$ is the (3, 2) element of the left-handed slepton soft mass matrix. $\tilde{\tau}_{L(R)}$ and $\tilde{\mu}_{L(R)}$ are the left-handed (right-handed) stau and smuon, respectively, and $\tilde{\nu}_\tau$ and $\tilde{\nu}_\mu$ the tau sneutrino and the mu sneutrino. \tilde{H}_1 and \tilde{H}_2 are Higgsino, \tilde{W} W -ino. The symbol μ is the Higgsino mass. The arrows represent the chirality.

take a basis where f_e is diagonal. The magnitudes of the off-diagonal components of $(m_{\tilde{L}}^2)$ and A_e are sensitive to f_{ν_i} and V_{D^2} .

As shown above, V_{D32} is expected to be of the order of one from the atmospheric neutrino observation. This leads to the nonvanishing $(m_{\tilde{L}}^2)_{32}$ and A_e^{32} , which result in a finite $\tau \rightarrow \mu \gamma$ decay rate via diagrams involving them. The dominant contributions are proportional to

$$(m_{\tilde{L}}^2)_{32} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{D33}^* V_{D32} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}}. \quad (17)$$

As will be shown, if f_{ν_3} is of the order of one, the branching ratio of $\tau \rightarrow \mu \gamma$ may reach the present experimental bound.

Moreover, if V_{D31} is finite, $(m_{\tilde{L}}^2)_{31}$ and $(m_{\tilde{L}}^2)_{21}$ are also large. They are approximately

$$(m_{\tilde{L}}^2)_{31} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{D33}^* V_{D31} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}},$$

$$(m_{\tilde{L}}^2)_{21} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{D32}^* V_{D31} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}}. \quad (18)$$

This fact implies a sizable $\mu \rightarrow e \gamma$ rate because the amplitudes proportional to $(m_{\tilde{L}}^2)_{23}(m_{\tilde{L}}^2)_{31}$ or $(m_{\tilde{L}}^2)_{21}$ are dominant.

²If U is not a unit matrix, the off-diagonal components for $(m_{\tilde{L}}^2)$ and A_e become

$$(m_{\tilde{L}}^2)_{ij} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{Dki}^* V_{Dlj} f_{\nu_k} f_{\nu_l} U_{km}^* U_{lm} \log \frac{M_{\text{grav}}}{M_{\nu_m}},$$

$$A_e^{ij} \simeq -\frac{3}{8\pi^2} a_0 f_{e_i} V_{Dki}^* V_{Dlj} f_{\nu_k} f_{\nu_l} U_{km}^* U_{lm} \log \frac{M_{\text{grav}}}{M_{\nu_m}}. \quad (16)$$

Then they are insensitive to the detail of U since the dependence on M_{ν_i} is logarithmic.

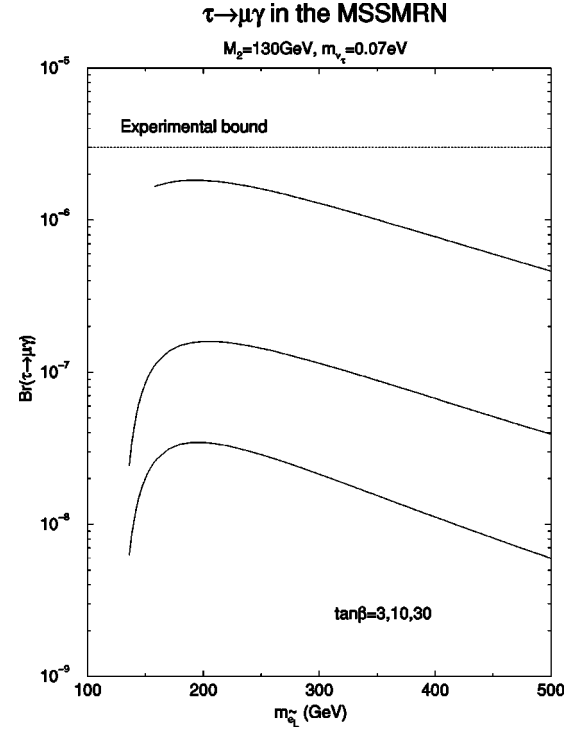


FIG. 2. Dependence of the branching ratio of $\tau \rightarrow \mu \gamma$ on the left-handed selectron mass $m_{\tilde{e}_L}$ in the MSSM with the right-handed neutrinos. We take the tau neutrino mass 0.07 eV and $V_{D32} = -0.71$, which are suggested by the atmospheric neutrino result. M_{ν_3} is fixed at $\approx 10^{14}$ GeV by imposing a condition $f_{\nu_3} = f_{\nu_3}$ at the gravitational scale. The dotted line shown in the figure is the present experimental bound. We set the W -ino mass $M_2 = 130$ GeV, and the Higgsino mass parameter μ positive. The mu neutrino mass is neglected. We take $\tan \beta = 3, 10, 30$. The larger $\tan \beta$ corresponds to the upper curve.

When V_{D21} is also of the order of one to explain the solar neutrino anomaly, an extra contribution to $(m_{\tilde{L}}^2)_{21}$ has to be taken into account as

$$(m_{\tilde{L}}^2)_{21} \simeq -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) \left(V_{D32}^* V_{D31} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}} + V_{D22}^* V_{D21} f_{\nu_2}^2 \log \frac{M_{\text{grav}}}{M_{\nu_2}} \right). \quad (19)$$

The experimental upper bound on the branching ratio of $\mu \rightarrow e \gamma$ is so severe that the predicted branching ratio may reach it even if f_{ν_2} is $\mathcal{O}(10^{-1})$.

A. The branching ratio of $\tau \rightarrow \mu \gamma$

Let us discuss the branching ratios of the LFV rare processes in the MSSMRN. First, $\tau \rightarrow \mu \gamma$. The amplitude of the $e_i^+ \rightarrow e_j^+ \gamma$ ($i > j$) takes a form

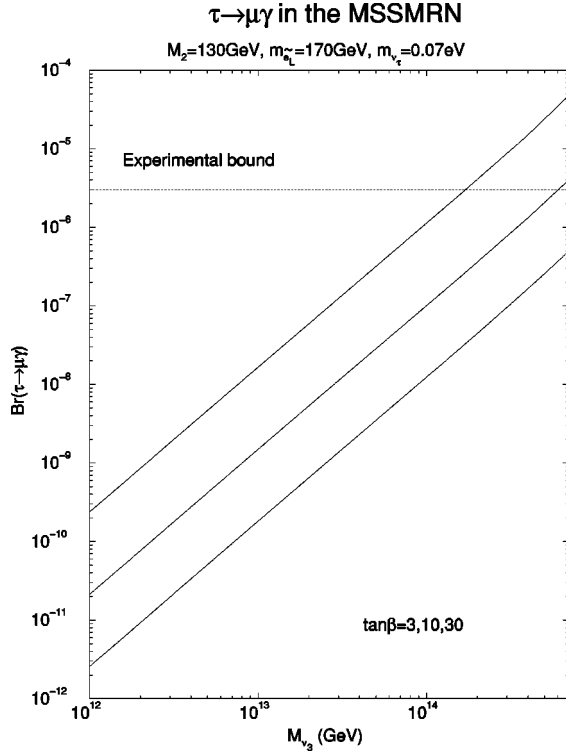


FIG. 3. Dependence of the branching ratio of $\tau \rightarrow \mu \gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} in the MSSM with the right-handed neutrinos. The input parameters are the same as those of Fig. 2 except that in this figure we take $m_{\tilde{e}_L} = 170$ GeV and that we do not impose the condition $f_{u_3} = f_{\nu_3}$ but treat M_{ν_3} as an independent variable. The dotted line shown in the figure is the present experimental bound. Here also the larger $\tan \beta$ corresponds to the upper curve.

$$T = e \epsilon^{\alpha*}(q) \bar{\nu}_i(p) i \sigma_{\alpha\beta} q^\beta (A_L^{(ij)} P_L + A_R^{(ij)} P_R) v_j(p - q), \quad (20)$$

where p and q are momenta of e_i and photon, and the decay rate is given by

$$\Gamma(e_i \rightarrow e_j \gamma) = \frac{e^2}{16\pi} m_{e_i}^3 (|A_L^{(ij)}|^2 + |A_R^{(ij)}|^2). \quad (21)$$

Here, we neglect the mass of e_j . The amplitude is not invariant for the $SU(2)_L$ and $U(1)_Y$ symmetry and the chiral symmetry of leptons. Then the coefficients $A_L^{(ij)}$ and $A_R^{(ij)}$ are proportional to the charged lepton masses. Since in the MSSMRN the mismatch between the left-handed slepton and the charged slepton mass eigenstates is induced, $A_L^{(ij)}$ is much larger than $A_R^{(ij)}$ since $A_R^{(ij)}$ is suppressed by m_{e_j}/m_{e_i} compared with $A_L^{(ij)}$. Also, when $\tan \beta (\equiv v_2/v_1)$ is large, the contribution to $A_L^{(ij)}$ proportional to $f_{e_i} v_2 (= -\sqrt{2} m_{e_i} \tan \beta)$ becomes dominant. In the MSSMRN, the dominant contribution to $\tau \rightarrow \mu \gamma$ is from the diagram of Figs. 1(a) and 1(b) and its expression is

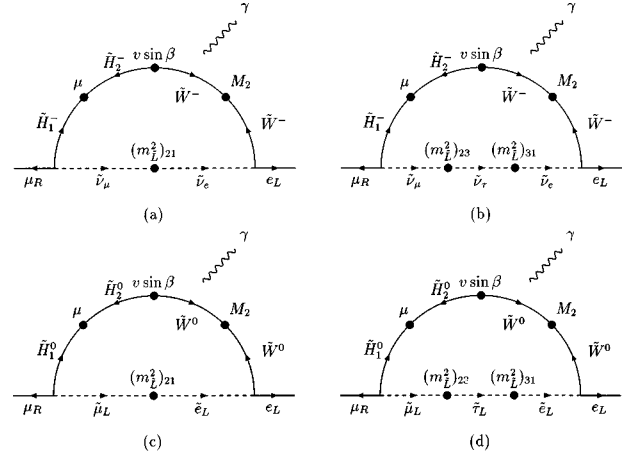


FIG. 4. Candidates of the Feynman diagrams which give dominant contributions to $\mu^+ \rightarrow e^+ \gamma$ when $\tan \beta \gtrsim 1$ and the off-diagonal elements of the right-handed soft mass matrix are negligible, as in the MSSM with the right-handed neutrinos. In the diagrams, $(m_L^2)_{ij}$ is the (i,j) element of the left-handed slepton mass matrix. $\tilde{\nu}_{L(R)}$ is the left-handed (right-handed) selectron and $\tilde{\nu}_e$ is the electron sneutrino. Other symbols are the same as those in Fig. 1.

$$A_L^{(\tau\mu)} \simeq m \frac{\alpha_2}{4\pi} \mu M_2 \tan \beta (m_L^2)_{32} \times D \left[D \left[\frac{1}{m^2} \left\{ f_{c2} \left(\frac{M^2}{m^2} \right) - \frac{1}{4} f_{n2} \left(\frac{M^2}{m^2} \right) \right\}; M^2 \right] \times (M_2^2, \mu^2); m^2 \right] (m_{\tilde{\nu}_\mu}^2, m_{\tilde{\nu}_\tau}^2), \quad (22)$$

which comes from the $SU(2)_L$ interaction. The functions $f_{n2}(x)$ and $f_{c2}(x)$ are defined in Appendix C and the operator $D[f(x); x](x_1, x_2)$ to a function $f(x)$ is defined by

$$D[f(x); x](x_1, x_2) \equiv \frac{1}{x_1 - x_2} (f(x_1) - f(x_2)). \quad (23)$$

Here, for a demonstrational purpose, we take a limit where the SUSY breaking scale is much larger than the W and Z gauge boson masses and $\tan \beta \gtrsim 1$. This equation can be derived from the mass-insertion formula represented in Appendix C. The LFV A term cannot give a dominant contribution when $\tan \beta \gtrsim 1$.

In Fig. 2 we show the branching ratio of $\tau \rightarrow \mu \gamma$ as a function of the left-handed selectron mass ($m_{\tilde{e}_L}$). Here, $m_{\nu_\tau} = 0.07$ eV, $V_{D33} = V_{D22} = -V_{D32} = V_{D23} = 1/\sqrt{2}$, and we assume that f_{ν_3} is as large as the Yukawa coupling constant for the top quark at the gravitational scale. This corresponds to $M_{\nu_3} \sim 10^{14}$ GeV. Also, we impose the radiative breaking condition of the $SU(2)_L \times U(1)_Y$ gauge symmetry with $\tan \beta = 3, 10, 30$ and the Higgsino mass parameter μ positive. In our calculation we considered the experimental constraints from the negative results of the SUSY particle search.

Though we do not assume the GUT's, we take the W -ino mass (M_2) 130 GeV and determine the other gaugino masses by the GUT relation for the gaugino masses. We use the formula for $\tau \rightarrow \mu \gamma$ in Ref. [9] for the numerical calculation.

The branching ratio is reduced where the left-handed selectron mass is comparable to the W -ino mass. This is because the slepton masses are almost determined by the radiative correction from the gaugino masses, and m_0^2 , which $(m_L^2)_{32}$ is proportional to, is negligible in the region. As mentioned above, the branching ratio is proportional to $\tan^2 \beta$ [see Eq. (22)], and the line for $\tan \beta=30$ is close to the experimental bound, $\text{Br}(\tau \rightarrow \mu \gamma) < 3.0 \times 10^{-6}$ [14].

In Fig. 3 we present the dependence of the branching ratio of $\tau \rightarrow \mu \gamma$ on M_{ν_3} . Here, we take $m_{\tilde{e}_L} = 170$ GeV, and the other SUSY breaking parameters are the same as in Fig. 2. The branching ratio is proportional to $M_{\nu_3}^2$ since we fix $m_{\nu_\tau} = 0.07$ eV. If 10^{-8} can be reached in the future experiments, we can probe $M_{\nu_3} > 10^{13}(10^{14})$ GeV for $\tan \beta = 30(3)$.³

B. The branching ratio of $\mu \rightarrow e \gamma$

Next, we discuss $\mu \rightarrow e \gamma$ in the MSSMRN. The forms of the amplitude and the event rate are the same as those of $\tau \rightarrow \mu \gamma$ [Eqs. (20), (21)]. This process has two types of the contribution, depending on the structure of the Yukawa coupling for the neutrino masses. One is the diagrams where $(m_L^2)_{21}$ or A_e^{21} is inserted, and another is those that $(m_L^2)_{32}$ or A_e^{32} and $(m_L^2)_{13}$ or A_e^{13} are inserted. Then the dominant contributions [Figs. 4(a)–4(d)] are the following:

$$\begin{aligned}
 A_L^{(\mu e)} &= -m_{\mu 4} \frac{\alpha_2}{\pi} M_2 \mu \tan \beta \\
 &\times D \left[\left[(m_L^2)_{21} + \frac{(m_L^2)_{23}(m_L^2)_{31}}{m_{\nu_\tau}^2 - m_{\nu_\tau}^2} \right] \right. \\
 &\times \frac{1}{m_{\nu_\tau}^2} \left\{ g_{c2} \left(\frac{M^2}{m_{\nu_\tau}^2} \right) - \frac{1}{4} g_{n2} \left(\frac{M^2}{m_{\nu_\tau}^2} \right) \right\} \\
 &- \left[\frac{(m_L^2)_{23}(m_L^2)_{31}}{(m_{\nu_\tau}^2 - m_{\nu_\tau}^2)^2} \right] \frac{1}{m_{\nu_\tau}^2} \left\{ f_{c2} \left(\frac{M^2}{m_{\nu_\tau}^2} \right) \right. \\
 &\left. \left. - \frac{1}{4} f_{n2} \left(\frac{M^2}{m_{\nu_\tau}^2} \right) \right\} \right]; M^2 \left(M_2^2, \mu^2 \right). \quad (24)
 \end{aligned}$$

Here, we take a limit where the SUSY breaking scale is much larger than the W and Z gauge boson masses and

³An alternative way to prove $(m_L^2)_{32}$ is to search for the slepton oscillation [15,16].

$\mu \rightarrow e \gamma$ in the MSSMRN with the MSW large angle solution

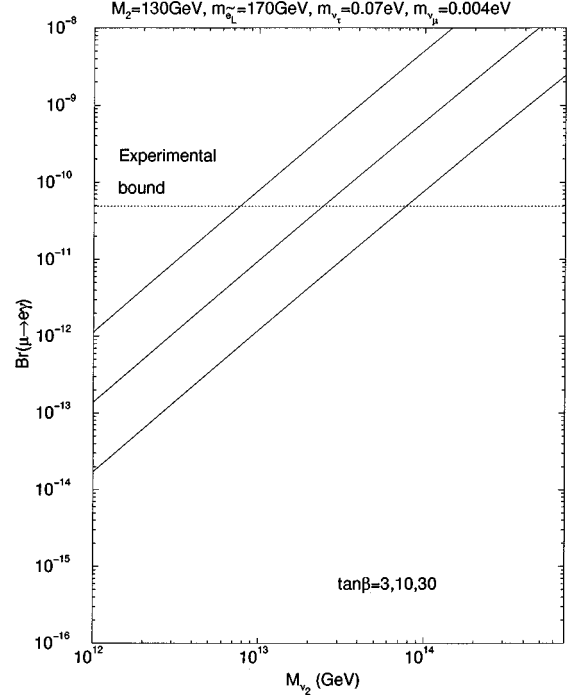


FIG. 5. Dependence of the branching ratio of $\mu \rightarrow e \gamma$ on the second-generation right-handed neutrino Majorana mass M_{ν_2} in the MSSM with the right-handed neutrinos under the assumption of the MSW large angle solution with $V_{D31} = 0$. We take $V_{D21} = -0.42$ and the mu neutrino mass as 0.004 eV, as suggested by the MSW large angle solution. The dotted line shown in the figure is the present experimental bound. Other input parameters are the W -ino mass 130 GeV, the left-handed selectron 170 GeV, the tau neutrino mass 0.07 eV, and $\tan \beta = 3, 10$, and 30. The larger $\tan \beta$ corresponds to the upper curve.

$\tan \beta \gtrsim 1$, again. We also assumed the mass degeneracy between the first- and the second-generation left-handed sleptons as

$$m_{\tilde{e}_L}^2 = m_{\tilde{\mu}_L}^2 = m_{\tilde{\nu}_\tau}^2 = m_{\tilde{\nu}_\mu}^2 \equiv m_{\tilde{\nu}}^2. \quad (25)$$

The functions $f_{c2,n2}(x)$ and $g_{c2,n2}(x)$ are defined in Appendix C.

As mentioned above, if the solar neutrino anomaly comes from the MSW effect or the vacuum oscillation with the large angle, V_{D21} is expected to be large. This leads to non-vanishing $(m_L^2)_{21}$. In Fig. 5, under the condition that

$$V_D = \begin{pmatrix} 0.91 & 0.35 & 0.24 \\ -0.42 & 0.72 & 0.55 \\ 0 & -0.60 & 0.80 \end{pmatrix} \quad (26)$$

and $m_{\nu_\mu} = 0.004$ eV [17] we show the branching ratio of $\mu \rightarrow e \gamma$ as a function of M_{ν_2} . This corresponds to the MSW solution with the large mixing. Here we take $V_{D31} = 0$ and we will discuss a case with finite V_{D31} later. The input parameters are taken to be the same as in Fig. 3. For $\tan \beta = 30(3)$, the branching ratio reaches the experimental bound [$\text{Br}(\mu \rightarrow e \gamma) < 4.9 \times 10^{-11}$ [14]] when $M_{\nu_2} \simeq 8 \times 10^{12}(8$

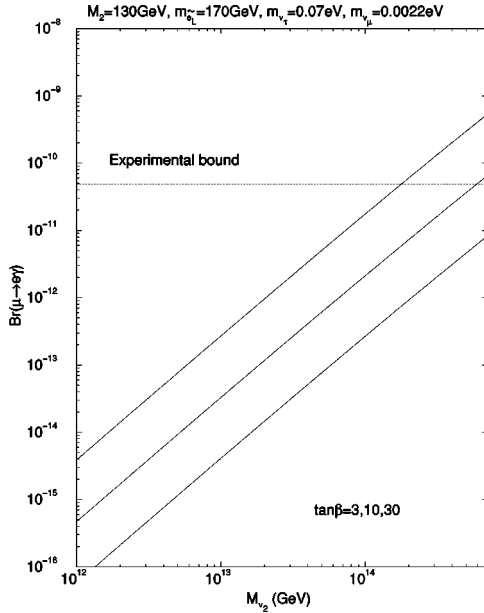
$\mu \rightarrow e\gamma$ in the MSSMRN with the MSW small angle solution

FIG. 6. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the second-generation right-handed neutrino Majorana mass M_{ν_2} in the MSSM with the right-handed neutrinos under the assumption of the MSW small angle solution with $V_{D31}=0$. We take $V_{D21}=-0.04$ and the mu neutrino mass as 0.0022 eV, which are suggested by the MSW small angle solution if the mixing comes from V_D . Other input parameters are the same as those in Fig. 5. The dotted line shown in the figure is the present experimental bound. $\tan \beta=3, 10$, and 30, and the larger $\tan \beta$ corresponds to the upper curve.

$\times 10^{13}$ GeV. This corresponds to $f_{\nu_2} \approx 0.03(0.11)$. Future experiments are expected to reach 10^{-14} [18]. This corresponds to $M_{\nu_2} \approx 10^{11}(10^{12})$ GeV.

If the solar neutrino anomaly comes from the MSW solution with the small mixing, we cannot distinguish whether the mixing comes from V_D or V_M . If it comes from V_D , the branching ratio is smaller by about 1/100 compared with that in the MSW solution with the large mixing, as shown in Fig. 6. In Fig. 6 we assume that

$$V_D = \begin{pmatrix} 1 & 0.04 & 0.03 \\ -0.04 & 0.79 & 0.59 \\ 0 & -0.60 & 0.80 \end{pmatrix} \quad (27)$$

and $m_{\nu_\mu} = 0.0022$ eV [17]. Other input parameters are the same as Fig. 5.

In Fig. 7 we take

$$V_D = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \\ -\frac{1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \\ 0 & -\frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad (28)$$

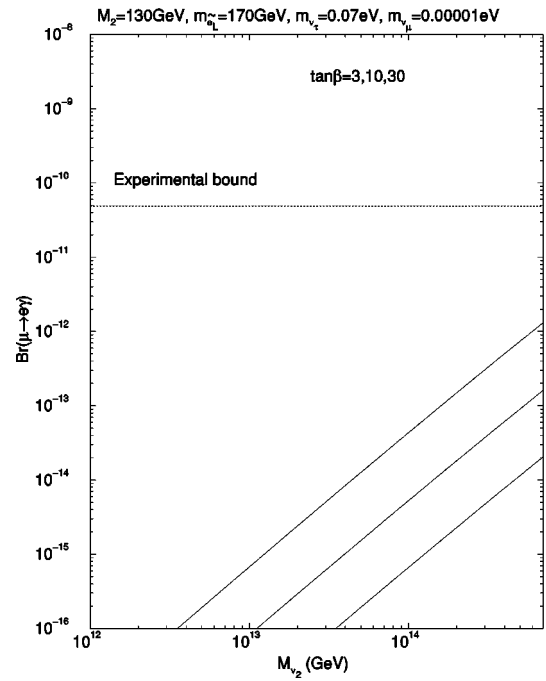
 $\mu \rightarrow e\gamma$ in the MSSMRN with the ‘just so’ solution

FIG. 7. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the second-generation right-handed neutrino Majorana mass M_{ν_2} in the MSSM with the right-handed neutrinos under the assumption of the ‘just so’ solution with $V_{D31}=0$. We take $V_{D21}=-0.71$ and the mu neutrino mass as 1.0×10^{-5} eV, as suggested by the ‘just so’ solution. Other input parameters are the same as those in Fig. 5. The dotted line shown in the figure is the present experimental bound. $\tan \beta=3, 10$, and 30, and the larger $\tan \beta$ corresponds to the upper curve.

and $m_{\nu_\mu} = 1.0 \times 10^{-5}$ eV [19]. Other parameters are the same as in Fig. 5. This corresponds to the ‘just so’ solution for the solar neutrino anomaly. Since the mu neutrino mass is smaller, the branching ratio is suppressed by 10^{-5} compared with that in the MSW solution with the large mixing.

Next we discuss the branching ratio of $\mu \rightarrow e\gamma$ when V_{D31} is finite. In Figs. 8 and 9 we show the branching ratio as a function of V_{D31} and M_{ν_3} for $\tan \beta=3$ and 30. Here we assume that f_{ν_2} is negligibly small. The other parameters are the same as in Fig. 3. The branching ratio is almost proportional to $V_{D31}^2 M_{\nu_3}^2$. Compared with this figure to Fig. 5, when $M_{\nu_2} = M_{\nu_3}$, the contribution from V_{D31} is negligible in the MSW solution with the large mixing angle unless V_{D31} is larger than $10^{-2.5}$. On the other hand, it can be dominant in the ‘just so’ solution even if $V_{D31} \sim 10^{-4}$.

Finally we consider the $\mu^+ \rightarrow e^+ e^- e^+$ process and the μ - e conversion on ${}^{48}\text{Ti}$. For these processes the penguin type diagrams dominate over the others, so the behavior of the decay rate is similar to that of $\mu \rightarrow e\gamma$. For the $\mu \rightarrow 3e$ process the following approximate relation holds between the branching ratios of the two processes,

$$\text{Br}(\mu \rightarrow 3e) \approx \frac{\alpha}{8\pi} \frac{8}{3} \left(\log \frac{m_\mu^2}{m_e^2} - \frac{11}{4} \right) \text{Br}(\mu \rightarrow e\gamma) \quad (29)$$

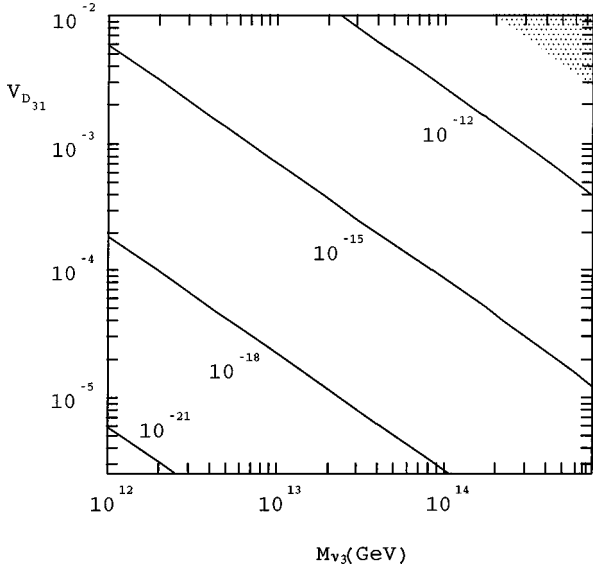


FIG. 8. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} and $V_{D_{31}}$ in the MSSM with the right-handed neutrinos. Here the tau neutrino mass is 0.07 eV and $V_{D_{32}} = -0.71$, as suggested by the atmospheric neutrino result. We neglect f_{ν_2} here. The curves mean the contours on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-21} , 10^{-18} , 10^{-15} , and 10^{-12} , respectively. The shaded region is already excluded experimentally. $\tan \beta$ is set to be 3. The W -ino mass is 130 GeV, the left-handed selectron mass 170 GeV, and the Higgsino mass parameter μ positive.

$$\simeq 7 \times 10^{-3} \text{Br}(\mu \rightarrow e\gamma). \quad (30)$$

For the μ - e conversion rate $\Gamma(\mu \rightarrow e)$ a similar relation holds at $\tan \beta > 1$ region,

$$\Gamma(\mu \rightarrow e) \simeq 16\alpha^4 Z_{\text{eff}}^4 |F(q^2)|^2 \text{Br}(\mu \rightarrow e\gamma). \quad (31)$$

Here Z is the proton number in the nucleus, and Z_{eff} is the effective charge, $F(q^2)$ the nuclear form factor at the momentum transfer q . For ${}^{48}_{22}\text{Ti}$, $Z_{\text{eff}} = 17.6$ and $F(q^2 \simeq -m_\mu^2) \simeq 0.54$ [20,21]. We express the magnitude of the μ - e conversion with the normalization the muon capture rate in Ti nucleus. Then the normalized conversion rate $R(\mu^- \rightarrow e^-; {}^{48}_{22}\text{Ti})$ is approximately

$$R(\mu^- \rightarrow e^-; {}^{48}_{22}\text{Ti}) \simeq 6 \times 10^{-3} \text{Br}(\mu \rightarrow e\gamma). \quad (32)$$

The future experiment for the μ - e conversion is planned to reach $R(\mu^- \rightarrow e^-; {}^{48}_{22}\text{Ti}) < 10^{-18}$ [22].

III. THE LEPTON FLAVOR VIOLATION IN THE SU(5) SUSY GUT WITH RIGHT-HANDED NEUTRINOS

In the SUSY GUT the gauge coupling unification is predicted, and the predicted weak mixing angle is consistent with the experimental data at the 1% level of accuracy. Moreover if the unified gauge group is SO(10), the right-handed neutrinos are introduced automatically into the matter multiplet. However, in order to accommodate the observed large mixing angle in the framework of the SO(10)

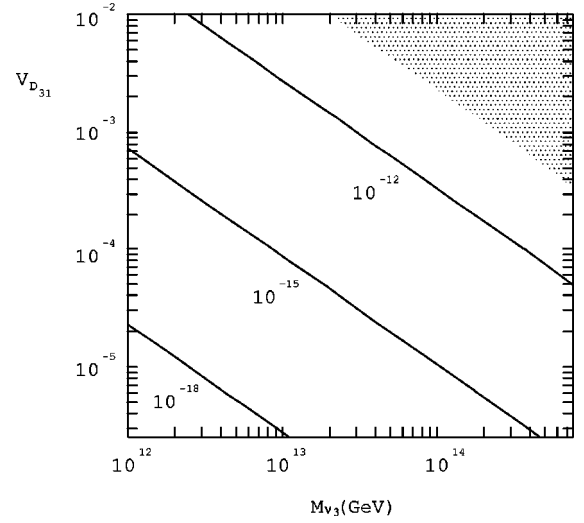


FIG. 9. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} and $V_{D_{31}}$ in the MSSM with the right-handed neutrinos. The input parameters are the same as those in Fig. 8 except that we take $\tan \beta = 30$ here. The curves mean the contours on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-18} , 10^{-15} , and 10^{-12} , respectively. The shaded region is already excluded experimentally.

SUSY GUT, one needs unnatural extension of the simplest version of the SO(10) SUSY GUT. Hence in this article we do not discuss the SO(10) SUSY GUT. Here we investigate the SU(5) SUSY GUT with the right-handed neutrinos as one of the extension of the MSSMRN in which the small neutrino mass is naturally obtained and the large neutrino mixing angle is possible without unnatural fine-tuning. We here call this model as SU(5)RN, for brevity. After introducing the model we estimate the off-diagonal elements of the slepton soft mass matrices using the one-loop level renormalization group equations (RGE's) under an assumption of the minimal SUGRA scenario. With them we study the LFV processes $\tau \rightarrow \mu\gamma$ and $\mu \rightarrow e\gamma$. After that we comment on the $b \rightarrow s\gamma$ branching ratio. We show that the LFV rates in this model is larger in general than those in the MSSMRN model, due to the fact that in this model the right-handed slepton mass matrix also can have non-negligible off-diagonal elements, in addition to the left-handed one [23,24,25].

First we introduce the model. This model has three families of matter multiplets ψ_i , ϕ_i , and η_i , which are **10**, **5***, and **1** dimension representations of SU(5), respectively. ψ_i contains the quark doublet, the charged lepton singlet, and the up-type quark singlet, while ϕ_i the down-type quark singlet and the lepton doublet and η_i the right-handed neutrino, respectively. The model has **5** and **5*** dimension representation Higgs multiplets, H and \bar{H} . H consists of the MSSM Higgs multiplet H_2 and a colored Higgs multiplet H_C , and \bar{H} another MSSM Higgs multiplet H_1 and another colored Higgs multiplet \bar{H}_C . The GUT gauge symmetry is spontaneously broken into the SM one at the GUT scale $M_{\text{GUT}} \simeq 2 \times 10^{16}$ GeV. Above the GUT scale the superpotential W of the matter sector of this model is

$$W = \frac{1}{4} f_{u_{ij}} \psi_i^{AB} \psi_j^{CD} H^E \epsilon_{ABCDE} + \sqrt{2} f_{d_{ij}} \psi_i^{AB} \phi_{jA} \bar{H}_B \\ + f_{\nu_{ij}} \eta_i \phi_{jA} H^A + \frac{1}{2} M_{\eta_i \eta_j} \eta_i \eta_j,$$

where A, B, \dots are indices of SU(5) and run from 1 to 5. We also introduce the soft SUSY breaking terms associated with the GUT multiplets. The relevant part of them is⁴

$$-\mathcal{L}_{\text{SUSY breaking}} = (m_\psi^2)_{ij} \tilde{\psi}_i^\dagger \tilde{\psi}_j + (m_\phi^2)_{ij} \tilde{\phi}_i^\dagger \tilde{\phi}_j + (m_\eta^2)_{ij} \tilde{\eta}_i^\dagger \tilde{\eta}_j \\ + m_h^2 h^\dagger h + m_{\bar{h}}^2 \bar{h}^\dagger \bar{h} + \left\{ \frac{1}{4} A_{u_{ij}} \tilde{\psi}_i \tilde{\psi}_j h \right. \\ \left. + \sqrt{2} A_{d_{ij}} \tilde{\psi}_i \tilde{\phi}_j \bar{h} + A_{\nu_{ij}} \tilde{\eta}_i \tilde{\phi}_j h + \text{H.c.} \right\}, \quad (33)$$

where $\tilde{\psi}_i$, $\tilde{\phi}_i$, and $\tilde{\eta}_i$ are the scalar components of the ψ_i , ϕ_i , and η_i chiral multiplets, respectively, and h and \bar{h} are the Higgs bosons. In the minimal SUGRA scenario these coefficients are given at the gravitational scale as

$$(m_\psi^2)_{ij} = (m_\phi^2)_{ij} = (m_\eta^2)_{ij} = \delta_{ij} m_0^2, \\ m_h^2 = m_{\bar{h}}^2 = m_0^2, \\ A_{u_{ij}} = f_{u_{ij}} a_0, \quad A_{d_{ij}} = f_{d_{ij}} a_0, \quad A_{\nu_{ij}} = f_{\nu_{ij}} a_0. \quad (34)$$

At the GUT scale we choose a basis where the up-type quark and the neutrino Yukawa coupling matrices are diagonalized as

$$f_{u_{ij}} = f_{u_i} e^{i\phi_{u_i}} \delta_{ij}, \\ f_{d_{ij}} = (V_{\text{KM}}^*)_{ik} f_{d_k} (V_D^\dagger)_{kj}, \\ f_{\nu_{ij}} = f_{\nu_i} e^{i\phi_{\nu_i}} \delta_{ij}, \quad (35)$$

where f_{ψ_i} ($\psi = u, d, \nu$) are the eigenvalues of $f_{\psi_{ij}}$, respectively, V_{KM} the Kobayashi-Maskawa matrix at the GUT scale, and V_D a unitary matrix which describes the generation mixing in the lepton sector. ϕ_{ψ_i} ($\psi = u, \nu$) are phase factors which satisfy $\phi_{u_1} + \phi_{u_2} + \phi_{u_3} = 0$ and $\phi_{\nu_1} + \phi_{\nu_2} + \phi_{\nu_3} = 0$. However these phases are completely irrelevant for our discussion below.

At the GUT scale the Yukawa coupling constants responsible to the down-type quark masses and those responsible to the charged lepton masses are supposed to unify as

$$f_{d_i} = f_{e_i}. \quad (36)$$

⁴For simplicity we neglect the Yukawa coupling $\lambda H \Sigma \bar{H}$ and the soft SUSY-breaking parameters associated with it, where Σ is an adjoint representation Higgs multiplet causing the breaking $\text{SU}(5)_{\text{GUT}} \rightarrow \text{SU}(3)_c \times \text{SU}(2)_L \times \text{U}(1)_Y$.

This relation is consistent with the particle spectrum at the low energy only for the third-generation. In order to explain the fermion masses of the first- and the second-generations, one has to consider the effect of the nonrenormalizable terms also. At that time those terms can be another source of LFV [26,27], but we do not take them into account for simplicity.

Below the GUT scale we take the basis in which the Yukawa coupling constant matrix responsible for charged lepton masses is diagonalized. The basis we take at low energy region is related to that of GUT multiplets by the following embedding:

$$\psi_i = \{Q_i, e^{-i\phi_{u_i}} \bar{U}_i, (V_{\text{KM}})_{ij} \bar{E}_j\}, \\ \phi_i = \{V_{Dij} \bar{D}_j, V_{Dij} L_j\}, \\ \eta_i = \{e^{-i\phi_{\nu_i}} \bar{N}_i\}. \quad (37)$$

Then the superpotential W is expanded in terms of the MSSM fields as

$$W = f_{u_i} Q_i \bar{U}_i H_2 + (V_{\text{KM}}^*)_{ij} f_{d_j} Q_i \bar{D}_j H_1 + f_{d_i} \bar{E}_i L_i H_1 \\ - f_{\nu_i} V_{Dij} \bar{N}_i L_j H_2 + f_{u_j} (V_{\text{KM}})_{ji} \bar{E}_i \bar{U}_j H_C \\ - \frac{1}{2} f_{u_i} e^{i\phi_{u_i}} Q_i Q_i H_C + (V_{\text{KM}}^*)_{ij} f_{d_j} e^{-i\phi_{u_i}} \bar{U}_i \bar{D}_j \bar{H}_C \\ - (V_{\text{KM}}^*)_{ij} f_{d_j} Q_i L_j \bar{H}_C + f_{\nu_i} V_{Dij} \bar{N}_i \bar{D}_j H_C + \frac{1}{2} M_{\nu_i \nu_j} \bar{N}_i \bar{N}_j. \quad (38)$$

Here we should notice that the fifth term of the right-hand side of the above equation is no longer generation-diagonal. This is nothing but a direct consequence of the GUT unification, that is, one of the central goals of the grand unification is to embed the leptons and the quarks into the same multiplet, which forces the mixing in the quark sector related to that of the lepton sector. No redefinition of \bar{E} in generation-space can eliminate this mixing, as can be seen from Eq. (38). This mixing causes the off-diagonal elements of (m_ν^2) via radiative corrections.

As for the origin of the observed mixing angle between the left-handed neutrinos a parallel discussion to that of the previous section applies. The Majorana mass matrix in Eq. (38) has an intergenerational mixing as

$$M_{\nu_i \nu_j} = U_{ik}^* M_{\nu_k} U_{kj}^\dagger. \quad (39)$$

The mass matrix of the left-handed neutrinos (m_ν) is then

$$(m_\nu)_{ij} = V_{Dik}^\dagger (\bar{m}_\nu)_{kl} V_{Dlj}, \quad (40)$$

where

$$(\bar{m}_\nu)_{ij} = m_{\nu_i D} [M^{-1}]_{ij} m_{\nu_j D} \\ \equiv V_{Mik}^\dagger m_{\nu_k} V_{Mkj}. \quad (41)$$

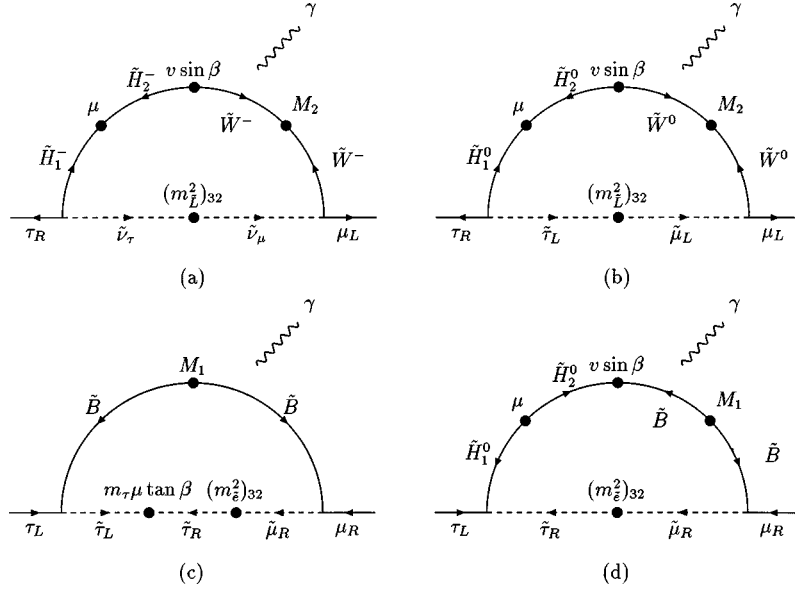


FIG. 10. Candidates of the Feynman diagrams which give dominant contributions to $\tau^+ \rightarrow \mu^+ \gamma$ when $\tan \beta \geq 1$ and $(m_{\tilde{e}}^2)_{32}$ is not negligible. The arrows represent the chirality.

Here also $m_{\nu_{iD}} = f_{\nu_i} v \sin \beta / \sqrt{2}$, the same notation as that of the previous section. The discussion in the previous section shows that the large mixing angle from V_M requires a fine-tuning between the elements of $M_{\nu_{ij}}$ [Eq. (12)]. Therefore as for the large mixing angle between neutrinos it is natural that its origin is in the mixing matrix V_D in Eq. (38). Here we assume $U_{ij} = \delta_{ij}$, for simplicity, which means that the mixing comes only from V_D .

Now we evaluate the off-diagonal elements of the slepton mass matrices at the low energy. As stated above, both the left- and right-handed slepton's ones have non-negligible off-diagonal elements at one-loop level. Assuming $m_{\nu_e} \ll m_{\nu_\mu} \ll m_{\nu_\tau}$ we neglect f_{ν_1} , and also f_{u_1} and f_{u_2} to obtain approximate formulas for the off-diagonal elements of the slepton mass matrices as

$$(m_{\tilde{e}}^2)_{ij} \approx -\frac{3}{8\pi^2} f_{u_3}^2 (V_{\text{KM}})_{3i} (V_{\text{KM}}^*)_{3j} (3m_0^2 + a_0^2) \log \frac{M_{\text{grav}}}{M_{\text{GUT}}}, \quad (42)$$

$$(m_{\tilde{L}}^2)_{ij} \approx -\frac{1}{8\pi^2} \left(f_{\nu_3}^2 V_{D3i}^* V_{D3j} \log \frac{M_{\text{grav}}}{M_{\nu_3}} + f_{\nu_2}^2 V_{D2i}^* V_{D2j} \log \frac{M_{\text{grav}}}{M_{\nu_2}} \right) (3m_0^2 + a_0^2), \quad (43)$$

$$A_e^{ij} \approx -\frac{3}{8\pi} a_0 \left(f_{e_i} V_{D3i}^* V_{D3j} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}} + f_{e_i} V_{D2i}^* V_{D2j} f_{\nu_2}^2 \log \frac{M_{\text{grav}}}{M_{\nu_2}} + 3f_{e_j} V_{\text{KM}3j}^* V_{\text{KM}3i} f_{u_3}^2 \log \frac{M_{\text{grav}}}{M_{\text{GUT}}} \right), \quad (44)$$

for $i \neq j$. These formulas are obtained by a logarithmic approximation from the RGE's (given in Appendix D).

Now we study the individual LFV processes. First we concentrate on the $\tau \rightarrow \mu \gamma$ decay. For $\tau \rightarrow \mu \gamma$ the most important contribution is from diagrams which involves $(m_{\tilde{L}}^2)_{32}$ and W -inos [that is, diagrams of Figs. 10(a) and 10(b)], which is the common feature with the MSSMRN case. These diagrams dominate because the large V_{D32} element, suggested by the atmospheric neutrino anomaly, enhances $(m_{\tilde{L}}^2)_{32}$ as

$$(m_{\tilde{L}}^2)_{32} \approx -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) V_{D33}^* V_{D32} f_{\nu_3}^2 \log \frac{M_{\text{grav}}}{M_{\nu_3}}, \quad (45)$$

which is the same situation as in the MSSMRN case. The main difference from the MSSMRN case is a presence of $(m_{\tilde{e}}^2)_{32}$, but the contribution to the $\tau \rightarrow \mu \gamma$ is too small at the broad parameter region to be comparable to those from Figs. 10(a) and 10(b), because $(m_{\tilde{e}}^2)_{32}$ is suppressed by small $(V_{\text{KM}})_{32}$ [28]. Our result of numerical calculation, Fig. 11, indeed shows that almost the same situation as in the MSSMRN case is realized. In the figure we plot the dependence of the branching ratio of $\tau \rightarrow \mu \gamma$ on the third generation right-handed Majorana mass M_{ν_3} for $\tan \beta = 3, 10,$ and 30 . The upper curve corresponds to larger $\tan \beta$. We take the B -ino mass as 65 GeV, the right-handed selectron mass 160 GeV, and the tau neutrino mass 0.07 eV, as expected from the atmospheric neutrino result. We take $a_0 = 0$ for simplicity. The figure shows us that the branching ratio of $\tau \rightarrow \mu \gamma$ is nearly proportional to the square of M_{ν_3} . At the right-hand side of each curve the Yukawa coupling constant f_{ν_3} blows up below the gravitational scale, so the perturbative treatment is no longer valid in this region. In the region near $M_{\nu_3} \approx 10^{14}$ GeV the branching ratio is close to or even be-

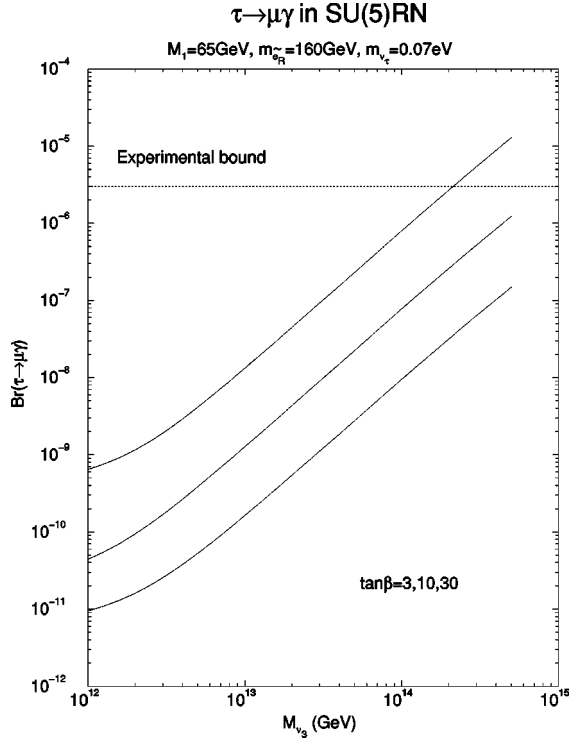


FIG. 11. Dependence of the branching ratio of $\tau \rightarrow \mu \gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} in the SU(5) SUSY GUT with the right-handed neutrinos. Here the tau neutrino mass m_{ν_τ} is 0.07 eV and $V_{D32} = -0.71$, as suggested by the atmospheric neutrino result. We take the B -ino mass M_1 65 GeV, the right-handed selectron mass $m_{\tilde{e}_R}$ 160 GeV. The three curves correspond to the case where $\tan \beta = 3, 10$, and 30, respectively. The branching ratio becomes larger for larger $\tan \beta$ value.

yond the current experimental bound, $\text{Br}(\tau \rightarrow \mu \gamma) < 3.0 \times 10^{-6}$ [14]. At relatively small M_{ν_3} region ($M_{\nu_3} \lesssim 4 \times 10^{12}$ GeV) the contribution from the right-handed slepton mass matrix [the diagrams shown in Figs. 10(c) and 10(d)] tends to dominate, and the curves of the branching ratio show deviation from simple straight lines.

Next the $\mu \rightarrow e \gamma$ process. Since we know $V_{D32} = \mathcal{O}(1)$ from the atmospheric neutrino result we can calculate $(m_{\tilde{L}}^2)_{23}$ as

$$(m_{\tilde{L}}^2)_{23} \approx -\frac{1}{8\pi^2} f_{\nu_3}^2 V_{D32}^* V_{D33} (3m_0^2 + a_0^2) \log \frac{M_{\text{grav}}}{M_{\nu_3}}. \quad (46)$$

On the other hand $(m_{\tilde{e}}^2)_{31}$ is determined from the GUT symmetry as

$$(m_{\tilde{e}}^2)_{31} \approx -\frac{3}{8\pi^2} f_{u_3}^2 (V_{\text{KM}})_{33} (V_{\text{KM}}^*)_{31} (3m_0^2 + a_0^2) \log \frac{M_{\text{grav}}}{M_{\text{GUT}}}. \quad (47)$$

Then we can definitely calculate the diagram shown in Fig. 12(e), which contributes to $A_R^{(\mu e)}$ defined in Eq. (20). This contribution is proportional to m_τ , and it tends to dominate

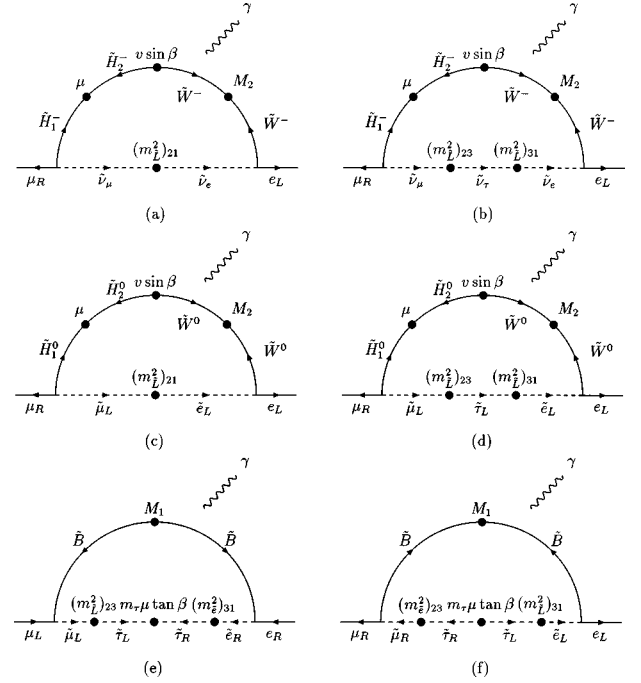


FIG. 12. Candidates of the Feynman diagrams which give dominant contributions to $\mu^+ \rightarrow e^+ \gamma$ when $\tan \beta \geq 1$ and the off-diagonal elements of $(m_{\tilde{e}}^2)$ are non-negligible. The arrows represent the chirality.

over the other contribution to $A_R^{(\mu e)}$. We need to know V_{D31} and V_{D21} to calculate $A_L^{(\mu e)}$. However, we can evaluate only the lower bound of the branching ratio from the structure given in Eq. (21) even if we do not know them. The contribution from the diagram 12(e) can be so large that it reaches the present experimental bound at some parameter regions [28].

We can imagine some cases where much larger rate than this lower bound is predicted by the contribution from $A_L^{(\mu e)}$. One of such cases is that V_{D31} is large. In this case, the diagrams in Figs. 12(a)–12(d) and 12(f) are enhanced since $(m_{\tilde{L}}^2)_{31}$ and $(m_{\tilde{L}}^2)_{21}$ are proportional to V_{D31} as

$$(m_{\tilde{L}}^2)_{31} \approx -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) f_{\nu_3}^2 V_{D33}^* V_{D31} \log \frac{M_{\text{grav}}}{M_{\nu_3}}, \quad (48)$$

$$(m_{\tilde{L}}^2)_{21} \approx -\frac{1}{8\pi^2} (3m_0^2 + a_0^2) f_{\nu_3}^2 V_{D32}^* V_{D31} \log \frac{M_{\text{grav}}}{M_{\nu_3}}. \quad (49)$$

While the diagram 12(f) is enhanced by m_τ , the contribution is suppressed by $(m_{\tilde{e}}^2)_{32}$, which is proportional to $(V_{\text{KM}})_{32}$. As a result, the diagrams 12(a)–12(d) dominate since V_{D32} is large. Also, if f_{ν_2} and V_{D21} are non-negligibly large, $(m_{\tilde{e}}^2)$ is enhanced as

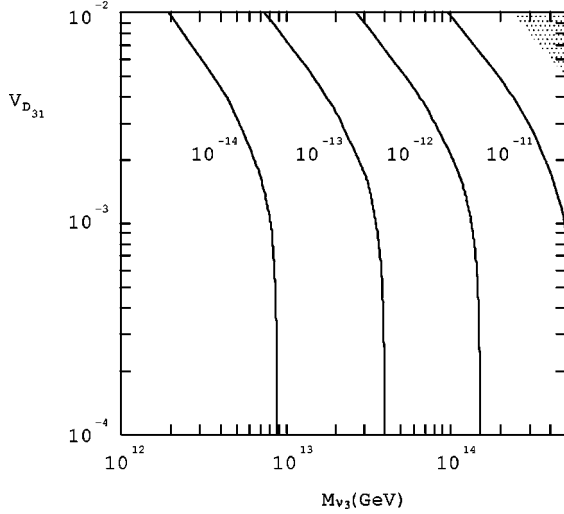


FIG. 13. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} and V_{D31} in the SU(5) SUSY GUT with the right-handed neutrinos. Here we take the tau neutrino mass m_{ν_τ} 0.07 eV and the mu neutrino mass is neglected. The curves mean the contours on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-14} , 10^{-13} , 10^{-12} , and 10^{-11} , respectively. The shaded region is already excluded experimentally. We take the B -ino mass M_1 65 GeV, the right-handed selectron mass $m_{\tilde{e}_R}$ 160 GeV, and $\tan\beta=3$.

$$(m_L^2)_{21} \approx -\frac{1}{8\pi^2} \left(f_{\nu_3}^2 V_{D32}^* V_{D31} \log \frac{M_{\text{grav}}}{M_{\nu_3}} + f_{\nu_2}^2 V_{D22}^* V_{D21} \log \frac{M_{\text{grav}}}{M_{\nu_2}} \right) (3m_0^2 + a_0^2), \quad (50)$$

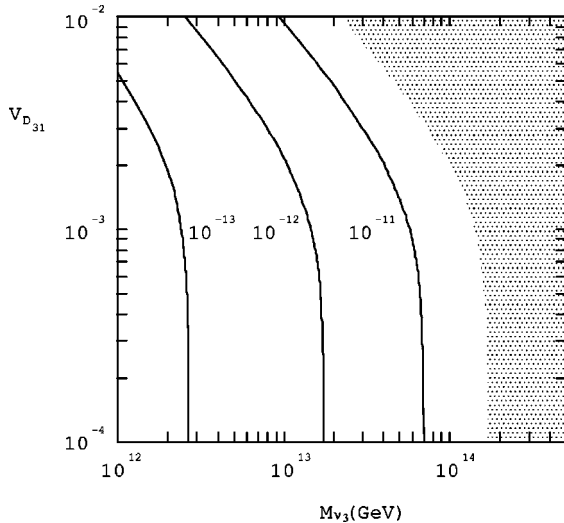


FIG. 14. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the third-generation right-handed neutrino Majorana mass M_{ν_3} and V_{D31} in the SU(5) SUSY GUT with the right-handed neutrinos. The curves mean the contours on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-13} , 10^{-12} , and 10^{-11} , respectively. The shaded region is already excluded experimentally. The input parameters are the same as those in Fig. 13 except that in this figure $\tan\beta=30$.

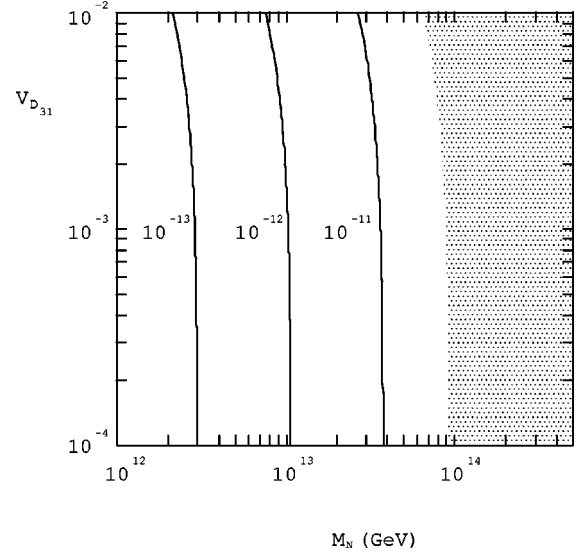


FIG. 15. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the typical right-handed neutrino Majorana mass M_N and V_{D31} in the SU(5) SUSY GUT with the right-handed neutrinos. We assume the MSW large angle solution, which suggests m_{ν_μ} to be 0.004 eV and $V_{D21} = -0.42$. We take the tau neutrino mass m_{ν_τ} 0.07 eV and $V_{D32} = -0.60$, as suggested by the atmospheric neutrino result. We assume the universality of the right-handed Majorana masses $M_{\nu_1} = M_{\nu_2} = M_{\nu_3} \equiv M_N$, for simplicity. The curves mean the contours on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-13} , 10^{-12} , and 10^{-11} , respectively. The shaded region is already excluded experimentally. We take the B -ino mass M_1 65 GeV, and the right-handed selectron mass $m_{\tilde{e}_R}$ 160 GeV. In this figure $\tan\beta=3$.

and the diagrams 12(a) and 12(c) dominate over the other contributions.

Now we examine these expectations numerically. First we set f_{ν_2} to zero and later investigate the nonzero f_{ν_2} case. We calculate the dependence of the branching ratio of $\mu \rightarrow e\gamma$ on M_{ν_3} and V_{D31} in Figs. 13 and 14. In the figures we take $V_{D32} = -1/\sqrt{2}$ and the tau neutrino mass 0.07 eV, as suggested by the atmospheric neutrino result. We neglect m_{ν_μ} here. Other parameters are the B -ino mass 65 GeV, the right-handed selectron mass 160 GeV, $\tan\beta=3$ and 30, and the Higgsino mass $\mu > 0$. From the Figs. 13 and 14 we can see that for $V_{D31} \lesssim 10^{-3}$ the diagram 12(e) dominates over the others. At relatively larger V_{D31} region ($V_{D31} \gtrsim 10^{-2.5}$) the diagrams 12(a)–12(d) become dominant, and the predicted rate is large enough to reach the experimental bound for $V_{D31} \approx 10^{-2}$ if $\tan\beta=30(3)$ and $M_{\nu_3} \gtrsim 10^{13.5}(10^{14.5})$ GeV.

Next let us consider the finite f_{ν_2} case. We show in Figs. 15 and 16 the dependence of the $\mu \rightarrow e\gamma$ branching ratio on V_{D31} and the typical right-handed neutrino Majorana mass M_N . In the figures the parameters that describe the MSW large angle solution are taken as the same as those we used in the MSSMRN case, and other parameters are taken as the same as in Figs. 13 and 14. We assume in the figure the universality of the right-handed Majorana mass, that is, $M_{\nu_1} = M_{\nu_2} = M_{\nu_3} (\equiv M_N)$. We can see from the figures that the enhancement due to the MSW large angle solution is so

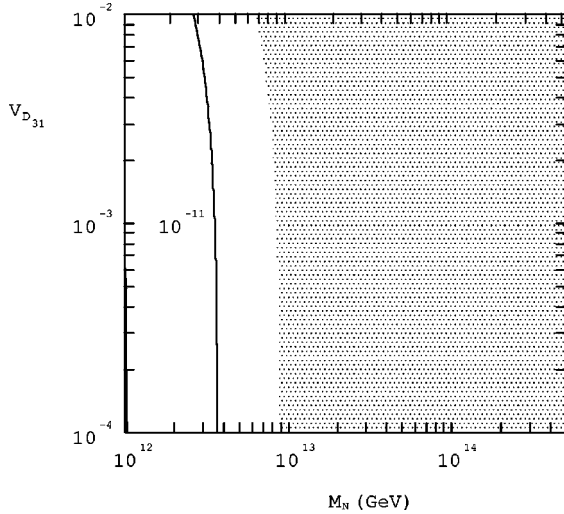


FIG. 16. Dependence of the branching ratio of $\mu \rightarrow e\gamma$ on the typical right-handed neutrino Majorana mass M_N and $V_{D_{31}}$ in the SU(5) SUSY GUT with the right-handed neutrinos. We assume the MSW large angle solution and the atmospheric neutrino result. All the input parameters are the same as those in Fig. 15 except that we take $\tan\beta=30$ in this figure. The curve means the contour on which the branching ratio of $\mu \rightarrow e\gamma$ is 10^{-11} . The shaded region is already excluded experimentally.

large that the dependence of the rate on $V_{D_{31}}$ is small, and that the almost same situation as in the MSSMRN is reproduced here again, as can be seen when comparing these figures with Fig. 5. Since the dominant contribution is determined by the second term of Eq. (50), the rate is almost determined from the value of M_{ν_2} because $f_{\nu_2}^2 \propto M_{\nu_2}$ for fixed m_{ν_μ} . Hence we can conclude from Figs. 15, 16 that for $\tan\beta=30(3)$ the excluded region of M_{ν_2} extends to $9 \times 10^{12}(1 \times 10^{14})$ GeV, at least for the parameters chosen in our calculation.

We also calculated for the parameters suggested from the MSW small angle solution and the ‘‘just so’’ solution. For the MSW small angle solution the difference from the $f_{\nu_2}=0$ case is at most the factor 2 enhancement, and for the ‘‘just so’’ solution no difference from the $f_{\nu_2}=0$ case can be seen.

Finally we comment on the $b \rightarrow s\gamma$ process. This model has a characteristic feature that the right-handed neutrinos couple to the right-handed down-type quarks with the large mixing through the ninth term of Eq. (38). This coupling induces the off-diagonal elements in the right-handed down-type squark soft mass matrix via the radiative corrections, which are as large as

$$(m_{d_R}^2)_{ij} \approx -\frac{1}{8\pi^2} f_{\nu_k}^2 V_{Dkj}^* V_{Dki} (3m_0^2 + a_0^2) \log \frac{M_{\text{grav}}}{M_{\text{GUT}}}. \quad (51)$$

This fact causes us an expectation that the $b \rightarrow s\gamma$ rate may become larger by the effect of the diagram which contains $(m_{d_R}^2)_{23}$, by an exact analogy to that of the lepton sector. We

examined whether this is true or not, and concluded that a tiny enhancement indeed occurs but it is too small to be experimentally distinguishable. This is because the heavy gluino suppresses the contribution.

IV. CONCLUSION

In this article taking the solar and the atmospheric neutrino experiment results into account we investigated the lepton flavor violating decay processes, such as $\mu \rightarrow e\gamma$ or $\tau \rightarrow \mu\gamma$ in the MSSM with the right-handed neutrinos (MSSMRN) and in the SU(5) SUSY GUT with the right-handed neutrinos [SU(5)RN].

In the MSSMRN we first studied the branching ratio of $\tau \rightarrow \mu\gamma$. It gets larger for larger $\tan\beta$ and almost proportional to $M_{\nu_3}^2$ for fixed m_{ν_τ} . A large rate naturally follows from the large mixing between ν_τ and ν_μ , suggested by the atmospheric neutrino result. If $\tan\beta=30(10)$ the rate reaches the current experimental bound for $M_{\nu_3} \sim 2 \times 10^{14}(6 \times 10^{14})$ GeV. We investigated the $\mu \rightarrow e\gamma$ rate under three kinds of the solar neutrino solutions, the MSW large and small angle solutions and the ‘‘just so’’ solution. We argued that the $\mu \rightarrow e\gamma$ depends on f_{ν_2} , and that especially the large $V_{D_{21}}$ and the large f_{ν_2} , which the MSW large angle solution suggests, naturally results in such a large rate that the predicted rate is beyond the current experimental bound if M_{ν_2} is larger than $8 \times 10^{12}(8 \times 10^{13})$ GeV for $\tan\beta=30(3)$. We also investigated the dependence of $\mu \rightarrow e\gamma$ on $V_{D_{31}}$. For $\tan\beta=30$ and $V_{D_{31}}=10^{-2}$, $M_{\nu_3} \approx 3 \times 10^{13}$ GeV is excluded for our input parameters.

In the SU(5)RN we calculated the $\tau \rightarrow \mu\gamma$ rate. Here also the large mixing between ν_τ and ν_μ leads to a large rate, which is almost the same situation as in the MSSMRN. For $\mu \rightarrow e\gamma$ in this model, we can predict the lower bound of the branching ratio of it. This lower bound is calculable from the value of m_{ν_τ} and the large mixing angle between ν_τ - ν_μ , suggested from the atmospheric neutrino result, and it turns out to be within accessible region by near future experiments. They are supposed to probe the $\mu \rightarrow e\gamma$ branching ratio to 10^{-14} level [18], and then the region $M_{\nu_3} > 10^{13}(10^{12})$ GeV can be probed for $\tan\beta=3(30)$ and $m_{\nu_\tau}=0.07$ eV. We considered the relation between $\mu \rightarrow e\gamma$ and three kinds of the solar neutrino solutions. The large rate naturally follows from the MSW large angle solution, similarly to the MSSMRN case, while in the MSW small angle solution and in the ‘‘just so’’ solution there is only a little difference from the $f_{\nu_2}=0$ case.

If the LFV processes are discovered or the experimental bounds are improved by near future experiments, the interesting insight on the lepton sector will be obtained. The best effort to implement it is strongly desired.

ACKNOWLEDGMENTS

The authors would like to thank K. Tobe for his participation to the preparation of Appendices B and C at the early stage. One of the authors (D.N.) would like to express his

gratitude to T. Yanagida for his continuous support, and thanks the members of the theory group of KEK for the hospitality extended to him during his stay at KEK. This work is partially supported by Grants-in-aid for Science and Culture Research for Monbusho, No. 10740133 and ‘‘Priority Area: Supersymmetry and Unified Theory of Elementary Particles (#707).’’

APPENDIX A: DEFINITIONS AND CONVENTIONS IN THE MSSM

In this appendix we collect our notations of the MSSM used in our article. The superpotential of the MSSM W_{MSSM} is defined as^{5,6}

$$W_{\text{MSSM}} \equiv f_{e_{ij}} H_1 \bar{E}_i L_j + f_{d_{ij}} H_1 Q_i \bar{D}_j + f_{u_{ij}} H_2 Q_i \bar{U}_j + \mu H_1 H_2. \quad (\text{A4})$$

We use i and j as the generation indices running from 1 to 3. \bar{E} , \bar{D} , and \bar{U} are the superfields associated with the right-handed electron e_R , the right-handed down-type quark d_R and the right-handed up-type quark u_R , respectively. Q and L are ones associated with the quark doublet q_L and the lepton doublet l_L defined by

$$q_L = \begin{pmatrix} u_L \\ d_L \end{pmatrix}, \quad l_L = \begin{pmatrix} \nu \\ e_L \end{pmatrix}, \quad (\text{A5})$$

and H_1 and H_2 the Higgs doublets whose $\text{SU}(2)_L$ components we denote as

$$H_1 = \begin{pmatrix} H_1^0 \\ H_1^- \end{pmatrix}, \quad H_2 = \begin{pmatrix} H_2^+ \\ H_2^0 \end{pmatrix}. \quad (\text{A6})$$

The scalar components of H_1^0 and H_2^0 develop the vacuum expectation values (VEV's) as

$$\langle H_1^0 \rangle \equiv \frac{v_1}{\sqrt{2}}, \quad \langle H_2^0 \rangle \equiv \frac{v_2}{\sqrt{2}}, \quad (\text{A7})$$

which satisfy $v_1^2 + v_2^2 = v^2$ with $v \simeq 246$ GeV. We define the ratio of these VEV's as $\tan \beta$,

⁵Supersymmetric couplings and masses of chiral multiplets are given as

$$-\mathcal{L} = \int d^2\theta W + \text{H.c.} \quad (\text{A1})$$

Our convention is unusual in order to keep the chargino mass matrix in accordance with the Haber-Kane convention [29].

⁶We implicitly assume the contraction convention over $\text{SU}(2)_L$ doublet indices ($a, b, \dots = 1, 2$) of two doublets A and B

$$A = \begin{pmatrix} A^1 \\ A^2 \end{pmatrix}, \quad B = \begin{pmatrix} B^1 \\ B^2 \end{pmatrix} \quad (\text{A2})$$

as

$$AB \equiv \epsilon_{ab} A^a B^b, \quad (\text{A3})$$

where ϵ_{ab} is an antisymmetric tensor with $\epsilon_{12} = -\epsilon_{21} = 1$.

$$\tan \beta \equiv \frac{v_2}{v_1}. \quad (\text{A8})$$

The Yukawa couplings are given by the fermion masses and the Kobayashi-Maskawa matrix as

$$\begin{aligned} f_{e_{ij}} &= -\sqrt{2} \frac{m_{e_i}}{v \cos \beta} \delta_{ij}, \\ f_{d_{ij}} &= -\sqrt{2} \frac{m_{d_i}}{v \cos \beta} \delta_{ij}, \\ f_{u_{ij}} &= \sqrt{2} \frac{m_{u_i}}{v \sin \beta} (V_{\text{KM}})_{ij}. \end{aligned} \quad (\text{A9})$$

We also introduce the soft SUSY breaking terms in the Lagrangian,

$$\begin{aligned} -\mathcal{L}_{\text{SUSY breaking}} &= (m_L^2)_{ij} \tilde{l}_L^i \tilde{l}_L^j + (m_{\tilde{e}}^2)_{ij} \tilde{e}_{Ri}^* \tilde{e}_{Rj} + (m_{\tilde{Q}}^2)_{ij} \tilde{q}_{Li}^\dagger \tilde{q}_{Lj} \\ &+ (m_{\tilde{u}}^2)_{ij} \tilde{u}_{Ri}^* \tilde{u}_{Rj} + (m_{\tilde{d}}^2)_{ij} \tilde{d}_{Ri}^* \tilde{d}_{Rj} + \tilde{m}_{h1}^2 h_1^\dagger h_1 \\ &+ \tilde{m}_{h2}^2 h_2^\dagger h_2 + (A_u^{ij} h_2 \tilde{q}_{Li} \tilde{u}_{Rj}^* + A_d^{ij} h_1 \tilde{q}_{Li} \tilde{d}_{Rj}^* \\ &+ A_e^{ij} h_1 \tilde{e}_{Ri}^* \tilde{l}_L^j + B_h h_1 h_2 + \text{H.c.}) \\ &+ \left(\frac{1}{2} M_1 \tilde{B}_L \tilde{B}_L + \frac{1}{2} M_2 \tilde{W}_L^a \tilde{W}_L^a \right. \\ &\left. + \frac{1}{2} M_3 \tilde{g}_L^a \tilde{g}_L^a + \text{H.c.} \right). \end{aligned} \quad (\text{A10})$$

Here the first seven terms are the soft SUSY breaking masses for the doublet-slepton \tilde{l}_L , the right-handed charged slepton \tilde{e}_R , the doublet-squark \tilde{q}_L , the right-handed up-type (down-type) squark \tilde{u}_R (\tilde{d}_R), and the Higgs bosons. \tilde{B} , \tilde{W} , and \tilde{g} stand for B -ino, W -ino, and gluino respectively, and the superscript a is the gauge group index for each corresponding gauge group.

Now we discuss the slepton mass matrices. Let \tilde{e}_{Li} and \tilde{e}_{Ri} be the superpartners of the left-handed electron e_{Li} and the right-handed electron e_{Ri} , respectively. Then the slepton mass matrix $(\tilde{m}_{\tilde{e}}^2)_{ij}$ is

$$\begin{aligned} -\mathcal{L} &= (\tilde{e}_{Li}^\dagger, \tilde{e}_{Ri}^\dagger) (\tilde{m}_{\tilde{e}}^2)_{ij} \begin{pmatrix} \tilde{e}_{Lj} \\ \tilde{e}_{Rj} \end{pmatrix} \\ &\equiv (\tilde{e}_{Li}^\dagger, \tilde{e}_{Ri}^\dagger) \begin{pmatrix} (m_L^2)_{ij} & (m_{LR}^{2\text{T}})_{ij} \\ (m_{LR}^2)_{ij} & (m_R^2)_{ij} \end{pmatrix} \begin{pmatrix} \tilde{e}_{Lj} \\ \tilde{e}_{Rj} \end{pmatrix}. \end{aligned} \quad (\text{A11})$$

Here m_L^2 and m_R^2 are 3×3 Hermitian matrices and m_{LR}^2 is a 3×3 matrix. They are given as

$$(m_L^2)_{ij} = (m_{\tilde{L}}^2)_{ij} + m_{e_i}^2 \delta_{ij} + m_Z^2 \delta_{ij} \cos 2\beta \left(-\frac{1}{2} + \sin^2 \theta_W \right), \quad (\text{A12})$$

$$(m_R^2)_{ij} = (m_e^2)_{ij} + m_{e_i}^2 \delta_{ij} - m_Z^2 \delta_{ij} \cos 2\beta \sin^2 \theta_W, \quad (\text{A13})$$

$$(m_{LR}^2)_{ij} = A_e^{ij} v \cos \beta / \sqrt{2} - m_{e_i} \mu \delta_{ij} \tan \beta. \quad (\text{A14})$$

As for the neutrinos we should notice that there is no right-handed sneutrino in the MSSM. Let $\tilde{\nu}_i$ be the superpartner of the left-handed neutrino ν_i . The mass matrix of sneutrino $(\tilde{m}_{\tilde{\nu}}^2)_{ij}$ is

$$-\mathcal{L} = \tilde{\nu}_i^\dagger (\tilde{m}_{\tilde{\nu}}^2)_{ij} \tilde{\nu}_j, \quad (\tilde{m}_{\tilde{\nu}}^2)_{ij} = (m_L^2)_{ij} + \frac{1}{2} m_Z^2 \delta_{ij} \cos 2\beta. \quad (\text{A15})$$

Next we discuss the mass matrix of gauginos. First we consider chargino mass matrix \mathcal{M}_C . It is a 2×2 matrix that appears in the chargino mass terms,

$$-\mathcal{L}_m = (\tilde{W}_R^- \tilde{H}_{2R}^-) \begin{pmatrix} M_2 & \sqrt{2} m_W \cos \beta \\ \sqrt{2} m_W \sin \beta & \mu \end{pmatrix} \begin{pmatrix} \tilde{W}_L^- \\ \tilde{H}_{1L}^- \end{pmatrix} + \text{H.c.} \quad (\text{A16})$$

$$\mathcal{M}_N = \begin{pmatrix} M_1 & 0 & -m_Z \sin \theta_W \cos \beta & m_Z \sin \theta_W \sin \beta \\ 0 & M_2 & m_Z \cos \theta_W \cos \beta & -m_Z \cos \theta_W \sin \beta \\ -m_Z \sin \theta_W \cos \beta & m_Z \cos \theta_W \cos \beta & 0 & -\mu \\ m_Z \sin \theta_W \sin \beta & -m_Z \cos \theta_W \sin \beta & -\mu & 0 \end{pmatrix}. \quad (\text{A21})$$

The diagonalization is done by a real orthogonal matrix O_N ,

$$O_N \mathcal{M}_N O_N^T = \text{diag}(M_{\tilde{\chi}_1^0}, \dots, M_{\tilde{\chi}_4^0}). \quad (\text{A22})$$

The mass eigenstates are given by

$$\tilde{\chi}_{AL}^0 = (O_N)_{AB} \tilde{\chi}_{BL}^0 \quad (A, B = 1, \dots, 4), \quad (\text{A23})$$

where

$$\tilde{\chi}_{AL}^0 = (\tilde{B}_L, \tilde{W}_L^0, \tilde{H}_{1L}^0, \tilde{H}_{2L}^0). \quad (\text{A24})$$

We have thus Majorana spinors

$$\tilde{\chi}_A^0 = \tilde{\chi}_{AL}^0 + \tilde{\chi}_{AR}^0 \quad (A = 1, \dots, 4) \quad (\text{A25})$$

with mass $M_{\tilde{\chi}_A^0}$.

Now we give the interaction Lagrangian of lepton-slepton-chargino (-neutralino) in a basis of the slepton weak eigenstate and the chargino (neutralino) mass eigenstate. By writing the interactions in this basis we can get a transparent view in the discussion of the multimass insertion technique discussed in Appendices B and C. The Lagrangian is

\mathcal{M}_C is diagonalized by 2×2 real orthogonal matrices O_L and O_R as

$$O_R \mathcal{M}_C O_L^T = \text{diag}(M_{\tilde{\chi}_1^-}, M_{\tilde{\chi}_2^-}). \quad (\text{A17})$$

Define the mass eigenstates $\tilde{\chi}_{A(L,R)}$ ($A = 1, 2$) by

$$\begin{pmatrix} \tilde{\chi}_{1L}^- \\ \tilde{\chi}_{2L}^- \end{pmatrix} = O_L \begin{pmatrix} \tilde{W}_L^- \\ \tilde{H}_{1L}^- \end{pmatrix}, \quad \begin{pmatrix} \tilde{\chi}_{1R}^- \\ \tilde{\chi}_{2R}^- \end{pmatrix} = O_R \begin{pmatrix} \tilde{W}_R^- \\ \tilde{H}_{2R}^- \end{pmatrix}. \quad (\text{A18})$$

Then

$$\tilde{\chi}_A^- = \tilde{\chi}_{AL}^- + \tilde{\chi}_{AR}^- \quad (A = 1, 2) \quad (\text{A19})$$

forms a Dirac fermion with mass $M_{\tilde{\chi}_A^-}$.

Finally we consider neutralinos. The mass matrix of the neutralino sector is given by

$$-\mathcal{L}_m = \frac{1}{2} (\tilde{B}_L \tilde{W}_L^0 \tilde{H}_{1L}^0 \tilde{H}_{2L}^0) \mathcal{M}_N \begin{pmatrix} \tilde{B}_L \\ \tilde{W}_L^0 \\ \tilde{H}_{1L}^0 \\ \tilde{H}_{2L}^0 \end{pmatrix} + \text{H.c.}, \quad (\text{A20})$$

where

$$-\mathcal{L}_{\text{int}} = \tilde{\nu}_i^\dagger \tilde{\chi}_A^- (C_{LR}^{A(i)} P_R + C_{LL}^{A(i)} P_L) e_i + \tilde{e}_{Li}^\dagger \tilde{\chi}_A^0 (N_{LR}^{A(i)} P_R + N_{LL}^{A(i)} P_L) e_i + \tilde{e}_{Ri}^\dagger \tilde{\chi}_A^0 (N_{RR}^{A(i)} P_R + N_{RL}^{A(i)} P_L) e_i + \text{H.c.}, \quad (\text{A26})$$

where the coefficients are

$$C_{LL}^{A(i)} = g_2 (O_R)_{A1},$$

$$C_{LR}^{A(i)} = -\frac{\sqrt{2} m_{e_i}}{v \cos \beta} (O_L)_{A2},$$

$$N_{LL}^{A(i)} = \frac{g_2}{\sqrt{2}} [-(O_N)_{A2} - (O_N)_{A1} \tan \theta_W],$$

$$N_{RL}^{A(i)} = \frac{\sqrt{2} m_{e_i}}{v \cos \beta} (O_N)_{A3},$$

$$N_{LR}^{A(i)} = \frac{\sqrt{2} m_{e_i}}{v \cos \beta} (O_N)_{A3},$$

$$N_{RR}^{A(i)} = \sqrt{2} g_2 (O_N)_{A1} \tan \theta_W. \quad (\text{A27})$$

APPENDIX B: MULTIMASS INSERTION TECHNIQUE

Mass insertion technique is useful to understand lepton flavor violating processes in the MSSM since only off-diagonal elements of the slepton mass matrices are source of the flavor violation. In this appendix we introduce the multimass insertion technique with the nondegenerate masses. The multimass inserted diagrams may not be necessarily suppressed than the single-mass inserted ones when the slepton mass matrix has several small flavor-violating elements. Also, though in the previous mass insertion formulas the degeneracy of the slepton masses is sometimes assumed, such a degeneracy is not necessarily maintained at low energy even if the universal scalar mass hypothesis is assumed in the higher energy scale. Our rule to derive the mass-inserted amplitudes is very simple. We can derive them by taking the finite difference on amplitudes which have no mass-insertion. In this section we derive general formulas of $e_i^+ \rightarrow e_j^+ \gamma$ keeping in mind that we apply them to more realistic cases of $\mu^+ \rightarrow e^+ \gamma$ and $\tau^+ \rightarrow \mu^+ \gamma$ in the next section.

We refer to the slepton mass matrices as $(\bar{m}_{\tilde{l}}^2)$,

$$-\mathcal{L} = \sum_{\tilde{l}=\bar{e},\bar{\nu}} (\bar{m}_{\tilde{l}}^2)_{ij} \tilde{l}_i^\dagger \tilde{l}_j, \quad (\text{B1})$$

where

$$\begin{aligned} \bar{e}_i &= (\bar{e}_L, \bar{\mu}_L, \bar{\tau}_L, \bar{e}_R, \bar{\mu}_R, \bar{\tau}_R)^\top \quad [\text{for } \tilde{l}_i = \bar{e}_i \ (i=1-6)], \\ \bar{\nu}_i &= (\bar{\nu}_e, \bar{\nu}_\mu, \bar{\nu}_\tau)^\top \quad [\text{for } \tilde{l}_i = \bar{\nu}_i \ (i=1-3)]. \end{aligned} \quad (\text{B2})$$

The explicit forms of these matrices are given in Appendix A. Here we assume that all the off-diagonal elements of the slepton mass matrices, both flavor violating and conserving ones, are much smaller than the diagonal elements $[(\bar{m}_{\tilde{l}}^2)_{ii} \gg (\bar{m}_{\tilde{l}}^2)_{jk} \ (j \neq k)].$

In the mass-insertion technique the internal slepton lines are classified to two types at one loop level. First type is a slepton line on which the momentum of slepton is not changed as

$$\begin{aligned} & \begin{array}{c} \tilde{l}_1 \quad \tilde{l}_2 \quad \tilde{l}_3 \quad \dots \quad \tilde{l}_N \\ \text{---} \bullet \text{---} \bullet \text{---} \bullet \text{---} \bullet \text{---} \text{---} \\ (\bar{m}_{\tilde{l}}^2)_{l_1 l_2} \quad (\bar{m}_{\tilde{l}}^2)_{l_2 l_3} \quad (\bar{m}_{\tilde{l}}^2)_{l_3 l_4} \quad \dots \quad (\bar{m}_{\tilde{l}}^2)_{l_{N-1} l_N} \\ \longrightarrow \\ k \end{array} \\ &= \frac{1}{k^2 - \bar{m}_{\tilde{l}_1}^2} (\bar{m}_{\tilde{l}}^2)_{l_1 l_2} \frac{1}{k^2 - \bar{m}_{\tilde{l}_2}^2} (\bar{m}_{\tilde{l}}^2)_{l_2 l_3} \frac{1}{k^2 - \bar{m}_{\tilde{l}_3}^2} (\bar{m}_{\tilde{l}}^2)_{l_3 l_4} \dots (\bar{m}_{\tilde{l}}^2)_{l_{N-1} l_N} \frac{1}{k^2 - \bar{m}_{\tilde{l}_N}^2}, \end{aligned} \quad (\text{B3})$$

where we call the diagonal components of the slepton mass matrices, $(\bar{m}_{\tilde{l}}^2)_{ii}$, as $\bar{m}_{\tilde{l}_i}^2$. We would like to consider the $e_i^+ \rightarrow e_j^+ \gamma$ process, and so in the above figure we mean that an antislepton is going from left to right with a momentum k . The product of propagators in the above equation is referred to as F_N ,

$$F_N(\bar{m}_{\tilde{l}_1}^2, \bar{m}_{\tilde{l}_2}^2, \dots, \bar{m}_{\tilde{l}_N}^2) \equiv \frac{1}{k^2 - \bar{m}_{\tilde{l}_1}^2} \frac{1}{k^2 - \bar{m}_{\tilde{l}_2}^2} \dots \frac{1}{k^2 - \bar{m}_{\tilde{l}_N}^2}. \quad (\text{B4})$$

The second type is a slepton line where the momentum is changed, by an emission of a photon with an outgoing momentum p , as

$$\begin{aligned}
&= \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} \frac{1}{k^2 - \bar{m}_{l_1}^2} (\bar{m}_{l_1}^2)_{l_1 l_2} \frac{1}{k^2 - \bar{m}_{l_2}^2} (\bar{m}_{l_2}^2)_{l_2 l_3} \cdots (\bar{m}_{l_{N-1}}^2)_{l_{N-1} l_N} \frac{1}{k^2 - \bar{m}_{l_N}^2} \\
&+ \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} (\bar{m}_{l_1}^2)_{l_1 l_2} \frac{1}{(k+p)^2 - \bar{m}_{l_2}^2} \frac{1}{k^2 - \bar{m}_{l_2}^2} (\bar{m}_{l_2}^2)_{l_2 l_3} \cdots (\bar{m}_{l_{N-1}}^2)_{l_{N-1} l_N} \frac{1}{k^2 - \bar{m}_{l_N}^2} \\
&+ \cdots \\
&+ \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} (\bar{m}_{l_1}^2)_{l_1 l_2} \frac{1}{(k+p)^2 - \bar{m}_{l_2}^2} (\bar{m}_{l_2}^2)_{l_2 l_3} \cdots (\bar{m}_{l_{N-1}}^2)_{l_{N-1} l_N} \frac{1}{(k+p)^2 - \bar{m}_{l_N}^2} \frac{1}{k^2 - \bar{m}_{l_N}^2}.
\end{aligned} \tag{B5}$$

The sum of product of propagators in this equation is referred to as G_N ,

$$\begin{aligned}
G_N(\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2) &\equiv \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} \frac{1}{k^2 - \bar{m}_{l_1}^2} \frac{1}{k^2 - \bar{m}_{l_1}^2} \cdots \frac{1}{k^2 - \bar{m}_{l_N}^2} + \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} \frac{1}{(k+p)^2 - \bar{m}_{l_2}^2} \frac{1}{k^2 - \bar{m}_{l_2}^2} \cdots \frac{1}{k^2 - \bar{m}_{l_N}^2} \\
&+ \cdots + \frac{1}{(k+p)^2 - \bar{m}_{l_1}^2} \frac{1}{(k+p)^2 - \bar{m}_{l_2}^2} \cdots \frac{1}{(k+p)^2 - \bar{m}_{l_N}^2} \frac{1}{k^2 - \bar{m}_{l_N}^2}.
\end{aligned} \tag{B6}$$

These functions F_N and G_N can be given as a finite difference of F_{N-1} and G_{N-1} ,

$$\begin{aligned}
F_N(\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2) &= D[F_{N-1}(\bar{m}_{l_1}^2, \bar{m}_{l_3}^2, \dots, \bar{m}_{l_N}^2); \bar{m}_{l_2}^2] \\
&\times (\bar{m}_{l_1}^2, \bar{m}_{l_2}^2),
\end{aligned} \tag{B7}$$

$$\begin{aligned}
G_N(\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2) &= D[G_{N-1}(\bar{m}_{l_2}^2, \bar{m}_{l_3}^2, \dots, \bar{m}_{l_N}^2); \bar{m}_{l_1}^2] \\
&\times (\bar{m}_{l_1}^2, \bar{m}_{l_2}^2),
\end{aligned} \tag{B8}$$

where

$$D[f(x); x](x_1, x_2) \equiv \frac{1}{x_1 - x_2} [f(x_1) - f(x_2)]. \tag{B9}$$

Then, by taking finite difference in sequence, each scalar line can be represented as a linear combination of flavor conserving scalar lines, F_1 or G_1 ,

$$\begin{aligned}
F_N(\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2) &= D^{N-1}[F_1(m^2); m^2] \\
&\times (\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2), \\
G_N(\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2) &= D^{N-1}[G_1(m^2); m^2] \\
&\times (\bar{m}_{l_1}^2, \bar{m}_{l_2}^2, \dots, \bar{m}_{l_N}^2),
\end{aligned} \tag{B10}$$

where

$$D^{N-1}[f(x); x](x_1, x_2, \dots, x_N) \equiv \sum_{i=1}^N \left(\prod_{j \neq i} \frac{1}{x_i - x_j} \right) f(x_i). \tag{B11}$$

Because of this fact, we can get amplitudes with any masses inserted from the corresponding flavor-conserving diagrams.

Before we derive amplitudes of $e_i^+ \rightarrow e_j^+ \gamma (i > j)$, we introduce a mass function I_N^M as

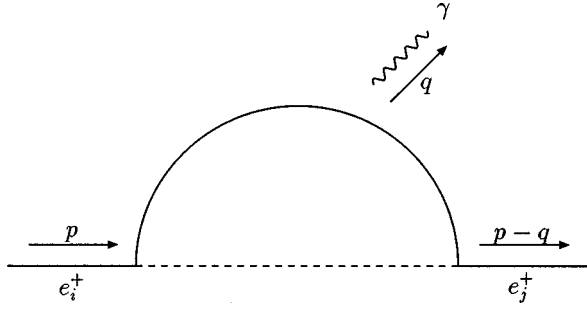


FIG. 17. Assignment of the momenta to the external leptons and the external photon in a lepton flavor violating diagram. An anticharged lepton e_i^+ going into the left vertex with momentum p is annihilated there, and a photon with an outgoing momentum q and an anticharged lepton e_j^+ with an outgoing momentum $p-q$ are emitted.

$$iI_N^M(m_1^2, m_2^2, \dots, m_N^2) = \int \frac{d^4k}{(\pi)^2} k^{2M} \prod_i^N \frac{1}{k^2 - m_i^2}. \quad (\text{B12})$$

This function I_N^M can be reduced to I_{N-1}^M or I_N^{M-1} by the following rules:

$$I_N^M(m_1^2, m_2^2, \dots, m_N^2) = D[I_{N-1}^M(m^2, m_3^2, \dots, m_N^2); m^2](m_1^2, m_2^2), \quad (\text{B13})$$

$$I_N^M(m_1^2, m_2^2, \dots, m_N^2) = m_N^2 I_N^{M-1}(m_1^2, m_2^2, \dots, m_N^2) + I_{N-1}^{M-1}(m_1^2, m_2^2, \dots, m_{N-1}^2), \quad (\text{B14})$$

and then, all I_N^M for $N > 2 + M$ can be derived from I_1^0 ,

$$I_1^0(m^2) = \left(1 - \log \frac{m^2}{\Lambda^2}\right) m^2, \quad (\text{B15})$$

$$(-1)^{L+M+N} D^L [I_N^M(\underbrace{m^2, m^2, \dots, m^2}_{N-P}, m_1^2, \dots, m_P^2); m^2](M_1^2, M_2^2, \dots, M_{L+1}^2) > 0. \quad (\text{B17})$$

Then, we can discuss about relative signs between several diagrams definitely by using I_N^M and the finite differences. Furthermore, since the mass dimension of I_N^M is $(4 + 2M - 2N)$, we can derive the following relation:

$$(2 + M - N) I_N^M(m_1^2, m_2^2, \dots, m_N^2) = \sum_{i=1}^N m_i^2 I_{N+1}^M(m_1^2, \dots, m_i^2, m_i^2, \dots, m_N^2) \quad (\text{B18})$$

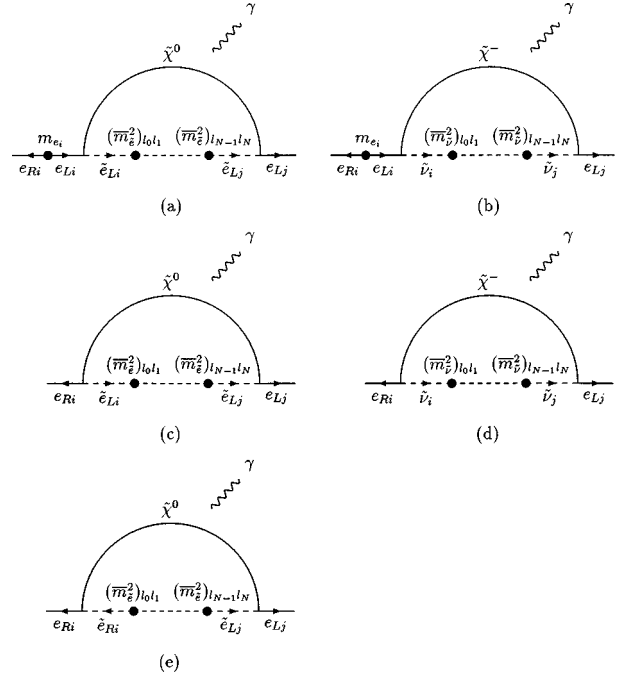


FIG. 18. Patterns of the chirality flips in the lepton flavor violating diagrams in the $e_i^+ \rightarrow e_j^+ \gamma$ decay ($i > j$). In the diagrams (a) and (b) the lepton chirality is flipped on the external lines, while in (c) and (d) it is flipped at a vertex of lepton-slepton-neutralino (-chargino). In (e) it flips on the internal slepton line. Chirality flip on the internal slepton line does not occur in the diagram with a virtual chargino because of the absence of the right-handed sneutrino at the low energy region.

where Λ is a renormalization point. Also, the signs of I_N^M and $D^L[I_N^M]$ are definitely determined as

$$(-1)^{N+M} I_N^M(m_1^2, m_2^2, \dots, m_N^2) > 0, \quad (\text{B16})$$

for $N > 2 + M$, since

$$\frac{d}{dx} \{x^{-2-M+N} I_N^M(xm_1^2, xm_2^2, \dots, xm_N^2)\} = 0. \quad (\text{B19})$$

Because of Eqs. (B13), (B14), (B18) there are some ways to represent one function, for example,

$$I_5^2(m^2, m^2, m^2, M^2, M^2) = -3m^2 I_5^1(m^2, m^2, m^2, m^2, M^2). \quad (\text{B20})$$

We will derive amplitudes of $e_i^+ \rightarrow e_j^+ \gamma$ ($i > j$) by the mass-insertion technique. The amplitude is generally written as

$$T = e \epsilon^{\alpha*} \bar{v}_i(p) i \sigma_{\alpha\beta} q^\beta (A_L^{(ij)} P_L + A_R^{(ij)} P_R) v_j(p-q). \quad (\text{B21})$$

Here, e is the electric charge, ϵ^* the photon polarization vector, v_i and v_j the wave functions for the external leptons. Assignment of momenta of the external fields is shown in Fig. 17. Since above term is violating the lepton chirality, it is convenient to decompose $A_L^{(ij)}$ and $A_R^{(ij)}$ as

$$\begin{aligned} A_L^{(ij)} &= A_L^{(ij)}|_{c1} + A_L^{(ij)}|_{c2} + A_L^{(ij)}|_{n1} + A_L^{(ij)}|_{n2} + A_L^{(ij)}|_{n3}, \\ A_R^{(ij)} &= A_R^{(ij)}|_{c1} + A_R^{(ij)}|_{c2} + A_R^{(ij)}|_{n1} + A_R^{(ij)}|_{n2} + A_R^{(ij)}|_{n3}. \end{aligned} \quad (\text{B22})$$

$A_{L,R}^{(ij)}|_{n1,c1}$ come from diagrams in which the lepton chirality is flipped on the external lines [Figs. 18(a) and 18(b)]. $A_{L,R}^{(ij)}|_{n2,c2}$ are contributions from the diagrams in which the chirality is flipped on a vertex of lepton-slepton-neutralino (-chargino) [Figs. 18(c) and 18(d)]. $A_{L,R}^{(ij)}|_{n3}$ are contributions from those with the chirality flip on the internal slepton line [Fig. 18(e)]. Subscripts c and n represent that they are from chargino and neutralino diagrams, respectively.

The neutralino contribution to $A_L^{(ij)}|_{n1}$, derived from a diagram where the chirality of lepton is flipped in the external lepton line, is given by

$$\begin{aligned} A_L^{(ij)}|_{n1} &= -\frac{1}{6(4\pi)^2} m_{e_i} N_{LL}^{A(i)*} N_{LL}^{A(j)} \sum_{N=1}^{\infty} \sum_{l_1, \dots, l_{N-1}} (\bar{m}_{\tilde{e}}^2)_{l_0 l_1} (\bar{m}_{\tilde{e}}^2)_{l_1 l_2} \cdots (\bar{m}_{\tilde{e}}^2)_{l_{N-1} l_N} \\ &\quad \times D^N [I_5^2(\bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, M_{\tilde{\chi}^0}^2, M_{\tilde{\chi}^0}^2), \bar{m}_{\tilde{e}}^2] (\bar{m}_{\tilde{e}_{l_0}}^2, \bar{m}_{\tilde{e}_{l_1}}^2, \dots, \bar{m}_{\tilde{e}_{l_N}}^2) \quad (l_0 = i, l_N = j), \\ A_R^{(ij)}|_{n1} &= (A_L^{(ij)}|_{n1})_{L \leftrightarrow R, l_0 = i+3, l_N = j+3}, \end{aligned} \quad (\text{B23})$$

where m_{e_i} is the i th generation charged lepton mass, and the explicit form of I_5^1 is

$$\begin{aligned} I_5^2(\bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, M_{\tilde{\chi}^0}^2, M_{\tilde{\chi}^0}^2) &= -\frac{1}{\bar{m}_{\tilde{e}}^2} \frac{1}{2(1-x)^4} (1-6x+3x^2 \\ &\quad + 2x^3 - 6x^2 \log x) \end{aligned} \quad (\text{B24})$$

with $x = M_{\tilde{\chi}^0}^2 / \bar{m}_{\tilde{e}}^2$. $N_{LL}^{A(i)}$ is a coupling constant between slepton and neutralino, and our definition of couplings of slepton to neutralino (or chargino) is

$$\begin{aligned} -\mathcal{L}_{\text{int}} &= \bar{\nu}_i^\dagger \bar{\chi}_A^0 (C_{LR}^{A(i)} P_R + C_{LL}^{A(i)} P_L) e_i + \bar{e}_i^\dagger \bar{\chi}_A^0 (N_{LR}^{A(i)} P_R \\ &\quad + N_{LL}^{A(i)} P_L) e_i + \bar{e}_i^\dagger \bar{\chi}_A^0 (N_{RR}^{A(i)} P_R + N_{RL}^{A(i)} P_L) e_i + \text{H.c.} \end{aligned} \quad (\text{B25})$$

Here, $C_{LL}^{A(i)}$, $N_{LL}^{A(i)}$, and $N_{RR}^{A(i)}$ correspond to the (lepton-

chirality conserving) gaugino interactions, and $C_{LR}^{A(i)}$, $N_{RL}^{A(i)}$, and $N_{LR}^{A(i)}$ are the (lepton-chirality violating) Higgsino interactions. These coupling constants are represented by the MSSM parameters in Appendix A.

Sign of the term with N off-diagonal inserted-masses in $A_L^{(ij)}|_{n1}$ is definitely determined by signs of coupling constants of slepton to neutralino and the inserted masses, and is independent of the diagonal slepton masses and neutralino masses since

$$\begin{aligned} (-1)^{N+1} D^N [I_5^2(\bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, M_{\tilde{\chi}^0}^2, M_{\tilde{\chi}^0}^2), \bar{m}_{\tilde{e}}^2] \\ \times (\bar{m}_{\tilde{e}_0}^2, \bar{m}_{\tilde{e}_1}^2, \dots, \bar{m}_{\tilde{e}_N}^2) > 0. \end{aligned} \quad (\text{B26})$$

A contribution from a neutralino diagram where the lepton chirality is flipped in the vertex of slepton-lepton-neutralino is given as

$$\begin{aligned} A_L^{(ij)}|_{n2} &= -\frac{1}{2(4\pi)^2} M_{\tilde{\chi}^0} N_{LR}^{A(i)*} N_{LL}^{A(j)} \sum_{N=1}^{\infty} \sum_{l_1, \dots, l_{N-1}} (\bar{m}_{\tilde{e}}^2)_{l_0 l_1} (\bar{m}_{\tilde{e}}^2)_{l_1 l_2} \cdots (\bar{m}_{\tilde{e}}^2)_{l_{N-1} l_N} \\ &\quad \times D^N [I_4^1(\bar{m}_{\tilde{e}}^2, \bar{m}_{\tilde{e}}^2, M_{\tilde{\chi}^0}^2, M_{\tilde{\chi}^0}^2); \bar{m}_{\tilde{e}}^2] (\bar{m}_{\tilde{e}_{l_0}}^2, \bar{m}_{\tilde{e}_{l_1}}^2, \dots, \bar{m}_{\tilde{e}_{l_N}}^2) \quad (l_0 = i, l_N = j), \\ A_R^{(ij)}|_{n2} &= (A_L^{(ij)}|_{n2})_{L \leftrightarrow R, l_0 = i+3, l_N = j+3}. \end{aligned} \quad (\text{B27})$$

The explicit form of I_4^1 in Eq. (B27) is

$$I_4^1(\bar{m}_e^2, \bar{m}_e^2, M_{\tilde{\chi}^0}^2, M_{\tilde{\chi}^0}^2) = -\frac{1}{\bar{m}_e^2} \frac{1}{(1-x)^3} (1-x^2 + 2x \log x) \quad (\text{B28})$$

with $x = M_{\tilde{\chi}^0}^2 / \bar{m}_e^2$, and the sign of $D^N[I_4^1]$ is $(-1)^{N+1}$ from Eq. (B17).

A contribution from a neutralino diagram where the lepton chirality is flipped in the internal slepton line is given as

$$\begin{aligned} A_L^{(ij)}|_{n3} &= -\frac{1}{2(4\pi)^2} M_{\tilde{\chi}_A^0} N_{RR}^{A(i)*} N_{LL}^{A(j)} \sum_{N=1}^{\infty} \sum_{l_1, \dots, l_{N-1}} (\bar{m}_e^2)_{l_0 l_1} (\bar{m}_e^2)_{l_1 l_2} \cdots (\bar{m}_e^2)_{l_{N-1} l_N} \\ &\quad \times D^N[I_4^1(\bar{m}_e^2, \bar{m}_e^2, M_{\tilde{\chi}_A^0}^2, M_{\tilde{\chi}_A^0}^2); \bar{m}_e^2] (\bar{m}_{e_{l_0}}^2, \bar{m}_{e_{l_1}}^2, \dots, \bar{m}_{e_{l_N}}^2) \quad (l_0 = i+3, l_N = j), \\ A_R^{(ij)}|_{n3} &= (A_L^{(ij)}|_{n3})_{L \rightarrow R, l_0 = i, l_N = j+3}. \end{aligned} \quad (\text{B29})$$

A contribution from a chargino diagram where the lepton chirality is flipped in the external lepton line is given as

$$\begin{aligned} A_L^{(ij)}|_{c1} &= \frac{1}{6(4\pi)^2} m_{e_i} C_{LL}^{A(i)*} C_{LL}^{A(j)} \sum_{N=1}^{\infty} \sum_{l_1, \dots, l_{N-1}} (\bar{m}_{\tilde{\nu}}^2)_{l_0 l_1} (\bar{m}_{\tilde{\nu}}^2)_{l_1 l_2} \cdots (\bar{m}_{\tilde{\nu}}^2)_{l_{N-1} l_N} \\ &\quad \times D^N[I_5^2(\bar{m}_{\tilde{\nu}}^2, \bar{m}_{\tilde{\nu}}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2); \bar{m}_{\tilde{\nu}}^2] (\bar{m}_{\tilde{\nu}_{l_0}}^2, \bar{m}_{\tilde{\nu}_{l_1}}^2, \dots, \bar{m}_{\tilde{\nu}_{l_N}}^2) \quad (l_0 = i, l_N = j), \\ A_R^{(ij)}|_{c1} &= \mathcal{O}(m_{e_j}). \end{aligned} \quad (\text{B30})$$

The explicit form of I_5^2 in Eq. (B30) is

$$I_5^2(\bar{m}_{\tilde{\nu}}^2, \bar{m}_{\tilde{\nu}}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2) = -\frac{1}{\bar{m}_{\tilde{\nu}}^2} \frac{1}{2(1-x)^4} (2 + 3x - 6x^2 + x^3 + 6x \log x) \quad (\text{B31})$$

with $x = M_{\tilde{\chi}_A^-}^2 / \bar{m}_{\tilde{\nu}}^2$, and the sign of $D^N[I_5^2]$ is $(-1)^{N+1}$.

A contribution from a chargino diagram where the lepton chirality is flipped in the vertex of lepton-sneutrino-chargino is given as

$$\begin{aligned} A_L^{(ij)}|_{c2} &= \frac{1}{(4\pi)^2} M_{\tilde{\chi}_A^-} C_{LR}^{A(i)*} C_{LL}^{A(j)} \sum_{N=1}^{\infty} \sum_{l_1, \dots, l_{N-1}} (\bar{m}_{\tilde{\nu}}^2)_{l_0 l_1} (\bar{m}_{\tilde{\nu}}^2)_{l_1 l_2} \cdots (\bar{m}_{\tilde{\nu}}^2)_{l_{N-1} l_N} \\ &\quad \times D^N[I_4^1(\bar{m}_{\tilde{\nu}}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2); \bar{m}_{\tilde{\nu}}^2] (\bar{m}_{\tilde{\nu}_{l_0}}^2, \bar{m}_{\tilde{\nu}_{l_1}}^2, \dots, \bar{m}_{\tilde{\nu}_{l_N}}^2) \quad (l_0 = i, l_N = j), \\ A_R^{(ij)}|_{c2} &= \mathcal{O}(m_{e_j}). \end{aligned} \quad (\text{B32})$$

The explicit form of I_4^1 in Eq. (B32) is

$$\begin{aligned} I_4^1(\bar{m}_{\tilde{\nu}}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2, M_{\tilde{\chi}_A^-}^2) &= \frac{1}{\bar{m}_{\tilde{\nu}}^2} \frac{1}{2(1-x)^3} \\ &\quad \times (3 - 4x + x^2 + 2 \log x) \end{aligned} \quad (\text{B33})$$

with $x = M_{\tilde{\chi}_A^-}^2 / \bar{m}_{\tilde{\nu}}^2$, and the sign of $D^N[I_4^1]$ is $(-1)^{N+1}$.

APPENDIX C: APPLICATION OF MASS INSERTION FORMULAS TO $\tau^+ \rightarrow \mu^+ \gamma$ AND $\mu^+ \rightarrow e^+ \gamma$

In this section we apply the formulas derived in the previous section to $\tau^+ \rightarrow \mu^+ \gamma$ and $\mu^+ \rightarrow e^+ \gamma$ processes. We neglect $m_\mu(m_e)$ in the calculation of $\tau^+ \rightarrow \mu^+ \gamma$ (μ^+

$\rightarrow e^+ \gamma$). First we consider $\tau^+ \rightarrow \mu^+ \gamma$. The whole contribution to $\tau^+ \rightarrow \mu^+ \gamma$ can be written as

$$\begin{aligned} A_L^{(\tau\mu)} &= A_L^{(\tau\mu)}|_{c1} + A_L^{(\tau\mu)}|_{c2} + A_L^{(\tau\mu)}|_{n1} + A_L^{(\tau\mu)}|_{n2} + A_L^{(\tau\mu)}|_{n3}, \\ A_R^{(\tau\mu)} &= A_R^{(\tau\mu)}|_{c1} + A_R^{(\tau\mu)}|_{c2} + A_R^{(\tau\mu)}|_{n1} + A_R^{(\tau\mu)}|_{n2} + A_R^{(\tau\mu)}|_{n3}, \end{aligned} \quad (\text{C1})$$

in the same notation as the previous section. In many models the dominant contributions to $\tau \rightarrow \mu \gamma$ are from the diagrams with single insertion of $(m_L^2)_{32}$, $(m_e^2)_{32}$, A_e^{32} , or A_e^{23} . Among these lepton flavor violating coupling constants only $(m_L^2)_{32}$ and $(m_e^2)_{32}$ are important for $\tan \beta \gtrsim 1$. We here show explicit expressions of the diagrams with single off-diagonal slepton mass matrix element insertion for $\tan \beta \gtrsim 1$. From the

formula we derived in the previous Appendix, the contributions from diagrams with the lepton chirality flipped in the external line [Figs. 18(a) and 18(b)] are given as

$$A_L^{(\tau\mu)}|_{c1} = -m_\tau \frac{1}{6} \frac{\alpha_2}{4\pi} (O_R)_{A1}^2 \left[\frac{(m_L^2)_{32}}{m_{\tilde{\nu}_\mu}^2 - m_{\tilde{\nu}_\tau}^2} \right] \left\{ \frac{1}{m_{\tilde{\nu}_\mu}^2} f_{c1}(x_{A\tilde{\nu}_\mu}) - \frac{1}{m_{\tilde{\nu}_\tau}^2} f_{c1}(x_{A\tilde{\nu}_\tau}) \right\}, \quad (C2)$$

$$A_R^{(\tau\mu)}|_{c1} = \mathcal{O}(m_\mu), \quad (C3)$$

$$A_L^{(\tau\mu)}|_{n1} = m_\tau \frac{1}{12} \frac{\alpha_2}{4\pi} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W)^2 \left[\frac{(m_L^2)_{32}}{m_{\tilde{\mu}_L}^2 - m_{\tilde{\tau}_L}^2} \right] \left\{ \frac{1}{m_{\tilde{\mu}_L}^2} f_{n1}(x_{A\tilde{\mu}_L}) - \frac{1}{m_{\tilde{\tau}_L}^2} f_{n1}(x_{A\tilde{\tau}_L}) \right\}, \quad (C4)$$

$$A_R^{(\tau\mu)}|_{n1} = m_\tau \frac{1}{3} \frac{\alpha_Y}{4\pi} (O_N)_{A1}^2 \left[\frac{(m_e^2)_{32}}{m_{\tilde{\mu}_R}^2 - m_{\tilde{\tau}_R}^2} \right] \left\{ \frac{1}{m_{\tilde{\mu}_R}^2} f_{n1}(x_{A\tilde{\mu}_R}) - \frac{1}{m_{\tilde{\tau}_R}^2} f_{n1}(x_{A\tilde{\tau}_R}) \right\}, \quad (C5)$$

where $x_{A\tilde{l}} = M_{\tilde{\chi}_A}^2 / m_{\tilde{l}}^2$ with $l = \nu_\mu, \nu_\tau, \mu_L, \tau_L, \mu_R,$ and τ_R and $\tilde{\chi}_A = \tilde{\chi}_A^-$ and $\tilde{\chi}_A^0$. The coefficient functions $f_{c1}(x)$ and $f_{n1}(x)$ are given as

$$f_{c1}(x) \equiv \frac{1}{2(1-x)^4} (2 + 3x - 6x^2 + x^3 + 6x \log x), \quad (C6)$$

$$f_{n1}(x) \equiv \frac{1}{2(1-x)^4} (1 - 6x + 3x^2 + 2x^3 - 6x^2 \log x), \quad (C7)$$

which are positive definite and monotonically decreasing. Following coefficient functions are also defined to be positive definite and monotonically decreasing.

The diagrams where the lepton chirality is flipped in the vertices give the following contributions:

$$A_L^{(\tau\mu)}|_{c2} = m_\tau \frac{\alpha_2}{4\pi} \frac{M_{\tilde{\chi}_A^-}}{\sqrt{2} m_W \cos \beta} (O_R)_{A1} (O_L)_{A2} \left[\frac{(m_L^2)_{32}}{m_{\tilde{\nu}_\mu}^2 - m_{\tilde{\nu}_\tau}^2} \right] \times \left\{ \frac{1}{m_{\tilde{\nu}_\mu}^2} f_{c2}(x_{A\tilde{\nu}_\mu}) - \frac{1}{m_{\tilde{\nu}_\tau}^2} f_{c2}(x_{A\tilde{\nu}_\tau}) \right\}, \quad (C8)$$

$$A_R^{(\tau\mu)}|_{c2} = \mathcal{O}(m_\mu), \quad (C9)$$

$$A_L^{(\tau\mu)}|_{n2} = -m_\tau \frac{1}{4} \frac{\alpha_2}{4\pi} \frac{M_{\tilde{\chi}_A^0}}{m_Z \cos \theta_W \cos \beta} (O_N)_{A3} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W) \left[\frac{(m_L^2)_{32}}{m_{\tilde{\mu}_L}^2 - m_{\tilde{\tau}_L}^2} \right] \left\{ \frac{1}{m_{\tilde{\mu}_L}^2} f_{n2}(x_{A\tilde{\mu}_L}) - \frac{1}{m_{\tilde{\tau}_L}^2} f_{n2}(x_{A\tilde{\tau}_L}) \right\}, \quad (C10)$$

$$A_R^{(\tau\mu)}|_{n2} = m_\tau \frac{1}{2} \frac{\alpha_Y}{4\pi} \frac{M_{\tilde{\chi}_A^0}}{m_Z \sin \theta_W \cos \beta} (O_N)_{A3} (O_N)_{A1} \left[\frac{(m_e^2)_{32}}{m_{\tilde{\mu}_R}^2 - m_{\tilde{\tau}_R}^2} \right] \left\{ \frac{1}{m_{\tilde{\mu}_R}^2} f_{n2}(x_{A\tilde{\mu}_R}) - \frac{1}{m_{\tilde{\tau}_R}^2} f_{n2}(x_{A\tilde{\tau}_R}) \right\}, \quad (C11)$$

where the functions $f_{c2}(x)$ and $f_{n2}(x)$ are defined as

$$f_{c2}(x) \equiv -\frac{1}{2(1-x)^3} (3 - 4x + x^2 + 2 \log x), \quad (C12)$$

$$f_{n2}(x) \equiv \frac{1}{(1-x)^3} (1 - x^2 + 2x \log x). \quad (C13)$$

Finally the contribution from the diagrams in which the lepton chirality is flipped on the internal slepton lines are

$$A_L^{(\tau\mu)}|_{n3} = -\frac{1}{4} \frac{\alpha_2}{4\pi} (O_N)_{A1} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W) \tan \theta_W (m_{LR}^2)_{33} (m_L^2)_{32} M_{\tilde{\chi}_A^0} \times \left\{ \frac{1}{m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\mu}_L}^2} f_{n2}(x_{A\tilde{\tau}_R}) + \frac{1}{m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\mu}_L}^2} f_{n2}(x_{A\tilde{\tau}_L}) + \frac{1}{m_{\tilde{\mu}_L}^2} \frac{1}{m_{\tilde{\mu}_L}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\mu}_L}^2 - m_{\tilde{\tau}_L}^2} f_{n2}(x_{A\tilde{\mu}_L}) \right\}, \quad (C14)$$

$$\begin{aligned}
A_R^{(\tau\mu)}|_{n3} = & -\frac{1}{2} \frac{\alpha_2}{4\pi} (O_N)_{A1} ((O_N)_{A2} \\
& + (O_N)_{A1} \tan \theta_W) \tan \theta_W (m_{LR}^2)_{33} (m_{\tilde{e}}^2)_{32} M_{\tilde{\chi}_A^0} \left\{ \frac{1}{m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\mu}_R}^2} f_{n2}(x_{A\tilde{\tau}_L}) \right. \\
& \left. + \frac{1}{m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\mu}_R}^2} f_{n2}(x_{A\tilde{\tau}_R}) + \frac{1}{m_{\tilde{\mu}_R}^2} \frac{1}{m_{\tilde{\mu}_R}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\mu}_R}^2 - m_{\tilde{\tau}_R}^2} f_{n2}(x_{A\tilde{\mu}_R}) \right\}. \quad (C15)
\end{aligned}$$

Next we present the mass insertion formula for $\mu^+ \rightarrow e^+ \gamma$. Here we assume mass degeneracy between selectron and smuon, then

$$m_{\tilde{\nu}_e} = m_{\tilde{\nu}_\mu} \equiv m_{\tilde{\nu}},$$

$$m_{\tilde{e}_L} = m_{\tilde{\mu}_L} \equiv m_{\tilde{l}_L},$$

$$m_{\tilde{e}_R} = m_{\tilde{\mu}_R} \equiv m_{\tilde{l}_R}.$$

The contributions from diagrams with the lepton chirality flipped in the external line [Figs. 18(a) and 18(b)] are given as

$$A_L^{(\mu e)}|_{c1} = m_\mu \frac{1}{6} \frac{\alpha_2}{4\pi} (O_R)_{A1}^2 \left\{ \left[(m_{\tilde{l}}^2)_{21} + \frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{m_{\tilde{\nu}}^2 - m_{\tilde{\nu}_\tau}^2} \right] \frac{1}{m_{\tilde{\nu}}^4} g_{c1}(x_{A\tilde{\nu}}) - \left[\frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{(m_{\tilde{\nu}}^2 - m_{\tilde{\nu}_\tau}^2)^2} \right] \frac{1}{m_{\tilde{\nu}_\tau}^2} f_{c1}(x_{A\tilde{\nu}_\tau}) \right\}, \quad (C16)$$

$$A_R^{(\mu e)}|_{c1} = \mathcal{O}(m_e), \quad (C17)$$

$$\begin{aligned}
A_L^{(\mu e)}|_{n1} = & -m_\mu \frac{1}{12} \frac{\alpha_2}{4\pi} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W)^2 \left\{ \left[(m_{\tilde{l}}^2)_{21} + \frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{m_{\tilde{l}_L}^2 - m_{\tilde{\tau}_L}^2} \right] \frac{1}{m_{\tilde{l}_L}^4} g_{n1}(x_{A\tilde{l}_L}) \right. \\
& \left. - \left[\frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{(m_{\tilde{l}_L}^2 - m_{\tilde{\tau}_L}^2)^2} \right] \frac{1}{m_{\tilde{\tau}_L}^2} f_{n1}(x_{A\tilde{\tau}_L}) \right\}, \quad (C18)
\end{aligned}$$

$$A_R^{(\mu e)}|_{n1} = -\frac{1}{3} m_\mu \frac{\alpha_Y}{4\pi} (O_N)_{A1}^2 \left\{ \left[(m_{\tilde{e}}^2)_{21} + \frac{(m_{\tilde{e}}^2)_{23}(m_{\tilde{e}}^2)_{31}}{m_{\tilde{l}_R}^2 - m_{\tilde{\tau}_R}^2} \right] \frac{1}{m_{\tilde{l}_R}^4} g_{n1}(x_{A\tilde{l}_R}) - \left[\frac{(m_{\tilde{e}}^2)_{23}(m_{\tilde{e}}^2)_{31}}{(m_{\tilde{l}_R}^2 - m_{\tilde{\tau}_R}^2)^2} \right] \frac{1}{m_{\tilde{\tau}_R}^2} f_{n1}(x_{A\tilde{\tau}_R}) \right\}, \quad (C19)$$

with

$$g_{c1}(x) \equiv f_{c1}(x) + x f'_{c1}(x),$$

$$g_{n1}(x) \equiv f_{n1}(x) + x f'_{n1}(x). \quad (C20)$$

Here primes on $f_{c1,n1}(x)$ mean differentiation about x .

The diagrams where the lepton chirality is flipped in the vertices give the following contributions:

$$A_L^{(\mu e)}|_{c2} = -m_\mu \frac{\alpha_2}{4\pi} \frac{M_{\tilde{\chi}_A^-}}{\sqrt{2} m_W \cos \beta} (O_R)_{A1} (O_L)_{A2} \left\{ \left[(m_{\tilde{l}}^2)_{21} + \frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{m_{\tilde{\nu}}^2 - m_{\tilde{\nu}_\tau}^2} \right] \frac{1}{m_{\tilde{\nu}}^4} g_{c2}(x_{A\tilde{\nu}}) - \left[\frac{(m_{\tilde{l}}^2)_{23}(m_{\tilde{l}}^2)_{31}}{(m_{\tilde{\nu}}^2 - m_{\tilde{\nu}_\tau}^2)^2} \right] \frac{1}{m_{\tilde{\nu}_\tau}^2} f_{c2}(x_{A\tilde{\nu}_\tau}) \right\}, \quad (C21)$$

$$A_R^{(\mu e)}|_{c2} = \mathcal{O}(m_e), \quad (C22)$$

$$\begin{aligned}
A_L^{(\mu e)}|_{n2} = & m_\mu \frac{1}{4} \frac{\alpha_2}{4\pi} \frac{M_{\tilde{\chi}_A^0}}{m_Z \cos \theta_W \cos \beta} (O_N)_{A3} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W) \left\{ \left[(m_{\tilde{L}}^2)_{21} + \frac{(m_{\tilde{L}}^2)_{23}(m_{\tilde{L}}^2)_{31}}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} \right] \frac{1}{m_{\tilde{L}}^2} g_{n2}(x_{A\tilde{L}}) \right. \\
& \left. - \left[\frac{(m_{\tilde{L}}^2)_{23}(m_{\tilde{L}}^2)_{31}}{(m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2)^2} \right] \frac{1}{m_{\tilde{\tau}_L}^2} f_{n2}(x_{A\tilde{\tau}_L}) \right\}, \tag{C23}
\end{aligned}$$

$$\begin{aligned}
A_R^{(\mu e)}|_{n2} = & -m_\mu \frac{1}{2} \frac{\alpha_Y}{4\pi} \frac{M_{\tilde{\chi}_A^0}}{m_Z \sin \theta_W \cos \beta} (O_N)_{A3} (O_N)_{A1} \left\{ \left[(m_{\tilde{e}}^2)_{21} + \frac{(m_{\tilde{e}}^2)_{23}(m_{\tilde{e}}^2)_{31}}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} \right] \frac{1}{m_{\tilde{L}}^2} g_{n2}(x_{A\tilde{I}_R}) \right. \\
& \left. - \left[\frac{(m_{\tilde{e}}^2)_{23}(m_{\tilde{e}}^2)_{31}}{(m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2)^2} \right] \frac{1}{m_{\tilde{\tau}_R}^2} f_{n2}(x_{A\tilde{\tau}_R}) \right\}. \tag{C24}
\end{aligned}$$

Here the functions $g_{c2}(x)$ and $g_{n2}(x)$ are defined, similarly to $g_{c1,n1}(x)$, as

$$\begin{aligned}
g_{c2}(x) & \equiv f_{c2}(x) + x f'_{c2}(x), \\
g_{n2}(x) & \equiv f_{n2}(x) + x f'_{n2}(x). \tag{C25}
\end{aligned}$$

The contributions from Fig. 18(e) in which the lepton chirality is flipped in the slepton lines are the following:

$$\begin{aligned}
A_L^{(\mu)}|_{n3} = & \frac{1}{2} \frac{\alpha_2}{4\pi} M_{\tilde{\chi}_A^0} (O_N)_{A1} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W) \tan \theta_W \\
& \times \left\{ - (m_{\tilde{e}}^2)_{23} (m_{LR}^2)_{33} (m_{\tilde{L}}^2)_{31} \left(\frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_L}^2} f_{n2}(x_{A\tilde{I}_R}) \right. \right. \\
& + \frac{1}{m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{L}}^2} f_{n2}(x_{A\tilde{\tau}_L}) + \frac{1}{m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{L}}^2} f_{n2}(x_{A\tilde{\tau}_R}) \\
& \left. \left. + \frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} f_{n2}(x_{A\tilde{I}_L}) \right) + (m_{LR}^2)_{22} (m_{\tilde{L}}^2)_{21} \left(- \frac{1}{m_{\tilde{L}}^2} \frac{1}{(m_{\tilde{L}}^2 - m_{\tilde{I}_L}^2)^2} f_{n2}(x_{A\tilde{I}_R}) \right. \right. \\
& + \frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_R}^2} g_{n2}(x_{A\tilde{I}_L}) \left. \left. + (m_{LR}^2)_{22} (m_{\tilde{L}}^2)_{23} (m_{\tilde{L}}^2)_{31} \left(- \frac{1}{m_{\tilde{L}}^2} \frac{1}{(m_{\tilde{L}}^2 - m_{\tilde{I}_L}^2)^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} f_{n2}(x_{A\tilde{I}_R}) \right. \right. \right. \\
& \left. \left. \left. - \frac{1}{m_{\tilde{\tau}_L}^2} \frac{1}{(m_{\tilde{\tau}_L}^2 - m_{\tilde{I}_L}^2)^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{I}_R}^2} f_{n2}(x_{A\tilde{\tau}_L}) + \frac{1}{m_{\tilde{I}_L}^2} \frac{1}{m_{\tilde{I}_L}^2 - m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{I}_L}^2 - m_{\tilde{\tau}_L}^2} g_{n2}(x_{A\tilde{I}_L}) \right) \right\}, \tag{C26}
\end{aligned}$$

$$\begin{aligned}
A_R^{(\mu e)}|_{n3} = & \frac{1}{2} \frac{\alpha_2}{4\pi} M_{\tilde{\chi}_A^0} (O_N)_{A1} ((O_N)_{A2} + (O_N)_{A1} \tan \theta_W) \tan \theta_W \\
& \times \left\{ - (m_{\tilde{L}}^2)_{23} (m_{LR}^2)_{33} (m_{\tilde{e}}^2)_{31} \left(\frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_R}^2} f_{n2}(x_{A\tilde{I}_L}) \right. \right. \\
& + \frac{1}{m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{I}_L}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{L}}^2} f_{n2}(x_{A\tilde{\tau}_R}) + \frac{1}{m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{I}_L}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{\tau}_R}^2} \frac{1}{m_{\tilde{\tau}_L}^2 - m_{\tilde{L}}^2} f_{n2}(x_{A\tilde{\tau}_L}) \\
& \left. \left. + \frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_L}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} f_{n2}(x_{A\tilde{I}_R}) \right) + (m_{LR}^2)_{22} (m_{\tilde{e}}^2)_{21} \left(- \frac{1}{m_{\tilde{L}}^2} \frac{1}{(m_{\tilde{L}}^2 - m_{\tilde{I}_L}^2)^2} f_{n2}(x_{A\tilde{I}_L}) \right. \right. \\
& + \frac{1}{m_{\tilde{L}}^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{I}_R}^2} g_{n2}(x_{A\tilde{I}_R}) \left. \left. + (m_{LR}^2)_{22} (m_{\tilde{e}}^2)_{23} (m_{\tilde{e}}^2)_{31} \left(- \frac{1}{m_{\tilde{L}}^2} \frac{1}{(m_{\tilde{L}}^2 - m_{\tilde{I}_L}^2)^2} \frac{1}{m_{\tilde{L}}^2 - m_{\tilde{\tau}_R}^2} f_{n2}(x_{A\tilde{I}_L}) \right. \right. \right. \\
& \left. \left. \left. - \frac{1}{m_{\tilde{\tau}_R}^2} \frac{1}{(m_{\tilde{\tau}_R}^2 - m_{\tilde{I}_R}^2)^2} \frac{1}{m_{\tilde{\tau}_R}^2 - m_{\tilde{I}_L}^2} f_{n2}(x_{A\tilde{\tau}_R}) + \frac{1}{m_{\tilde{I}_R}^2} \frac{1}{m_{\tilde{I}_R}^2 - m_{\tilde{I}_L}^2} \frac{1}{m_{\tilde{I}_R}^2 - m_{\tilde{\tau}_R}^2} g_{n2}(x_{A\tilde{I}_R}) \right) \right\}. \tag{C27}
\end{aligned}$$

APPENDIX D: RENORMALIZATION GROUP EQUATIONS

In this Appendix we show the one-loop level renormalization group equations (RGE's) relevant to our discussion in the text. In the equations below we use a shorthand notation, that is, for matrices in generation space A and B we define $\{A, B\}$ by

$$\{A, B\}_{ij} \equiv \sum_k (A_{ik}B_{kj} + B_{ik}A_{kj}). \quad (\text{D1})$$

1. RGE's for the MSSMRN model

First we devote ourselves to the MSSMRN model. The RGE's for the gauge coupling constants and the gaugino masses are unchanged from the MSSM since the right-handed neutrinos are singlet under the standard model gauge group. First we list the RGE's of the Yukawa coupling constants:

$$16\pi^2\mu \frac{d}{d\mu} f_{e_{ij}} = \left\{ -\frac{9}{5}g_1^2 - 3g_2^2 + 3\text{Tr}(f_d f_d^\dagger) + \text{Tr}(f_e f_e^\dagger) \right\} f_{e_{ij}} + 3(f_e f_e^\dagger f_e)_{ij} + (f_e f_e^\dagger f_\nu)_{ij}, \quad (\text{D2})$$

$$16\pi^2\mu \frac{d}{d\mu} f_{\nu_{ij}} = \left\{ -\frac{3}{5}g_1^2 - 3g_2^2 + 3\text{Tr}(f_u f_u^\dagger) + \text{Tr}(f_\nu f_\nu^\dagger) \right\} f_{\nu_{ij}} + 3(f_\nu f_\nu^\dagger f_\nu)_{ij} + (f_\nu f_\nu^\dagger f_e)_{ij}. \quad (\text{D3})$$

Next RGE's of the soft massess and A parameters:

$$16\pi^2\mu \frac{d}{d\mu} (m_L^2)_{ij} = -\left(\frac{6}{5}g_1^2 |M_1|^2 + 6g_2^2 |M_2|^2 \right) \delta_{ij} - \frac{3}{5}g_1^2 S \delta_{ij} + \{m_L^2, f_e^\dagger f_e + f_\nu^\dagger f_\nu\}_{ij} + 2(f_e^\dagger m_e^2 f_e + \tilde{m}_{h1}^2 f_e^\dagger f_e + A_e^\dagger A_e)_{ij} \\ + 2(f_\nu^\dagger m_\nu^2 f_\nu + \tilde{m}_{h2}^2 f_\nu^\dagger f_\nu + A_\nu^\dagger A_\nu)_{ij}, \quad (\text{D4})$$

$$16\pi^2\mu \frac{d}{d\mu} (m_e^2)_{ij} = -\frac{24}{5}g_1^2 |M_1|^2 \delta_{ij} + \frac{6}{5}g_1^2 S \delta_{ij} + 2\{m_e^2, f_e f_e^\dagger\}_{ij} + 4(f_e m_L^2 f_e^\dagger + \tilde{m}_{h1}^2 f_e f_e^\dagger + A_e A_e^\dagger)_{ij}, \quad (\text{D5})$$

$$16\pi^2\mu \frac{d}{d\mu} (m_\nu^2)_{ij} = 2\{m_\nu^2, f_\nu f_\nu^\dagger\}_{ij} + 4(f_\nu m_L^2 f_\nu^\dagger + \tilde{m}_{h2}^2 f_\nu f_\nu^\dagger + A_\nu A_\nu^\dagger)_{ij}, \quad (\text{D6})$$

$$16\pi^2\mu \frac{d}{d\mu} A_{e_{ij}} = \left\{ -\frac{9}{5}g_1^2 - 3g_2^2 + 3\text{Tr}(f_d^\dagger f_d) + \text{Tr}(f_e^\dagger f_e) \right\} A_{e_{ij}} + 2\left\{ -\frac{9}{5}g_1^2 M_1 - 3g_2^2 M_2 + 3\text{Tr}(f_d^\dagger A_d) + \text{Tr}(f_e^\dagger A_e) \right\} f_{e_{ij}} \\ + 4(f_e f_e^\dagger A_e)_{ij} + 5(A_e f_e^\dagger f_e)_{ij} + 2(f_e f_\nu^\dagger A_\nu)_{ij} + (A_e f_\nu^\dagger f_\nu)_{ij}, \quad (\text{D7})$$

$$16\pi^2\mu \frac{d}{d\mu} A_{\nu_{ij}} = \left\{ -\frac{3}{5}g_1^2 - 3g_2^2 + 3\text{Tr}(f_u^\dagger f_u) + \text{Tr}(f_\nu^\dagger f_\nu) \right\} A_{\nu_{ij}} + 2\left\{ -\frac{3}{5}g_1^2 M_1 - 3g_2^2 M_2 + 3\text{Tr}(f_u^\dagger A_u) + \text{Tr}(f_\nu^\dagger A_\nu) \right\} f_{\nu_{ij}} \\ + 4(f_\nu f_\nu^\dagger A_\nu)_{ij} + 5(A_\nu f_\nu^\dagger f_\nu)_{ij} + 2(f_\nu f_e^\dagger A_e)_{ij} + (A_\nu f_e^\dagger f_e)_{ij}, \quad (\text{D8})$$

where

$$S = \text{Tr}(m_Q^2 + m_d^2 - 2m_u^2 - m_L^2 + m_e^2) - \tilde{m}_{h1}^2 + \tilde{m}_{h2}^2. \quad (\text{D9})$$

Here g_1 is the U(1) gauge coupling constant in the GUT convention, which is related to the U(1)_Y gauge coupling constant g_Y by $g_Y^2 = (3/5)g_1^2$.

2. RGE's for the SU(5)RN model

Next we list the RGE's for the SU(5)RN model. The superpotential of the matter sector is

$$W = \frac{1}{4} f_{u_{ij}} \psi_i^{AB} \psi_j^{CD} H^E \epsilon_{ABCDE} + \sqrt{2} f_{d_{ij}} \psi_i^{AB} \phi_{jA} \bar{H}_B + f_{\nu_{ij}} \eta_i \phi_{jA} H^A + \frac{1}{2} M_{\eta_i \eta_j} \eta_i \eta_j,$$

and the soft SUSY breaking terms are

$$\begin{aligned}
-\mathcal{L}_{\text{SUSY breaking}} = & (m_{\tilde{\psi}}^2)_{ij} \tilde{\psi}_i^\dagger \tilde{\psi}_j + (m_{\tilde{\phi}}^2)_{ij} \tilde{\phi}_i^\dagger \tilde{\phi}_j + (m_{\tilde{\eta}}^2)_{ij} \tilde{\eta}_i^\dagger \tilde{\eta}_j + m_h^2 h^\dagger h + m_{\tilde{h}}^2 \tilde{h}^\dagger \tilde{h} \\
& + \left\{ \frac{1}{4} A_{u_{ij}} \tilde{\psi}_i \tilde{\psi}_j h + \sqrt{2} A_{d_{ij}} \tilde{\psi}_i \tilde{\phi}_j \tilde{h} + A_{\nu_{ij}} \tilde{\eta}_i \tilde{\phi}_j h + \text{H.c.} \right\} + \frac{1}{2} M_5 \lambda_{5L} \lambda_{5L} + \text{H.c.}
\end{aligned} \quad (\text{D10})$$

We denote the $SU(5)_{\text{GUT}}$ gauge coupling constant as g_5 , the $SU(5)_{\text{GUT}}$ gaugino λ_5 , and its soft Majorana mass M_5 . We neglect a coupling $\lambda H \Sigma \tilde{H}$, as stated in the text.

First is the RGE's for the dimensionless coupling constants:

$$16\pi^2 \mu \frac{d}{d\mu} g_5 = -3g_5^3, \quad (\text{D11})$$

$$16\pi^2 \mu \frac{d}{d\mu} f_{d_{ij}} = \left[-\frac{84}{5} g_5^2 + 4 \text{Tr}(f_{d_{ij}}^\dagger f_{d_{ij}}) \right] f_{d_{ij}} + 6(f_{d_{ij}} f_{d_{ij}}^\dagger)_{ij} + 3(f_{u_{ij}} f_{u_{ij}}^\dagger)_{ij} + (f_{d_{ij}} f_{d_{ij}}^\dagger)_{ij}, \quad (\text{D12})$$

$$16\pi^2 \mu \frac{d}{d\mu} f_{u_{ij}} = \left[-\frac{96}{5} g_5^2 + 3 \text{Tr}(f_{u_{ij}}^\dagger f_{u_{ij}}) + \text{Tr}(f_{\nu_{ij}}^\dagger f_{\nu_{ij}}) \right] f_{u_{ij}} + 6(f_{u_{ij}} f_{u_{ij}}^\dagger)_{ij} + 2(f_{d_{ij}} f_{d_{ij}}^\dagger)_{ij} + 2(f_{u_{ij}} f_{d_{ij}}^* f_{d_{ij}}^\dagger)_{ij}, \quad (\text{D13})$$

$$16\pi^2 \mu \frac{d}{d\mu} f_{\nu_{ij}} = \left[-\frac{48}{5} g_5^2 + 3 \text{Tr}(f_{u_{ij}}^\dagger f_{u_{ij}}) + \text{Tr}(f_{\nu_{ij}}^\dagger f_{\nu_{ij}}) \right] f_{\nu_{ij}} + 6(f_{\nu_{ij}} f_{\nu_{ij}}^\dagger)_{ij} + 4(f_{u_{ij}} f_{d_{ij}}^\dagger)_{ij}. \quad (\text{D14})$$

Next is for the soft masses:

$$16\pi^2 \mu \frac{d}{d\mu} M_5 = -6g_5^2 M_5, \quad (\text{D15})$$

$$\begin{aligned}
16\pi^2 \mu \frac{d}{d\mu} m_h^2 = & -\frac{96}{5} g_5^2 M_5^2 + 6 \text{Tr}(f_{u_{ij}} f_{u_{ij}}^\dagger) m_h^2 + 6 \text{Tr}(f_{u_{ij}} m_{\tilde{\psi}}^2 f_{u_{ij}}^\dagger + f_{u_{ij}}^\dagger m_{\tilde{\psi}}^{2T} f_{u_{ij}} + A_u A_u^\dagger) + 2 \text{Tr}(f_{u_{ij}} f_{\nu_{ij}}^\dagger) m_h^2 \\
& + 2 \text{Tr}(f_{\nu_{ij}} m_{\tilde{\phi}}^2 f_{\nu_{ij}}^\dagger + f_{\nu_{ij}}^\dagger m_{\tilde{\eta}}^2 f_{\nu_{ij}}^* + A_\nu A_\nu^\dagger),
\end{aligned} \quad (\text{D16})$$

$$16\pi^2 \mu \frac{d}{d\mu} m_{\tilde{h}}^2 = -\frac{96}{5} g_5^2 M_5^2 + 8 \text{Tr}(f_{d_{ij}} f_{d_{ij}}^\dagger) m_{\tilde{h}}^2 + 8 \text{Tr}(f_{d_{ij}} m_{\tilde{\phi}}^2 f_{d_{ij}}^\dagger + f_{d_{ij}}^\dagger m_{\tilde{\psi}}^{2T} f_{d_{ij}} + A_d A_d^\dagger), \quad (\text{D17})$$

$$\begin{aligned}
16\pi^2 \mu \frac{d}{d\mu} (m_{\tilde{\psi}}^2)_{ij} = & -\frac{144}{5} g_5^2 M_5^2 \delta_{ij} + \{m_{\tilde{\psi}}^2, 2f_{d_{ij}}^* f_{d_{ij}}^\dagger + 3f_{u_{ij}}^* f_{u_{ij}}^\dagger\}_{ij} + 4[(f_{d_{ij}}^* f_{d_{ij}}^\dagger)_{ij} m_{\tilde{h}}^2 + (f_{d_{ij}}^* m_{\tilde{\phi}}^{2T} f_{d_{ij}}^\dagger + A_d^* A_d^\dagger)_{ij}] \\
& + 6[(f_{u_{ij}}^* f_{u_{ij}}^\dagger)_{ij} m_{\tilde{h}}^2 + (f_{u_{ij}}^* m_{\tilde{\psi}}^{2T} f_{u_{ij}}^\dagger + A_u^* A_u^\dagger)_{ij}],
\end{aligned} \quad (\text{D18})$$

$$\begin{aligned}
16\pi^2 \mu \frac{d}{d\mu} (m_{\tilde{\phi}}^2)_{ij} = & -\frac{96}{5} g_5^2 M_5^2 \delta_{ij} + \{m_{\tilde{\phi}}^2, 4f_{d_{ij}}^\dagger f_{d_{ij}} + f_{\nu_{ij}}^\dagger f_{\nu_{ij}}\}_{ij} + 8[(f_{d_{ij}}^\dagger f_{d_{ij}})_{ij} m_{\tilde{h}}^2 + (f_{d_{ij}}^\dagger m_{\tilde{\psi}}^{2T} f_{d_{ij}} + A_d^\dagger A_d)_{ij}] \\
& + 2[(f_{\nu_{ij}}^\dagger f_{\nu_{ij}})_{ij} m_{\tilde{h}}^2 + (f_{\nu_{ij}}^\dagger m_{\tilde{\eta}}^{2T} f_{\nu_{ij}} + A_\nu^\dagger A_\nu)_{ij}],
\end{aligned} \quad (\text{D19})$$

$$16\pi^2 \mu \frac{d}{d\mu} (m_{\tilde{\eta}}^2)_{ij} = 5\{m_{\tilde{\eta}}^2, f_{\nu_{ij}}^* f_{\nu_{ij}}^\dagger\}_{ij} + 10[(f_{\nu_{ij}}^* f_{\nu_{ij}}^\dagger)_{ij} m_{\tilde{h}}^2 + (f_{\nu_{ij}}^* m_{\tilde{\phi}}^{2T} f_{\nu_{ij}}^\dagger + A_\nu^* A_\nu^\dagger)_{ij}]. \quad (\text{D20})$$

Finally A terms:

$$\begin{aligned}
16\pi^2 \mu \frac{d}{d\mu} A_{d_{ij}} = & \left[-\frac{84}{5} g_5^2 + 4 \text{Tr}(f_{d_{ij}}^\dagger f_{d_{ij}}) \right] A_{d_{ij}} + 2 \left[-\frac{84}{5} g_5^2 M_5 + 4 \text{Tr}(f_{d_{ij}}^\dagger A_{d_{ij}}) \right] f_{d_{ij}} + 10(f_{d_{ij}} f_{d_{ij}}^\dagger A_{d_{ij}})_{ij} + 3(f_{u_{ij}} f_{u_{ij}}^\dagger A_{d_{ij}})_{ij} + 8(A_{d_{ij}} f_{d_{ij}}^\dagger)_{ij} \\
& + 6(A_{u_{ij}} f_{u_{ij}}^\dagger)_{ij} + (A_{d_{ij}} f_{\nu_{ij}}^\dagger)_{ij} + 2(f_{d_{ij}} f_{\nu_{ij}}^\dagger A_{\nu_{ij}})_{ij},
\end{aligned} \quad (\text{D21})$$

$$\begin{aligned}
16\pi^2 \mu \frac{d}{d\mu} A_{u_{ij}} = & \left[-\frac{96}{5} g_5^2 + 3 \text{Tr}(f_{u_{ij}}^\dagger f_{u_{ij}}) + \text{Tr}(f_{\nu_{ij}}^\dagger f_{\nu_{ij}}) \right] A_{u_{ij}} + 2 \left[-\frac{96}{5} g_5^2 M_5 + 3 \text{Tr}(f_{u_{ij}}^\dagger A_{u_{ij}}) + \text{Tr}(f_{\nu_{ij}}^\dagger A_{\nu_{ij}}) \right] f_{u_{ij}} + 2(f_{d_{ij}} f_{d_{ij}}^\dagger A_{u_{ij}})_{ij} \\
& + 9(f_{u_{ij}} f_{u_{ij}}^\dagger A_{u_{ij}})_{ij} + 2(A_{u_{ij}} f_{d_{ij}}^* f_{d_{ij}}^\dagger)_{ij} + 4(A_{d_{ij}} f_{d_{ij}}^\dagger)_{ij} + 9(A_{u_{ij}} f_{u_{ij}}^\dagger)_{ij} + 4(f_{u_{ij}} f_{d_{ij}}^* A_{d_{ij}}^\dagger)_{ij},
\end{aligned} \quad (\text{D22})$$

$$\begin{aligned}
16\pi^2\mu\frac{d}{d\mu}A_{v_{ij}} = & \left[-\frac{48}{5}g_5^2 + 3\text{Tr}(f_u^\dagger f_u) + \text{Tr}(f_w^\dagger f_w) \right] A_{v_{ij}} + 2 \left[-\frac{48}{5}g_5^2 M_5 + 3\text{Tr}(f_u^\dagger A_u) + \text{Tr}(f_w^\dagger A_w) \right] f_{v_{ij}} + 7(f_w f_w^\dagger A_w)_{ij} \\
& + 4(A_w f_w^\dagger f_w)_{ij} + 11(A_w f_w^\dagger f_w)_{ij} + 8(f_w f_w^\dagger A_w)_{ij}.
\end{aligned} \tag{D23}$$

-
- [1] J. Ellis and D. V. Nanopoulos, *Phys. Lett.* **110B**, 44 (1982); I-Hsiu Lee, *ibid.* **138B**, 121 (1984); *Nucl. Phys.* **B246**, 120 (1984).
- [2] L. J. Hall, V. A. Kostelecky, and S. Raby, *Nucl. Phys.* **B267**, 415 (1986).
- [3] Super-Kamiokande Collaboration, Y. Fukuda *et al.*, *Phys. Rev. Lett.* **81**, 1562 (1998); T. Kajita, hep-ex/9810001.
- [4] K. S. Hirata *et al.*, *Phys. Lett. B* **205**, 416 (1988); **280**, 146 (1992); Y. Fukuda *et al.*, *ibid.* **335**, 237 (1994); D. Casper *et al.*, *Phys. Rev. Lett.* **66**, 2561 (1991); R. Becker-Szendy *et al.*, *Phys. Rev. D* **46**, 3720 (1992).
- [5] CHOOZ Collaboration, M. Apollonio *et al.*, *Phys. Lett. B* **420**, 397 (1998).
- [6] T. Yanagida, in *Proceedings of the Workshop on Unified Theory and Baryon Number of the Universe*, Tsukuba, Japan, 1979, edited by O. Sawada and A. Sugamoto (KEK, Tsukuba, 1979), p. 95; M. Gell-Mann, P. Ramond, and R. Slansky, in *Supergravity*, Proceedings of the Workshop, Stony Brook, New York, 1979, edited by P. van Nieuwenhuizen and D. Freedman (North-Holland, Amsterdam, 1979).
- [7] F. Borzumati and A. Masiero, *Phys. Rev. Lett.* **57**, 961 (1986).
- [8] J. Hisano, T. Moroi, K. Tobe, M. Yamaguchi, and T. Yanagida, *Phys. Lett. B* **357**, 579 (1995).
- [9] J. Hisano, T. Moroi, K. Tobe, and M. Yamaguchi, *Phys. Rev. D* **53**, 2442 (1996).
- [10] R. Davis, Jr., D. S. Harmer, and K. C. Hoffman, *Phys. Rev. Lett.* **20**, 1205 (1968); K. S. Hirata *et al.*, *ibid.* **63**, 16 (1989); **65**, 1297 (1990).
- [11] S. P. Mikheyev and A. Y. Smirnov, *Yad. Fiz.* **42**, 1441 (1985) [*Sov. J. Nucl. Phys.* **42**, 913 (1985)]; *Nuovo Cimento C* **9**, 17 (1986); L. Wolfenstein, *Phys. Rev. D* **17**, 2369 (1978).
- [12] B. Pontecorvo, *Zh. Eksp. Teor. Fiz.* **53**, 1717 (1967) [*Sov. Phys. JETP* **26**, 984 (1968)]; S. M. Bilenky and B. Pontecorvo, *Phys. Rep.* **41**, 225 (1978); V. Barger, R. J. N. Phillips, and K. Whisnant, *Phys. Rev. D* **24**, 538 (1981); S. L. Glashow and L. M. Krauss, *Phys. Lett. B* **190**, 199 (1987).
- [13] J. Bahcall, P. I. Krastev, and A. Yu. Smirnov, *Phys. Rev. D* **58**, 096016 (1998).
- [14] Particle Data Group, C. Caso *et al.*, *Eur. Phys. J. C* **3**, 1 (1998).
- [15] N. Arkani-Hamed, H.-C. Cheng, J. L. Feng, and L. J. Hall, *Phys. Rev. Lett.* **77**, 1937 (1996); *Nucl. Phys.* **B505**, 7 (1997).
- [16] J. Hisano, M. M. Nojiri, Y. Shimizu, and M. Tanaka, hep-ph/9808410.
- [17] S. M. Bilenky and C. Giunti, hep-ph/9802201, version 2.
- [18] Y. Kuno and Y. Okada, *Phys. Rev. Lett.* **77**, 434 (1996); Y. Kuno, A. Maki, and Y. Okada, *Phys. Rev. D* **55**, 2517 (1997); Y. Kuno, KEK-Report-97-59.
- [19] V. Barger, S. Pakvasa, T. J. Weiler, and K. Whisnant, *Phys. Lett. B* **437**, 107 (1998).
- [20] J. C. Sen, *Phys. Rev.* **113**, 679 (1959); K. W. Ford and J. G. Wills, *Nucl. Phys.* **35**, 295 (1962); R. Pla and J. Bernabeu, *Ann. Phys. (Leipzig)* **67**, 455 (1971); H. C. Chiang, E. Oset, T. S. Kosmal, A. Faessler, and J. D. Vergados, *Nucl. Phys.* **A559**, 526 (1993).
- [21] J. Bernabeu, E. Nardi, and D. Tommasini, *Nucl. Phys.* **B409**, 63 (1993), and references therein.
- [22] MECO Collaboration, Proposal to Brookhaven National Laboratory AGS (1997).
- [23] R. Barbieri and L. J. Hall, *Phys. Lett. B* **338**, 212 (1994).
- [24] R. Barbieri, L. Hall, and A. Strumia, *Nucl. Phys.* **B445**, 219 (1995).
- [25] J. Hisano, T. Moroi, K. Tobe, and M. Yamaguchi, *Phys. Lett. B* **391**, 341 (1997); **397**, 357(E) (1997).
- [26] N. Arkani-Hamed, H. C. Cheng, and L. J. Hall, *Phys. Rev. D* **53**, 413 (1996).
- [27] J. Hisano, D. Nomura, Y. Okada, Y. Shimizu, and M. Tanaka, *Phys. Rev. D* **58**, 116010 (1998).
- [28] J. Hisano, D. Nomura, and T. Yanagida, *Phys. Lett. B* **437**, 351 (1998).
- [29] H. E. Haber and G. L. Kane, *Phys. Rep.* **117**, 75 (1985).